# BASICS OF QCD FOR THE LHC 

Lecture IV

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## LECTURES

I. Intro and QCD fundamentals
2. QCD in the final state
3. QCD in the initial state

## 4. From accurate QCD to useful QCD

5. Advanced QCD with applications at the LHC

## LHC MASTER FORMULA



## HOW DO WE MAKE PREDICTIONS?

## I. Fixed order computations: from LO to NNLO <br> 2. Parton showers and fully exclusive simulations

In practice we use public codes, which are often very-loosely called Monte Carlo's, that implement various results/approaches. In general the predictions of NLO and NNLO calculations are given in terms of distributions of infrared safe observables (histograms), while proper Monte Carlo Generators give out events. Keep this difference in mind!

## LHC MASTER FORMULA

$$
\sigma_{X}=\sum_{a, b} \int_{0}^{1} d x_{1} d x_{2} f_{a}\left(x_{1}, \mu_{F}^{2}\right) f_{b}\left(x_{2}, \mu_{F}^{2}\right) \times \hat{\sigma}_{a b \rightarrow X}\left(x_{1}, x_{2}, \alpha_{S}\left(\mu_{R}^{2}\right), \frac{Q^{2}}{\mu_{F}^{2}}, \frac{Q^{2}}{\mu_{R}^{2}}\right)
$$

Two ingredients necessary:
I. Parton Distribution functions (from exp, but evolution from th).
2. Short distance coefficients as an expansion in $\boldsymbol{\alpha}_{\mathrm{S}}$ (from th).

$$
\hat{\sigma}_{a b \rightarrow X}=\sigma_{0}+\alpha_{S} \sigma_{1}+\alpha_{S}^{2} \sigma_{2}+\ldots
$$

Leading order
Next-to-leading order
Next-to-next-to-leading order

## Predictions at LO

How do we calculate a LO cross section for 3 jets at the LHC?
I. Identify all subprocesses ( $\mathrm{gg} \rightarrow \mathrm{ggg}, \mathrm{qg} \rightarrow \mathrm{qgg} . .$. ) in:

$$
\sigma(p p \rightarrow 3 j)=\sum_{i j k} \int f_{i}\left(x_{1}\right) f_{j}\left(x_{2}\right) \hat{\sigma}\left(i j \rightarrow k_{1} k_{2} k_{3}\right)
$$

II. For each one, calculate the amplitude:

$$
\mathcal{A}(\{p\},\{h\},\{c\})=\sum_{i} D_{i}
$$

III. Square the amplitude, sum over spins \& color, integrate over the phase space (D ~ 3n)

$$
\hat{\sigma}=\frac{1}{2 \hat{s}} \int d \Phi_{p} \sum_{h, c}|\mathcal{A}|^{2}
$$

## HOW DIFFICULT IS IT TO CALCULATE $|A|^{2}$ FOR ARBITRARY PROCESSES?

Consider a simple 5 gluon amplitude:




There are 25 diagrams with a complicated tensor structure, so you get....
$\mathrm{A}(\mathrm{k} 1, \mathrm{e} 1, \mathrm{k} 2, \mathrm{e} 2, \mathrm{k} 3, \mathrm{e} 3, \mathrm{k} 4, \mathrm{e} 4, \mathrm{k} 5, \mathrm{ot})=-\operatorname{Tr}(\mathrm{Ta} 1, \mathrm{Ta} 2, \mathrm{Ta} 3, \mathrm{Ta} 4, \mathrm{Ta} 5)^{*}\left(1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{k} 1 . \mathrm{e}^{*} \mathrm{e} 1 . \mathrm{e} 3^{*} \mathrm{e} 4 . \mathrm{e} 5-\mathrm{den}\left(2^{*} \mathrm{k} 1 \mathrm{k} 1 \mathrm{k} 2\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*} \mathrm{e} 1 . \mathrm{e} 4^{*} \mathrm{e} 3 . \mathrm{e} 5\right.$

$-1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{k}^{2} \mathrm{en}^{*} \mathrm{e} 2 . \mathrm{e} 3^{*} \mathrm{e} 4 . \mathrm{e} 5+\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{e} 2 . \mathrm{e} 4^{*} \mathrm{e} 3 . \mathrm{e} 5-1 / 2^{*} \mathrm{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{k} 2 . e 1^{*} \mathrm{e} 2 . \mathrm{e} 5^{*} \mathrm{e} 3 . \mathrm{e} 4+1 / 4^{*} \mathrm{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{k} \mathrm{O}_{\mathrm{e}} 3^{*} \mathrm{e} 1 . \mathrm{e} 2^{*} \mathrm{e} 4 . \mathrm{e} 5$
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## $+\operatorname{Tr}(\mathrm{Ta} 1, \mathrm{Ta} 2, \mathrm{Ta} 3, \mathrm{Ta} 4, \mathrm{Ta} 5) *(1 / 2 * \operatorname{den}(2 * \mathrm{k} 1 . \mathrm{k} 2) * \mathrm{k} 1 . \mathrm{e} 2 * \mathrm{e} 1 . \mathrm{e} 3 * \mathrm{e} 4 . \mathrm{e} 5$


$-\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e}^{*} \mathrm{k} 1 . \mathrm{e}^{*} \mathrm{k} 4 . \mathrm{e} 3^{*} \mathrm{e} 1 . e 4+1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*} \mathrm{k} 2 . \mathrm{k} 3^{*} \mathrm{e} 1.5^{*} \mathrm{e} 3 . \mathrm{e} 4-1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*}$ $\mathrm{k} 2 . \mathrm{k} 4^{*} \mathrm{e} 1 . \mathrm{e}^{*} \mathrm{e} 3 . \mathrm{e} 4-\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*} \mathrm{k} 2 . \mathrm{e}^{*} \mathrm{k} 3 . \mathrm{e}^{*} \mathrm{e} 1 . \mathrm{e} 5+\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*} \mathrm{k} 2 . \mathrm{e} 4^{*} \mathrm{k} 4 . \mathrm{e}^{*} \mathrm{e} 1 . \mathrm{e} 5-1 / 2^{*} \mathrm{den}\left(2^{*}\right.$ $\mathrm{k} 1 . \mathrm{k} 2)^{*} \mathrm{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . e 2^{*} \mathrm{k} 2 . \mathrm{e}^{*} \mathrm{k} 3 . \mathrm{e} 1^{*} \mathrm{e} 3 . \mathrm{e} 4+\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . e 2^{*} \mathrm{k} 2 . \mathrm{e} 5^{*} \mathrm{k} 3 . e 4^{*} \mathrm{e} 1 . e 3+1 / 2^{*} \mathrm{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . e 2^{*}$ $\mathrm{k} 2 . e 5^{*} \mathrm{k} 4 . \mathrm{e} 1^{*} \mathrm{e} 3 . e 4-\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . e 2^{*} \mathrm{k} 2 . \mathrm{e} 5^{*} \mathrm{k} 4 . \mathrm{e}^{*} \mathrm{e} 1 . e 4-1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*} \mathrm{k} 3 . \mathrm{k} 5^{*} \mathrm{e} 1 . e 5^{*} \mathrm{e} 3 . e 4+$ $\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*} \mathrm{k} 3 . \mathrm{e} 1^{*} \mathrm{k} 3 . \mathrm{e} 4^{*} \mathrm{e} 3 . \mathrm{e} 5-\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*} \mathrm{k} 3 . \mathrm{e} 1^{*} \mathrm{k} 4 . \mathrm{e} 3^{*} \mathrm{e} 4 . \mathrm{e} 5+\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 1 . \mathrm{e} 2^{*}$ $\mathrm{k} 3 . \mathrm{e} 1^{*} \mathrm{k} 4 . \mathrm{e}^{*} \mathrm{e} 3 . \mathrm{e} 4-\operatorname{den}\left(2^{*} \mathrm{k} \mathrm{m}\right.$
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\operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 3^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{e} 2 . e 5^{*}$ $\mathrm{e} 3 . \mathrm{e} 4+1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 3^{*} \mathrm{k} 3 . \mathrm{e} 4^{*} \mathrm{e} 1 . \mathrm{e} 2^{*} \mathrm{e} 3 . \mathrm{e} 5-1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 3^{*} \mathrm{k} 4 . \mathrm{e}^{*} \mathrm{e} 1 . \mathrm{e}^{*} \mathrm{e} 4 . \mathrm{e} 5+1 / 2^{*} \mathrm{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*}$ $\operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 3^{*} \mathrm{k} 4 . \mathrm{e}^{*} \mathrm{e} 1.2^{*} \mathrm{e} 3 . \mathrm{e} 4+1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 4^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{e} 2 . \mathrm{e} 5^{*} \mathrm{e} 3 . \mathrm{e} 4+1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 4^{*} \mathrm{k} 3 . \mathrm{e} 4^{*}$ e1.e2*e3.e5-1/2*den(2*k1.k2)*den(2*k3.k4)*k2.k4*k3.e5*e1.e2*e3.e4-1/2*den(2*k1.k2)*den(2*k3.k4)*k2.k4*k4.e3*e1.e2*e4.e5-1/2*den(2* $\mathrm{k} 1 . \mathrm{k} 2)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 5^{*} \mathrm{k} 3 . \mathrm{e}^{*} \mathrm{e} 1 . \mathrm{e} 2^{*} \mathrm{e} 3 . \mathrm{e} 5+1 / 4^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 5^{*} \mathrm{k} 3 . \mathrm{e} 5^{*} \mathrm{e} 1 . \mathrm{e} 2^{*} \mathrm{e} 3 . \mathrm{e} 4+1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 5^{*}$ $\mathrm{k} 4 . \mathrm{e}^{*} \mathrm{e} 1 . \mathrm{e}^{*} \mathrm{e} 4 . \mathrm{e} 5-1 / 4^{*} \mathrm{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{k} 5^{*} \mathrm{k} 4.5^{*} \mathrm{e} 1 . \mathrm{e}^{*} \mathrm{e} 3 . \mathrm{e} 4+\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \mathrm{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{k} 2 . \mathrm{e} 3^{*} \mathrm{k} 3 . \mathrm{e} 4^{*} \mathrm{e} 2 . \mathrm{e} 5-$ $\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{k} 2 . \mathrm{e} 4^{*} \mathrm{k} 4 . \mathrm{e} 3^{*} \mathrm{e} 2 . \mathrm{e} 5+1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{k} 2 . \mathrm{e} 5^{*} \mathrm{k} 3 . \mathrm{e} 2^{*} \mathrm{e} 3 . \mathrm{e} 4-\operatorname{den}\left(2^{*} \mathrm{k} 1 \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{e} 1^{*}$ $\mathrm{k} 2.5^{*} \mathrm{k} 3 . \mathrm{e} 4^{*} \mathrm{e} 2 . \mathrm{e} 3-1 / 2^{*} \operatorname{den}\left(2^{*} \mathrm{k} 1 \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{k} 2 . \mathrm{e} 5^{*} \mathrm{k} 4 . \mathrm{e}^{*} \mathrm{e} 3 . \mathrm{e} 4+\operatorname{den}\left(2^{*} \mathrm{k} 1 . \mathrm{k} 2\right)^{*} \operatorname{den}\left(2^{*} \mathrm{k} 3 . \mathrm{k} 4\right)^{*} \mathrm{k} 2 . \mathrm{e} 1^{*} \mathrm{k} 2 . e 5^{*} \mathrm{k} 4 . \mathrm{e} 3^{*} \mathrm{e} 2 . e 4$

## HOW DIFFICULT IS IT TO CALCULATE $|A|^{2}$ FOR ARBITRARY PROCESSES?

Solution

- Work always at the amplitude level (not squared)
- Keep track of all the quantum numbers, (momenta, spin and color)
- Organize them in efficient way, by choosing appropriate basis


## HOW DIFFICULT IS IT TO CALCULATE $|A|^{2}$ FOR ARBITRARY PROCESSES?

Calculate helicity amplitudes, ie amplitudes for gluons and quarks in a definite helicity states. For massless quarks this amounts to condering chirality states:

$$
u_{ \pm}(k)=\frac{1}{2}\left(1 \pm \gamma_{5}\right) u(k)
$$

External gluons you always think them as attached to a quark-anti-quark pair with a definite (yet arbitrary) polarization vectors:

$$
\varepsilon_{\mu}^{+}(k ; q)=\frac{\left\langle q^{-}\right| \gamma_{\mu}\left|k^{-}\right\rangle}{\sqrt{2}\langle q k\rangle}, \quad \varepsilon_{\mu}^{-}(k, q)=\frac{\left\langle q^{+}\right| \gamma_{\mu}\left|k^{+}\right\rangle}{\sqrt{2}[k q]}
$$

It's just a more sophisticated version of the circular polarization. Choosing appropriately the gauge vector, expressions simplify dramatically.

## HOW DIFFICULT IS IT TO CALCULATE $|A|^{2}$ FOR ARBITRARY PROCESSES?

Inspired by the way gauge theories appear as the zero-slope limits of (open) string theories, it has been suggested to decompose the full amplitude as a sum of gauge invariant subamplitudes times color coefficients:

$$
\mathcal{A}_{n}\left(g_{1}, \ldots, g_{n}\right)=g^{n-2} \sum_{\sigma \in S_{n-1}} \operatorname{Tr}\left(\mathbf{t}^{a_{1}} \mathbf{t}^{a_{\sigma_{2}}} \ldots \mathbf{t}^{a_{\sigma_{n}}}\right) A_{n}\left(1, \sigma_{2}, \ldots, \sigma_{n}\right)
$$

where the formula $i f a b c=\operatorname{Tr}\left(\mathrm{t}^{\mathrm{a}},\left[\mathrm{t}^{\mathrm{b}}, \mathrm{t}^{\mathrm{c}}\right]\right)$ has been repeatedly used to reduce the $\mathrm{f}^{\prime} \mathrm{s}$ into traces of lambdas and the Fierz identities to cancel traces of length $1<n$. Analogously for quarks:

$$
\mathcal{A}_{n}\left(q_{1}, g_{2}, \ldots, g_{n-1}, \bar{q}_{n}\right)=g^{n-2} \sum_{\sigma \in S_{n-2}}\left(\mathbf{t}^{a_{\sigma_{2}} \cdots} \mathbf{t}^{a_{\sigma_{n-1}}}\right)_{j}^{i} A_{n}\left(1_{q}, \sigma_{2}, \ldots, \sigma_{n-2}, n_{\bar{q}}\right)
$$

The $A_{n}$ are $M U C H$ simpler objects to calculate, with many less diagrams...

## HOW DIFFICULT IS IT TO CALCULATE $|\mathbf{A}|^{2}$ FOR ARBITRARY PROCESSES?

| n | full Amp | partial Amp |
| :---: | :---: | :---: |
| 4 | 4 | 3 |
| 5 | 25 | 10 |
| 6 | 220 | 36 |
| 7 | 2485 | 133 |
| 8 | 34300 | 501 |
| 9 | 559405 | 1991 |
| 10 | 10525900 | 7335 |
| 11 | 224449225 | 28199 |
| 12 | 5348843500 | 108280 |

$$
(2 n)!\quad 3.8^{n}
$$

## HOW DIFFICULT IS IT TO CALCULATE $|A|^{2}$ FOR ARBITRARY PROCESSES?

Feynman diagrams are not efficient because the same subdiagrams are recomputed over and over. Solution: cash them! In other words use recursive relations.

For the color-ordered subamplitudes for $n$ gluons, such relations (called Berends-Giele) are very easy:


Off-shell amplitudes with max n - I number of legs !

How difficult is it to calculate $|A|^{2}$ FOR ARBITRARY PROCESSES?

| n | full Amp | partial Amp | BG |
| :---: | :---: | :---: | :---: |
| 4 | 4 | 3 | 3 |
| 5 | 25 | 10 | 10 |
| 6 | 220 | 36 | 35 |
| 7 | 2485 | 133 | 70 |
| 8 | 34300 | 501 | 126 |
| 9 | 559405 | 1991 | 210 |
| 10 | 10525900 | 7335 | 330 |
| 11 | 224449225 | 28199 | 495 |
| 12 | 5348843500 | 108280 | 715 |
| $(2 n)!$ |  |  |  |

The factorial growth is tamed to a polynomial one!
Note, however, one still needs to sum over color, an operation which sets the complexity back to exponential.

## HOW DIFFICULT IS TO CALCULATE $|A|^{2}$ FOR ARBITRARY PROCESSES?

Problem of generating the matrix elements for any process of interest has been solved in full generality and it has been automatized!

More than that, also the integration over phase of such matrix elements can be achieved in an automatic way (non-trivial problem not discussed here)!

Several public tools exist:
CompHEP/CalcHEP/MadGraph/SHERPA/Whizard/....




## LO PREDICTIONS : FINAL REMARKS

$$
\sigma_{X}=\sum_{a, b} \int_{0}^{1} d x_{1} d x_{2} f_{a}\left(x_{1}, \mu_{F}^{2}\right) f_{b}\left(x_{2}, \mu_{F}^{2}\right) \times \hat{\sigma}_{a b \rightarrow X}\left(x_{1}, x_{2}, \alpha_{S}\left(\mu_{R}^{2}\right), \frac{Q^{2}}{\mu_{F}^{2}}, \frac{Q^{2}}{\mu_{R}^{2}}\right)
$$

- By calculating the short distance coefficient at tree-level we obtain the first estimate of rates for inclusive final states.
- Even at LO extra radiation is included: it is described by the PDF's in the initial state and by the definition of a final state parton, which at LO represents all possible final state evolutions.
- Due to the above approximations a cross section at LO can strongly depend on the factorization and renormalization scales.
- Predictions can be systematically improved, at NLO and NNLO, by including higher order corrections in the short distance and in the evolution of the PDF's.


## Predictions at NLO

$$
\begin{gathered}
\sigma_{X}=\sum_{a, b} \int_{0}^{1} d x_{1} d x_{2} f_{a}\left(x_{1}, \mu_{F}^{2}\right) f_{b}\left(x_{2}, \mu_{F}^{2}\right) \times \hat{\sigma}_{a b \rightarrow X}\left(x_{1}, x_{2}, \alpha_{S}\left(\mu_{R}^{2}\right), \frac{Q^{2}}{\mu_{F}^{2}}, \frac{Q^{2}}{\mu_{R}^{2}}\right) \\
\hat{\sigma}_{a b \rightarrow X}=\sigma_{0}+\alpha_{S} \sigma_{1}+\alpha_{S}^{2} \sigma_{2}+\ldots
\end{gathered}
$$

## Why?

I. First order where scale dependences are compensated by the running of $\boldsymbol{\alpha}_{S}$ and the evolution of the PDF's: FIRST RELIABLE ESTIMATE OF THE TOTAL CROSS SECTION.
2. The impact of extra radiation is included. For example, jets now have a structure.
3. New effects coming up from higher order terms (e.g., opening up of new production channels or phase space dimensions) can be evaluated.


## PREDICTIONS AT NLO

How? I. Get the "ingredients"
2. "Method" to combine them to calculate infrared observables

## Ingredients:



Virtual part


Real emission part


Born

$$
\int \mathrm{d} \sigma^{(\mathrm{NLO})} O(\Phi)=\int \mathrm{d} \Phi_{B} V\left(\Phi_{B}\right) O\left(\Phi_{B}\right)+\int \mathrm{d} \Phi_{R} R\left(\Phi_{R}\right) O\left(\Phi_{R}\right) \int \mathrm{d} \Phi_{B} B\left(\Phi_{B}\right) O\left(\Phi_{B}\right)
$$

Loops have been for long the bottleneck of NLO computations, with their calculations taking years of manual and symbolic work to get the correct results.

## Predictions At NLO



Thanks to new amazing results, some of them inspired by string theory developments, now the computation of loops has been extended to high-multiplicity processes or/and automated.

## PREDICTIONS AT NLO

How? I. Get the "ingredients"
2. "Method" to combine them to calculate infrared observables

## Method: Universal subtraction

$$
\begin{aligned}
\int \mathrm{d} \sigma^{(\mathrm{NLO})} O(\Phi) & =\int \mathrm{d} \Phi_{B} V\left(\Phi_{B}\right) O\left(\Phi_{B}\right)+\int \mathrm{d} \Phi_{R} R\left(\Phi_{R}\right) O\left(\Phi_{R}\right) \int \mathrm{d} \Phi_{B} B\left(\Phi_{B}\right) O\left(\Phi_{B}\right) \\
& =\int \mathrm{d} \Phi_{B}\left[B\left(\Phi_{B}\right)+V\left(\Phi_{B}\right)+\int \mathrm{d} \Phi_{R \mid B} S\left(\Phi_{R}\right)\right] O\left(\Phi_{B}\right) \\
& +\int \mathrm{d} \Phi_{R}\left[R\left(\Phi_{R}\right) O\left(\Phi_{R}\right)-S\left(\Phi_{R}\right) O\left(\Phi_{B}\right)\right]
\end{aligned}
$$

Local universal counterterms have been identified whose integral on the extra radiation variable is analytically known and that can be used to make reals and virtuals separately finite.

## Predictions at NLO

MCFM: downloadable general purpose NLO code [Campbell \& Ellis+ collaborators]

| Final state | Notes | Reference |
| :---: | :---: | :---: |
| W/Z |  |  |
| diboson <br> (W/Z/Y) | photon fragmentation, <br> anomalous couplings | hep-ph/9905386, <br> arXiv:1105.0020 |
| Wbb | massless b-quark <br> massive b quark | hep-ph/9810489 <br> arXiv:1011.6647 |
| Zbb | massless b-quark | hep-ph/0006304 |
| $W / Z+\mathrm{I}$ jet |  | hep-ph/0202176, <br> hep-ph/0308195 |
| W/Z+2 jets | massive c-quark | hep-ph/0506289 |
| Zb | 5-flavour scheme | hep-ph/0312024 |
| Zb+jet | 5-flavour scheme | hep-ph/0510362 |


| Final state | Notes | Reference |
| :---: | :---: | :---: |
| H (gluon fusion) |  |  |
| $\mathrm{H}+\mathrm{I}$ jet (g.f.) | effective coupling |  |
| $\mathrm{H}+2$ jets (g.f.) | effective coupling | hep-ph/0608194, <br> arXiv:1001.4495 |
| $\mathrm{WH} / \mathrm{ZH}$ |  |  |
| H (WBF) |  | hep-ph/0403194 |
| Hb | 5-flavour scheme | hep-ph/0204093 |
| t | s- and t-channel (5F), <br> top decay included | hep-ph/0408158 |
| t | t-channel (4F) | arXiv:0903.0005, <br> arXiv:0907.3933 |
| Wt | 5-flavour scheme | hep-ph/0506289 |
| top pairs | top decay included |  |

$\sim 30$ processes
First results implemented in 1998 ...this is 13 years worth of work of several people ( $\sim 4 \mathrm{M} \$$ )
Cross sections and parton-level distributions at NLO are provided
One general framework. However, each process implemented by hand

## Predictions at NLO

Completely automatically generated NLO codes for a variety of processes via MadLoop+MadFKS

Total sample cross sections at the LHC for 26 sample procs

Code generation time:
a few hours
Running time:
two weeks on a cluster

|  | Process | $\mu$ | $n_{l f}$ | Cross section (pb) |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | LO | NLO |
| a. 1 | $p p \rightarrow t \bar{t}$ | $m_{\text {top }}$ | 5 | $123.76 \pm 0.05$ | $162.08 \pm 0.12$ |
| a. 2 | $p p \rightarrow t j$ | $m_{\text {top }}$ | 5 | $34.78 \pm 0.03$ | $41.03 \pm 0.07$ |
| a. 3 | $p p \rightarrow t j j$ | $m_{\text {top }}$ | 5 | $11.851 \pm 0.006$ | $13.71 \pm 0.02$ |
| a. 4 | $p p \rightarrow t \bar{b} j$ | $m_{\text {top }} / 4$ | 4 | $25.62 \pm 0.01$ | $30.96 \pm 0.06$ |
| a. 5 | $p p \rightarrow t \bar{b} j j$ | $m_{\text {top }} / 4$ | 4 | $8.195 \pm 0.002$ | $8.91 \pm 0.01$ |
| b. 1 | $p p \rightarrow\left(W^{+} \rightarrow\right) e^{+} \nu_{e}$ | $m_{W}$ | 5 | $5072.5 \pm 2.9$ | $6146.2 \pm 9.8$ |
| b. 2 | $p p \rightarrow\left(W^{+} \rightarrow\right) e^{+} \nu_{e} j$ | $m_{W}$ | 5 | $828.4 \pm 0.8$ | $1065.3 \pm 1.8$ |
| b. 3 | $p p \rightarrow\left(W^{+} \rightarrow\right) e^{+} \nu_{e} j j$ | $m_{W}$ | 5 | $298.8 \pm 0.4$ | $300.3 \pm 0.6$ |
| b. 4 | $p p \rightarrow\left(\gamma^{*} / Z \rightarrow\right) e^{+} e^{-}$ | $m_{Z}$ | 5 | $1007.0 \pm 0.1$ | $1170.0 \pm 2.4$ |
| b. 5 | $p p \rightarrow\left(\gamma^{*} / Z \rightarrow\right) e^{+} e^{-} j$ | $m_{Z}$ | 5 | $156.11 \pm 0.03$ | $203.0 \pm 0.2$ |
| b. 6 | $p p \rightarrow\left(\gamma^{*} / Z \rightarrow\right) e^{+} e^{-} j j$ | $m_{Z}$ | 5 | $54.24 \pm 0.02$ | $56.69 \pm 0.07$ |
| c. 1 | $p p \rightarrow\left(W^{+} \rightarrow\right) e^{+} \nu_{e} b \bar{b}$ | $m_{W}+2 m_{b}$ | 4 | $11.557 \pm 0.005$ | $22.95 \pm 0.07$ |
| c. 2 | $p p \rightarrow\left(W^{+} \rightarrow\right) e^{+} \nu_{e} t \bar{t}$ | $m_{W}+2 m_{\text {top }}$ | 5 | $0.009415 \pm 0.000003$ | $0.01159 \pm 0.00001$ |
| c. 3 | $p p \rightarrow\left(\gamma^{*} / Z \rightarrow\right) e^{+} e^{-} b \bar{b}$ | $m_{Z}+2 m_{b}$ | 4 | $9.459 \pm 0.004$ | $15.31 \pm 0.03$ |
| c. 4 | $p p \rightarrow\left(\gamma^{*} / Z \rightarrow\right) e^{+} e^{-} t \bar{t}$ | $m_{Z}+2 m_{\text {top }}$ | 5 | $0.0035131 \pm 0.0000004$ | $0.004876 \pm 0.000002$ |
| c. 5 | $p p \rightarrow \gamma t \bar{t}$ | $2 m_{\text {top }}$ | 5 | $0.2906 \pm 0.0001$ | $0.4169 \pm 0.0003$ |
| d. 1 | $p p \rightarrow W^{+} W^{-}$ | $2 m_{W}$ | 4 | $29.976 \pm 0.004$ | $43.92 \pm 0.03$ |
| d. 2 | $p p \rightarrow W^{+} W^{-} j$ | $2 m_{W}$ | 4 | $11.613 \pm 0.002$ | $15.174 \pm 0.008$ |
| d. 3 | $p p \rightarrow W^{+} W^{+} j j$ | $2 m_{W}$ | 4 | $0.07048 \pm 0.00004$ | $0.1377 \pm 0.0005$ |
| e. 1 | $p p \rightarrow H W^{+}$ | $m_{W}+m_{H}$ | 5 | $0.3428 \pm 0.0003$ | $0.4455 \pm 0.0003$ |
| e. 2 | $p p \rightarrow H W^{+} j$ | $m_{W}+m_{H}$ | 5 | $0.1223 \pm 0.0001$ | $0.1501 \pm 0.0002$ |
| e. 3 | $p p \rightarrow H Z$ | $m_{Z}+m_{H}$ | 5 | $0.2781 \pm 0.0001$ | $0.3659 \pm 0.0002$ |
| e. 4 | $p p \rightarrow H Z j$ | $m_{Z}+m_{H}$ | 5 | $0.0988 \pm 0.0001$ | $0.1237 \pm 0.0001$ |
| e. 5 | $p p \rightarrow H t \bar{t}$ | $m_{\text {top }}+m_{H}$ | 5 | $0.08896 \pm 0.00001$ | $0.09869 \pm 0.00003$ |
| e. 6 | $p p \rightarrow H b \bar{b}$ | $m_{b}+m_{H}$ | 4 | $0.16510 \pm 0.00009$ | $0.2099 \pm 0.0006$ |
| e. 7 | $p p \rightarrow H j j$ | $m_{H}$ | 5 | $1.104 \pm 0.002$ | $1.036 \pm 0.002$ |

## Predictions At NLO

- NLO calculations have historically presented two types of challenges: the loop calculations and the construction of a numerical code resilient to the cancellation of the divergences.
- Both issues have now basically solved in general and many NLO calculations can now be done in an automatic way.
- Several public codes that compute IR-safe quantities (cross sections, jet rates, ...) at the parton level are available.
- Be careful : NLO codes are NOT event generators!!


## Predictions at NLO

Calling a code "a NLO code" is an abuse of language and can be confusing. A NLO calculation always refers to an IR-safe observable, when the genuine $\boldsymbol{\alpha}_{\mathrm{s}}$ corrections to this observable on top of the LO estimate are known.

An NLO code will, in general, be able to produce results for several quantities and distributions, only some of which will be at NLO accuracy.

Example: Jet rates in the JADE algorithm:

$$
\begin{aligned}
\sigma_{2 j} & =\sigma^{\text {Born }}\left(1-\frac{\alpha_{S} C_{F}}{\pi} \log ^{2} y+\ldots\right) \\
\sigma_{3 j} & =\sigma^{\operatorname{Born}} \frac{\alpha_{S} C_{F}}{\pi} \log ^{2} y+\ldots
\end{aligned}
$$

$\sigma_{2 j}$ is NLO, while $\sigma_{3 j}$ is just LO!


## Predictions at NLO

Calling a code "a NLO code" is an abuse of language and can be confusing.
A NLO calculation always refers to an IR-safe observable, when the genuine $\boldsymbol{\alpha}_{\mathrm{s}}$ corrections to this observable on top of the LO estimate are known.

An NLO code will, in general, be able to produce results for several quantities and distributions, only some of which will be at NLO accuracy.

Example: Suppose we use the NLO code for pp $\rightarrow \mathrm{tt}$


## EXPERIENCE A 'SIMPLE" NLO CALCULATION YOURSELF

## $P P \longrightarrow$ WIGS + X AT NEO

- LO : I-loop calculation and HEFT
- NLO in the HEFT
- Virtual corrections and renormalization
- Real corrections and IS singularities
- Cross sections at the LHC


## $\mathrm{PP} \longrightarrow \mathrm{H}+\mathrm{X}$ AT $\mathrm{LO}_{p}$

This is a "simple" $2 \rightarrow \mid$ process.
However, at variance with $\mathrm{pp} \rightarrow \mathrm{W}$, the LO order process already proceeds through a loop.

In this case, this means that the loop calculation ${ }_{b, \nu}$ has to give a finite result! Let's do the calculation!

$q$

$$
i \mathcal{A}=-\left(-i g_{s}\right)^{2} \operatorname{Tr}\left(t^{a} t^{b}\right)\left(\frac{-i m_{t}}{v}\right) \int \frac{d^{d} \ell}{(2 \pi)^{n}} \frac{T^{\mu \nu}}{\operatorname{Den}}(i)^{3} \epsilon_{\mu}(p) \epsilon_{\nu}(q)
$$

where

$$
\operatorname{Den}=\left(\ell^{2}-m_{t}^{2}\right)\left[(\ell+p)^{2}-m_{t}^{2}\right]\left[(\ell-q)^{2}-m_{t}^{2}\right]
$$

We combine the denominators into one by using $\quad \frac{1}{A B C}=2 \int_{0}^{1} d x \int_{0}^{1-x} \frac{d y}{[A x+B y+C(1-x-y)]^{3}}$

$$
\frac{1}{\mathrm{Den}}=2 \int d x d y \frac{1}{\left[\ell^{2}-m_{t}^{2}+2 \ell \cdot(p x-q y)\right]^{3}}
$$

## $\mathrm{PP} \longrightarrow \mathrm{H}+\mathrm{X}$ AT $\mathrm{LO}_{p}$

We shift the momentum:
$\ell^{\prime}=\ell+p x-q y$
$\frac{1}{\text { Den }} \rightarrow 2 \int d x d y \frac{1}{\left[\ell^{\prime 2}-m_{t}^{2}+M_{H}^{2} x y\right]^{3}}$.


And now the tensor in the numerator:

$$
\begin{aligned}
T^{\mu \nu} & \left.=\operatorname{Tr}\left[\left(\ell+m_{t}\right) \gamma^{\mu}\left(\ell+p+m_{t}\right)\left(\ell-q+m_{t}\right) \gamma^{\nu}\right)\right] \\
& =4 m_{t}\left[g^{\mu \nu}\left(m_{t}^{2}-\ell^{2}-\frac{M_{H}^{2}}{2}\right)+4 \ell^{\mu} \ell^{\nu}+p^{\nu} q^{\mu}\right]
\end{aligned}
$$

where I used the fact that the external gluons are on-shell. This trace is proportional to mt !
This is due to the spin flip caused by the scalar coupling.
Now we shift the loop momentum also here, we drop terms linear in the loop momentum (they are odd and vanish) and

## $\mathrm{PP} \longrightarrow \mathrm{H}+\mathrm{X}$ AT $\mathrm{LO}_{p}$

We perform the tensor decomposition using:

$$
\int d^{d} k \frac{k^{\mu} k^{\nu}}{\left(k^{2}-C\right)^{m}}=\frac{1}{d} g^{\mu \nu} \int d^{d} k \frac{k^{2}}{\left(k^{2}-C\right)^{m}}
$$

So I can write an expression which depends only $b, \nu$ on scalar loop integrals:

$$
\begin{aligned}
i \mathcal{A} & =-\frac{2 g_{s}^{2} m_{t}^{2}}{v} \delta^{a b} \int \frac{d^{d} \ell^{\prime}}{(2 \pi)^{d}} \int d x d y\left\{g^{\mu \nu}\left[m^{2}+\ell^{\prime 2}\left(\frac{4-d}{d}\right)+M_{H}^{2}\left(x y-\frac{1}{2}\right)\right]\right. \\
& \left.+p^{\nu} q^{\mu}(1-4 x y)\right\} \frac{2 d x d y}{\left(\ell^{\prime 2}-m_{t}^{2}+M_{H}^{2} x y\right)^{3}} \epsilon_{\mu}(p) \epsilon_{\nu}(q)
\end{aligned}
$$

There's a term which apparently diverges....??
Ok, Let's look the scalar integrals up in a table (or calculate them!)

## $\mathrm{PP} \longrightarrow \mathrm{H}+\mathrm{X}$ AT $\mathrm{LO}_{p}$

$$
\begin{array}{lll}
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{k^{2}}{\left(k^{2}-C\right)^{3}}=\frac{i}{32 \pi^{2}}(4 \pi)^{\epsilon} \frac{\Gamma(1+\epsilon)}{\epsilon}(2-\epsilon) C^{-\epsilon} & a, \mu & 00000 \\
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{\left(k^{2}-C\right)^{3}}=-\frac{i}{32 \pi^{2}}(4 \pi)^{\epsilon} \Gamma(1+\epsilon) C^{-1-\epsilon} . & \ell+p
\end{array}
$$ a very simple final result!!

$$
\mathcal{A}(g g \rightarrow H)=-\frac{\alpha_{S} m_{t}^{2}}{\pi v} \delta^{a b}\left(g^{\mu \nu} \frac{M_{H}^{2}}{2}-p^{\nu} q^{\mu}\right) \int d x d y\left(\frac{1-4 x y}{m_{t}^{2}-m_{H}^{2} x y}\right) \epsilon_{\mu}(p) \epsilon_{\nu}(q) .
$$

Comments:

* The final dependence of the result is $\mathrm{mt}^{2}$ : one from the Yukawa coupling, one from the spin flip.
* The tensor structure could have been guessed by gauge invariance.
* The integral depends on $m_{t}$ and $m_{h}$.


## LO CROSS SECTION

$$
\begin{aligned}
& \sigma(p p \rightarrow H)=\int_{\tau_{0}}^{1} d x_{1} \int_{\tau_{0} / x_{1}}^{1} d x_{2} g\left(x_{1}, \mu_{f}\right) g\left(x_{2}, \mu_{f}\right) \hat{\sigma}(g g \rightarrow H) \\
& x_{1} \equiv \sqrt{\tau} e^{y} \quad x_{2} \equiv \sqrt{\tau} e^{-y} \quad \tau=x_{1} x_{2} \quad \tau_{0}=M_{H}^{2} / S \quad z=\tau_{0} / \tau
\end{aligned}
$$

$$
=\frac{\alpha_{S}^{2}}{64 \pi v^{2}}\left|I\left(\frac{M_{H}^{2}}{m^{2}}\right)\right|^{2} \tau_{0} \int_{\log \sqrt{\tau_{0}}}^{-\log \sqrt{\tau_{0}}} d y g\left(\sqrt{\tau_{0}} e^{y}\right) g\left(\sqrt{\tau_{0}} e^{-y}\right)
$$

The hadronic cross section can be expressed a function of the gluon-gluon luminosity.

$I(x)$ has both a real and imaginary part, which develops at $m_{h}=2 m_{t}$.

This causes a bump in the cross section.

## PP $\rightarrow \mathrm{H}+\mathrm{X}$ @ $\mathbf{N L O}$

At NLO we have to include an extra parton (virtual or real).

The virtuals will become a two-loop calculation!!
Can we avoid that?


Let's consider the case where the Higgs is light:

$$
\begin{aligned}
\mathcal{A}(g g \rightarrow H) & =-\frac{\alpha_{S} m_{t}^{2}}{\pi v} \delta^{a b}\left(g^{\mu \nu} \frac{M_{H}^{2}}{2}-p^{\nu} q^{\mu}\right) \int d x d y\left(\frac{1-4 x y}{m_{t}^{2}-m_{H}^{2} x y}\right) \epsilon_{\mu}(p) \epsilon_{\nu}(q) . \\
& \xrightarrow{m} M_{H}-\frac{\alpha_{S}}{3 \pi v} \delta^{a b}\left(g^{\mu \nu} \frac{M_{H}^{2}}{2}-p^{\nu} q^{\mu}\right) \epsilon_{\mu}(p) \epsilon_{\nu}(q) .
\end{aligned}
$$

This looks like a local vertex, ggH.
The top quark has disappeared from the low energy theory but it has left something behind (non-decoupling).

## Higgs effective field Theory



## LO CROSS SECTION: FULL VS HEFT

$$
\sigma(p p \rightarrow H)=\int_{\tau_{0}}^{1} d x_{1} \int_{\tau_{0} / x_{1}}^{1} d x_{2} g\left(x_{1}, \mu_{f}\right) g\left(x_{2}, \mu_{f}\right) \hat{\sigma}(g g \rightarrow H)
$$

The accuracy of the calculation in the HEFT calculation can be directly assessed by taking the limit $\mathrm{m} \rightarrow \infty$.

For light Higgs is better than 10\%.


So, if we are interested in a light Higgs we use the HEFT and simplify our life. If we do so, the NLO calculation becomes a standard I-loop calculation, similar to Drell-Yan at NLO.

We can (try to) do it!!

## Virtual contributions



Out of 8 diagrams, only two are non-zero (in dimensional regularization), a bubble and a triangle.

They can be easily written down by hand.

Then the integration over the tensor decomposition into scalar integrals and loop integration has to be performed.

One also have to consider that the coefficient of the HEFT receive corrections which have to be included in the result.

The result is:

$$
\begin{aligned}
& \sigma_{\text {virt }}=\sigma_{0} \delta(1-z)\left[1+\frac{\alpha_{S}}{2 \pi} C_{A}\left(\frac{\mu^{2}}{m_{H}^{2}}\right)^{\epsilon} c_{\Gamma}\left(-\frac{2}{\epsilon^{2}}+\frac{11}{3}+\pi^{2}\right)\right], \\
& \sigma_{\text {Born }}=\frac{\alpha_{S}^{2}}{\pi} \frac{m_{H}^{2}}{576 v^{2} s}\left(1+\epsilon+\epsilon^{2}\right) \mu^{2 \epsilon} \delta(1-z) \equiv \sigma_{0} \delta(1-z) \quad z=m_{H}^{2} / s
\end{aligned}
$$

## REAL CONTRIBUTIONS I

$$
\begin{aligned}
& \overline{|\mathcal{M}|^{2}}=\frac{4}{81} \frac{\alpha_{S}^{3}}{\pi v^{2}} \frac{\left(\hat{u}^{2}+\hat{t}^{2}\right)-\epsilon(\hat{u}+\hat{t})^{2}}{\hat{s}} \\
& \text { Integrating over phase space (cms angle theta) } \\
& \hat{t}=-\hat{s}(1-z)(1-\cos \theta) / 2 \\
& \hat{u}=-\hat{s}(1-z)(1+\cos \theta) / 2 \\
& \sigma_{\text {real }}(q \bar{q})=\sigma_{0} \frac{\alpha_{S}}{2 \pi} \frac{64}{27} \frac{(1-z)^{3}}{z} \quad \text { finite! } \\
& \text { H } \\
& \overline{|\mathcal{M}|^{2}}=-\frac{1}{54(1-\epsilon)} \frac{\alpha_{S}^{3}}{\pi v^{2}} \frac{\left(\hat{u}^{2}+\hat{s}^{2}\right)-\epsilon(\hat{u}+\hat{s})^{2}}{\hat{t}} \\
& \text { Integrating over the D-dimensional phase space the } \\
& \text { collinear singularity manifests a pole in I/eps } \\
& \sigma_{\text {real }}=\sigma_{0} \frac{\alpha_{S}}{2 \pi} C_{F}\left(\frac{\mu^{2}}{m_{H}^{2}}\right)^{\epsilon} c_{\Gamma}\left[-\frac{1}{\epsilon} p_{g q}(z)+\frac{(1-z)(7 z-3)}{2 z}+p_{g q}(z) \log \frac{(1-z)^{2}}{z}\right] \\
& \sigma^{\overline{\mathrm{MS}}}(q g)=\sigma_{\text {real }}+\sigma_{\text {c.t. }}^{\text {coll. }} \\
& \sigma_{\text {c.t. }}^{\text {coll. }}=\sigma_{0} \frac{\alpha_{S}}{2 \pi}\left[\left(\frac{\mu^{2}}{\mu_{F}^{2}}\right)^{\epsilon} \frac{c_{\Gamma}}{\epsilon} P_{g q}(z)\right] \\
& =\sigma_{0} \frac{\alpha_{S}}{2 \pi} C_{F}\left[p_{g q}(z) \log \frac{m_{H}^{2}}{\mu_{F}^{2}}+p_{g q}(z) \log \frac{(1-z)^{2}}{z}+\frac{(1-z)(7 z-3)}{2 z}\right]
\end{aligned}
$$

## REAL CONTRIBUTIONS II



## FINAL RESULTS = YOU MADE IT!!

$$
\sigma(p p \rightarrow H)=\sum_{i j} \int_{\tau_{0}}^{1} d x_{1} \int_{\tau_{0} / x_{1}}^{1} d x_{2} f_{i}\left(x_{1}, \mu_{f}\right) f_{j}\left(x_{2}, \mu_{f}\right) \hat{\sigma}(i j)\left[\mu_{f} / m_{h}, \mu_{r} / m_{h}, \alpha_{S}\left(\mu_{r}\right)\right]
$$

The final cross section is the sum of three channels: q qbar, q g, and g g.

The short distance cross section at NLO depends explicitly on the subtraction scales (renormalization and factorization).

The explicit integration over the pdf's is trivial (just mind the plus distributions).

The result is that the corrections are huge!
$K$ factor is $\sim 2$ and scale dependence not really very much improved.


Is perturbation theory valid? NNLO is mandatory...

## FINAL RESULTS = YOU MADE IT!!

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$$

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## Predictions at NNLO

$$
\begin{gathered}
\sigma_{X}=\sum_{a, b} \int_{0}^{1} d x_{1} d x_{2} f_{a}\left(x_{1}, \mu_{F}^{2}\right) f_{b}\left(x_{2}, \mu_{F}^{2}\right) \times \hat{\sigma}_{a b \rightarrow X}\left(x_{1}, x_{2}, \alpha_{S}\left(\mu_{R}^{2}\right), \frac{Q^{2}}{\mu_{F}^{2}}, \frac{Q^{2}}{\mu_{R}^{2}}\right) \\
\hat{\sigma}_{a b \rightarrow X}=\sigma_{0}+\alpha_{S} \sigma_{1}+\alpha_{S}^{2} \sigma_{2}+\ldots
\end{gathered}
$$

## Why?

- A NNLO computation gives control on the uncertainties of a perturbative calculation.
- It's "mandatory" if NLO corrections are very large to check the behaviour of the perturbative series
- It's the best we have! It is needed for Standard Candles and for really exploiting all the available information, for example that of NNLO PDF's.



## Drell-Yan Predictions at NLO



- At LO the W has no pt, therefore the pt of the lepton has a sharp cutoff.
- The "K-factor" looks like enormous at high PT. When this happens it means that the observable you are looking at it is actually at LO not at NLO!
- It is important to keep the spin correlations of the lepton in the calculation.


## Drell-Yan Predictions AT NNLO

- Virtual-Virtual : O(I00) terms
- Real-Virtual : O(300) terms

- Real-Real : O(500) terms



## Drell-Yan Predictions AT NNLO


[TH = Anastasiou, Dixon, Melnikov, Petriello. 2004]

- Impressive improvement of the scale dependence.
- High-pt end of the electron and extra jet known at NLO accuracy


## Higgs predictions at NNLO




- The perturbative series stabilizes.
- NLO estimation of higher orders effects by scale uncertainty works reasonably well


## Higgs predictions At NNLO


be careful : just illustrative example, not very precise

## Higgs predictions at 7 TeV




## PREDICTIONS AT NNLO : FINAL REMARKS

- Handful of precious predictions at NNLO now available for Higgs and Drell-Yan processes at the parton level for distributions.
- Others (VV, ttbar) in progress and in sight.


## NNLO stays to the LHC era <br> as <br> NLO stayed to the Tevatron era

