

VAN ALLEN BELT INSTABILITY AS MONITOR FOR EARTH SEISMICITY



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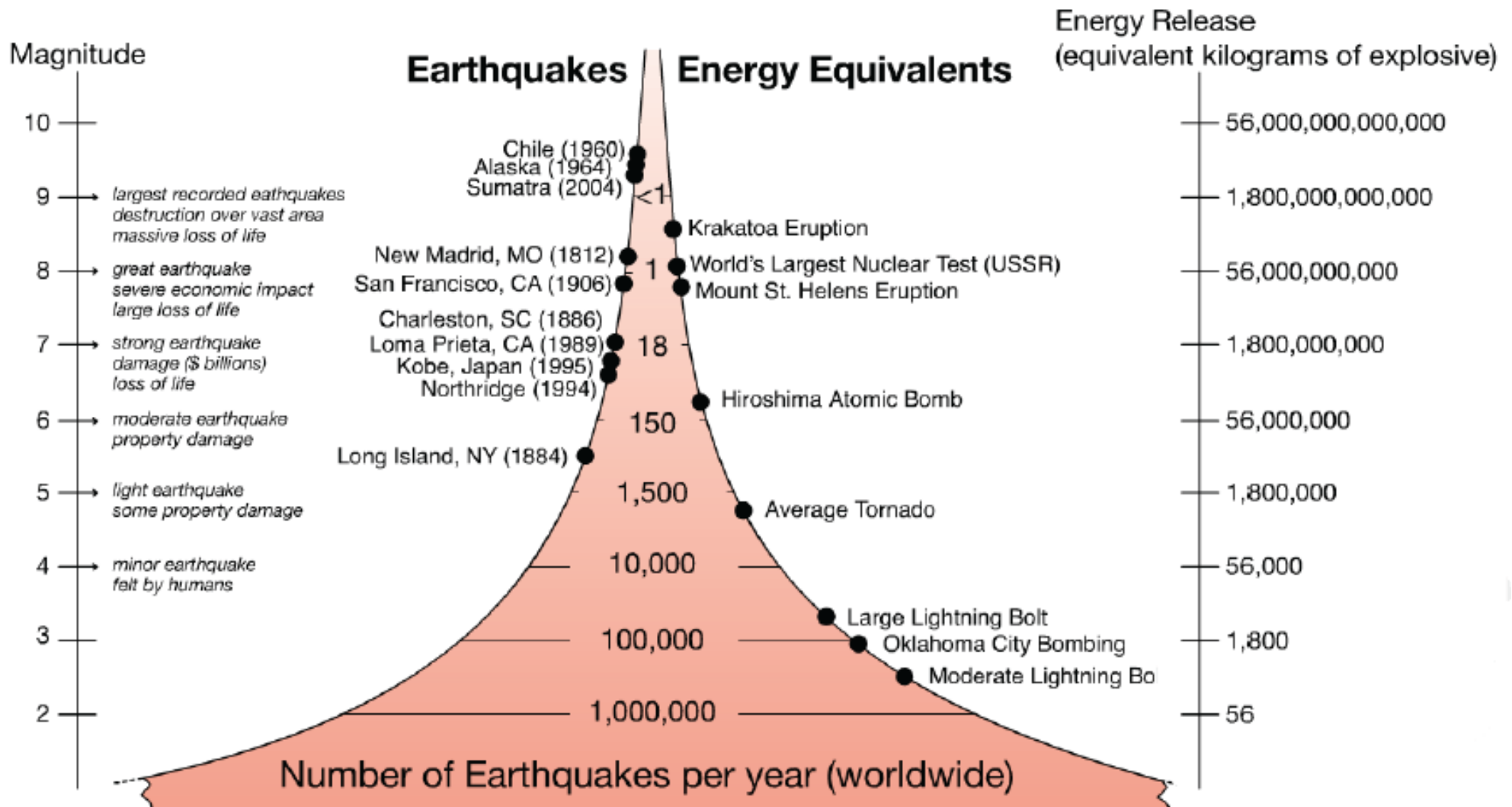
V. Vitale
F. Ambroglini
W. Burger
C. Fidani

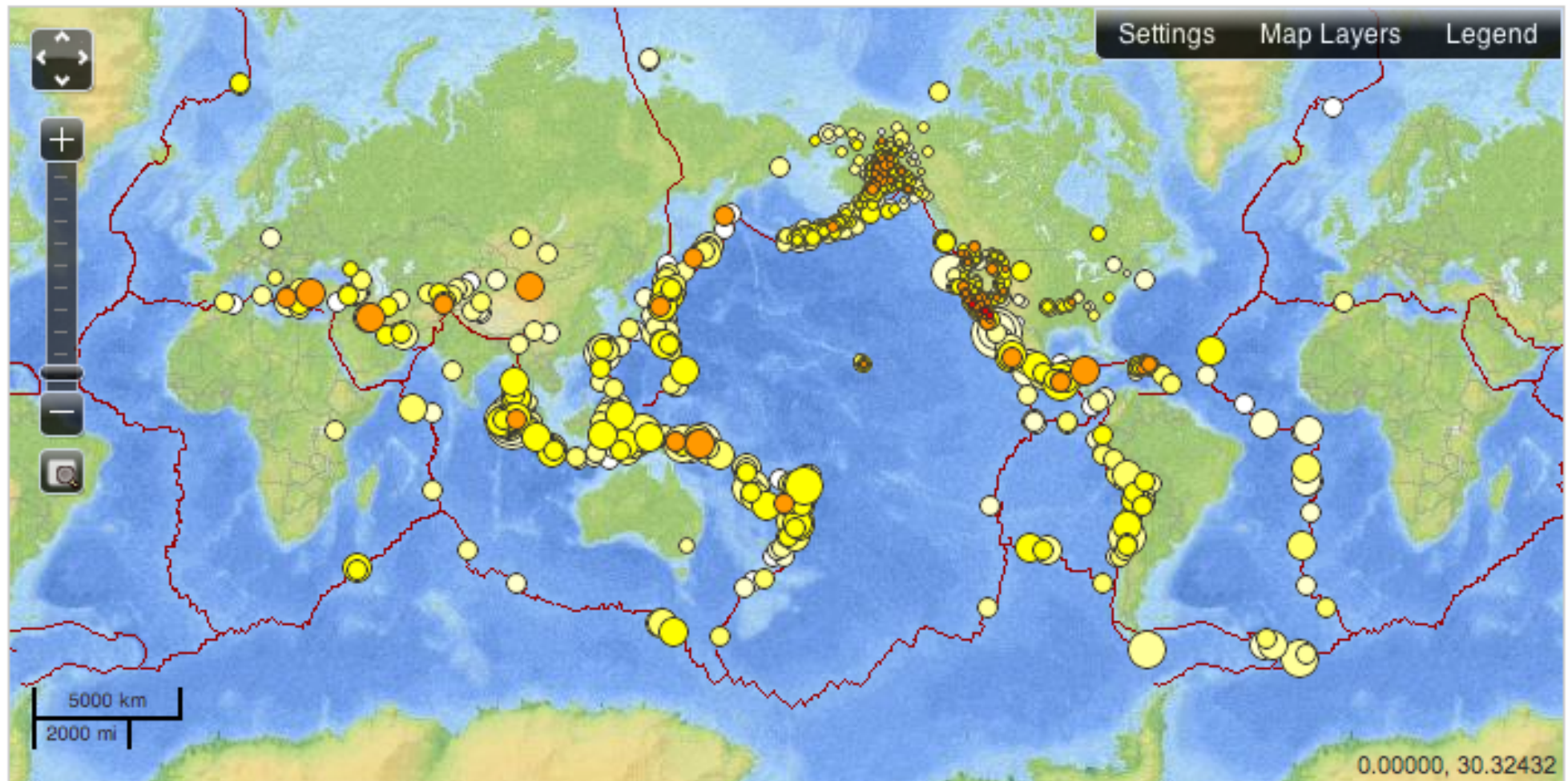
SPACEPART12 November 7, 2012, CERN

Introduction

Earthquakes are a very complex phenomenon, with a long preparation phase and involving measurable changes of physical parameters over huge surfaces together with large energy releases.

Energy Scale of Earthquakes





The Distribution Map of Seismic Belts

On each day there are about two Earthquakes with Magnitude $M > 5$

Every two day there is a $M > 6$ Earthquake

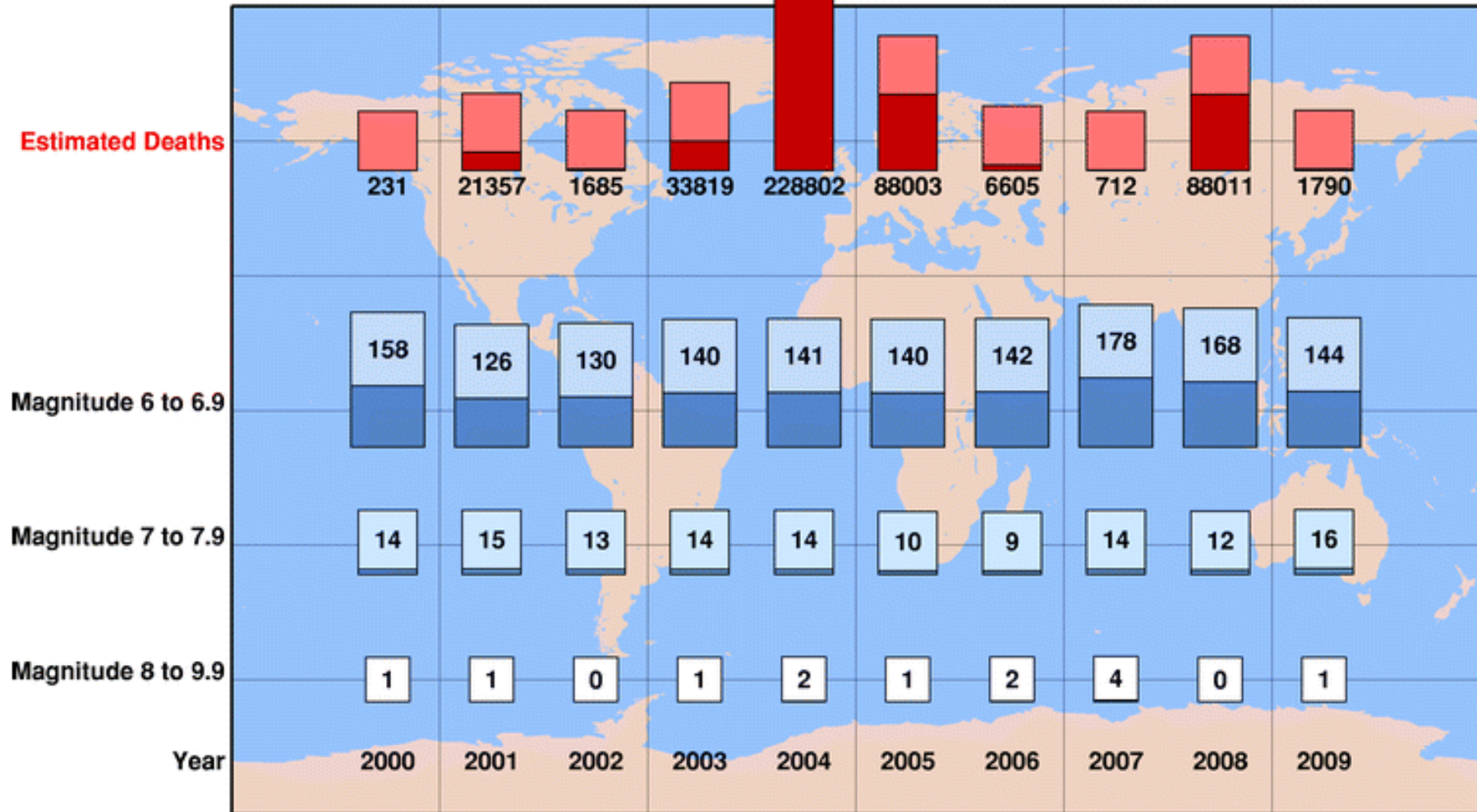
However due to the their very low rate, ground based, short term Earthquakes forecast efforts have been so far unsuccessful

Loss of lives is likely the most
damaging consequence of a
large Earthquakes

but

Earthquakes do not kill,
crumbling buildings do !

Worldwide Earthquakes: 2000 - 2009 *



2012 05 03

* Located by the USGS National Earthquake Information Center

On each day there are about two Earthquakes with Magnitude $M > 5$
 Every two day there is a $M > 6$ Earthquake

Space is a privileged point of observation

Worldwide on each day there are
2 Earthquakes with $M > 5$
every two days there is
1 Earthquake with $M > 6$

Large observational statistics can be
accumulated and studied

2) Earthquake-Particle Precipitation Correlations

a) NOAA data analysis

R. B. , V. Vitale 2012

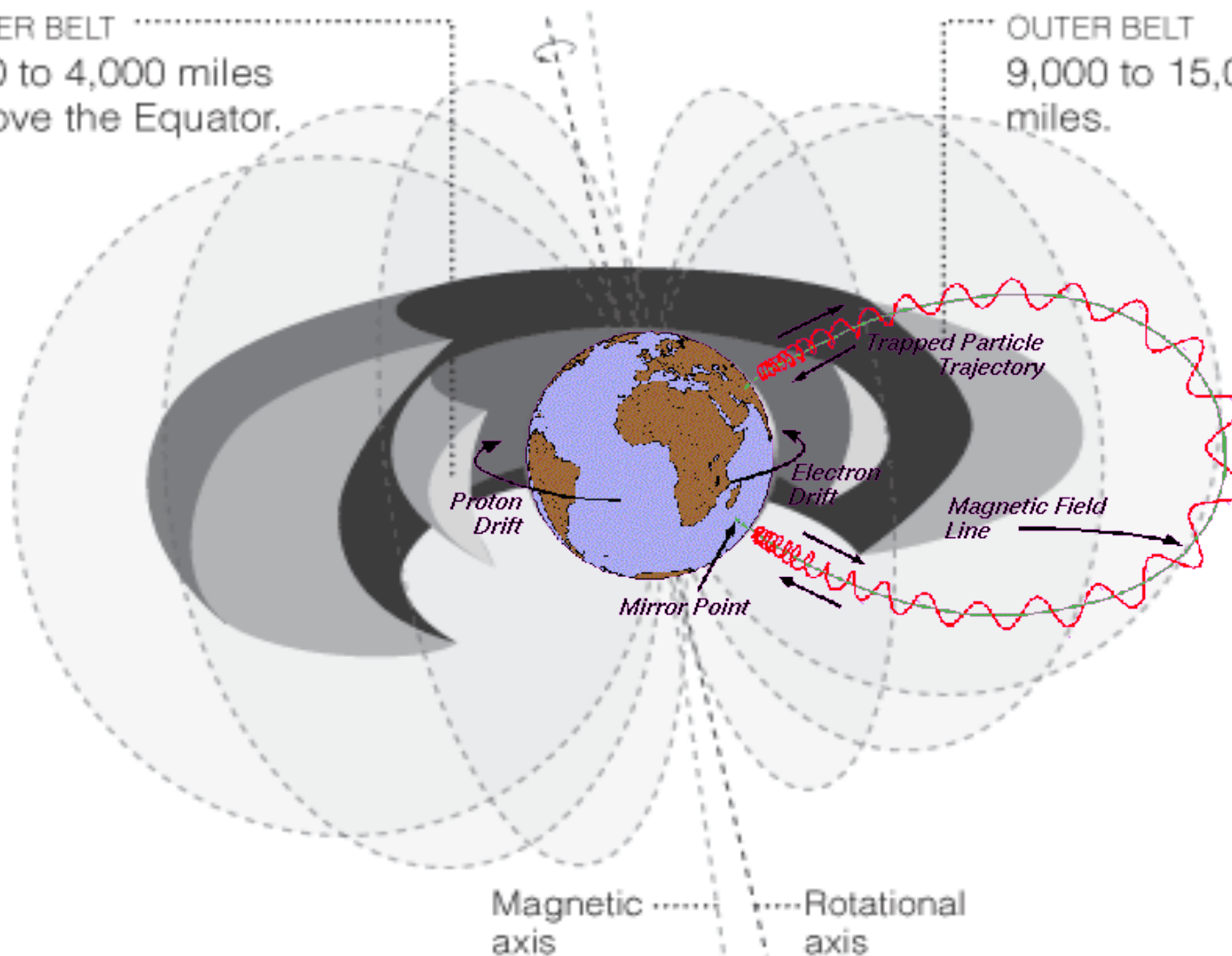
Analysis of Earthquake-Particle
Precipitation Correlations using
NOAA-15,16,17 and 18 data
(more than 14 years of data)

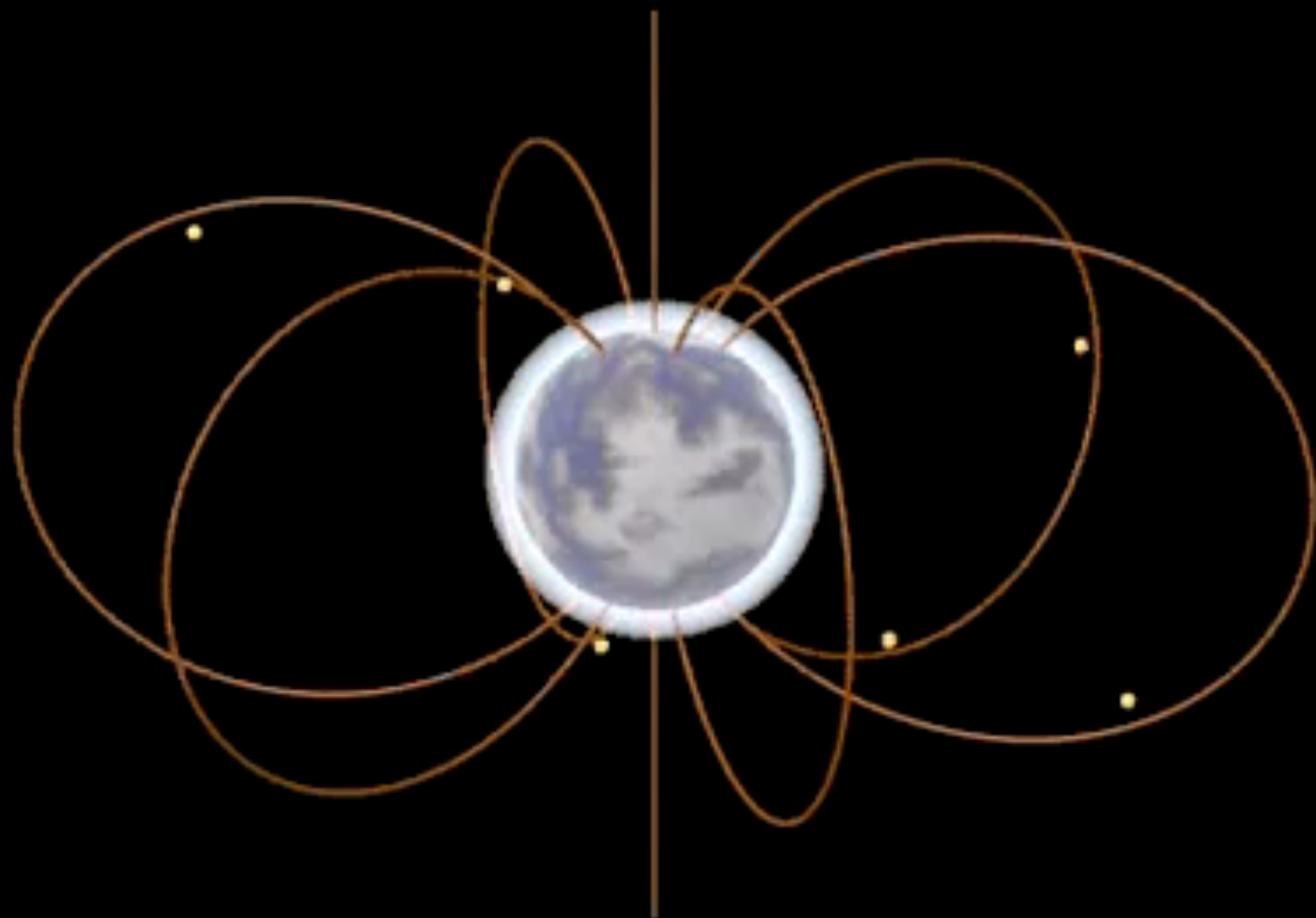
Van Allen Belts

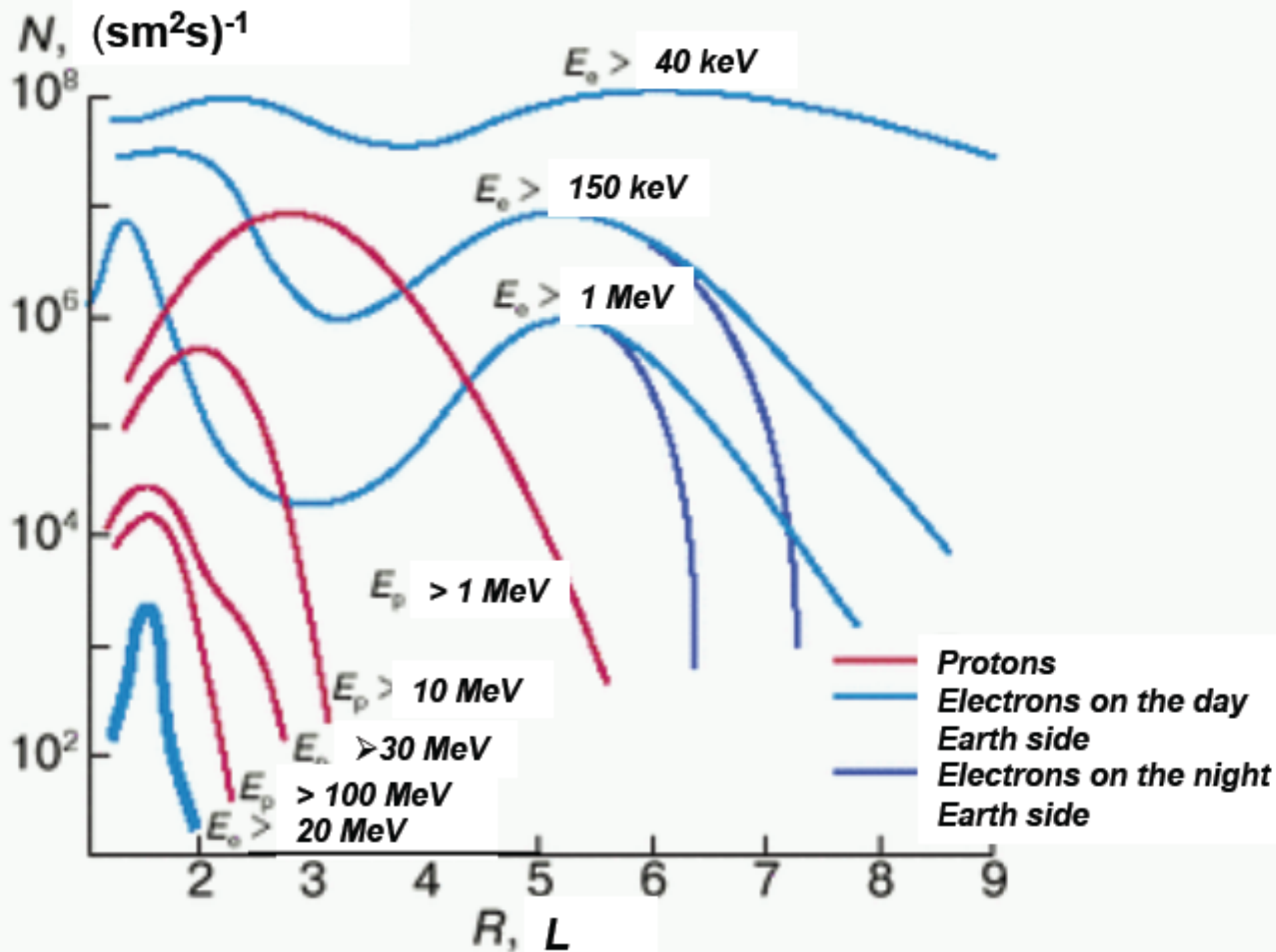
In 1958, Dr. James A. Van Allen, an American astrophysicist, discovered two belts of charged particles circling the planet, trapped by the Earth's magnetic field.

INNER BELT
400 to 4,000 miles
above the Equator.

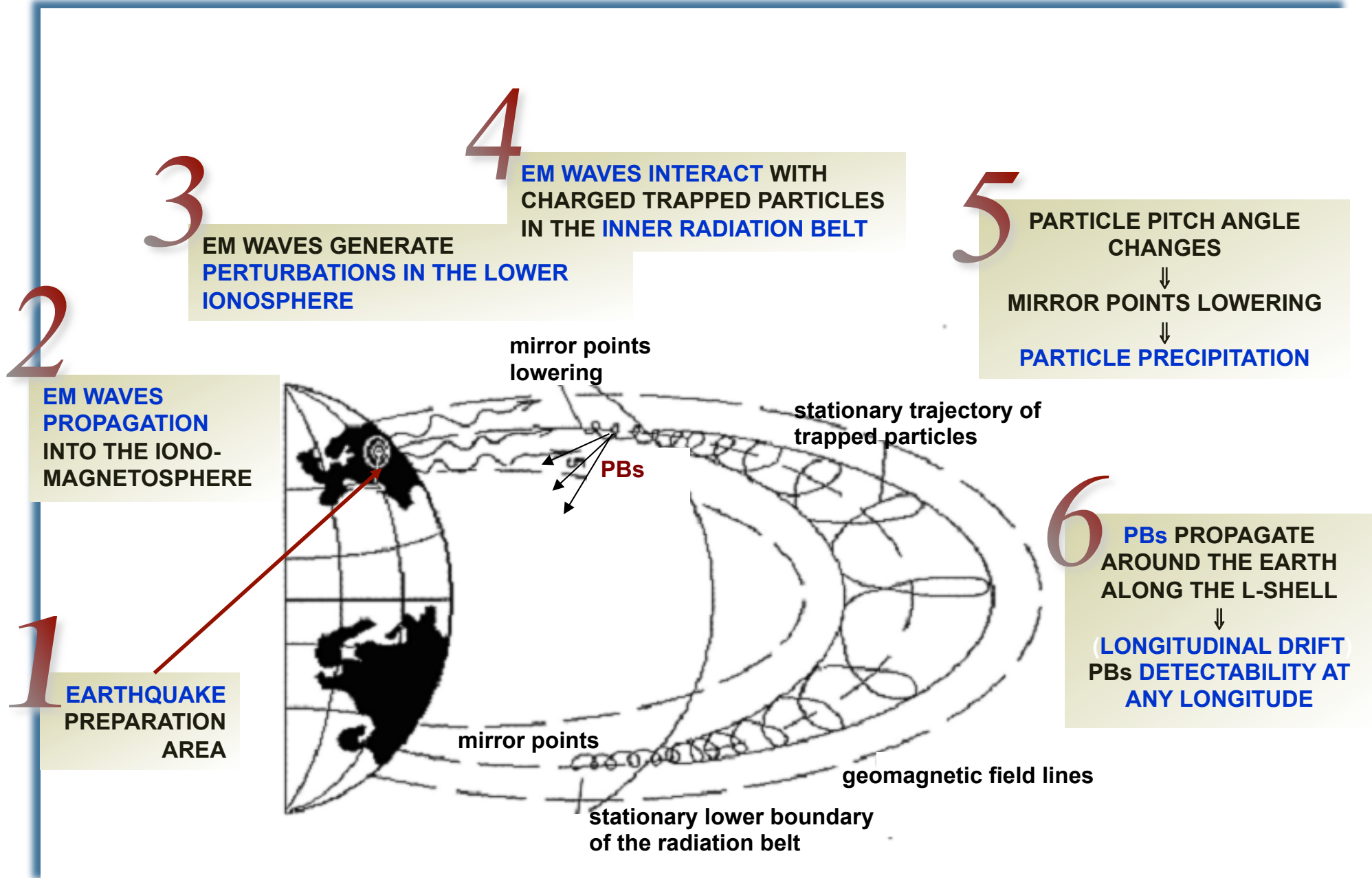
OUTER BELT
9,000 to 15,000
miles.







Schematic representation in a meridian plane of the trapped particle trajectories



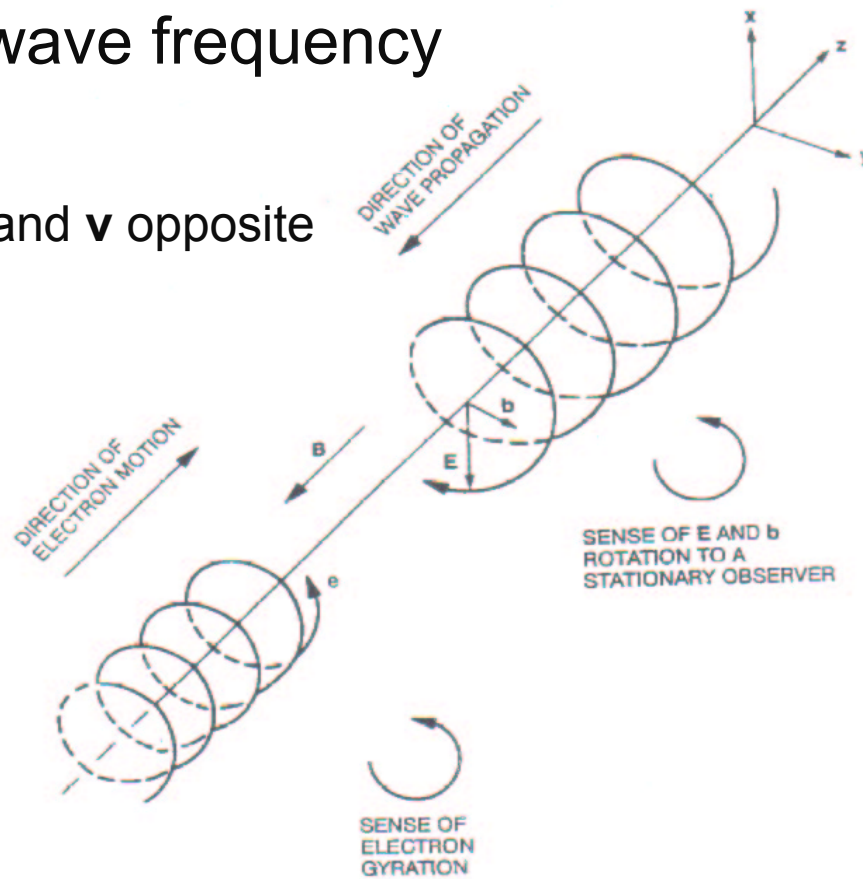
Wave-particle interaction

Cyclotron resonance - Whistler mode

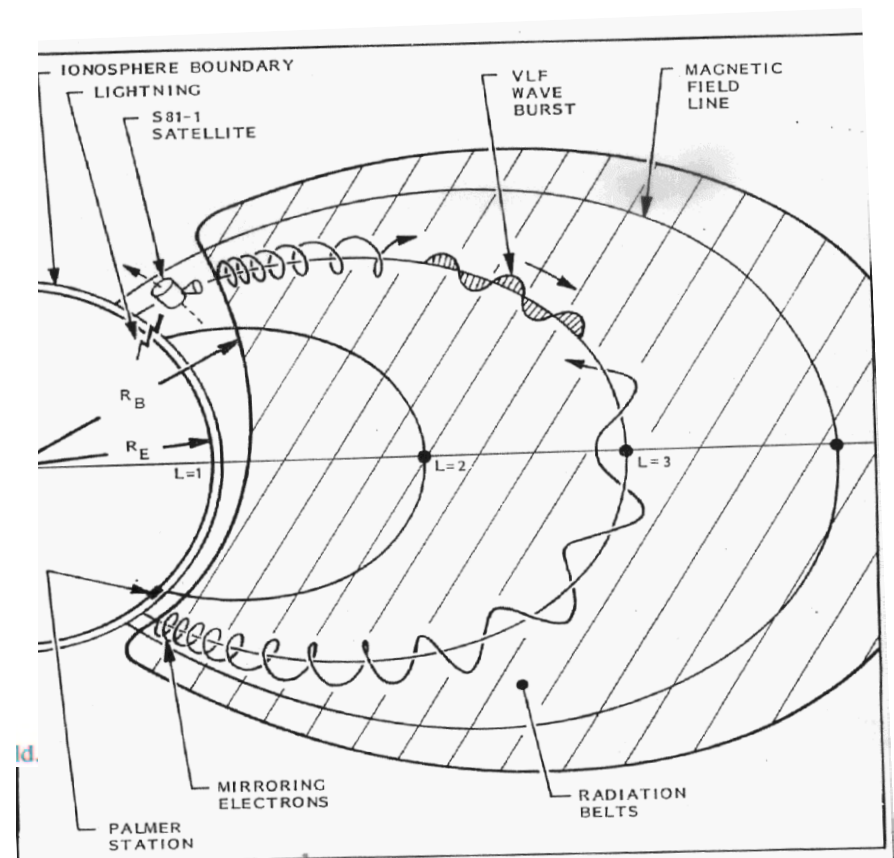
Cumulative deflection from many interaction with VLF (3-30 kHz period 10-45 s) circularly polarized waves would force electrons into the loss cone i.e pitch angle diffusion

Doppler-shifted wave frequency

k and v opposite



Process efficiency measured with VLF transmitters, viable but low (Imhof,82)



Wave-particle interaction

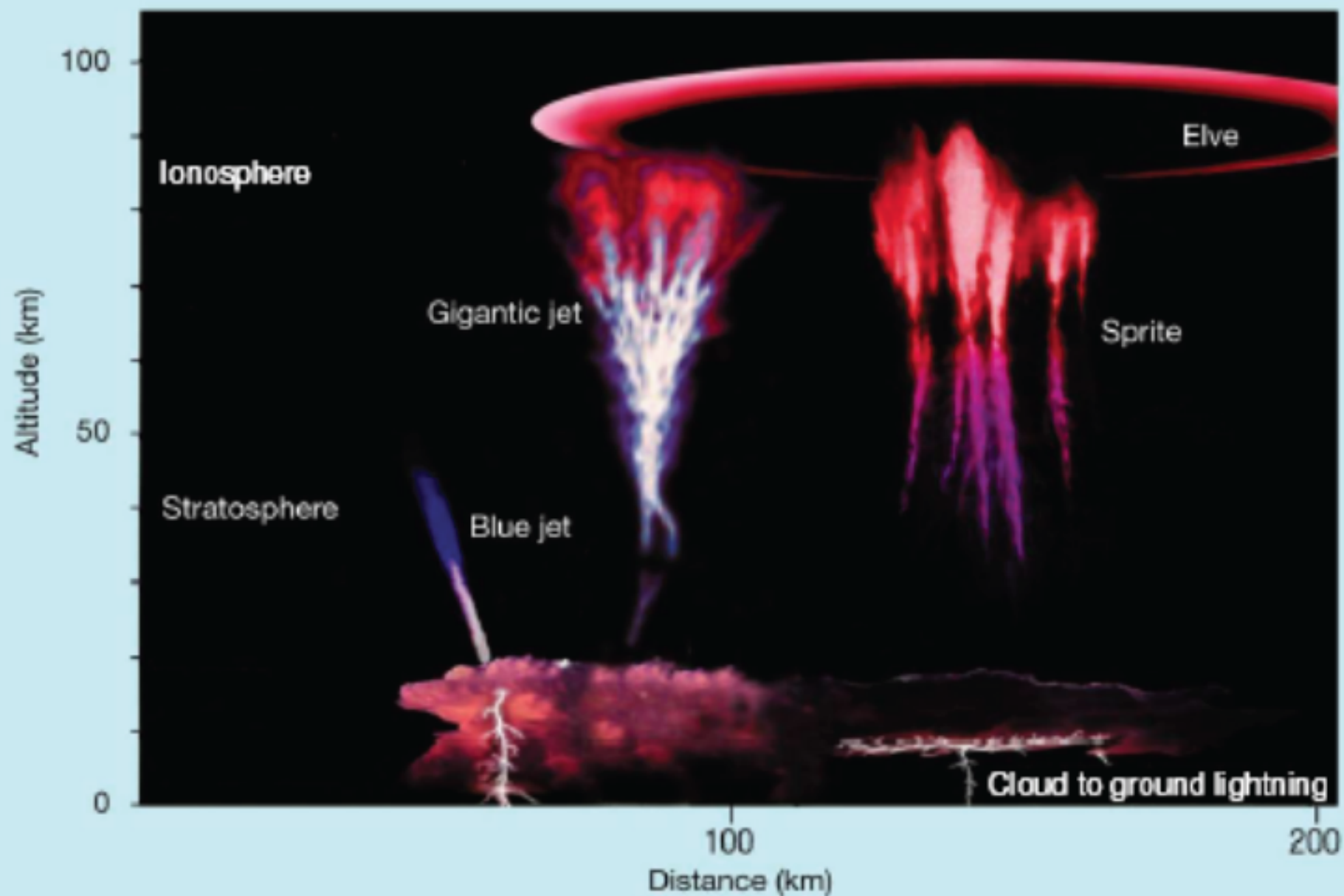
Bounce resonance - Alfvén mode

Magnetosonic micropulsation or electrostatic ULF waves (< 3Hz period 300 s) can interact resonantly with particles during their bouncing motion

drifting $d\lambda/dt = 5 \text{ deg/s}$ interacting with an ULF wave of field $E = 10^{-1} \text{ V/m}$ with typical extent of interaction region in longitude $\Delta\lambda = 12^\circ$ and in latitude $\delta\varphi = 5^\circ$ change in pitch angle $\Delta\alpha_{\text{bounce}} = 10^\circ$

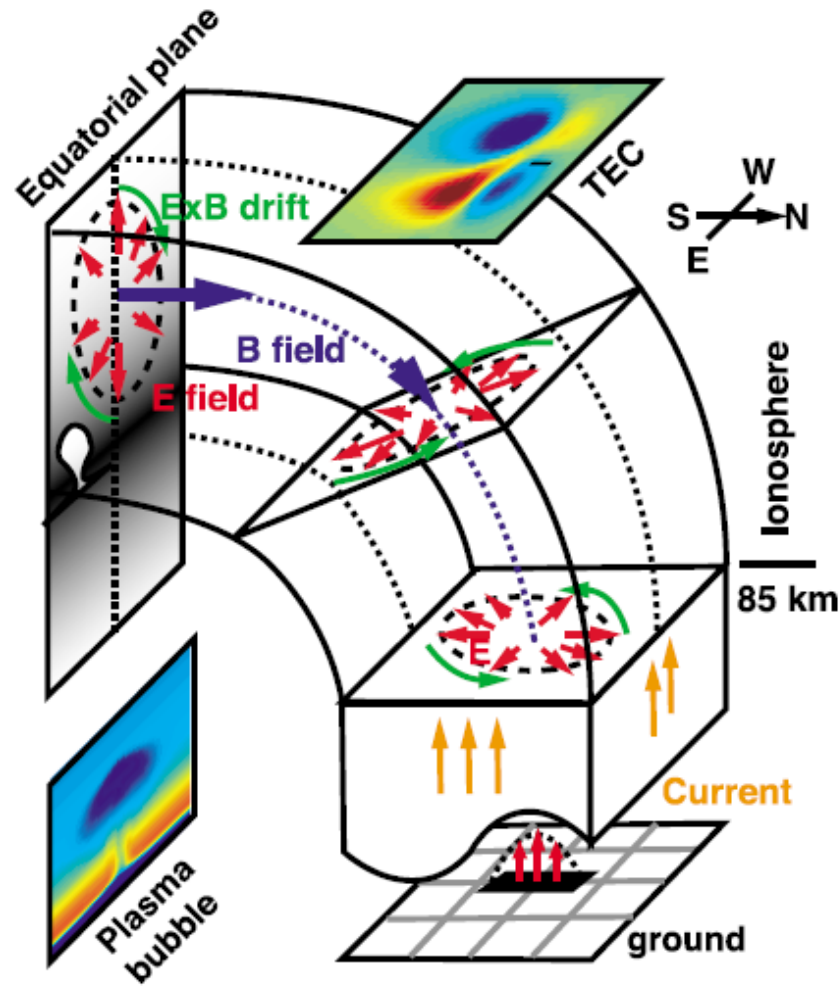
Process efficiency viable between ULF waves and protons of $E > 10 \text{ MeV}$ and electrons of $E > 1 \text{ MeV}$ (Aleshina 92)

High-altitude electric discharges



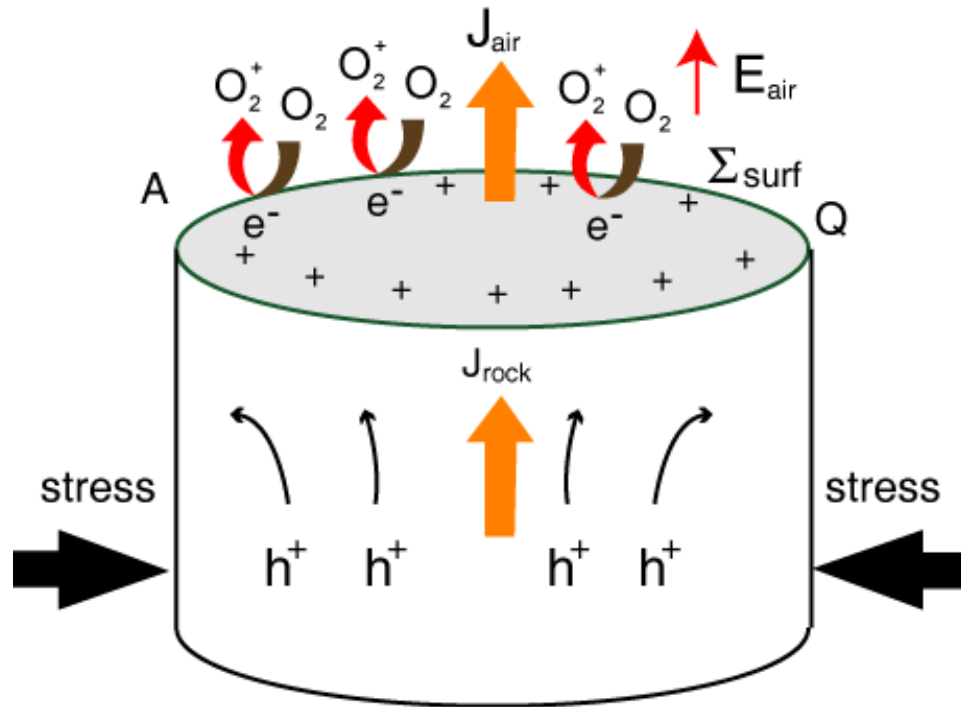
[Lyons et al., BAMS, 84, 445, 2003; Pasko, Nature, 423, 927, 2003]

Electric coupling between the ionosphere and surface charges



C. L. Kuo, J. D. Huba, G. Joyce, L. C. Lee, J. Geophys. Res., 116, A10317, 2011.

A Coupling Model Earthquake Region



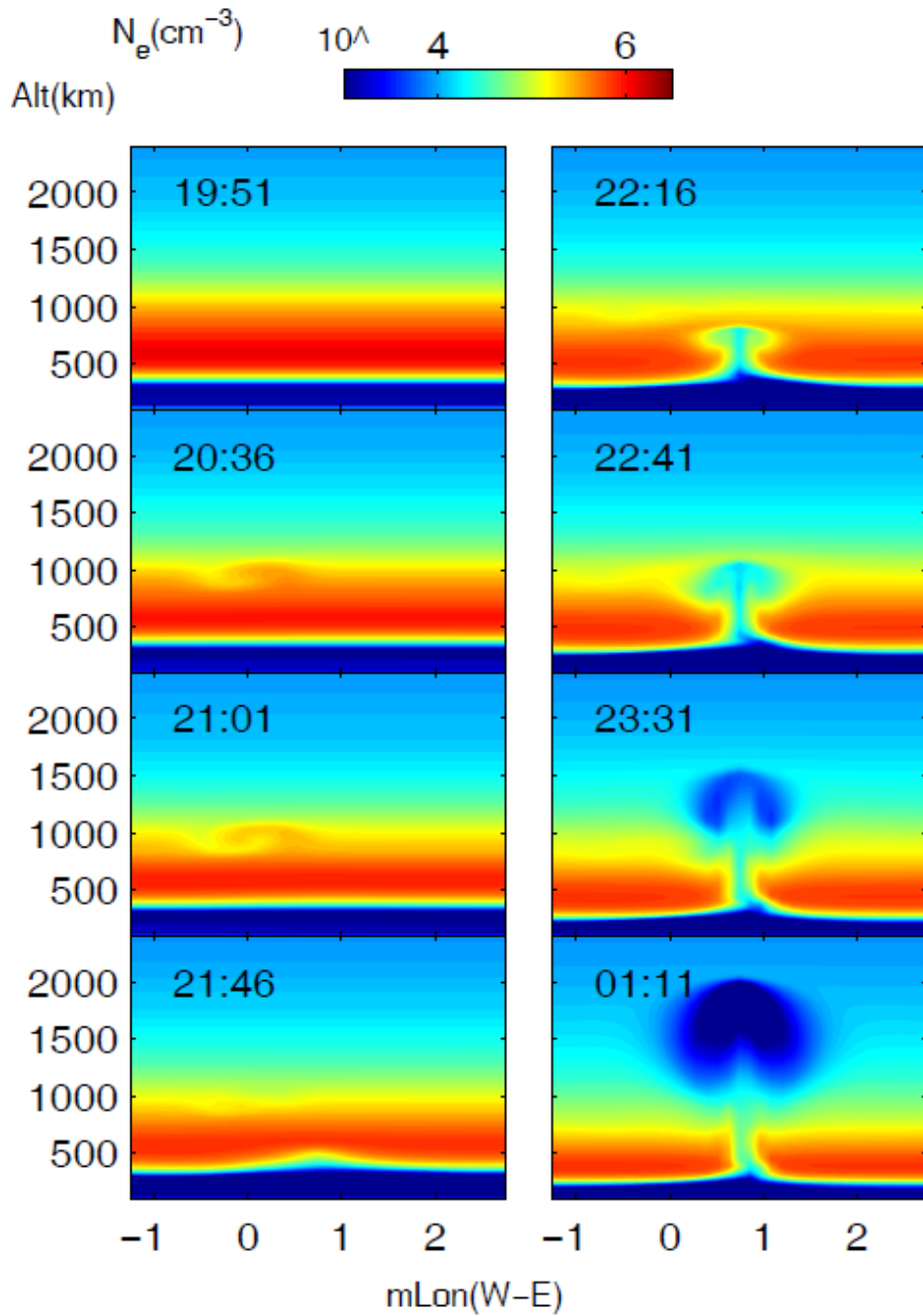
The surface charge density Σ_{surf} and total charge amount Q are accumulated by positive charges from the rock surface and O₂⁺ ions in the air. The upward electric field E_{air} associated with the positive surface charges will drive the upward current J_{air} .

$$E_{air} = \left(\Sigma_{surf} / \epsilon_0 \right) \hat{z} \quad J_{air} = \sigma_{air} E_{air}$$

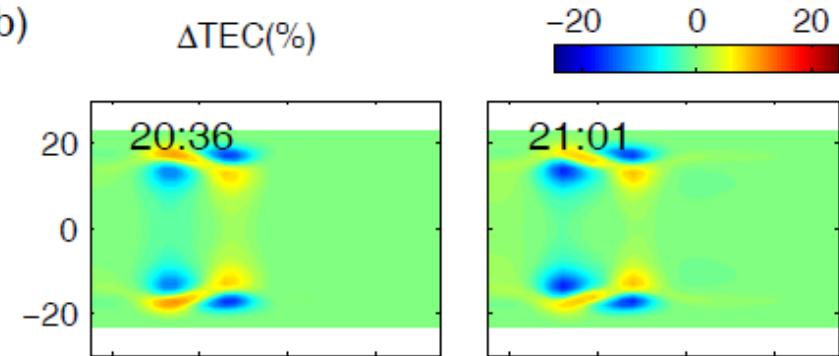
C. L. Kuo, J. D. Huba, G. Joyce, L. C. Lee, J. Geophys. Res., 116, A10317, 2011.

Nighttime case ($J_{rock} = 0.2 \mu\text{Am}^{-2}$)

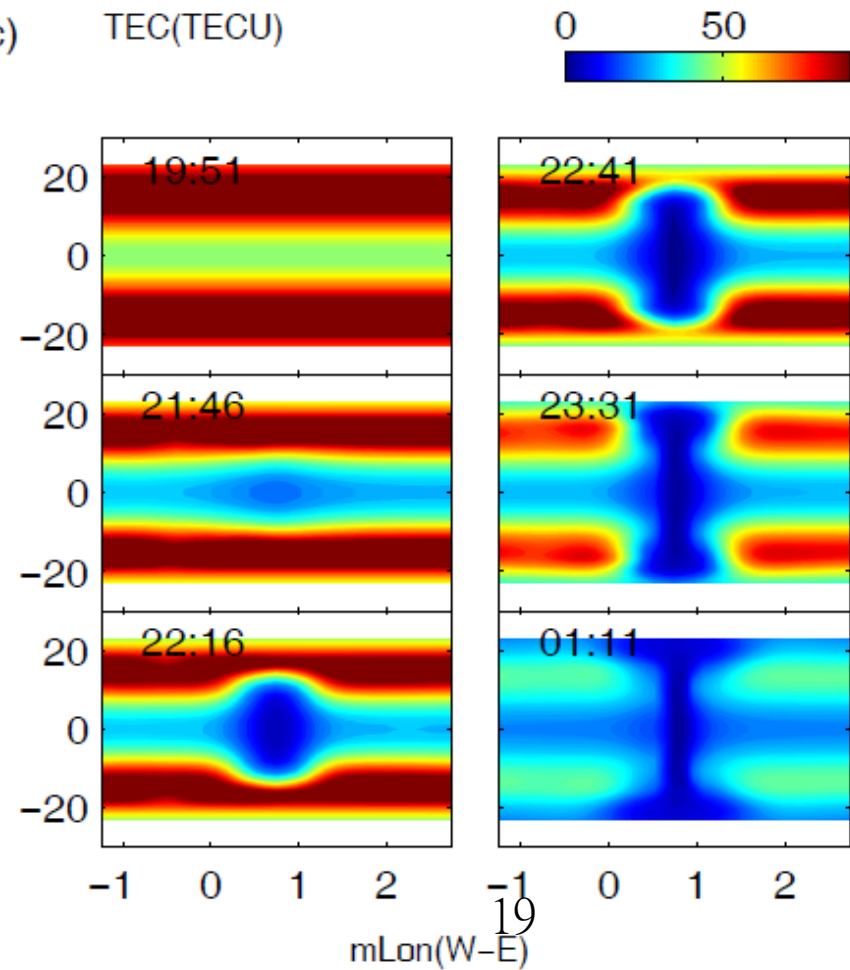
(a)



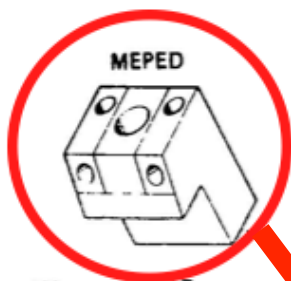
(b)



(c)



Electron and Proton
detectors



NOAA Satellites



NOAA-15 14.0
years

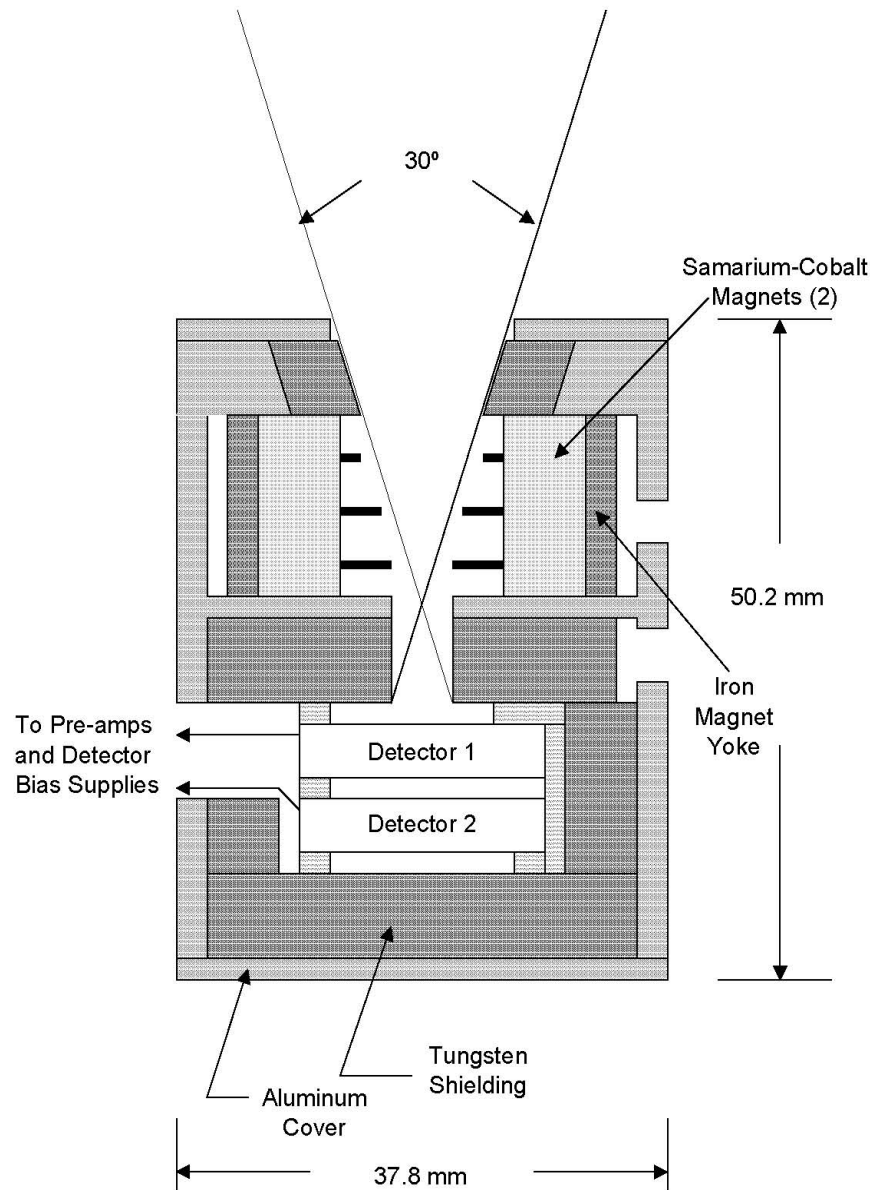
NOAA-16 12.5
years

NOAA-17 11.0
years

NOAA-18 8.0
years

NOAA-19 4.5
years

POES MEPED



- Telescope at 0deg oriented with central axis 9deg from direction $-X$ in $(-X, -Z)$ plane

- Telescope at 90deg rotated 9deg from $+Y$ direction in the YZ plane

- At geomagnetic latitude > 30 deg is it 0deg telescopes see particles in the loss cone while with 90deg telescope trapped ones

- 30deg Collimator

- Aluminum and tungsten foil avoid electrons with energies < 6 MeV e/o protons < 90 MeV to cross the detector

DETECTORS PARAMETERS

detectors channels	energy range keV	contaminants
e1	30-2500	210-2700keV protons
e2	100-2500	280-2700keV protons
e3	300-2500	440-2700keV protons
p1	30-80	
p2	80-240	
p3	240-800	
p4	800-2500	
p5	2500-6900	
p6	>6900	>700keV electrons
p6 _{omni}	>16000	>800keV electrons
p7 _{omni}	>36000	

Table 2: MEPED Telescopes Parameters. Electron detectors have energy channels identifier starting with *e*, protons detector with *p*. In the second columns the channel energy range is reported and in the third one the known contaminants. Data are taken from [14] and [16].

[14] Evans, D.S and M.S. Greer, (2000) Polar Orbiting Environmental Satellite Space Environment Monitor-2: Instrument Descriptions and Archive Data Documentation, NOAA Technical Memorandum OAR SEC-93, Boulder, CO.

[16] C.J. Rodger et al, 2010 Journal of Geophysical Research, 115, A04202

Particle Bursts

Particle Bursts are significant fluctuation
in the counting rate

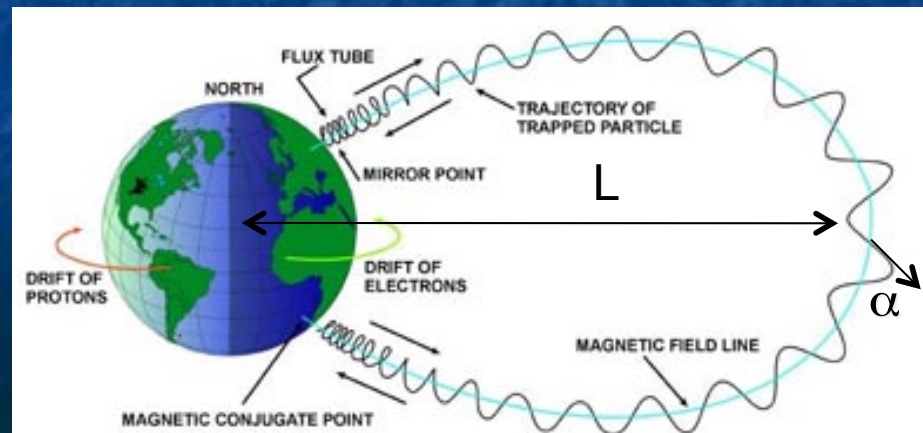
“significant” can be quantified by studying the counting rate statistics (7σ)

As the particle motion is strongly variable along the satellite orbit we choose to study the counting rate statistics in the invariant space (L, α, B)

L : L-shell

α : Pitch angle

B : Geomagnetic field



**Geomagnetic
field line**



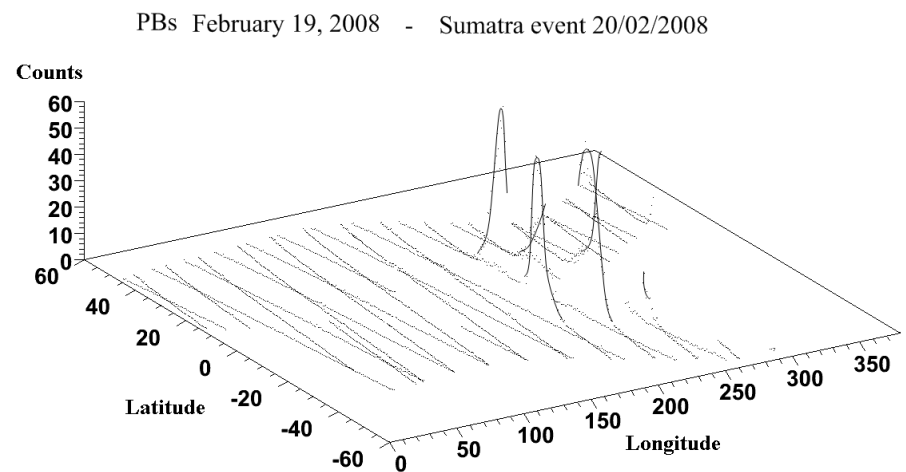
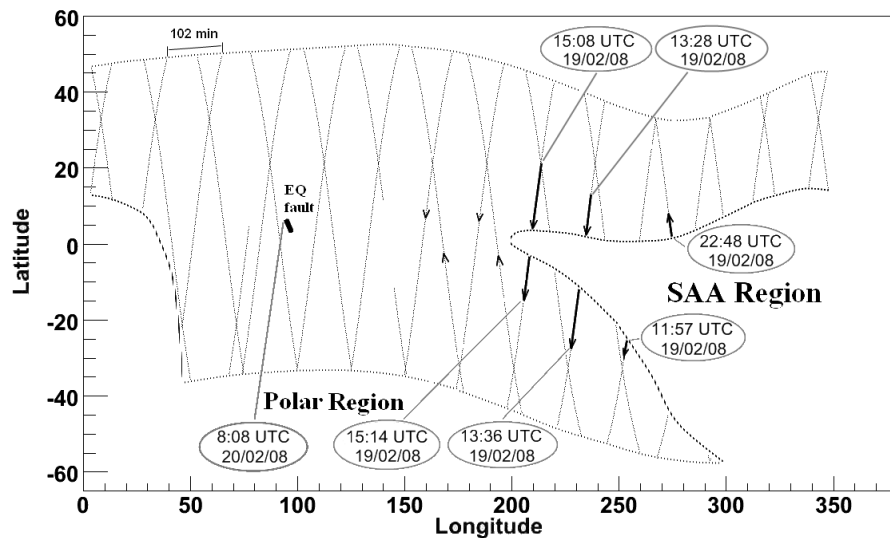
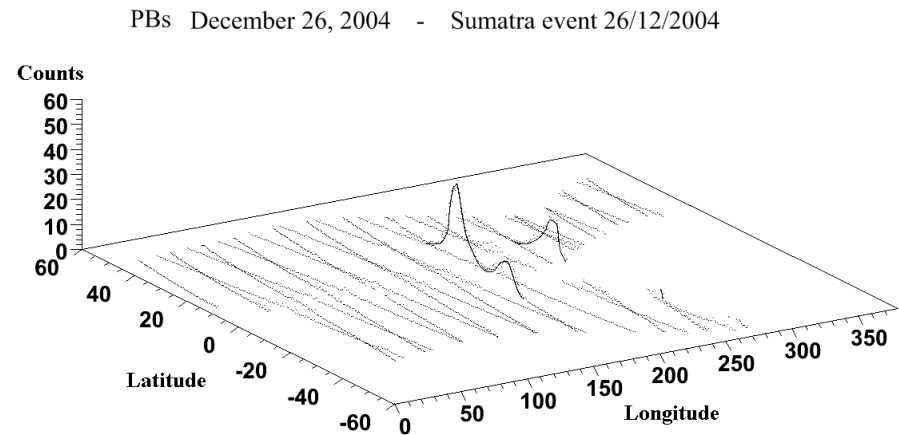
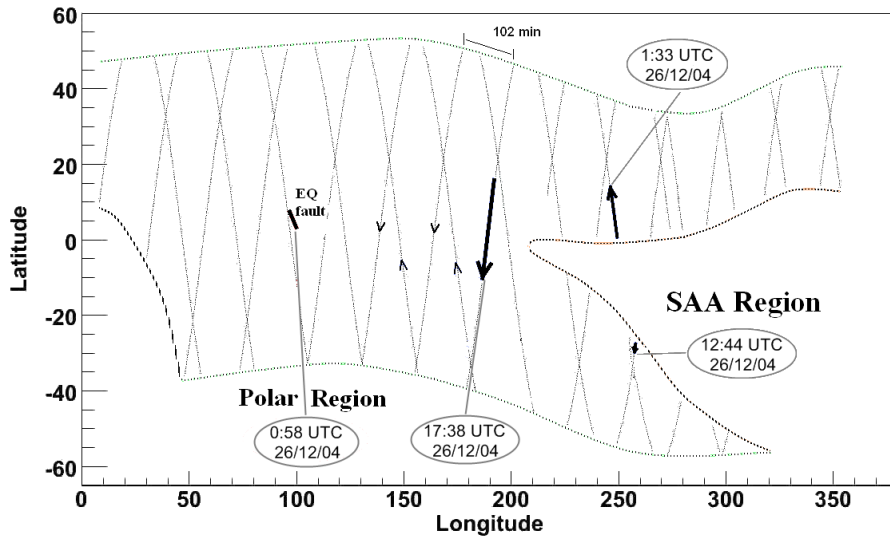
NOAA satellite

**Particle
trajectory**



Examples of Particle Bursts

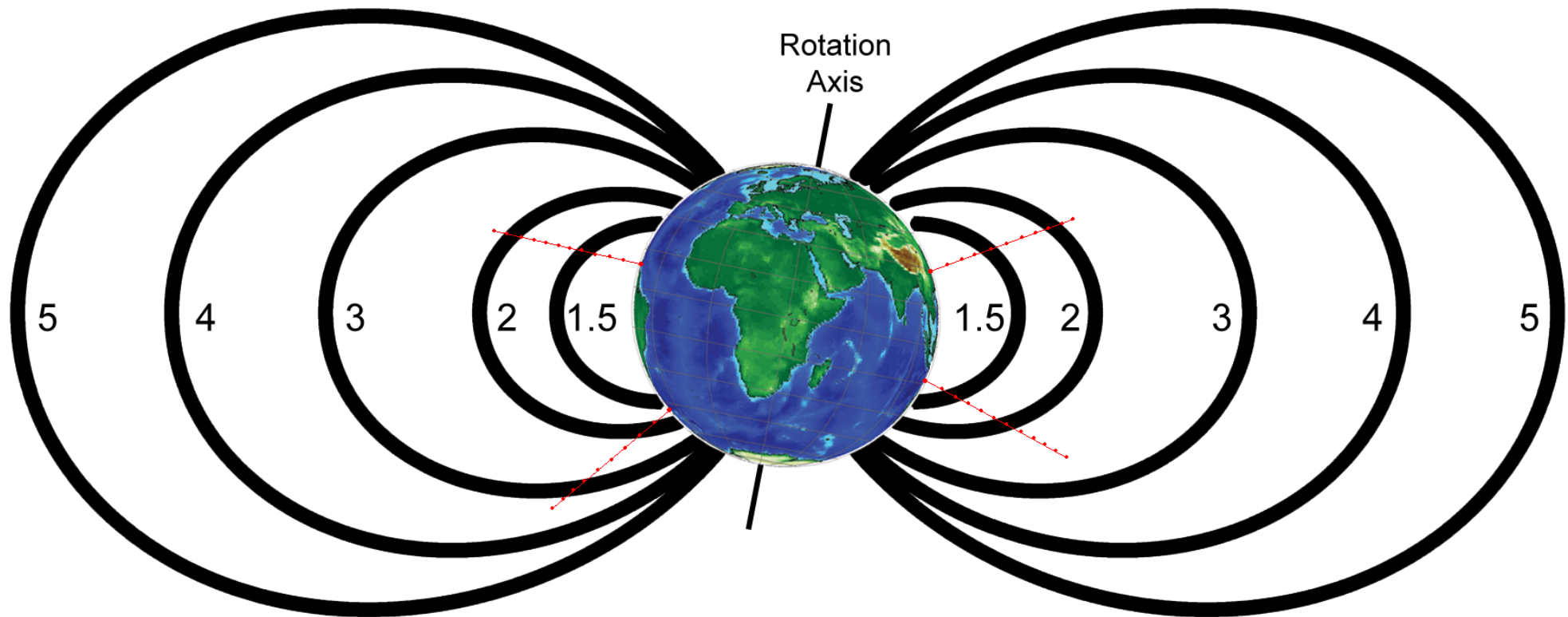
(C. Fidani and R.B., NHESS, 8, 1277-1291,2008)



Search for correlations with Earthquakes by first eliminating spurious effects

- 1) Solar influence (SID)
- 2) Earthquake clustering
- 3) SAA masking
- 4) Particle Burst merging

Correlation with the earthquakes



$L_{\text{Earthquake}}$: L -shell where earthquake interact with the Particle Burst

For a given a projection altitude we considered only the Earthquakes with :

$$|L_{\text{Earthquake}} - L_{\text{ParticleBurst}}| < 0.1$$

(Aleksandrin et al., 2003; Sgrigna et al. 2005)

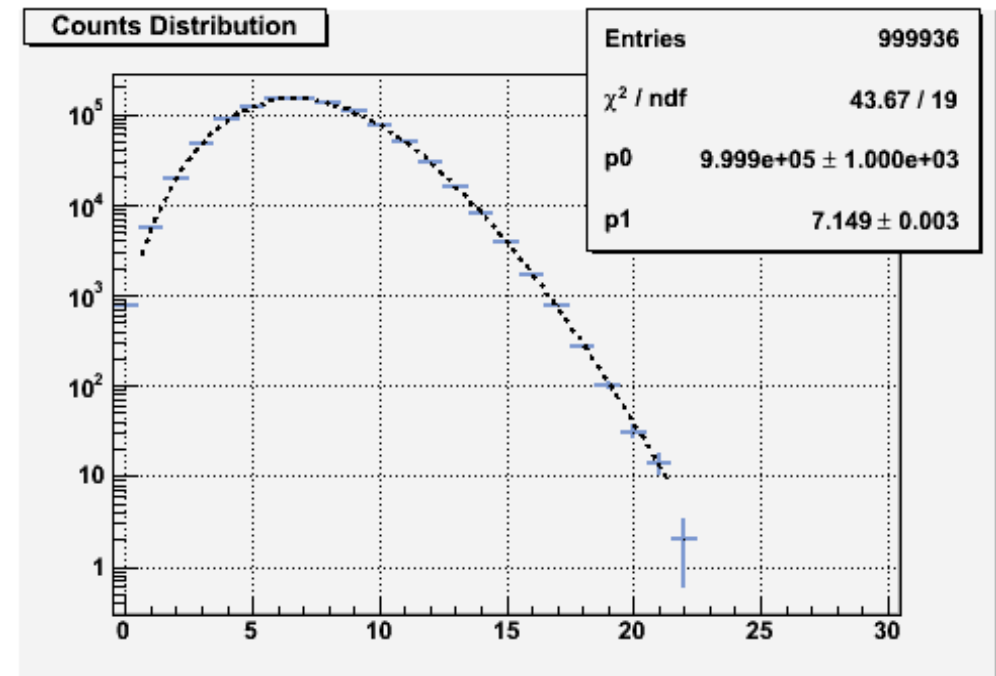
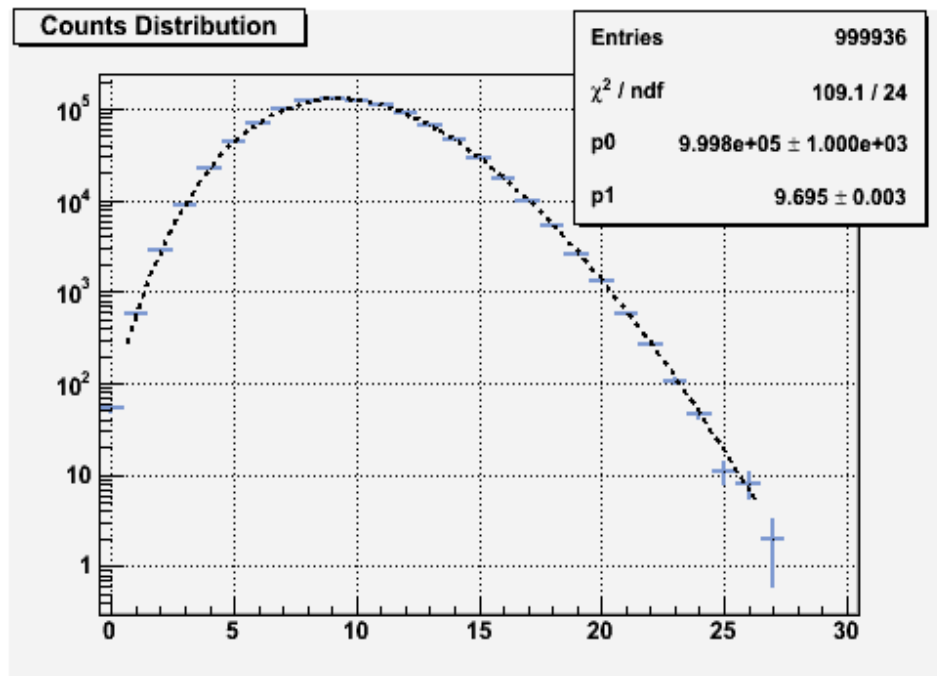
SINGLE SATELLITES RESULTS

Satellite	C	μ	$(C-\mu)/\sqrt{\mu}$
15	7	3.25 ± 0.19	2.08
16	7	2.00 ± 0.16	3.53
17	7	2.21 ± 0.13	3.23
18	4	0.98 ± 0.12	3.05

Table 5: Contributions of each satellite to the found excess. Here are reported identification number of the satellite in use, the number of counts in -1.25 hours bin (C), the mean value of in the Δt distribution (μ) and significance of the counts in the -1.25 hours bin $(S-M)/\sqrt{\mu}$.

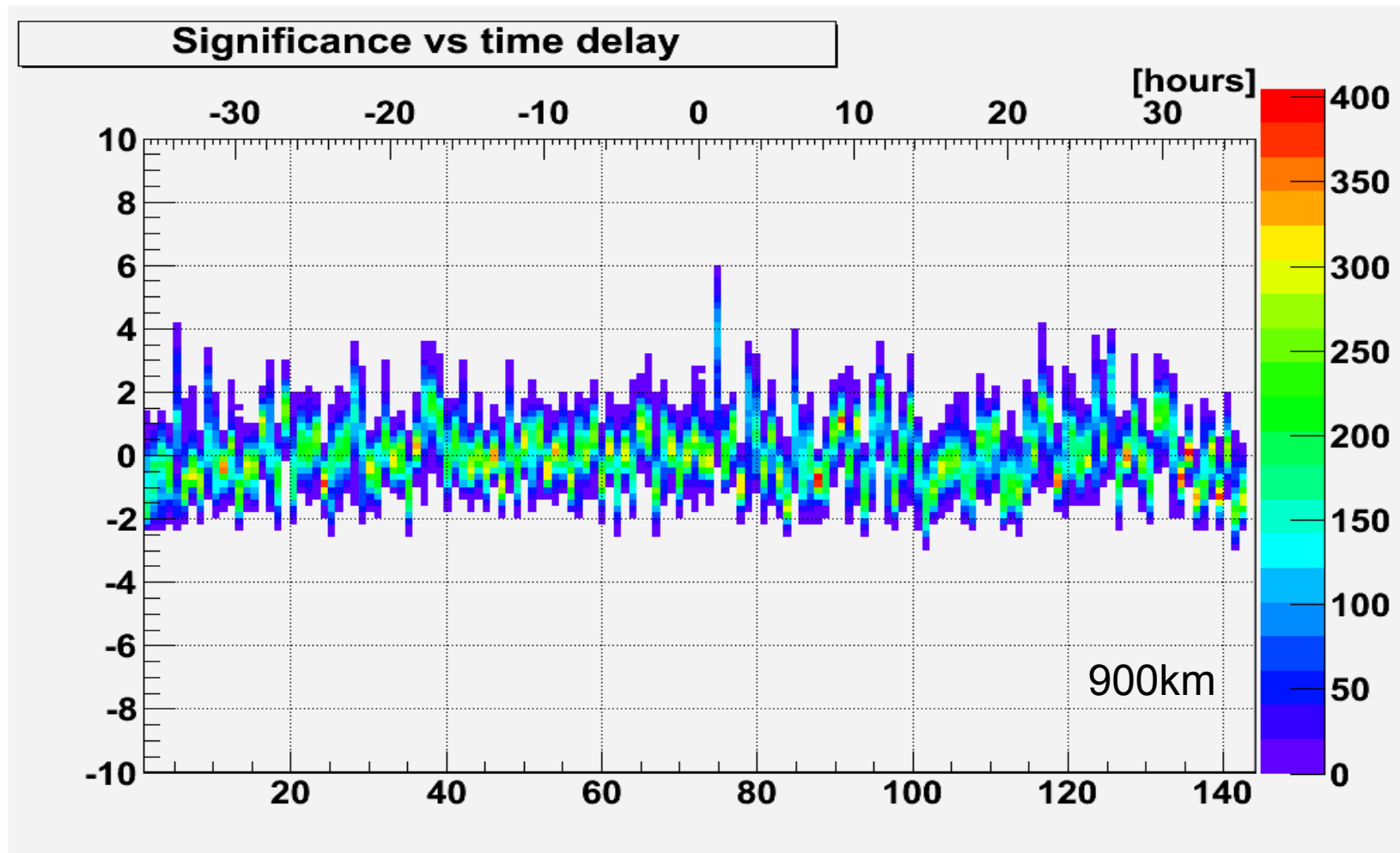
- All the data samples (satellites) contribute to the excess

Statistical significance



- Can random samples of EQs generate the found excess ?
- The analysis was repeated 2 X 6944 times with simulated EQs
- For each repetition approx 18k Earthquakes are simulated
- For test 1 only the time of EQs is simulated randomly within the relevant time range
- For test 2 time latitude and longitude are simulated
- Test 1 (2) reaches a maximum of 27(22) counts, eq a 5.5 s.d. event, in 1 million trials
- No systematic deviation from the expected Poisson statistics is found

Stability analysis



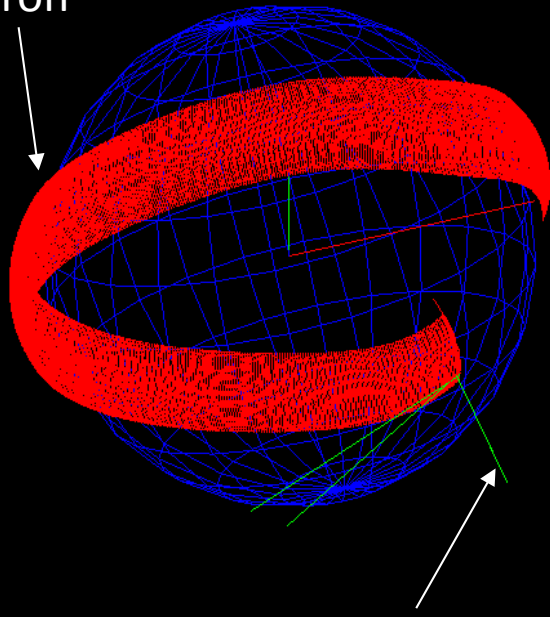
- Significances reached in each delay channel, with the 1715 analysis repetition, with P5=900km. Channel 75 @ $t = -1.25$ hours

SPES: PLANETOCOSMICS

Example of particle propagation with IGRF magnetic field and NRLMSISE00 atmosphere model

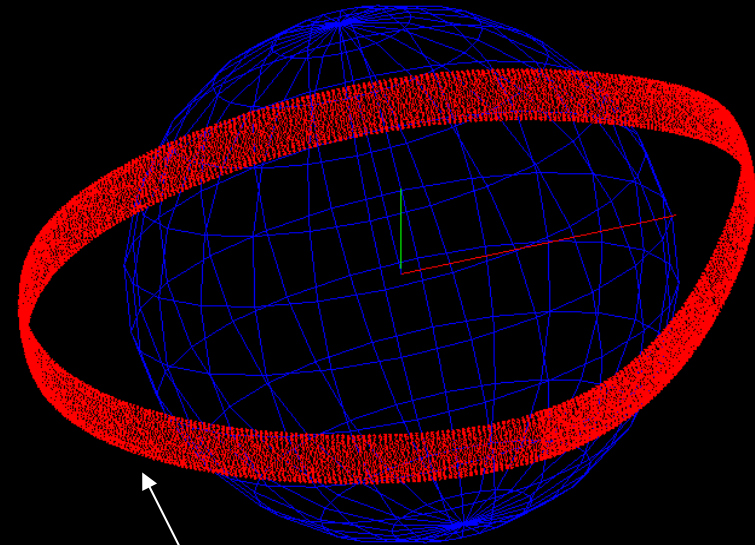
e- 100 MeV Lat 0° Long 0° Altitude 500 km

electron



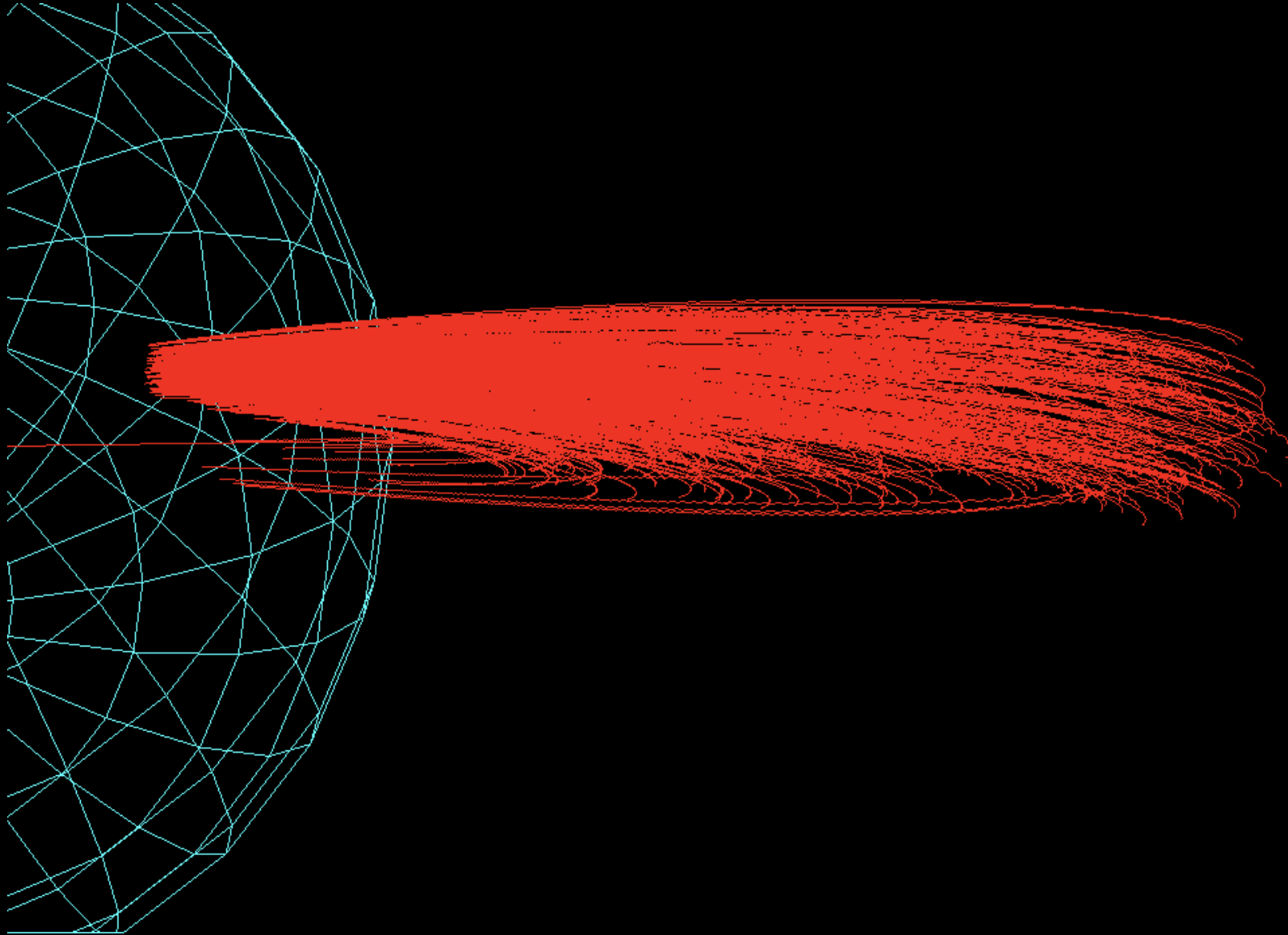
gamma

e- 100 MeV Lat 0° Long 0° Altitude 2000 km



electron

SPES events



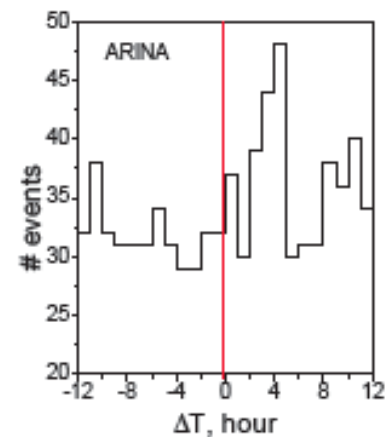
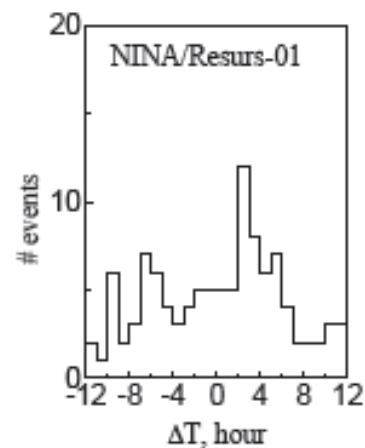
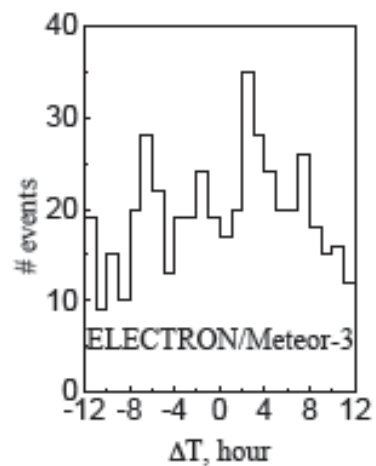
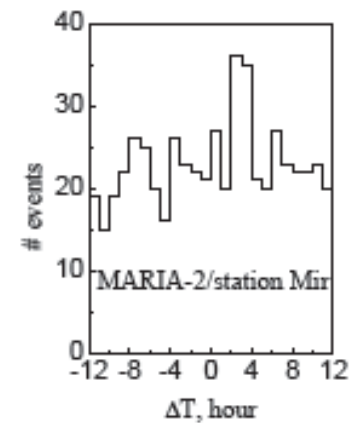
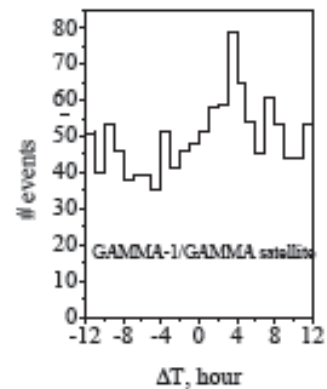
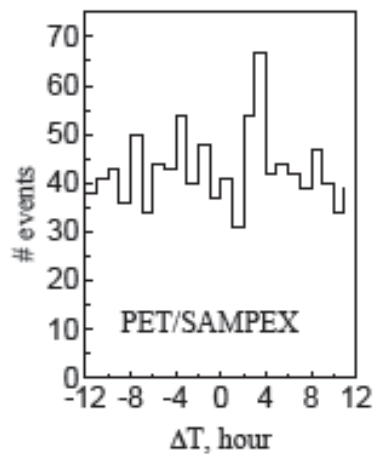
Result for NOAA analysis

We established a 1,25 hours correlation (precursor) between electron Particle Bursts and Earthquakes. For the low energies covered by the NOAA detectors ($E \sim 300$ keV), a 5,7 sigma correlation is observed, concentrated on a small region near the South Atlantic Anomaly.

The correlation appears when integrating 14 years of NOAA 15, 16, 17 and 18 data, and involve 25 Earthquakes with $M > 5$.

**Correlations have been reported by
different space experiments**

**INTERKOSMOS-BULGARIA-1300,
METEOR-03, MIR, GAMMA, RESOURCE 01,
RESOURCE DSK, SAMPEX, NOAA**



S.V.Aleksandrin, A.M.Galper, S.V.Koldashov et al. *Annales Geophysical*, 2003, 21, 597.

These detectors, however, were not optimized for Earthquakes monitoring, had limited sensitivity and were not able to exploit all the informations available to determine the Earthquakes geographical position

An optimized Particle Spectrometer for Earthquake precursors localization

R. B., W. Burger, et. Al. NIM A, 617, 467, 2009

Viewing angle : zenith

Energy range: <50 MeV (e) , <300 MeV (p)

Single particle detection → backtracking for E.Q. localization

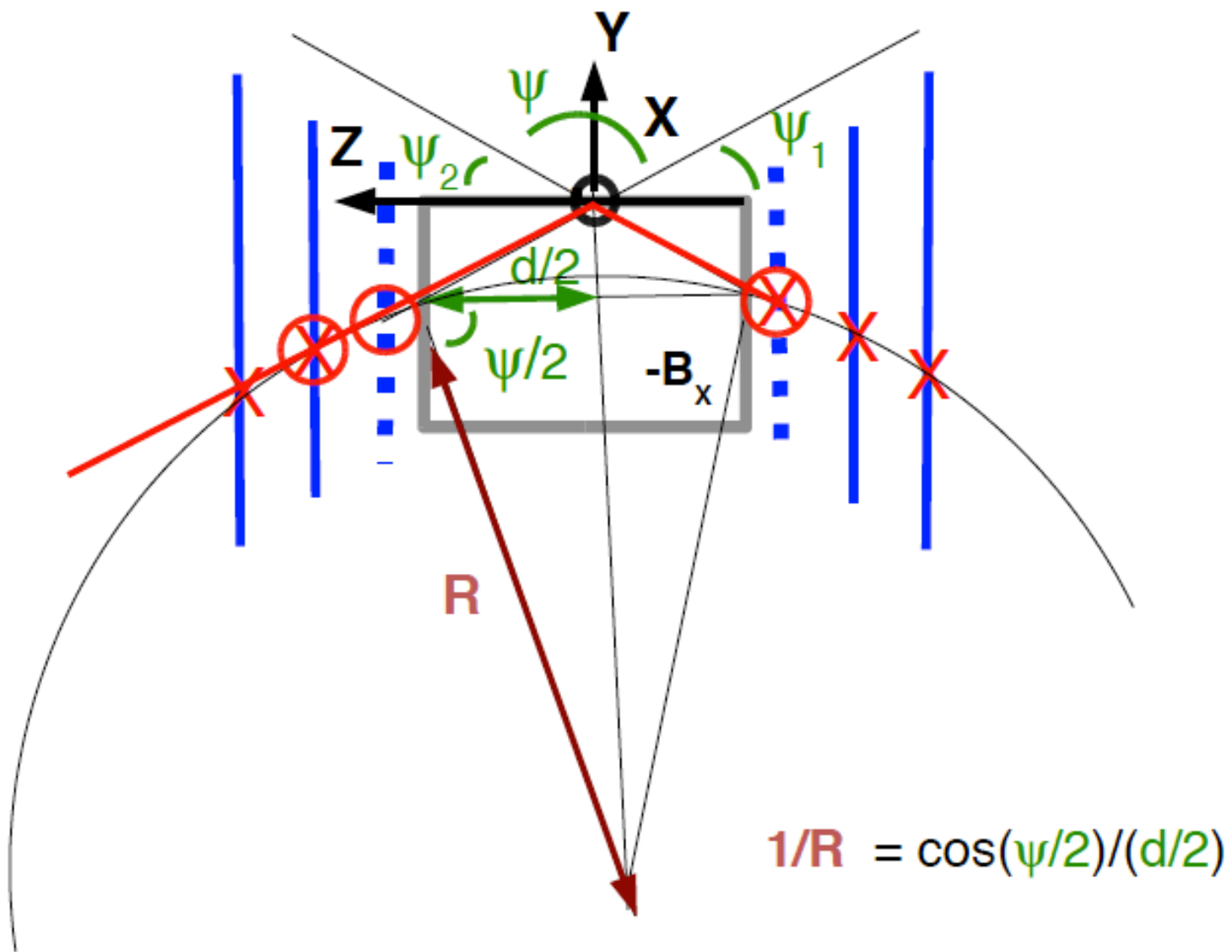
Pitch angle resolution : < 10 degree;

Acceptance, Energy resolution: the best possible

Energy vs Localization dependance of precipitating particles based on existing experience

NOAA: 0.3 MeV Correlation strongly orbit dependent;

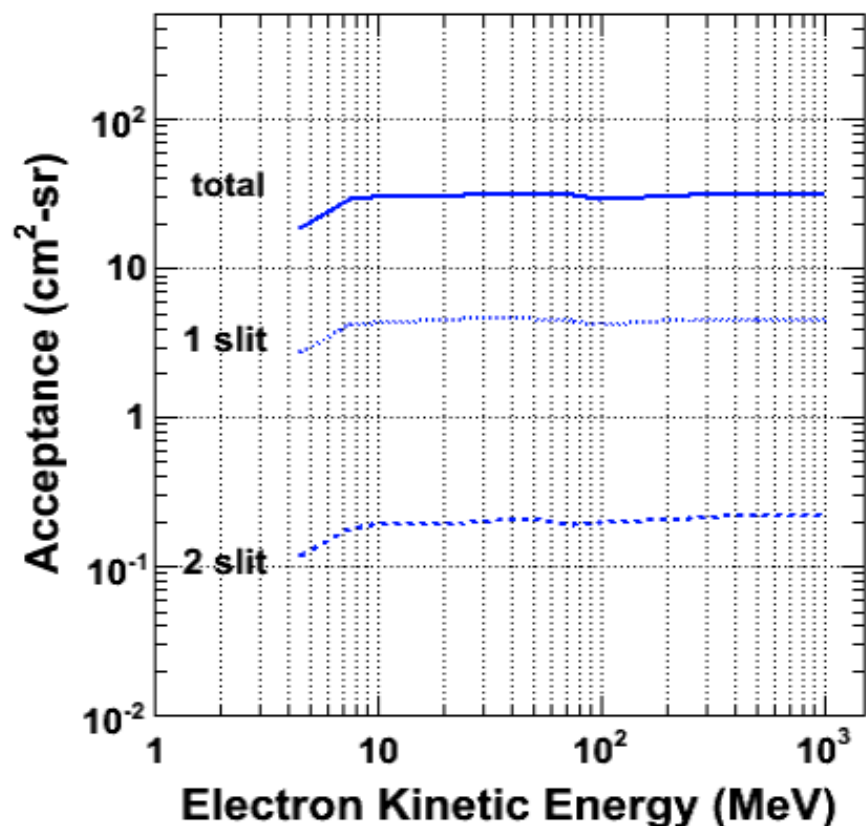
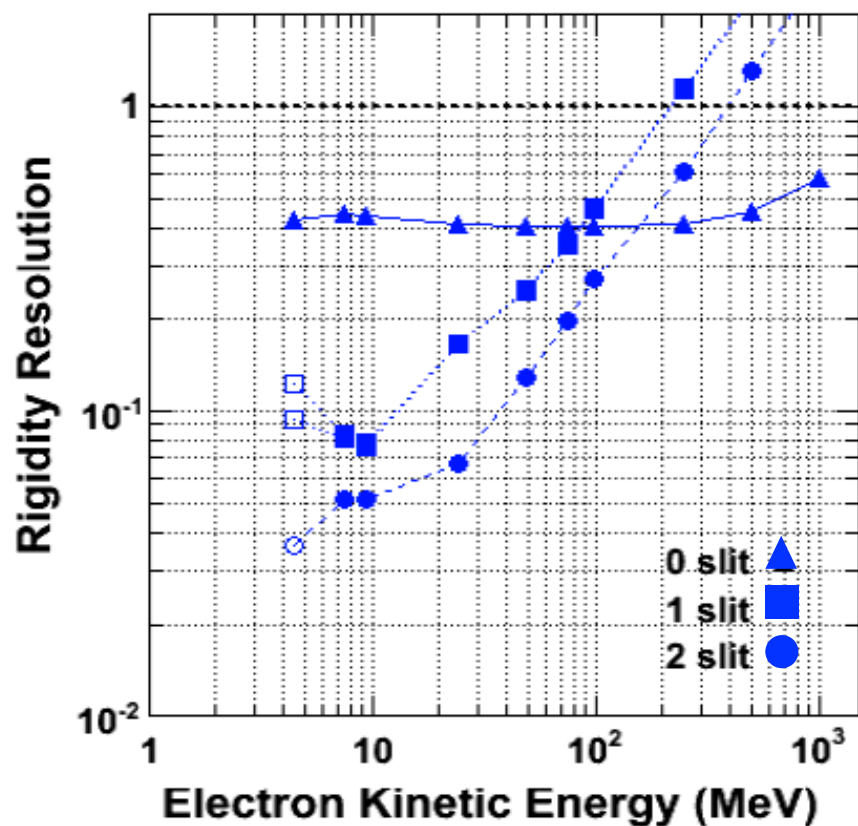
SAMPEX: 4-20 MeV Correlation slightly orbit dependent;



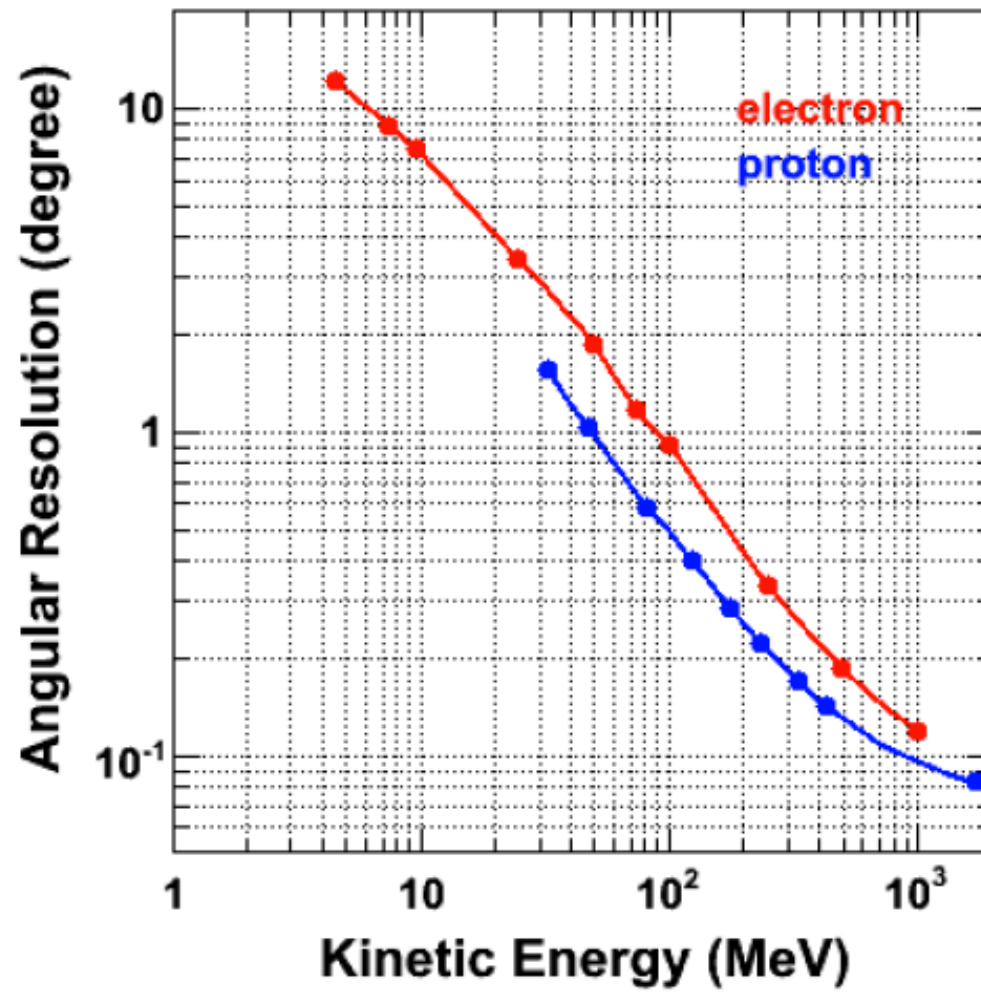
1 Slit Event 3 point reconstruction

Electron Performance

Resolution and Acceptance



Angular Resolution



Low Energy Particle Spectrometer

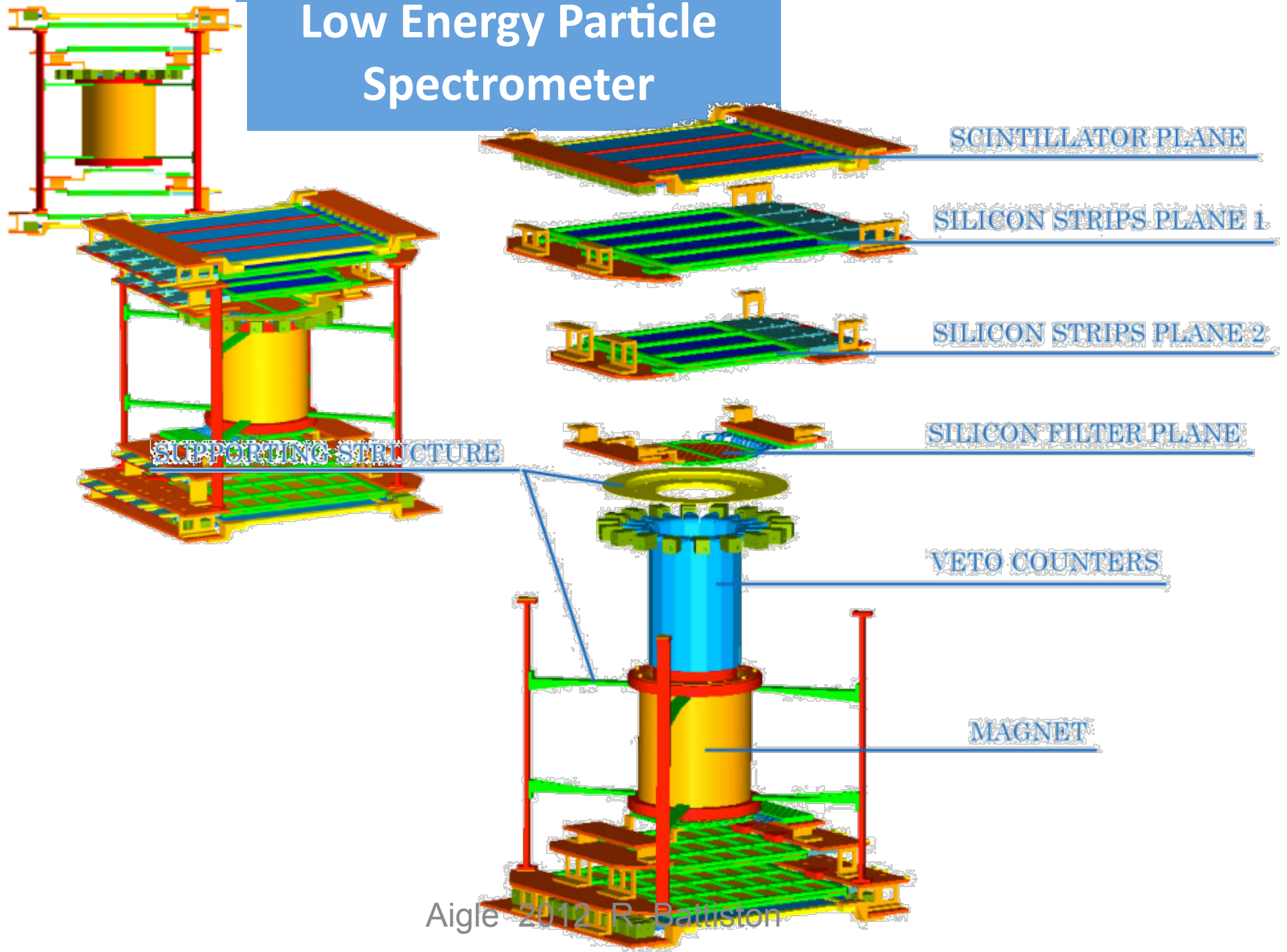


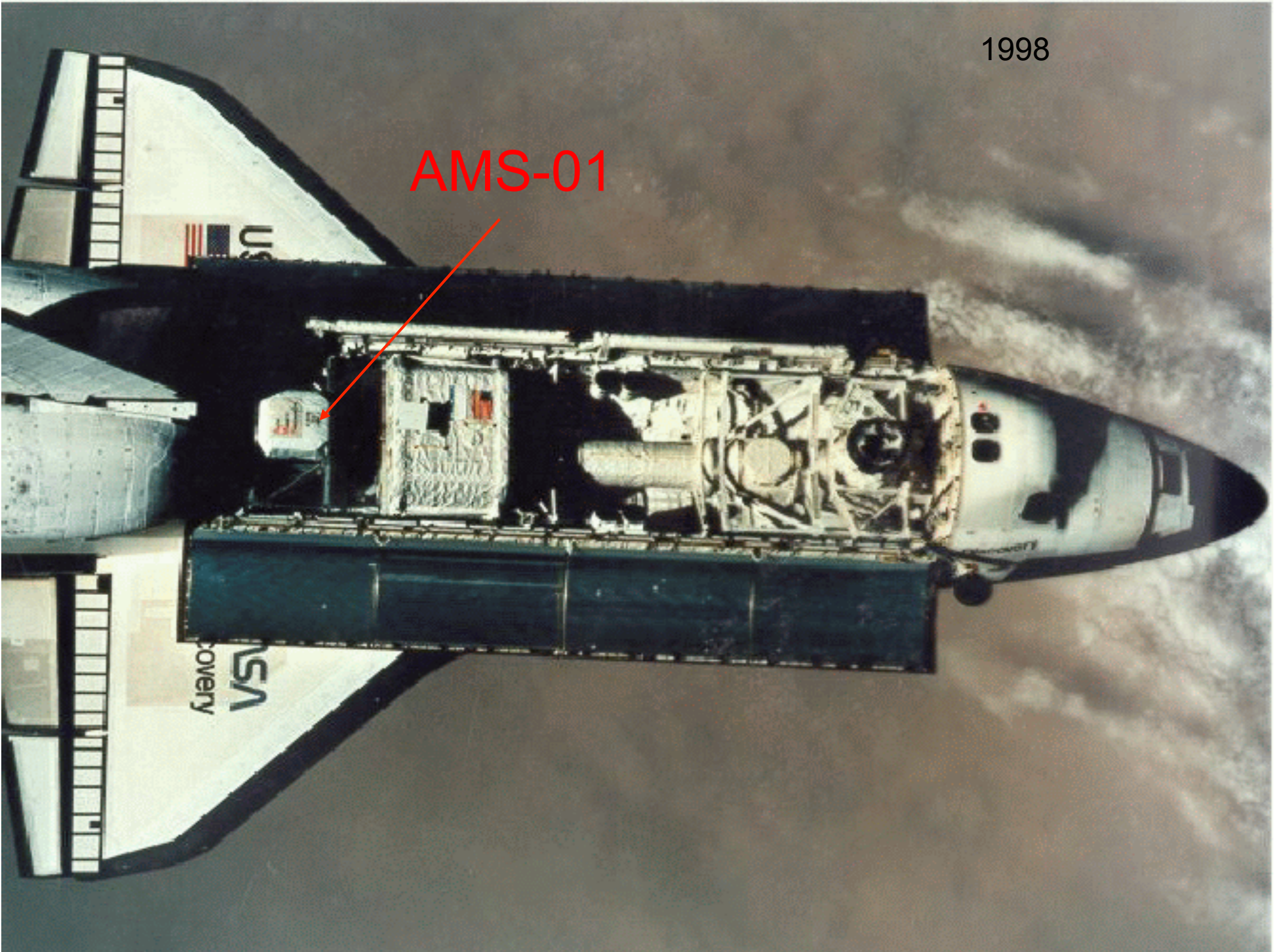
Table 1: Performance compared with existing satellite payload particle detectors.

Detector	Geometric Factor <i>cm²-sr</i>	Aperture	Pitch Angle	Energy Range	
				Electron <i>MeV</i>	Proton <i>MeV</i>
NOAA MEPED	0.1	30 ⁰ no <u>angular res.</u>	0-90 ⁰ 90-0 ⁰	0.03-2.5	0.03-6.9
AIGLON (Si)	4.35 30	70 ⁰ (Δ 13.4 ⁰ Δ 1.7 ⁰) ²	0-90 ⁰	5-25 ¹	30-100 ¹

¹energy range with resolution better $\leq 15\%$
² limiting values of the angular resolution at the lowest energies

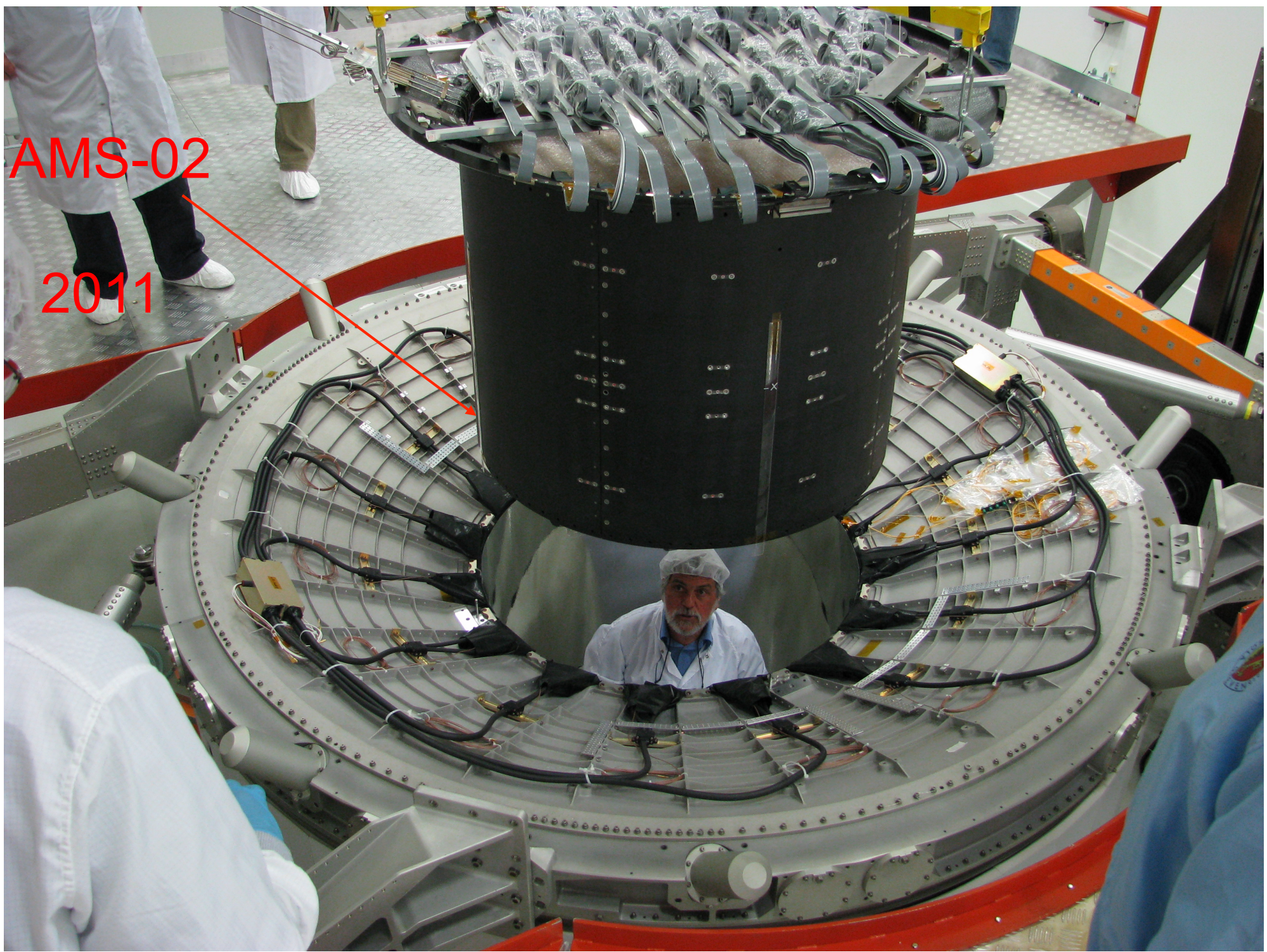
1998

AMS-01



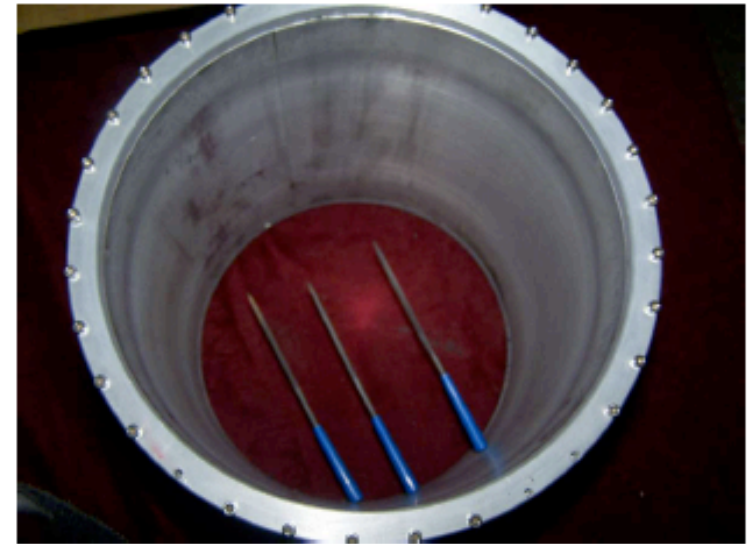
AMS-02

2011



NdFeB Magnets

Feasibility study magnets

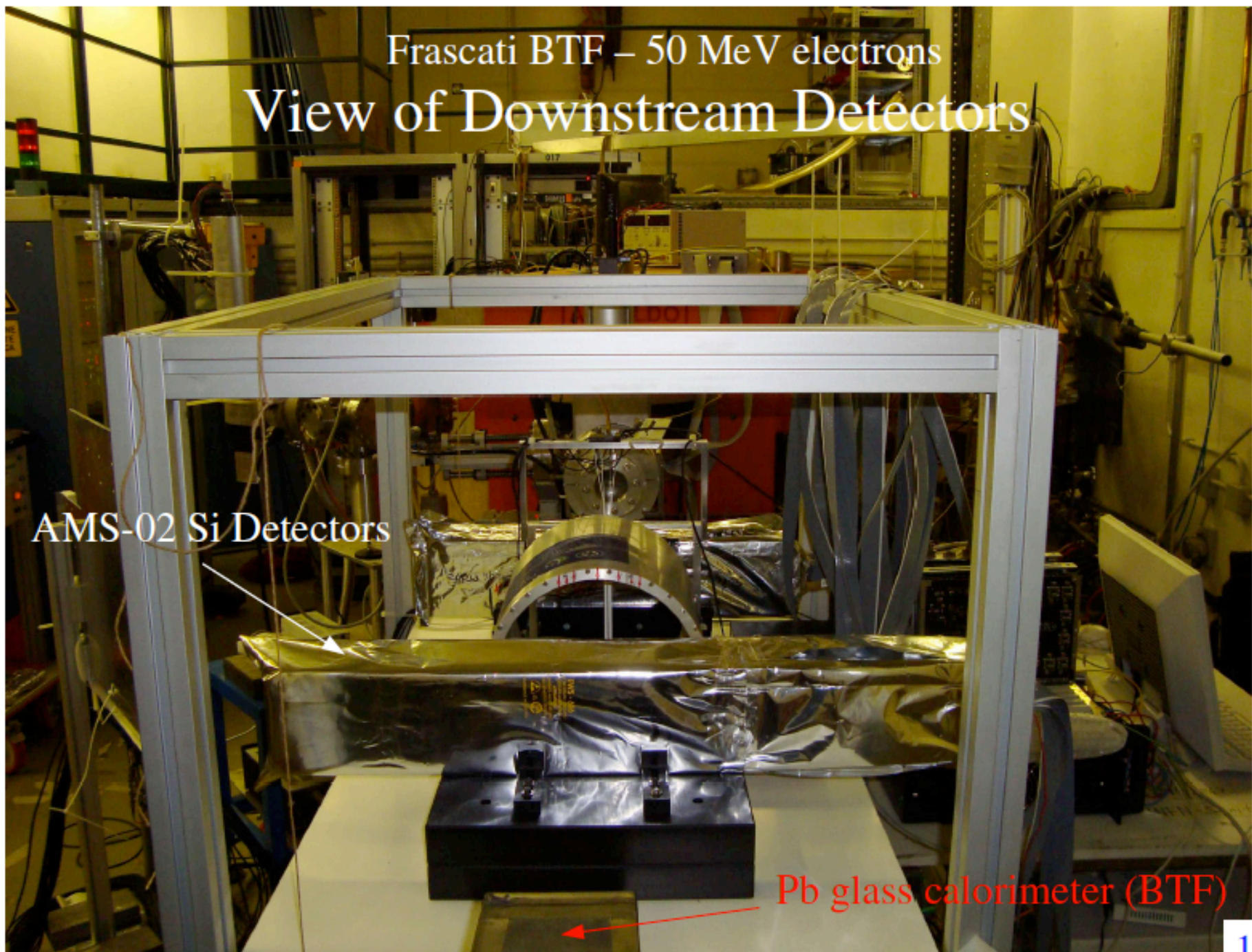


B = 0.522 kG
mass = 10.2 kg

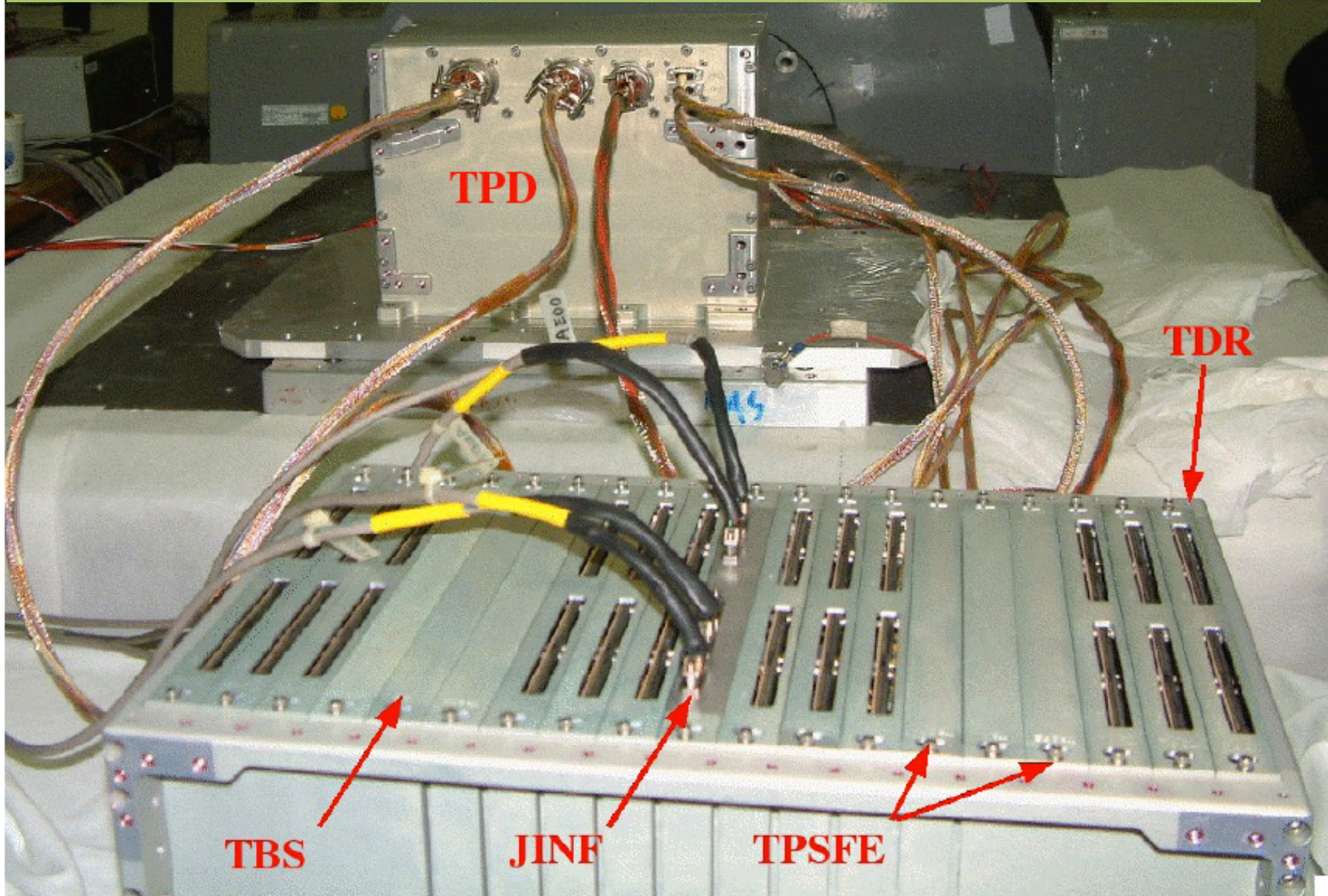
Frascati BTF – 50 MeV electrons
View of Downstream Detectors

AMS-02 Si Detectors

Pb glass calorimeter (BTF)



Existing tracker electronics system now operating on the ISS for 1,5 years



5) Conclusions

Observations from space support the evidence for detectable precursor phenomena influencing the magnetosphere producing precipitating particle burst taking place 1-3 hours before major earthquakes

In order to localize the position of the precursor effects a low energy cosmic ray detector is required

This design is derived from the technology used successfully for the AMS detector on the ISS

Thanks for your attention