

Latest results from LHCf on very forward particle production at LHC

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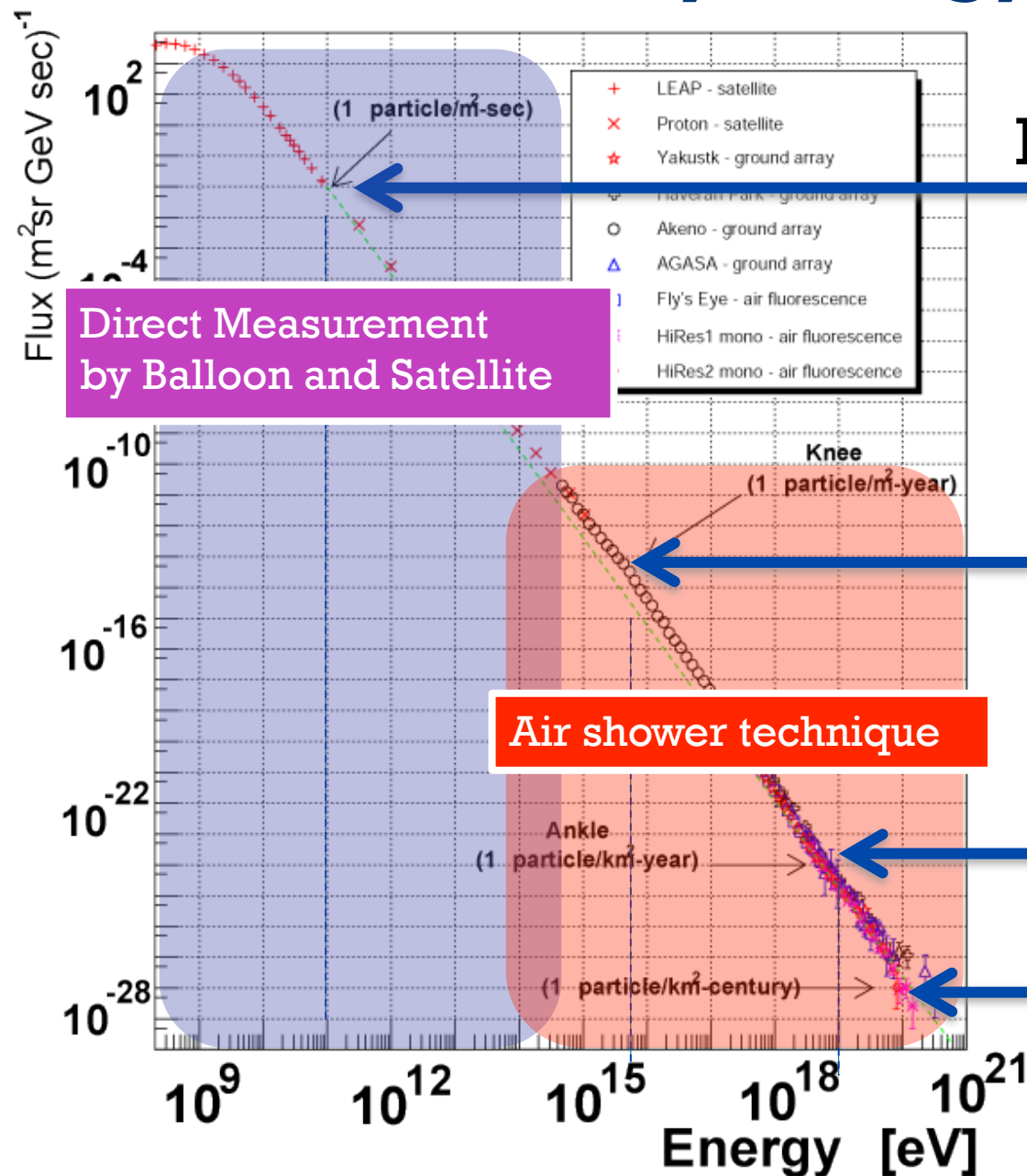
Latest results from the LHC
CERN, July 12th, 2012



Physics Motivations

Impact on High Energy Cosmic Ray Physics

+ *The Cosmic Rays energy spectrum*



1 particle/ m^2/sec

1 particle/ m^2/year

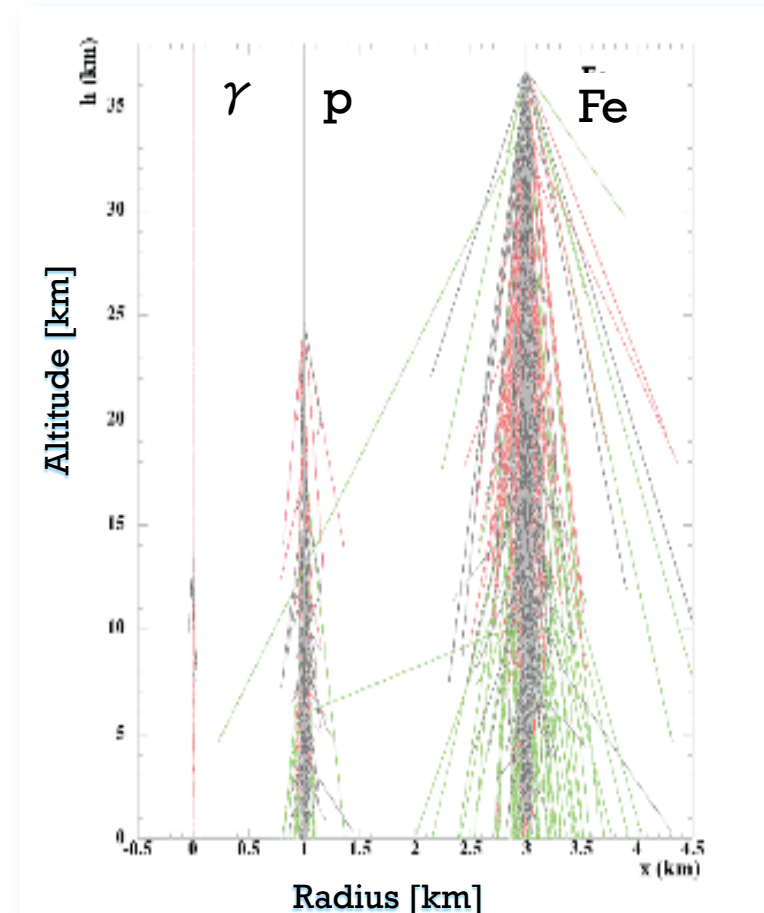
1 particle/ km^2/year

1 particle/ $\text{km}^2/\text{century}$

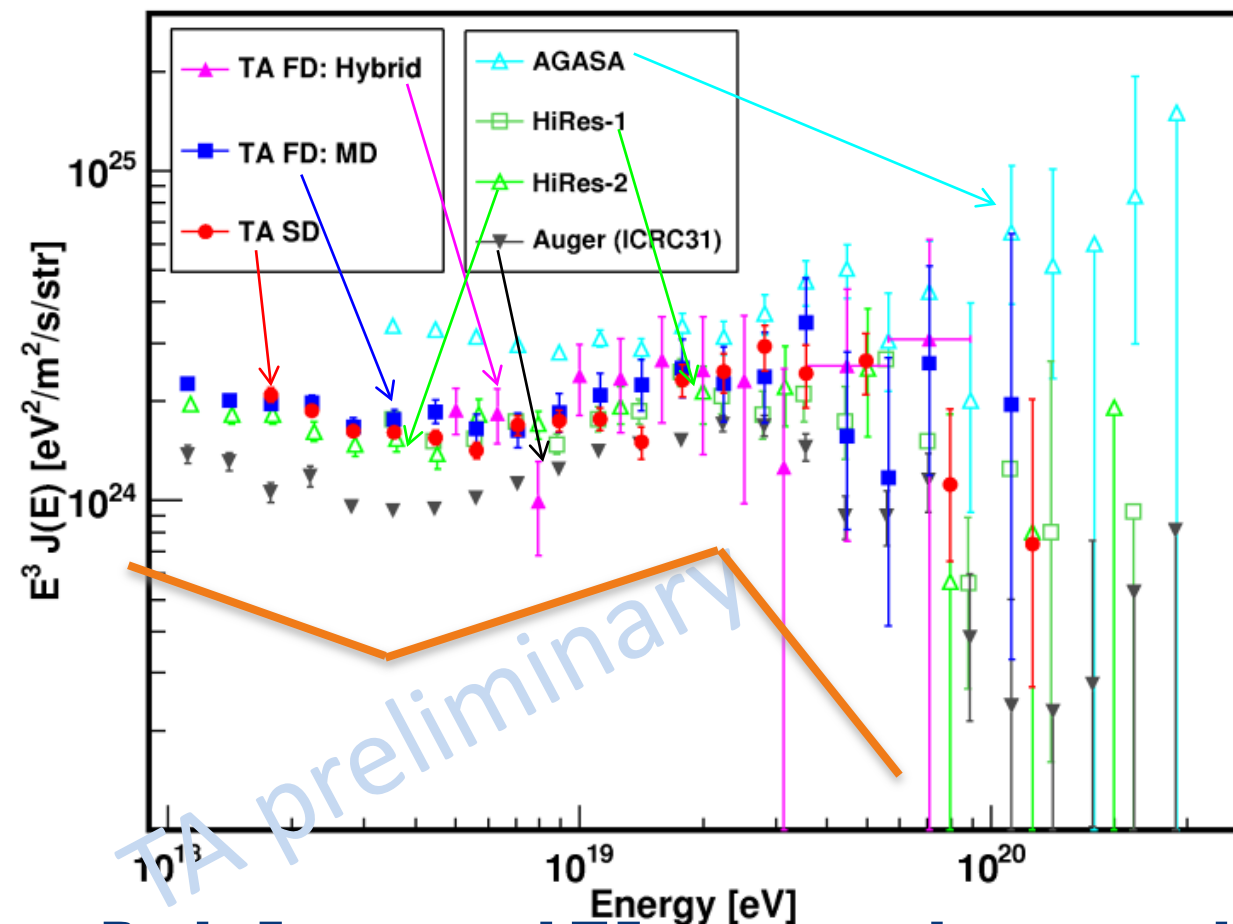
+ *Indirect measurement of cosmic rays*



- Composition and energy of cosmic rays affect the generation of Extended Air Showers
- Precise understanding of high-energy cosmic ray should be achieved with the indirect measurement technique
- Comparison between the MC simulation of EAS and observation is necessary to infer the information on the primary CR from the measurement of the shower
- **Largest systematic uncertainty of indirect measurement is caused by the finite understanding of the hadronic interaction of cosmic ray in atmosphere**



+ Example I: the VHECR Energy spectrum



Auger and Telescope Array are currently taking data

Auger has the highest statistics.

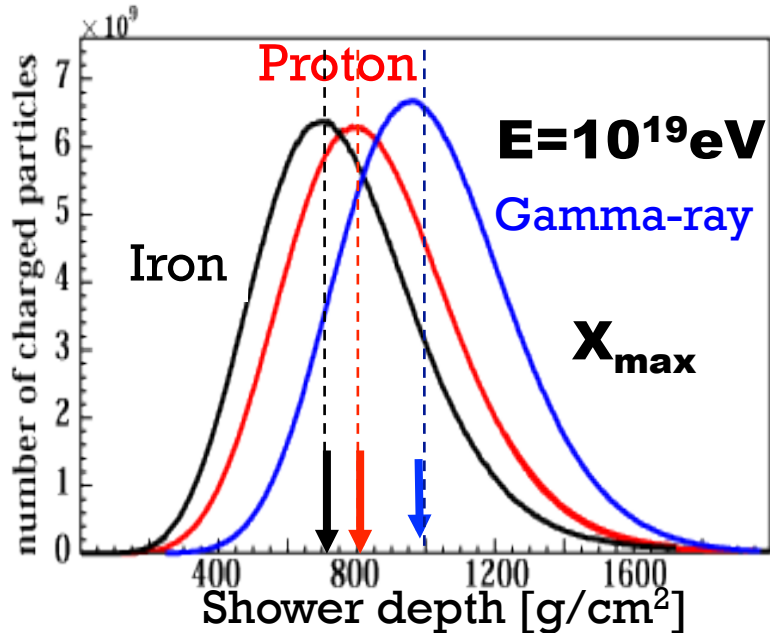
Two kinks: the ankle and the GZK cutoff

Both Auger and TA spectra shows two kinks... **BUT...**
The fluxes are significantly different!
Energy scale problem? → Precise knowledge of the hadronic interaction mechanism at VHE is necessary

+ Example II: Chemical Composition

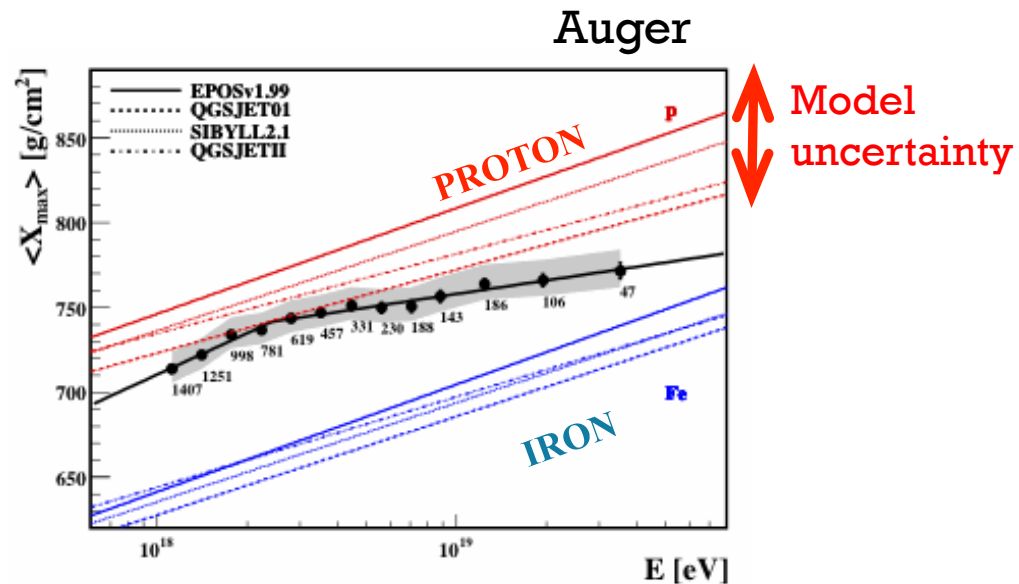
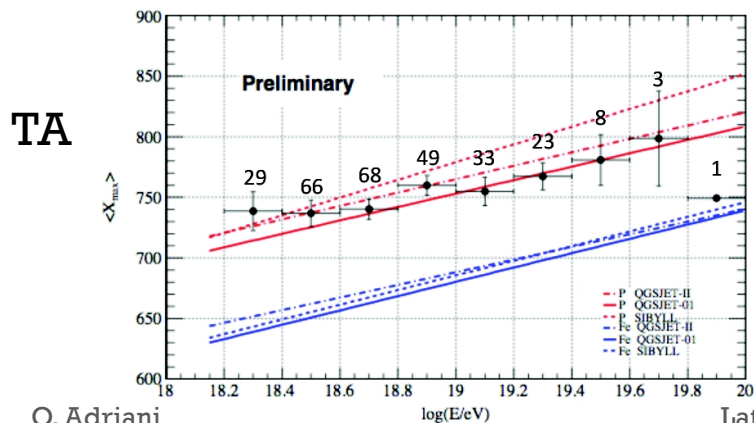


- The Chemical composition can be inferred from the depth of the maximum of the shower (X_{\max})



BUT....

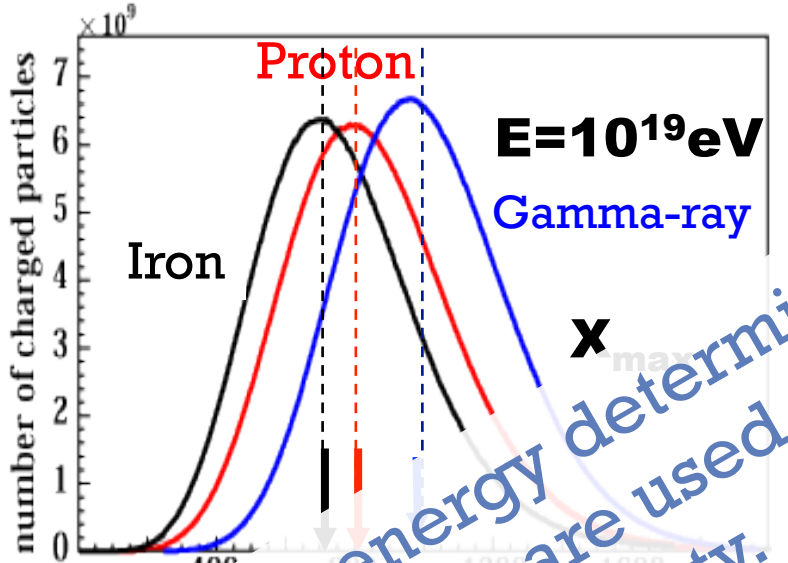
The results depend on the precise knowledge of the hadronic interaction mechanism!!!!



+ Example II: Chemical Composition



- The Chemical composition can be inferred from the depth of the maximum of the shower (X_{\max})



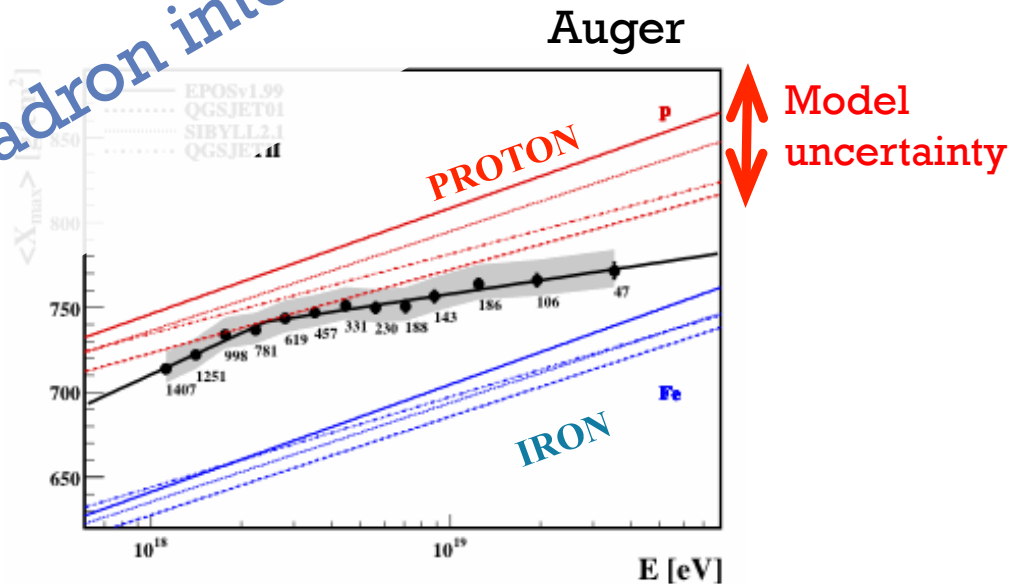
BUT....

The

Both in the energy determination and X_{\max} prediction the results depend on the precise knowledge of the hadronic interaction mechanism!!!!

MC simulations are used, and are one of the greater sources of uncertainty.

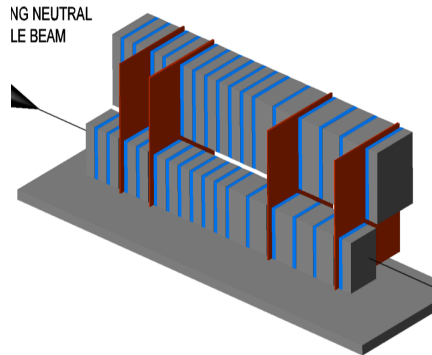
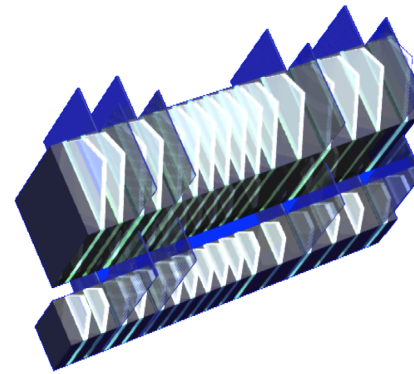
Experimental tests of hadron interaction models are necessary! → LHCf



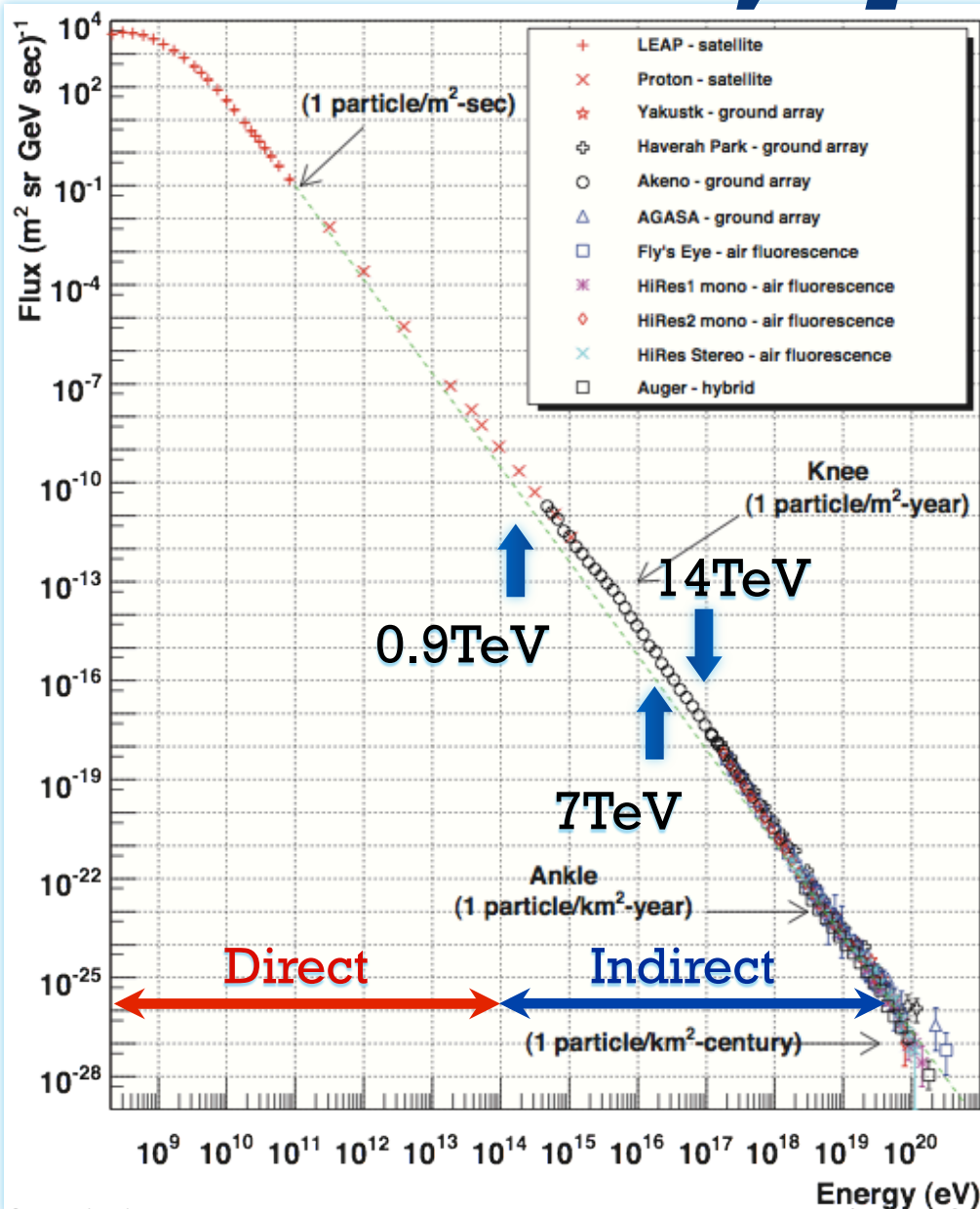


Why LHCf @ LHC?

The experimental set-up

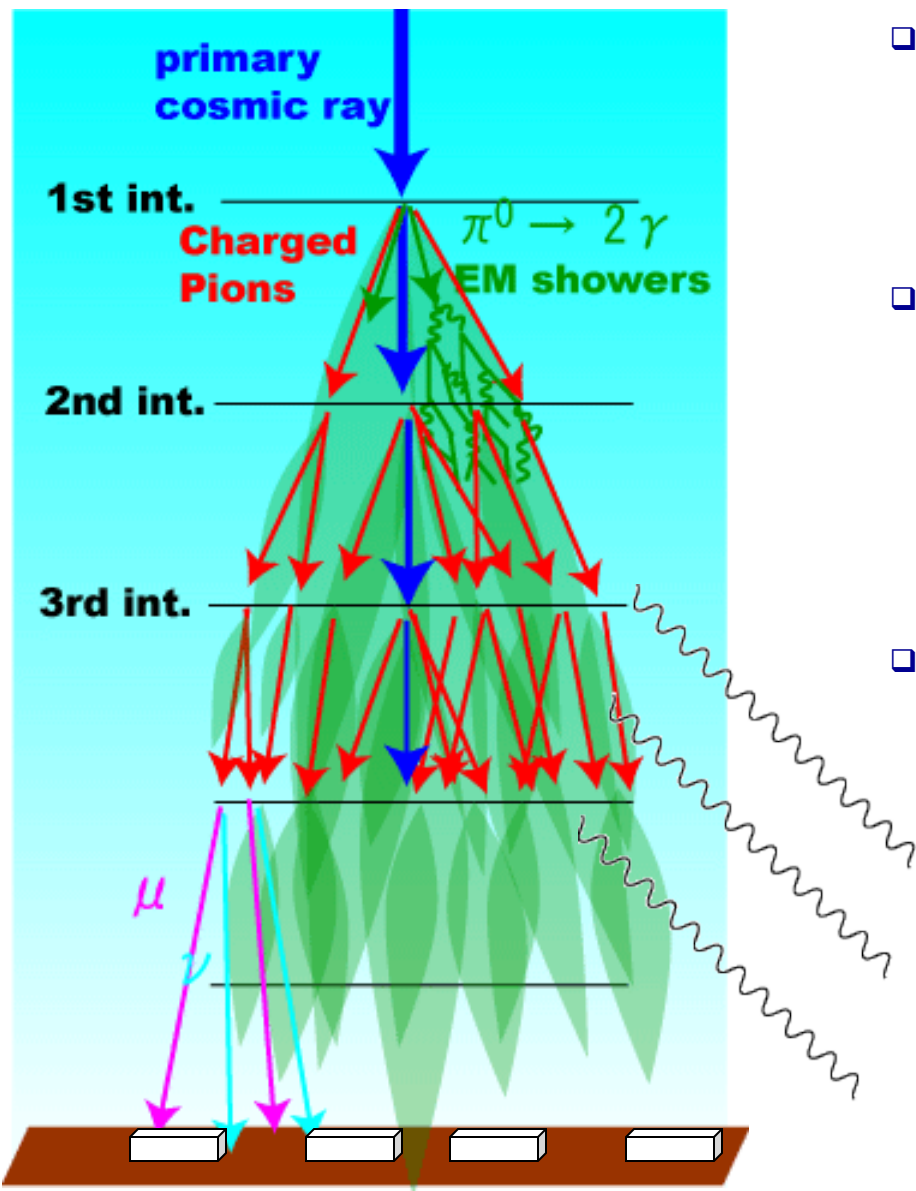
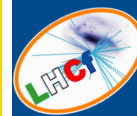


+ *The Cosmic Ray spectrum*

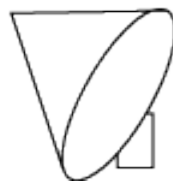


LHC Energy range is very interesting for Cosmic Ray Physics!

+ *Air Shower development*



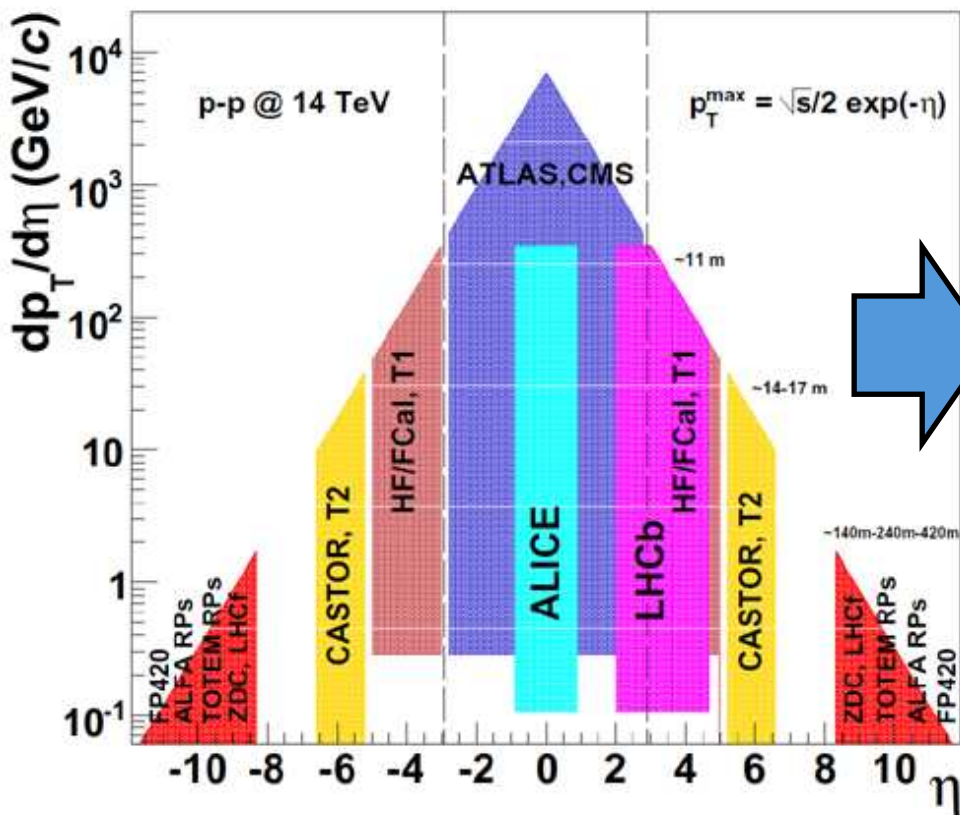
- **Total cross section**
 - Large $\sigma \rightarrow$ rapid development
 - Small $\sigma \rightarrow$ deep penetrating
- **Multiplicity (N)**
 - Large N \rightarrow rapid development
large number of muons
 - Small N \rightarrow deep penetrating
small number of muons
- **Inelasticity(k)/Secondary spectra**
 - Large k, Softer spectra
 \rightarrow rapid development
 - Small k, harder spectra
 \rightarrow deep penetrating



+ *LHC experiments*



Whole pseudo-rapidity is covered by the different LHC experiments

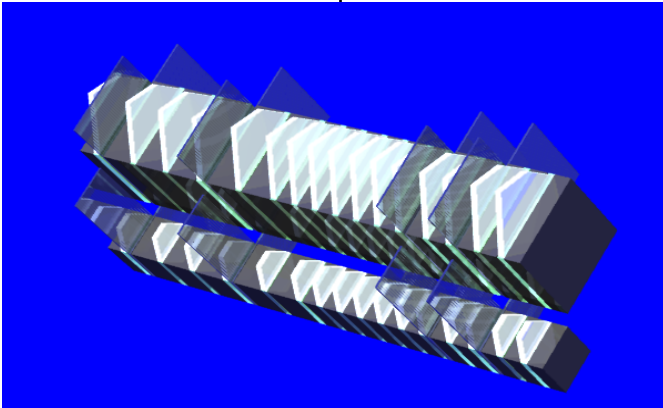
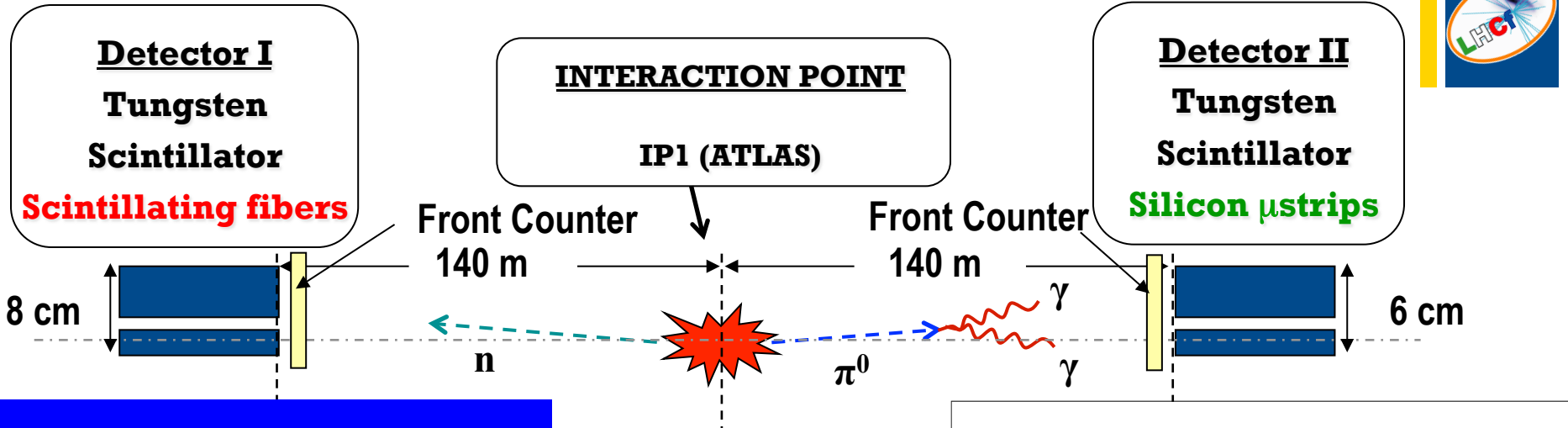


Key parameters
for air shower developments

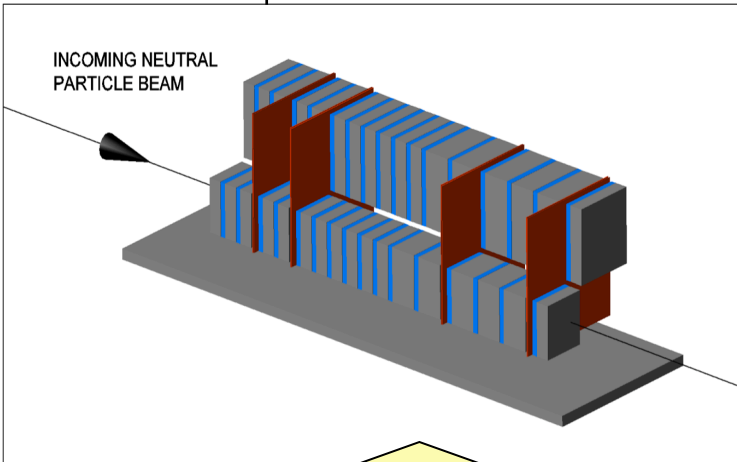
- Total cross section
↔ TOTEM, ATLAS, CMS
- Multiplicity
↔ Central detectors
- Inelasticity/Secondary spectra
↔ Forward calorimeters
LHCf, ZDCs

R. Orava, (2005)

+ LHCf: location and detector layout



$$44 X_0, \\ 1.6 \lambda_{\text{int}}$$



Arm#1 Detector
20mmx20mm+40mmx40mm
4 X-Y SciFi tracking layers

Expected Performance
Energy resolution
< 5% for γ
~30% for neutrons
Position resolution
~ 100 μ m

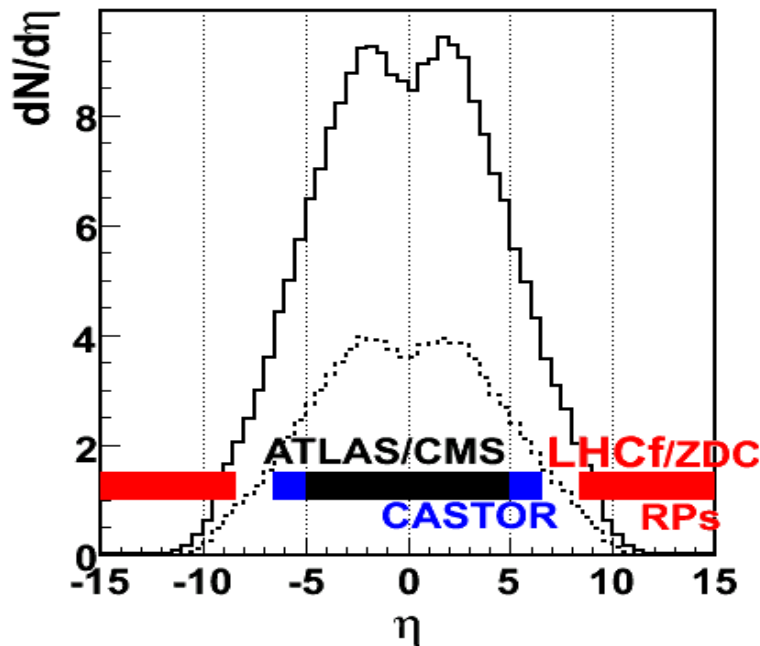
Arm#2 Detector
25mmx25mm+32mmx32mm
4 X-Y Silicon strip layers

+ *What LHCf can measure?*



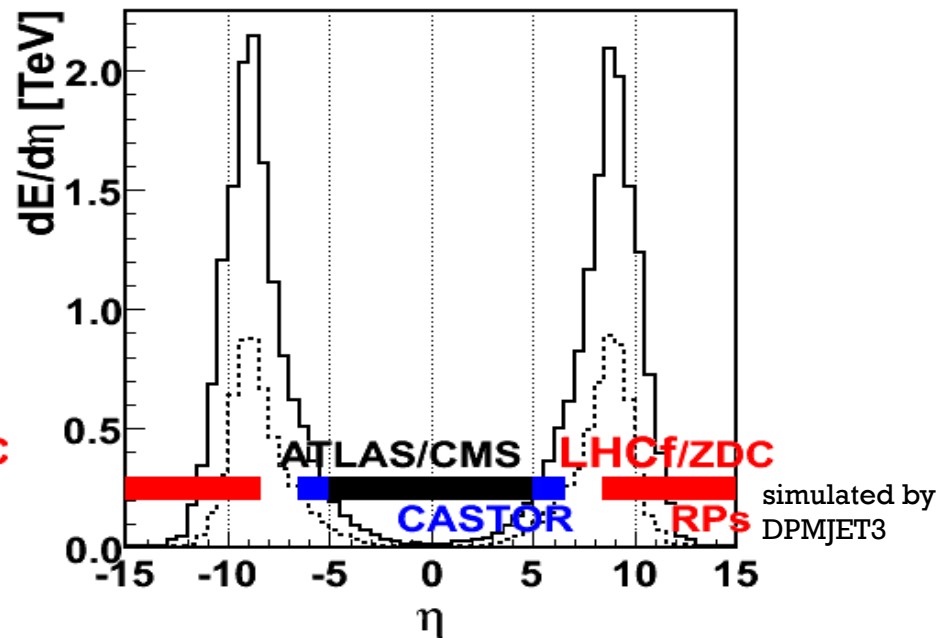
Energy spectra and Transverse momentum distribution of photons, neutrons and π^0 in the pseudorapidity region $\eta > 8.4$

Multiplicity@14TeV



Low multiplicity !!

Energy Flux @14TeV



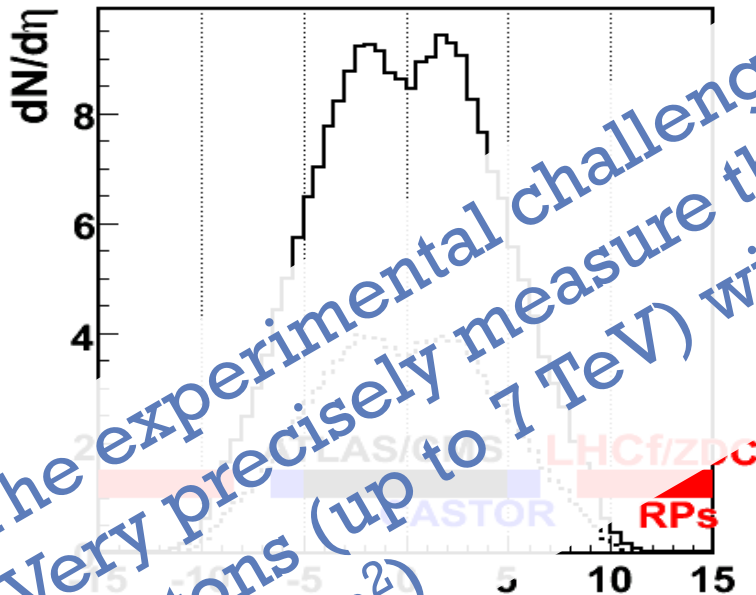
High energy flux !!

+ *What LHCf can measure?*



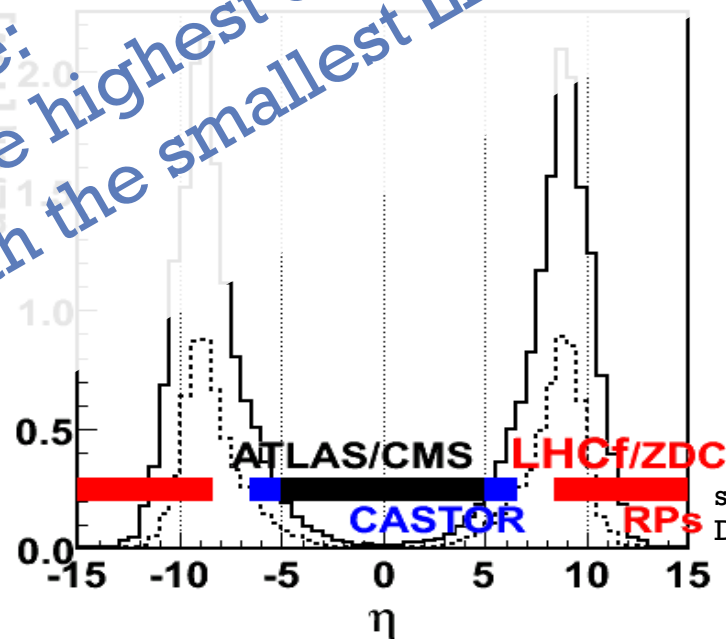
Energy spectra and Transverse momentum distribution of photons, neutrons and π^0 in the pseudorapidity region > 8.4

Multiplicity@14TeV



The experimental challenge:
Very precisely measure the highest energy LHC photons (up to 7 TeV) with the smallest LHC detector (~few cm^2)

Low multiplicity !!



High energy flux !!

simulated by
DPMJET3



LHCf main physics results

“Measurement of zero degree single photon energy spectra for $\sqrt{s} = 7$ TeV proton-proton collisions at LHC”

PLB 703 (2011) 128

“Measurement of zero degree single photon energy spectra for $\sqrt{s} = 900$ GeV proton-proton collisions at LHC”

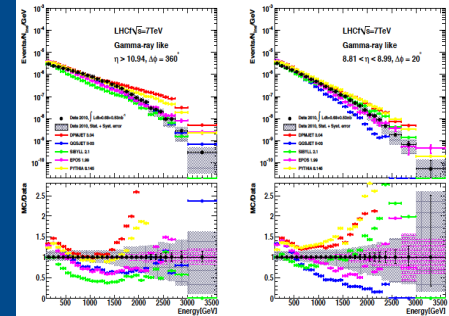
Submitted to PLB

CERN-PH-EP-2012-048

“Measurement of forward neutral pion transverse momentum spectra for $\sqrt{s} = 7$ TeV proton-proton collisions at LHC”

Submitted to PRD

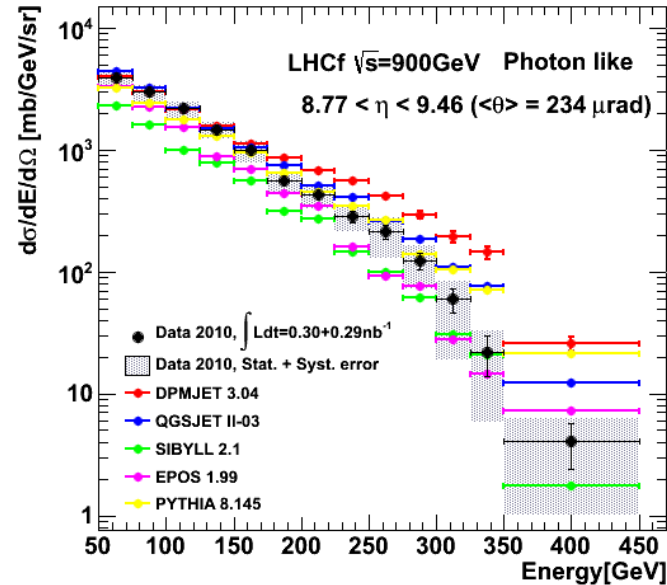
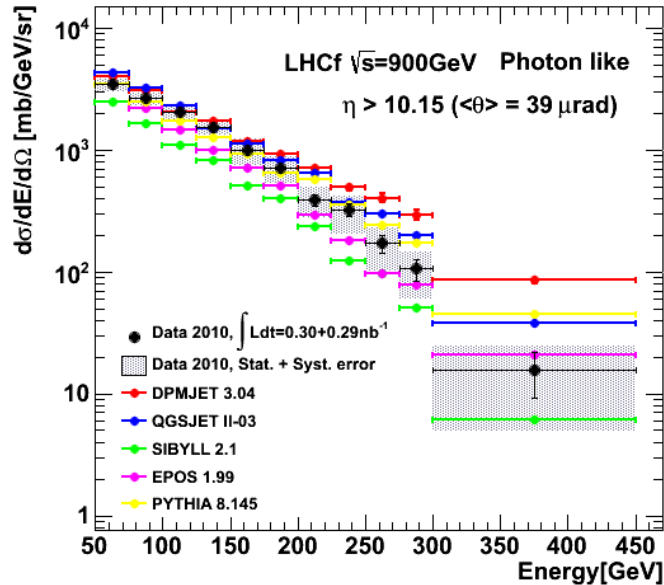
CERN-PH-EP-2012-145



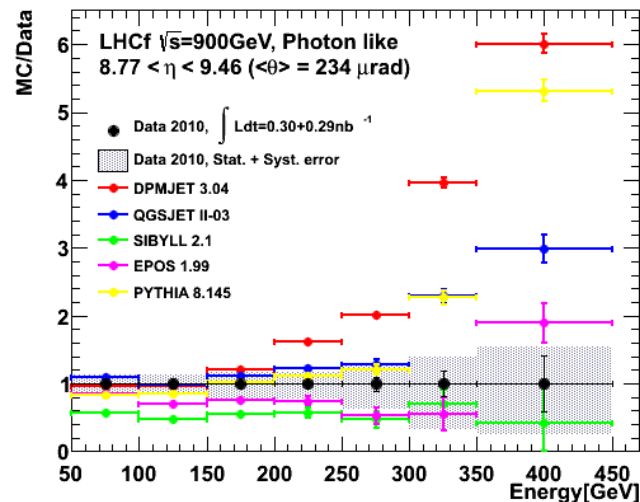
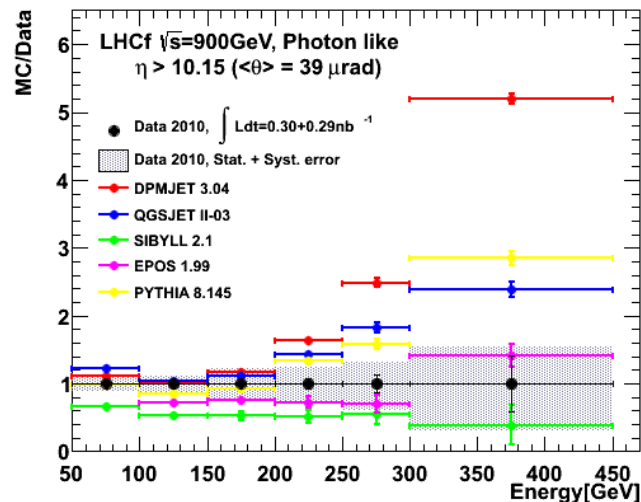
+ Inclusive γ spectrum at 900 GeV



Data



MC/Data



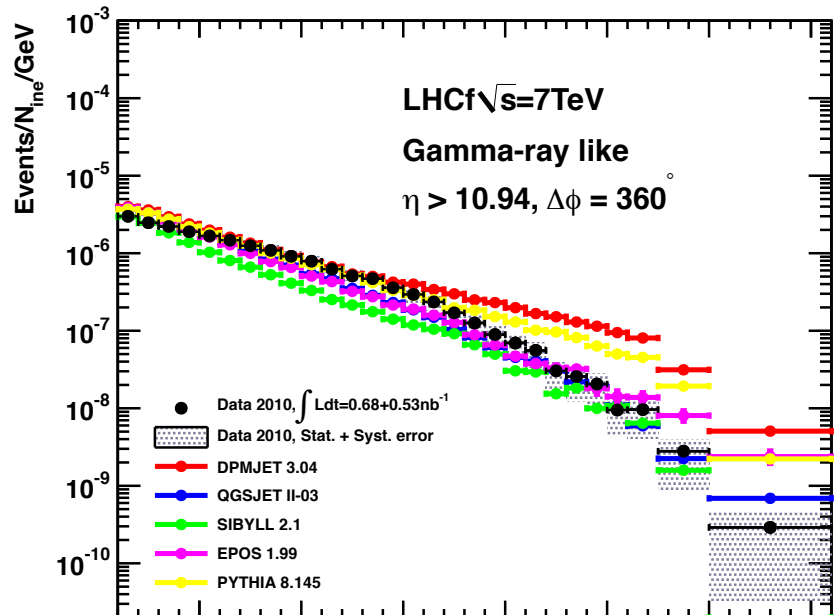
DPMJET
3.04
SIBYLL
2.1
EPOS
1.99
PYTHIA
8.145
QGSJET
II-03

Gray hatch :
Systematic
Errors

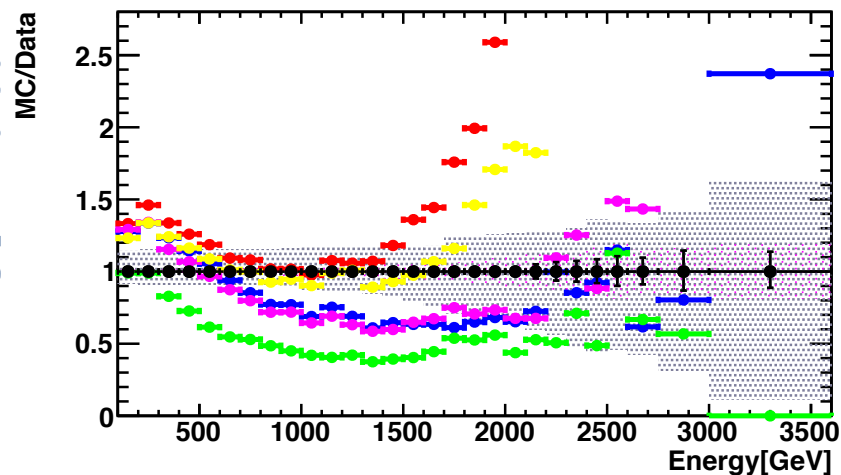
+ Inclusive γ spectrum at 7 TeV



Data

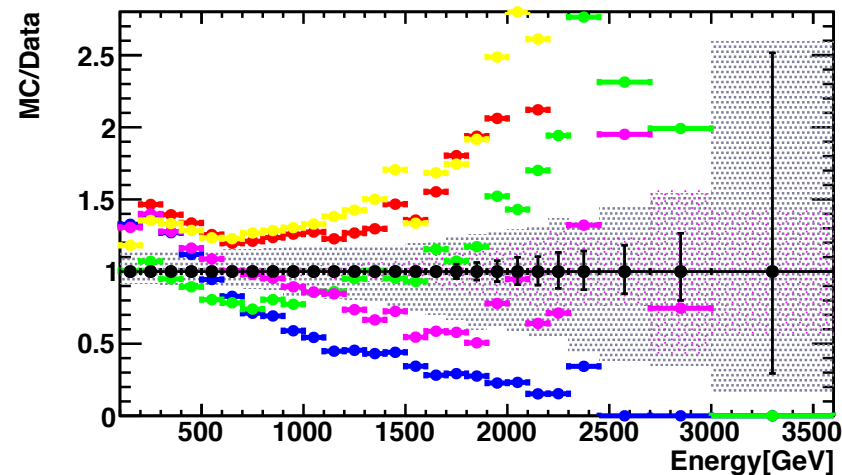
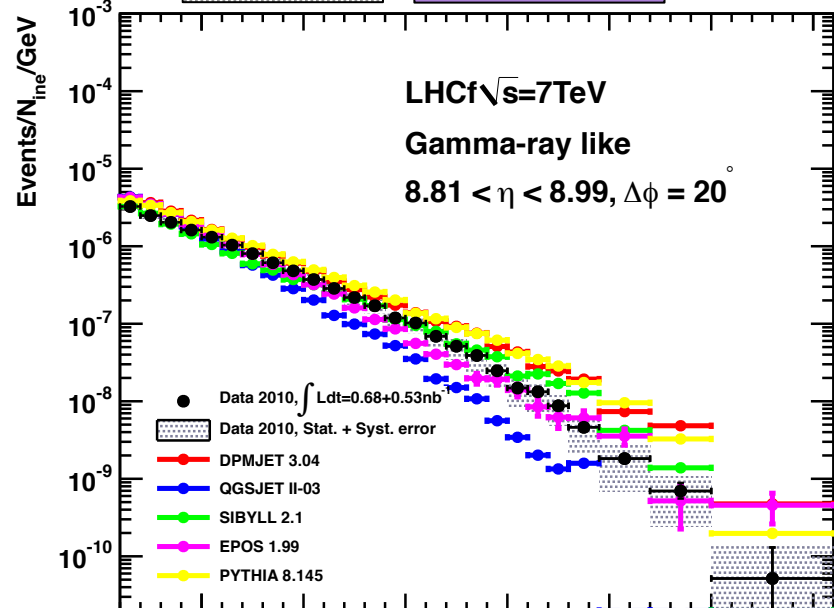


MC/Data



Gray hatch :
Systematic
Errors

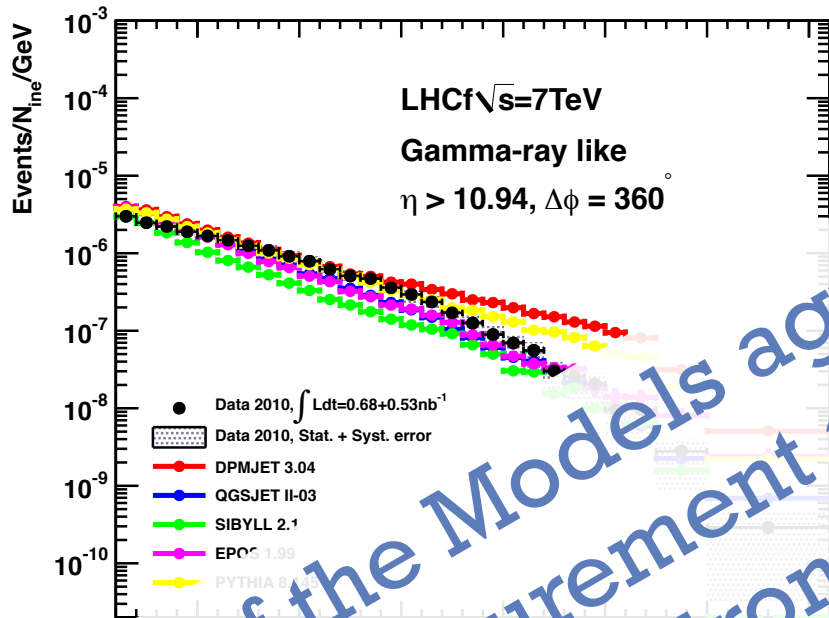
Magenta hatch:
MC Statistical
errors



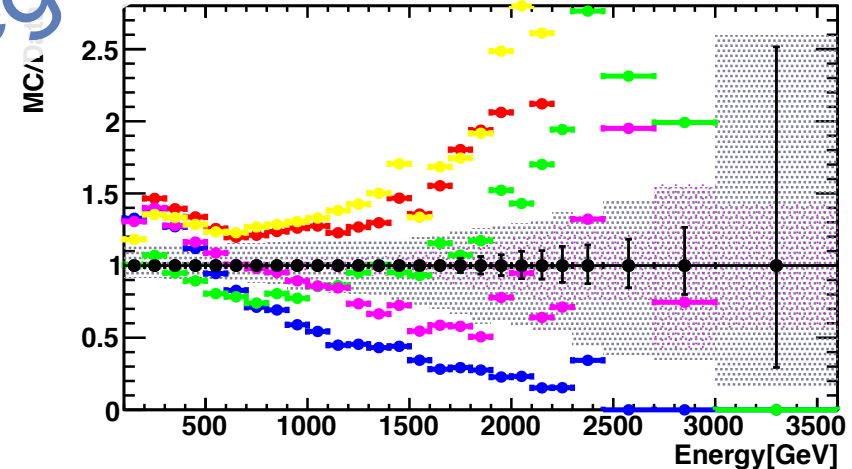
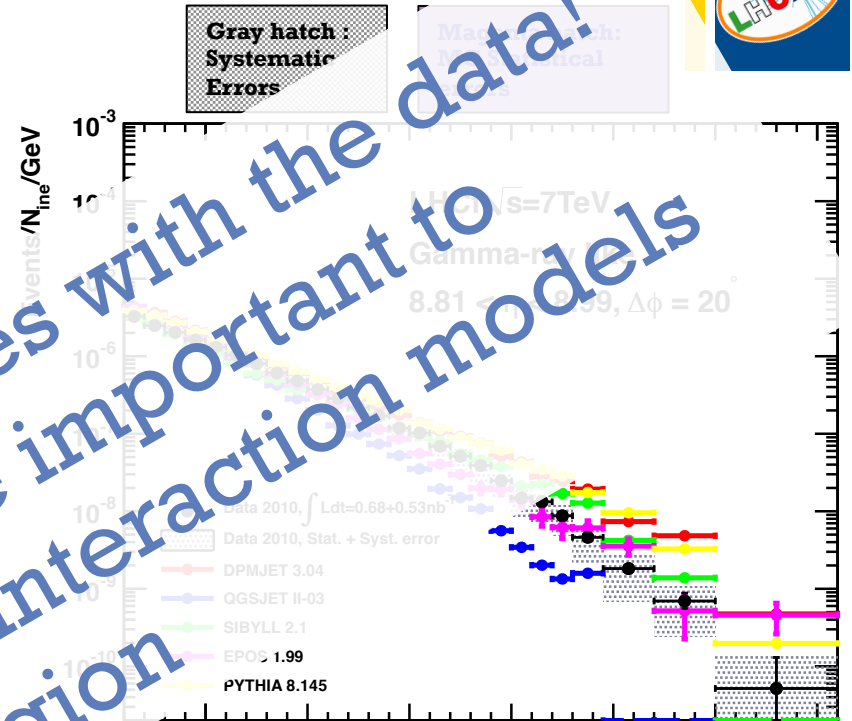
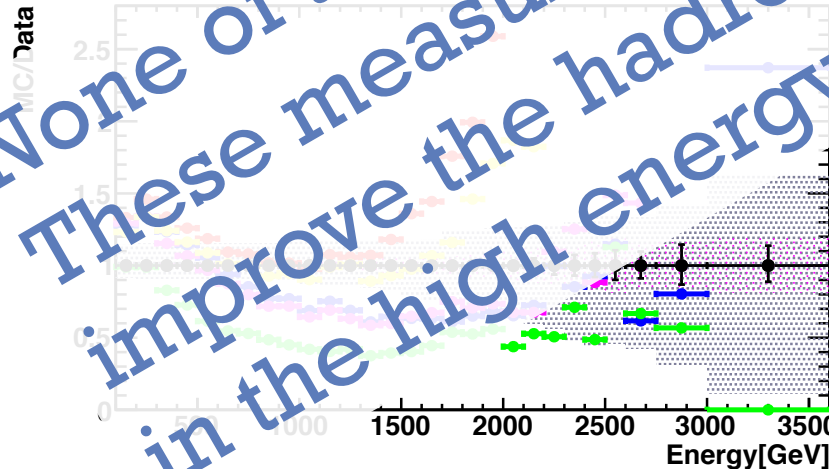
+ Inclusive γ spectrum at 7 TeV



Data



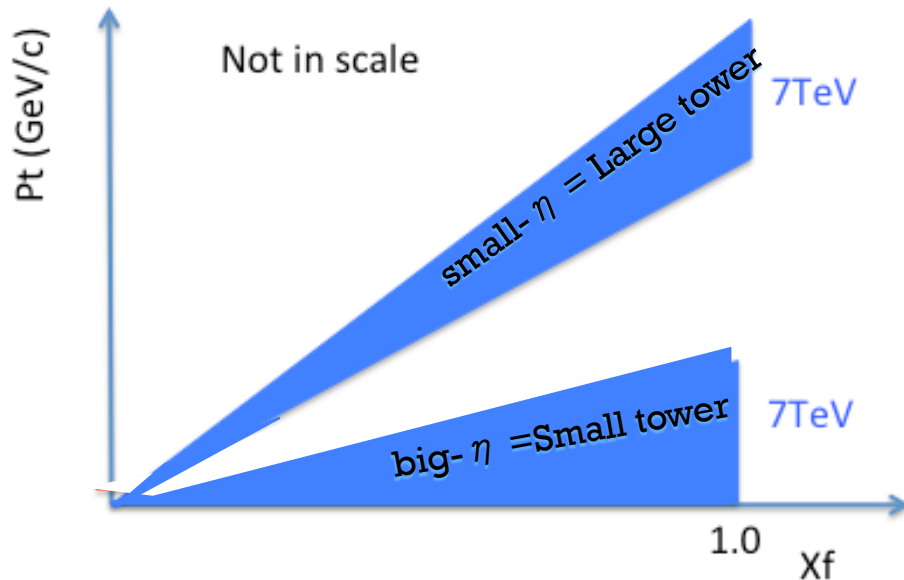
MC/Data



+ *Comparison between 900 GeV and 7 TeV spectra*



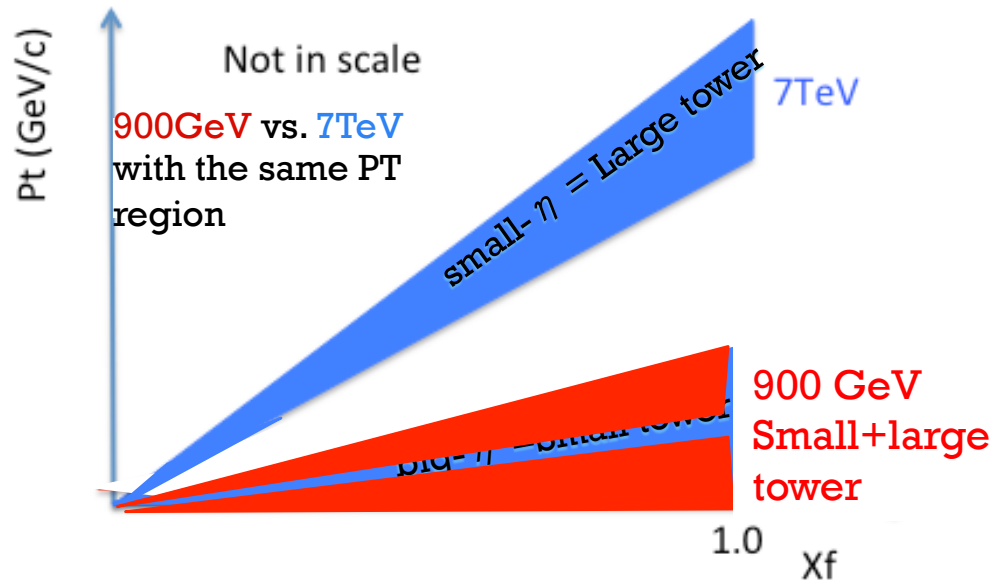
*Coverage of the photon spectra
in the plane Feynman- X vs P_T*



+ Comparison between 900 GeV and 7 TeV spectra



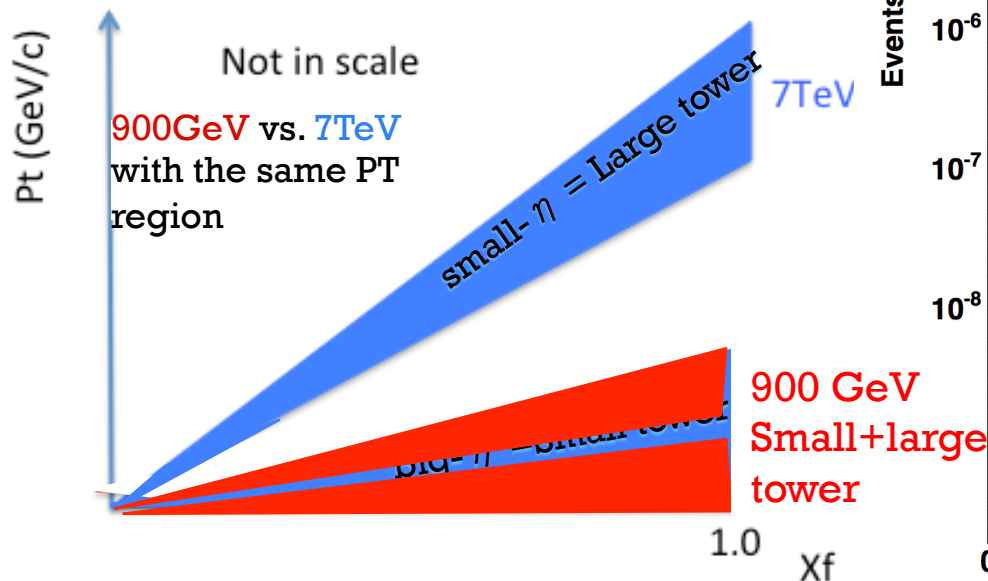
Coverage of the photon spectra
in the plane Feynman- X vs P_T



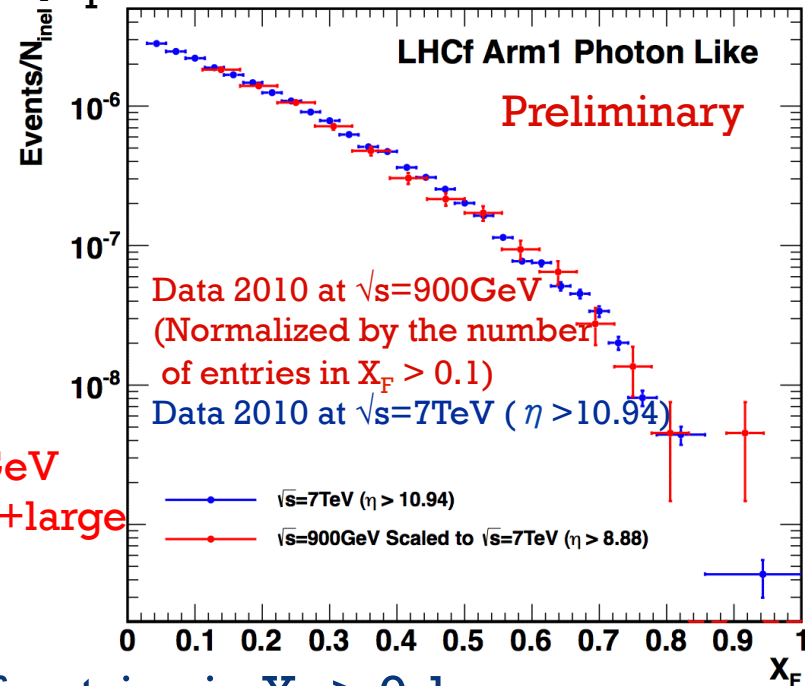
+ Comparison between 900 GeV and 7 TeV spectra



Coverage of the photon spectra in the plane Feynman- X vs P_T



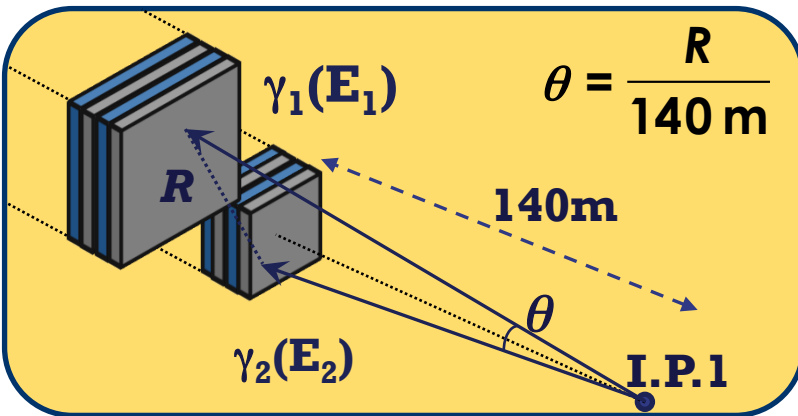
X_F spectra : 900GeV data vs. 7TeV data



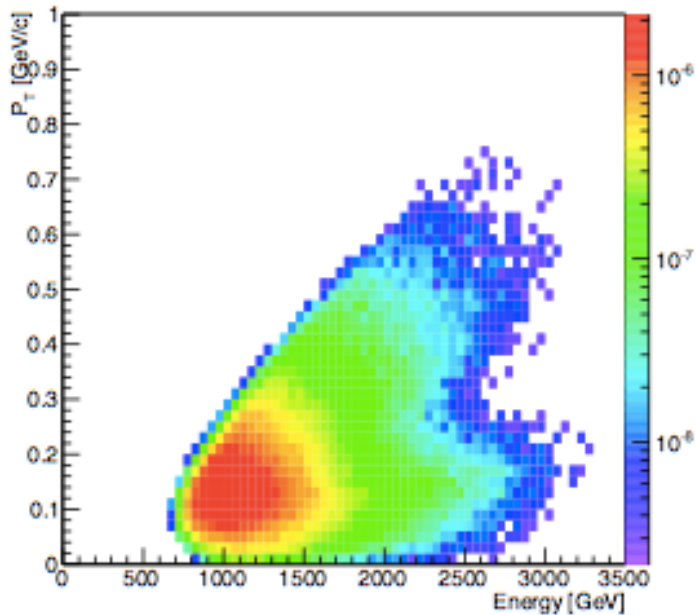
- ✓ Normalized by the number of entries in $X_F > 0.1$
- ✓ No systematic error is considered in both collision energies.

Good agreement of X_F spectrum shape between 900 GeV and 7 TeV.
→ weak dependence of $\langle p_T \rangle$ on E_{CMS}

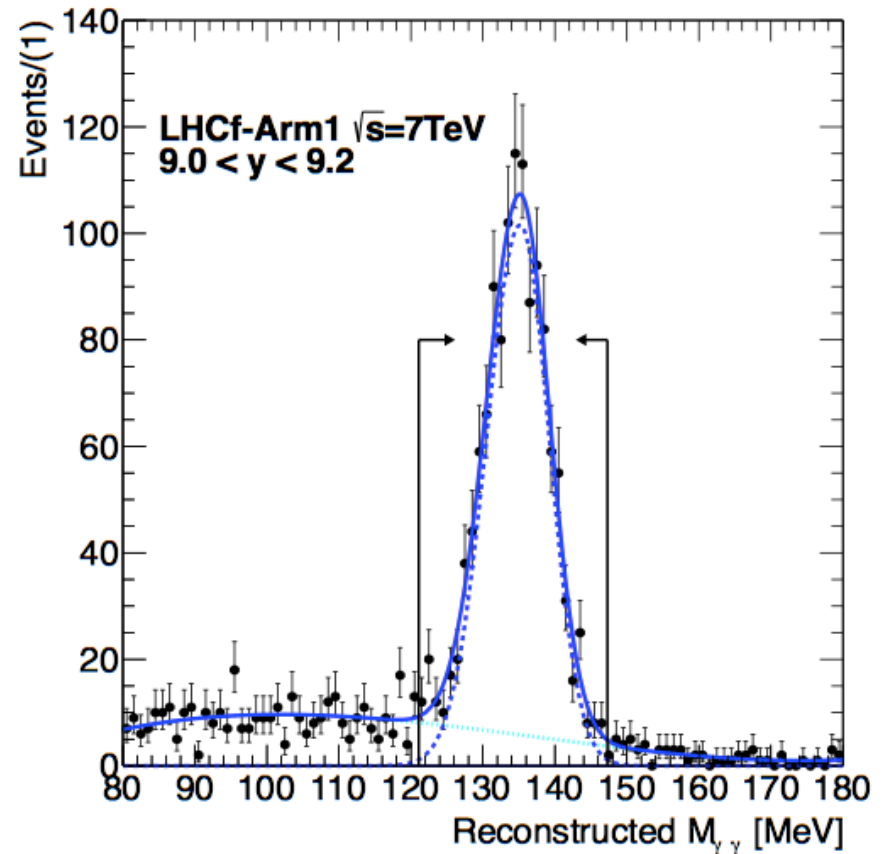
+ 7 TeV π^0 analysis



Type-I sample

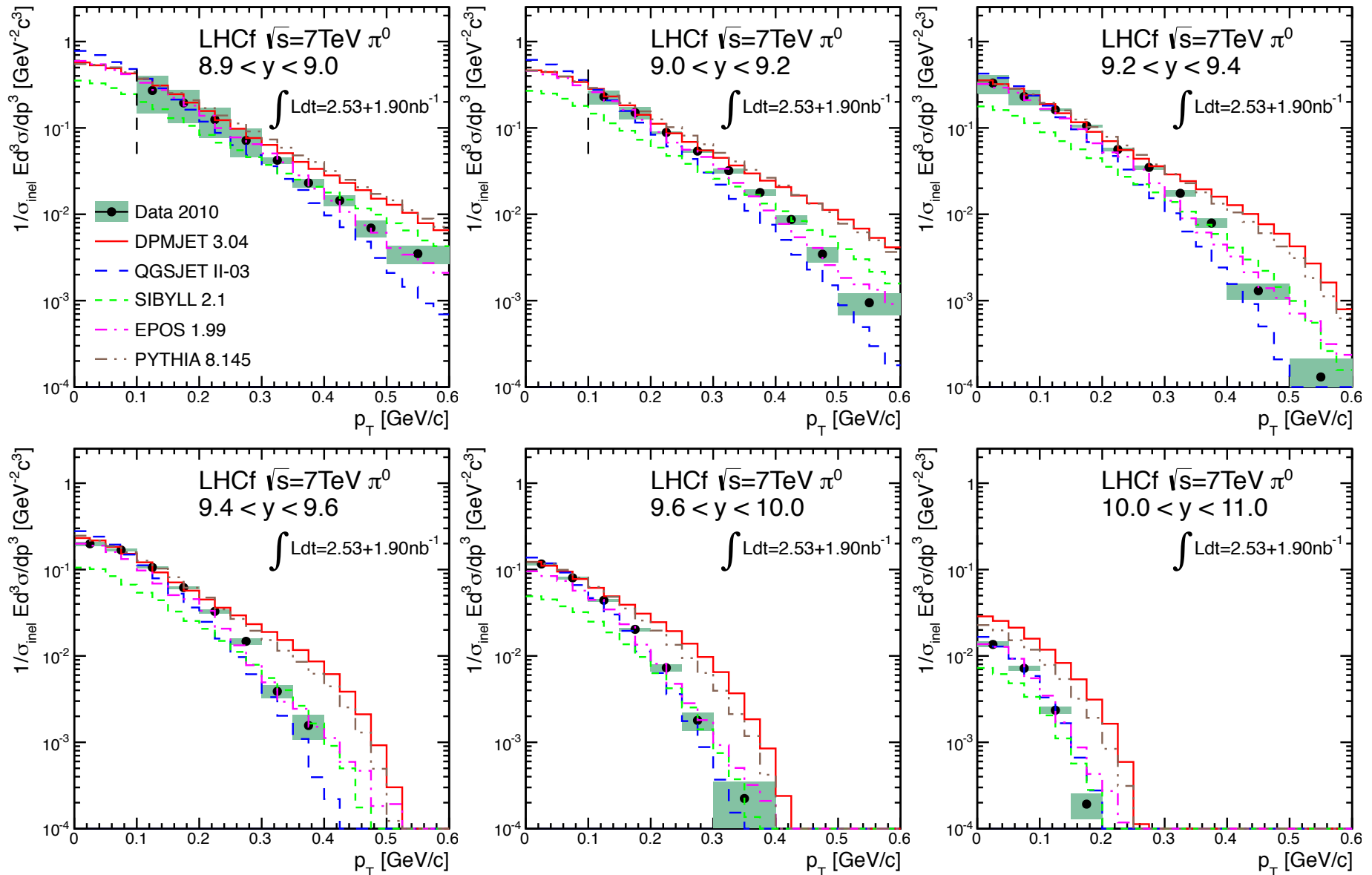


Mass, energy and transverse momentum are reconstructed from the energies and impact positions of photon pairs measured by each calorimeter



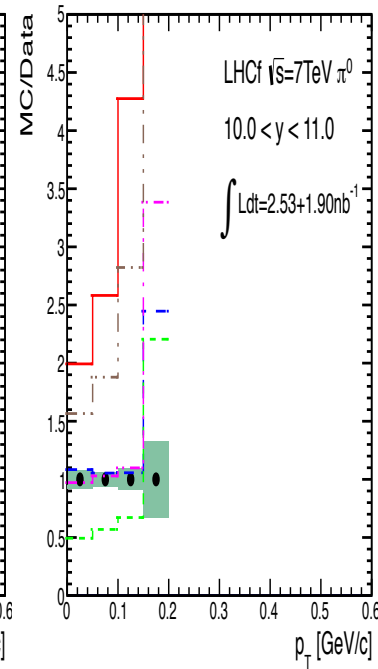
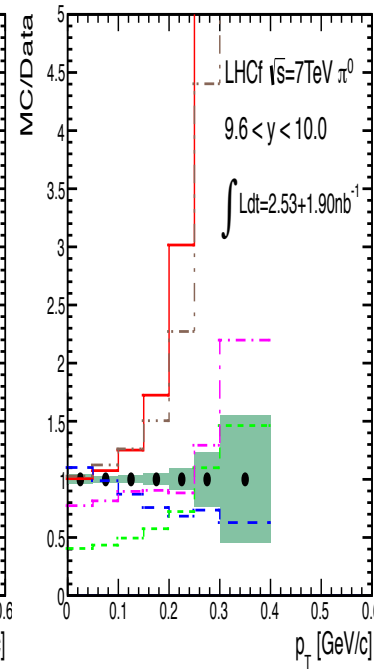
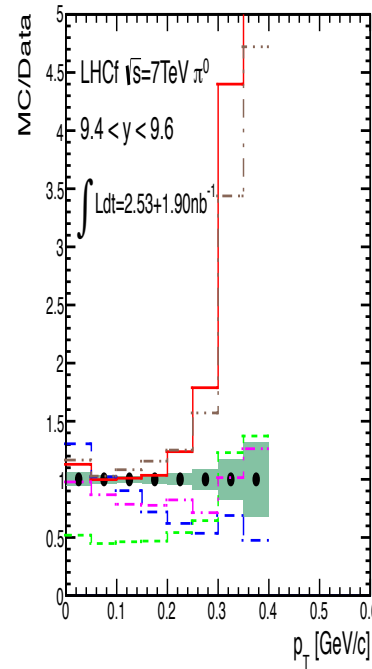
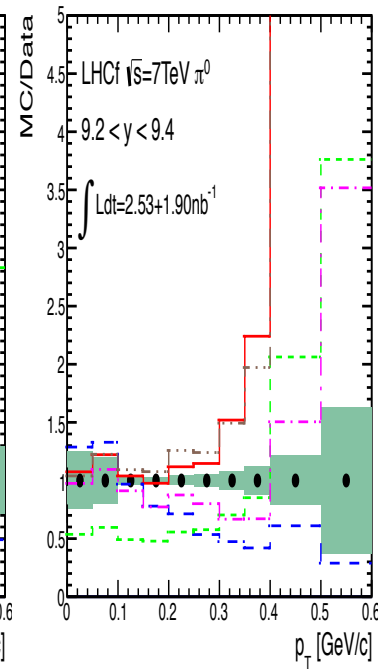
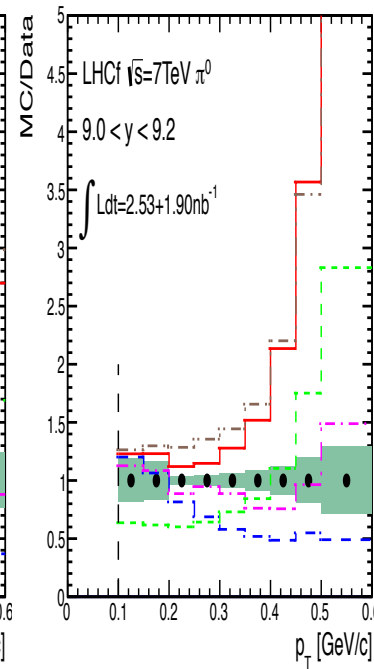
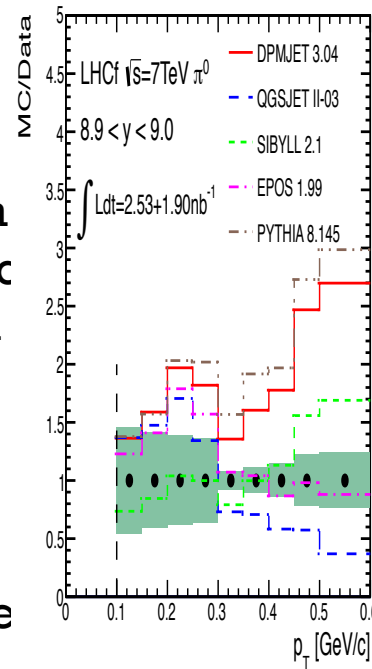
+ π^0 Data vs MC:

P_T spectra for different rapidity bins



+ π^0 Data vs MC

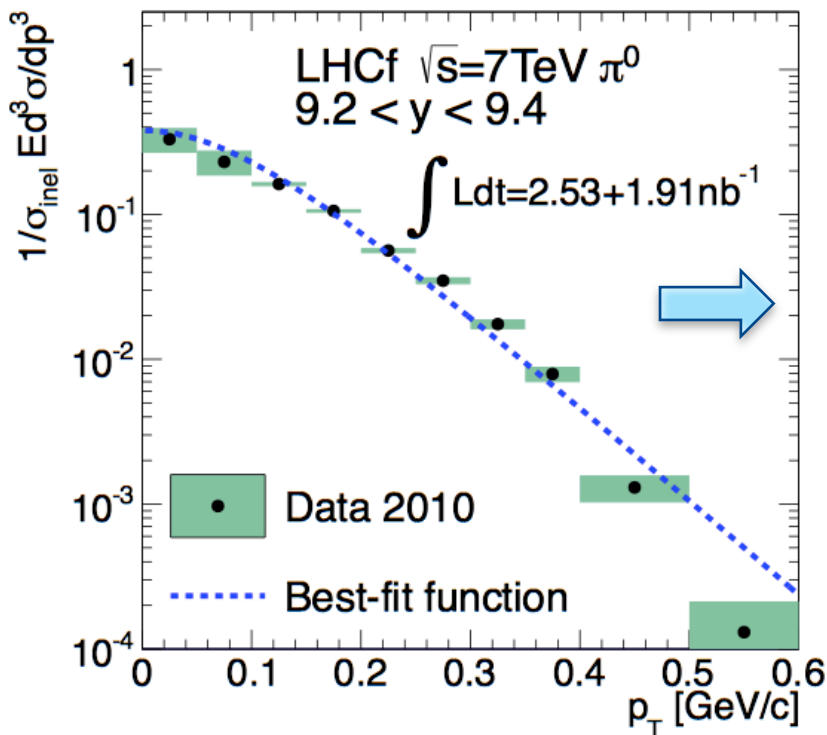
- **dpmjet 3.04** & **pythia 8.145** show overall agreement with LHCf data for $9.2 < y < 9.6$ and $p_T < 0.25$ GeV/c while the expected π^0 production rates by both models exceed the LHCf data as p_T becomes large
- **sibyll 2.1** predicts harder pion spectra than data, but the expected π^0 yield is generally small
- **qgsjet II-03** predicts π^0 spectra softer than LHCf data
- **epos 1.99** shows the best overall agreement with the LHCf data.
 - behaves softer in the low p_T region, $p_T < 0.4$ GeV/c in $9.0 < y < 9.4$ and $p_T < 0.3$ GeV/c in $9.4 < y < 9.6$
 - behaves harder in the large p_T region.



+ π^0 analysis at $\sqrt{s}=7$ TeV



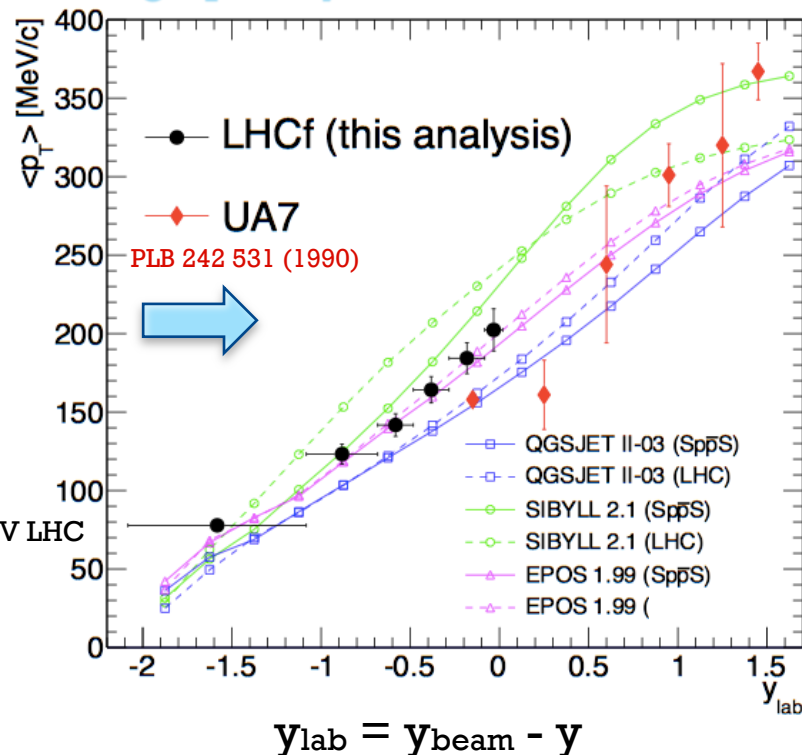
p_T spectra vs best-fit function



$\langle p_T \rangle$ for
different y
bins

$Y_{\text{Beam}}=6.5$ for SPS
 $Y_{\text{Beam}}=8.92$ for 7 TeV LHC

Average p_T vs y_{lab}



Systematic uncertainty of LHCf data is 5%.

Comparison with the UA7 data ($\sqrt{s}=630$ GeV) and MC simulations (QGSJET, SIBYLL, EPOS).

Two experimental data mostly appear to lie along a common curve

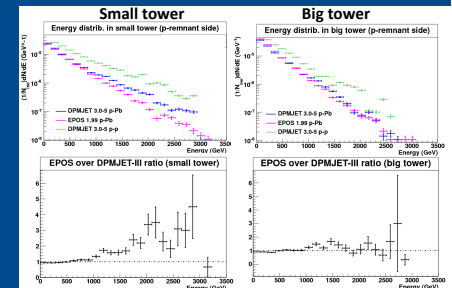
→ No evident dependence of $\langle p_T \rangle$ on E_{CMS} .

Smallest dependence on E_{CMS} is found in EPOS and it is consistent with LHCf and UA7.

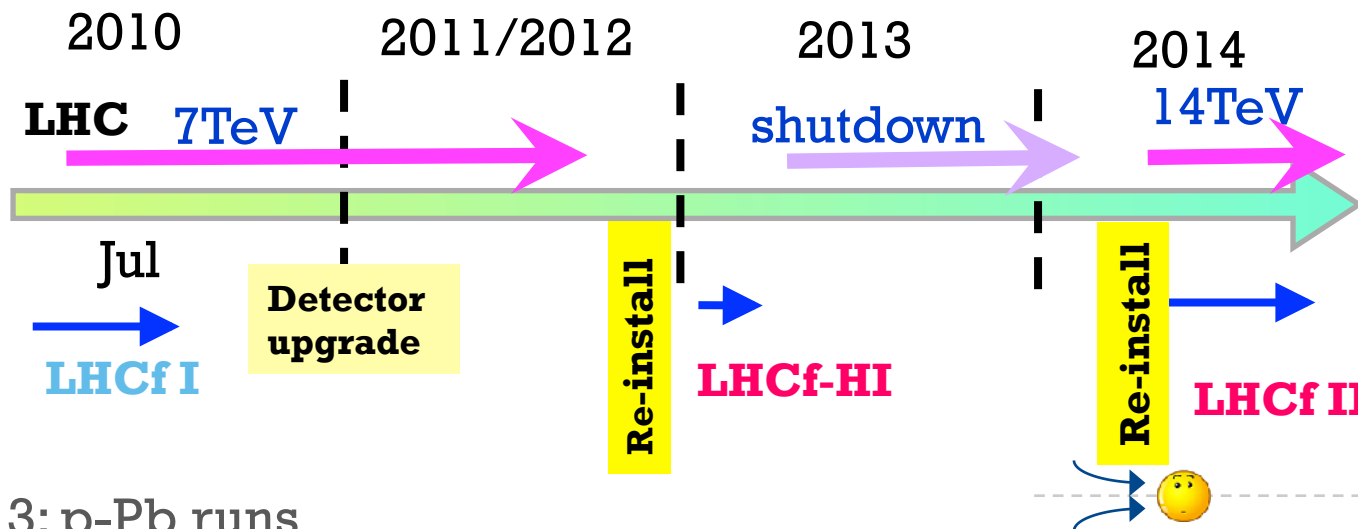
Large E_{CMS} dependence is found in SIBYLL



What's next and ... conclusions ...



+ LHCf Future PLANS: Ion run and 14 TeV

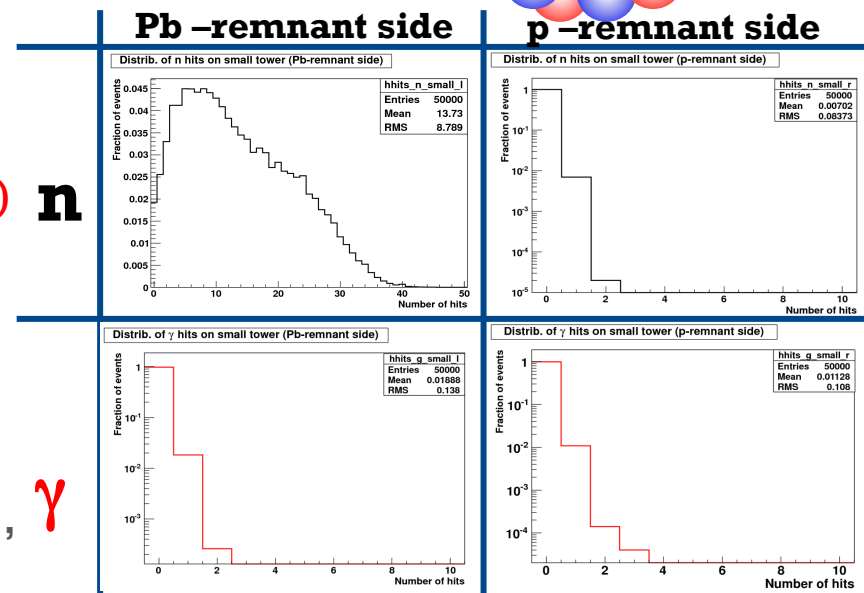


2013: p-Pb runs

- Interest in Ion runs
- Physics case study well motivated
- We will reinstall one ARM on p-remnant side during p-Pb run (beginning of 2013)** **n**

2014: 14 TeV p-p runs

- Necessary to complete the LHCf physics program
- The highest E_{lab} will be reached (10^{17} eV), **γ** to go closer to the VHECR region





Conclusions



- LHCf represent a very interesting and useful link between accelerator physics and cosmic ray physics
- The experiment has been carefully designed to precisely measure the highest energy LHC photons (up to 7 TeV), despite being the smallest LHC detector (few cm²)
- Nice physics results have been performed both on γ and π^0
- LHCf results have shown to be very useful for the very high energy cosmic rays models developer to improve their codes
- p/Pb run in 2013 and 14 TeV p/p run in 2014 are the next important steps