

Photon acceleration as a scattering process

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Outline

- **Historical background;**
- **Photon acceleration models;**
- **Acceleration in a laser wakefields;**
- **Ionization fronts and particle beams;**
- **Scattering by relativistic plasma bubbles;**
- **Photon acceleration in gravitational fields;**
- **Generalized Sachs-Wolfe effect;**
- **Conclusions.**

Historical background

- **Relativistic mirror (Einstein, 1905)**
- **Moving ionization fronts:(Semenova, 1967; Lampe et al., 1978).**
- **Moving nonlinear perturbation: adiabatic frequency shift (Mendonça, 1979).**
- **Photon acceleration in a laser wakefield (Wilks et al., 1989).**
- Time refraction (Mendonça, 2000).**



Experimental evidence

- **Laser self blue shift: flash ionization (Yablomovich, 1974, Wood et al., 1993).**
- **Microwave experiments: ionization fronts (Savage et al., 1992).**
- **2D optical experiments (Dias et al., 1997).**
- **Photon trapping in a wakefield (Murphy et al., 2006).**

Related optical phenomena:

self and cross-phase modulation, super-continuum (70's); Unruh radiation!! (2010)



Theoretical models

- **Classical model:**

Geometric optics of space and time varying media

- **Kinetic model**

Photon kinetic theory (laser as a photon gas, photon Landau damping);

- **Quantum model**

Full wave theory (similar to scattering theory in Quantum Mechanics);

- **Second-Quantization model**

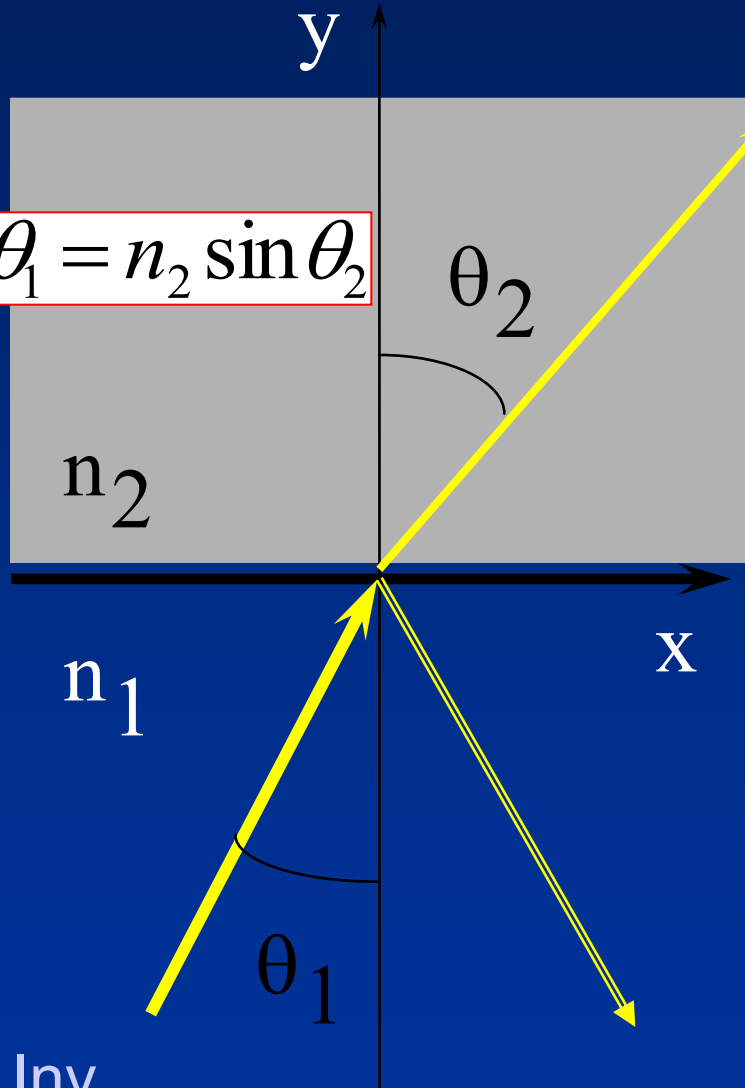
Quantum optics approach (theory of time refraction and temporal beam splitters);

- **Extension to other types of interaction**

neutrino-plasma physics, gravitational waves.

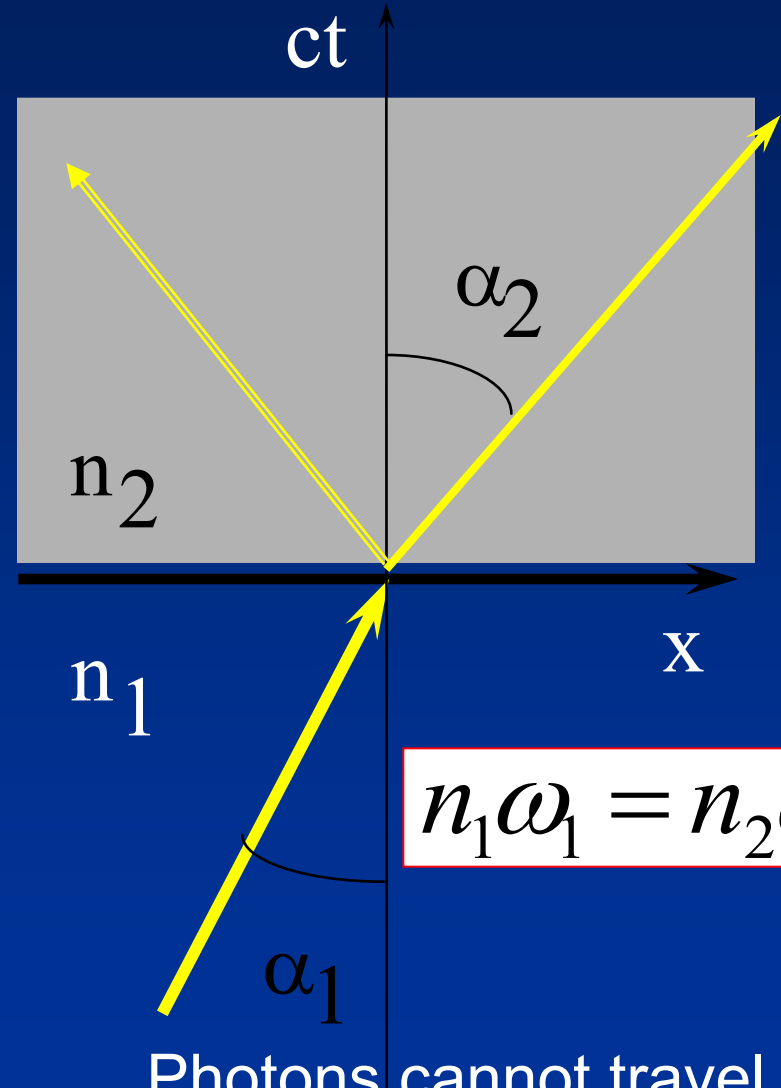
Refraction

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$



$W = \text{Inv}$

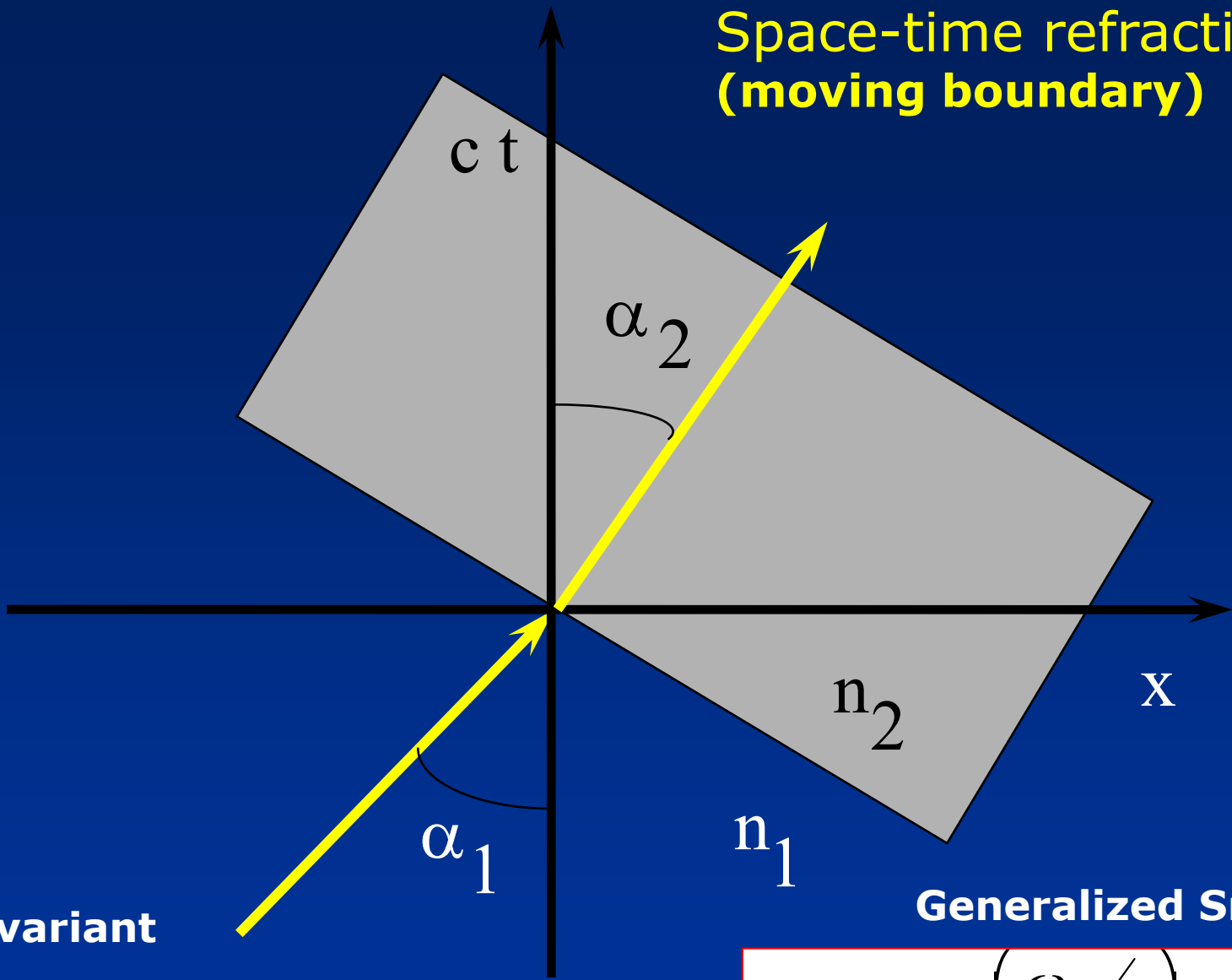
Time refraction



$$n_1 \omega_1 = n_2 \omega_2$$

Photons cannot travel
back in the past

Space-time refraction (moving boundary)



Invariant

$$\omega_j (1 - \beta n_j \cos \theta_j) = \text{Inv.}$$

Generalized Snell's law

$$n_1 \sin \theta_1 = \left(\frac{\omega_2}{\omega_1} \right) n_2 \sin \theta_2$$



Photon ray equations

- Photons \equiv relativistic particles with effective mass
- Photon Dynamics \equiv canonical equations of motion

$$\frac{d\mathbf{r}}{dt} = \frac{\partial \omega}{\partial \mathbf{k}} = \frac{\mathbf{k}}{\omega} c^2$$
$$\frac{d\mathbf{k}}{dt} = -\frac{\partial \omega}{\partial \mathbf{r}} = -\frac{1}{2\omega} \nabla \omega_p^2$$

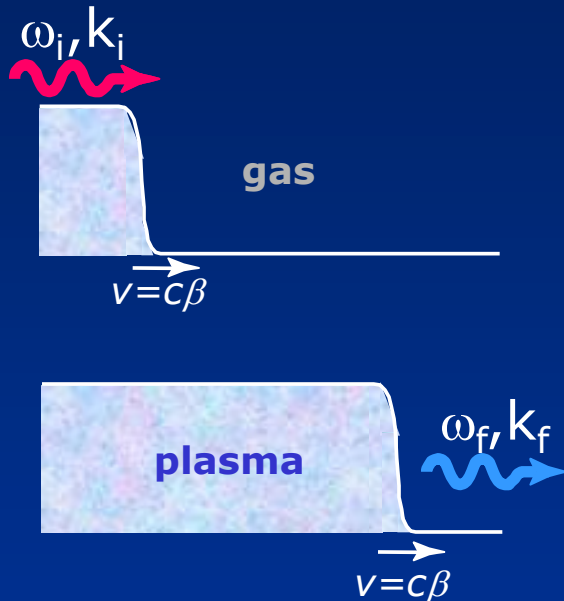
Hamiltonian

$$\omega = \sqrt{\mathbf{k}^2 c^2 + \omega_p^2}$$

Force acting on the photon: refraction (inhomogeneous plasma) and photon acceleration (non-stationary plasma).

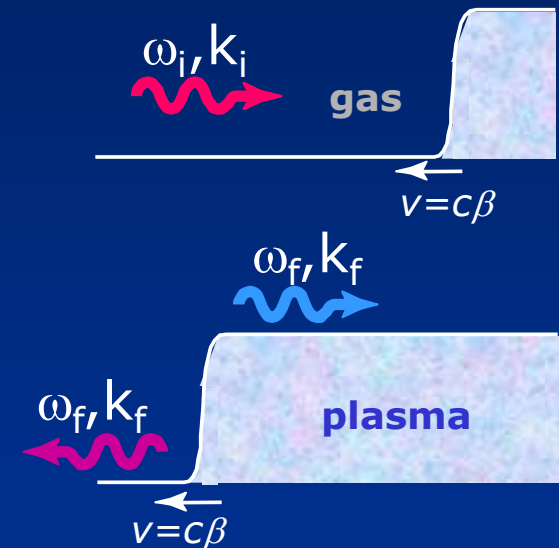
Photon interaction with ionization fronts

Co-propagation (-)



$$\Delta\omega \approx \frac{\omega_{po}^2}{2\omega_0} \frac{\beta}{1 \pm \beta}$$

Counter-propagation (+)



❖ Reflection in counter-propagation
(relativistic mirror)

$$\Delta\omega = \omega_0 \frac{2\beta}{1 - \beta}$$

❖ Plasma formation (flash ionization) $\beta \rightarrow \infty$



$$\Delta\omega \approx \frac{\omega_{po}^2}{2\omega_0}$$

Intense laser pulse in a gas jet

Relativistic ionization fronts

Shadow images of a relativistic front

t_0



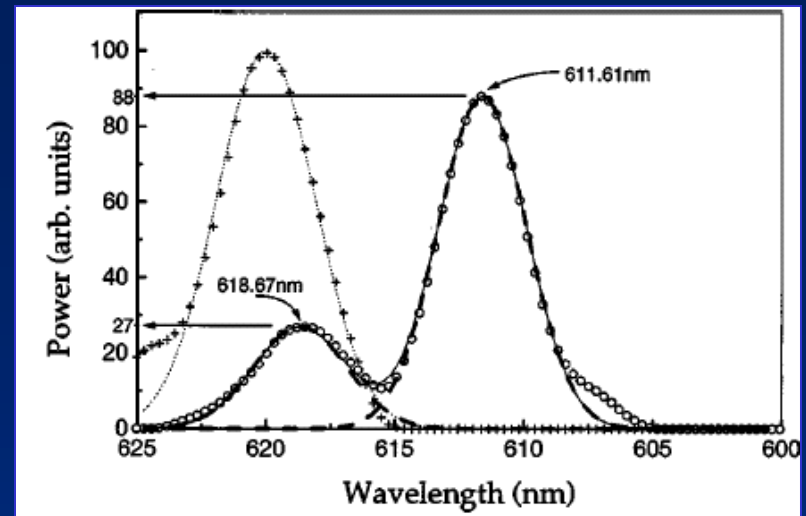
t_1



(Experiments done in collaboration with LOA, France)

Photon acceleration

Laser pulse
65 fs, 2.5 mJ @ 620 nm

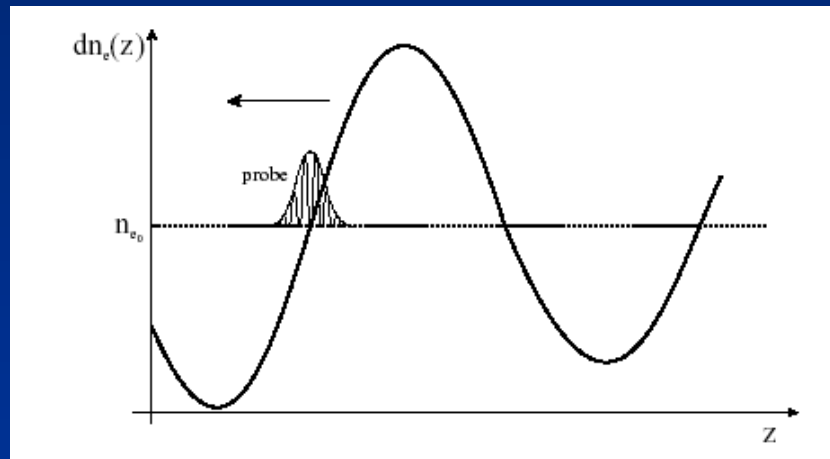


Simultaneous measurement:
 $\beta=0.942$, $n=4.26 \times 10^{19} \text{ cm}^{-3}$

Dias et al. PRL (1997)

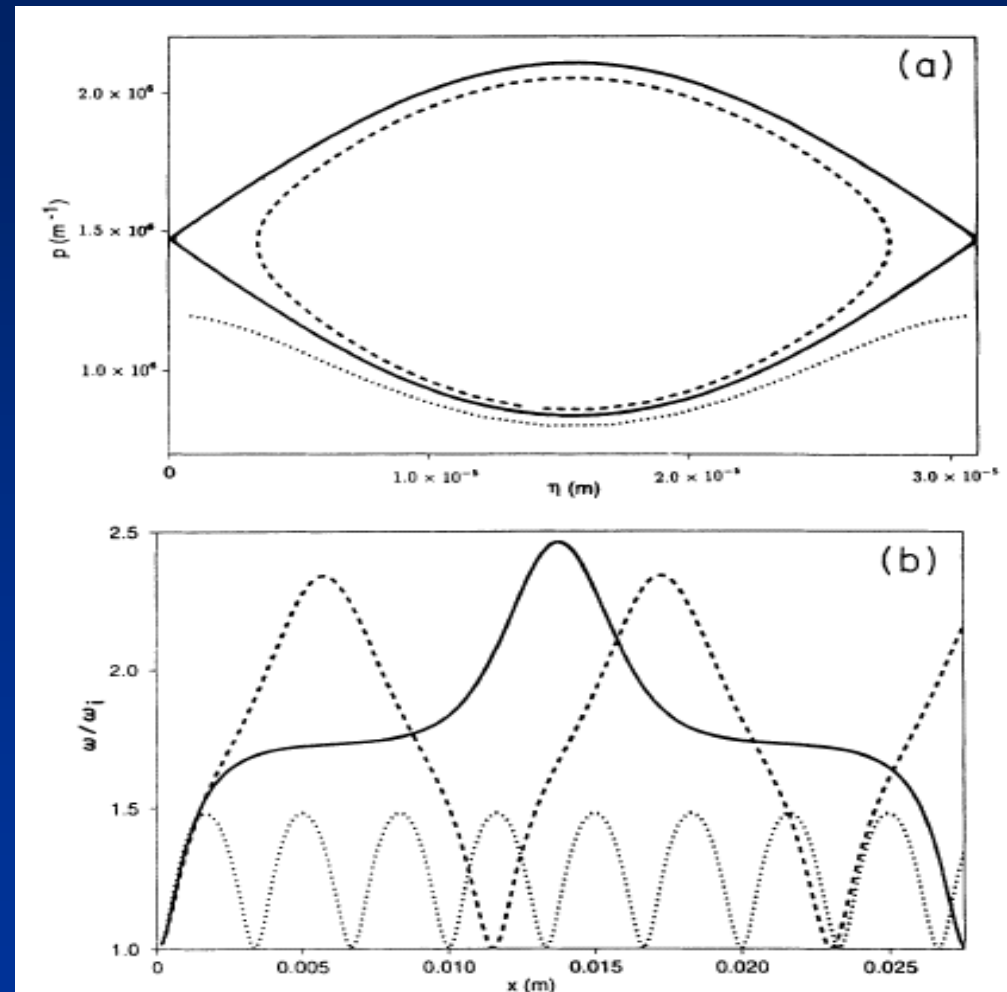
Photon dynamics in a wake field

Laser probe diagnostic for laser accelerators



Wakefield = relativistic electron plasma wave produced by a pump laser pulse

Phase space plot



frequency shift

Photon trapping in laser wakefield

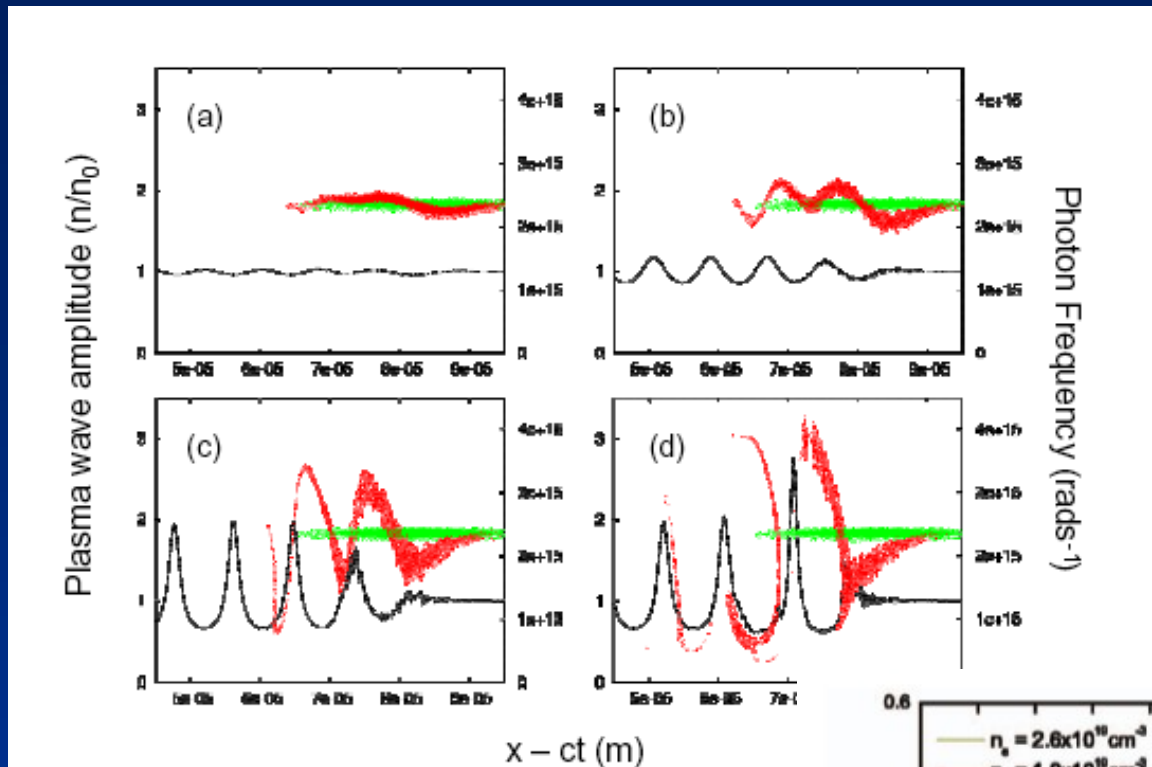
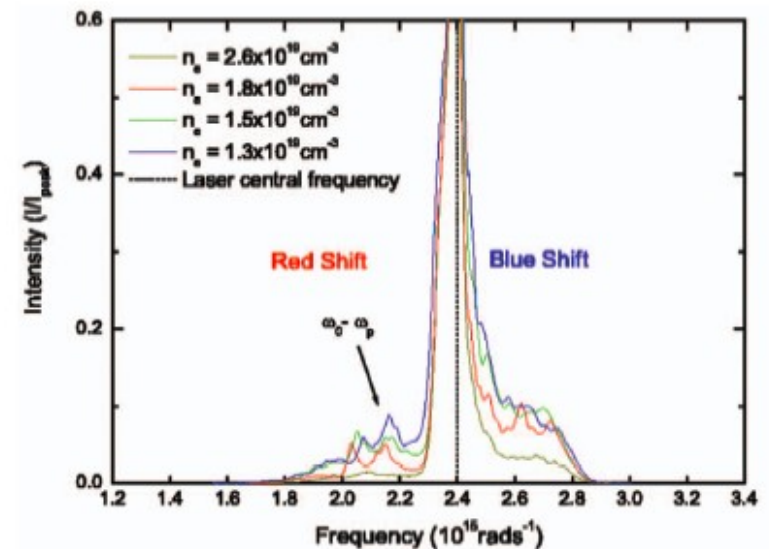


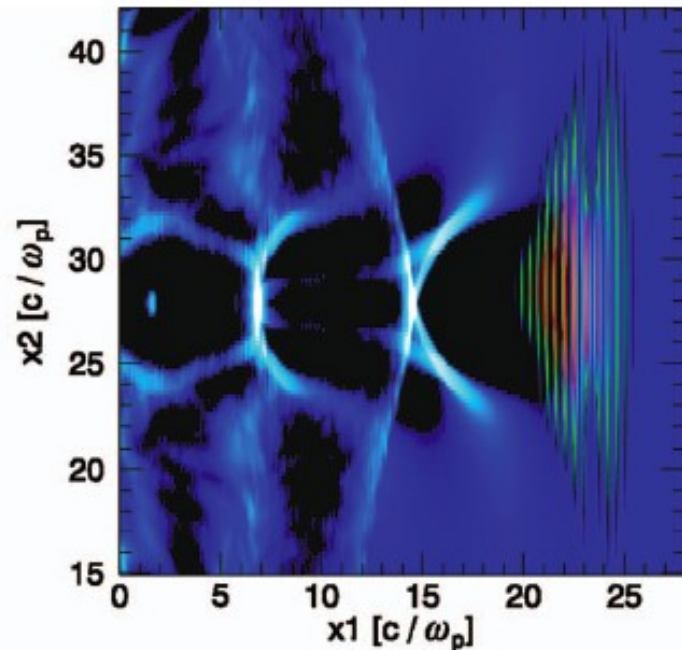
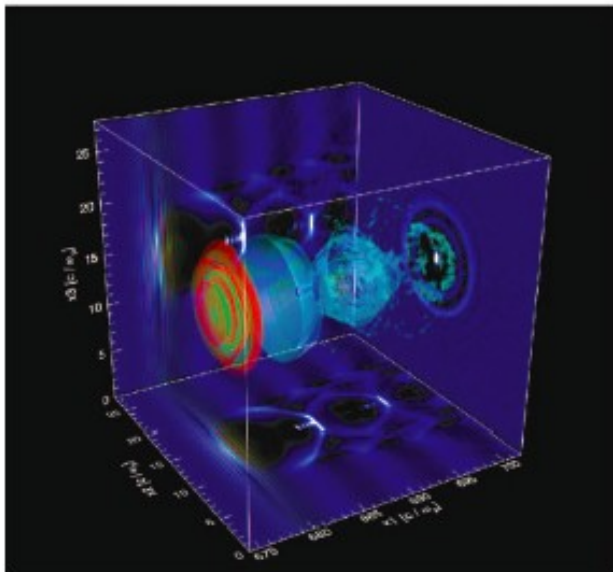
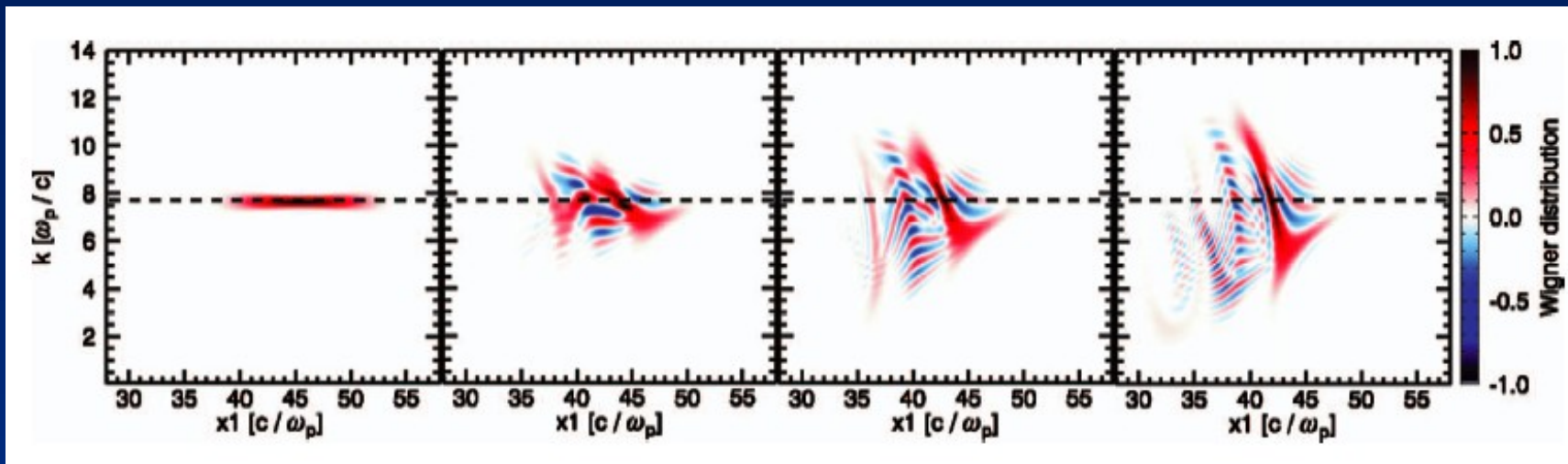
FIG. 2: Phase space plot when the laser pulse length is a) 200 b) 500 c) 750 and d) 1000 μm in a plasma with $n_0 = 3.6 \times 10^{19} \text{ cm}^{-3}$



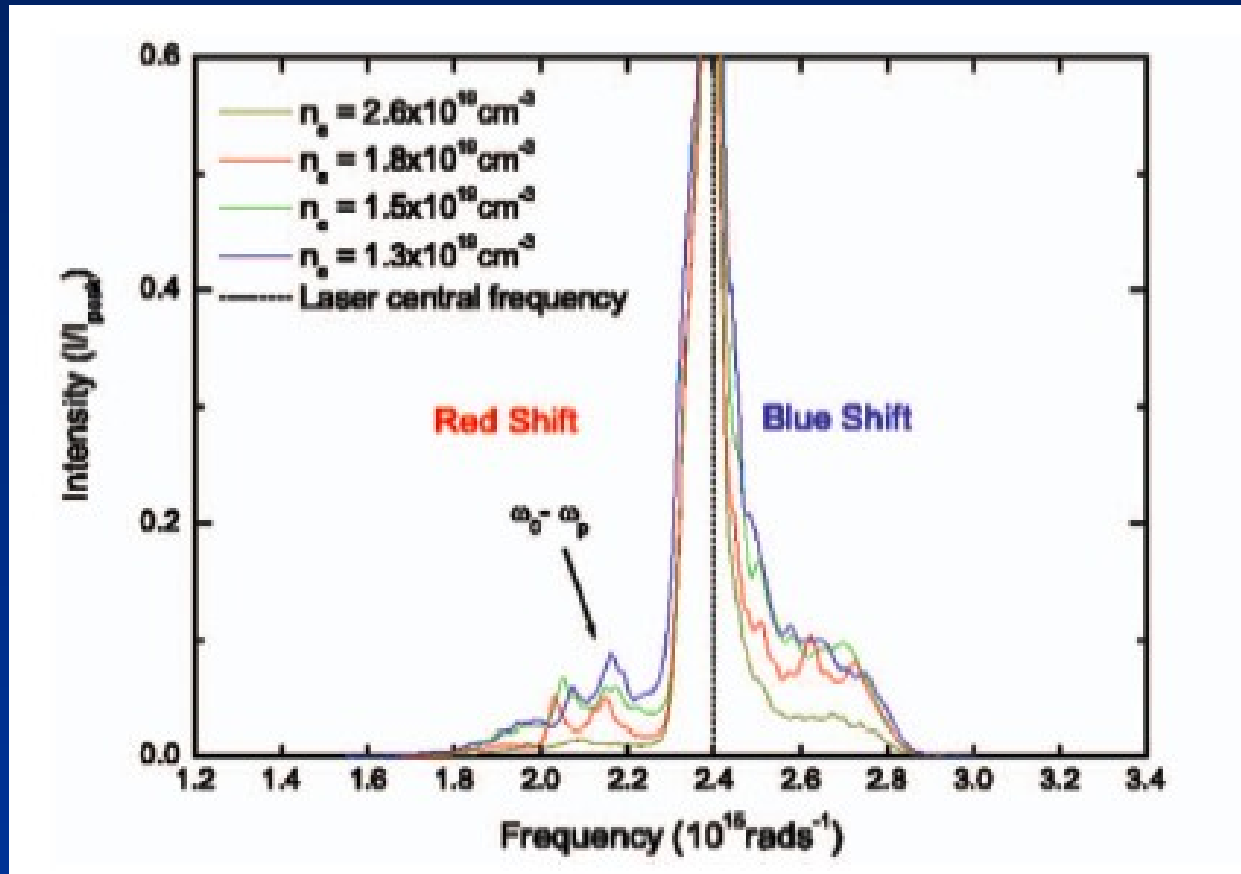
R. Trines (simulations)

C. Murphy (experiments)

Confirmed by PIC code simulations



Laser wake field experiments at RAL



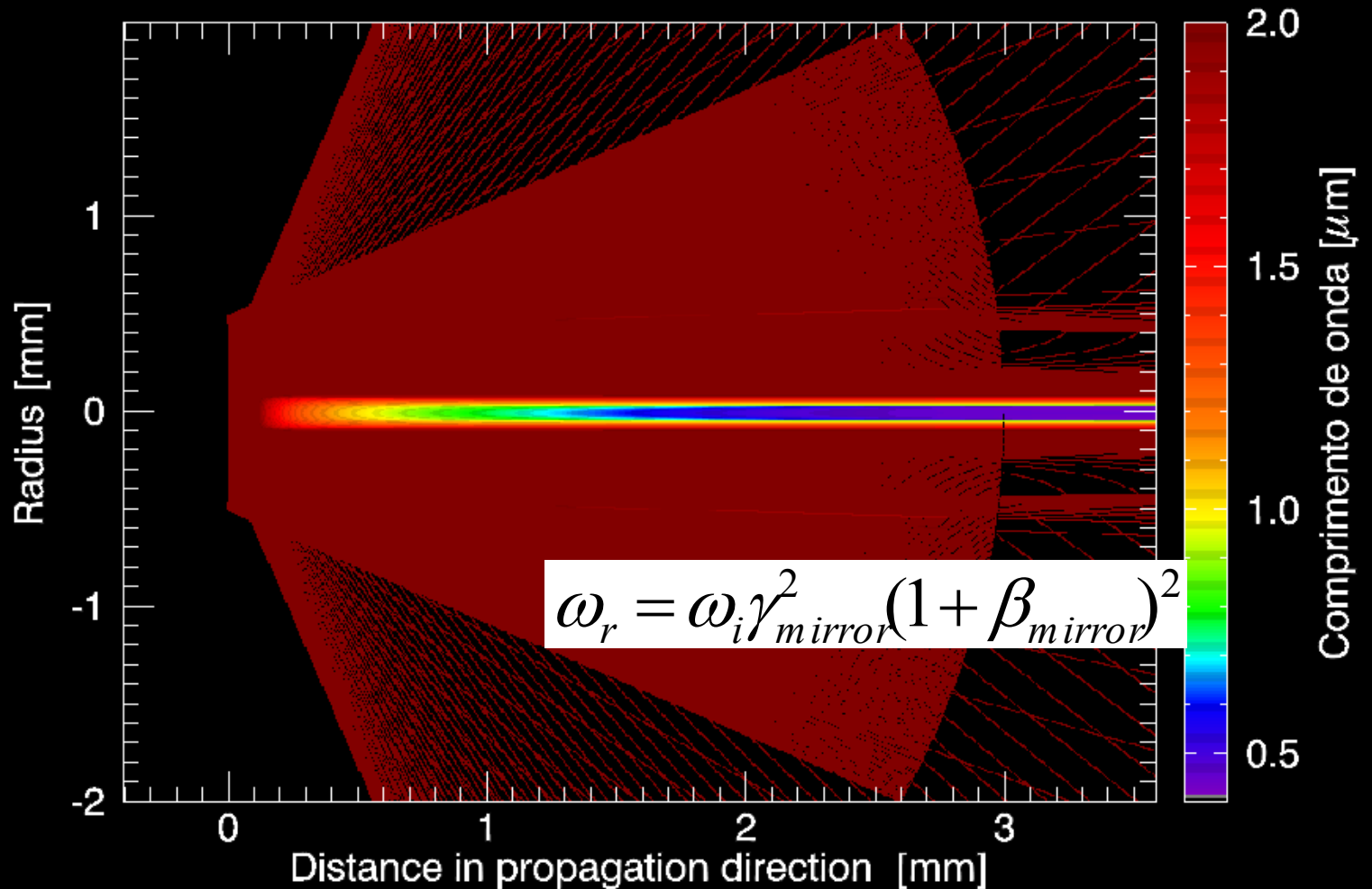
C. Murphy et al. PoP (2006)

Photon scattering by a ionization front

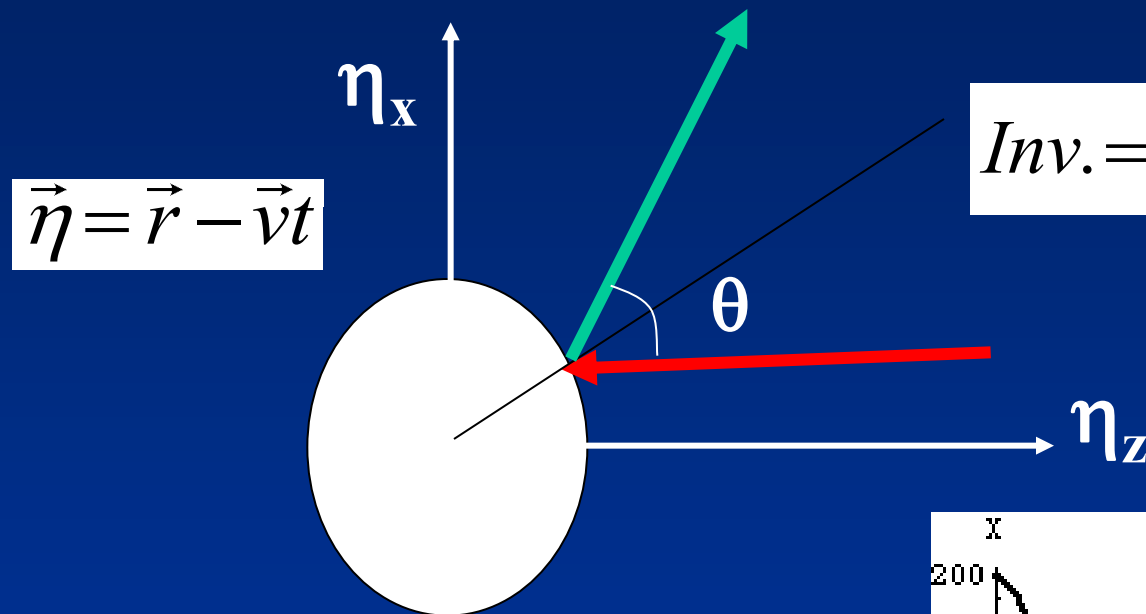
Relativistic mirror for THz radiation
Strathclyde laser

Beta
0.997

n_{Ar}
 $1 \times 10^{19} \text{ cm}^{-3}$



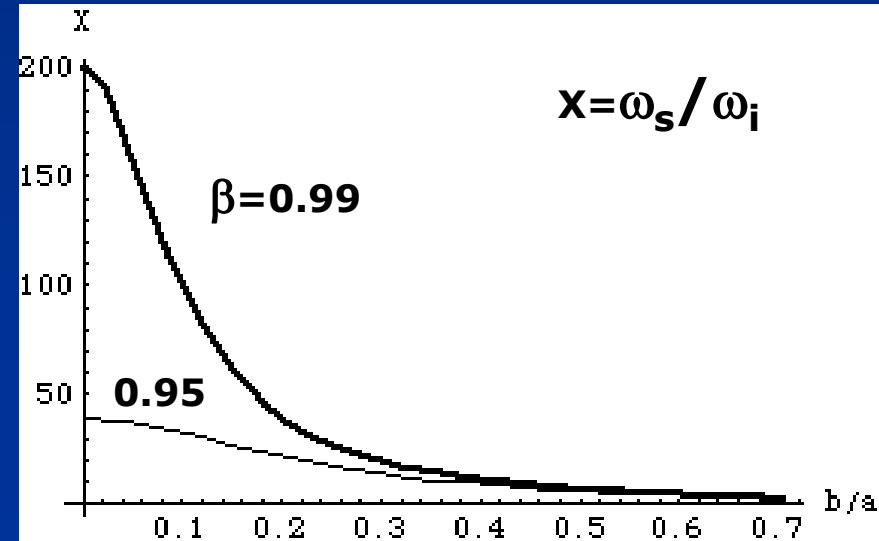
Photon Scattering by a relativistic plasma bubble



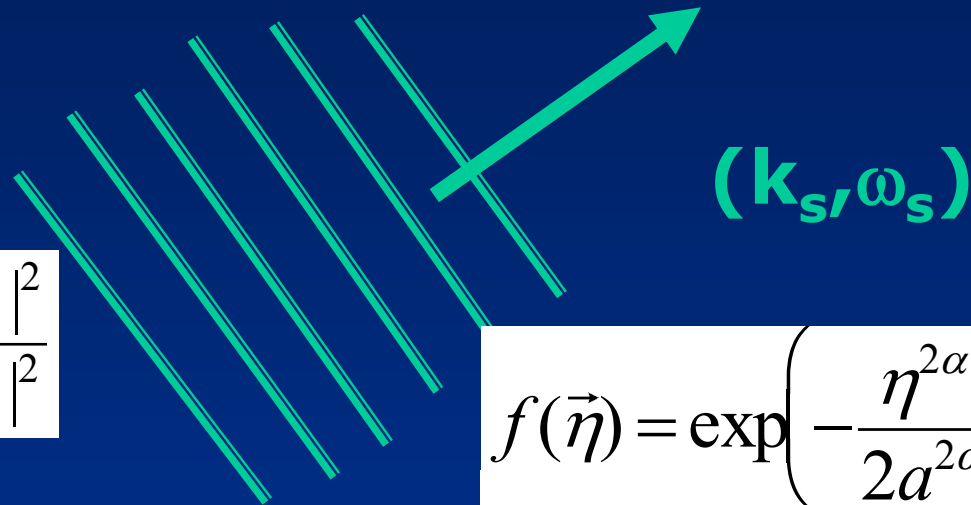
$$Inv. = \sqrt{k^2 c^2 + \omega_p^2(\vec{\eta})} - \vec{v} \cdot \vec{k}$$

Scattered frequency

$$\omega_s = \omega_0 \frac{1 + \beta \sqrt{1 + \omega_{p0}^2 / \omega_0^2}}{1 - \beta \cos \theta \sqrt{1 + \omega_{p0}^2 / \omega_s^2}}$$

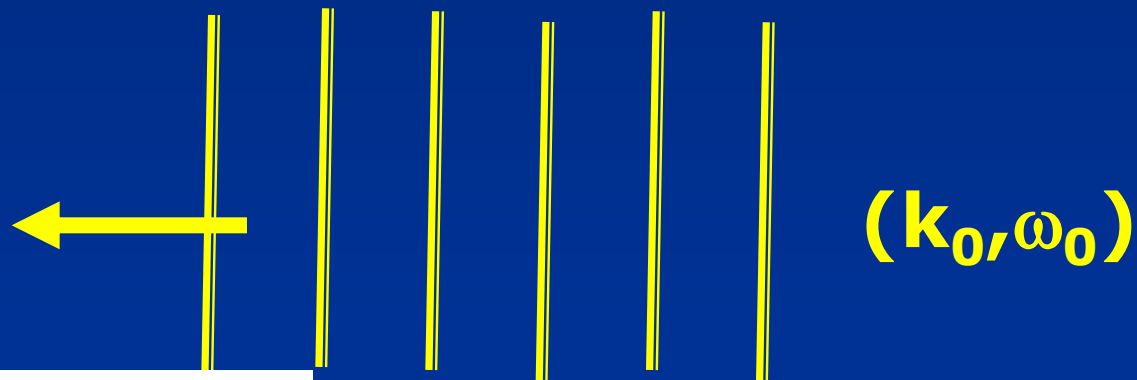
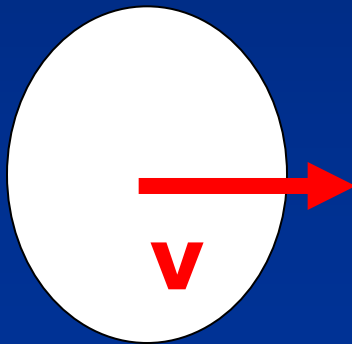


Photon-bubble scattering problem



$$S(\omega_0, \omega, \theta) = \frac{|E(\omega, \theta)|^2}{|E_0(\omega_0)|^2}$$

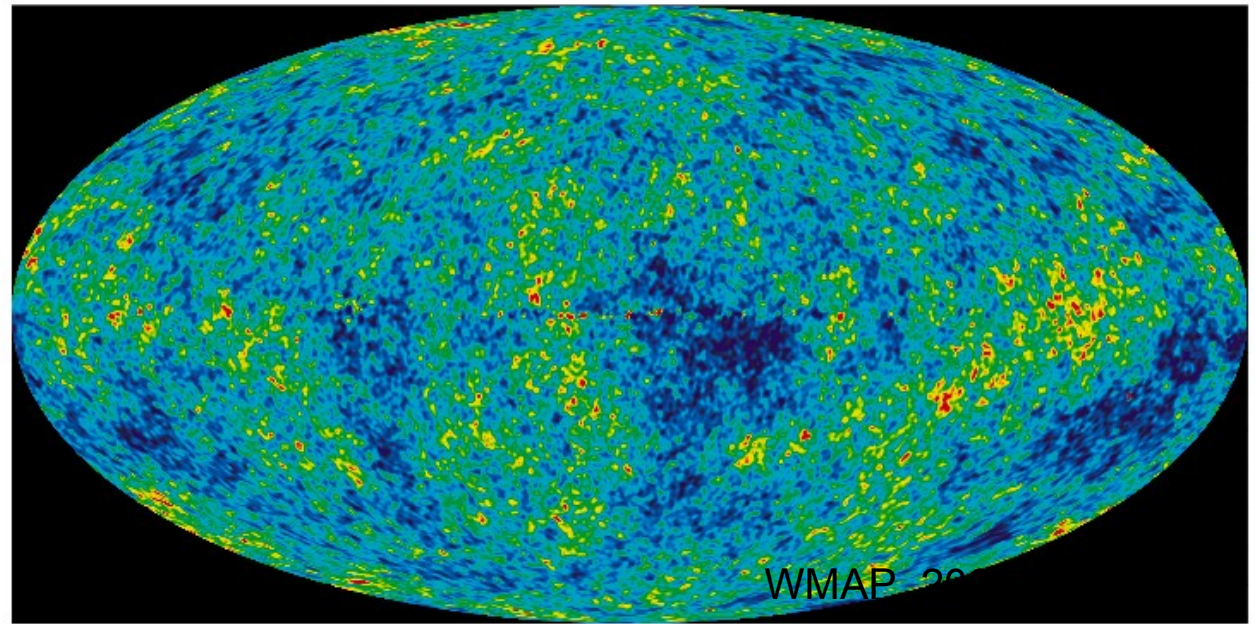
$$f(\vec{\eta}) = \exp\left(-\frac{\eta^{2\alpha}}{2a^{2\alpha}}\right), \quad \alpha = 1, 2$$



$$\omega_p^2 = \omega_{p0}^2 [1 - \mathcal{E}f(\vec{r} - \vec{v}t)]$$

Photon acceleration in a gravitational field scenario

Cosmic Temperature map



Sachs-Wolfe effect is photon acceleration

QuickTime™ and a
TIFF (Uncompressed) decompressor
are needed to see this picture.

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TIFF (Uncompressed) decompressor
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Low plasma density limit

QuickTime™ and a
TIFF (Uncompressed) decompressor
are needed to see this picture.

Low frequency shift limit

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Mendonça, Bingham and Wang, CQG (2008)

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Moving perturbations

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Dynamical invariant

TIFF (Uncompressed) decompressor
are needed to see this picture.

Generalized Sachs-Wolfe

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are needed to see this picture.

Purely vacuum perturbations

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This is not a Doppler correction!

Photons in a Gravitational wave

$$g^{ij} = \eta^{ij} + h^{ij}$$

Perturbed flat space time

$$\eta^{\alpha\beta} = -\delta^{\alpha\beta}, \quad \eta^{00} = 1$$

Weak gravitational wave

$$a = A \sin(k_0 x^0 + k_1 x^1 + \phi) = A \sin(qx - \Omega t + \phi)$$

$$h_{22} = -h_{33} = a$$

Photon dispersion relation

$$\omega = kc \left\{ 1 + \frac{a}{2} \left[\left(\frac{k_y}{k} \right)^2 - \left(\frac{k_z}{k} \right)^2 \right] \right\}$$

Nearly parallel photon propagation

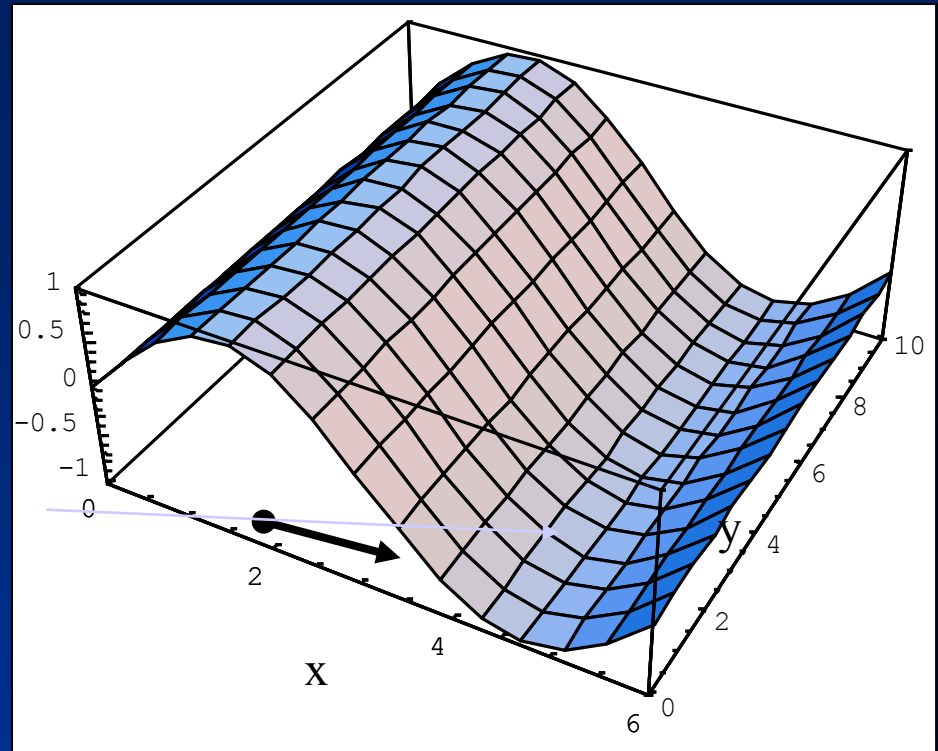
Perpendicular photon motion

$$y(t) = ck_y \int^t \frac{1 + a(t')}{k(t')} dt'$$

Parallel photon motion

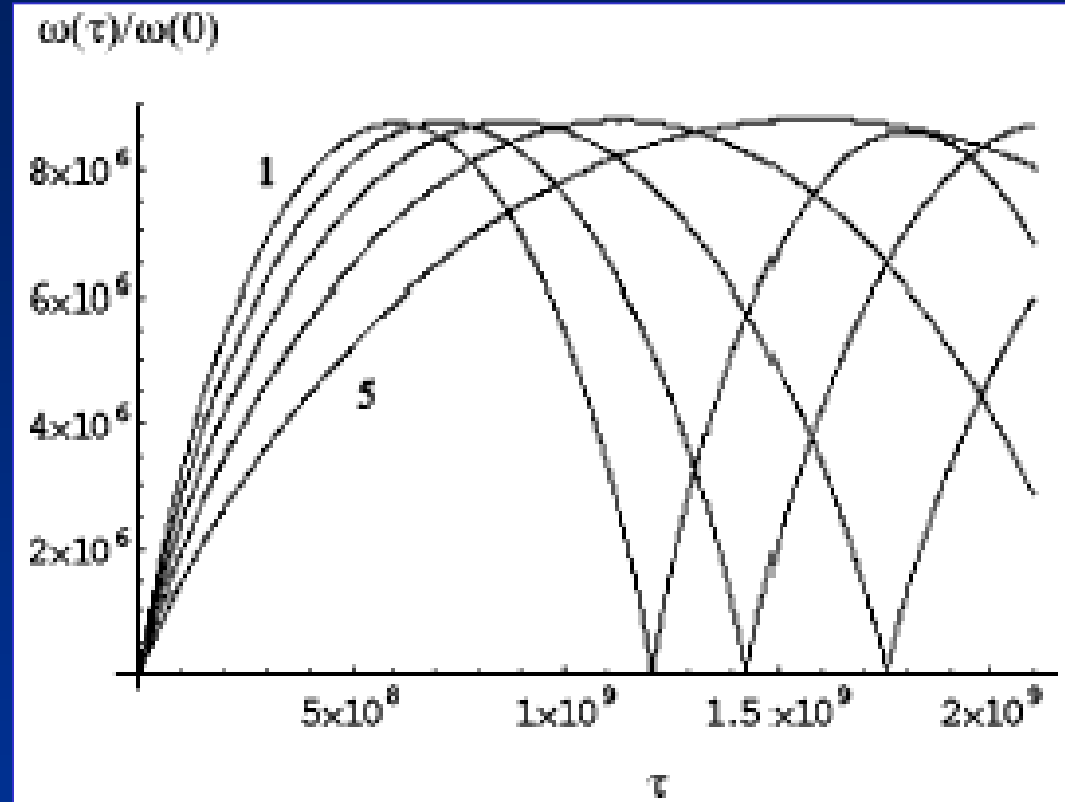
$$\frac{dx}{dt} = c \frac{k_x}{k}$$

$$\frac{dk_x}{dt} = -\frac{1}{2} \left(\frac{k_y}{k} \right)^2 \frac{\partial a}{\partial t}$$



Typical photon trajectories

$$\frac{d\omega}{dt} = c \frac{dk_x}{dt}$$



Photon acceleration by gravitational waves

Gravitational wave frequency: $\Omega = 10^4 \text{ s}^{-1}$

$$\tau = A c t / 2$$

and amplitude $A = 10^{-4}$

Conclusions

- **Photon acceleration is a first order effect;**
- **Experimental evidence: relativistic fronts, wakefields (plasmas), self and cross phase modulations (optics);**
- **It can be seen as space-time refraction;**
- **It can also be seen as a scattering process;**
- **Interaction with Gfields and Gwaves (γ ray bursts, Sachs-Wolfe effect);**
- **Applications: plasma diagnostics, ultra-short laser pulses, sub-cycle and attosecond optics, tunable radiation sources.**