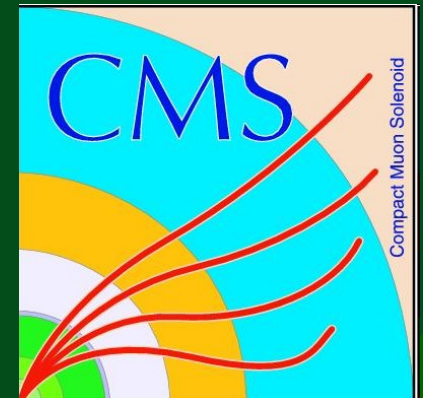


Searches: discovery techniques and/or limits in CMS



Amnon Harel

aharel@fnal.gov



PhysStat 2011

CERN, Geneva, Switzerland

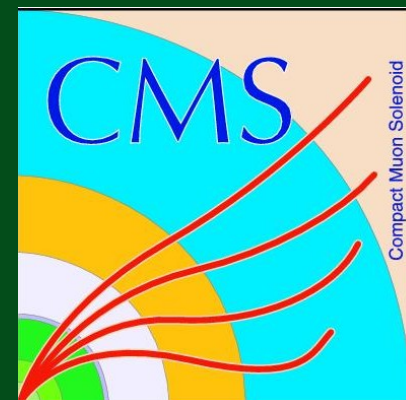
January 17th, 2011

Statistical methods in CMS searches to date



Amnon Harel

aharel@fnal.gov



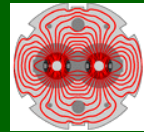
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Experimental Apparatus

The Large Hadron Collider

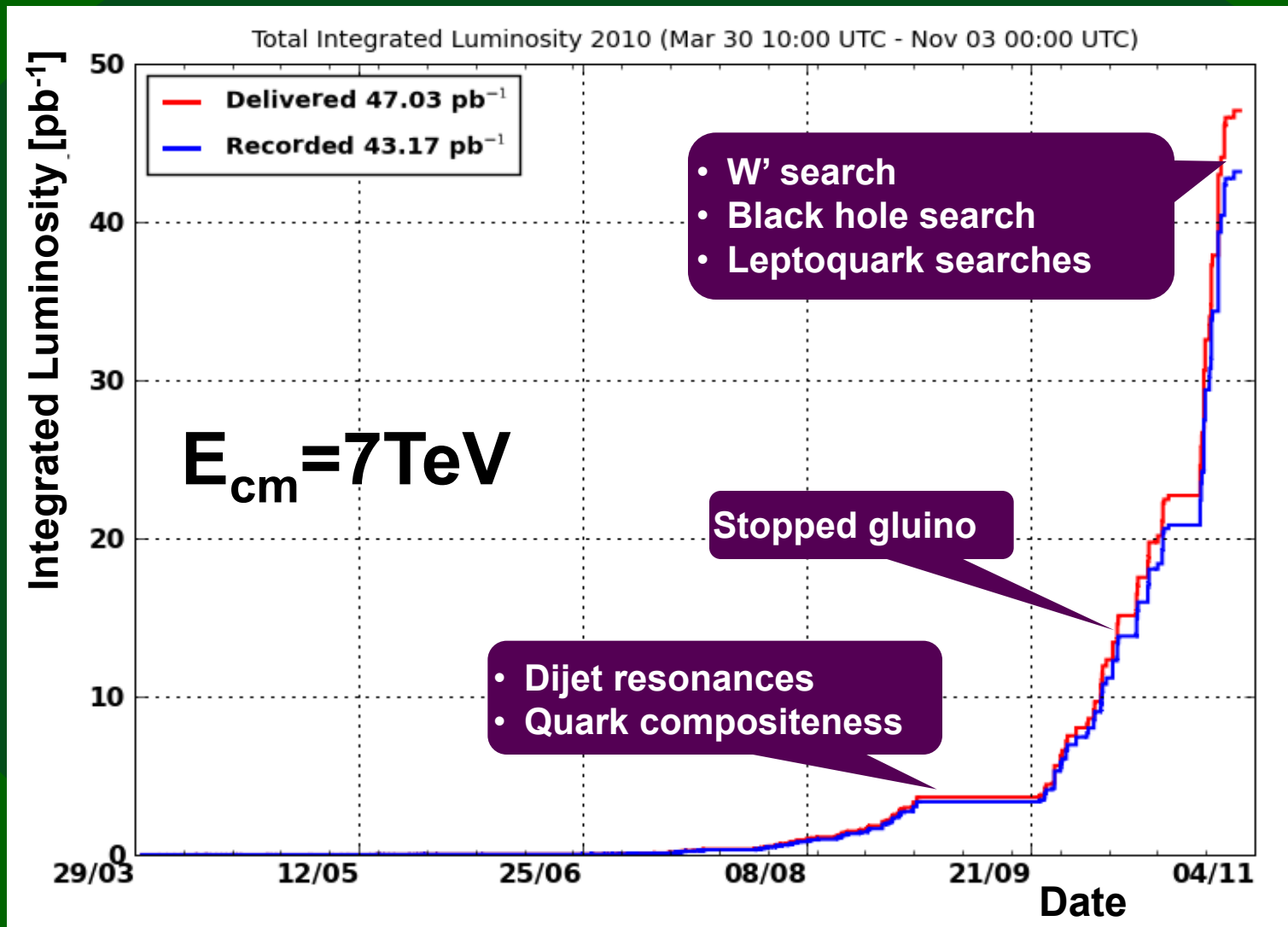


The collisions for today's talk

- pp
- $E_{c.m.} = 7\text{TeV}$
 - Energies up to 2TeV previously explored at Tevatron
 - Design: 14TeV
 - 1st collisions: 30 of March, 2010

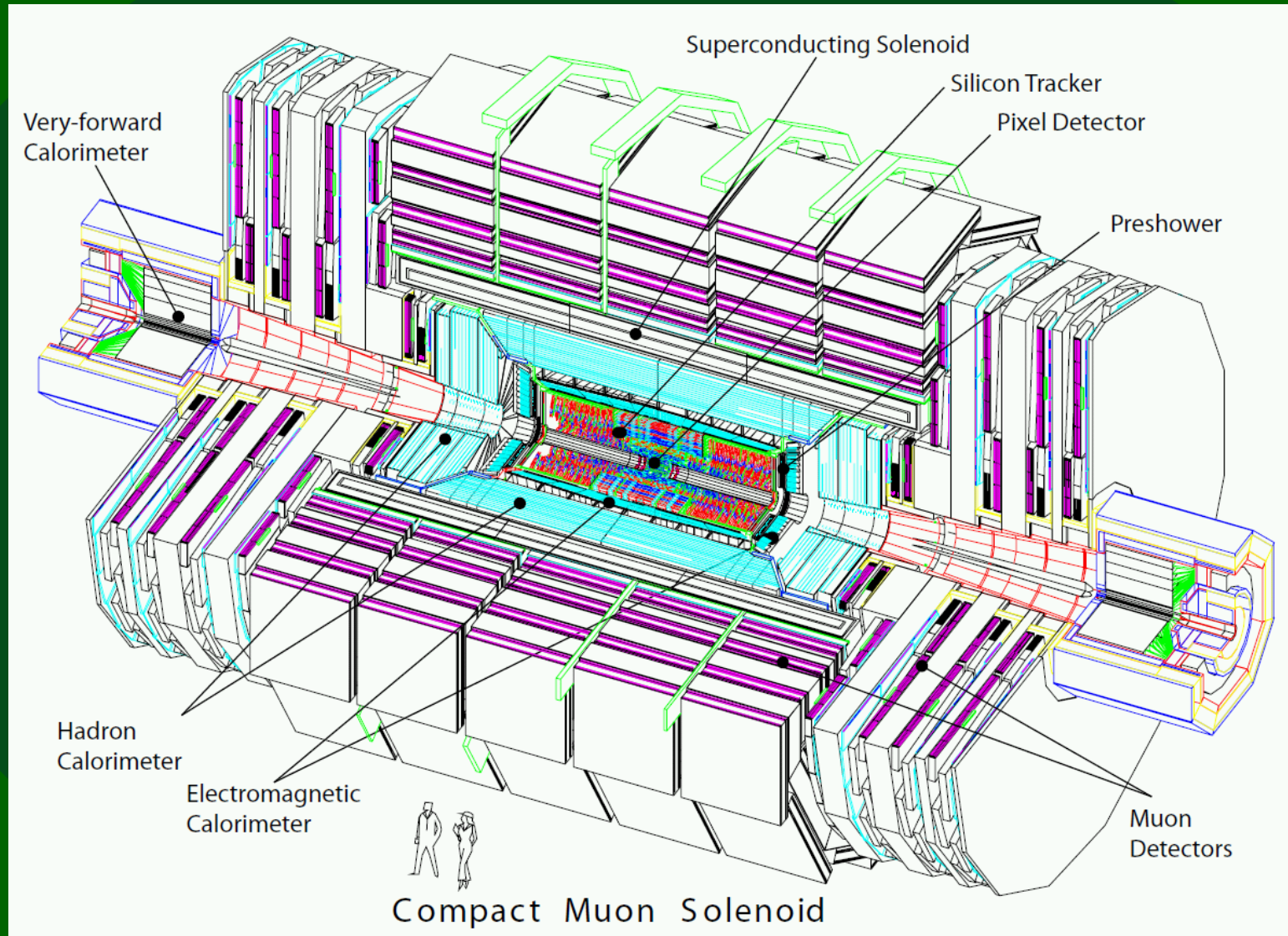


LHC performance



Peak luminosity: $L = 205 \mu\text{b}^{-1} \text{ s}^{-1}$

The CMS Detector



Design emphasizes tracking and EM calorimetry

The basic tools

How do we construct intervals?

Nuisance parameters

Bayesian approach

Integrated (marginalization)

Neyman construction
(frequentist, inversion of hypothesis test)

Multidimensional construction natural
But in practice, use a hybrid approach
(Cousins Highland) and integrate

Likelihood ratios

Minimization (Profiling, “MINOS”)

We'll discuss power constraints and other approaches later on.

The experimentalist perspective

Claiming a discovery first is the best case scenario.

But claiming a discovery is also the worst case scenario if you got it wrong.

Which of these statistical tools helps us get it right?

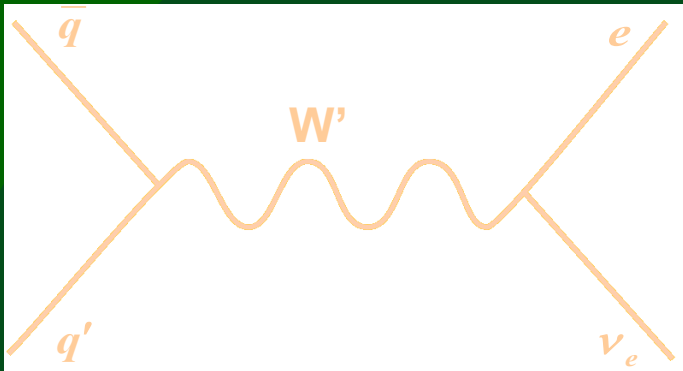
→ A “pragmatic” approach is typical.

No standard approach. Yet.

Bayesian

W' search

Experimental signature



- Isolated electron

- p_T imbalance (done with particle-flow energies, E_T^{miss})

Reference model

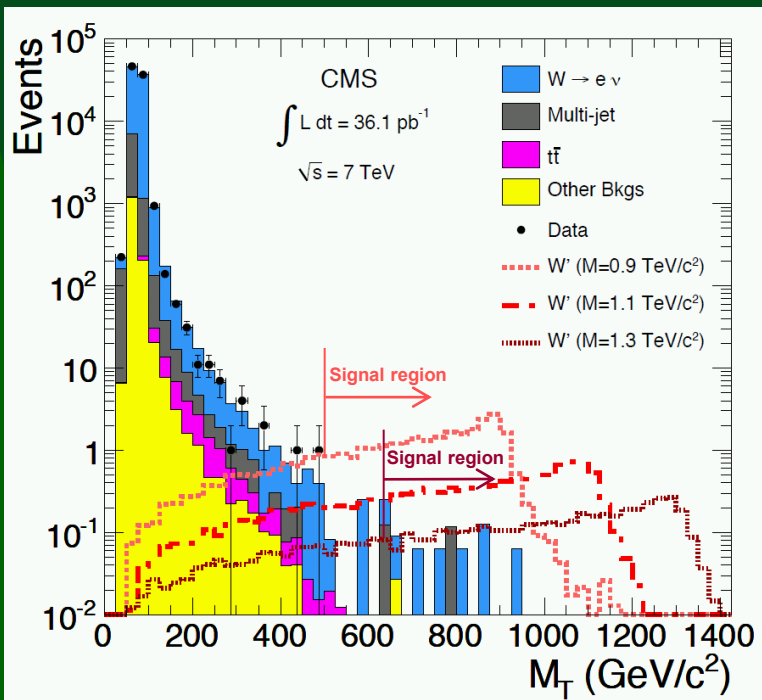
- W' has W-like fermionic couplings
- W' does not couple to other gauge bosons
- Tevatron limits: $m_{W'} > 1.1 \text{ TeV}$
- signal from Pythia, cross section scaled to NNLO

Observable

Number of events with $M_T > \text{threshold}$

$$\bullet (M_T)^2 = 2E_T^e E_T^{\text{miss}} (1 - \cos \Phi_{e E_T^{\text{miss}}})$$

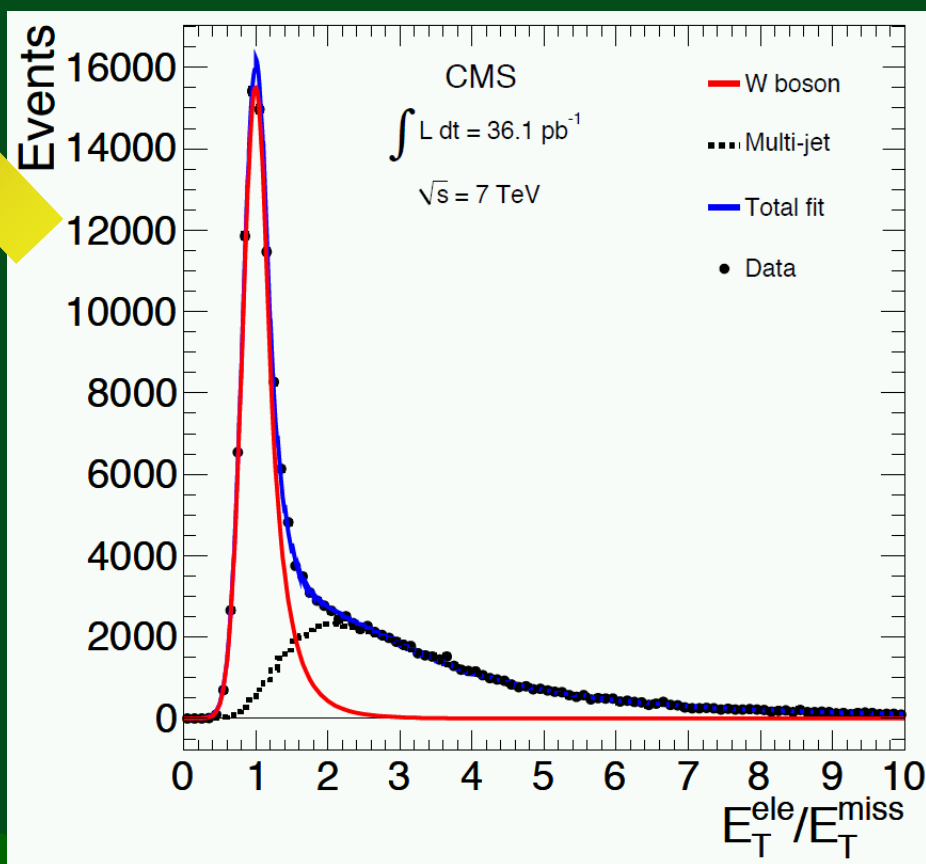
- Threshold chosen a-priori as a function of $m_{W'}$,



W' backgrounds

Estimating the main backgrounds:

- The two main backgrounds are W+jets and multijet production
- Data-driven techniques to estimate their shapes (unfortunately, beyond talk's scope)
- Normalizations from a fit to the (other-background subtracted) data



W' limit setting

The simplest scenario, as far as limit setting goes:

- A counting experiment (Poisson probability in each M_T bin)
- No interference between backgrounds and signal
- Systematic uncertainties factorize easily

$$N_{pred} = b + \mathcal{L} \epsilon \sigma_{eff}$$

Integrated luminosity
 \downarrow
 \uparrow
 Selection efficiency
 (for that bin)

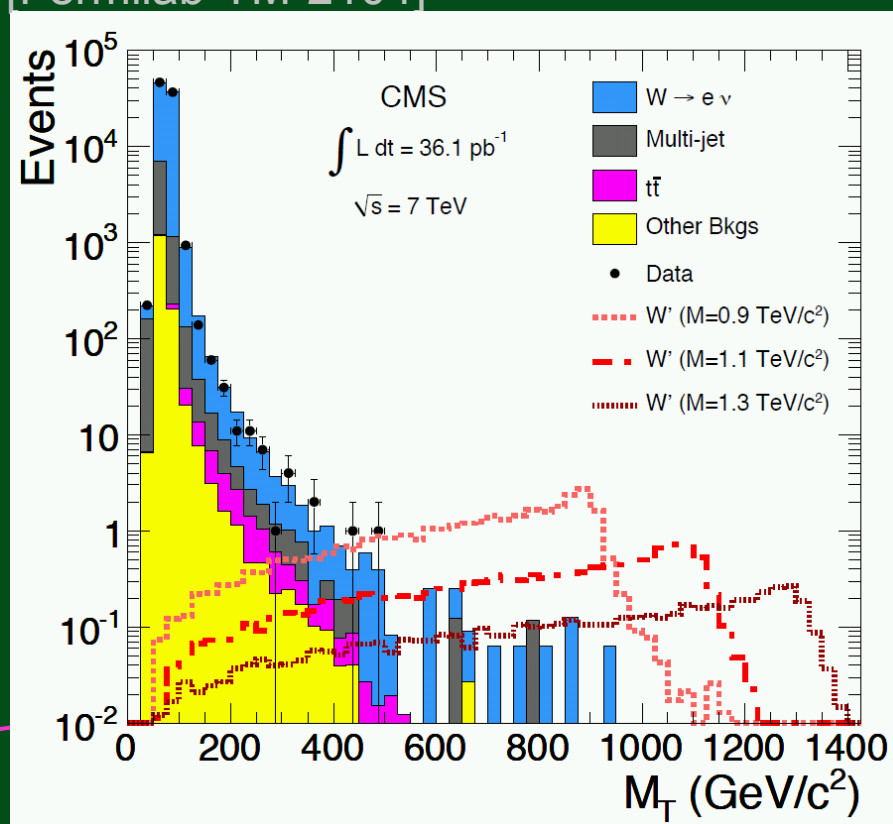
Use a simple Bayesian procedure [Fermilab-TM-2104]

- nuisance parameters are integrated out
- priors:

$$p(\sigma_{eff}) = \begin{cases} \text{const} & \text{if } 0 < \sigma_{eff} < \sigma_{max} \\ 0 & \text{otherwise} \end{cases}$$

- log normal priors for the nuisance parameters b, \mathcal{L}, ϵ
 - background uncertainty (e.g. fit results) summarized in one number
 - typical approximation

Rule out a W' with mass below 1.36 TeV at 95% CL



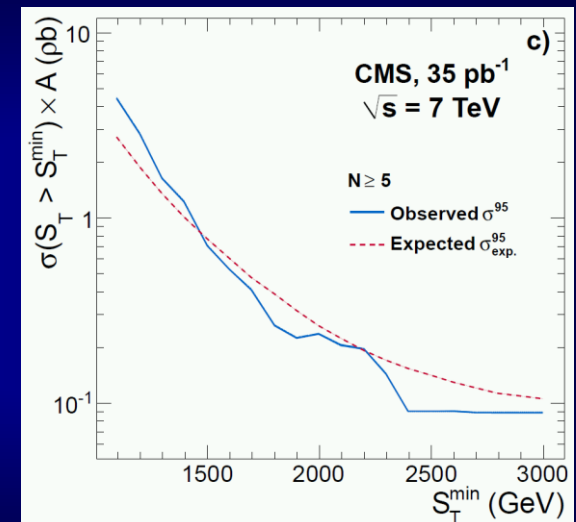
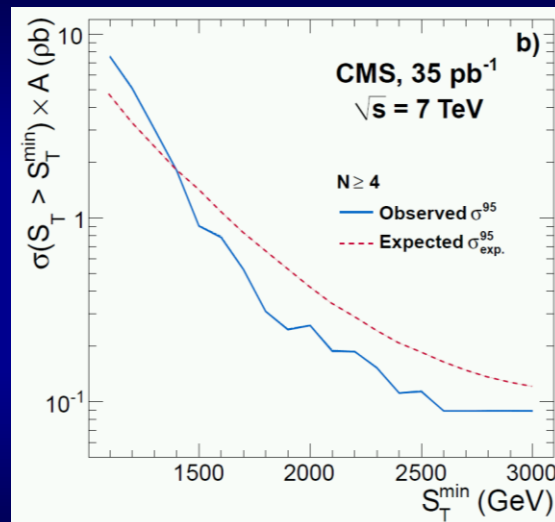
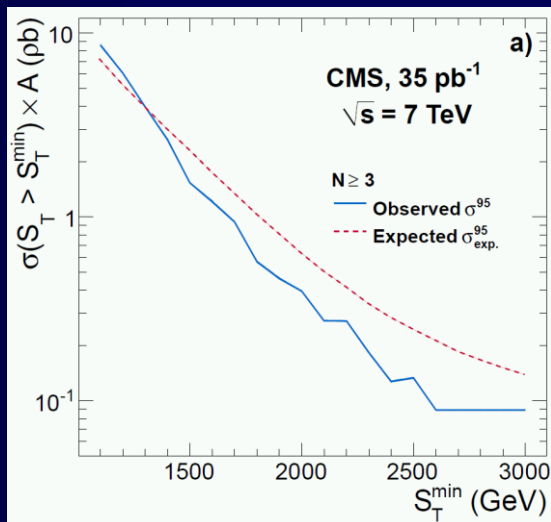
Covering all Bayeses

CMS searched for:

- 1st generation leptoquarks in the two electron + two jet final state
- 2nd generation leptoquarks in the two muon + two jet final state
- Microscopic black holes [see also]

With statistical treatments identical to that in the W' search

The black hole search contains also model independent limits:



Only 10% worse, as one object multiplicity (N) dominates the signal models

Dijet resonance search

Models

- dijet resonances common in new physics models
- eight specific models studies in paper
- signal models for qq , qg and gg final states



→ also model-independent limits

Experimental signature - the two jets...

- Single jet trigger required
- $|\eta_1, \eta_2| < 2.5$ and $|\Delta\eta| < 1.3$

Observable

Event counts as a function of dijet mass

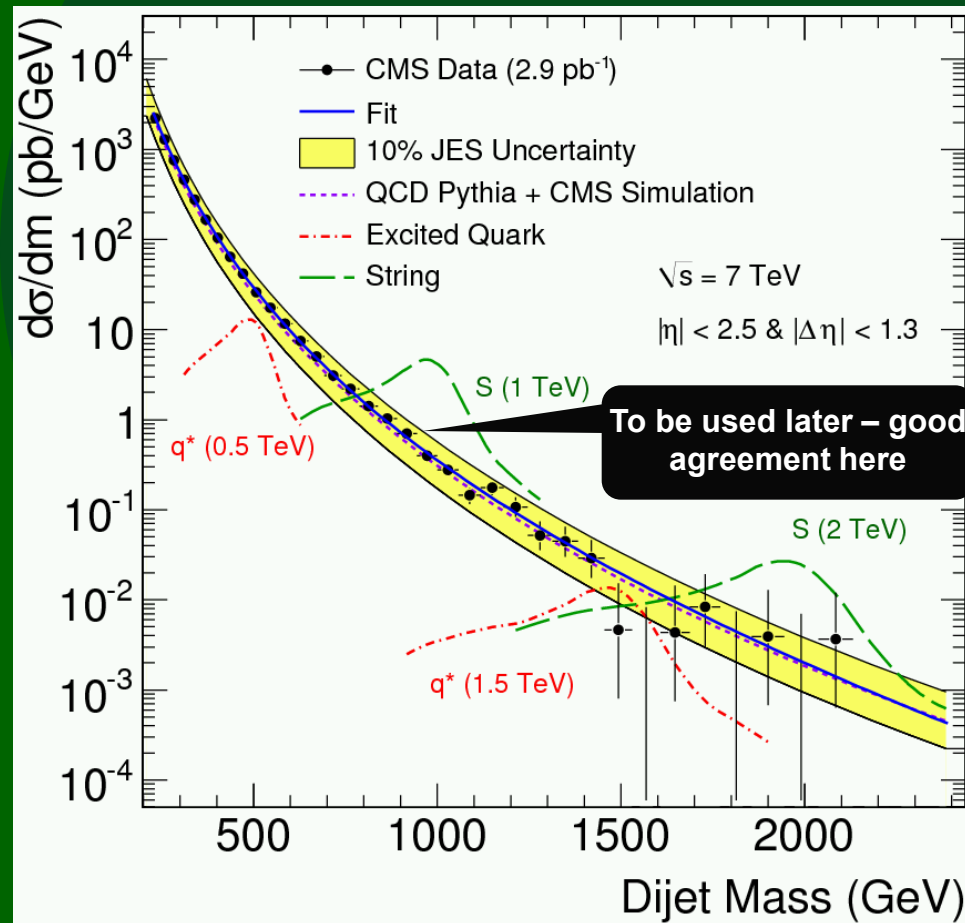
- binning predefined. Width \approx resolution.

Background estimate

By fit. Best is with 4 parameter function:

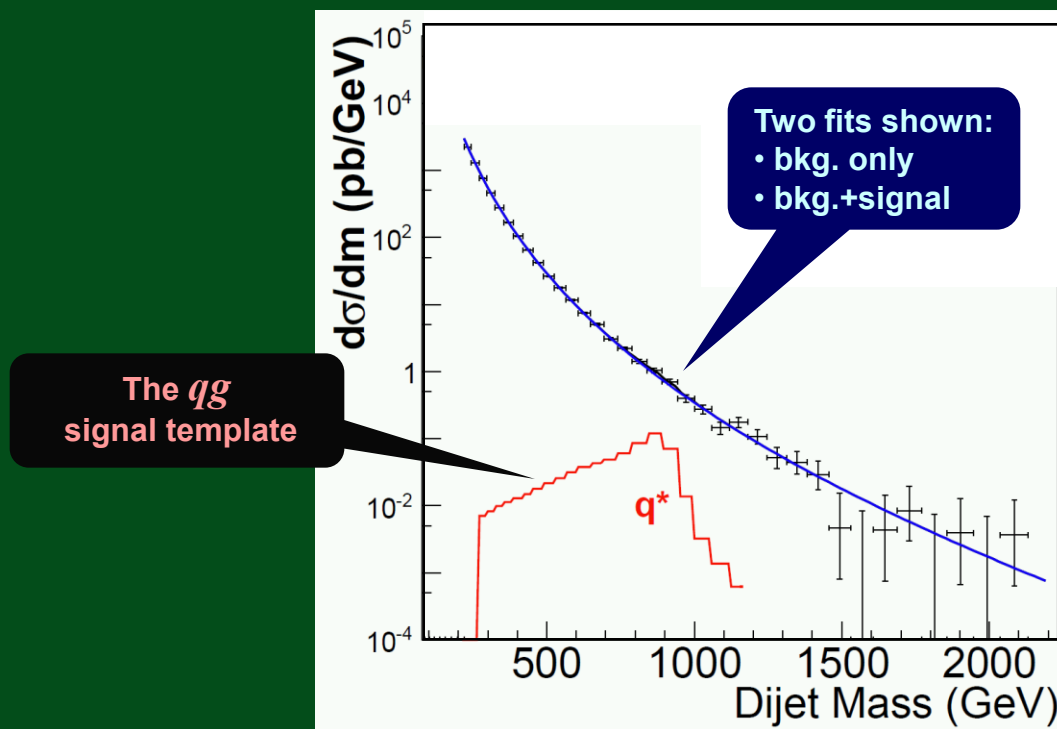
$$\frac{d\sigma}{dm} = p_0 \frac{\left(1 - \frac{m}{\sqrt{s}}\right)^{p_1}}{\left(\frac{m}{\sqrt{s}}\right)^{p_2 + p_3} \ln\left(\frac{m}{\sqrt{s}}\right)}$$

- “best” is out of 3 predefined functional forms
- $\chi^2 / \text{ndof} = 32 / 31$



Resonance significance

The JES uncertainty band may visually overstate the data – fit agreement, due to bin-to-bin correlations → quantify the biggest discrepancy in 0.5 – 2.0 TeV



- Fluctuation at mass of 900 GeV has local significance of 1.7σ from LLR ($\Delta\chi^2$)
- Significance reduced to 0.2σ accounting for “look elsewhere effect.”
 - p-value calibrated using ensemble tests
- No resonance observed → proceeding to set limits

Limit setting & systematics

Statistics-only Bayesian limit setting

- A counting experiment (Poisson probabilities)
- No interference between backgrounds and signal

For each m_{res} :

• For each m_{jj} bin:

$$N_{pred} = b + \mathcal{L} \epsilon \sigma_{eff}$$

Integrated luminosity \downarrow
 \mathcal{L}
 Signal template \uparrow
 ϵ

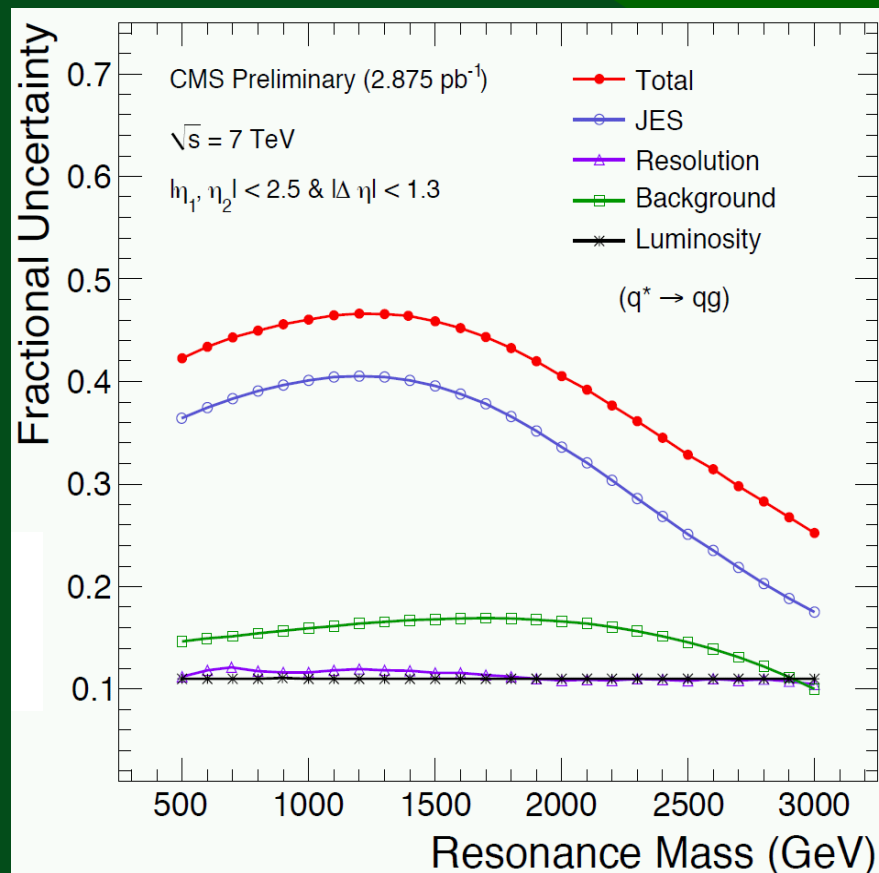
b from background + signal fits

• Uniform prior for the cross section:

$$p(\sigma_{eff}) = \begin{cases} \text{const} & \text{if } 0 < \sigma_{eff} < \sigma_{max} \\ 0 & \text{otherwise} \end{cases}$$

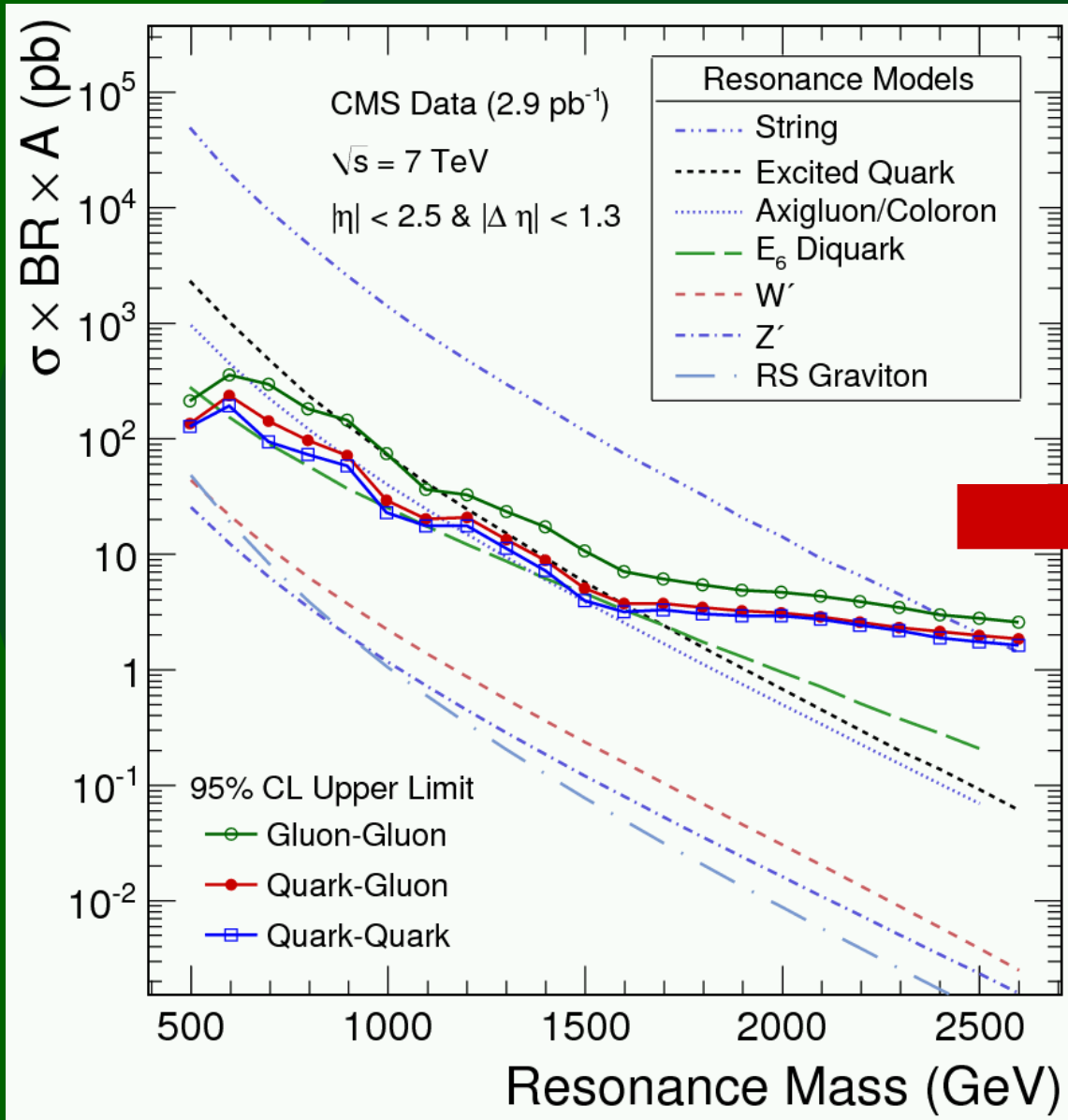
Incorporate systematic uncertainties at each mass by smearing the posterior $p(\sigma)$ with a Gaussian

- Approximate, but here, also conservative
- Verified frequentist coverage at 1TeV for σ =limit value
 - Without systematics, coverage ~95%
 - With JES systematic, coverage >98%



- Systematic uncertainties increase cross section limits by 15-50% depending on resonance mass and parton content
- Mass limits decrease by ~10% with inclusion of systematic uncertainties

Results



Model	Exclusion regions [TeV]
String resonance	0.50 – 2.50
Excited quark	0.50 – 1.58
Axigluon / Coloron	0.50 – 1.17 & 1.47 – 1.52
E_6 diquark	0.50-0.58, 0.97-1.08, & 1.45-1.60

CL_s

The CLs method

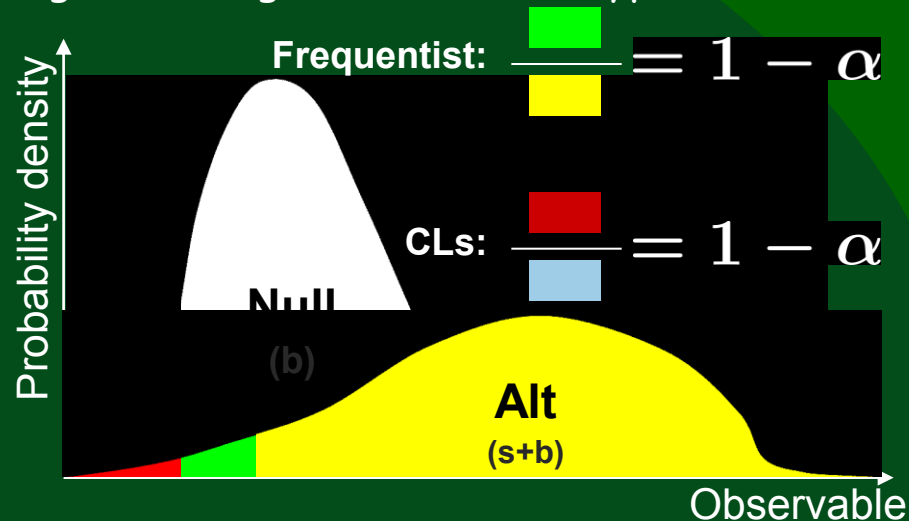
The CLs method is to exclude regions of phase space where



where α is the desired confidence level

- an LLR observable is recommended
- sometimes a statistics-only LLR used

E.g. excluding an alternative hypothesis

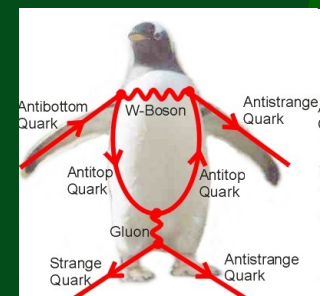


Criticism

- “What a horrible name!” Still, it’s just a name
- Fundamentally unsound – neither frequentist nor Bayesian

But

- Corresponds to frequentist limits when experiment is fully sensitive
- (Partially?) satisfies the real need for power constraints
 - graceful degradation
 - has been producing sensible results for over a decade

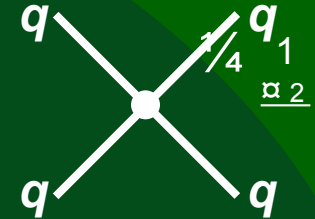


Quark compositeness

Model

- Quark compositeness will first appear as a contact interaction
- Model independent turn-on, details can vary at high energies

Contact interaction

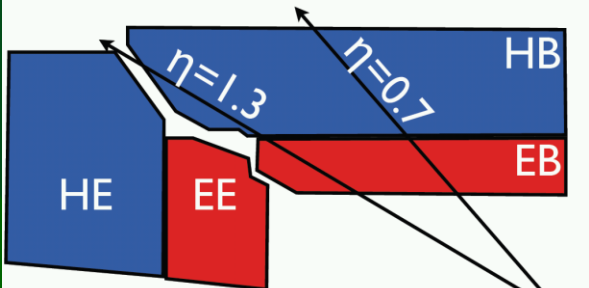


Experimental signature

Enhanced central dijet production

Observable

$$R_\eta = \frac{N(|\eta| < 0.7)}{N(0.7 < |\eta| < 1.3)} = \frac{N(\text{inner})}{N(\text{outer})}^*$$

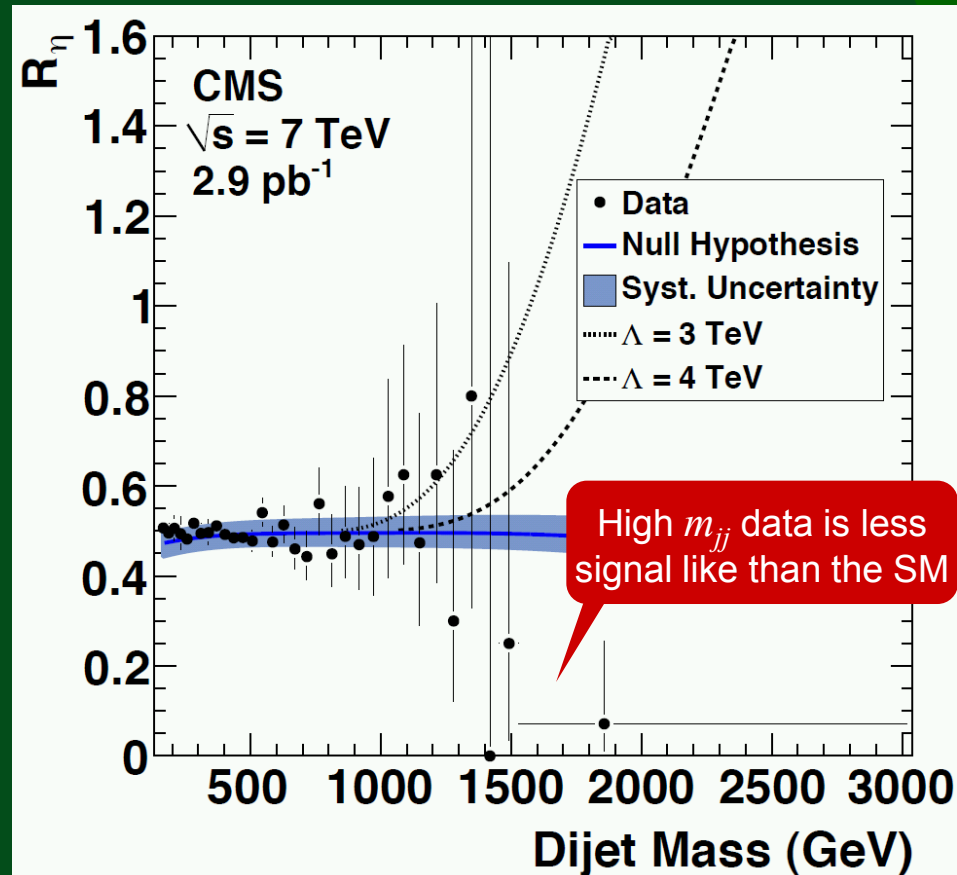


* Both jets inner or outer.

- “Ratio of Poisson means” problem
- m_{jj} binning as dijet resonance search

Background estimate

NLO calculations + NP corrections + a shift from a fit to data in a low m_{jj} region



R_η and SM

Shift from the normalization is:

Fit to data gave an offset of:

[Redacted]

- Two sided p-value (from ensemble testing with full systematic variations) is 0.29

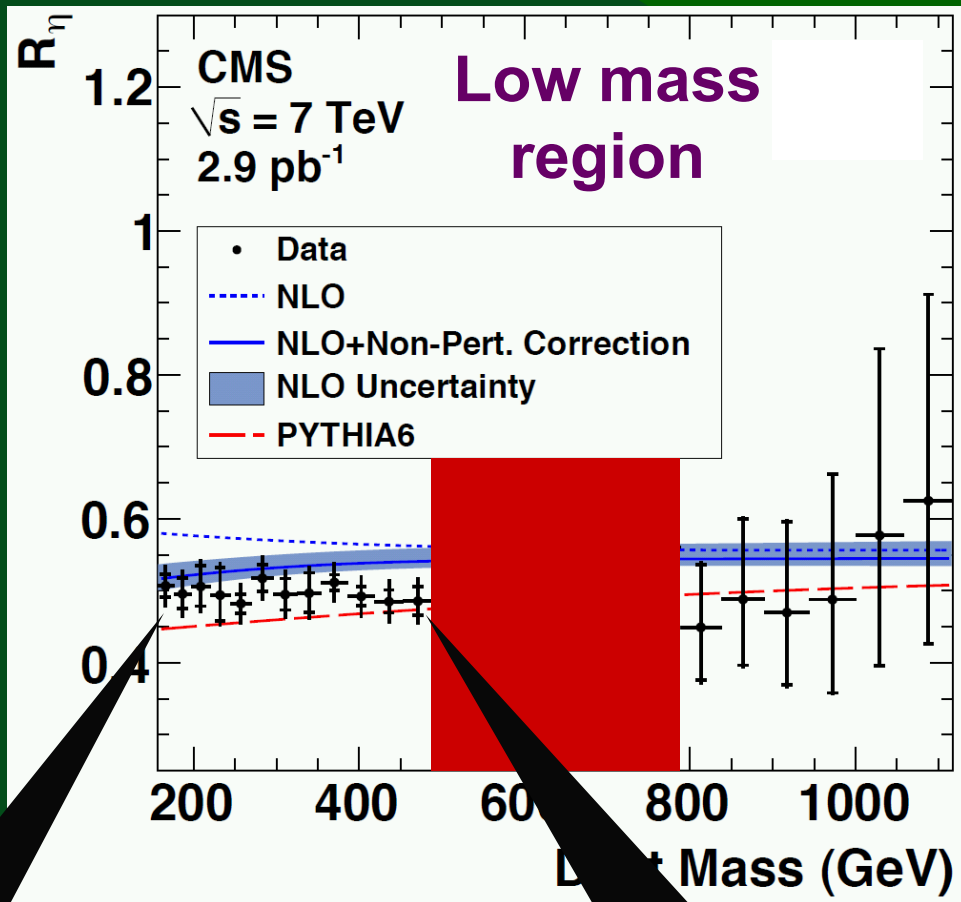
Overall, data consistent with SM

Offset for entire range:

$$-0.037 \pm 0.007(\text{stat.}) \pm 0.039(\text{syst.})$$

- Two sided p-value is 0.34

→ Proceeding to set limits on Δ



Vertical line:
total uncertainty

Horizontal tick:
statistical uncertainty

Note: Hard to estimate consistency with this visualization

Implementing CL_s



All probabilities conditioned on the observed total (inner + outer) number of events for that m_{jj} bin

- Standard and extremely useful treatment of “Ratio of Poisson means” problem

[Przyborowski and Wilenski, “Homogeneity of results in testing samples from Poisson series with an application to testing clover seeds for dodder”, Biometrika 31 (1940)]

Test statistics is the log likelihood ratio for SM and SM+CI hypotheses

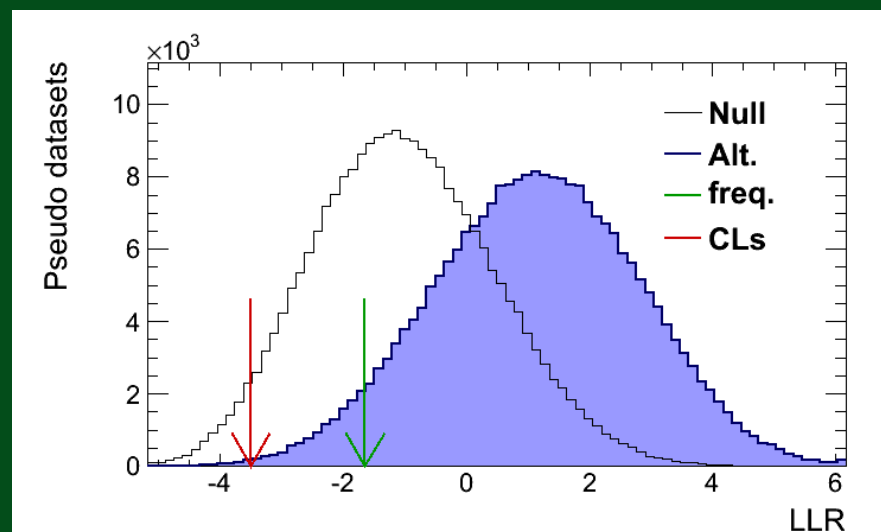
$$\mathcal{R}_{LL} = \ln \left(\frac{\mathcal{L}_{SM + \text{signal}}}{\mathcal{L}_{SM}} \right)$$

- Each Λ value evaluated separately
- Brute force integration of nuisance parameters (i.e., ensemble tests)
- Low R_η at high masses \rightarrow Low CL_b
 \rightarrow need to examine extreme tails of CL_{s+b}

A painful combination!

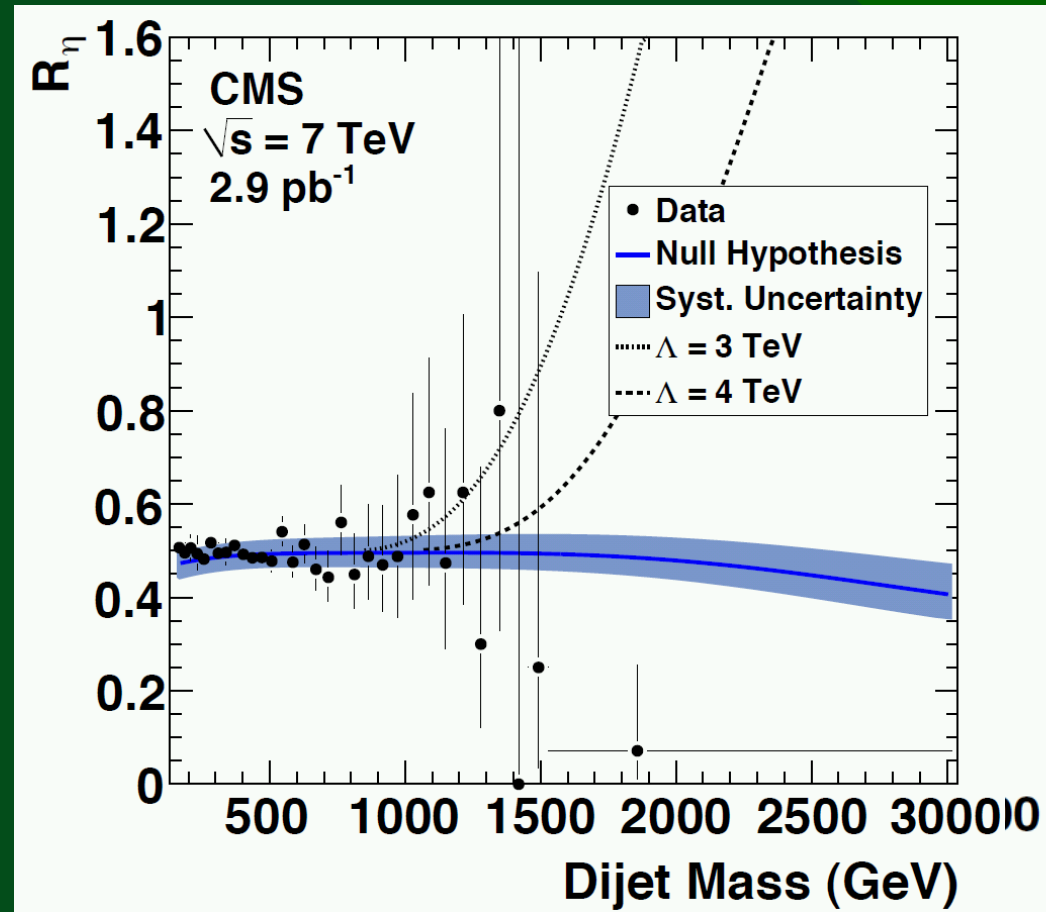
Stopping conditions:

- Λ value included/excluded at 2σ level
- CL_s value known at 0.5% accuracy



Limits

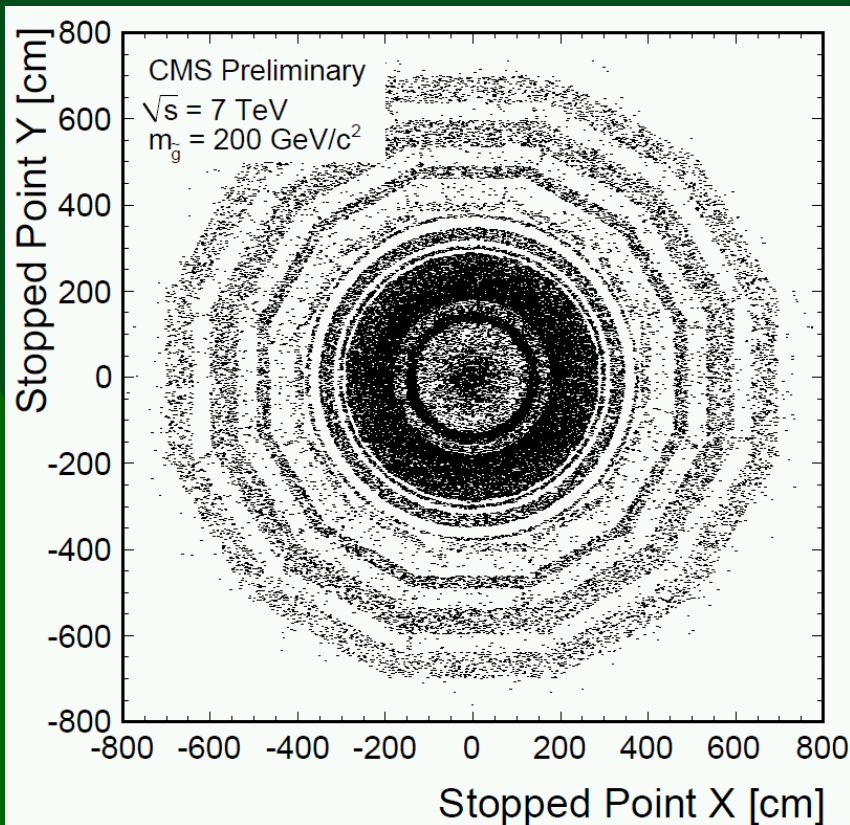
- Expected 95% CL exclusion is $\Lambda < 2.9 \text{ TeV}$
- Exclude all models with $\Lambda < 4 \text{ TeV}$ at 95% CL
 - Relevant tail probability (CL_{s+b}) is smaller than $1-\alpha$ by a factor of >100



Stopped gluinos

Model

- New, heavy, quasi-stable particles
 - In particular, split SUSY
- Searching for charged R-hadrons
 - Lifetime: 75ns – 10⁶s, m_{gluino} : 150-500 GeV
 - Up to 10⁶s, since still expect to see an event



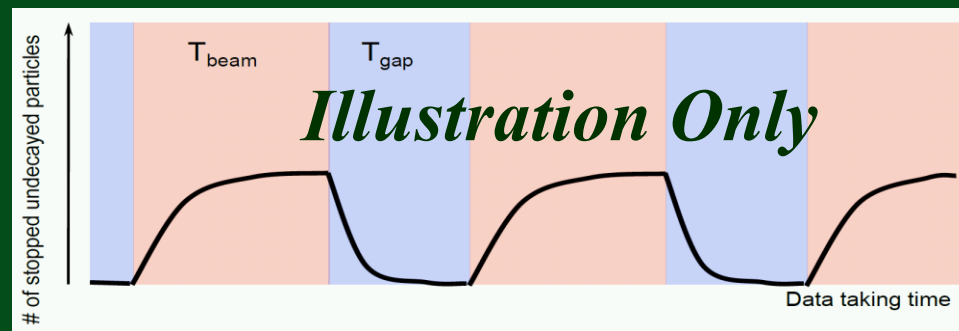
Experimental signature

Out of time energy deposits in calorimeters
+ vetoes on cosmic rays, beam halo, detector noise

Observable

Number of events with a lifetime window

- from 50ns to 1.256 · τ_{gluino}
- after each bunch crossing / fill



SM-background rate

For each of the two strongest cuts:

$$R_{\text{search}}^{\text{after}} = R_{\text{search}}^{\text{before}} \frac{R_{\text{control}}^{\text{after}}}{R_{\text{control}}^{\text{before}}}$$

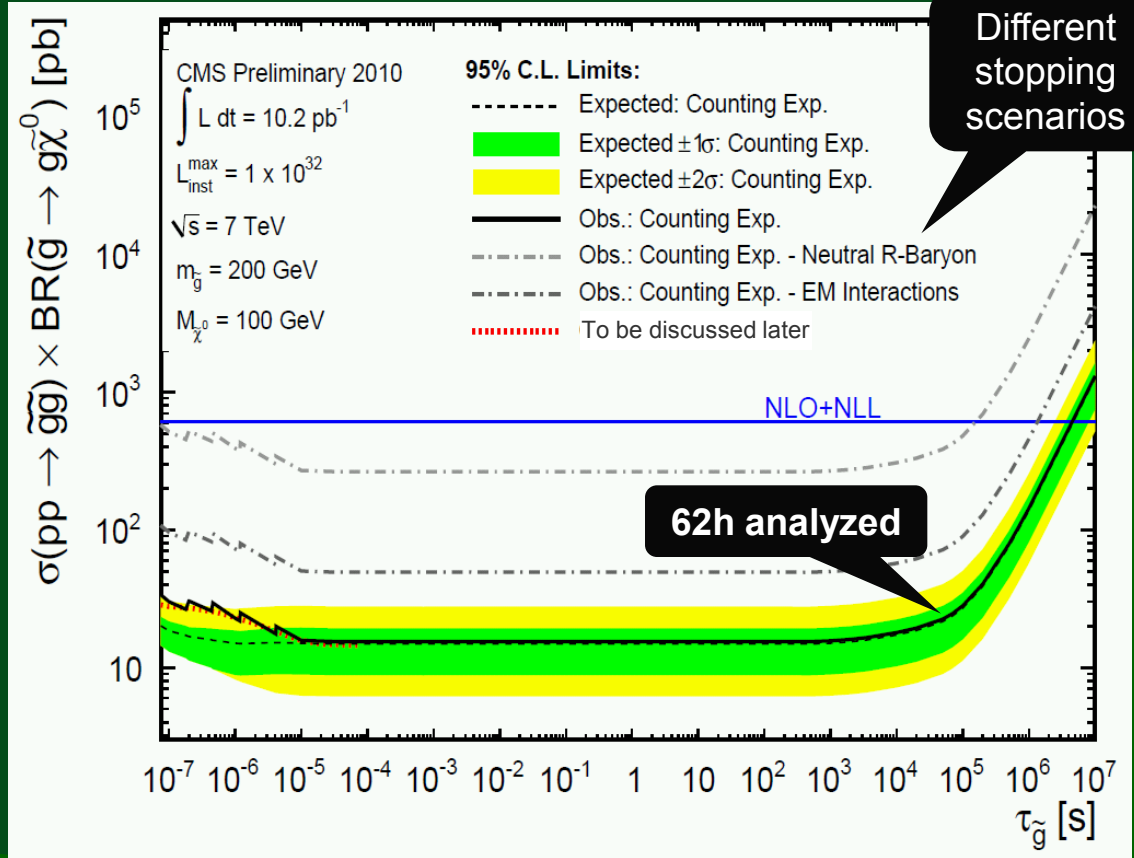
- strongest $\rightarrow R^{\text{after}} \ll R^{\text{before}}$
- Uncertainties from stability in time, timing simulation, integrated luminosity

CLs limits

Lifetime [s]	Expected background (\pm stat. \pm syst.)	Observed
10^{-7}	$0.8 \pm 0.2 \pm 0.2$	2
10^{-6}	$1.9 \pm 0.4 \pm 0.5$	3
$\geq 10^{-5}$	$4.9 \pm 1.0 \pm 1.3$	5

No indication of signal
 → proceed to set limits

- CLs method
- Nuisance parameters integrated out
 - Background rate 23%
 - Integrated luminosity 11%
 - JES 7%



Back to Bayes

Stopped gluinos – time profile

2nd Observable

The time of the selected events

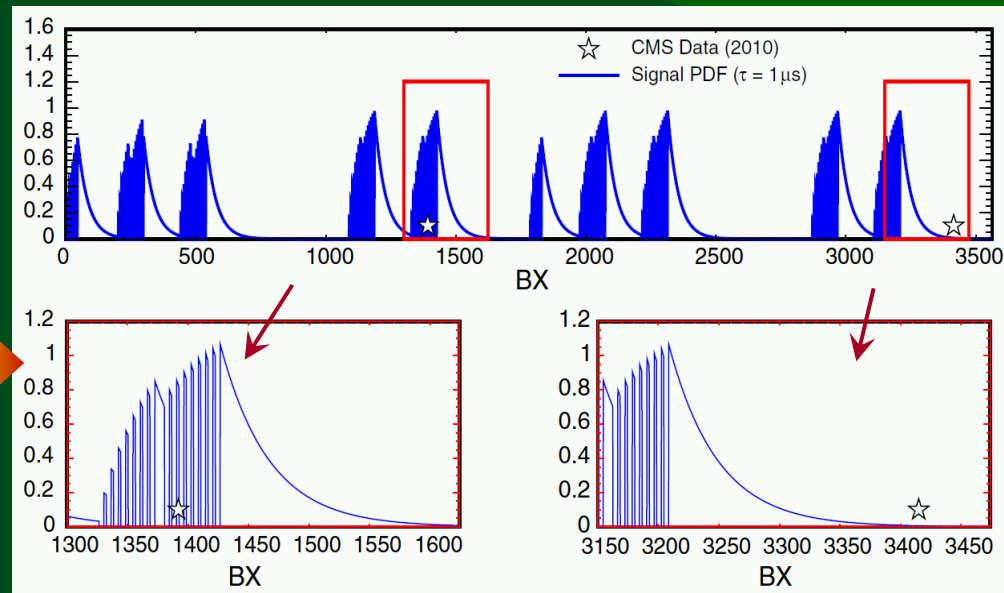
- same time windows

Signal shape

- depends on lifetime – 75ns to 10^{-4} s

Background shape flat

- instrumental noise dominates



No indication of signal \rightarrow proceed to set limits

- Calculate likelihood as a function of
 - background amount (per LHC filling scheme)
 - effective signal cross-section
- Calculate posterior probability using uniform priors in both
 - Limit from 95th quantile
- Nuisance parameters integrated out
 - Integrated luminosity 11%
 - JES 7%

Summary

CMS searches published using:

- Bayesian limits
 - prior constant as a function of searched for cross-section
- The CLs method
 - constraining the limits according to the power of the measurement is a must
 - Points to a need

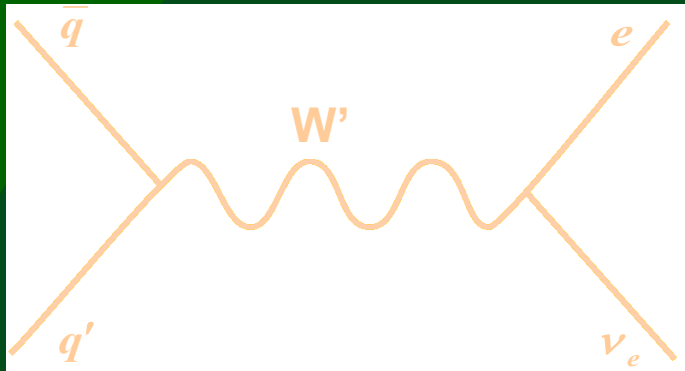
Many ideas, methods, and checks based on Tevatron experience

Will we have an LHC standard method?



Back up slides

W' search



Experimental signature

- Isolated electron
 - $p_T > 30 \text{ GeV}$, $|\eta| < 2.5$
 - Triggers the data acquisition
- p_T imbalance (done with energies though, E_T^{miss})
 - from the full-event reconstruction “particle flow”
 - $\Delta\Phi_{eE_T^{\text{miss}}} > 2.6$
 - $0.4 < \frac{E_T^e}{E_T^{\text{miss}}} < 1.5$

Reference model

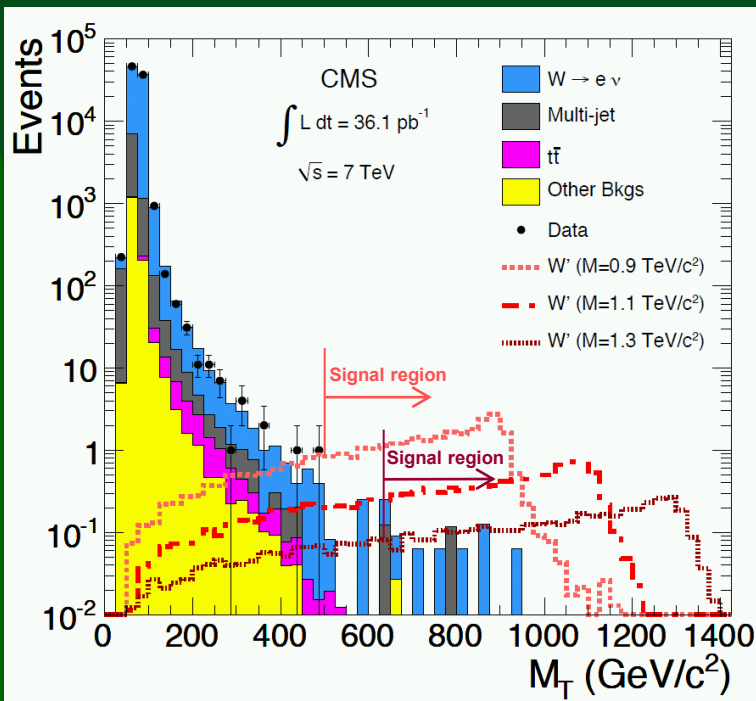
- W' has W-like fermionic couplings,
- W' does not couple to other gauge bosons
- Tevatron limits: $m_{W'} > 1.1 \text{ TeV}$
- signal from Pythia, cross section scaled to NNLO

Observable

Number of events with $M_T >$ threshold

$$(M_T)^2 = 2E_T^e E_T^{\text{miss}} (1 - \cos \Delta\Phi_{eE_T^{\text{miss}}})$$

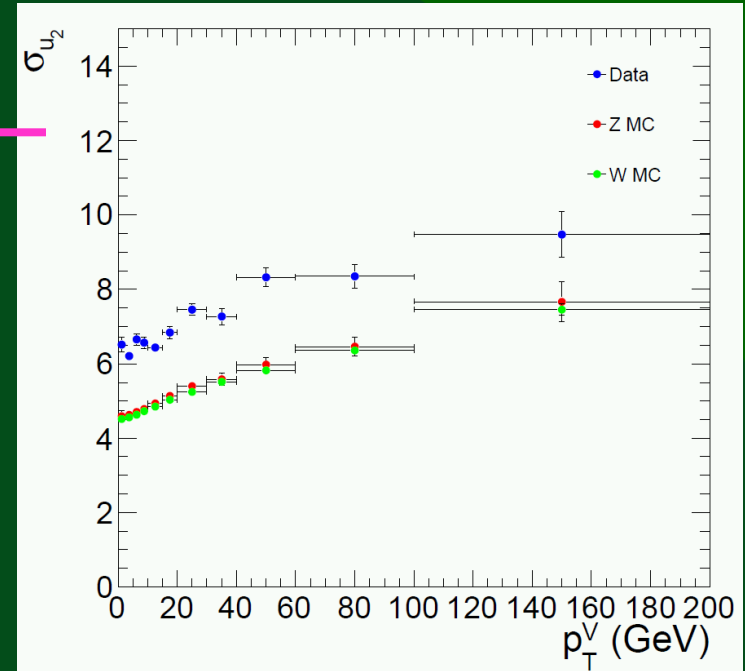
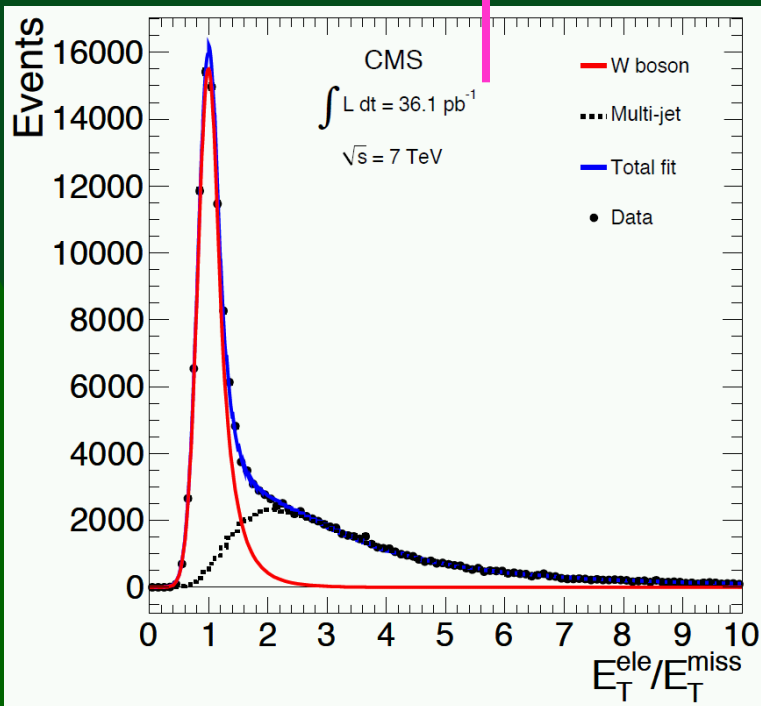
Threshold chosen a-priori as a function of $m_{W'}$,



W' backgrounds

Estimating the main backgrounds:

- Hadronic recoil of W taken from Z data
 - not a big effect
- Multijet spectra taken from sample with non-isolated e candidates
- Their normalizations from a fit to the other-background subtracted data



Black hole search

Model

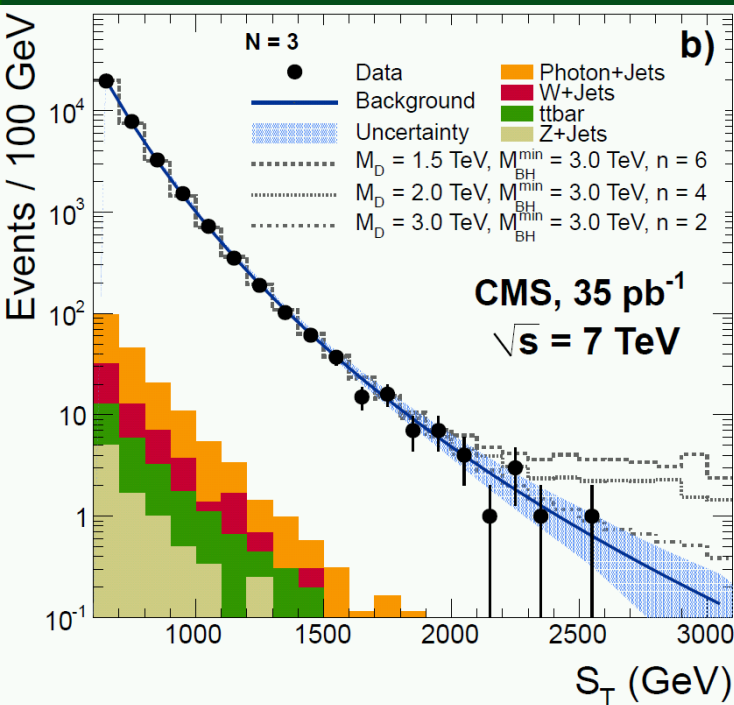
- ADD: n large flat extra dimensions
[Arkani-Hamed et al., PLB 429 (1998) 263]
→ lower Planck scale M_D
- Microscopic black holes in thermal equilibrium → decay via Hawking radiation
- Neglect energy untrapped by the black hole - too model dependent

Experimental signature

Hawking radiation – democratic production

- All objects selected with $p_T > 20 \text{ GeV}$
- Mostly quarks & gluons → jets
- Also electrons, photons and muons
- Objects separated by $\Delta R > 0.3$ (in Φ, η plane)
- Little graviton radiation expected
[Emparan et al., PRL 85 (2000) 499]

Lots (at least >2) of them with large total p_T !



Observables

N = Number of objects

S_T = scalar sum of p_T 's of objects with $p_T > 50 \text{ GeV}$

- low sensitivity to pile up & for QCD's ISR & FSR

Threshold chosen a-priori for each model

Signal

- BlackMax / CHARYBDIS2 generator

$$\sigma \approx \pi r_s^2 \propto \frac{1}{M_D^2} \left(\frac{M_{BH}}{M_D} \right)^{\frac{2}{2+n}}$$

Black hole limit setting

