

# Neutrino Physics

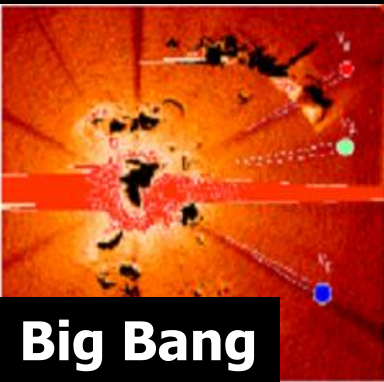
Zhi-zhong Xing  
(IHEP, Beijing)

- A1:** Neutrino's history & lepton families
- A2:** Dirac & Majorana neutrino masses
- B1:** Lepton flavor mixing & CP violation
- B2:** Neutrino oscillation phenomenology
- C1:** Seesaw & leptogenesis mechanisms
- C2:** Extreme corners in the neutrino sky

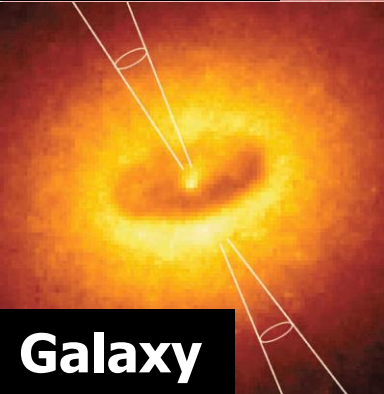


@ The 1<sup>st</sup> Asia-Europe-Pacific School of HEP, 10/2012, Fukuoka

# Neutrinos: How Elusive They Are?



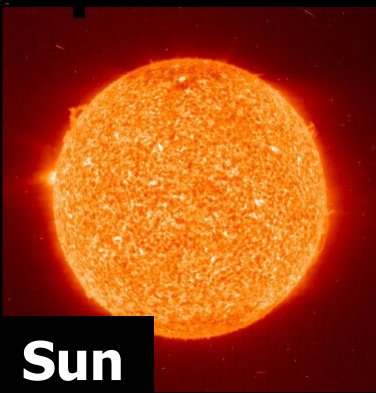
Big Bang



Galaxy



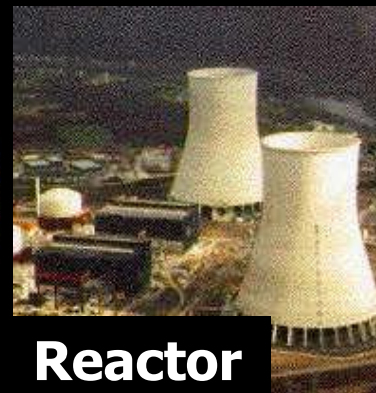
Supernova



Sun



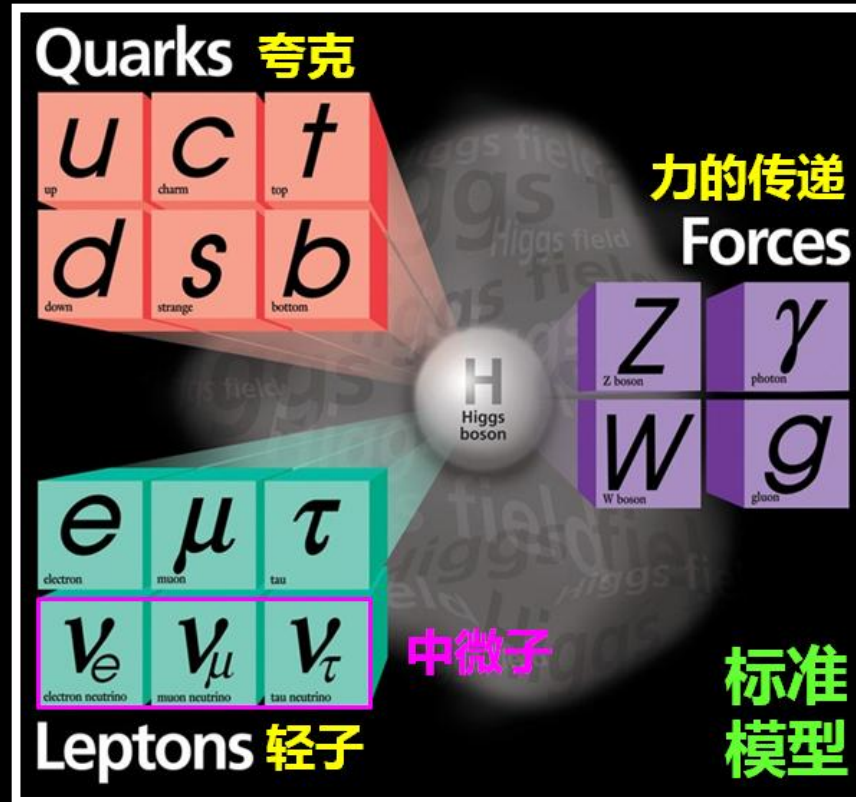
Earth



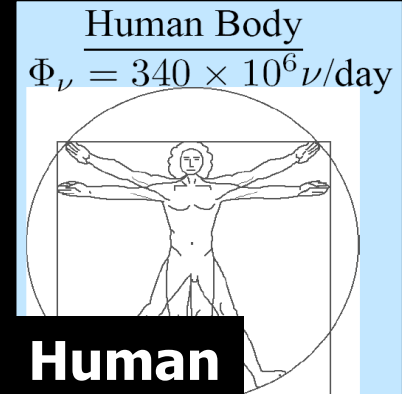
Reactor



Accelerator

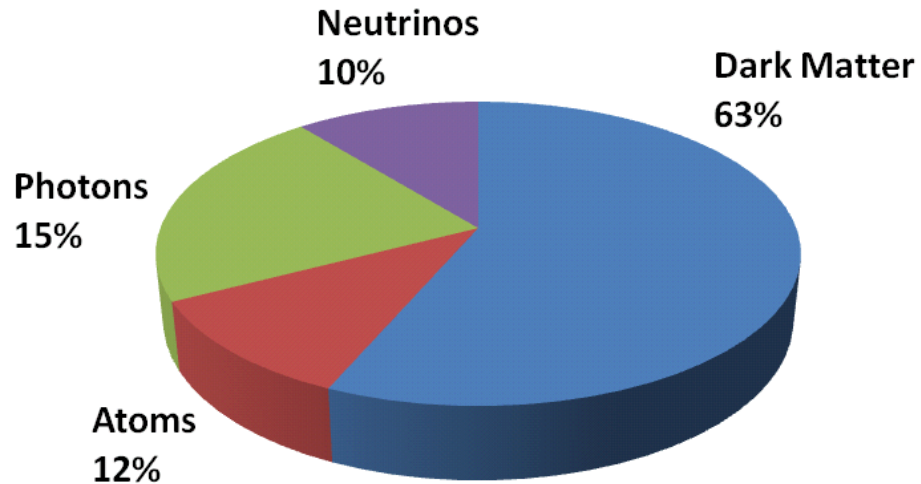


charge = 0  
spin =  $\frac{1}{2}$   
mass = 0  
speed = c

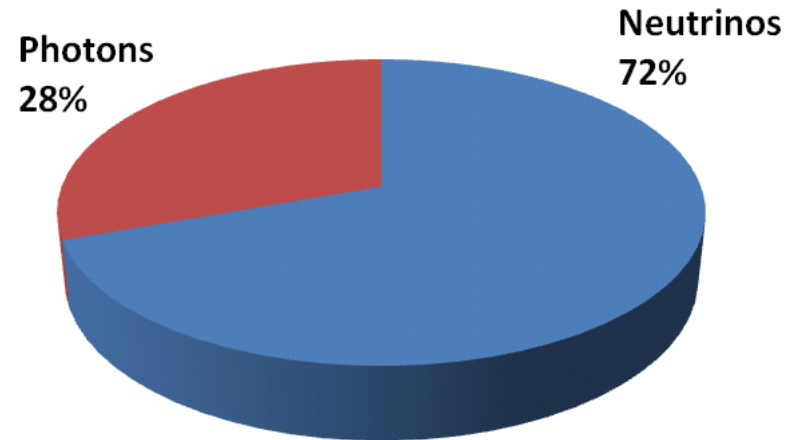




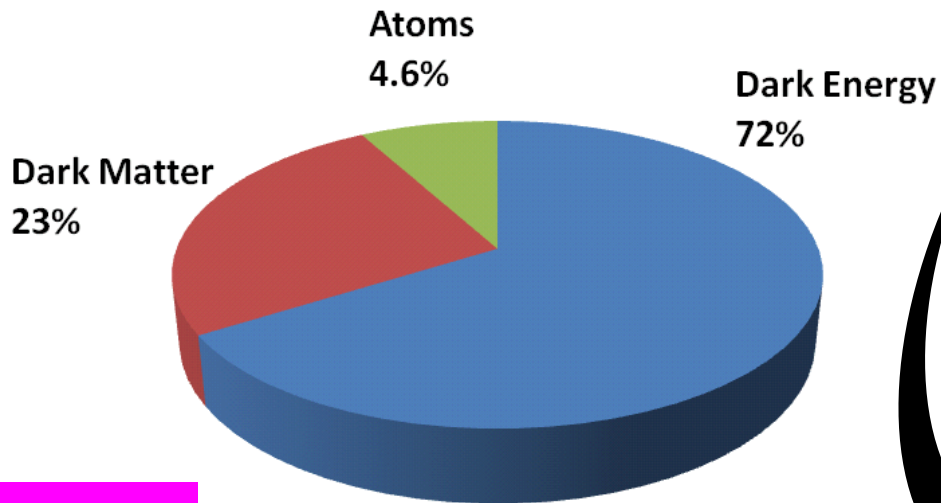
# NEUTRINOS



photon decoupling  
 $t = 380\,000$  years



neutrino decoupling  
 $t = 1$  second



Today  
 $t = 13.7$  billion years

A very important role  
in the evolution of  
the Universe



# Open Questions

the absolute  $\nu$  mass scale?  
the mass hierarchy problem?  
the flavor desert problem?  
how small is  $\theta_{13}$ ?  
leptonic CP violation?  
the Majorana nature?  
how many species? ...  
cosmic  $\nu$  background?  
supernova & stellar  $\nu$ 's?  
UHE cosmic  $\nu$ 's?  
warm dark matter?  
matter-antimatter asymmetry...

Daya Bay

# Exciting 2012

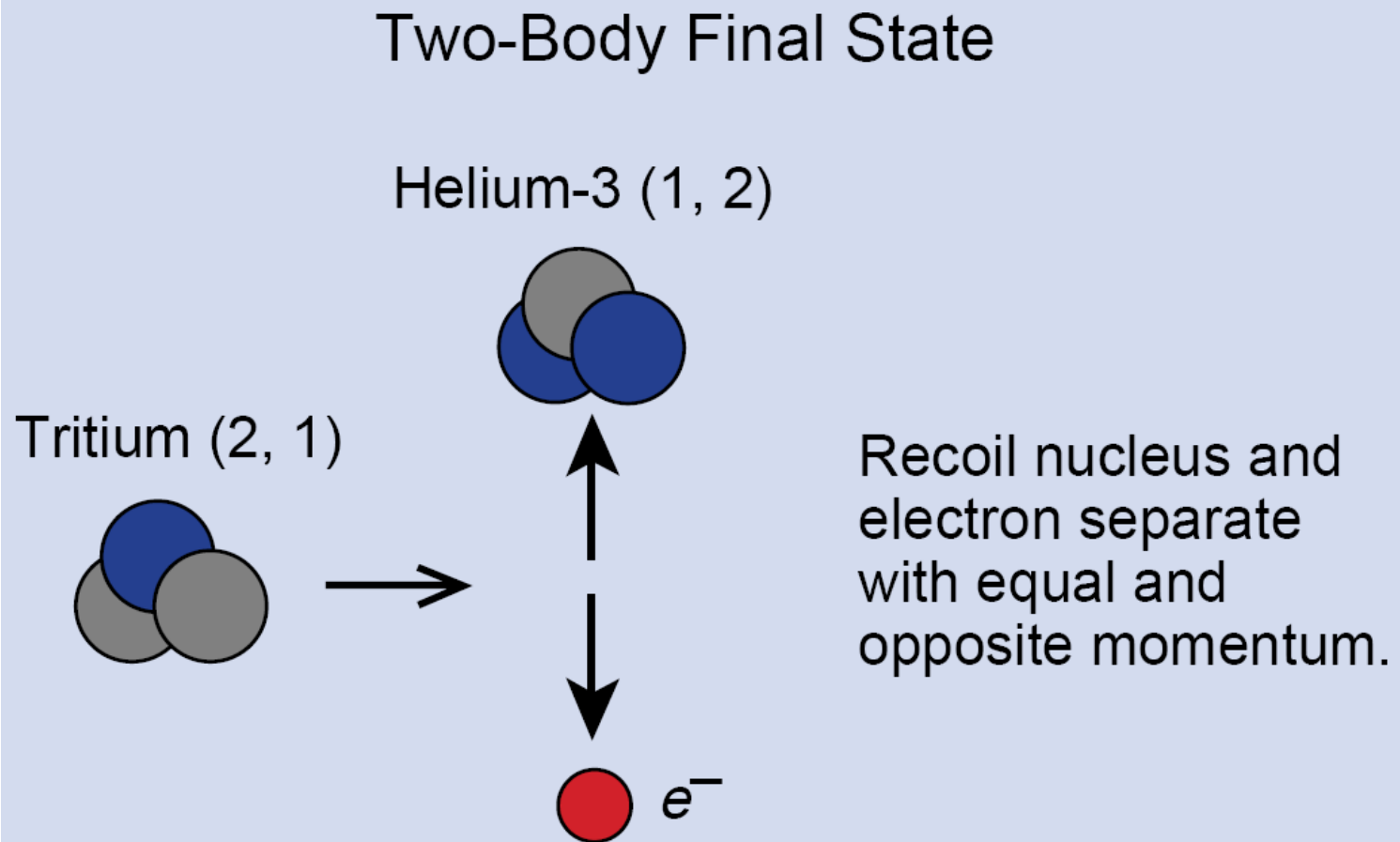


# Lecture A1

- ★ **Neutrino: Concept and Discovery**
- ★ **Lepton Flavors and Families**

# Beta Decay in 1930

6

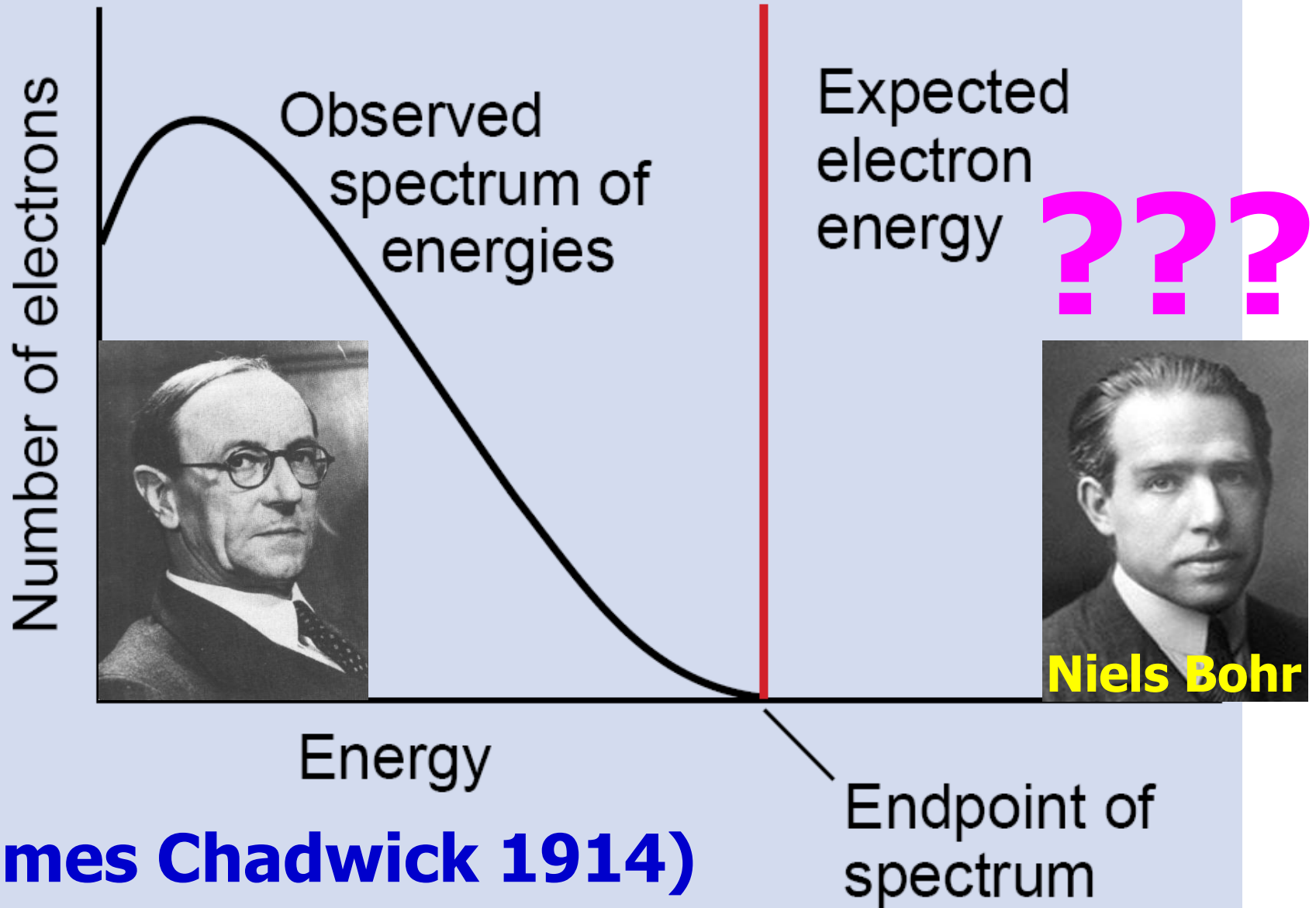


$$(N, Z) \rightarrow (N - 1, Z + 1) + e^{-},$$

where  $N$  = number of neutrons, and  
 $Z$  = number of protons.

# Energy Crisis?

7



# Desperate Remedy

8

**Wolfgang Pauli**  
**(1930)**



## The Desperate Remedy

4 December 1930  
Gloriastr.  
Zürich

Physical Institute of the  
Federal Institute of Technology (ETH)  
Zürich

Dear radioactive ladies and gentlemen,



# Pauli's Letter

9

As the bearer of these lines, to whom I ask you to listen graciously, will explain more exactly, considering the 'false' statistics of N-14 and Li-6 nuclei, as well as the continuous  $\beta$ -spectrum, I have hit upon a desperate remedy to save the "exchange theorem"\* of statistics and the energy theorem. Namely [there is] the possibility that there could exist in the nuclei electrically neutral particles that I wish to call neutrons,\*\* which have spin  $1/2$  and obey the exclusion principle, and additionally differ from light quanta in that they do not travel with the velocity of light: The mass of the neutron must be of the same order of magnitude as the electron mass and, in any case, not larger than 0.01 proton mass. The continuous  $\beta$ -spectrum would then become understandable by the assumption that in  $\beta$  decay a neutron is emitted together with the electron, in such a way that the sum of the energies of neutron and electron is constant.

But I don't feel secure enough to publish anything about this idea, so I first turn confidently to you, dear radioactives, with a question as to the situation concerning experimental proof of such a neutron, if it has something like about 10 times the penetrating capacity of a  $\gamma$  ray.

I admit that my remedy may appear to have a small *a priori* probability because neutrons, if they exist, would probably have long ago been seen. However, only those who wager can win, and the seriousness of the situation of the continuous  $\beta$ -spectrum can be made clear by the saying of my honored predecessor in office, Mr. Debye, who told me a short while ago in Brussels, "One does best not to think about that at all, like the new taxes." Thus one should earnestly discuss every way of salvation.—So, dear radioactives, put it to test and set it right.—Unfortunately, I cannot personally appear in Tübingen, since I am indispensable here on account of a ball taking place in Zürich in the night from 6 to 7 of December.—With many greetings to you, also to Mr. Back, your devoted servant,

W. Pauli

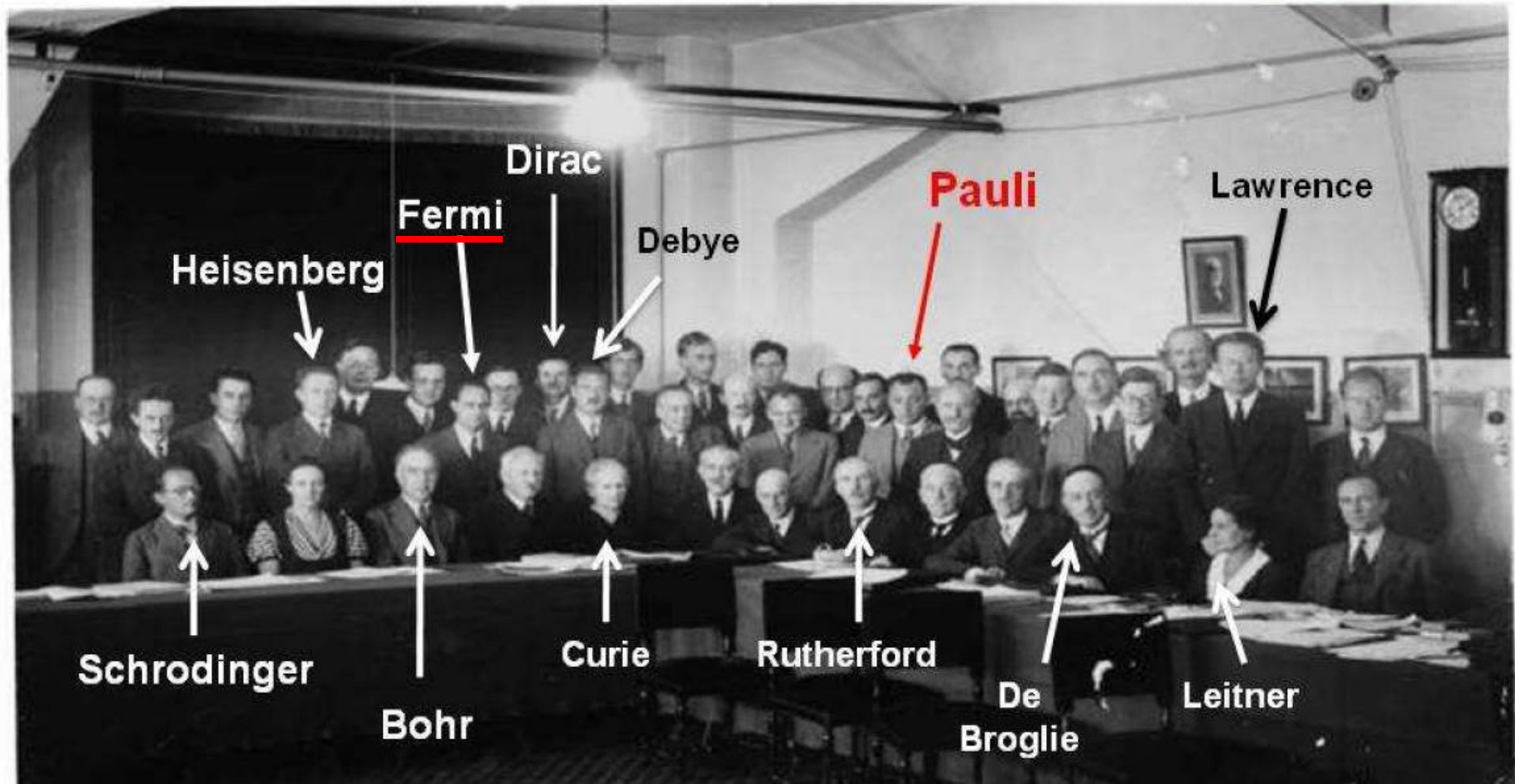
# Solvay 1933

11

Pauli gave a talk on his **neutrino** proposal in this congress.

INSTITUT INTERNATIONAL DE PHYSIQUE SOLVAY  
SEPTIÈME CONSEIL DE PHYSIQUE -- BRUXELLES, 22-29 OCTOBRE 1933

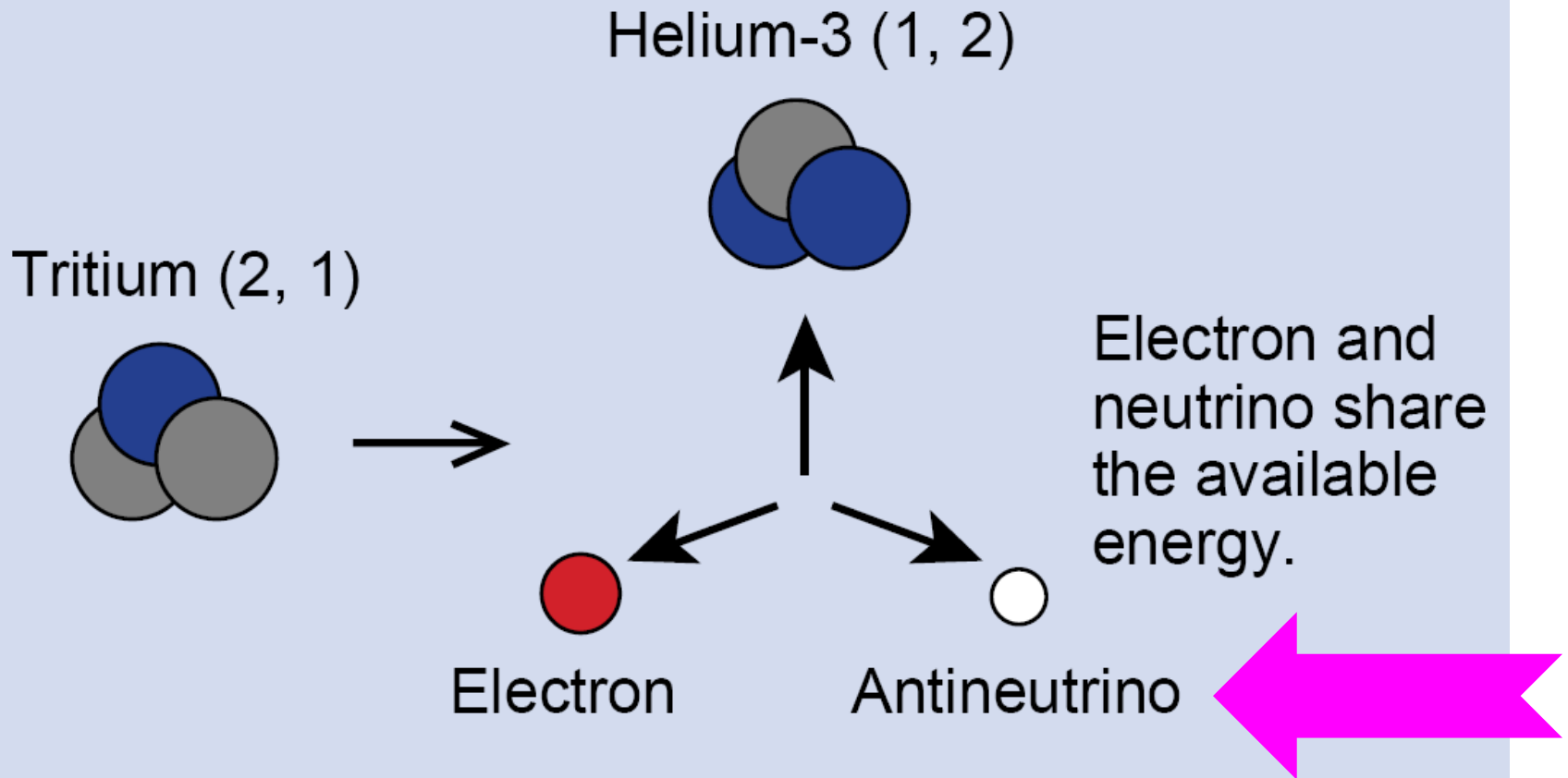
22 – 29 Octobre 1933



# True Picture of $\beta$ Decay

12

Three-Body Final State



$$(N, Z) \rightarrow (N - 1, Z + 1) + e^{-} + \bar{\nu} \text{ .}$$

# Fermi's Paper

13

## E. Fermi's publications on the Weak Interaction

**REJECTED**

E. Fermi, "Tentative Theory of Beta Rays"  
Letter Submitted to Nature (1933)

31 Dec, 1933



**Published first in this journal and later in Z. Phys. in 1934.**



# Fermi's Theory

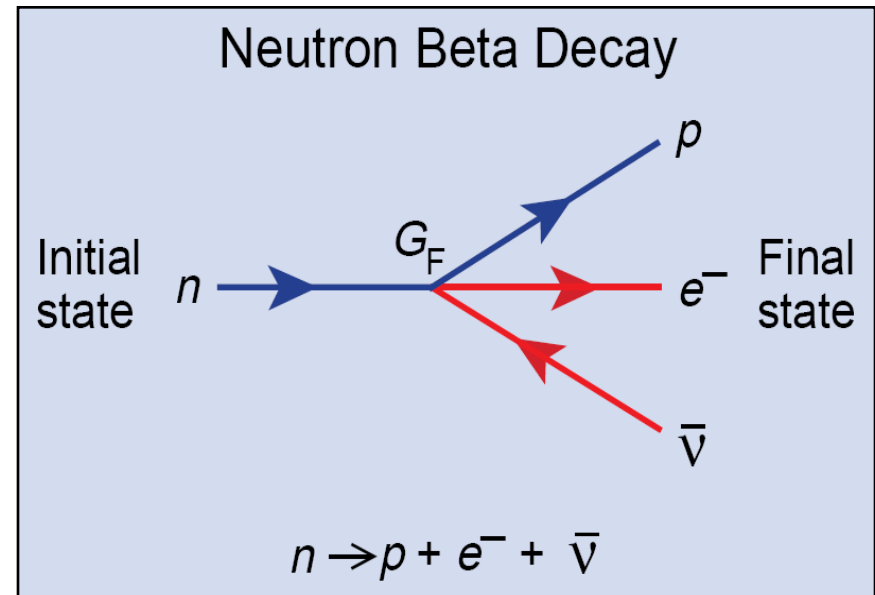
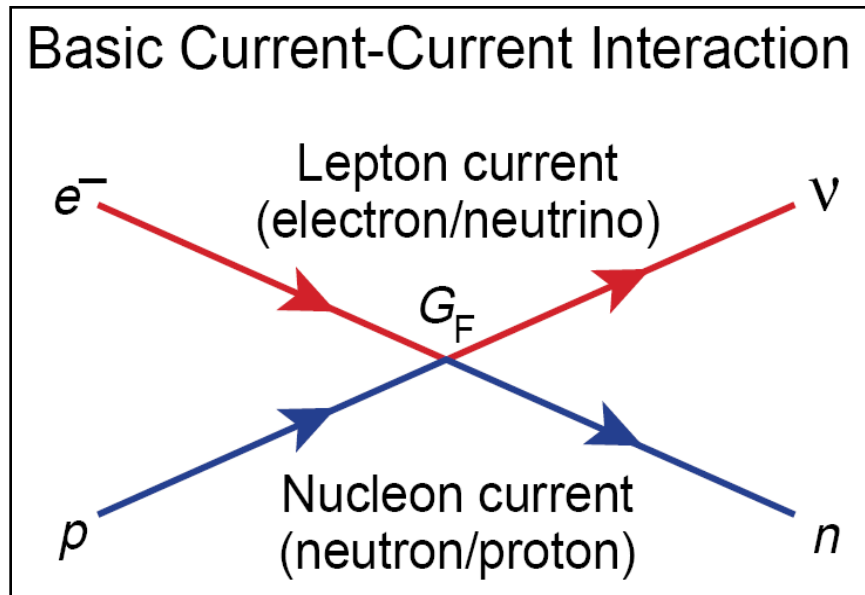
14

**Enrico Fermi (1933 & 1934) assumed a new force for  $\beta$  decay by combining 3 brand-new concepts:**

★ **Pauli's hypothesis: neutrinos**

★ **Dirac's thought: creation of particles**

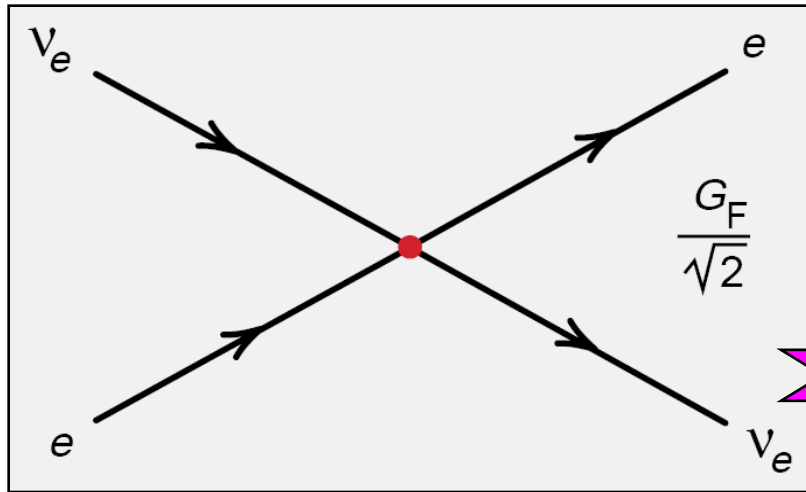
★ **Heisenberg's idea: neutron is related to proton**



# Weak Interactions

15

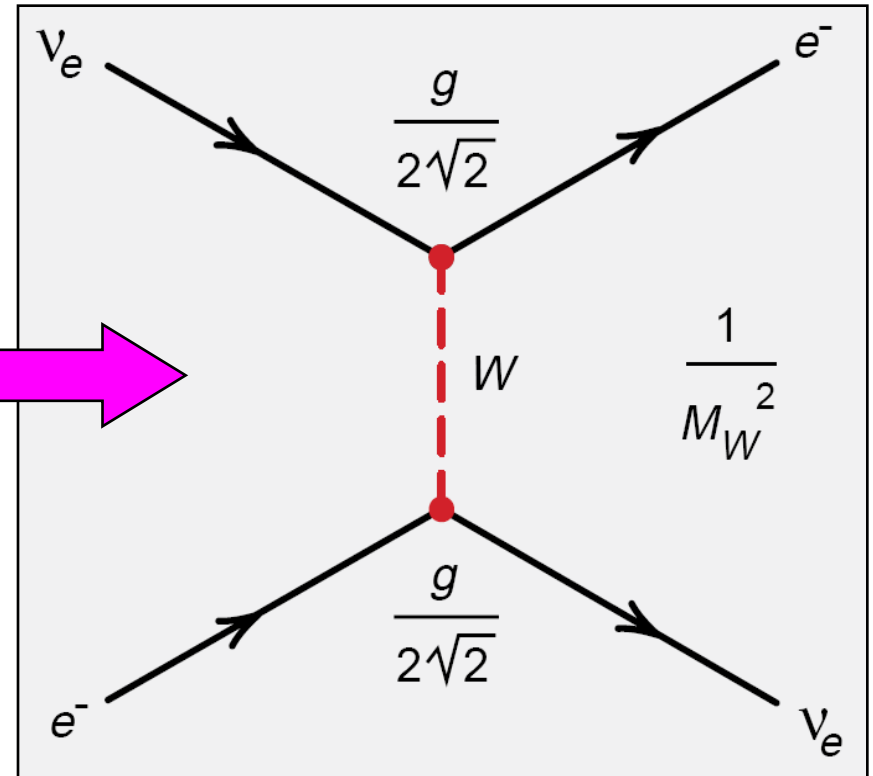
From Fermi's current-current interaction to weak charged-current gauge interactions



$$\frac{G_F}{\sqrt{2}} = \frac{g^2}{8M_W^2}$$



$$M_W = 80.4 \text{ GeV}$$

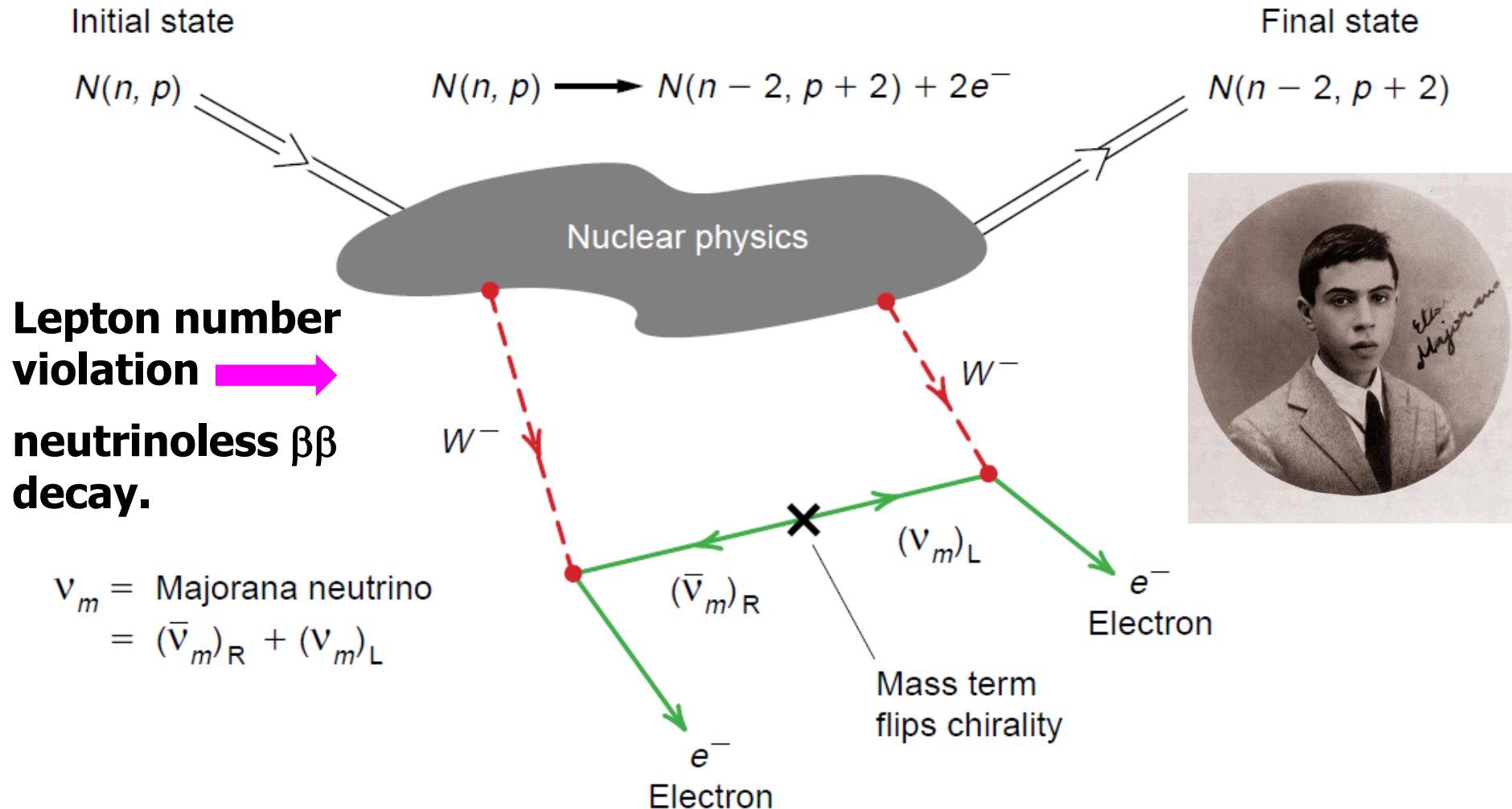


$$G_F = 1.66 \times 10^{-5} \text{ GeV}^{-2}$$

# Majorana Neutrinos

16

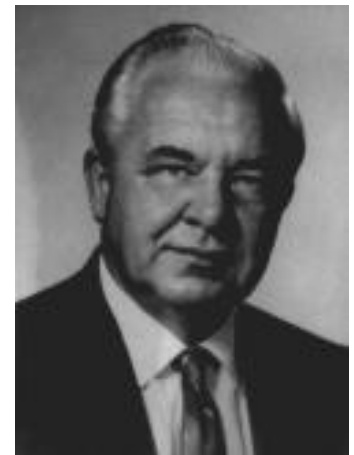
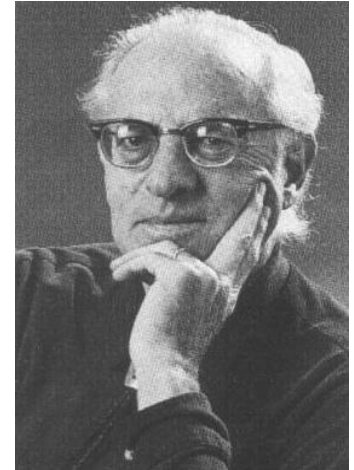
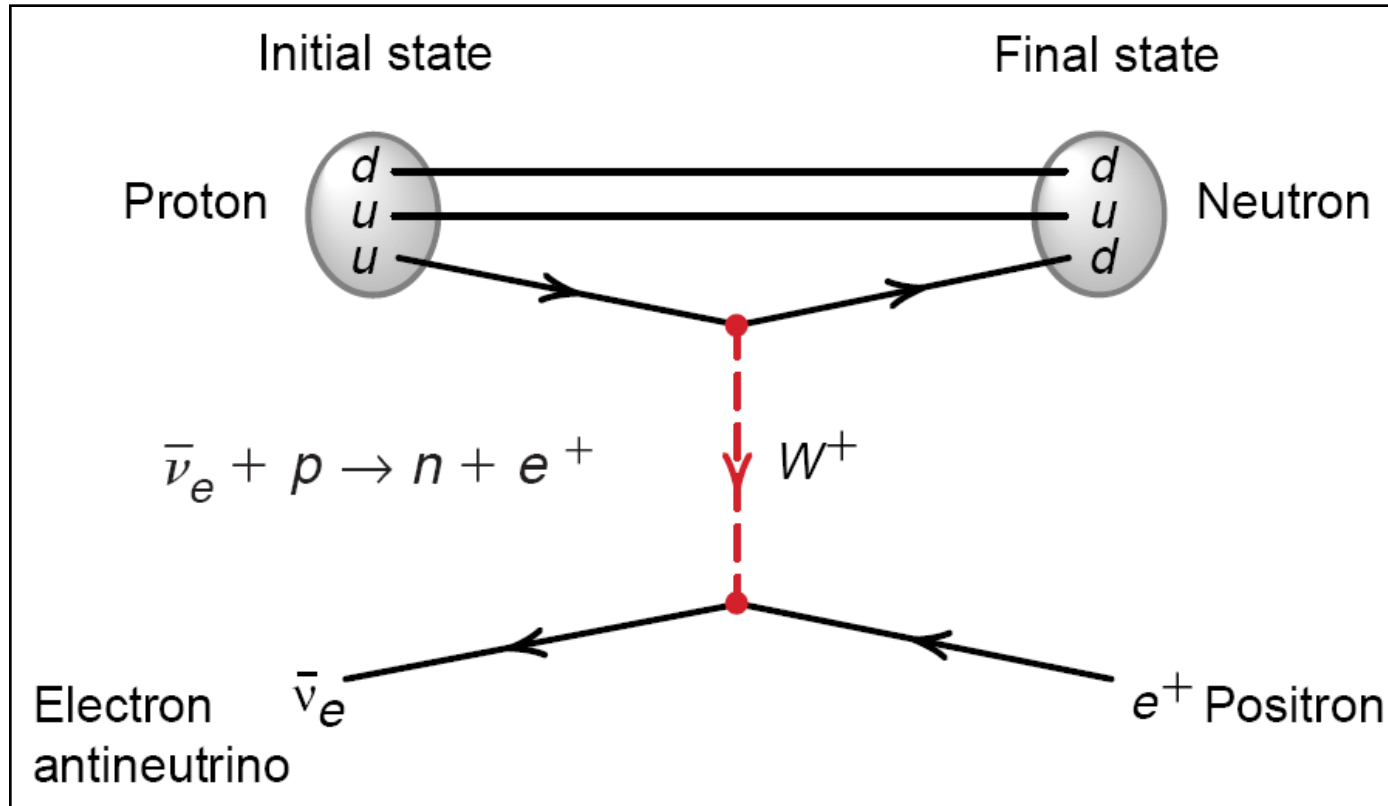
**Ettore Majorana, Fermi's PhD student, proposed a new idea in 1937: neutrino is its own antiparticle.**



# Impossible Challenge

17

An inverse  $\beta$  decay to detect neutrinos (Hans Bethe 1936)



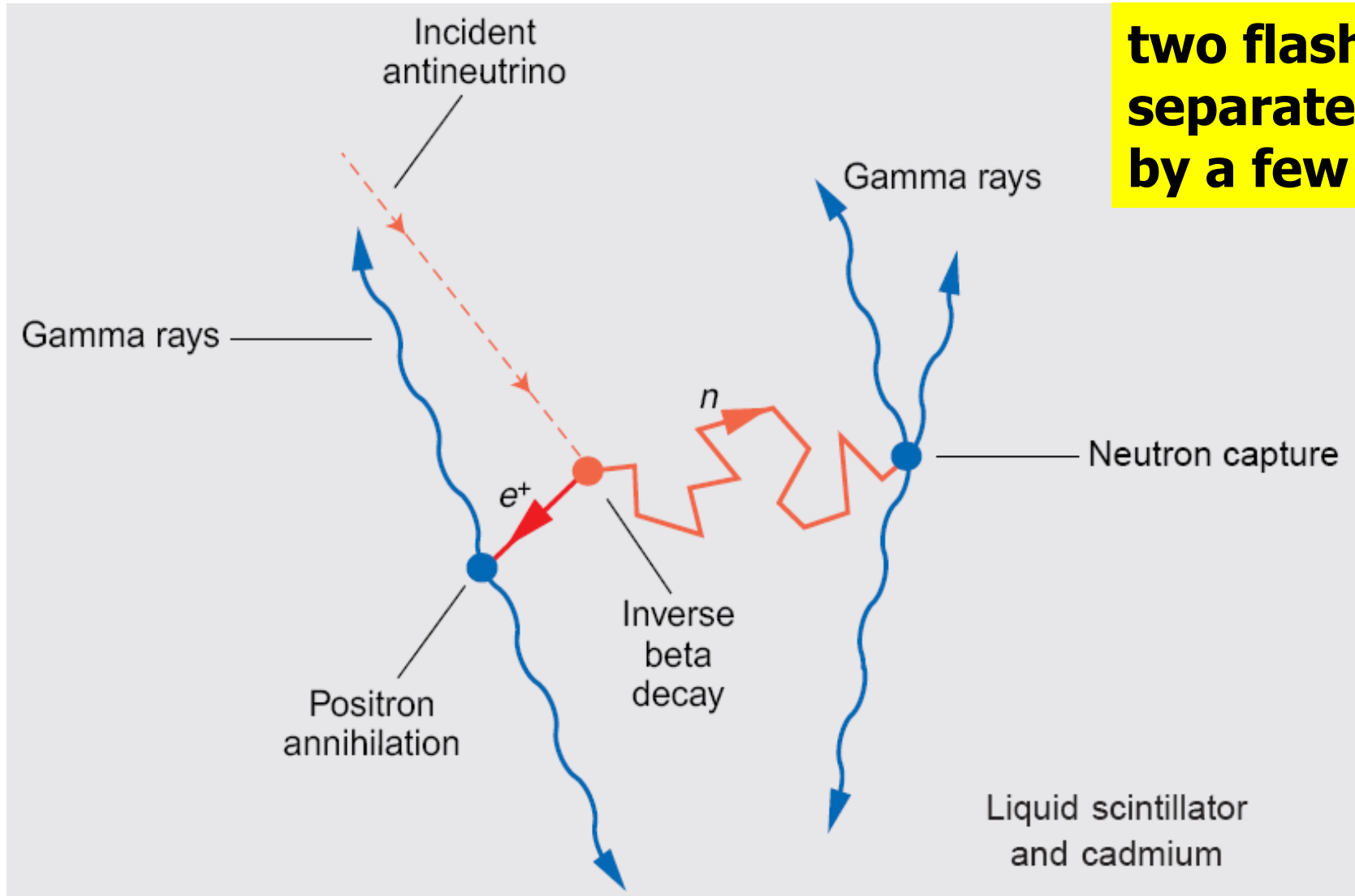
Very intense sources of neutrinos (1950's):  
**fission bombs** and **fission reactors**.

**Frederick Reines & Clyde Cowan's Project (1951).**

# Reactor Neutrinos

18

**Decision in 1952: neutrinos from a fission reactor.**



**two flashes  
separated  
by a few  $\mu s$**



# Positive Result?

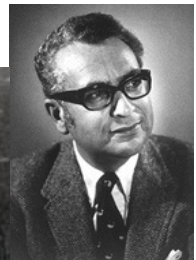
19

Reines and Cowan's telegram to Pauli on 14/06/1956:

**We're happy to inform you that we've definitely detected neutrinos from fission fragments by observing inverse  $\beta$  decay of protons. Observed cross section agrees well with expected  $6 \times 10^{-44} \text{ cm}^2$ .**

Such a theoretical value was based on a **parity-conserving** formulation of the  $\beta$  decay with **4** independent degrees of freedom for  **$\nu$ 's**.

This value **doubled** after the discovery of **parity violation** in 1957, leading to the two-component  **$\nu$**  theory in 1957 and the **V-A** theory in 1958.



# Nobel Prize

20

A new paper on this experiment published in Phys. Rev. in 1960 reported a cross section **twice** as large as that given in 1956.

Reines (1979): **our initial analysis grossly overestimated the detection efficiency with the result that the measured cross section was at first thought to be in good agreement with [the pre-parity violation] prediction.**



**Skepticism**  
**{over 39 years}**

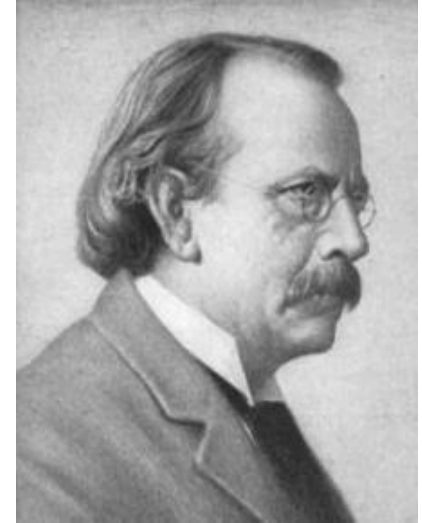
The **Nobel Prize** finally came to Frederick Reines in 1995!

# Electron & Its Neutrino

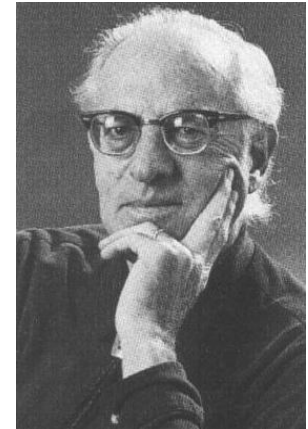
21

The **electron** was discovered in 1897 by **Joseph Thomson**.

The electron's anti-particle, **positron**, was predicted by **Paul Dirac** in 1928, and discovered by **Carl Anderson** in 1932.



In 1956 **Clyde Cowan** and **Frederick Reines** discovered the positron's partner, **electron antineutrino**.



# Muon

22

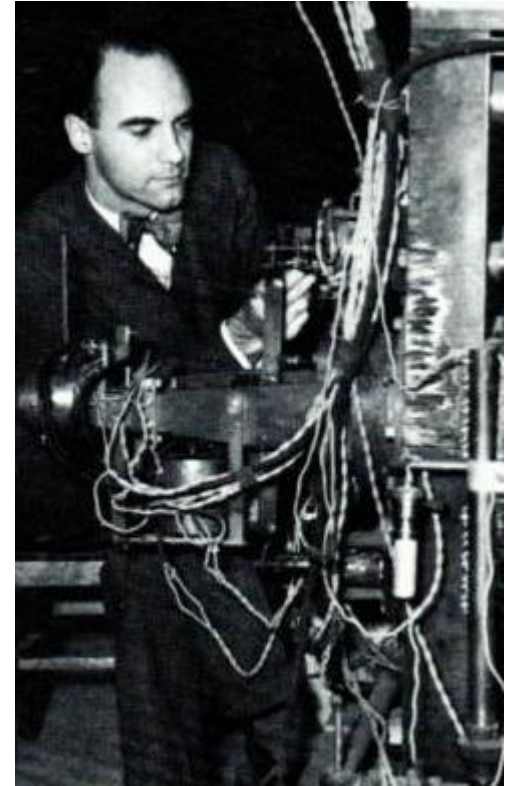
The **muon** particle, a sister of the electron, was discovered in 1936 by **Carl Anderson** and his first student **S. Neddermeyer**; and independently by **J. Street *et al.***

It was not **Hideki Yukawa's** "pion". And it was the first flavor puzzle.

**Isidor Rabi** famously asked:  
**Who** ordered that?



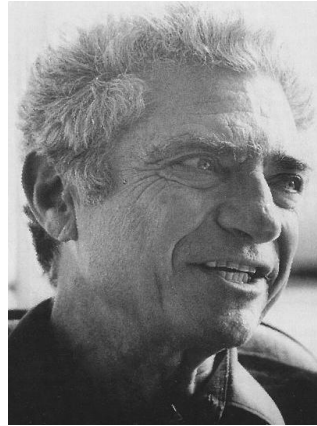
Isidor Isaac Rabi



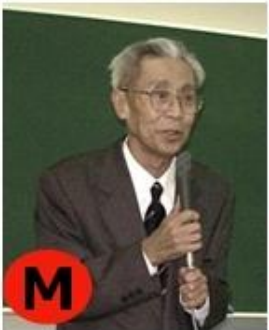
# Muon Neutrino

23

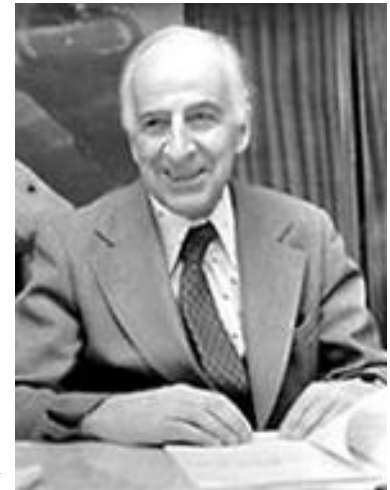
The **muon neutrino** was discovered by **Leon Lederman**, **Melvin Schwartz** and **Jack Steinberger** in 1962.



Neutrino flavor conversion was proposed by **Z. Maki**, **M. Nakagawa** and **S. Sakata** in 1962.



**Neutrinos** oscillate into **antineutrinos**: first proposed by **Bruno Pontecorvo** in 1957.



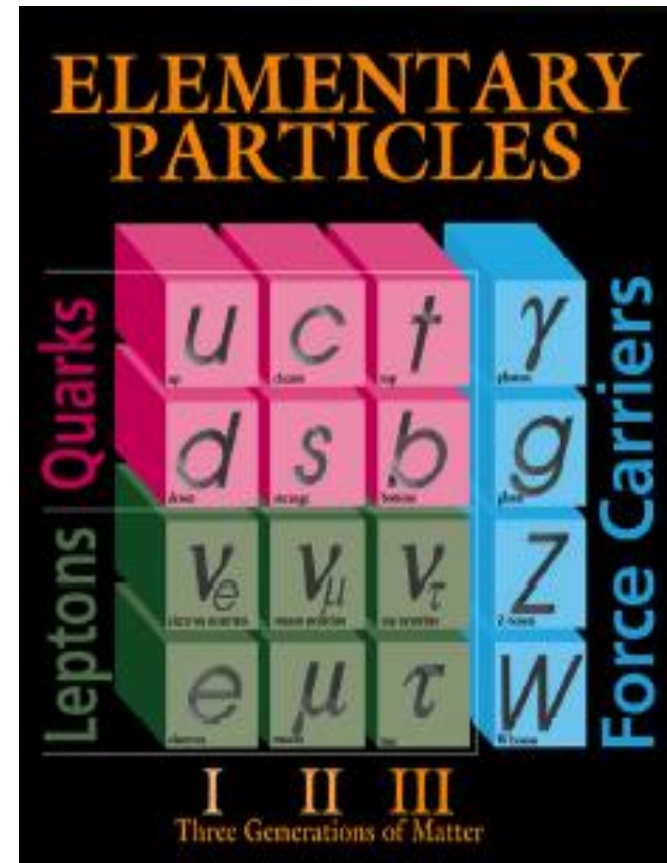
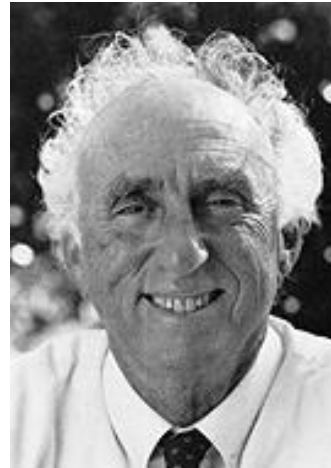
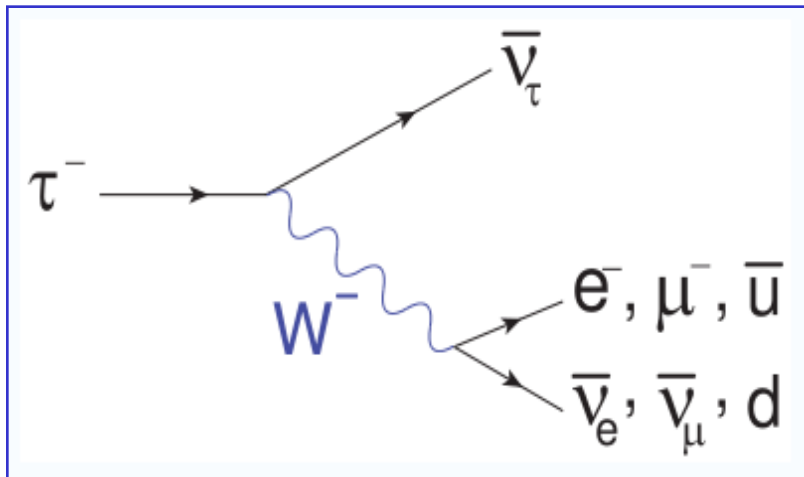


# Tau & Tau Neutrino

24

The **tau** particle was discovered by **Martin Perl** in 1975 via:

$$e^+ + e^- \rightarrow e^\pm + \mu^\mp + \text{undetected particles}$$



In 2000, the **tau neutrino** was finally discovered at the Fermilab.

The lepton family is complete!

# Leptons & Nobel Prizes

25

$e$	J.J. Thomson <b>1897</b> 😊	J.J. Thomson <b>1906</b> (NP)
$\nu_e$	C.L. Cowan et al. <b>1956</b>	F.J. Reines <b>1995</b> (NP)
$\mu$	J.C. Street et al. C.D. Anderson <b>1936</b> 😊	<b>1975 – 1936 = 1936 – 1897 = 39</b>
$\nu_\mu$	G. Danby et al. <b>1962</b>	M. Schwartz, L.M. Lederman, J. Steinberger <b>1988</b> (NP)
$\tau$	M.L. Perl et al. <b>1975</b> 😊	M.L. Perl <b>1995</b> (NP)
$\nu_\tau$	K. Kodama et al. <b>2000</b>	

**Antimatter: Positron.**

**Predicted by P.A.M. Dirac in 1928.**

**Discovered by C.D. Anderson in 1932; Nobel Prize in 1936.**

# A prediction?

26

In 1995 it was an Indian theorist who first discovered the **39-year gap** of charged leptons.

## 2) NOBEL LEPTONS.

By **K.V.L. Sarma** (**Tata Inst.**),. TIFR-TH-95-56, Dec 1995. 13pp.

Submitted to Curr. Sci.

e-Print Archive: **hep-ph/9512420**

$$1975 + 39 = 2014$$

A summary of the discoveries made in the world of leptons is given in Table 1. We see that the third generation has started getting Nobel prizes. It is amusing that the charged-leptons crop up with a **39-year gap** and may be the 4th one would show up in the year **2114**. For the present, the available experimental information implies that there are no charged leptons which are heavier than tau and lighter than 45 GeV.

**My contribution:** corrected **2114** to **2014**, so the discovery would be possible **100** years earlier (two years later)!

# Lecture A2

- ★ Dirac and Majorana Mass Terms
- ★ The Seesaw Mechanisms

# In the SM

28

All  $\nu$ 's are **massless** due to the model's simple structure:

----  $SU(2) \times U(1)$  **gauge symmetry and Lorentz invariance**

Fundamentals of a quantum field theory

---- Economical **particle content**:

No right-handed neutrino; only a single Higgs doublet

---- Mandatory **renormalizability**:

No dimension  $\geq 5$  operator ( **$B-L$**  conserved in the SM)

Neutrinos are **massless** in the SM: Natural or not?

**YES:** It's tooooooooo light and almost left-handed;

**NO:** No fundamental symmetry/conservation law.



# Some Notations

29

Define the **left-** and **right-**handed neutrino fields:

$$\nu_L = \begin{pmatrix} \nu_{eL} \\ \nu_{\mu L} \\ \nu_{\tau L} \end{pmatrix} \quad \blacksquare \quad N_R = \begin{pmatrix} N_{1R} \\ N_{2R} \\ N_{3R} \end{pmatrix}$$

**Extend the SM's  
particle content**

$$\begin{aligned} \psi_L &\equiv \frac{1 - \gamma_5}{2} \psi \\ \psi_R &\equiv \frac{1 + \gamma_5}{2} \psi \end{aligned}$$

Their **charge-conjugate counterparts** are defined below and transform as **right-** and **left-**handed fields, respectively:

$$(\nu_L)^c \equiv \mathcal{C} \overline{\nu_L}^T, \quad (N_R)^c \equiv \mathcal{C} \overline{N_R}^T$$

$$\overline{(\nu_L)^c} = (\nu_L)^T \mathcal{C}, \quad \overline{(N_R)^c} = (N_R)^T \mathcal{C}$$

$$(\nu_L)^c = (\nu^c)_R \text{ and } (N_R)^c = (N^c)_L \text{ hold} \quad \text{(can be proved easily)}$$

**Properties of the charge-conjugation matrix:**

$$\mathcal{C} \gamma_\mu^T \mathcal{C}^{-1} = -\gamma_\mu, \quad \mathcal{C} \gamma_5^T \mathcal{C}^{-1} = \gamma_5, \quad \mathcal{C}^{-1} = \mathcal{C}^\dagger = \mathcal{C}^T = -\mathcal{C}$$

They are from the requirement that **the charge-conjugated field** must satisfy the same **Dirac** equation ( $\mathcal{C} = i\gamma^2\gamma^0$  in the **Dirac** representation)

# Dirac Mass Term

30

A **Dirac** neutrino is described by a **4**-component spinor:  $\nu = \nu_L + N_R$

**Step 1:** the gauge-invariant **Dirac** mass term and **SSB**:

$$-\mathcal{L}_{\text{Dirac}} = \bar{\ell}_L Y_\nu \tilde{H} N_R + \text{h.c.}$$



$$-\mathcal{L}'_{\text{Dirac}} = \bar{\nu}_L M_D N_R + \text{h.c.}$$

$$M_D = Y_\nu \langle H \rangle \text{ with } \langle H \rangle \simeq 174 \text{ GeV}$$

**Step 2:** basis transformation:

$$V^\dagger M_D U = \widehat{M}_\nu \equiv \text{Diag}\{m_1, m_2, m_3\}$$

$$-\mathcal{L}'_{\text{Dirac}} = \bar{\nu}'_L \widehat{M}_\nu N'_R + \text{h.c.}$$

**Mass states link to flavor states:**

$$\nu'_L = V^\dagger \nu_L \text{ and } N'_R = U^\dagger N_R$$

$$\nu' = \nu'_L + N'_R = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

**Step 3:** physical mass term and kinetic term:

$$-\mathcal{L}'_{\text{Dirac}} = \bar{\nu}' \widehat{M}_\nu \nu' = \sum_{i=1}^3 m_i \bar{\nu}_i \nu_i$$

$$\mathcal{L}_{\text{kinetic}} = i\bar{\nu}_L \gamma_\mu \partial^\mu \nu_L + i\bar{N}_R \gamma_\mu \partial^\mu N_R = i\bar{\nu}' \gamma_\mu \partial^\mu \nu' = i \sum_{k=1}^3 \bar{\nu}_k \gamma_\mu \partial^\mu \nu_k$$

# Dirac Neutrino Mixing

31

**Weak charged-current interactions of leptons:**

$$\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} (\overline{e \ \mu \ \tau})_L \gamma^\mu \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}_L W_\mu^- + \text{h.c.}$$

**In the flavor basis**

$$\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} (\overline{e \ \mu \ \tau})_L \gamma^\mu V \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_L W_\mu^- + \text{h.c.}$$

**In the mass basis**

**Without loss of generality, one may choose **mass states=flavor states** for **charged leptons**. So **V** is just the **MNSP** matrix of neutrino mixing.**

**Both the mass & CC terms are invariant with respect to a **global** phase transformation: **lepton number (flavor) conservation (violation)**.**

$$l(x) \rightarrow e^{i\Phi} l(x)$$

$$\nu'_L(x) \rightarrow e^{i\Phi} \nu'_L(x)$$

$$N'_R(x) \rightarrow e^{i\Phi} N'_R(x)$$

	$e^-$	$\nu_e$	$e^+$	$\bar{\nu}_e$	$\mu^-$	$\nu_\mu$	$\mu^+$	$\bar{\nu}_\mu$	$\tau^-$	$\nu_\tau$	$\tau^+$	$\bar{\nu}_\tau$
$L$	+1	+1	-1	-1	+1	+1	-1	-1	+1	+1	-1	-1
$L_e$	+1	+1	-1	-1	0	0	0	0	0	0	0	0
$L_\mu$	0	0	0	0	+1	+1	-1	-1	0	0	0	0
$L_\tau$	0	0	0	0	0	0	0	0	+1	+1	-1	-1

# Majorana Mass Term (1)

32

A **Majorana** mass term can be obtained by introducing a **Higgs triplet** into the SM, writing out the gauge-invariant Yukawa interactions and Higgs potentials, and then integrating out heavy degrees of freedom (**type-II seesaw mechanism**):

$$-\mathcal{L}'_{\text{Majorana}} = \frac{1}{2} \overline{\nu}_L M_L (\nu_L)^c + \text{h.c.}$$

The **Majorana** mass matrix must be a **symmetric** matrix. It can be diagonalized by a **unitary** matrix

$$\overline{\nu}_L M_L (\nu_L)^c = [\overline{\nu}_L M_L (\nu_L)^c]^T = -\overline{\nu}_L C^T M_L^T \overline{\nu}_L^T = \overline{\nu}_L M_L^T (\nu_L)^c$$

**Diagonalization:**

$$-\mathcal{L}'_{\text{Majorana}} = \frac{1}{2} \overline{\nu}'_L \widehat{M}_\nu (\nu'_L)^c + \text{h.c.}$$

$$V^\dagger M_L V^* = \widehat{M}_\nu \equiv \text{Diag}\{m_1, m_2, m_3\}$$

$$\nu'_L = V^\dagger \nu_L \text{ and } (\nu'_L)^c = C \overline{\nu}'_L^T$$

**Physical mass term:**

$$-\mathcal{L}'_{\text{Majorana}} = \frac{1}{2} \overline{\nu}' \widehat{M}_\nu \nu' = \frac{1}{2} \sum_{i=1}^3 m_i \overline{\nu}_i \nu_i$$

$$\nu' = \nu'_L + (\nu'_L)^c = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\text{Majorana condition } (\nu')^c = \nu'$$

# Majorana Mass Term (2)

33

**Kinetic term** (you may prove  $\overline{(\psi_L)^c} \gamma_\mu \partial^\mu (\psi_L)^c = \overline{\psi_L} \gamma_\mu \partial^\mu \psi_L$ )

$$\mathcal{L}_{\text{kinetic}} = i \overline{\nu_L} \gamma_\mu \partial^\mu \nu_L = i \overline{\nu'_L} \gamma_\mu \partial^\mu \nu'_L = \frac{i}{2} \overline{\nu'} \gamma_\mu \partial^\mu \nu' = \frac{i}{2} \sum_{k=1}^3 \overline{\nu_k} \gamma_\mu \partial^\mu \nu_k$$

**Question:** why is there a factor **1/2** in the **Majorana** mass term?

**Answer:** it allows us to get the correct **Dirac** equation of motion.

**A proof:** write out the Lagrangian of free massive **Majorana** neutrinos:

$$\begin{aligned} \mathcal{L}_\nu &= i \overline{\nu_L} \gamma_\mu \partial^\mu \nu_L - \left[ \frac{1}{2} \overline{\nu_L} M_L (\nu_L)^c + \text{h.c.} \right] \\ &= i \overline{\nu'_L} \gamma_\mu \partial^\mu \nu'_L - \left[ \frac{1}{2} \overline{\nu'_L} \widehat{M}_\nu (\nu'_L)^c + \text{h.c.} \right] \\ &= \frac{1}{2} \left( i \overline{\nu'} \gamma_\mu \partial^\mu \nu' - \overline{\nu'} \widehat{M}_\nu \nu' \right) = -\frac{1}{2} \left( i \partial^\mu \overline{\nu'} \gamma_\mu \nu' + \overline{\nu'} \widehat{M}_\nu \nu' \right) \end{aligned}$$



**Euler-Lagrange equation:**

$$\partial^\mu \frac{\partial \mathcal{L}_\nu}{\partial (\partial^\mu \overline{\nu'})} - \frac{\partial \mathcal{L}_\nu}{\partial \overline{\nu'}} = 0$$



$$i \gamma_\mu \partial^\mu \nu' - \widehat{M}_\nu \nu' = 0$$

$$i \gamma_\mu \partial^\mu \nu_k - m_k \nu_k = 0$$



# Majorana Neutrino Mixing 34

Weak charged-current interactions of leptons:

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} (\overline{e \ \mu \ \tau})_{\text{L}} \gamma^{\mu} \begin{pmatrix} \nu_e \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix}_{\text{L}} W_{\mu}^{-} + \text{h.c.}$$

In the flavor basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} (\overline{e \ \mu \ \tau})_{\text{L}} \gamma^{\mu} V \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_{\text{L}} W_{\mu}^{-} + \text{h.c.}$$

In the mass basis

The **MNSP** matrix **V** contains 2 extra CP-violating phases.

Mass and CC terms are not simultaneously invariant under a global phase transformation --- **Lepton number violation**

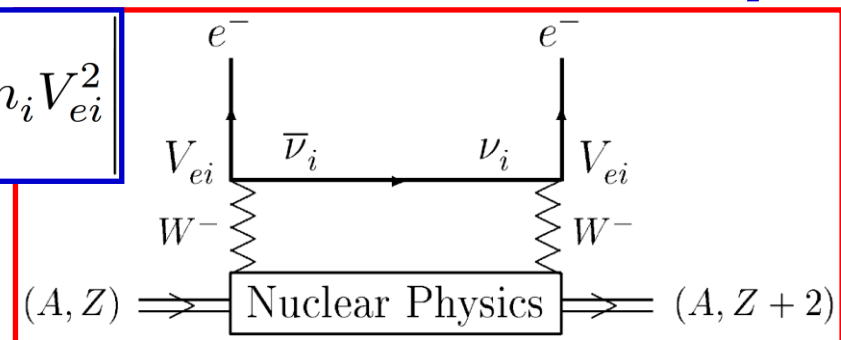
**Neutrinoless double-beta decay**

$$l(x) \rightarrow e^{i\Phi} l(x)$$

$$\nu'_{\text{L}}(x) \rightarrow e^{i\Phi} \nu'_{\text{L}}(x)$$

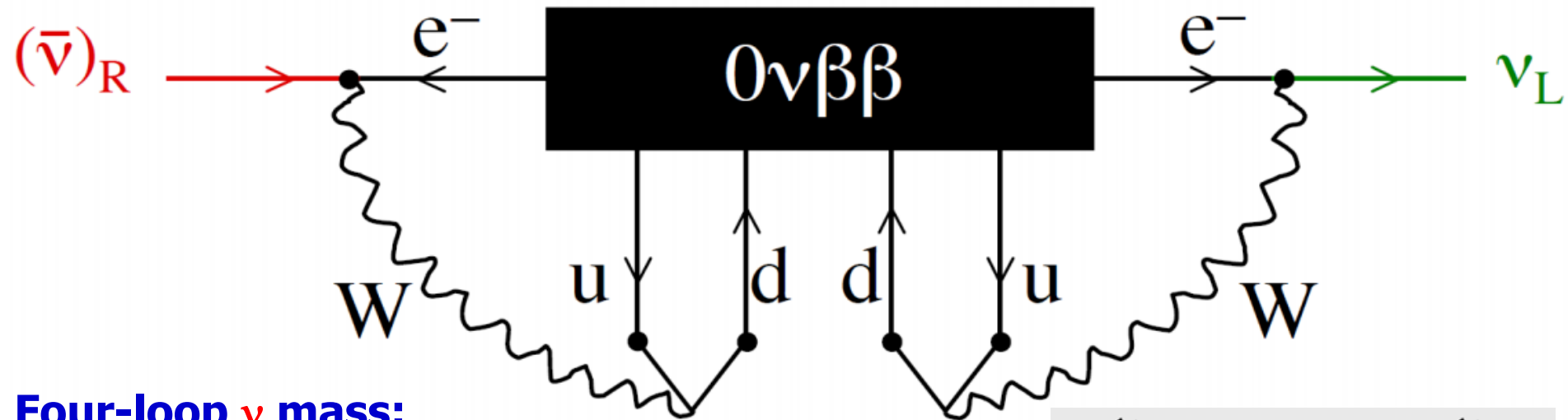
$$\overline{\nu'_{\text{L}}} \rightarrow e^{-i\Phi} \overline{\nu'_{\text{L}}} \text{ and } (\nu'_{\text{L}})^c \rightarrow e^{-i\Phi} (\nu'_{\text{L}})^c$$

$$\langle m \rangle_{ee} \equiv \left| \sum_i m_i V_{ei}^2 \right|$$



# Schechter-Valle Theorem 35

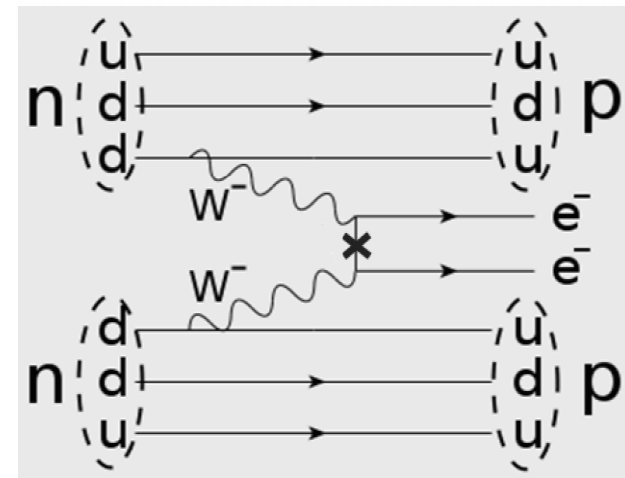
**THEOREM (1982):** if a  $0\nu\beta\beta$  decay happens, there must be an effective **Majorana** mass term.



**Four-loop  $\nu$  mass:**

$$\delta m_\nu = \mathcal{O}(10^{-24} \text{ eV}) \quad (\text{Duerr, Lindner, Merle, 2011})$$

**Note:** The **black box** can in principle have many different processes (new physics). Only in the simplest case, it is likely to constrain neutrino masses.



# Current Bounds

36

$$\langle m \rangle_{ee} \equiv \left| \sum_i m_i V_{ei}^2 \right|$$

**A recent review: W. Rodejohann, IJMPE 20 (2011) 1833**

Isotope	$T_{1/2}^{0\nu}$ [yrs]	Experiment	$ m_{ee} _{\min}^{\lim}$ [eV]	$ m_{ee} _{\max}^{\lim}$ [eV]	
$^{48}\text{Ca}$	$5.8 \times 10^{22}$	CANDLES	3.55	9.91	$\times 1.02$
$^{76}\text{Ge}$	$1.9 \times 10^{25}$	HDM	0.21	0.53	$\times 1.04$
	$1.6 \times 10^{25}$	IGEX	0.25	0.63	$\times 1.04$
$^{82}\text{Se}$	$3.2 \times 10^{23}$	NEMO-3	0.85	2.08	$\times 1.04$
$^{96}\text{Zr}$	$9.2 \times 10^{21}$	NEMO-3	3.97	14.39	$\times 1.06$
$^{100}\text{Mo}$	$1.0 \times 10^{24}$	NEMO-3	0.31	0.79	$\times 1.06$
$^{116}\text{Cd}$	$1.7 \times 10^{23}$	SOLOTVINO	1.22	2.30	$\times 1.06$
$^{130}\text{Te}$	$2.8 \times 10^{24}$	CUORICINO	0.27	0.57	$\times 1.09$
$^{136}\text{Xe}$	$1.6 \times 10^{25}$	EXO-200	0.15	0.36	$\times 1.10$
$^{150}\text{Nd}$	$1.8 \times 10^{22}$	NEMO-3	2.35	5.08	$\times 1.12$

# Hybrid Mass Term (1)

37

A **hybrid** mass term can be written out in terms of the left- and right-handed neutrino fields and their charge-conjugate counterparts:

$$\begin{aligned}
 -\mathcal{L}'_{\text{hybrid}} &= \overline{\nu_L} M_D N_R + \frac{1}{2} \overline{\nu_L} M_L (\nu_L)^c + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \text{h.c.} \\
 &= \frac{1}{2} \begin{bmatrix} \overline{\nu_L} & \overline{(N_R)^c} \end{bmatrix} \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} + \text{h.c.} ,
 \end{aligned}$$

**type-(I+II) seesaw**

**Here we have used**

**Diagonalization** by means of a **6×6 unitary** matrix:

$$\overline{(N_R)^c} M_D^T (\nu_L)^c = [(N_R)^T \mathcal{C} M_D^T \mathcal{C} \overline{\nu_L}^T]^T = \overline{\nu_L} M_D N_R$$

$$\begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* = \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix}$$

$$\widehat{M}_\nu \equiv \text{Diag}\{m_1, m_2, m_3\}, \quad \widehat{M}_N \equiv \text{Diag}\{M_1, M_2, M_3\}$$

**It is actually a Majorana mass term!**

$$-\mathcal{L}'_{\text{hybrid}} = \frac{1}{2} \begin{bmatrix} \overline{\nu'_L} & \overline{(N'_R)^c} \end{bmatrix} \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix} \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} + \text{h.c.}$$

**Majorana mass states**

$$\nu' = \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} + \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ N_1 \\ N_2 \\ N_3 \end{pmatrix}$$

$$(\nu')^c = \nu'$$

$$\nu'_L = V^\dagger \nu_L + S^\dagger (N_R)^c$$

$$N'_R = R^T (\nu_L)^c + U^T N_R$$

# Hybrid Mass Term (2)

38

**Physical mass term:**

$$-\mathcal{L}'_{\text{hybrid}} = \frac{1}{2} \overline{\nu'} \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix} \nu' = \frac{1}{2} \sum_{i=1}^3 (m_i \overline{\nu}_i \nu_i + M_i \overline{N}_i N_i)$$

**Kinetic term:**

$$\begin{aligned} \mathcal{L}_{\text{kinetic}} &= i \overline{\nu}_L \gamma_\mu \partial^\mu \nu_L + i \overline{N}_R \gamma_\mu \partial^\mu N_R \\ &= \frac{i}{2} \begin{bmatrix} \overline{\nu}_L & \overline{(N_R)^c} \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} \nu_L \\ (N_R)^c \end{bmatrix} + \frac{i}{2} \begin{bmatrix} (\nu_L)^c & \overline{N_R} \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} \\ &= \frac{i}{2} \begin{bmatrix} \nu'_L & \overline{(N'_R)^c} \end{bmatrix} \gamma_\mu \partial^\mu \begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} V & R \\ S & U \end{pmatrix} \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} \\ &\quad + \frac{i}{2} \begin{bmatrix} (\nu'_L)^c & \overline{N'_R} \end{bmatrix} \gamma_\mu \partial^\mu \begin{pmatrix} V & R \\ S & U \end{pmatrix}^T \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} \\ &= \frac{i}{2} \begin{bmatrix} \nu'_L & \overline{(N'_R)^c} \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} + \frac{i}{2} \begin{bmatrix} (\nu'_L)^c & \overline{N'_R} \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} \\ &= i \overline{\nu'_L} \gamma_\mu \partial^\mu \nu'_L + i \overline{N'_R} \gamma_\mu \partial^\mu N'_R \\ &= \frac{i}{2} \overline{\nu'} \gamma_\mu \partial^\mu \nu' = \frac{i}{2} \sum_{k=1}^3 (\overline{\nu}_k \gamma_\mu \partial^\mu \nu_k + \overline{N}_k \gamma_\mu \partial^\mu N_k) \end{aligned}$$

# Non-unitary Flavor Mixing 39

Weak charged-current interactions of leptons:

In the flavor basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)}_L \gamma^\mu \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}_L W_\mu^- + \text{h.c.}$$

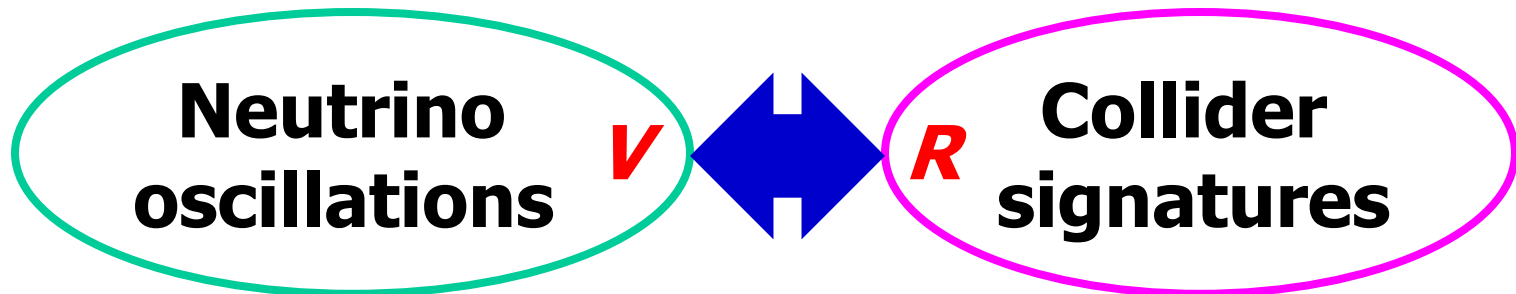
$$\nu_L = V \nu'_L + R(N'_R)^c$$

In the mass basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)}_L \gamma^\mu \left[ V \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_L + R \begin{pmatrix} N_1 \\ N_2 \\ N_3 \end{pmatrix}_L \right] W_\mu^- + \text{h.c.}$$

$V$  = non-unitary light neutrino mixing (**MNSP**) matrix  $VV^\dagger + RR^\dagger = 1$

$R$  = light-heavy neutrino mixing (**CC** interactions of heavy neutrinos)



**TeV seesaws** may bridge the gap between neutrino & collider physics.



# Seesaw Mechanisms (1)

40

A **hybrid** neutrino mass Lagrangian may contain three distinct terms:

$$\begin{aligned} -\mathcal{L}'_{\text{hybrid}} &= \overline{\nu}_L M_D N_R + \frac{1}{2} \overline{\nu}_L M_L (\nu_L)^c + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \text{h.c.} \\ &= \frac{1}{2} \begin{bmatrix} \overline{\nu}_L & \overline{(N_R)^c} \end{bmatrix} \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} + \text{h.c.} , \end{aligned}$$

- ♣ **Normal Dirac mass term**, proportional to the scale of electroweak symmetry breaking ( $\sim$  **174 GeV**);
- ♣ **Light Majorana mass term**, violating the SM gauge symmetry and having a scale much lower than **174 GeV**;
- ♣ **Heavy Majorana mass term**, originating from the SU(2)<sub>L</sub> singlet and having a scale much higher than **174 GeV**.

A strong hierarchy of **3** mass scales allows us to make approximation

$$\begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* = \begin{pmatrix} \widehat{M}_\nu & 0 \\ 0 & \widehat{M}_N \end{pmatrix}$$

# Seesaw Mechanisms (2)

41

The above **unitary** transformation leads to the following relationships:

$$\begin{aligned} R\widehat{M}_N &= M_L R^* + M_D U^* \\ S\widehat{M}_\nu &= M_D^T V^* + M_R S^* \end{aligned}$$

$$\begin{aligned} M_R &\gg M_D \gg M_L \\ R &\sim S \sim \mathcal{O}(M_D/M_R) \end{aligned}$$

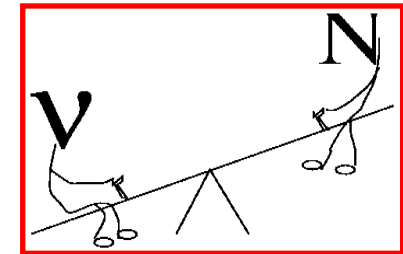
$$\begin{aligned} U\widehat{M}_N &= M_R U^* + M_D^T R^* \\ V\widehat{M}_\nu &= M_L V^* + M_D S^* \end{aligned}$$

$$\begin{aligned} U\widehat{M}_N U^T &= M_R (U U^\dagger)^T + M_D^T (R^* U^T) \approx M_R, \\ V\widehat{M}_\nu V^T &= M_L (V V^\dagger)^T + M_D (S^* V^T) \approx M_L + M_D (S^* V^T) \end{aligned}$$

$$S^* V^T = M_R^{-1} S\widehat{M}_\nu V^T - M_R^{-1} M_D^T (V V^\dagger)^T \approx -M_R^{-1} M_D^T$$

Then we arrive at the **type-(I+II)** seesaw formula:

$$M_\nu \equiv V\widehat{M}_\nu V^T \approx M_L - M_D M_R^{-1} M_D^T$$



**Type-I** seesaw limit:  $M_\nu \approx -M_D M_R^{-1} M_D^T$  (Fritzsch, Gell-Mann, Minkowski, 1975; Minkowski, 1977; ...)

**Type-II** seesaw limit:  $M_\nu = M_L$  (Konetschny, Kummer, 1977; ...)

# History of Seesaw

42

The **seesaw** idea **originally** appeared in a paper's **footnote**.



## **Seesaw**—A Footnote Idea:

**H. Fritzsch, M. Gell-Mann,  
P. Minkowski**, PLB 59 (**1975**) 256

This idea was very clearly elaborated by **Minkowski** in Phys. Lett. B 67 (**1977**) 421 ---- but it was unjustly forgotten until **2004**.



The idea was later on embedded into the **GUT** frameworks in **1979** and **1980**:

- **T. Yanagida** **1979**
- **M. Gell-Mann, P. Ramond, R. Slansky** **1979**
- **S. Glashow** **1979**
- **R. Mohapatra, G. Senjanovic** **1980**

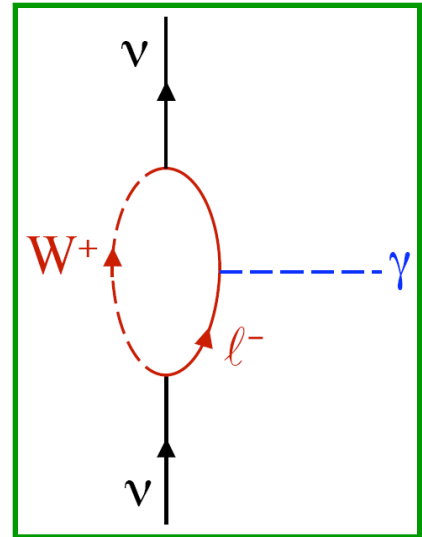
It was **Yanagida** who named this mechanism as "**seesaw**".

# Electromagnetic Properties 43

A neutrino does not have electric charges, but it has **electromagnetic interactions** with the photon via quantum loops.

Given the SM interactions, a **massive Dirac** neutrino can only have a tiny **magnetic** dipole moment:

$$\mu_\nu \sim \frac{3eG_F}{8\sqrt{2}\pi^2} m_\nu = 3 \times 10^{-20} \frac{m_\nu}{0.1 \text{ eV}} \mu_B$$



A **massive Majorana** neutrino can **not** have **magnetic** & **electric** dipole moments, as its antiparticle is itself.

**Proof:** **Dirac** neutrino's electromagnetic vertex can be parametrized as

$$\Gamma_\mu(p, p') = f_Q(q^2)\gamma_\mu + f_M(q^2)i\sigma_{\mu\nu}q^\nu + f_E(q^2)\sigma_{\mu\nu}q^\nu\gamma_5 + f_A(q^2)(q^2\gamma_\mu - q_\mu q^\nu\gamma_\nu)\gamma_5$$

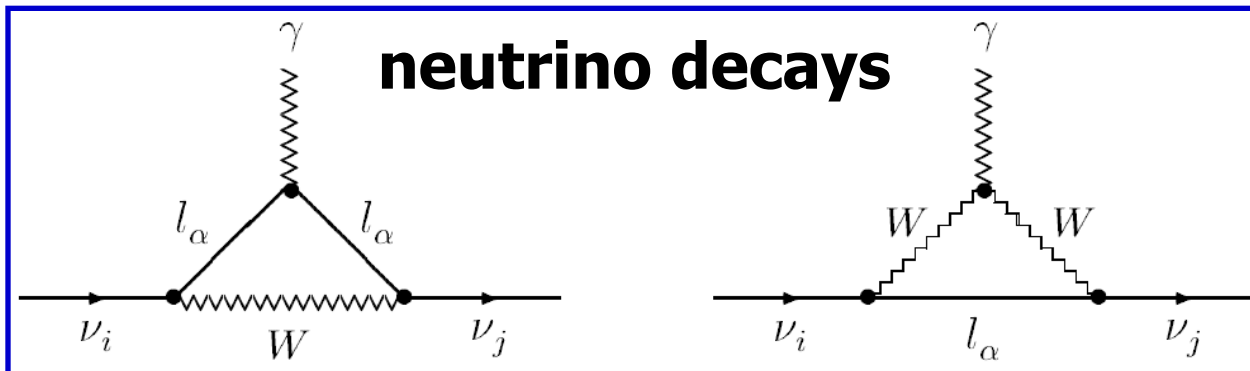
**Majorana  
neutrinos**

$$\bar{\psi}\Gamma_\mu\psi = \bar{\psi}^c\Gamma_\mu\psi^c = \psi^T\mathcal{C}\Gamma_\mu\mathcal{C}\bar{\psi}^T = \left(\psi^T\mathcal{C}\Gamma_\mu\mathcal{C}\bar{\psi}^T\right)^T = -\bar{\psi}\mathcal{C}^T\Gamma_\mu^T\mathcal{C}^T\psi = \bar{\psi}\mathcal{C}\Gamma_\mu^T\mathcal{C}^{-1}\psi$$

➡  $f_Q(q^2) = f_M(q^2) = f_E(q^2) = 0$  intrinsic property of **Majorana v's**.

# Transition Dipole Moments 44

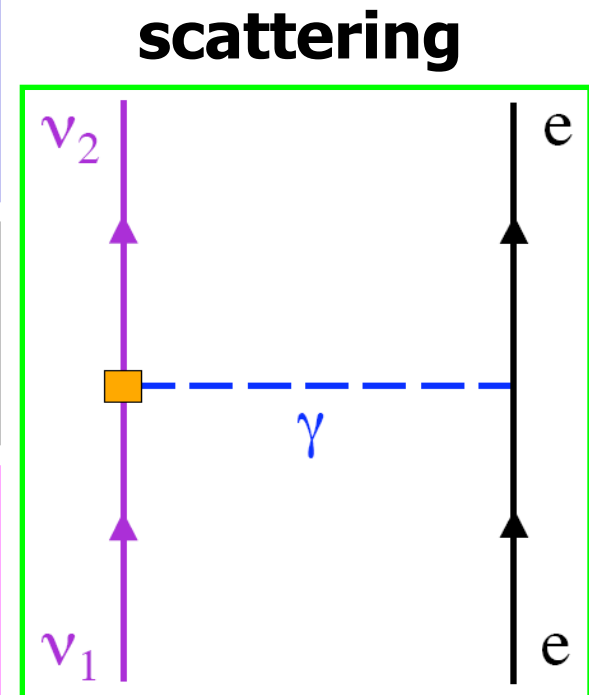
Both **Dirac** & **Majorana** neutrinos can have **transition** dipole moments (of a size comparable with  $\mu_\nu$ ) that may give rise to neutrino decays, scattering with electrons, interactions with external magnetic field & contributions to  $\nu$  masses. (**Data:**  $< \text{a few} \times 10^{-11}$  Bohr magneton).



$$\mu_{\text{eff}} \equiv \sqrt{|\mu_{ij}|^2 + |\epsilon_{ij}|^2}$$

$$\Gamma_{\nu_i \rightarrow \nu_j + \gamma} = 5.3 \times \left(1 - \frac{m_j^2}{m_i^2}\right)^3 \left(\frac{m_i}{1 \text{ eV}}\right)^3 \left(\frac{\mu_{\text{eff}}}{\mu_B}\right)^2 \text{ s}^{-1}$$

$$\frac{d\sigma'_\mu}{dT} = \frac{\alpha^2 \pi}{m_e^2} \sum_{k=1}^3 \left| \sum_{j=1}^3 e^{iq_j L} V_{ej} \left( i \frac{\mu_{jk}}{\mu_B} + \frac{\epsilon_{jk}}{\mu_B} \right) \right|^2 \left( \frac{1}{T} - \frac{1}{E_\nu} \right)$$



# Summary of Lecture A

45

- (A) Three reasons for neutrinos to be massless in the SM.
- (B) The **Dirac** mass term and **lepton number conservation**.
- (C) The **Majorana** mass term and **lepton number violation**.
  - the **Majorana** mass matrix must be **symmetric**;
  - factor **1/2** in front of the mass term makes sense.
- (D) The **hybrid** mass term and **seesaw mechanisms**.
  - light and heavy neutrinos are **Majorana** particles;
  - the **3×3 light** flavor mixing matrix is **non-unitary**;
  - light neutrino masses: the type-(I+II) seesaw.
- (E) **Electromagnetic dipole moment** of massive neutrinos.
  - **Dirac** neutrinos have **magnetic** dipole moments;
  - **Majorana** neutrinos have no dipole moments;
  - **Dirac & Majorana** neutrinos: **transition** moments.