

**Physics of BES III and
Particle Physics Programme in China**

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Outline

- 1. Introduction**
- 2. BEPC/BEPCII**
- 3. Daya Bay reactor neutrino experiment**
- 4. Particle astrophysics experiments**
 - a) Yanbajing cosmic ray observatory**
 - b) Space experiments**
- 5. Future plan**
- 6. Closing remarks**

1. Particle Physics Experiments in China

- **High energy frontier:**
 - LHC exp.: CMS, Atlas, LHCb, Alice
 - ILC R&D: 1.3GHz SC cavities, Positron source
- **Intensity frontier:**
 - Charm and tau physics at BEPCII
 - Daya Bay neutrino experiment
- **Cosmic frontier**
 - Yangbajing cosmic ray observatory
 - Space Experiments

2. Beijing Electron Positron Collider (BEPC)

BEPC: constructed 1984-1988

BESI: run from 1989-1998

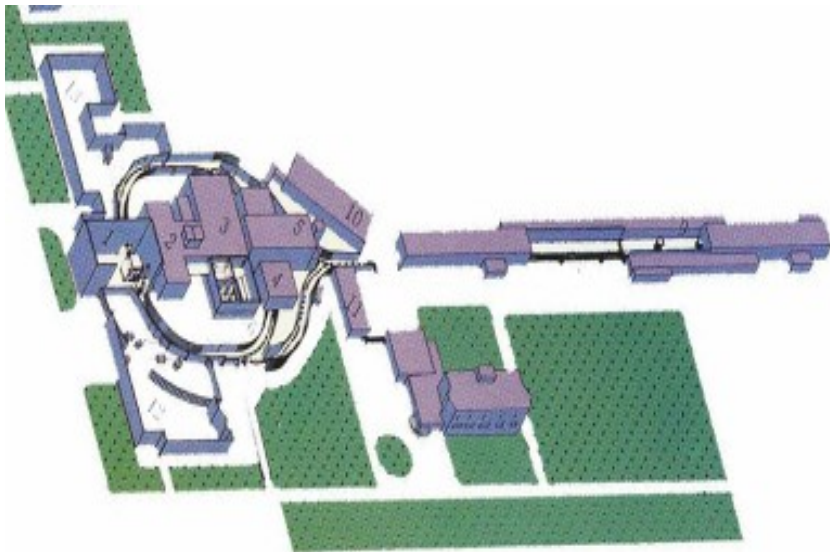
BESII: run from 1999-2004

BECPII upgraded 2004-2008

BESIII: run from 2008

$E_{\text{beam}} \sim 1-2.5 \text{ GeV}$

$E_{\text{beam}} \sim 1-2.3 \text{ GeV}$



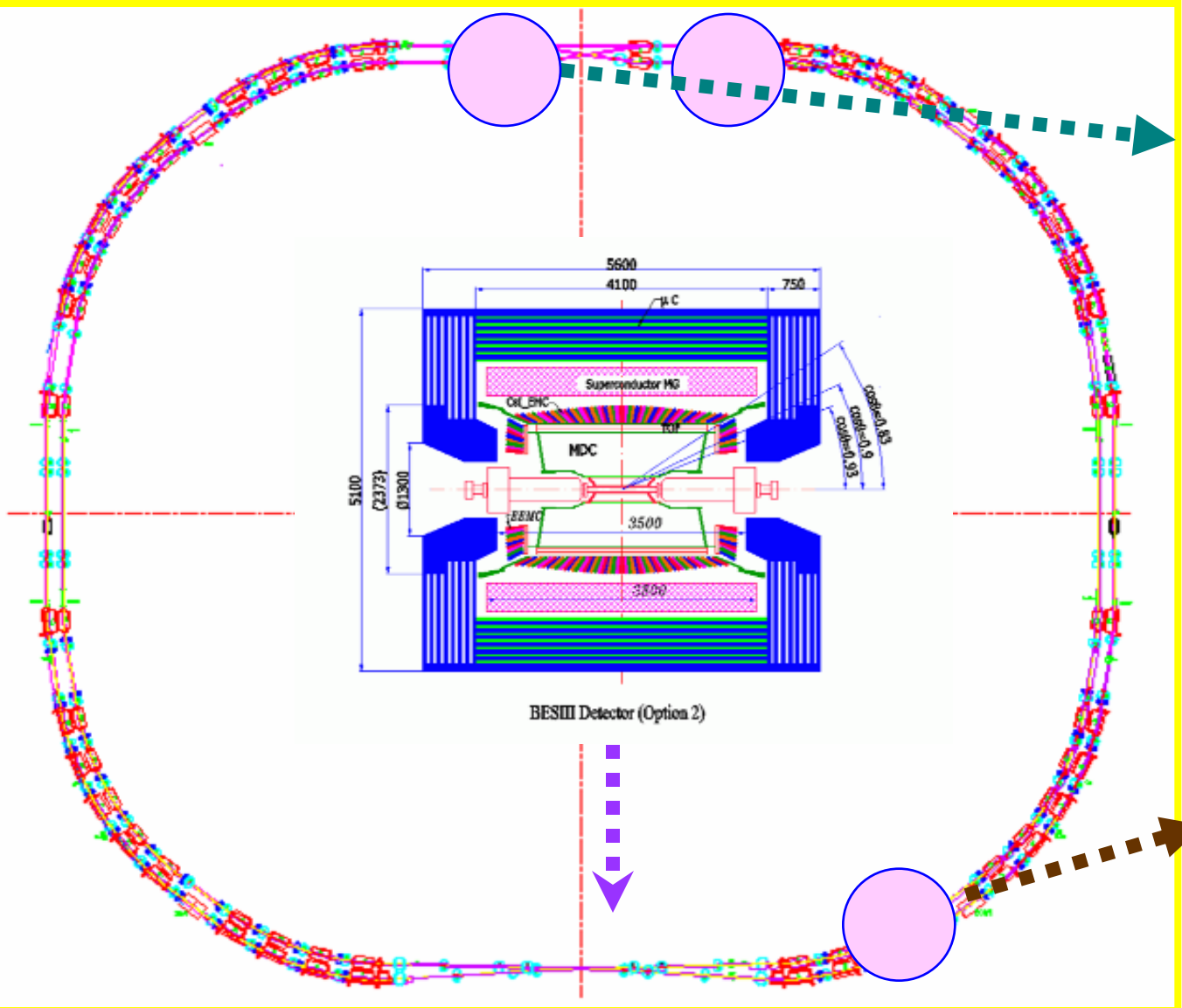
With BES I and BES II data

- precision measurement of τ mass: **10 times improved.**
Lepton universality!
- R measurements improve uncertainties by a factor of 2-3 ($\Delta R/R \sim 6\%$). Great impact to $M_H \cdot \alpha(M_Z^2)$, $g-2$
- Some new particles X(1835) observed. Hard to be interpreted as conventional hadrons. ppbar bound state?

Precision measurements on the Charm physics require high statistics and small sys. errors →

Major upgrade: BEPCII / BESIII (2004-2008)

BEPCII: a high luminosity double-ring collider



SC RF



Two rings

Storage Ring installation finished



The BESIII Detector

Magnet yoke

SC magnet, 1T

RPC: 9 layers

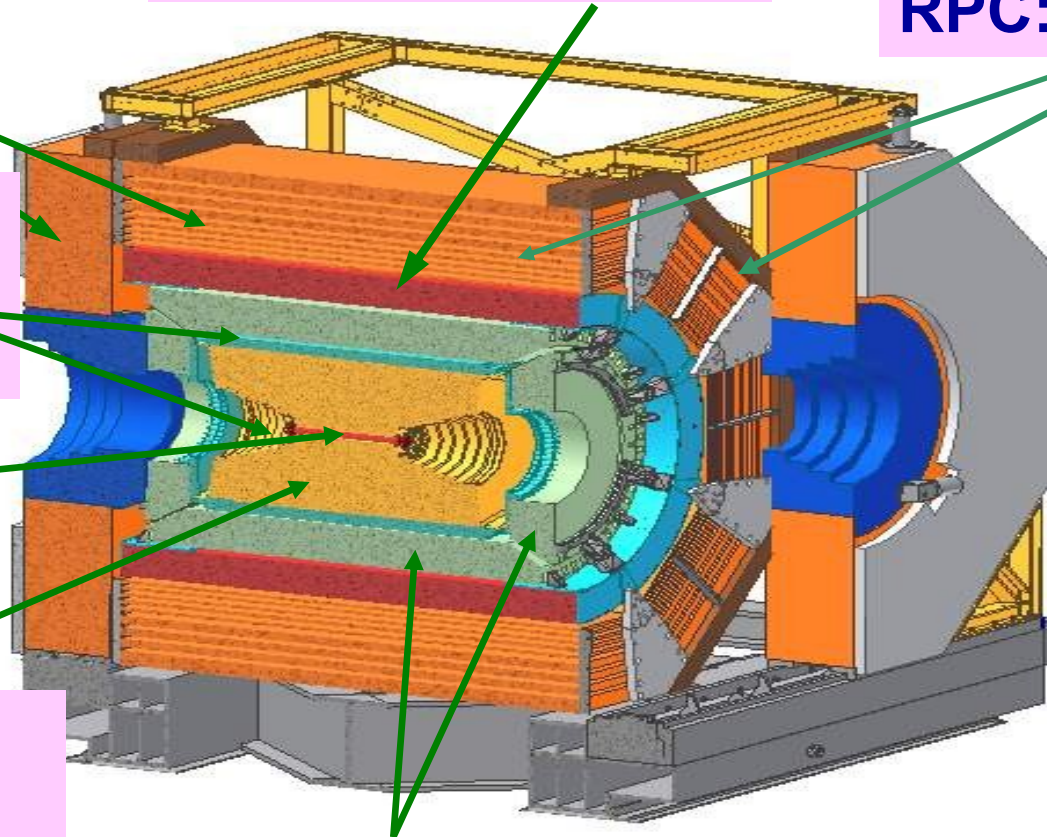
TOF,
 σ_T (ps) = 100 ps Barrel
110 ps Endcap

Be beam pipe

MDC,

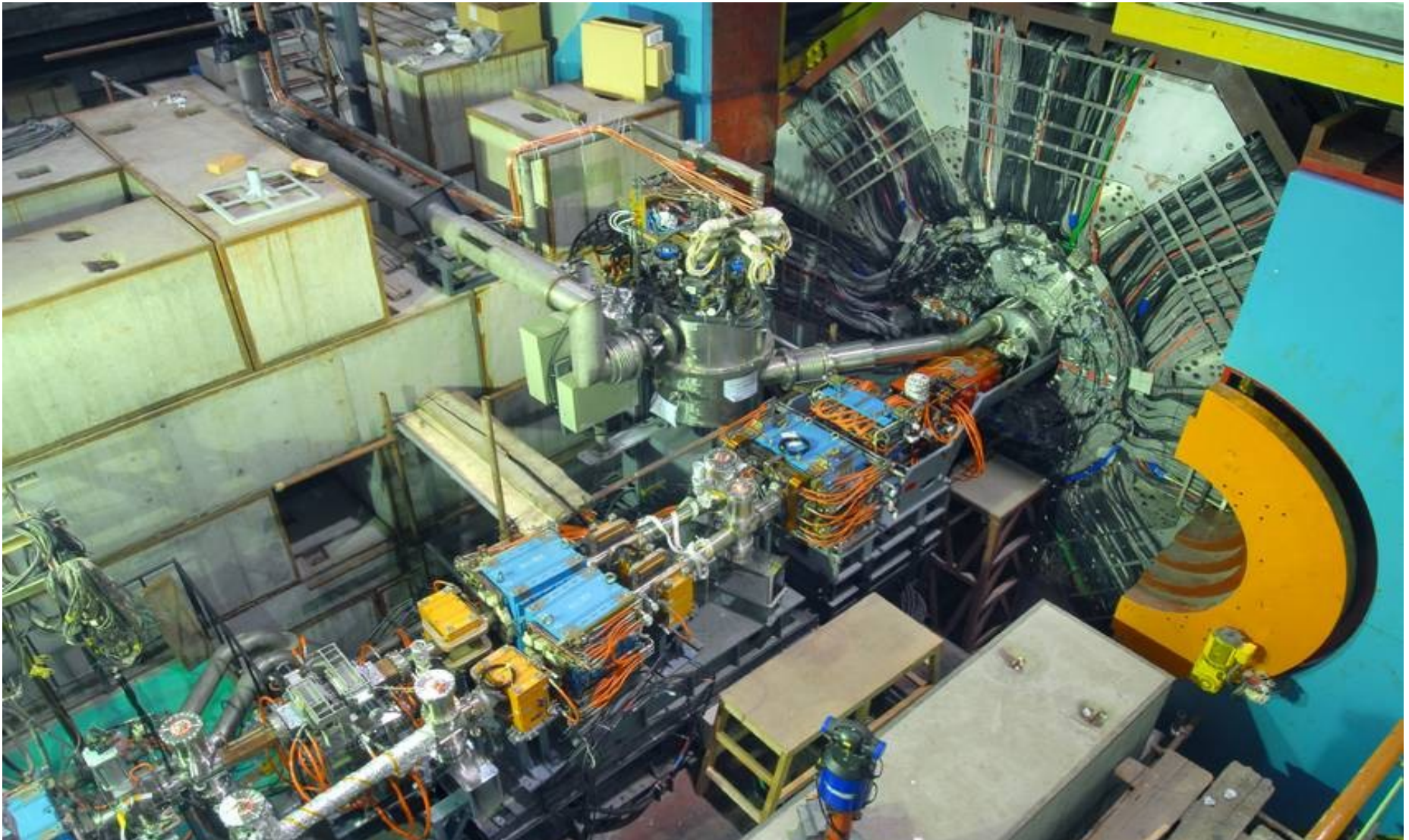
σ_{XY} (μm) = 130
 $\Delta P/P$ = 0.5 % (1 GeV)
 $\sigma_{dE/dx}$ = 6-7 %

CsI(Tl) calorimeter,
 $\Delta E/\sqrt{E}$ = 2.5 % (1 GeV)
 $\sigma_{z,\phi}(\text{cm})$ = $0.5\text{cm}/\sqrt{E}$



Joint Commissioning

- BESIII detector moved into the IR in May 2008
- Joint commissioning started 22 June 2008.



Performance of BEPCII injector linac

Parameters		Design	Accept test	BEPC
Energy (GeV)		1.89	1.89	1.30-1.55
Current (mA)	e+	37	66	~5
	e-	500	550	300
Emittance ($1 \hat{f}$) (mm-mrad)	e+	0.40	0.35 ~ 0.27	----
	e-	0.10	0.097~0.079	----
Energy spread (1σ) (%)	e+	0.50	0.371	~0.80
	e-	0.50	0.295	~0.80
Energy stability (%)		± 0.15	$\pm (0.050 \sim 0.035)$	----
Orbit stability (mm)		± 0.30	$\pm (0.119 \sim 0.058)$	----
Repeation rate		50	50	12.5
e ⁺ inj. rate (mA/min.)		50	61.5	1 ~ 3

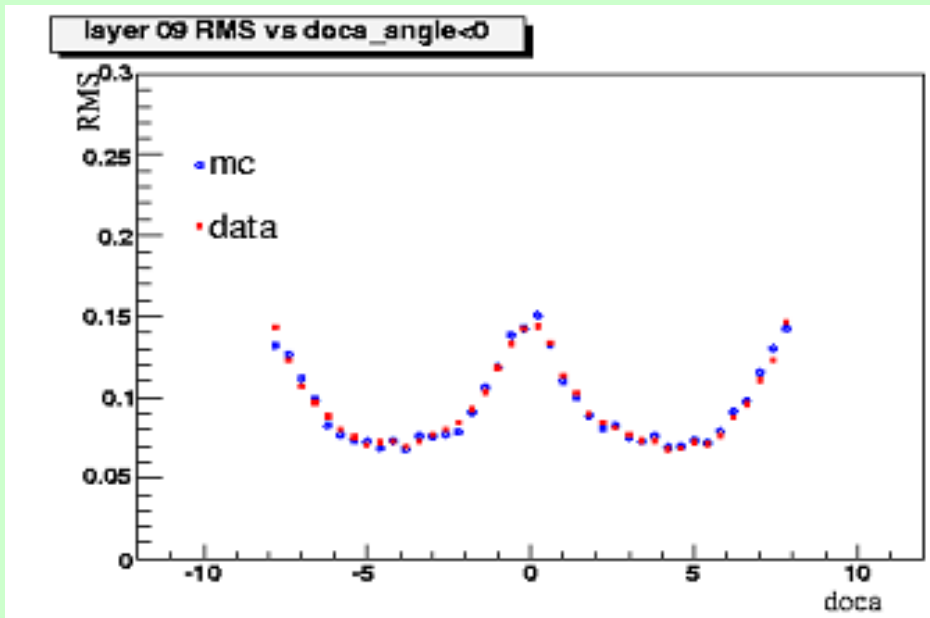
positron injection rate improved by > factor of 15

Performance of BEPCII Storage Ring

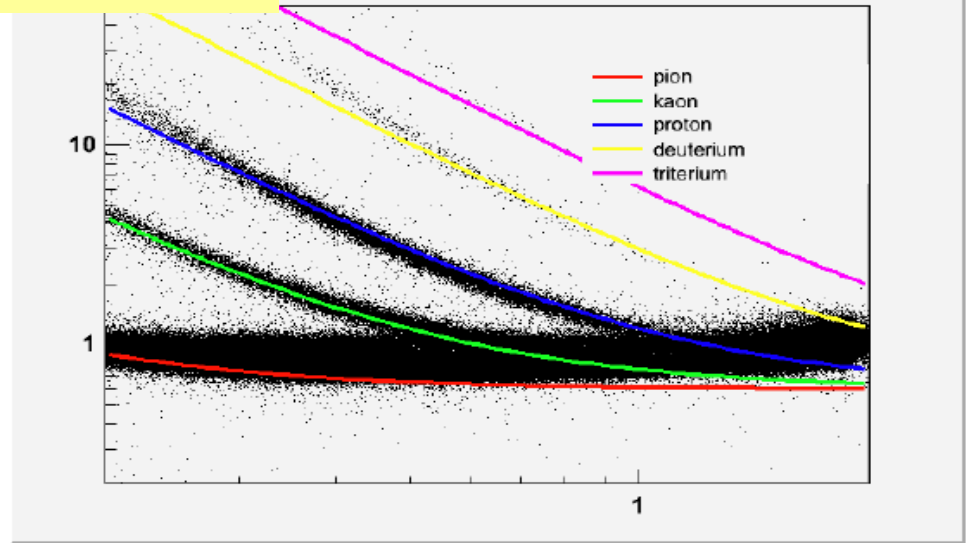
Parameters	Design	Accept test result	
		BER	BPR
Energy (GeV)	1.89	1.89	1.89
Beam current. (mA)	910	650	700
Bunch current. (mA)	9.8	>10	>10
Bunch number	93	93	93
RF voltage	1.5	1.6	1.6
ν_s @ $V_{RF}=1.5\text{MV}$	0.033	0.032	0.032
β_x^*/β_y^* (m)	1.0/0.015	~1.0/0.0135	~1.0/0.0135
Injection rate (mA/min)	200 e ⁻ /50 e ⁺	>200	>50
L ($10^{33}\text{cm}^{-2}\text{s}^{-1}$)	1.0	0.33	

Factor of 33 improvement in luminosity at May 2009

Excellent detector performance

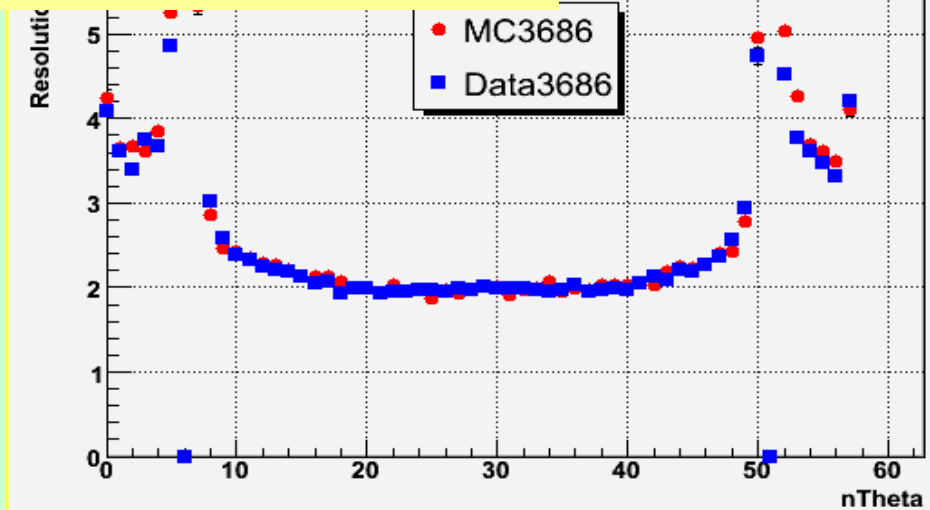


dE/dx vs P



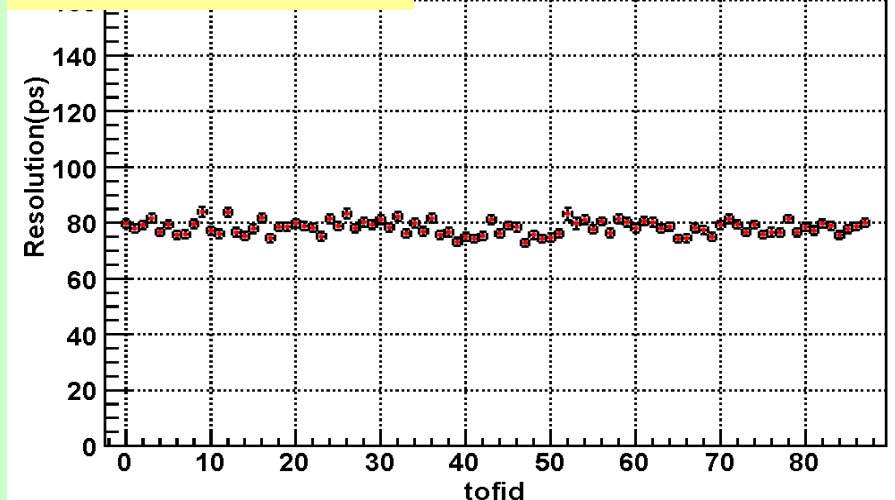
EMC energy resolution

Design: 2.5% @ 1 GeV

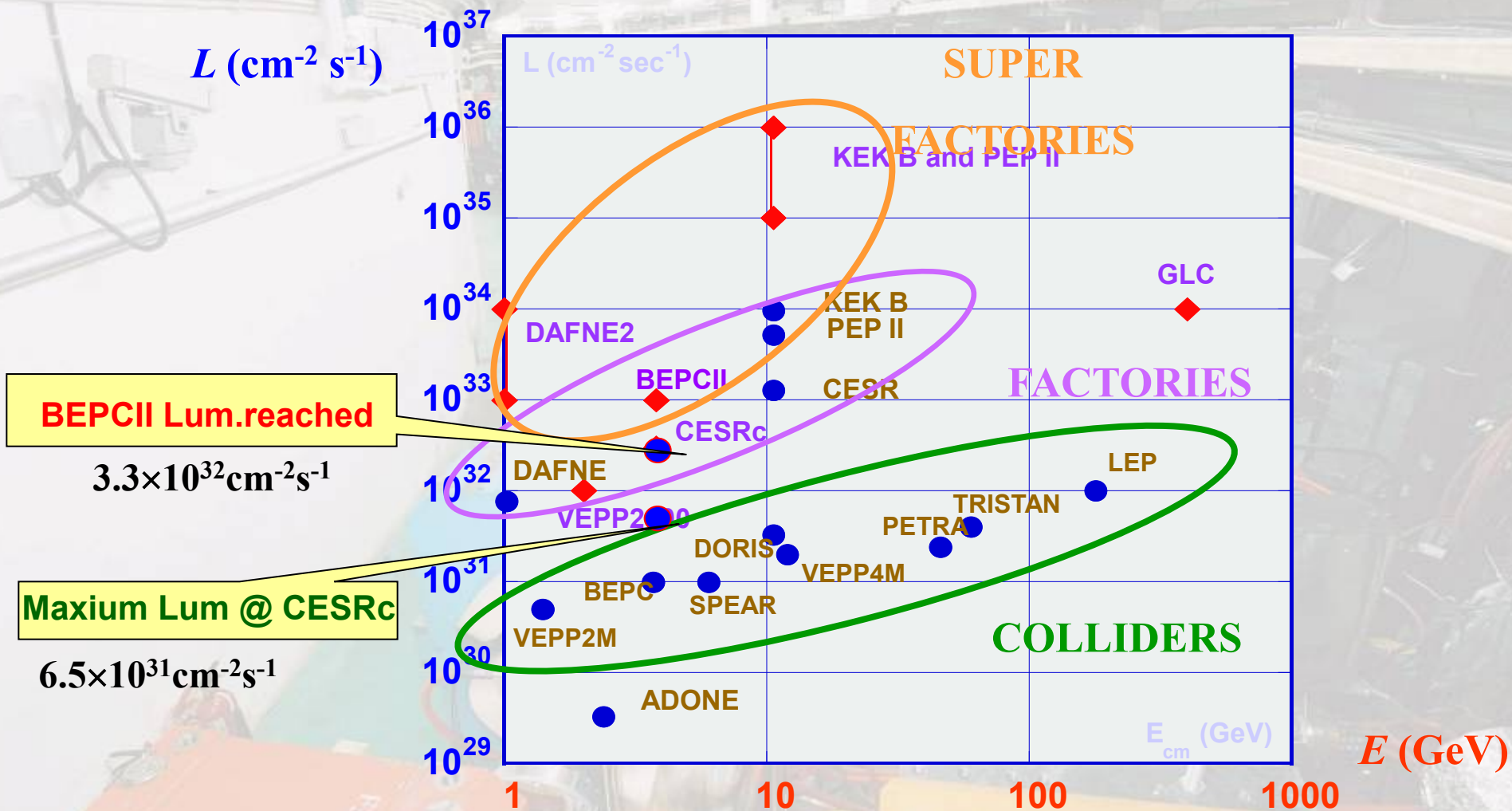


Time resolution

Design: 80ps



Electron Positron Colliders



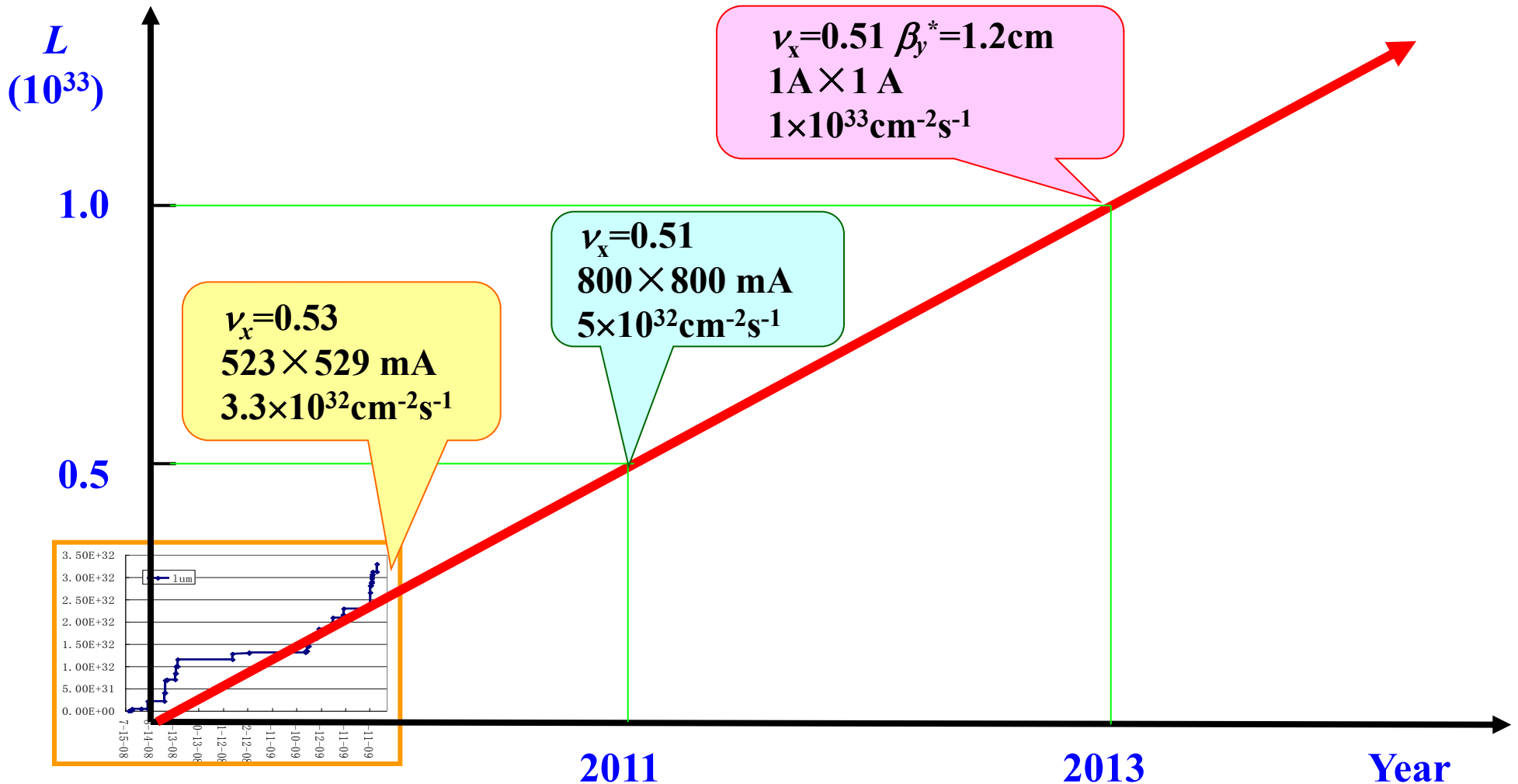
On going upgrades

- To increase single bunch current in collision
- To enhance beam-beam parameter towards 0.04
- To move horizontal tune to 0.51
- To increase colliding bunch number and pattern
- Squeeze β_y^*
- ...

Longterm upgrades (under discussion)

- Increasing beam energy? (2.3GeV now)
- Crab-waist for higher luminosity
- Collision with polarized beam
 - Physics requirements
 - e- beam? location for rotators ? □

BEPCII luminosity plan



BESIII collaboration

Political Map of the World, June 1999

US (6)

Univ. of Hawaii
Univ. of Washington
Carnegie Mellon Univ.
Univ. of Minnesota
Univ. of Rochester
Univ. of Indiana

EUROPE (10)

Germany: Univ. of Bochum,
Univ. of Giessen, GSI, Mainz, HIM
Russia: JINR, Dubna; BINP, Novosibirsk
Italy: Univ. of Torino, Frascati Lab
Netherland: KVI/Univ. of Groningen

Korea (1)

Souel Nat. Univ.

Japan (1)

Tokyo Univ.

Pakistan (1)

Univ. of Punjab

China(29)

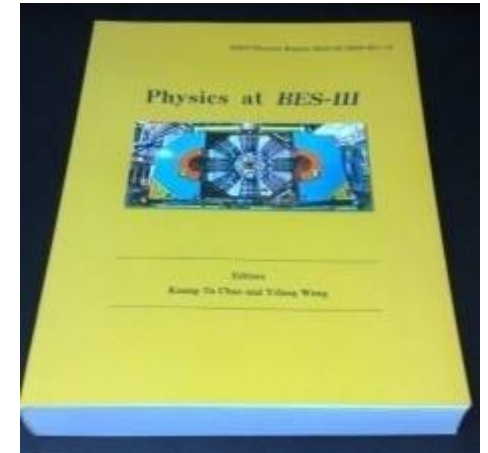
IHEP, CCAST, Shandong Univ.,
Univ. of Sci. and Tech. of China
Zhejiang Univ., Huangshan Coll.
Huazhong Normal Univ., Wuhan Univ.
Zhengzhou Univ., Henan Normal Univ.
Peking Univ., Tsinghua Univ. ,
Zhongshan Univ., Nankai Univ.
Shanxi Univ., Sichuan Univ
Suzhou Uni., Hangzhou Normal Uni.
Hunan Univ., Liaoning Univ.
Henan Uni. of Sci. & Tech.,
Nanjing Univ., Nanjing Normal Univ.
Guangxi Normal Univ., Guangxi Univ.
Hong Univ., Hong Kong Chinese Univ.

48 institutions

~ 300 collaborators

BESIII Physics: Precision tests of SM and search for new physics

- Light hadron spectroscopy
 - Full spectra: normal & exotic hadrons QCD
 - How quarks form hadron ? non-pQCD
- Charm physics
 - CKM matrix elements → SM and beyond
 - $D\bar{D}$ mixing and CPV → SM and beyond
- Charmonium physics
 - Spectroscopy and transition → pQCD & non-pQCD
 - New states above open charm thresholds → exotic hadrons ?
 - pQCD: $\rho\pi$ puzzle → a probe to non-pQCD or ?
- Tau physics and QCD
 - Precision measurement of the tau mass and R value
- Search for rare and forbidden decays

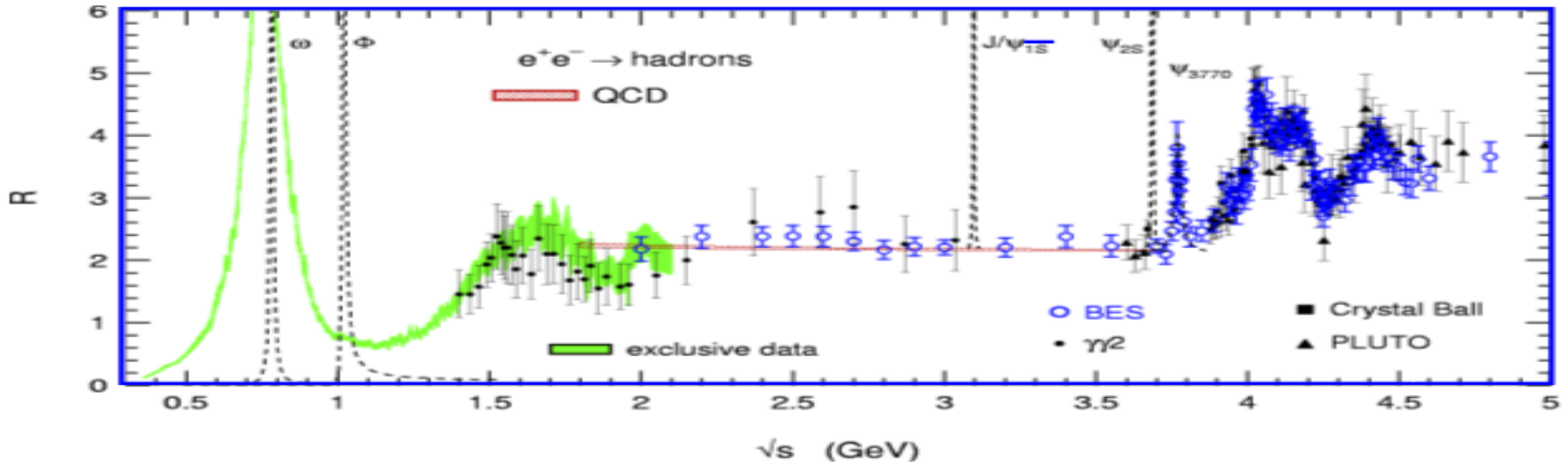


hep-ex/0809.1869

IJMP A V24, No 1(2009) supp

BESIII data taking status & plan

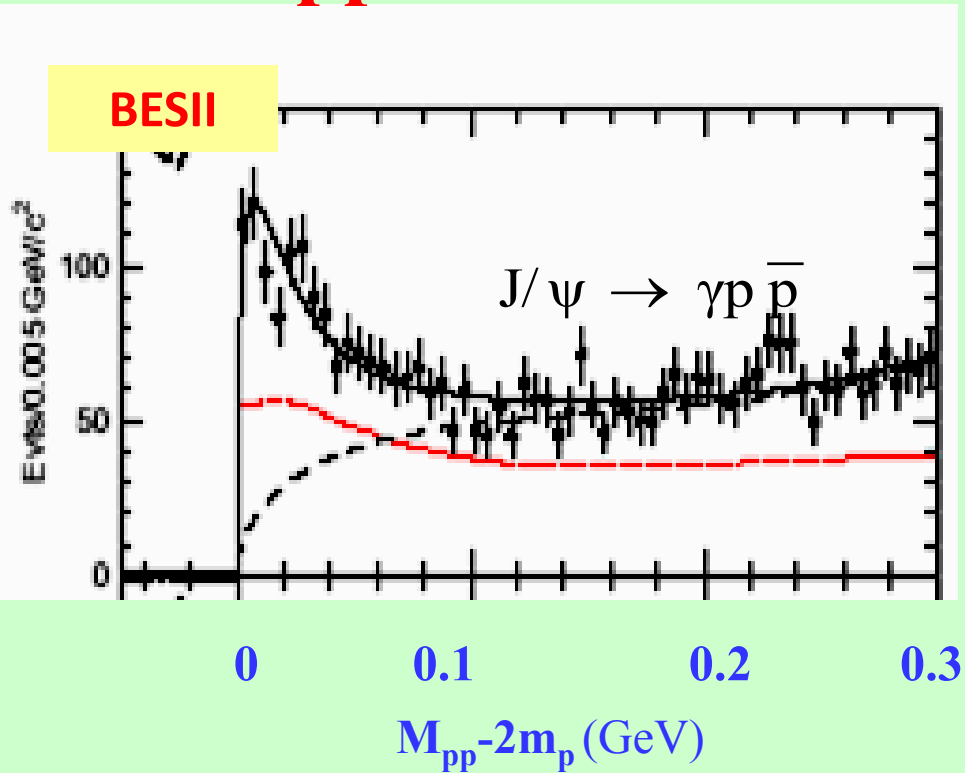
	Previous Data set	BESIII Near future
J/psi	BESII 58M	2009: 200M, 2012: 1 B
Psi'	CLEO: 28 M	2009: 100M
Psi''	CLEO: 0.8 /fb	2010: 0.9/fb, 2011-12: 3.5/fb
$\psi(4040)/\psi(4160)$ & scan	CLEO: 0.6/fb @ $\psi(4160)$	2011: 0.4/fb @ $\psi(4040)$ 2013: 4/fb
R scan & Tau	BESII	2014



BESIII: First physics results

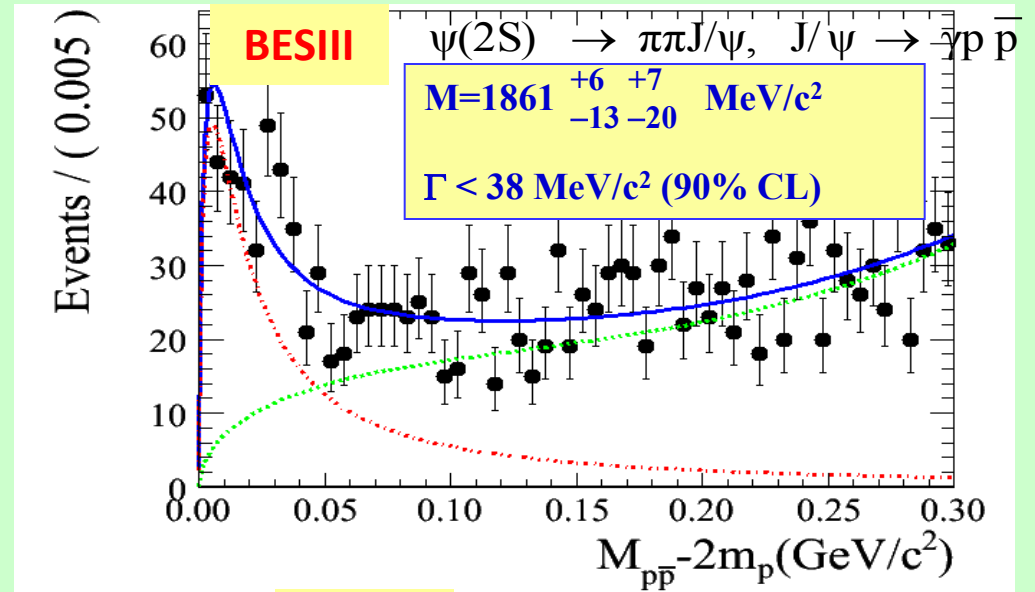
- **Light hadron physics**
 - Confirmation of BESII results: **threshold enhancement**
 $\gamma p p \bar{p}$, $\gamma \omega \phi$, $X(1835)$, ...
 - New resonances
 - New observations: e.g. a_0 - f_0 mixing
- **Charmonium physics**
 - Improved measurements: h_c , η_c , χ_{cJ} , , ...
 - New observations
 - χ_{cJ} decays
 - h_c decays

Confirmation of the BESII observation: $\bar{p}p$ threshold enhancement in J/ψ decays

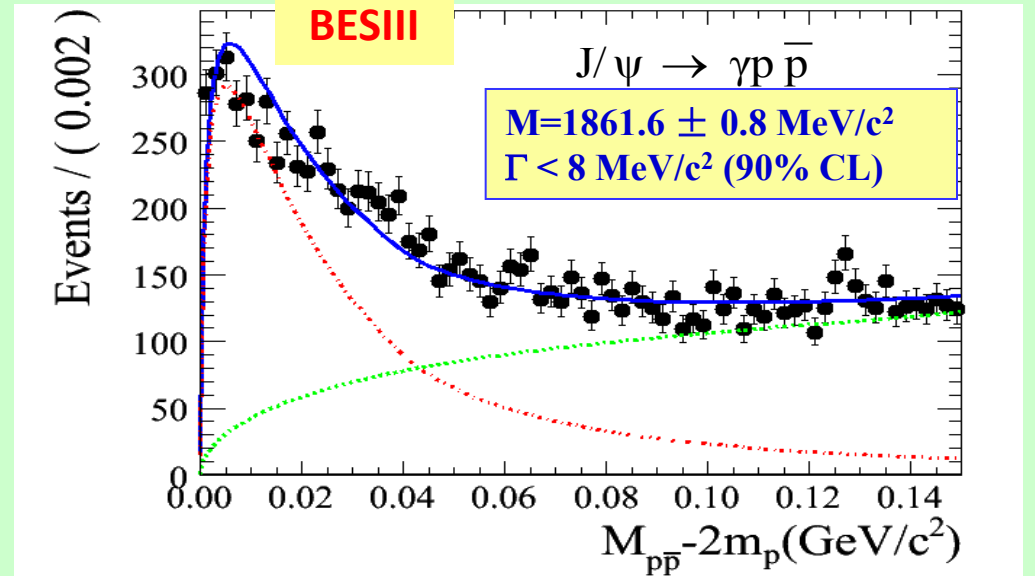


$M = 1859^{+3}_{-10} \text{ MeV}/c^2$
 $\Gamma < 30 \text{ MeV}/c^2$ (90% CL)

arXiv:1001.5328,
 Chinese Phys. C 34 (2010) 421

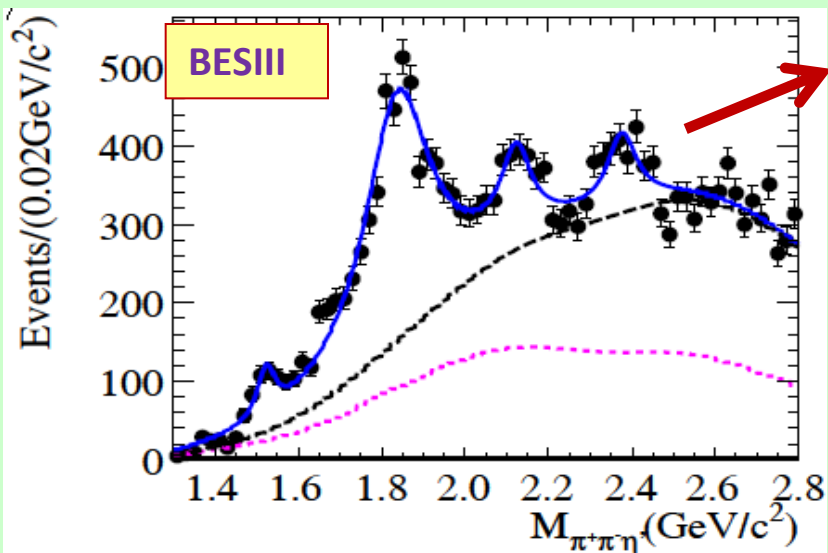


$M = 1861^{+6}_{-13} \text{ MeV}/c^2$
 $\Gamma < 38 \text{ MeV}/c^2$ (90% CL)



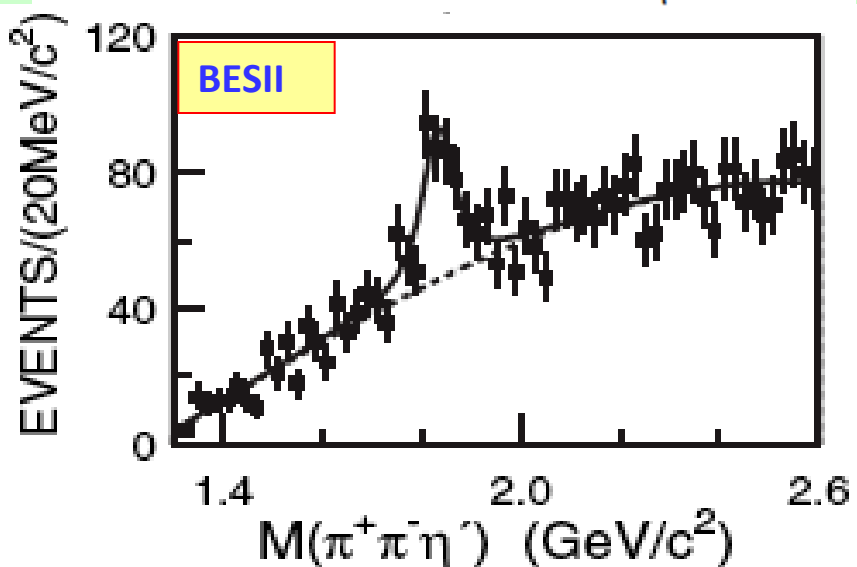
$M = 1861.6 \pm 0.8 \text{ MeV}/c^2$
 $\Gamma < 8 \text{ MeV}/c^2$ (90% CL)

Confirmation of BESII observation: X(1835) in $J/\psi \rightarrow \gamma \eta' \pi \pi$



Two new resonance

resonance	$M(\text{ MeV}/c^2)$	$\Gamma(\text{ MeV}/c^2)$	Stat. sig.
X(1835)	1838.1 ± 2.8	179.5 ± 9.1	$> 25\sigma$
X(2120)	2124.8 ± 5.6	101 ± 14	$> 7.2\sigma$
X(2370)	2371.0 ± 6.4	108 ± 15	$> 6.7\sigma$



significant $ce : 7.7 \sigma$

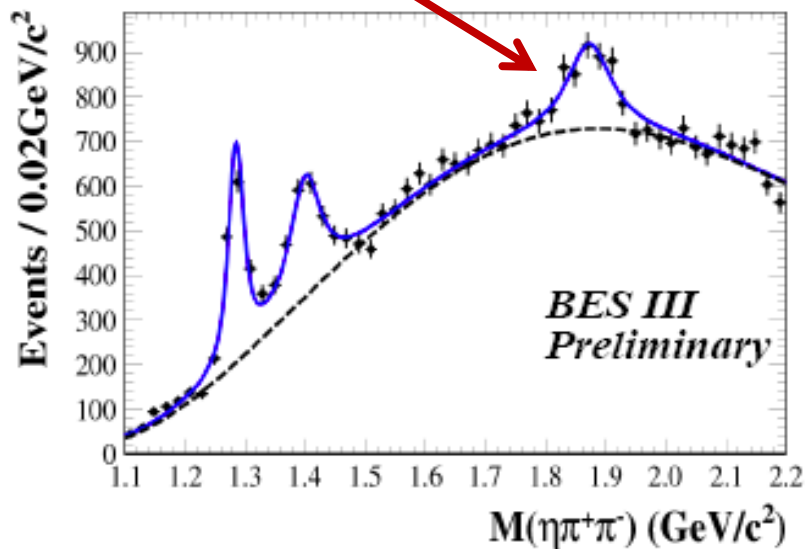
$M = 1833.7 \pm 6.1(\text{stat}) \pm 2.7(\text{syst}) \text{ MeV}$

$\Gamma = 67.7 \pm 20.3(\text{stat}) \pm 7.7(\text{syst}) \text{ MeV}$

To be submitted to PRL

Observation of $X(1870) \rightarrow a_0(980)\pi$ in $J/\psi \rightarrow \omega\pi^+\pi^-\eta$

X(1870)



Fit result (stat. sig. $\sim 7.7\sigma$)

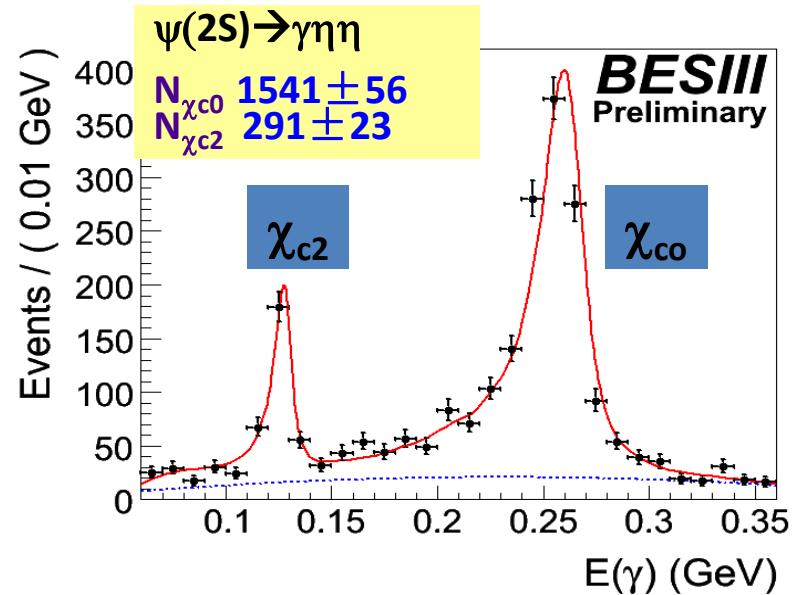
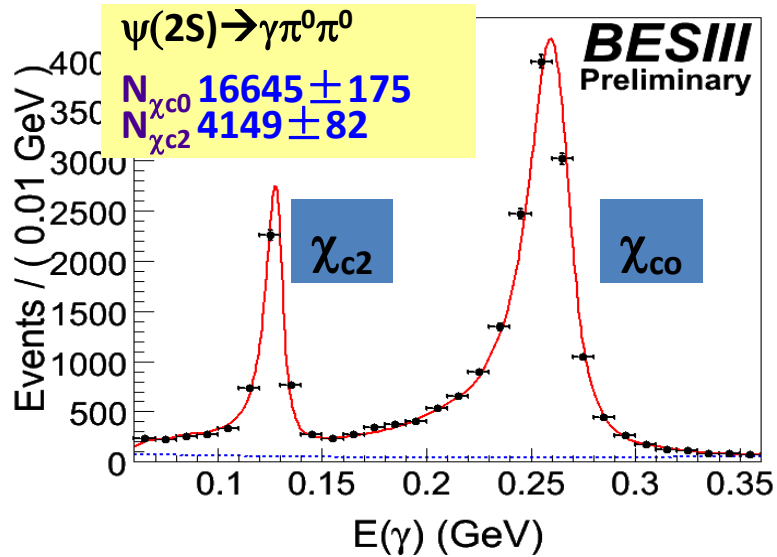
$$M = 1873 \pm 11 \text{ MeV}$$

$$\Gamma = 82 \pm 19 \text{ MeV}$$

Whether the X(1870) is the X(1835) or $\eta_2(1870)$, or a new resonance, further study is needed.

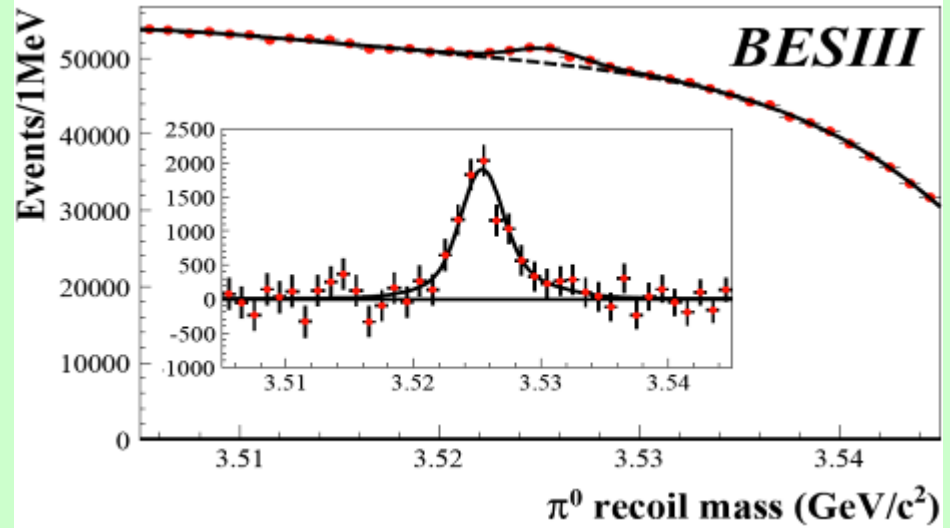
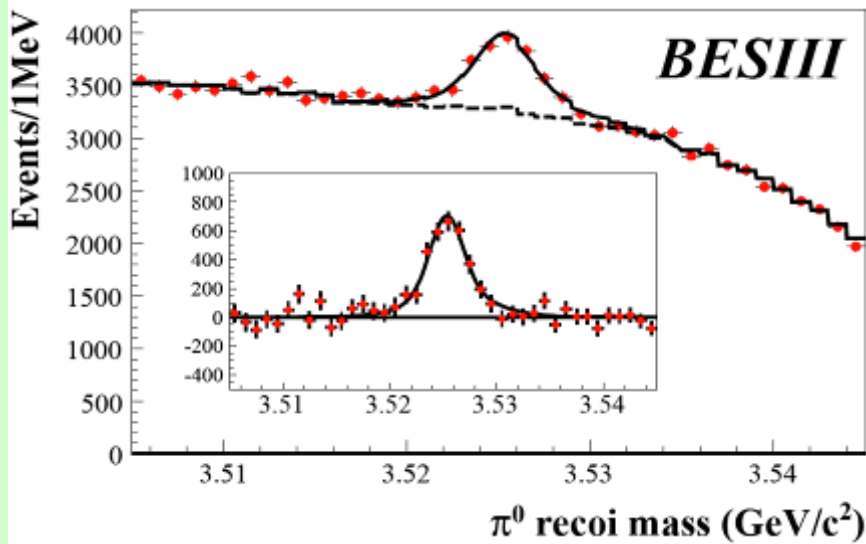
To be submitted to PRL

$\psi(2S) \rightarrow \gamma\pi^0\pi^0, \gamma\eta\eta$ ($\eta \rightarrow \gamma\gamma, \pi^0 \rightarrow \gamma\gamma$)



BR (10^{-3})		χ_{c0}	χ_{c2}
$\pi^0\pi^0$	BESIII	$3.23 \pm 0.03 \pm 0.23 \pm 0.14$	$0.88 \pm 0.02 \pm 0.06 \pm 0.04$
	PDG08	2.43 ± 0.20	0.71 ± 0.08
	CLEO-c	$2.94 \pm 0.07 \pm 0.35$	$0.68 \pm 0.03 \pm 0.08$
$\eta\eta$	BESIII	$3.44 \pm 0.10 \pm 0.24 \pm 0.20$	$0.65 \pm 0.04 \pm 0.05 \pm 0.03$
	PDG08	2.4 ± 0.4	< 0.5
	CLEO-c	$3.18 \pm 0.13 \pm 0.35$	$0.51 \pm 0.05 \pm 0.06$

Observation of h_c in $\psi(2S) \rightarrow \pi^0 h_c, h_c \rightarrow \gamma \eta_c$



$$M(h_c)^{\text{Inc}} = 3525.40 \pm 0.13 \pm 0.18 \text{ MeV}$$

$$\Gamma(h_c)^{\text{Inc}} = 0.73 \pm 0.45 \pm 0.28 \text{ MeV}$$

$$\begin{aligned} \text{Br}(\psi' \rightarrow \pi^0 h_c) \times \text{Br}(h_c \rightarrow \gamma \eta_c)^{\text{Inc}} \\ = (4.58 \pm 0.40 \pm 0.50) \times 10^{-4} \end{aligned}$$

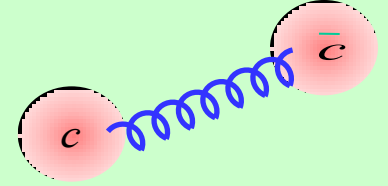
$$\text{Br}(\psi' \rightarrow \pi^0 h_c) = (8.4 \pm 1.3 \pm 1.0) \times 10^{-4}$$

$$\text{Br}(h_c \rightarrow \gamma \eta_c) = (54.3 \pm 6.7 \pm 5.2) \%$$

arXiv:1002.0501
Phys.Rev.Lett.
104(2010) 132002

BESIII measured for the first time
 $\Gamma(h_c)^{\text{Inc}}$, $\text{Br}(\psi' \rightarrow \pi^0 h_c)$ & $\text{Br}(h_c \rightarrow \gamma \eta_c)$

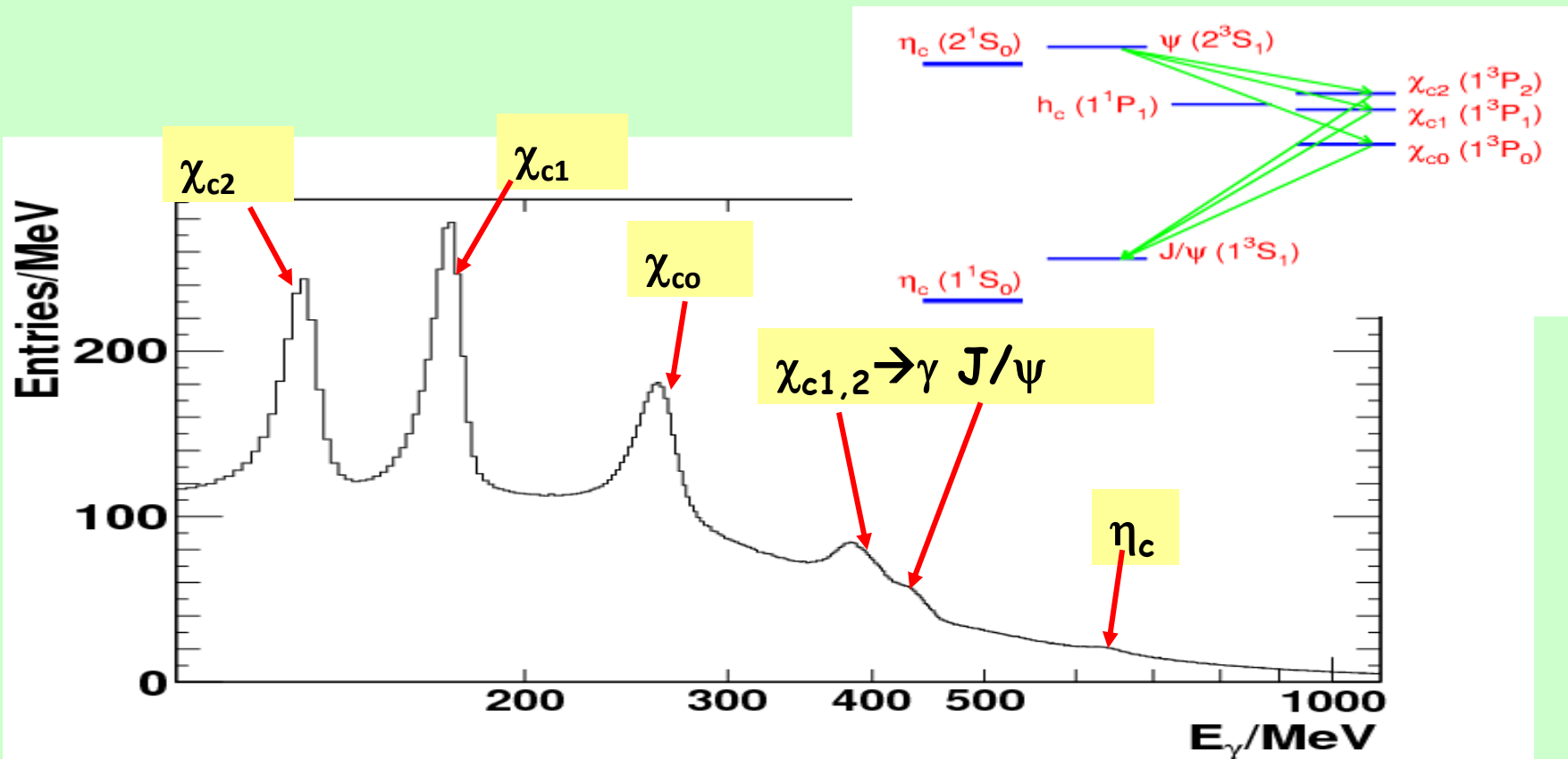
Charmonium physics



- Understand how quarks form hadron?
 - Production, decays, transition, spectrum

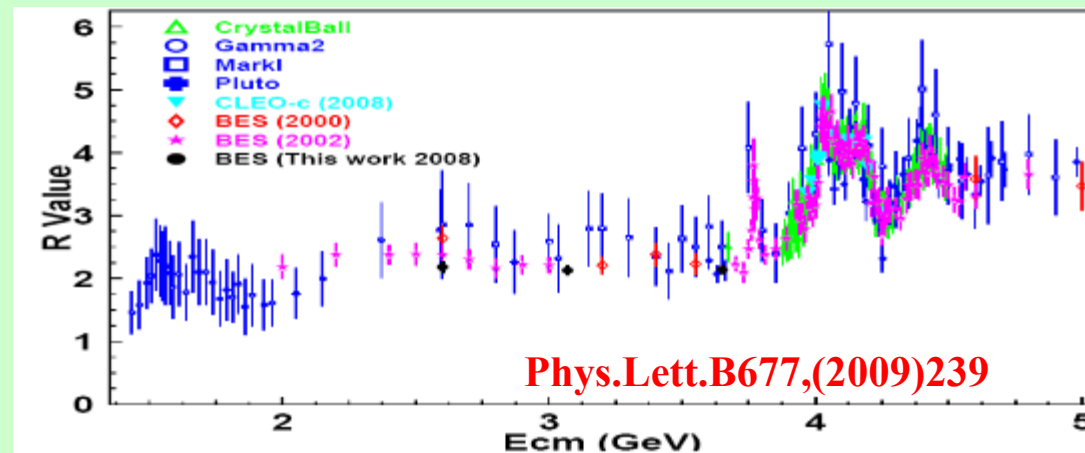
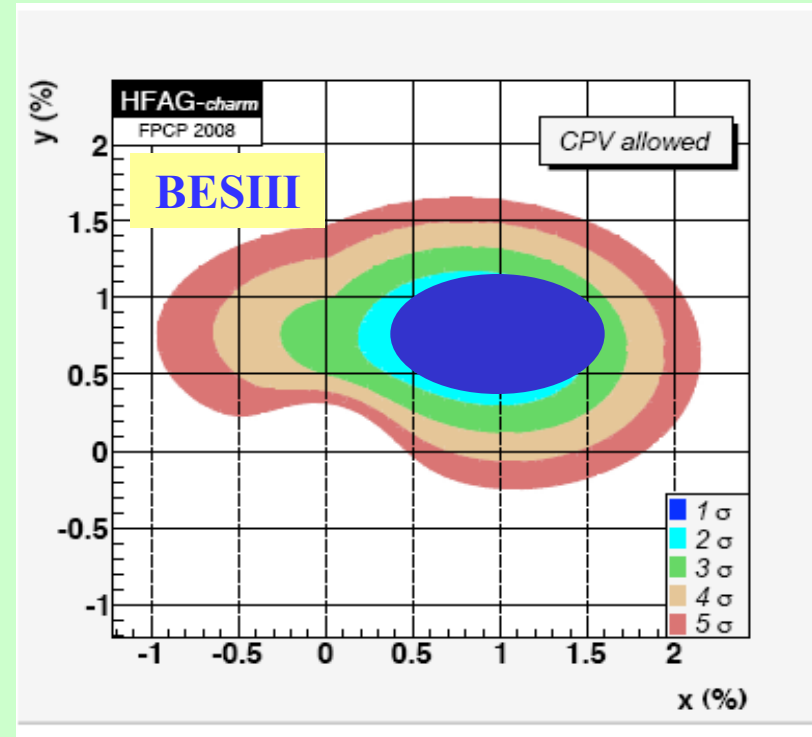
Examples of interesting/long standing issues:

- $\rho\pi$ puzzle, glueball
- Missing states ?
- Mixing states ?
- New states above open charm thre.(X,Y,Z,...)



Other physics programs at BESIII

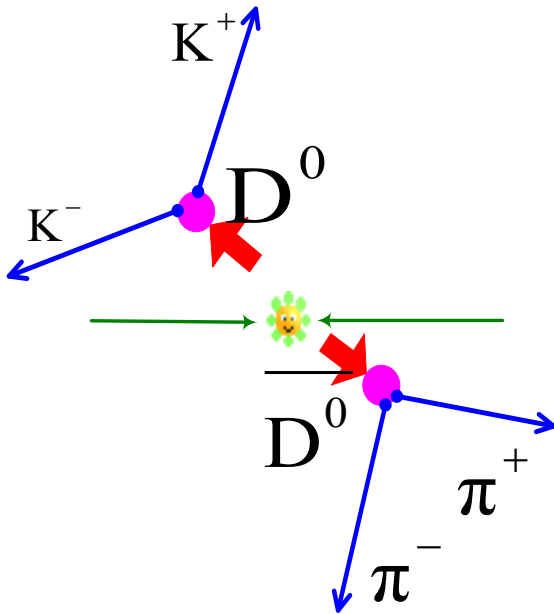
- Charm physics
 - DD mixing and CPV
 - CKM matrix elements
 - V_{cs} : 11% \rightarrow 1.2%
 - V_{cd} : 4% \rightarrow 1.4%
- QCD
 - pQCD \longleftrightarrow non-pQCD
 - Form factor of hadrons
 - measurement of R & α_s
- Tau physics
 - Tau mass
 - Tau decays:



Expected Exp. Accuracies in D decay measurements at BESIII

Observable	CKM	QCD	Lattice	Exp meas	Exp err
$Br(D \rightarrow l\nu)$	$ V_{cd} $	f_D	2%	$f_D V_{cd} $	1.1%
$Br(D_s \rightarrow l\nu)$	$ V_{cs} $	f_{D_s}	1.5%	$f_{D_s} V_{cs} $	0.7%
$\frac{Br(D_s \rightarrow l\nu)}{Br(D \rightarrow l\nu)}$	$\frac{ V_{cs} }{ V_{cd} }$	$\frac{f_{D_s}}{f_D}$	1%	$\left \frac{V_{cs} f_{D_s}}{V_{cd} f_D} \right $	0.8%
$d\Gamma(D^0 \rightarrow \pi^-)$	$ V_{cd} $	$F_{D \rightarrow \pi}(0)$	4%	$ V_{cd} F_{D \rightarrow \pi}(0)$	0.6%
$d\Gamma(D^0 \rightarrow K^-)$	$ V_{cs} $	$F_{D \rightarrow K}(0)$	3%	$ V_{cs} F_{D \rightarrow K}(0)$	0.5%
$d\Gamma(D_s \rightarrow K)$	$ V_{cd} $	$F_{D_s \rightarrow K}(0)$	2%	$ V_{cd} F_{D_s \rightarrow K}(0)$	1.2%
$d\Gamma(D_s \rightarrow \phi)$	$ V_{cs} $	$F_{D_s \rightarrow \phi}(0)$	1%	$ V_{cs} F_{D_s \rightarrow \phi}(0)$	0.8%

CP Violation at $\psi(3770)$ at BESIII



CP violating asymmetries can be measured by searching for events with two CP odd or two CP even final states:

$\pi^+\pi^-$, K^+K^- , $\pi^0\pi^0$, $K_S\pi^0$,

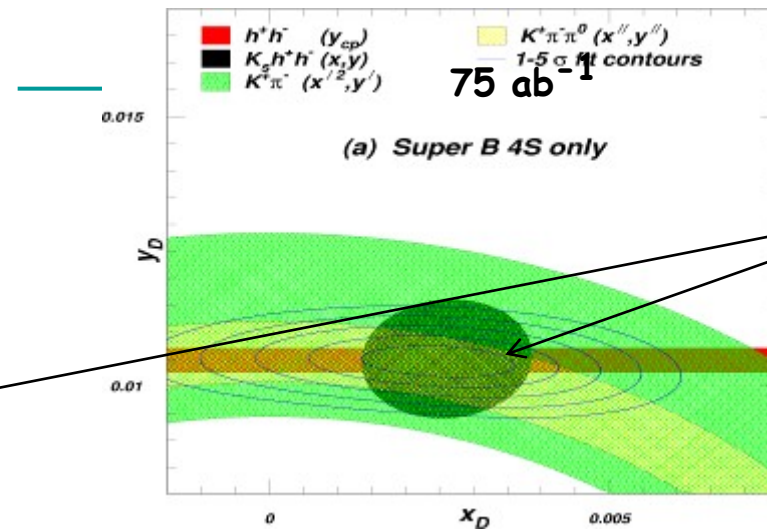
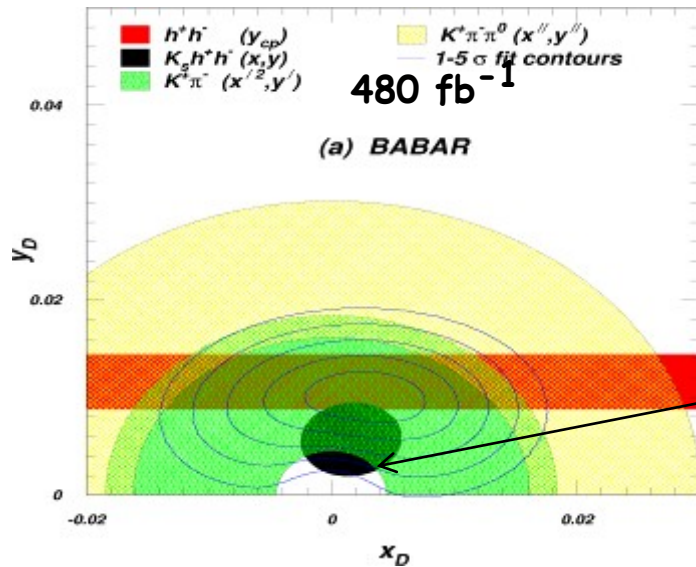
For the decay of $\psi'' \rightarrow f_1 f_2$

$$\text{CP}(f_1 f_2) = \text{CP}(f_1) \cdot \text{CP}(f_2) \cdot (-1)^L = -$$

$$\text{CP}(\psi'') = +$$

A_{CP} sensitivity : $\Delta A \sim 10^{-3}$

Project to $75\text{ab}^{-1}@Y(4S)$:



Golden channels

Min. χ^2 fits (blue contours)

$$x_D = (5.5 \pm_{1.2}^{1.3}) \times 10^{-3}$$

$$y_D = (8.3 \pm 1.3) \times 10^{-3}$$



$$x_D = (xxx^{+0.72}_{-0.75}) \times 10^{-3}$$

$$y_D = (xxx \pm 0.19) \times 10^{-3}$$

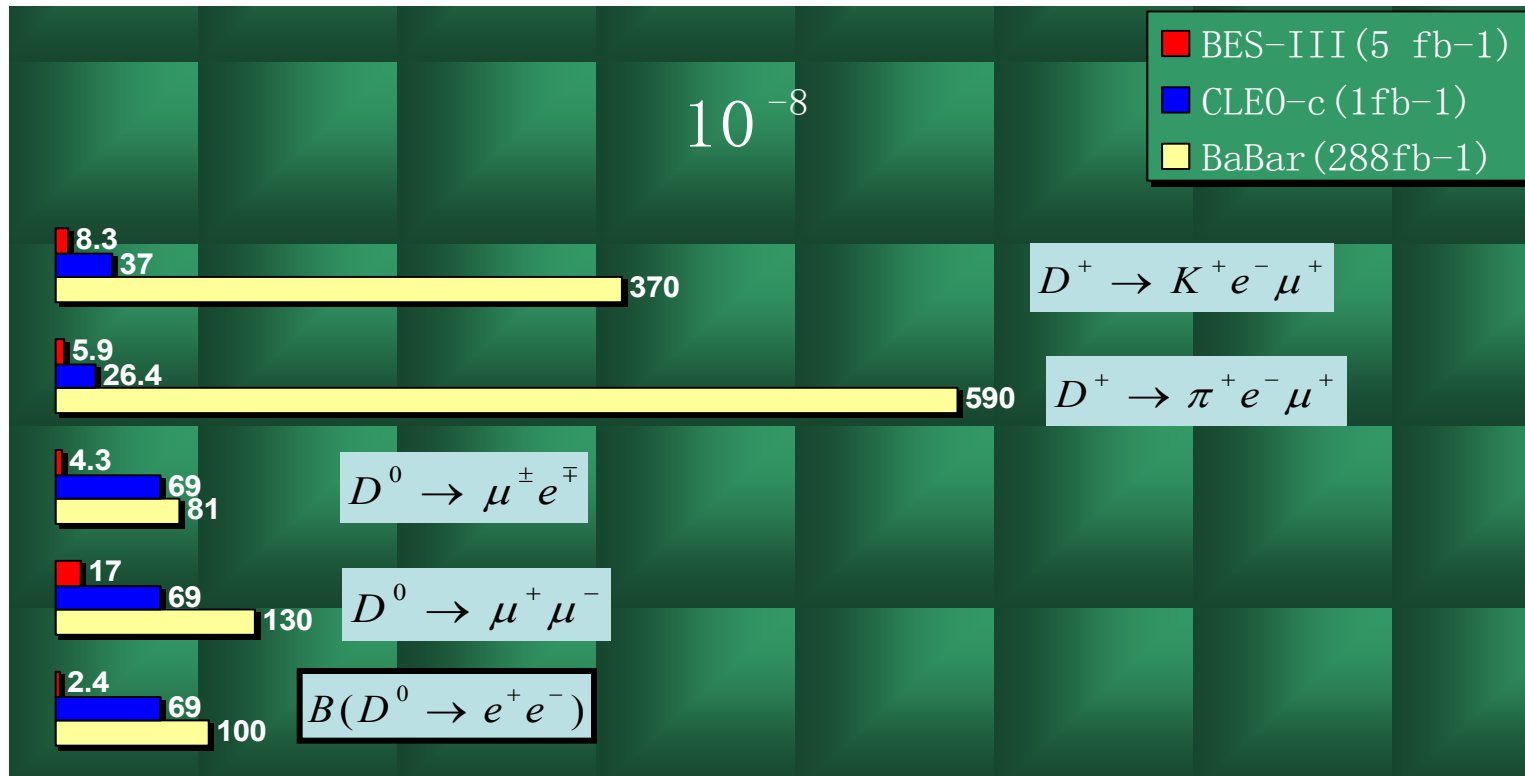
Uncertainties shrink: but are limited by the irreducible model uncertainty (biggest effect on x_D)

Strong phase measurement from $\psi(3770)$ can greatly reduce this.

$$x_D \rightarrow xxx \pm 2.0 \times 10^{-4} \quad y_D \rightarrow xxx \pm 1.2 \times 10^{-4}$$

Sensitivity of LFV

Improve the limits by more than order magnitude!



LFV and LNV are “smoking gun”, any indication of deviation from zero will indicate New Physics (NP).

3. Precision measurement of ν mixing θ_{13}

Daya Bay reactor ν experiment

- Daya Bay nuclear power plant: 4 reactor cores, 11.6 GW
2 more in 2011 for a total of 17.4 GW
- Mountains near by, easy to construct a lab with enough overburden to shield cosmic-ray backgrounds
- Tunnel construction finished. Begin data taking with Near sites middle 2011, Near-Far configuration beginning 2012
- Expect to reach sensitivity of 0.01 with 3 years of running.



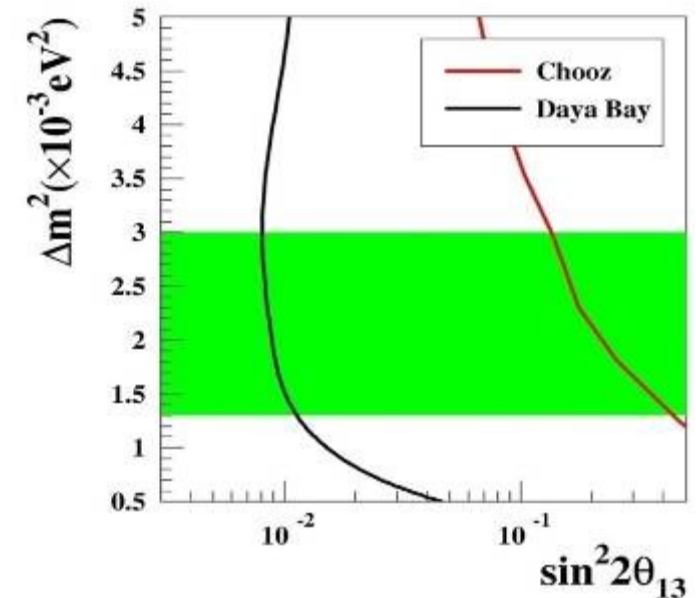
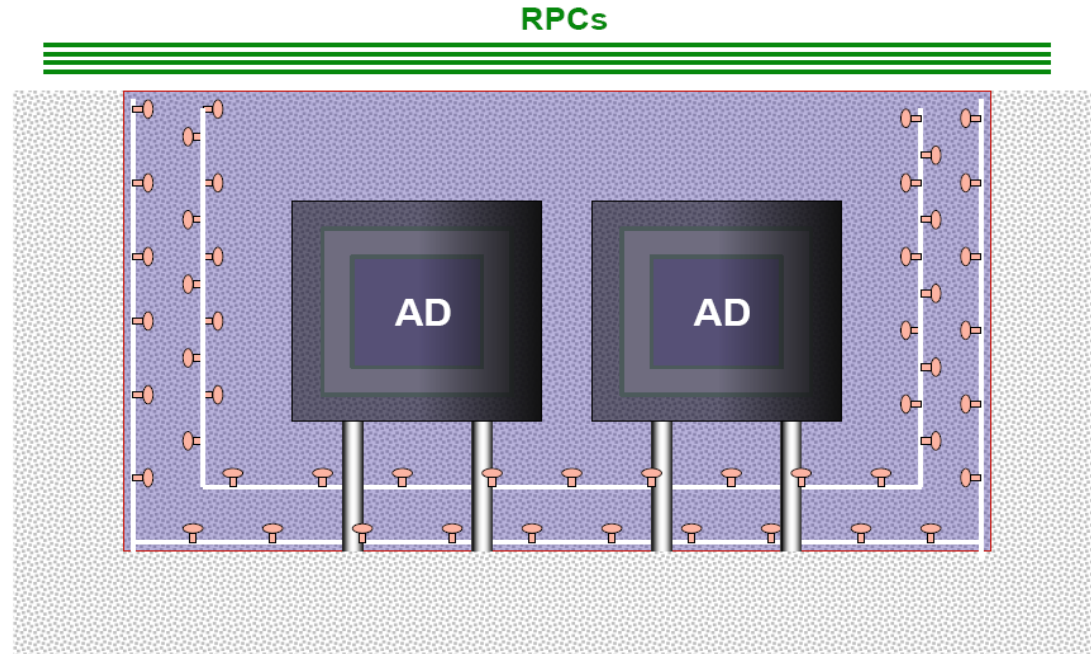
Experimental layout



- Identical detector at near and far site to perform relative measurement in order to cancel reactor related systematic error
- Experimental halls are connected by 3000m tunnel
- Signal rate:
~1200/day Near
~350/day Far
- Backgrounds:
B/S ~0.4% Near
B/S ~0.2% Far

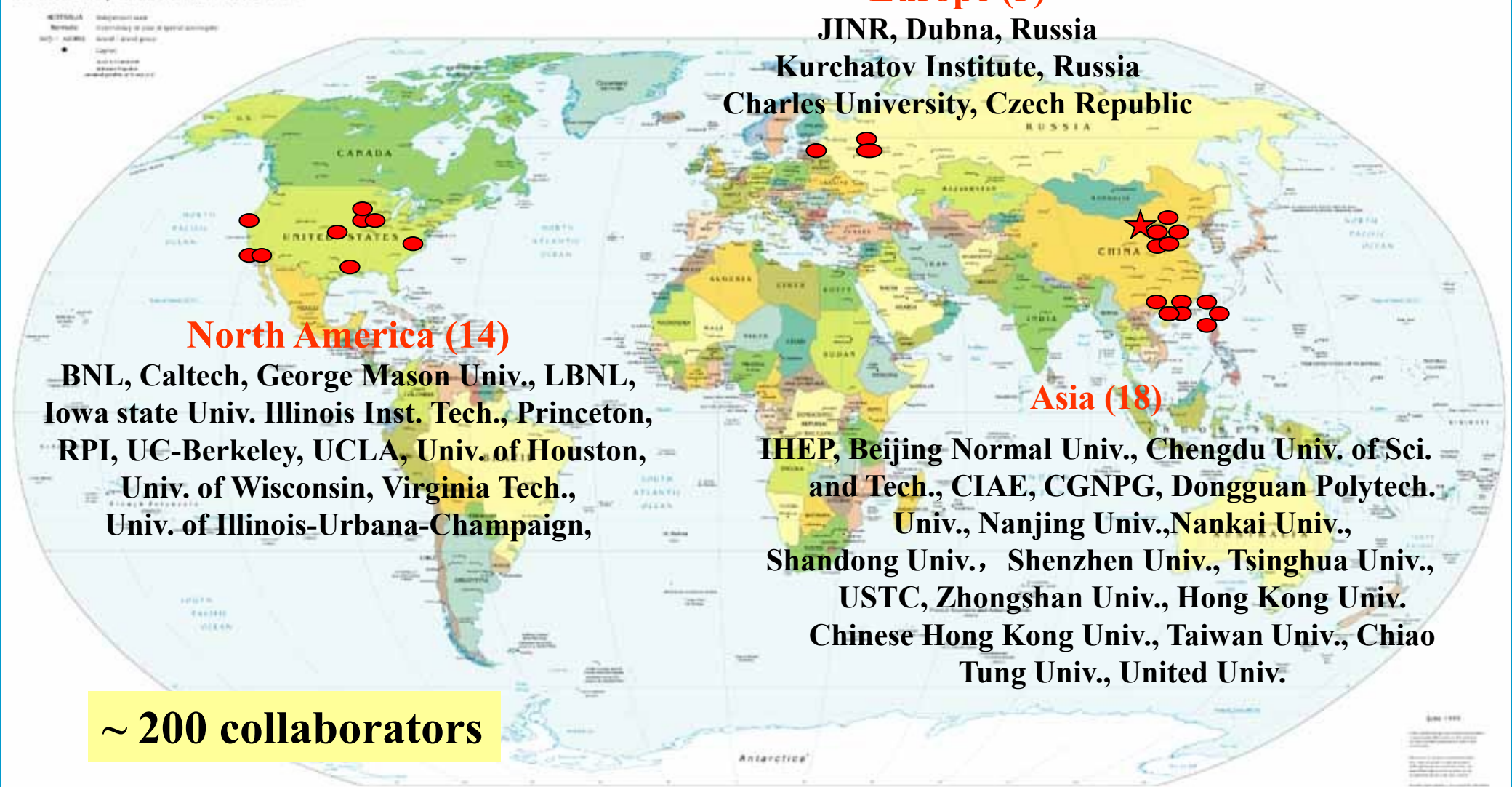
Daya Bay experiment

- Identical near and far detectors to cancel reactor-related errors
- Multiple modules for reducing detector-related errors and cross checks
- Overburden and shielding to reduce backgrounds
- Multi-layer neutrino detector:
 - Gd-loaded liquid scintillator
 - SSV + 2 AV + reflectors
- Multi- muon veto detector for reducing backgrounds and cross checks: RPC+water Cherenkov



Daya Bay collaboration

Political Map of the World, June 1999



Most of Civil Construction finished



Surface assembly bldg.



Control building



HVAC/Ventillation



Tunnel entrance



Tunnel



Experimental hall

Detector components production and assemble

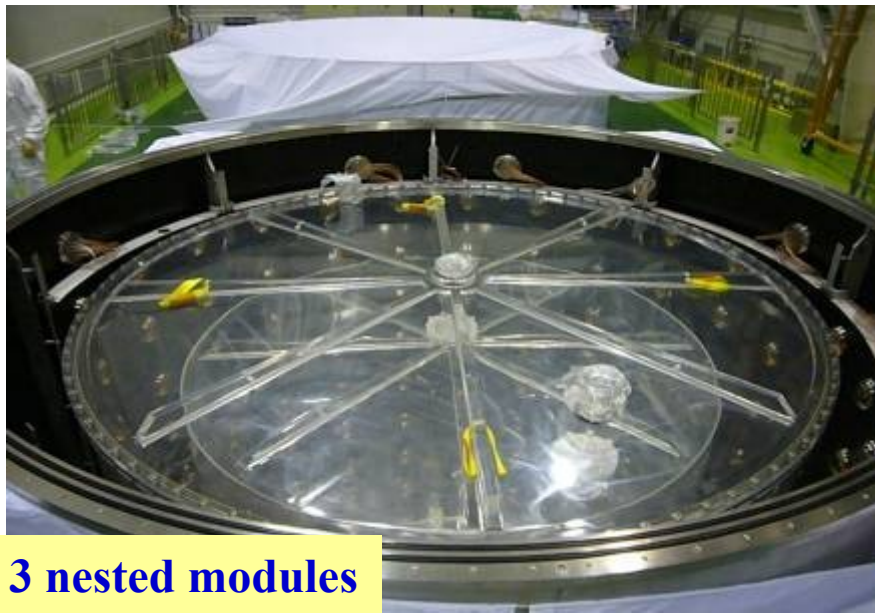
test module successful, two modules finished



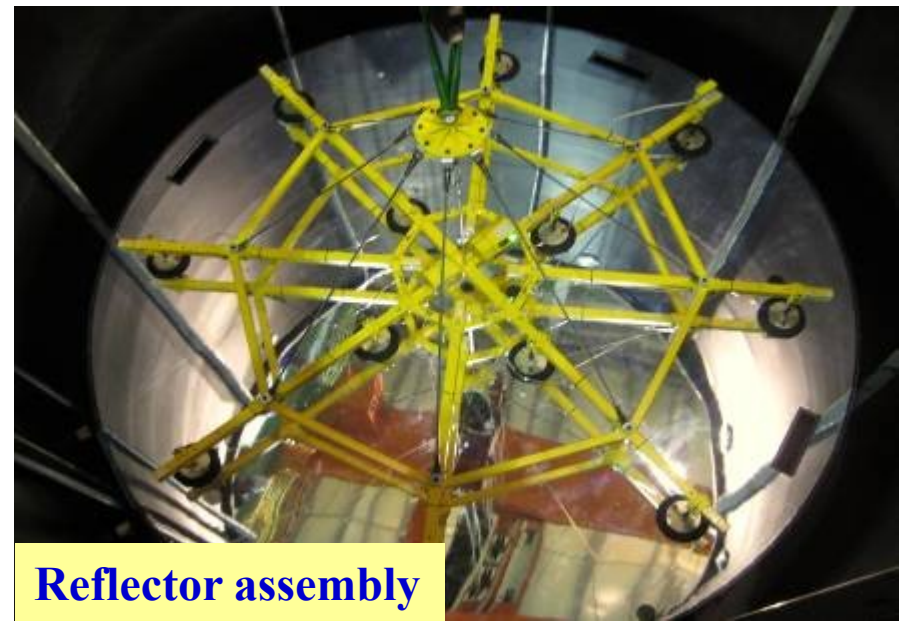
SSV under fabrication, tests include full load, negative pressure, laser survey, etc.



PMT ladder



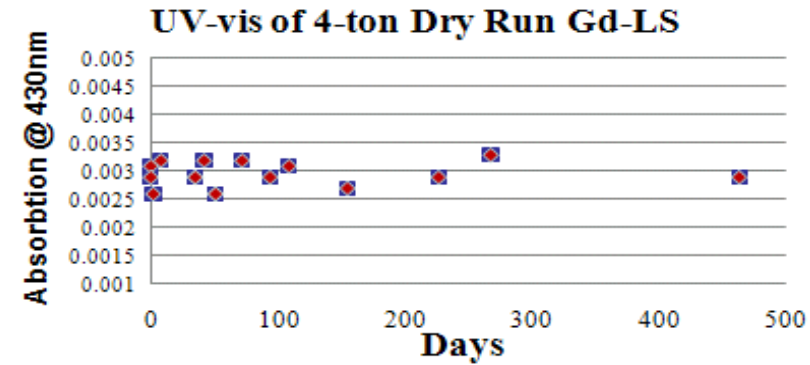
3 nested modules



Reflector assembly

Liquid Scintillator: production started

- ◆ Equipment successfully manufactured and tested at IHEP
- ◆ 4-ton 0.1% Gd-LS dry run successful
 - ⇒ The Gd-LS in the tank looks good up to now.
 - ⇒ High temperature aging tests show good stability
- ◆ All mixing equipments(including QA and TMHA purification system from BNL) installed and tested in Hall 5
- ◆ Mass production successfully started



4. Particle Astrophysics Experiments

- **Experiments finished :**
 - γ Bust detector @ Shenzhou-2 flown 2001, First Astronomy detector of China in space
 - Chinese Moon project: Change-1, flown 2007
 - L3 Cosmic
 - AMS01: permanent magnet
- **On-going Experiments:**
 - Yangbajing experiment Asy & Argo
 - Chinese Moon project: Chang'e-2: launch Oct. 2010
 - AMS02: permanent magnet, ECAL.
- **Approved Experiments:**
 - Hard X Ray Modulate Telescope: Launch by 2014
 - Chinese Moon project: Chang'e-3
 - SVOM: China -France (Satellite)

4a Yangbajing Cosmic Ray Observatory (Tibet a.s.l. 4300m)

IHEP-INFN Argo RPC

China-Japan Air Shower Array



Yangbajing: new anisotropy and corotation of GCR

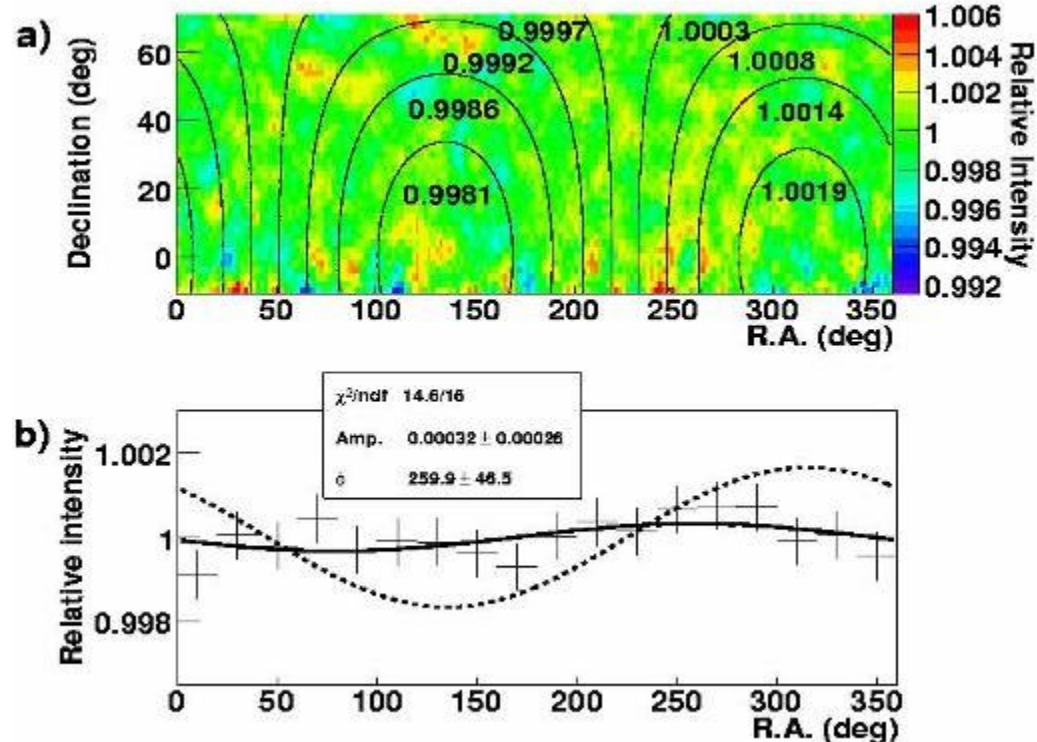
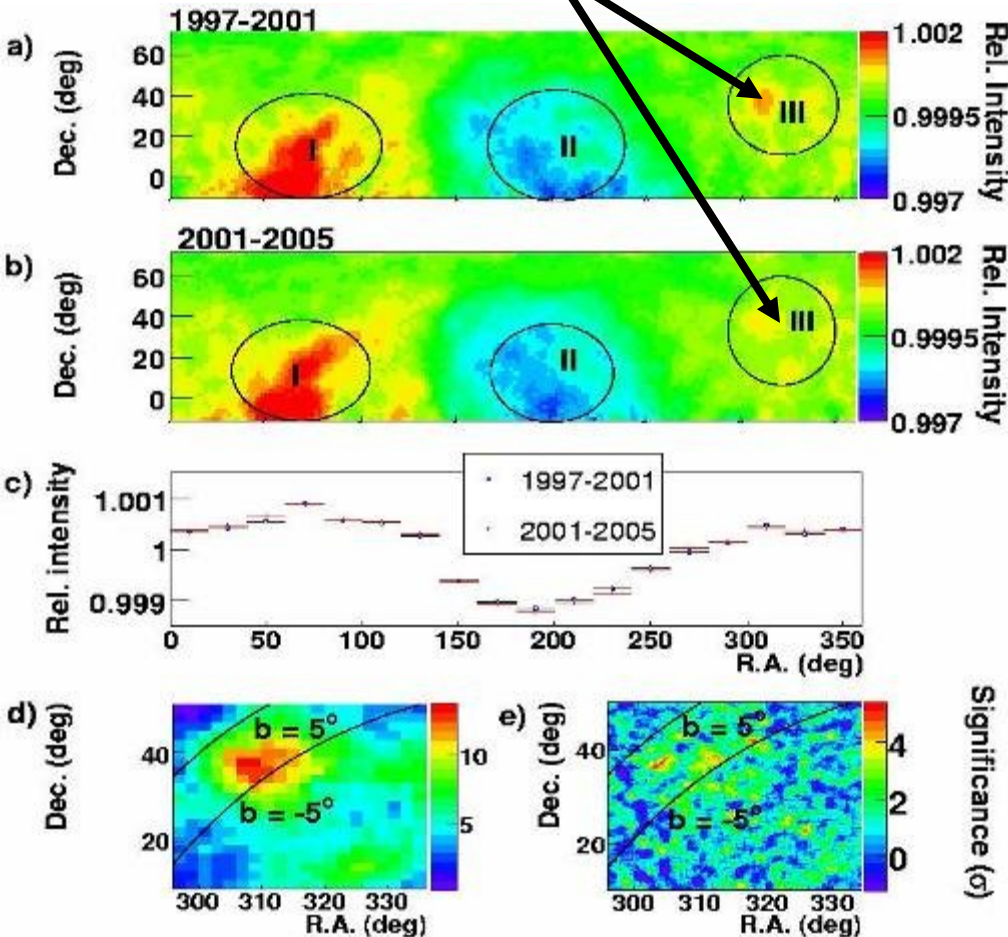
(Science 314(2006) 439-443)

Celestial Intensity map ($E \sim 3\text{TeV}$)

Intensity @ $E \sim 300\text{TeV}$

New anisotropy component

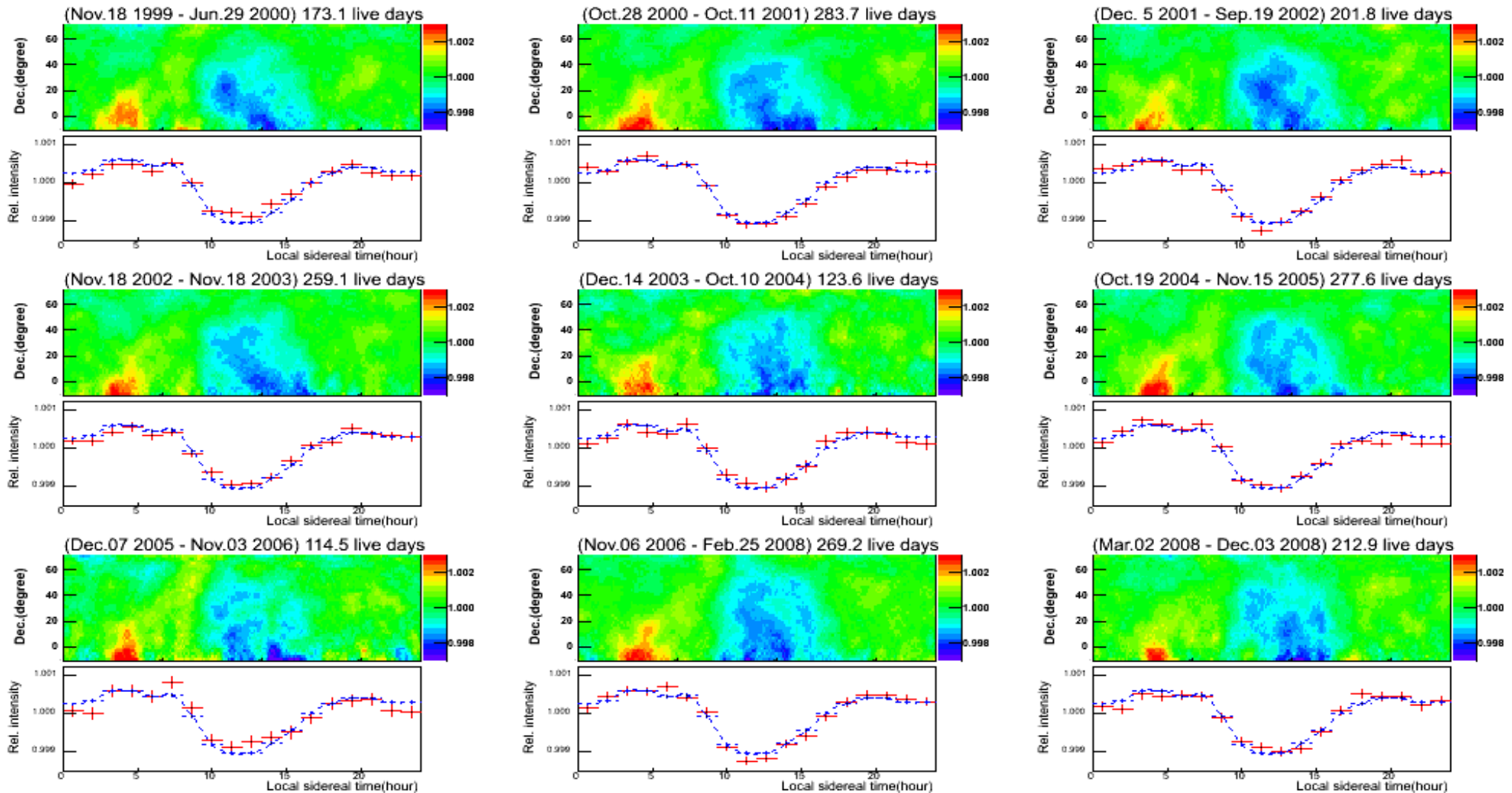
Amp=0.16% w/o corotation;
Observation: $0.03\% \pm 0.03\%$.

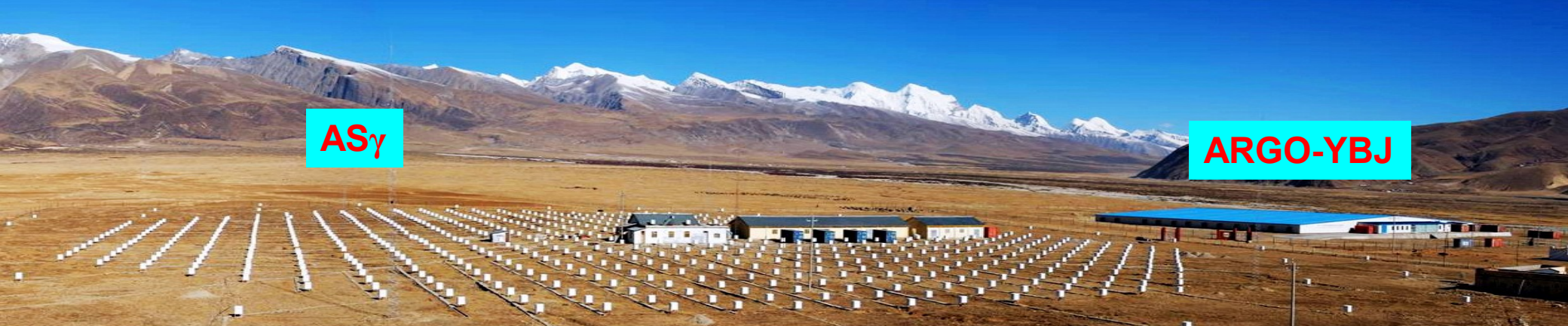


Yangbajing Asy: Stable Multi-TeV anisotropy

- insensitive to solar activity

ApJ 711(2010)119-124

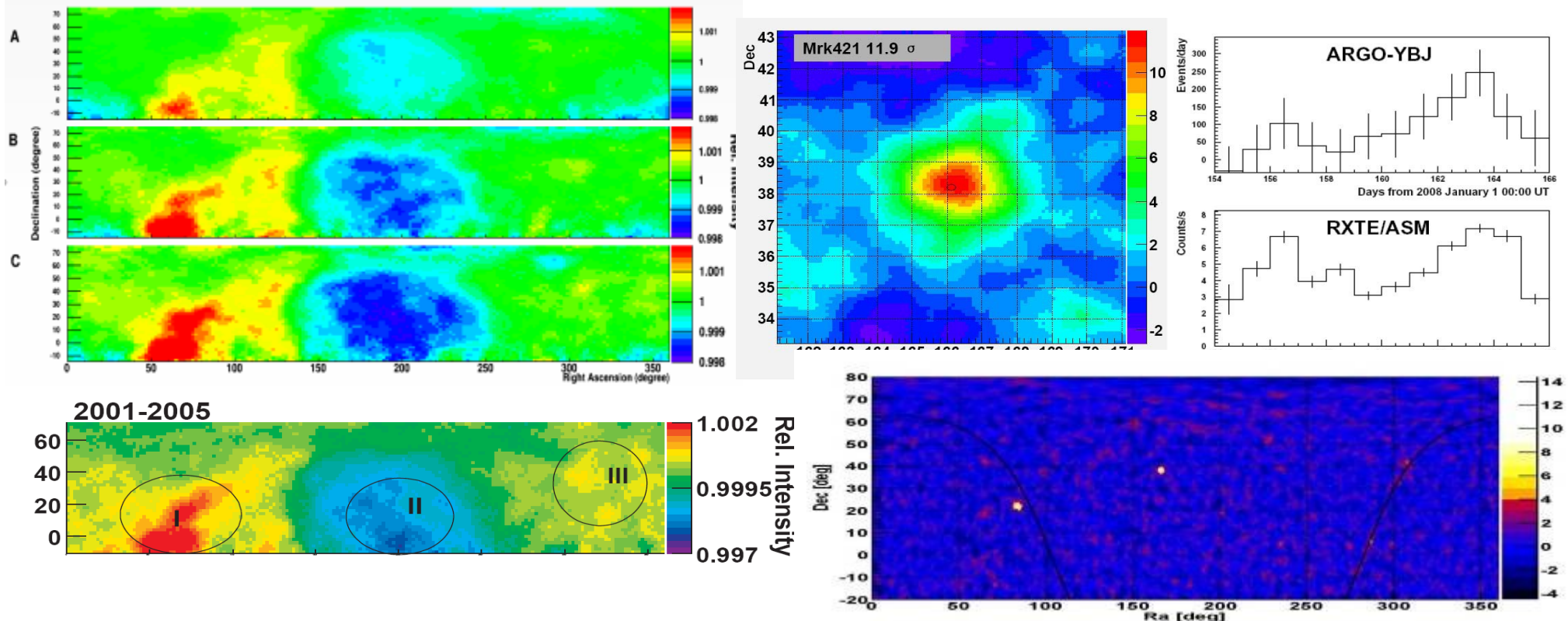




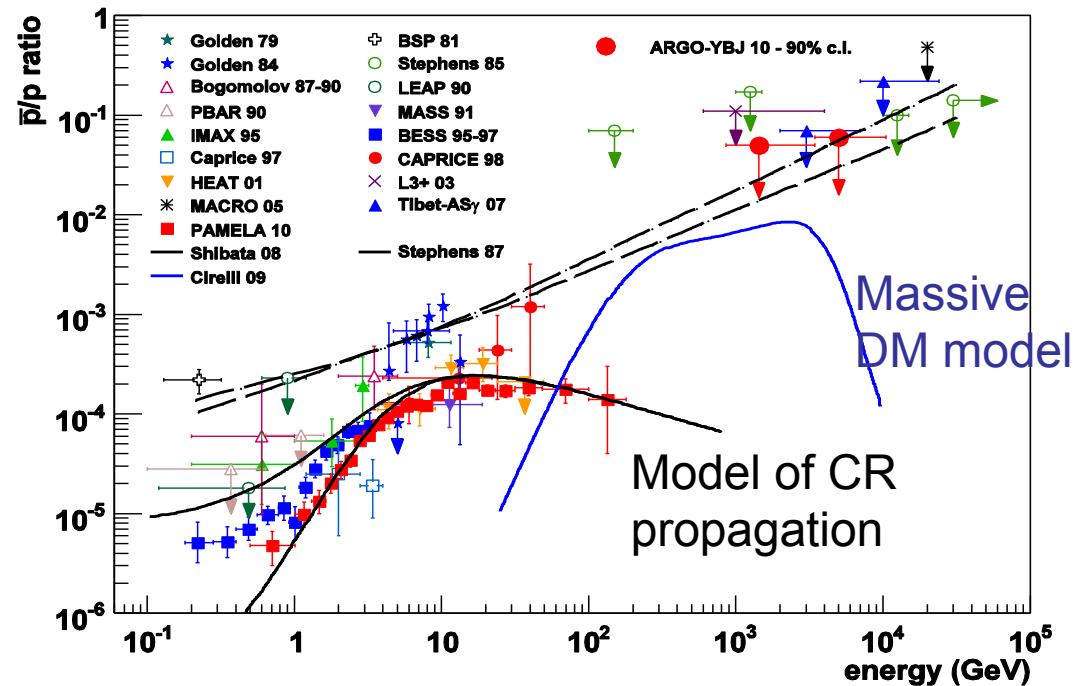
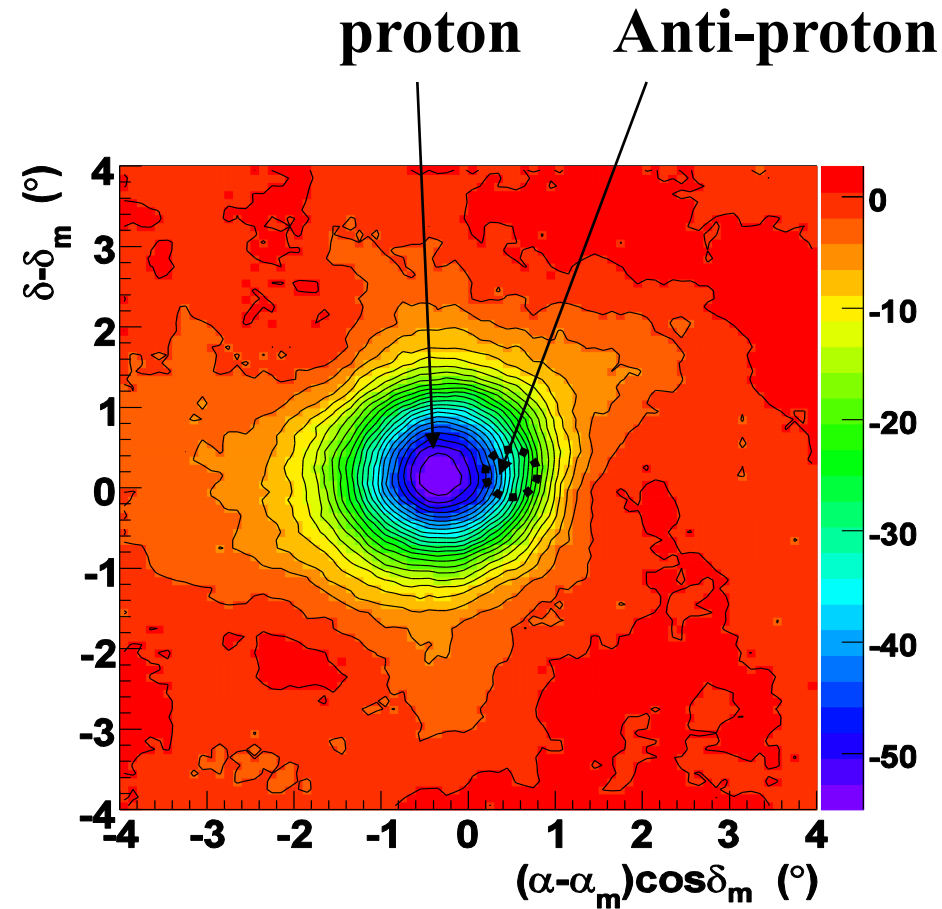
AS γ

ARGO-YBJ

Cosmic ray anisotropy and its energy dependence (top down: 0.5~5TeV)
 Long term monitoring for AGN Mrk421 bursts: accumulative significance 12S.D.
 Multi-wave-length analysis with X-ray observation demonstrates strong correlation

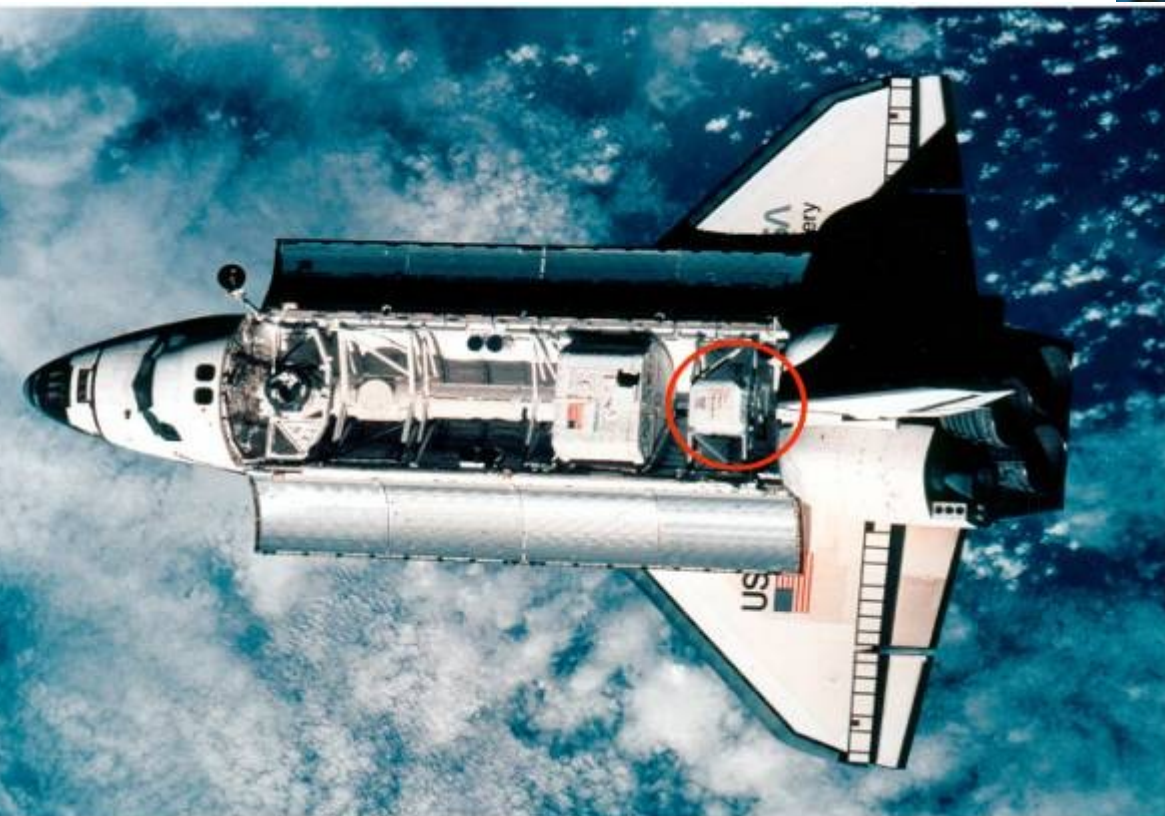
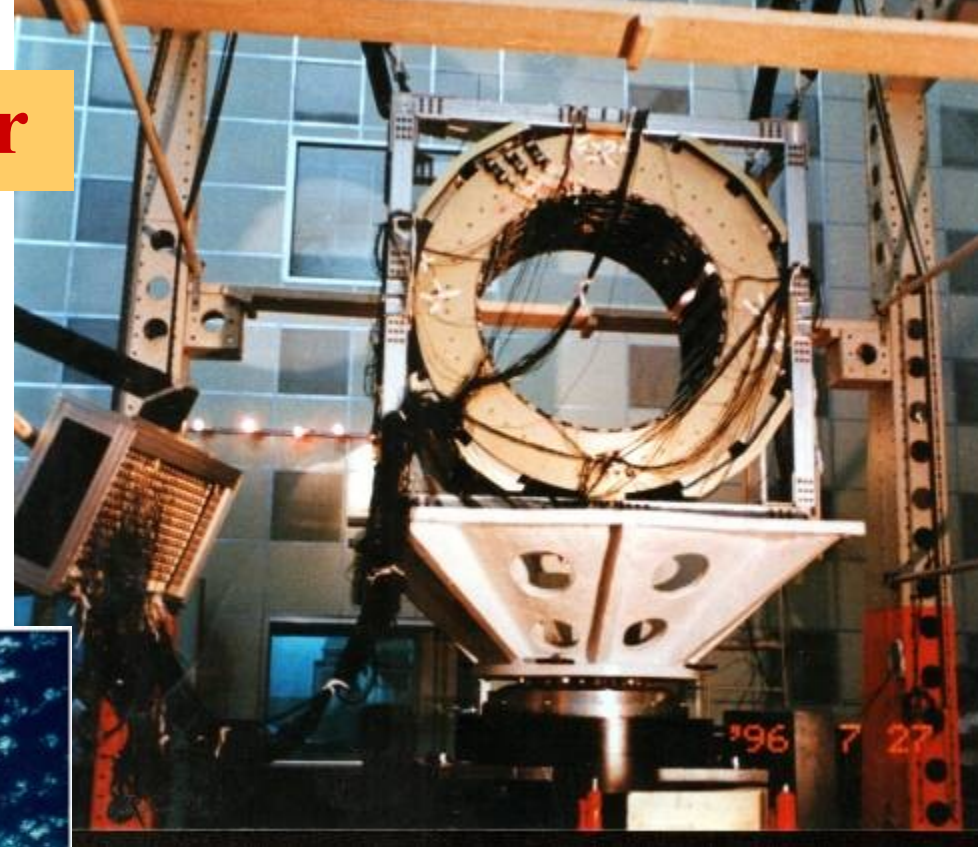


ARGO-YBJ : \bar{p}/p Ratio Upper Limit by using the displacement of moon shadow in geomagnetic fields



Alpha Magnetic Spectrometer

- Search for antimatter and dark matter
- precision measurement of isotopes



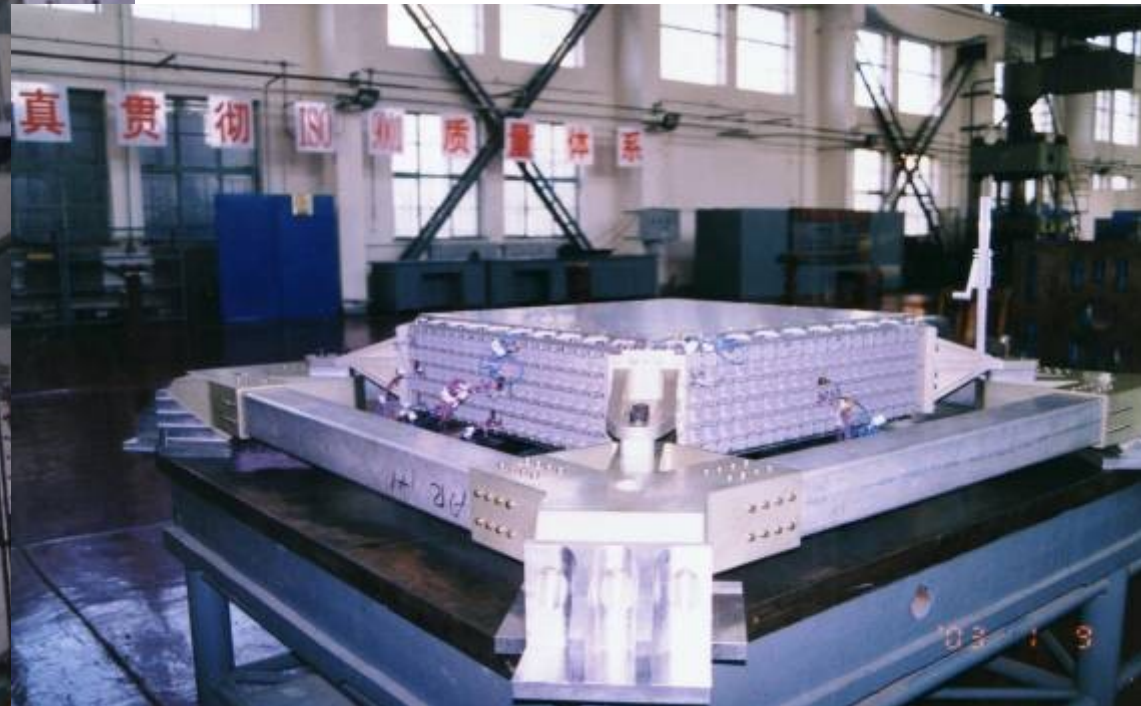
AMS01 permanent magnet and structure were built at Beijing, and became the first big magnet in space as payload of Discovery June 1998. It will be flown again by AMS02 Feb. 2011.



AMS02 ECAL: 700Kg IHEP LAPP and PISA

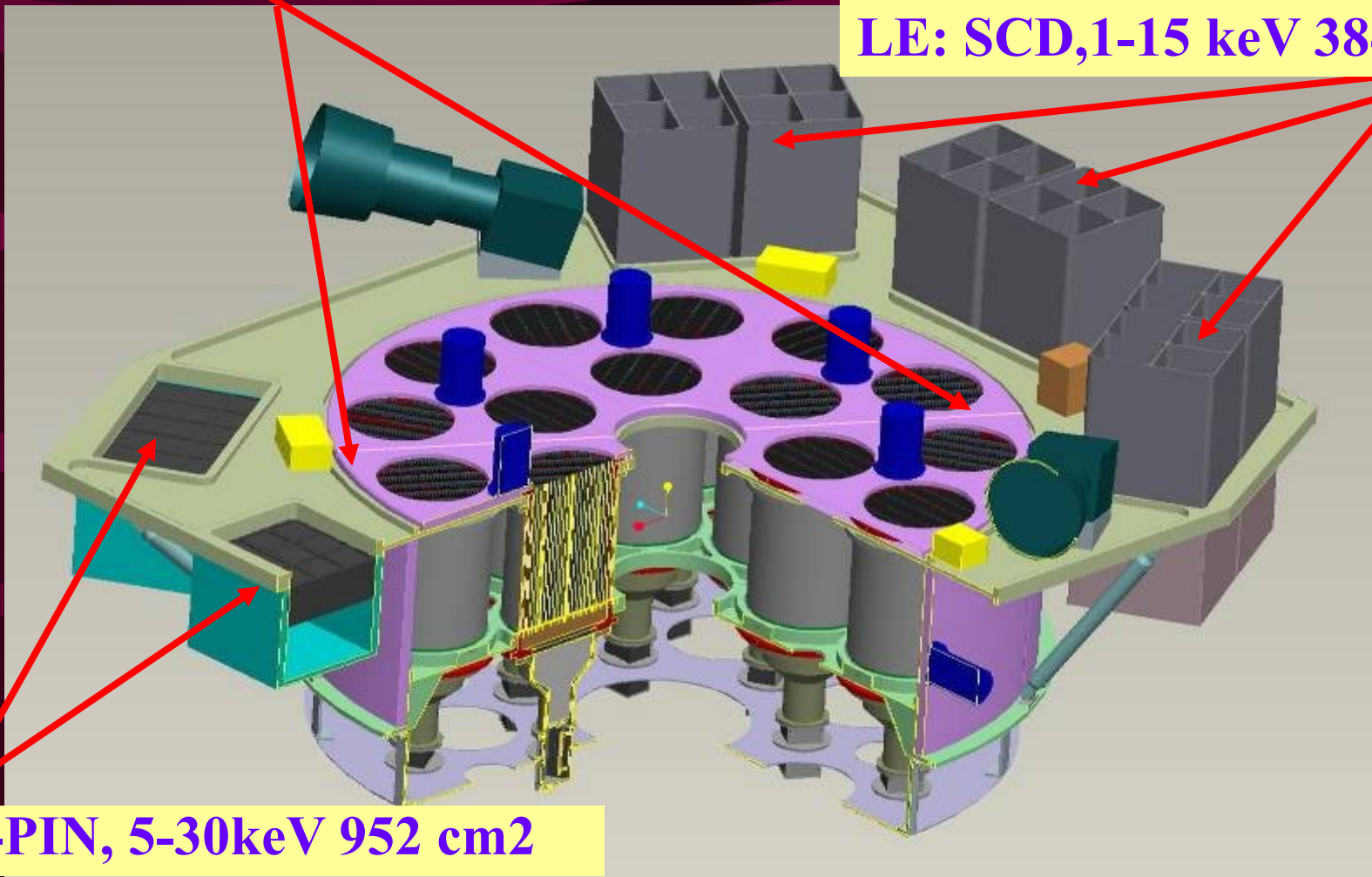
Space qualification at Beijing

ECAL assembling at IHEP



HE: NaI/CsI 20-250 keV 5000 cm²

LE: SCD, 1-15 keV 384 cm²



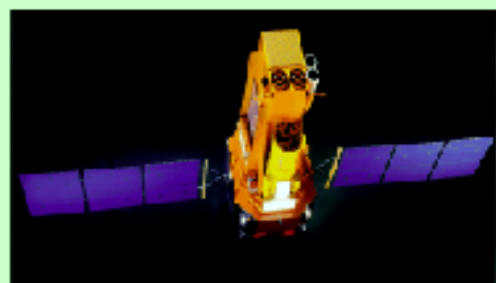
ME: Si-PIN, 5-30keV 952 cm²

Hard X-ray Modulation Telescope (HXMT)

To be launched 2014 Satellite 2700 kg

Comparison of HXMT and other two telescopes in the same energy band.

Integral/IBIS



HXMT/HE

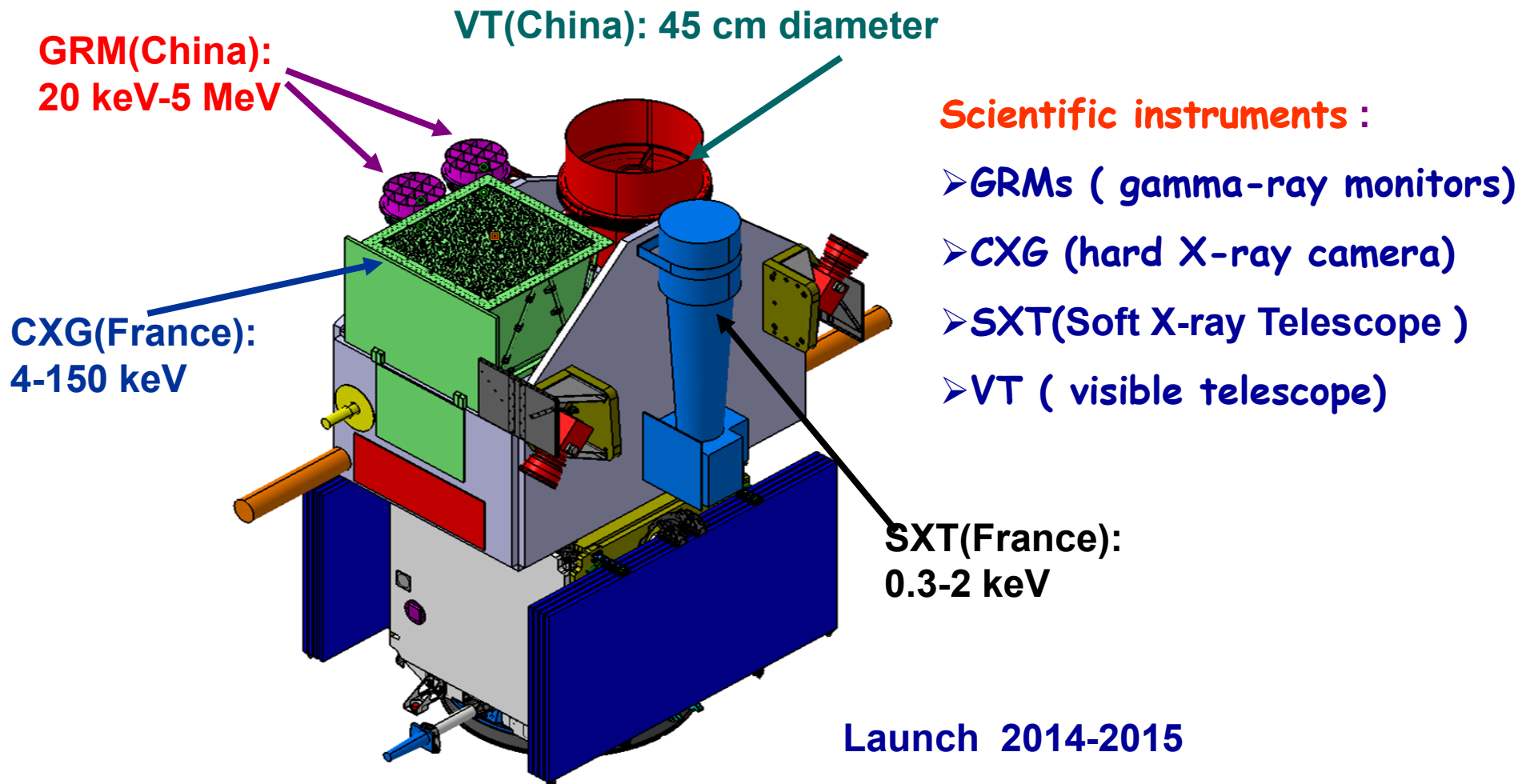


wift/BAT



Angular Resolution	12'	< 5'	14'
Source Location (20σ)	1'	< 1'	1'
Pointed Sensitivity (mCrab@100 keV)	3.8	0.5	9
Half Year Survey Sensitivity (mCrab)	2	0.5	1
Observation Capability			
All sky survey	ok	good	yes
Selected sky deep survey	good	good	bad
Narrow field pointing observation	bad	good	no

SVOM: multi-wavelength GRB project



China-France collaboration: approved 48

5. Future Plan

- **Intensity Frontier**
 - Charm physics @ BEPCII: next 10 years or more.
Future plan to be decided based on results of BESIII & LHC in ~ 5 - 6 years
 - Neutrino experiments: Daya Bay II . (> 2016)
- **Energy Frontier:** Intl. collab. @ LHC exp., ILC/CLIC...
- **Cosmic Frontier :**
 - Particle Astrophysics experiments in Space: Polar
 - Cosmic ray measurement : **LHAASO @ Yangbajing**
 - **Deep underground Lab.:** Jinping Mountain
 - **South pole Dome A:** IR and THz telescope
- **Large Scientific Facilities**
 - Chinese Spallation Neutron Source (CSNS): approved.
 - Accelerator Driven Sub-critical System: CAS is planning the ADS pilot effort.
 - Beijing Advance Light Source: construction scheduled at 2016

LHAASO Project: γ astronomy and origin of CR

Charge Particle Array
 μ detector Array

Water C Array

Wide FOV C-Telescope Array

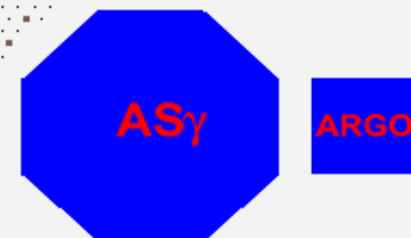
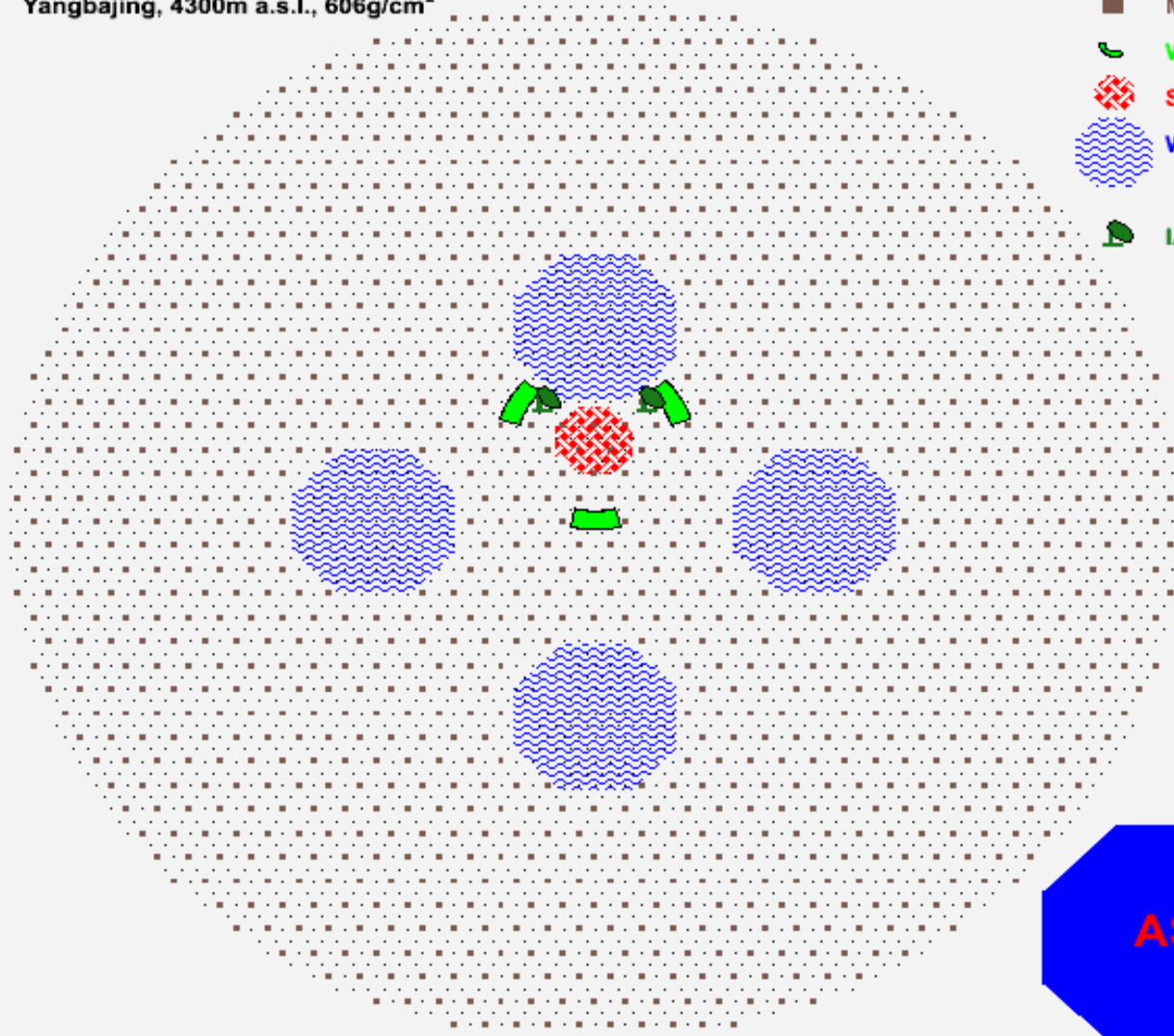
&

Core Detector Array

Large High Altitude Air Shower Observatory

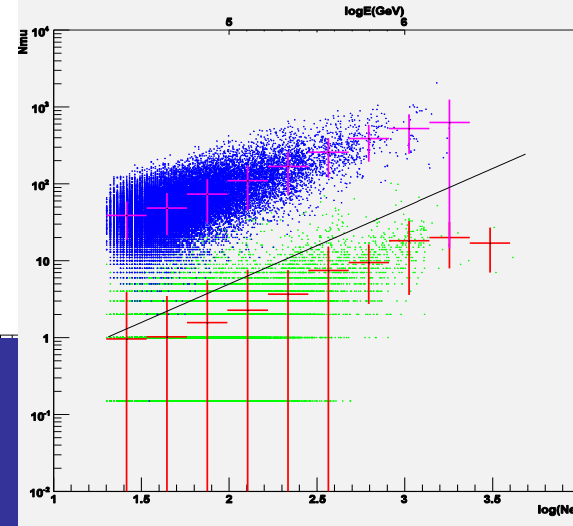
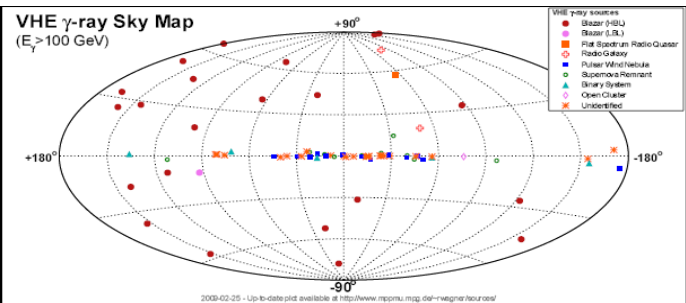
Yangbajing, 4300m a.s.l., 606g/cm²

- ED: 5137, 1m \times 1m \times 2cm
15m spacing
- MD: 1161, 6m \times 6m \times 2cm
30m spacing
- WFCAs: 3 \times 8, 16 \times 16pixels
130m spacing
- SCDA: 5000m² (Φ 80m)
- WCDA: 4 \times 900
 Φ 170m \times 4m
300m spacing
- IACs: 2
100m spacing

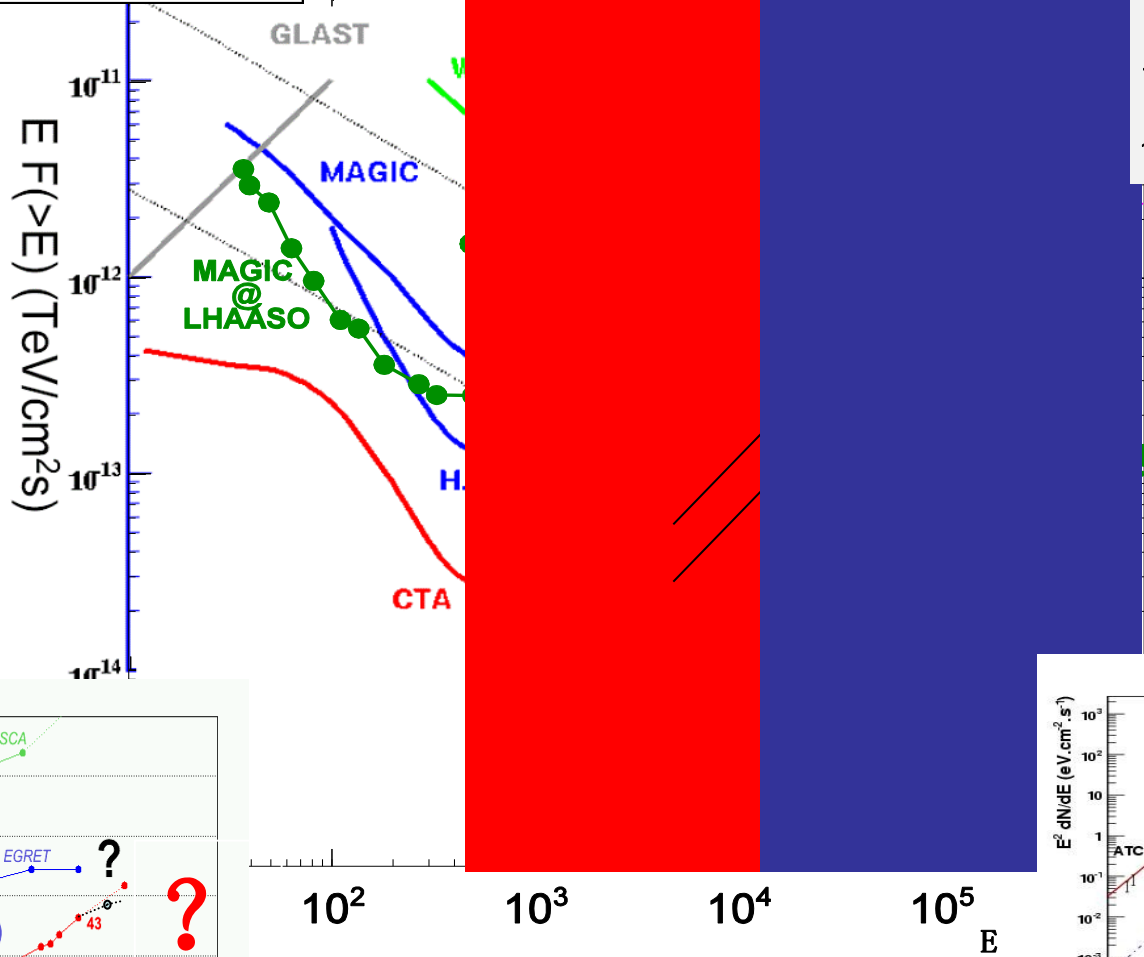


1000m

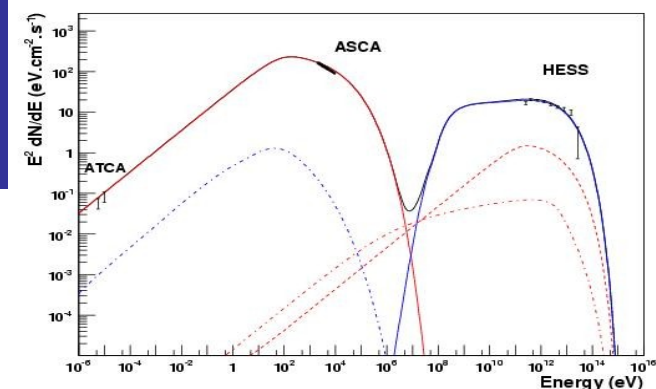
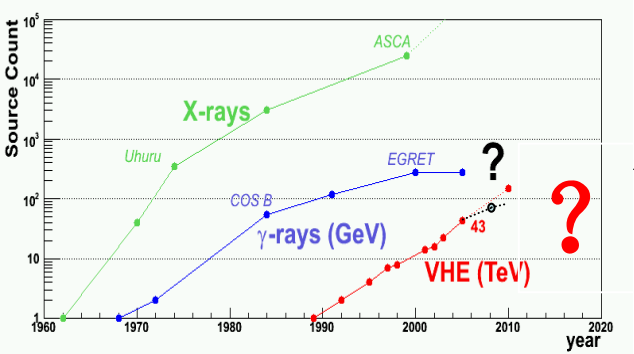
LHAASO sensitivity & scientific goals



- Surveying for γ ray sources
- ~ 1000 extragalactic sources are expected in 10 years



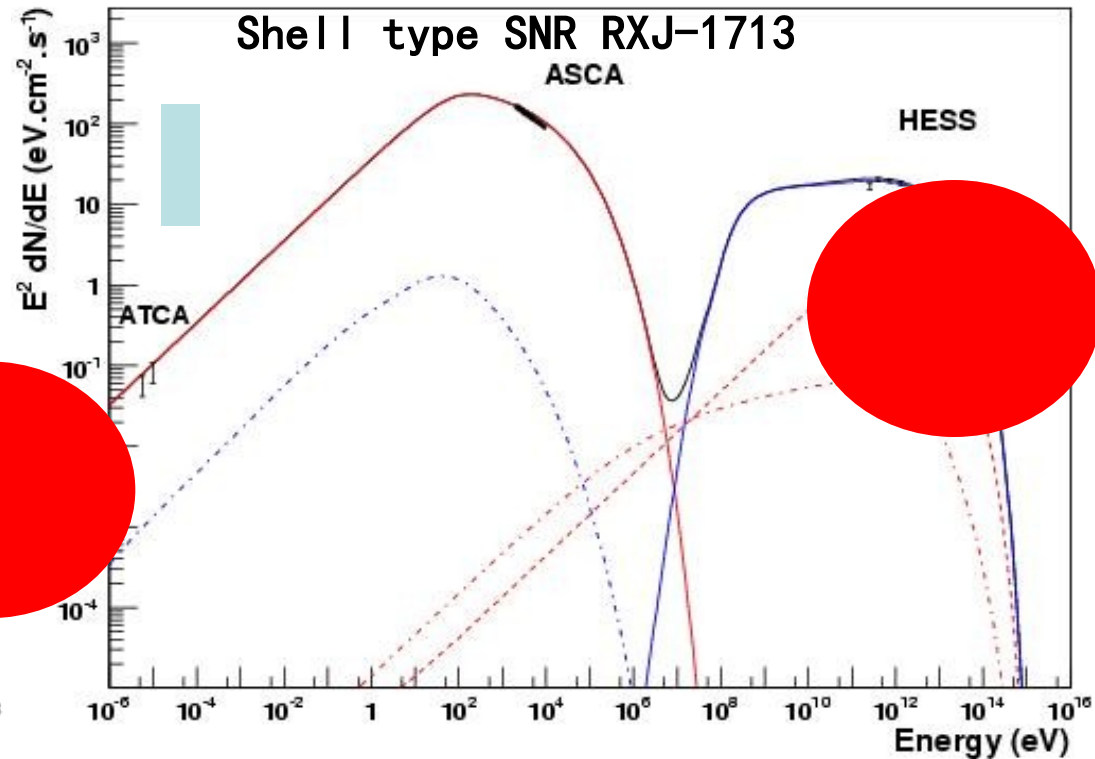
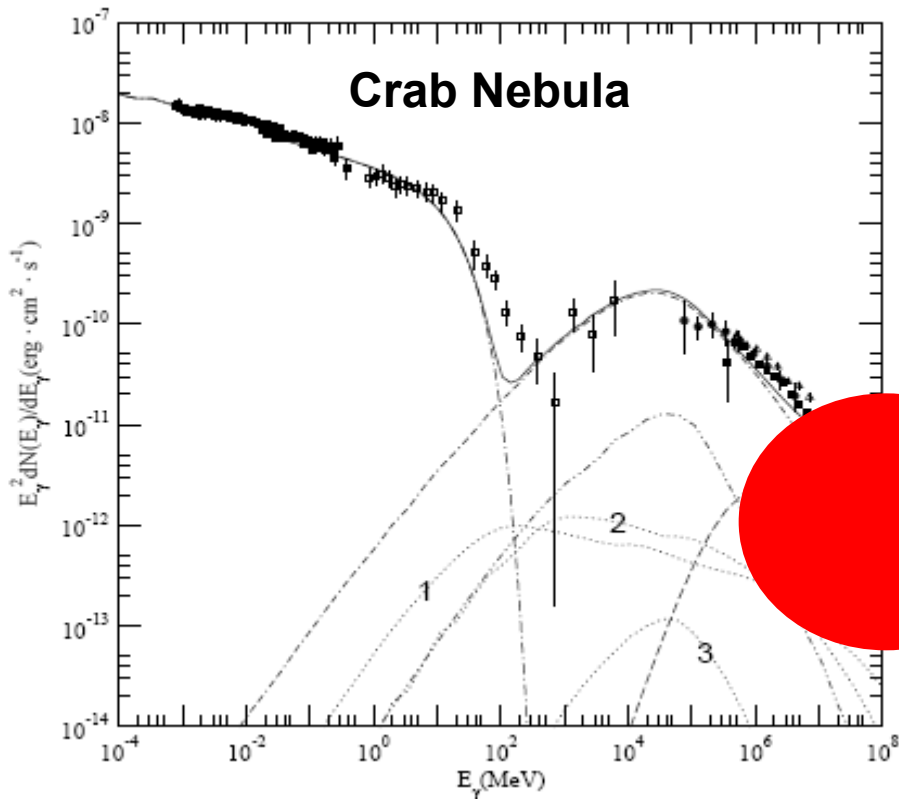
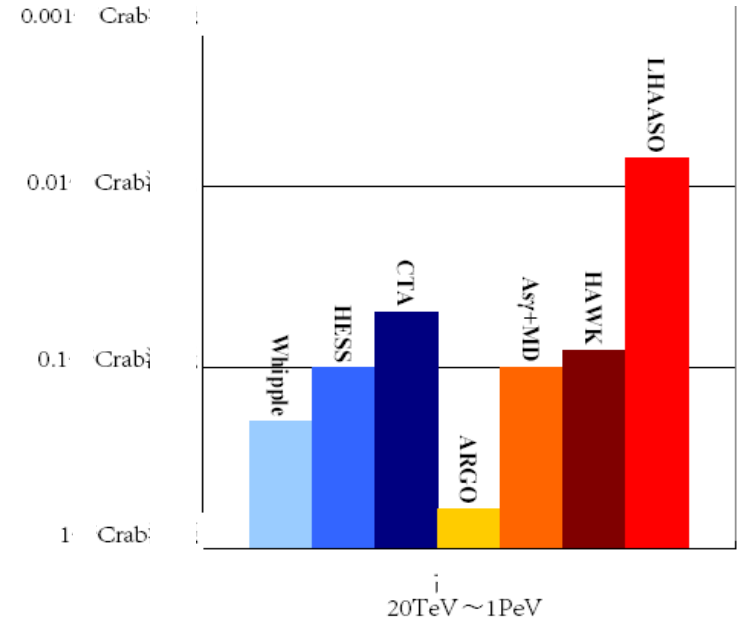
- Above 60 TeV CR BG-free (10^{-5})
- γ survival rate $\sim 99\%$
- Finding Cosmic ray origins.



Main Goal 1: CR origin

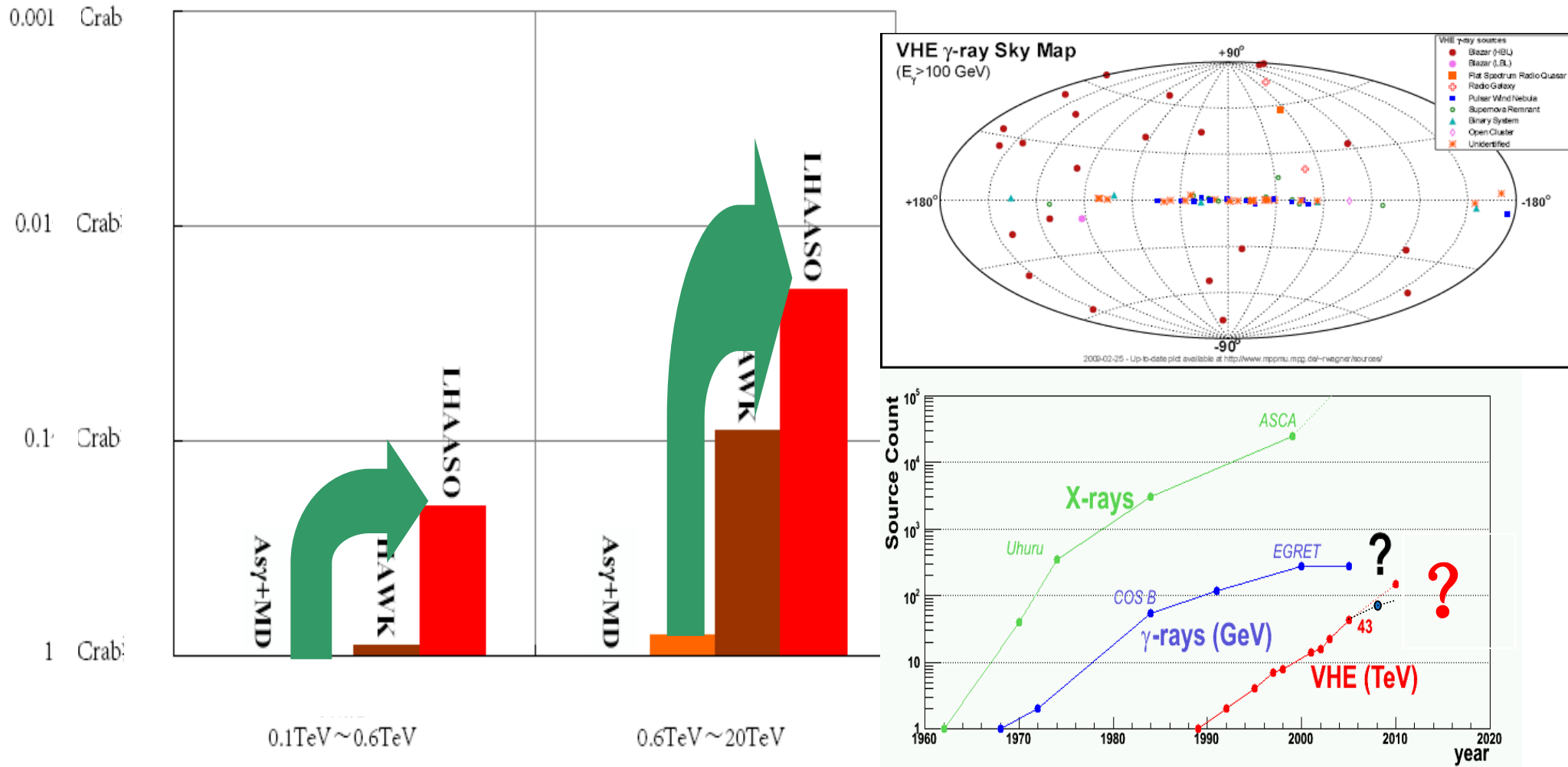
precise measurement of energy spectrum of γ sources to 1PeV

Smoking gun: photons from $\pi^0 \rightarrow 2\gamma$



Main Goal 2: all sky survey for more extragalactic gamma ray sources

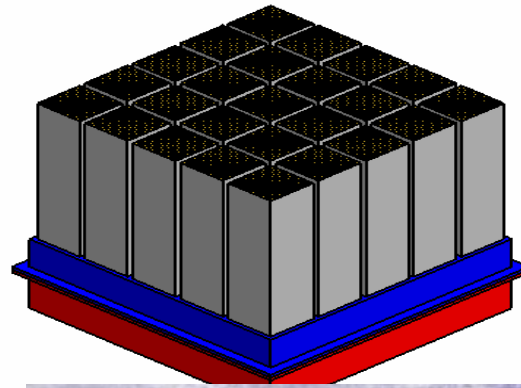
Very high surveying sensitivity: from 100 sources to 1000



POLAR: Proposal and R&D

- Instrument conception proposed by N. Produit, et al., NIM (2005)
- On board China's spacelab TG-2: launch time 2013-14 (Phase 2 of manned spacecraft)
- FOV of POLAR: $\sim\frac{1}{2}$ sky
- MDP is 10%: >10 GRBs per year down to 10% polarization;

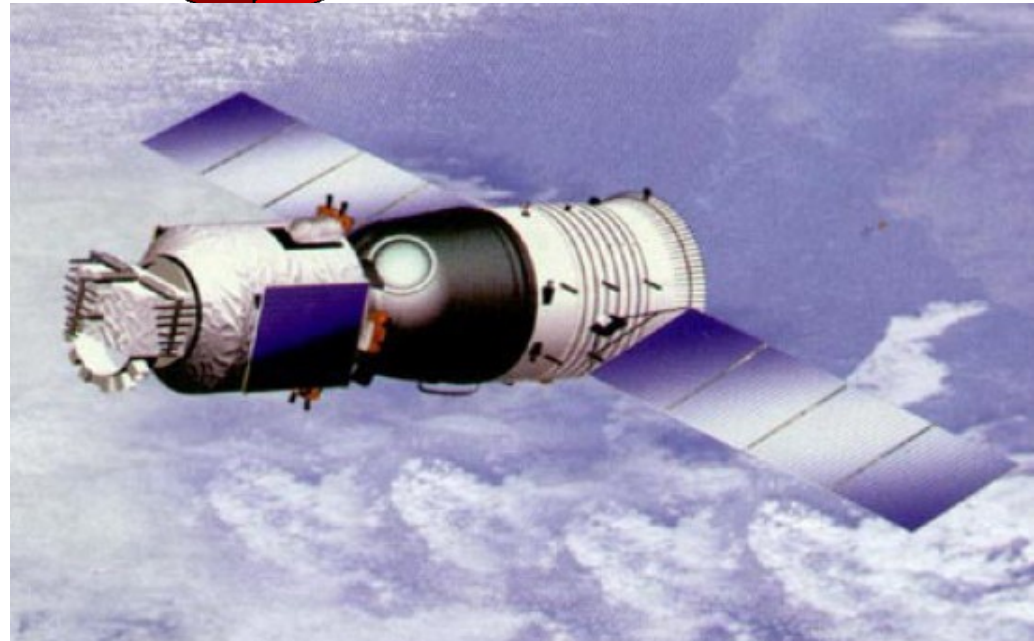
POLAR



Tian-Gong

天宫

Palace in
Heaven



China Deep Underground Lab. (CDUL) at Jingpin mountain

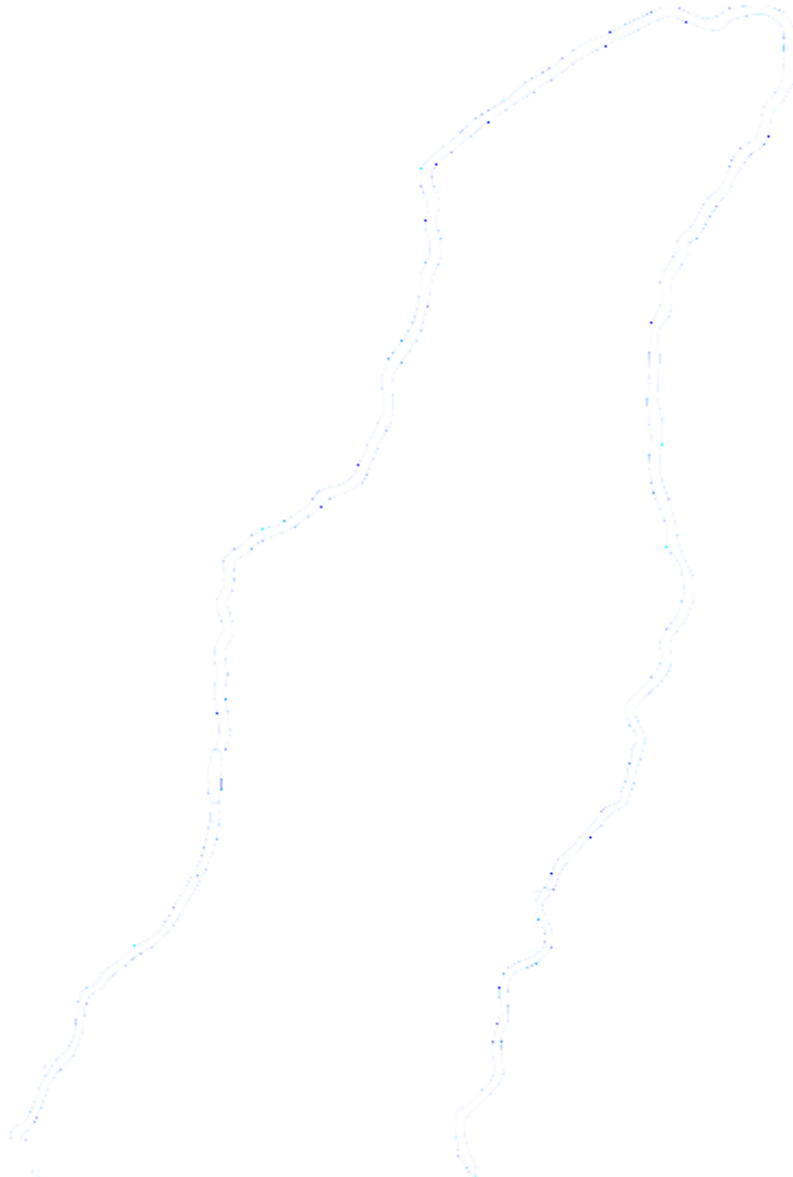


Yalong River and Jingpin Hydropower station



Yalong river locates the west of Sichuan, with length of 1571km and drop height of 4420m. 21 Hydropower stations are planned, with total electricity of 30GW.

Big U-turn at Jingpin mountain



- Tunnel length of 17km
 - 4 water tunnels: ϕ 13m
 - 2 traffic tunnels: ϕ 6m
 - drain tunnel ϕ 7.2m (only used during construction)
- max. overburden 2525m
- Rock radioactivity is very low.
- Horizontal Tunnel access
- Convenient transportation, good infrastructure

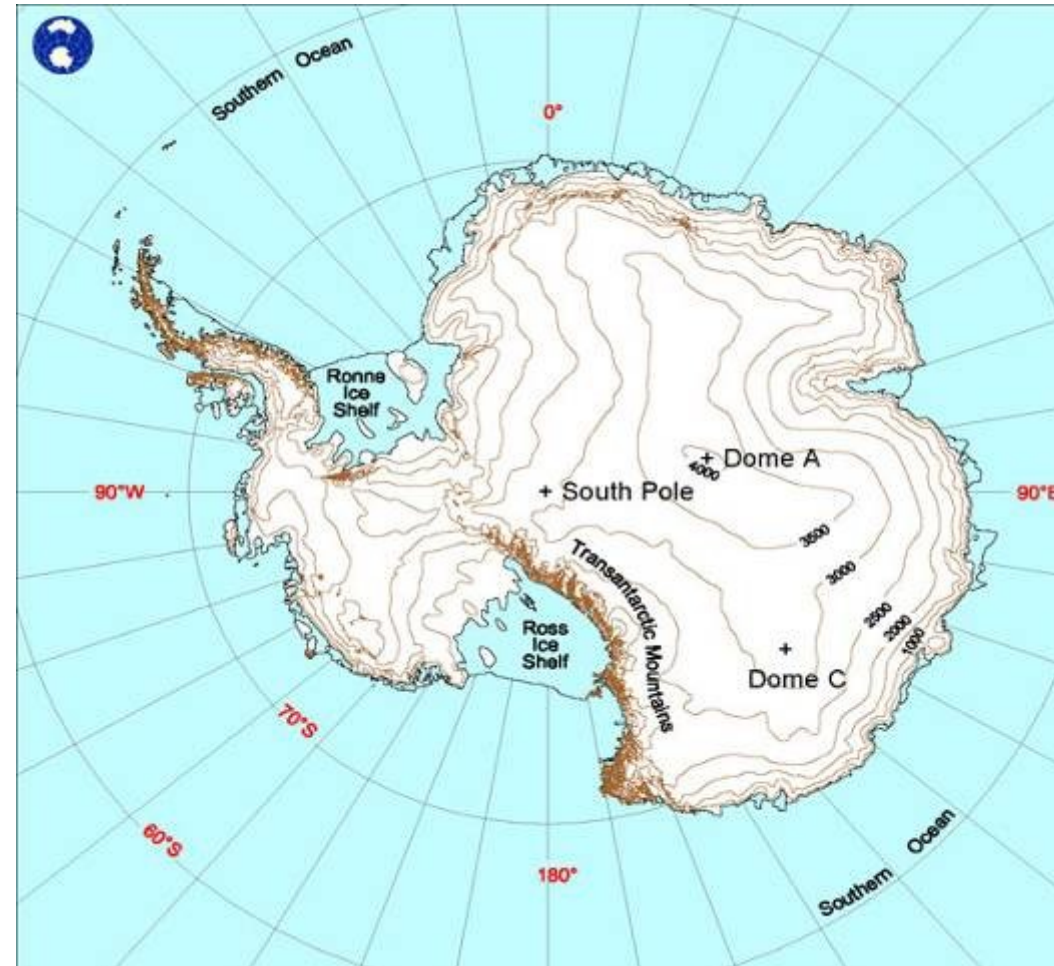
Deep underground Lab. @ Jingpin mountain

- **Unique opportunity for world class deep underground science and engineering lab for multiple science research.**
 - **Physics Exp. Proposals: dark matter search, double beta decay, proton decay, nuclear astrophysics reaction, gravitation wave...**
 - **Deep underground engineering**
 - **Geophysics and geology**
- **Two 6x6x20 m exp. halls near traffic tunnel B under construction: dark matter exp., & rock mechanics studies**
- **CAS proposal: large deep underground multiple discipline international Lab. . Under review.**

4m Telescope at Dome A of South Pole

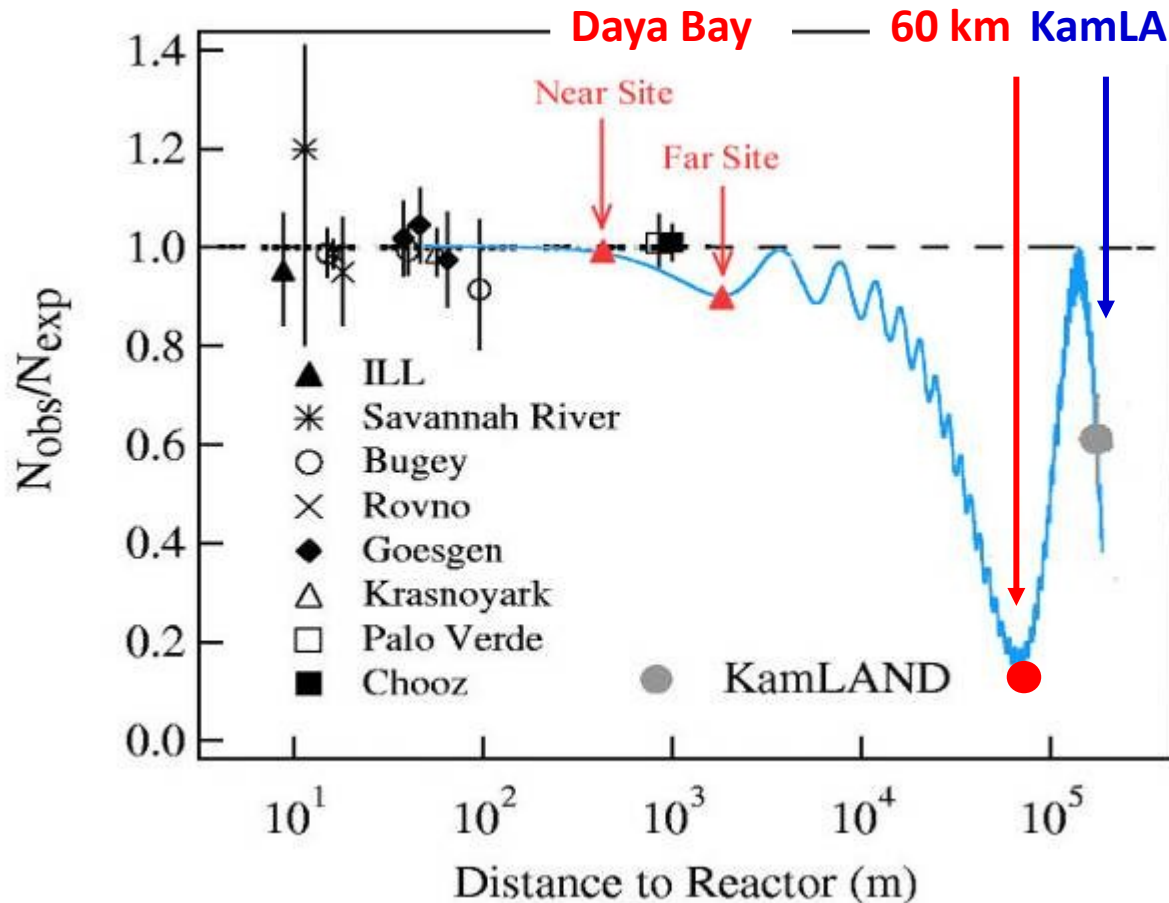
space-quality imaging with much lower cost

- 4100m, -83°C, Highest, Driest, Coldest on earth
- Very Quiet Atmosphere, boundary layer is 9m, seeing = 0.27 arcsec
- Very good transmission from 150 to 800 μm .
- infrared bg. factor 20-100 less
- perpetual nights > 135
- possible for maintenance
- Infrared and THz: most interesting
- 15cmx4 telescope since 2006
- proposal under review



Daya Bay-II Experiment

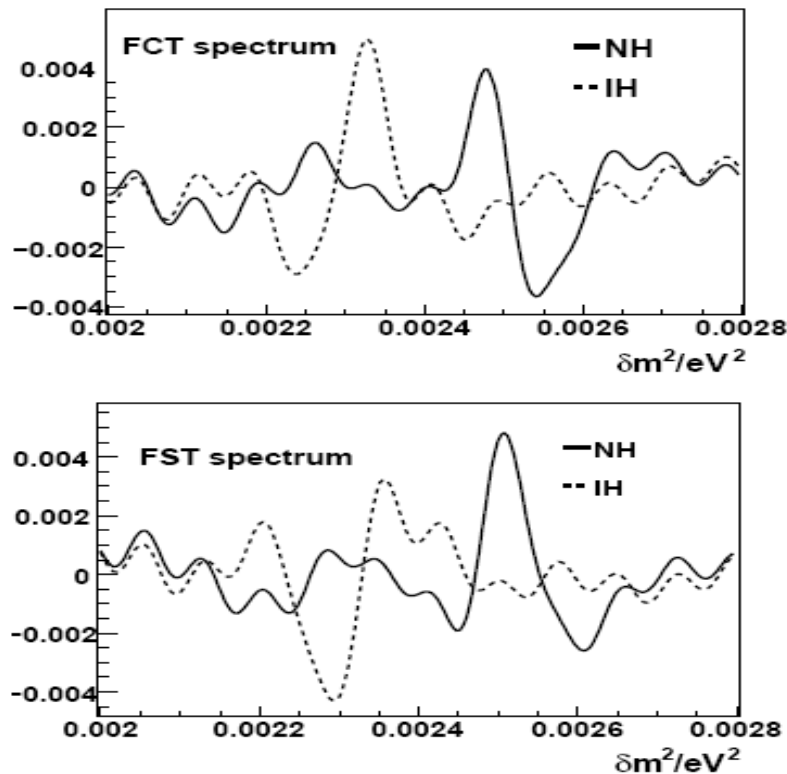
Giant Detector located at 60 km from Daya Bay reactors, the 1st maximum of θ_{12} oscillation.



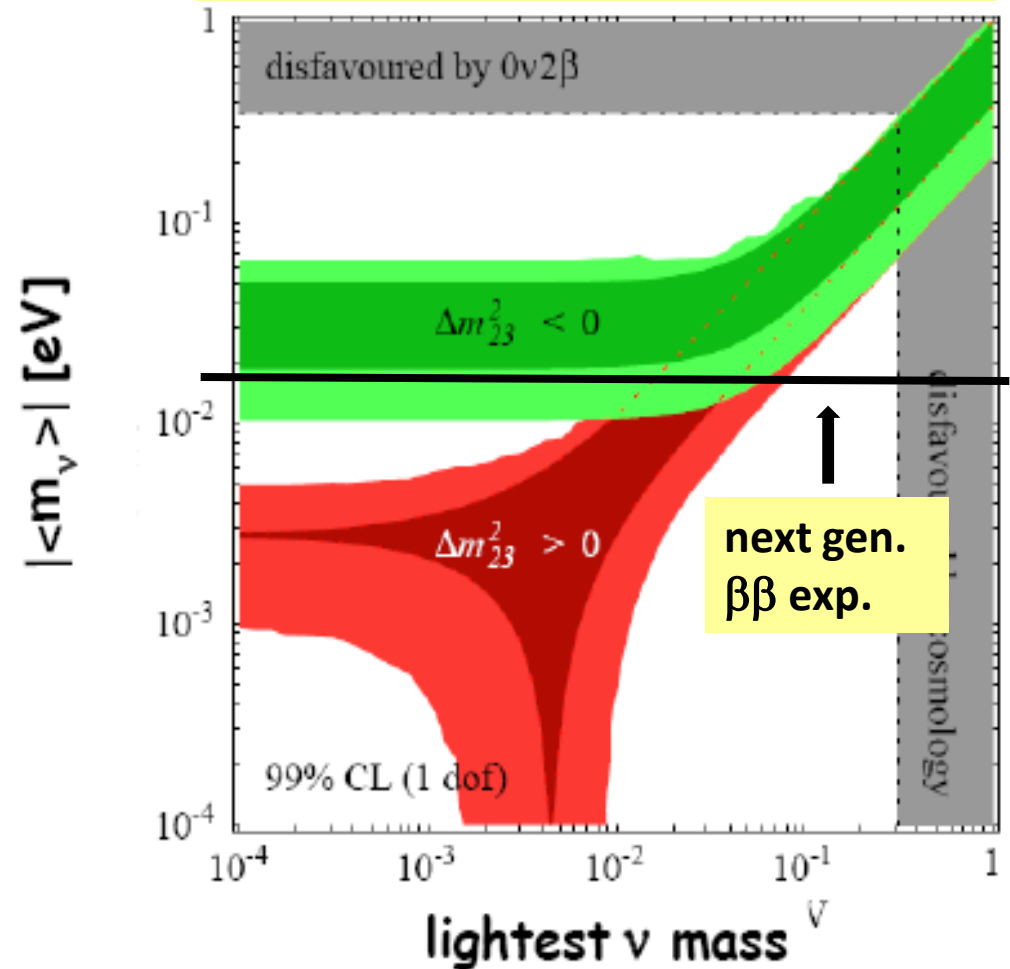
- ◆ 10-50 kton detector
- ◆ 2-3% energy resolution
- ◆ Rich physics possibilities
 - ⇒ Mass hierarchy
 - ⇒ Precision measurement of 4 mixing parameters
 - ⇒ Supernovae neutrino
 - ⇒ Geoneutrino
 - ⇒ Sterile neutrino
 - ⇒ Abnormal magnetic moment
 - ⇒ Discoveries with a high precision detector?

Mass hierarchy at reactors

- Effects of mass hierarchy can be seen from the distortion of neutrino energy spectrum at reactors after a F_0 transformation

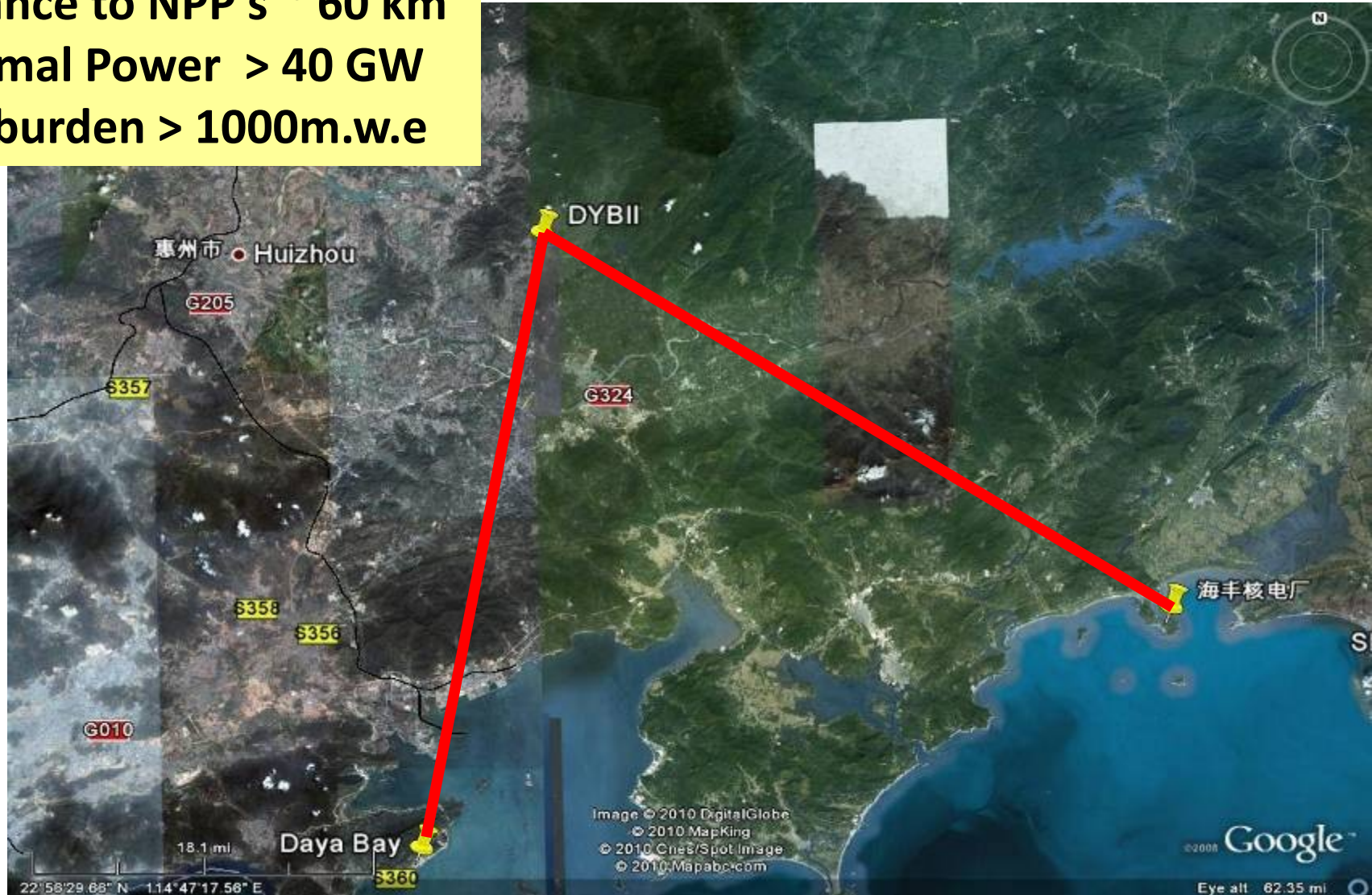


Mass hierarchy is crucial for the understanding of the Dirac/Majorana nature of neutrinos



Possible Site for Daya Bay II

Distance to NPP's ~ 60 km
Thermal Power > 40 GW
overburden > 1000m.w.e



Closing Remarks

- **BEPCII/BESIII will produce many interesting results in Charm and τ physics in next 10 years. The future plan will depend on the results from BESIII+LHC+SuperB... and to be decided within 5-6 years.**
- **Dayabay neutrino exp. Is under construction, and will start data taking next year.**
- **Experiments in the cosmic frontier are fast increasing.**
- **Future plan: next 5 years plan under review**
- **More international cooperation**
- **Good opportunity and Great challenges**