

Dark Matter Implications of the First CMS and ATLAS SUSY Results

Gregory Peim

Department of Physics, Northeastern University, Boston, MA 02115, USA

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- Based On Papers
 - **arXiv: 1103.1197:**
Sujeet Akula, Ning Chen, Daniel Feldman, Mengxi Liu, Zuowei Liu, Pran Nath, GP
 - **arXiv:1103.5061:**
SA, DF, ZL, PN, GP
- With Collaborators at
 - C.N. Yang Institute for Theoretical Physics at Stony Brook University
 - Michigan Center for Theoretical Physics at University of Michigan
 - Northeastern University

- 1 Motivation: LHC SUSY Results
- 2 Reach Dependence on A_0 and $\tan \beta$
- 3 Implications of LHC Results for mSUGRA Parameter Space
- 4 Dark Matter/LHC Implications for mSUGRA and NUSUGRA
- 5 Conclusion

Framework of analysis is mSUGRA

Model Parameters

- m_0 : scalar mass at GUT scale
- $m_{1/2}$: gaugino mass at GUT scale
- A_0 : Trilinear soft breaking parameter at GUT scale
- $\tan \beta$: ratio of the Higgs VEVs
- $\text{sign}(\mu)$: sign of the Higgs mixing parameter

mSUGRA with R-parity supplies a dark matter candidate, which is the lightest neutralino

Exp. Results: Reach for fix A_0 and $\tan\beta$

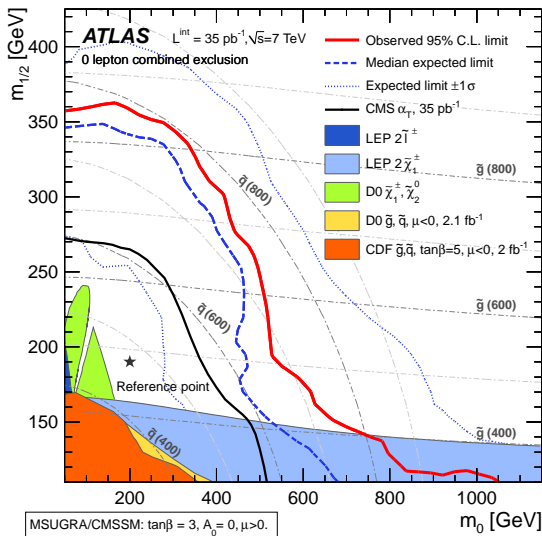
- How does the LHC reach change for different values of A_0 and $\tan\beta$?
- How much of the SUSY space has now been constrained?
- What is the effect on the dark matter direct detection search?

LHC SUSY Searches:

arXiv:1101.1628 (CMS)

arXiv:1102.2357 (ATLAS)

arXiv:1102.5290 (ATLAS; exhibited to right)



Identification (and Trigger) Criteria

- $H_T = \sum_{i=1}^{N_{jet}} p_T(j_i) > 250 \text{ GeV}$
- Jets (j_i): $p_T > 50 \text{ GeV}$ and $|\eta| < 3$
 - $|\eta(j_1)| < 2.5$
 - $p_T(j_1) > 100 \text{ GeV}$
 - $p_T(j_2) > 100 \text{ GeV}$
- Leptons (e, μ): $p_T > 10 \text{ GeV}$
- Photons: $p_T > 25 \text{ GeV}$

Event Selection

- $n(\ell) = 0$
- $n(\gamma) = 0$
- $n(j) \geq 2$
- $H_T > 350 \text{ GeV}$
- $\alpha_T = p_T^{\tilde{j}_2} / M_T > 0.55$
 - If $N_{jet} > 2$ then two pseudo-jets, \tilde{j}
 - \tilde{j} is the scalar sum of jets such that the p_T difference between \tilde{j} is minimized
- $\Delta R_{ecal} > 0.3$ and $\Delta\phi^*(j, \cancel{E}_T^{reco}) > 0.5$
- $R_{miss} = \cancel{E}_T^{reco} / \cancel{E}_T^{cal} < 1.25$

Identification Criteria

- Leptons: $p_T > 20$ GeV
 - $|\eta^e| < 2.47$ and vetoed if $1.37 < |\eta^e| < 1.52$
 - if $\Delta R(x, \mu) < 0.2$ then $p_T^x < 1.8$ GeV
 - $|\eta^\mu| < 2.4$
- Jets: $p_T > 20$ GeV and $|\eta| < 2.5$
 - jet discarded if $\Delta R(j, e) > 0.2$
 - $\Delta R(j, \ell) > 0.4$
- $\cancel{E}_T^{reco} > 125$ GeV
 - η^j extended to $|\eta^j| < 4.9$

Event Selection

- $n(\ell) = 1$
- $n(j) \geq 3$
 - $p_T(j_1) > 60$ GeV
 - $p_T(j_2) > 30$ GeV
 - $p_T(j_3) > 30$ GeV
- $\Delta\phi(j_i, \cancel{E}_T^{reco}) > 0.2$
 - for $i \in \{1, 2, 3\}$
- $m_T(\ell, \cancel{E}_T^{reco}) > 100$ GeV
- $\cancel{E}_T^{reco} > 0.25 m_{eff}$
 - $m_{eff} = \sum_{i=1}^3 p_T^{j_i} + p_T^\ell + \cancel{E}_T^{reco}$
- $m_{eff} > 500$
- Regions by m_T and \cancel{E}_T^{reco}

Identification Criteria

- Leptons: $p_T > 10$ GeV
 - $|\eta^e| < 2.47$ and vetoed if $1.37 < |\eta^e| < 1.52$
 - if $\Delta R(x, \mu) < 0.2$ then $p_T^x < 1.8$ GeV
 - $|\eta^\mu| < 2.4$
- Jets: $p_T > 20$ GeV and $|\eta| < 2.5$
 - jet discarded if $\Delta R(j, e) > 0.2$
 - $\Delta R(j, \ell) > 0.4$
- $\cancel{E}_T^{reco} > 100$ GeV
 - η^j extended to $|\eta^j| < 4.9$

Event Selection

- $n(\ell) = 0$
- $n(j) \geq i$
 - Region A and B: $i = 2$
 - Region C and D: $i = 3$
- $p_T^j > 120$ GeV
- $p_T^{jm} > 40$ GeV for $m \leq i$
- $\Delta\phi(j, \cancel{E}_T^{reco}) > 0.4$
- $(\cancel{E}_T^{reco} / m_{eff})^{\{A; \{C, D\}\}} > \{0.3; 0.25\}$
- $(m_{eff})^{\{\{A, C\}; D\}} > \{500 \text{ GeV}; 1000 \text{ GeV}\}$
- $(m_{T2})^{\{B\}} > 300$ GeV

Calculating allowed contributions of new physics to the observed number of events, N_{obs}

- Apply LHC cuts to SM background (b) and Signal (s)
- Calculate p -value; where $CL \equiv 1 - p$ and CL is the confidence level to upper limits
- α_T and 1 lepton search: apply a single-channel method
- 0 lepton search: we use a likelihood-based approach, which corresponds to a product of likelihoods of the form of Eq.(1) that is then maximized

$$p = \frac{\int_0^\infty db ds \sum_{n=0}^{N_{\text{obs}}} G(s; N_s, \delta N_s) G(b; N_b, \delta N_b) P(n | b + s)}{\int_0^\infty db \sum_{n=0}^{N_{\text{obs}}} G(b; N_b, \delta N_b) P(n | b)} \quad (1)$$

For more details see: arXiv:0901.0512; FERMILAB-CONF-00-048-E

p -value (Cont.)

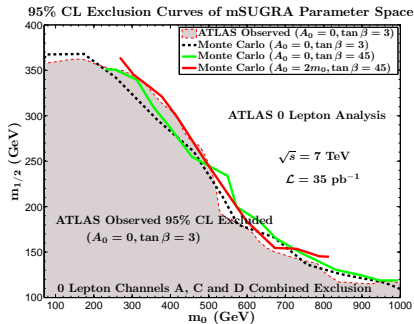
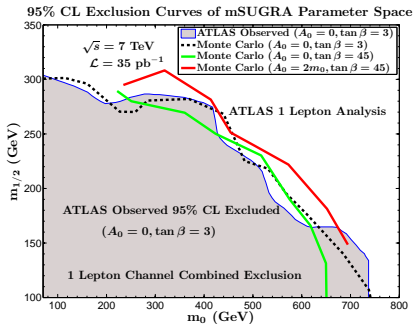
$$p = \frac{\int_0^\infty db ds \sum_{n=0}^{N_{\text{obs}}} G(s; N_s, \delta N_s) G(b; N_b, \delta N_b) P(n | b + s)}{\int_0^\infty db \sum_{n=0}^{N_{\text{obs}}} G(b; N_b, \delta N_b) P(n | b)}$$

- $G(x; N_x, \delta N_x)$ is the Gaussian probability density function of observing N_x events with a standard deviation of δN_x
- $P(n | x)$ is the Poisson distribution of observing n events given the events x .

For more details see: arXiv:0901.0512; FERMILAB-CONF-00-048-E

Vary A_0 and $\tan \beta$

Using the cuts reported in arXiv:1102.2357 (left) and arXiv:1102.5290 (right) we see good agreement with $A_0 = 0$ and $\tan \beta = 3$



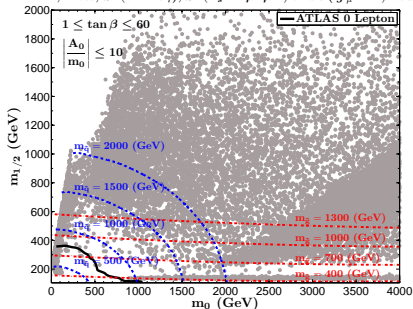
- Exclusion boundary does not change much when A_0 or $\tan \beta$ are varied
- Main changes are at the ends due to LEP, EWSB and $\tilde{\tau}_1$ LSP
- Consistent to use LHC exp. curve for random inputs

How does LHC data effect parameter space

Previous results only needed to pass REWSB, so what if we apply (non-LHC) collider, flavor and astrophysics constraints? (K. Nakamura *et al.*, and

arXiv: 1001.4538, hep-ph/0609232, hep-ph/0602001, 1006.3469, 0808.1297)

WMAP, Mass, $Br(b \rightarrow s\gamma)$, $Br(B_s \rightarrow \mu^+\mu^-)$ and $\delta(g_\mu - 2)$ bounds



$$0.0896 < \Omega_\chi h^2 < 0.1344$$

$$m_h > 93.5 \text{ GeV}$$

$$m_{\tilde{\tau}_1} > 81.9 \text{ GeV}$$

$$m_{\tilde{\chi}_1^\pm} > 103.5 \text{ GeV}$$

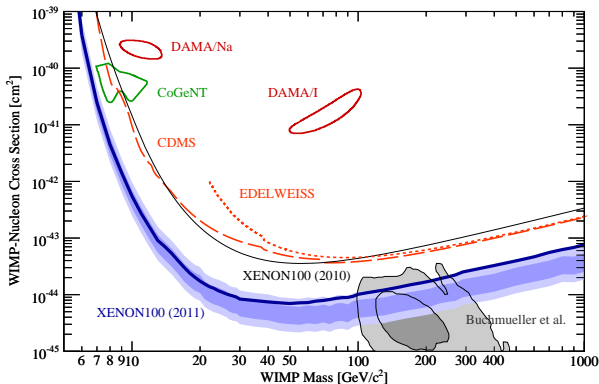
$$m_{\tilde{t}_1} > 100 \text{ GeV}$$

$$Br(B_s \rightarrow \mu^+\mu^-) \leq 4.2 \times 10^{-8}$$

$$(-11.4 \times 10^{-10}) \leq \delta(g_\mu - 2) \leq (9.4 \times 10^{-9})$$

$$(2.77 \times 10^{-4}) \leq Br(b \rightarrow s\gamma) \leq (4.37 \times 10^{-4})$$

Implications of LHC on Dark Matter Direct Detection



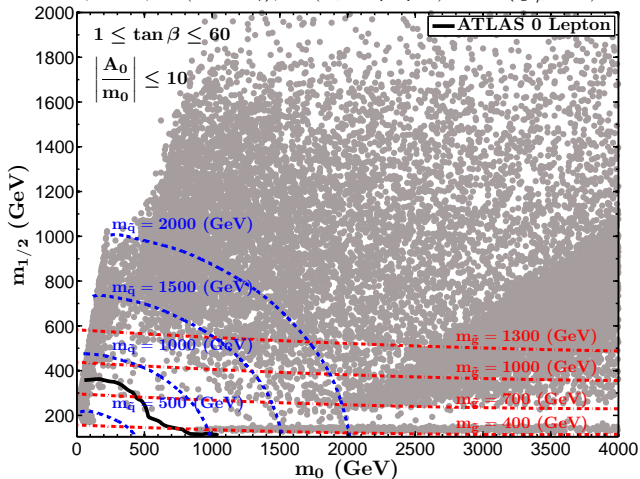
arXiv:1104.2549 (XENON)

What are the implications of the LHC Results on XENON-100?

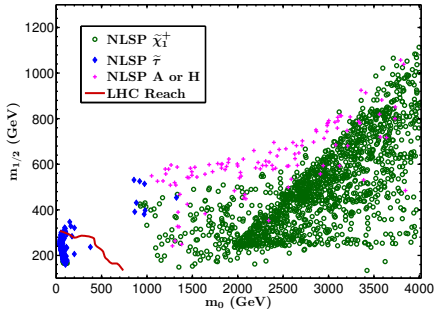
XENON-100 Results on mSUGRA Parameter Space

Before XENON-100 constraint is applied we have :

WMAP, Mass, $\mathcal{B}r(b \rightarrow s\gamma)$, $\mathcal{B}r(B_s \rightarrow \mu^+\mu^-)$ and $\delta(g_\mu - 2)$ bounds

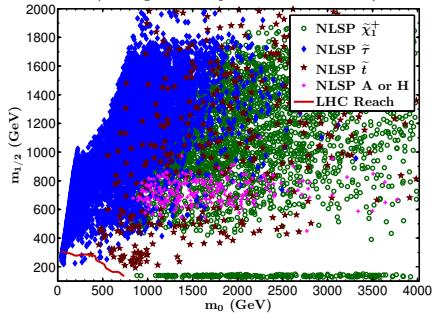


mSUGRA Models Above XENON-100 σ_{SI} Limit
(Passing Other Experimental Constraints)



ABOVE XENON-100

mSUGRA Models Below XENON-100 σ_{SI} Limit
(Passing Other Experimental Constraints)



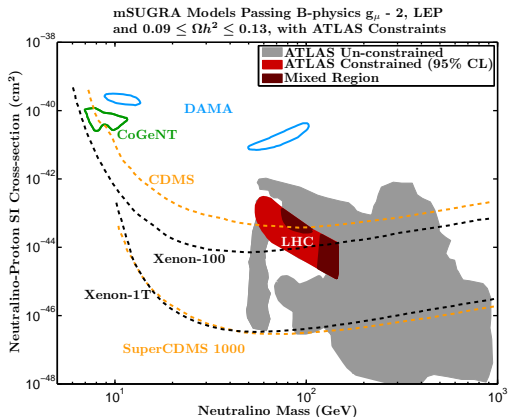
BELOW XENON-100

Models Most Constrained

- slepton coannihilation models (LHC and XENON)
- Hyperbolic Branch/Focus Point region (XENON)

There are significant uncertainties in the value of the strange quark form factors f_s as shown by arXiv:0907.4177. The uncertainty can give a significantly smaller σ_{SI} which would relax the constraints

LHC Constraint on Dark Matter Direct Detection

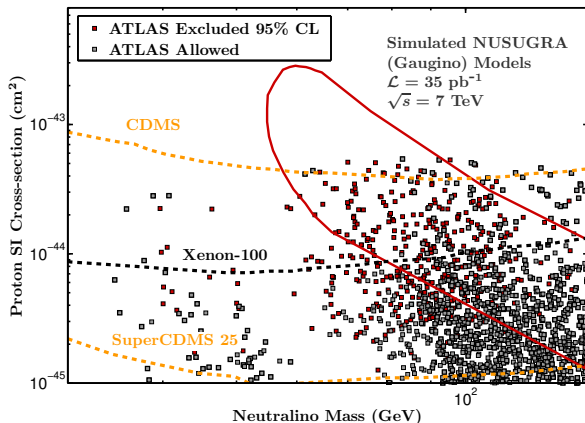


LHC SUSY Search

- **Models in the Red Region** have been **constrained**
- **~ 60%** of models in the **Maroon Regions** have been **constrained**
- **Models in the Gray Region** are **unconstrained**

ATLAS Excluded Region Repopulated by NUSUGRA

ATLAS 0 Lepton Analysis of NUSUGRA (Gaugino) Models
Consistent with B-physics, $g_\mu - 2$, LEP and $0.09 \leq \Omega h^2 \leq 0.13$



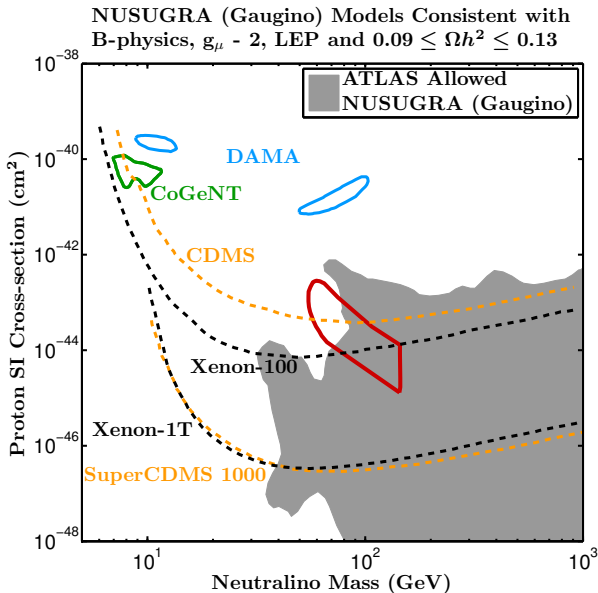
$m_\alpha = m_{1/2} (1 + \delta_\alpha)$ for $\alpha \in \{1, 2, 3\}$ and $-1 < \delta_\alpha < 1$
mSUGRA excluded region is now repopulated

Conclusion

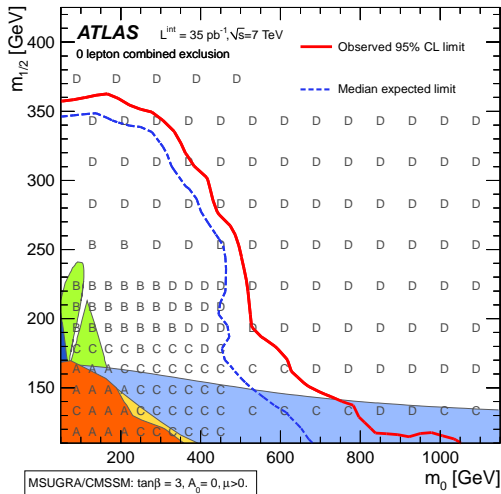
- With 35 pb^{-1} of integrated luminosity the LHC has probed a portion of the mSUGRA parameter space and has constrained a portion of the slepton coannihilation region
- Recent XENON-100 results have put additional constraints on the mSUGRA parameter space
 - Constrained mostly the HB/FP region
 - The neutralino-proton SI cross-section is sensitive to the strange quark form factors, which has large uncertainties.
- Repopulation of mSUGRA region excluded by LHC or direct detection experiments could hint at non-universalities

Additional Slides

NUSUGRA before LHC constraints are applied

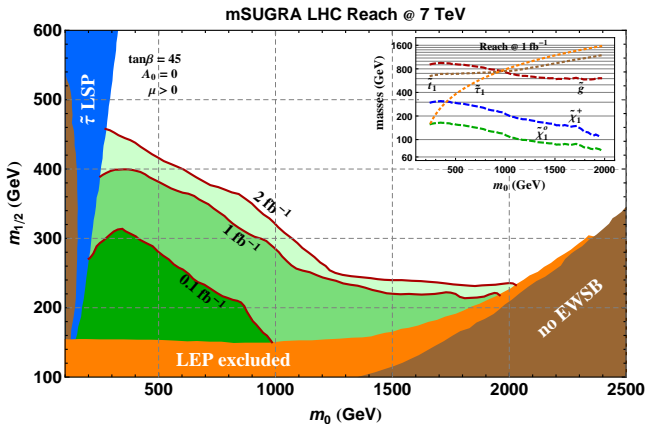


Points Marked By Channel that Best Excludes Region



http://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/PAPERS/susy-0lepton_01

Projected Reach at 7 TeV



arXiv:1008.3423

LHC: SM Background

SM process	Cross section (fb)	SM process	Cross section (fb)
QCD 2, 3, 4 jets	2.0×10^{10}	$t\bar{t}\bar{t}$	0.5
$Z/\gamma + b\bar{b} + \leq 2$ jets	2.8×10^3	$t\bar{t}b\bar{b}$	1.2×10^2
$b\bar{b} + \leq 2$ jets	9.5×10^7	$b\bar{b}b\bar{b}$	2.2×10^4
$Z/\gamma + \leq 3$ jets	6.2×10^6	$W^\pm W^\pm$	2.0×10^3
$W^\pm + \leq 3$ jets	1.9×10^7	$W^\pm Z$	1.1×10^3
$W^\pm + t\bar{t} + \leq 2$ jets	70	$\gamma + \leq 3$ jets	1.5×10^7
$Z/\gamma + t\bar{t} + \leq 2$ jets	56	$t\bar{t} + \leq 2$ jets	1.6×10^5
$W^\pm + b\bar{b} + \leq 2$ jets	3.2×10^3	ZZ	7.3×10^2
$W^\pm + t\bar{b}(\bar{t}b) + \leq 2$ jets	2.4×10^2		

SM background was generated at 7 TeV using MadGraph 4.4 (J. Alwall *et al.*).

For more detail see arXiv:1008.3423