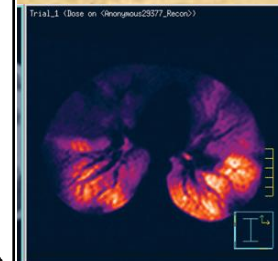
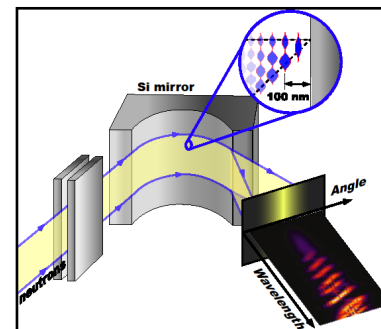
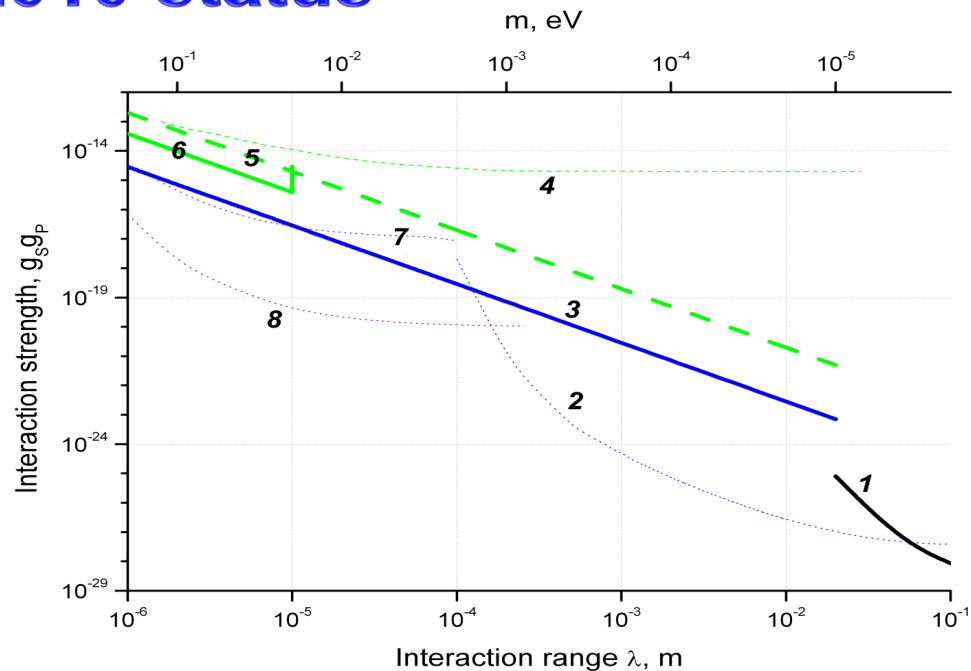
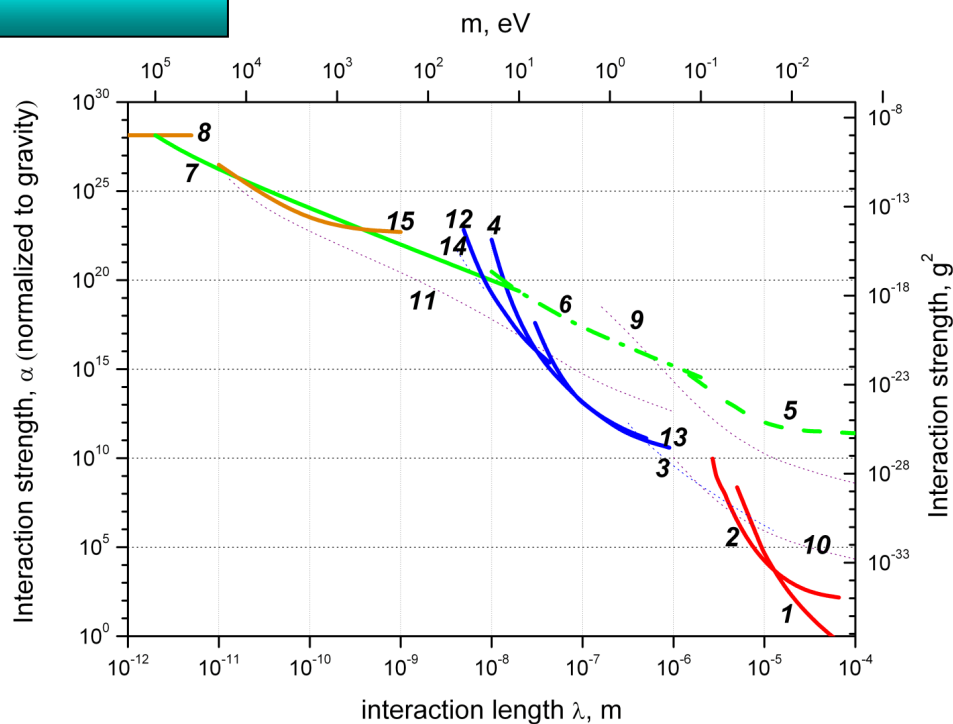


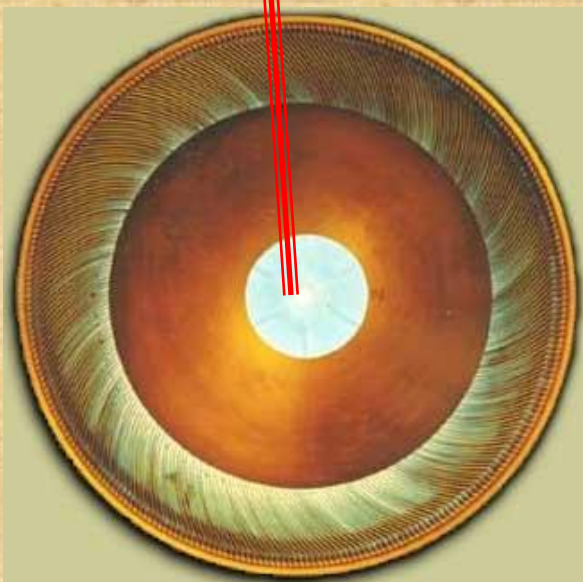
1. **Constraints** for spin-independent, spin-dependent, chameleon-like extra short-range forces are being **improved** over a broad range of distances.
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Let's start from the conclusion GRANIT-2010 status

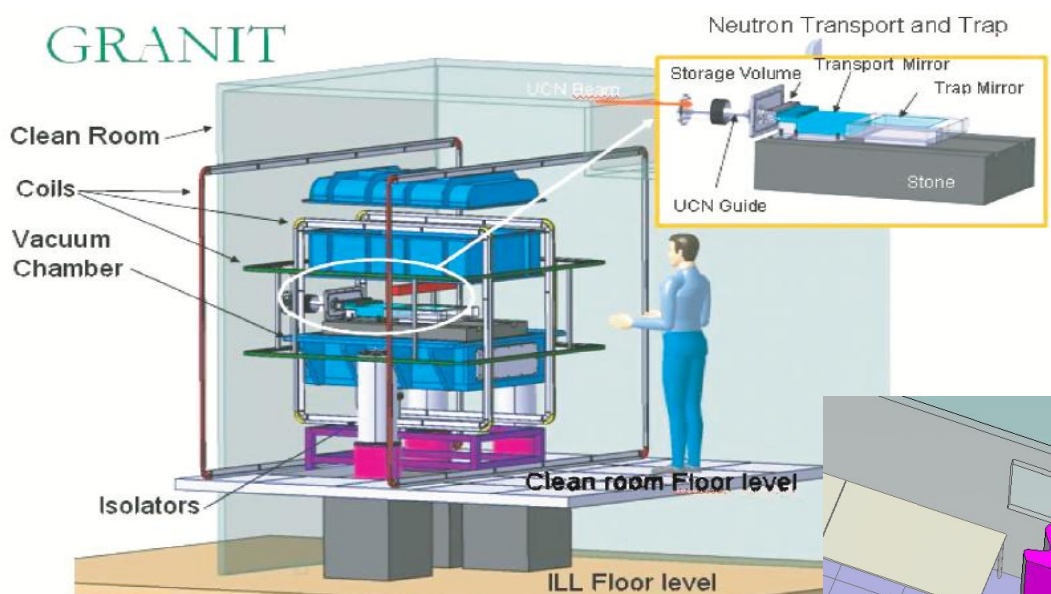


GRANIT facility

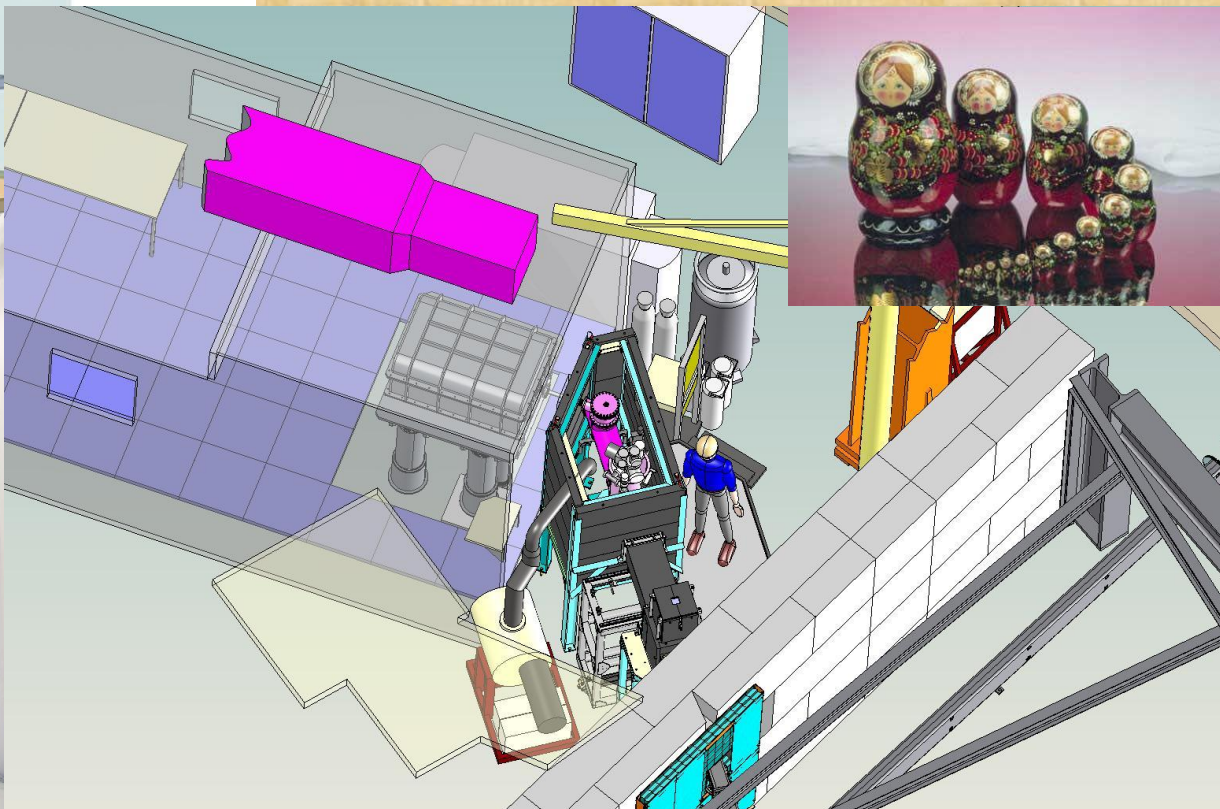
Located at the Institut Laue-Langevin
in Grenoble, France (largest nuclear
research reactor in the world)



GRANIT



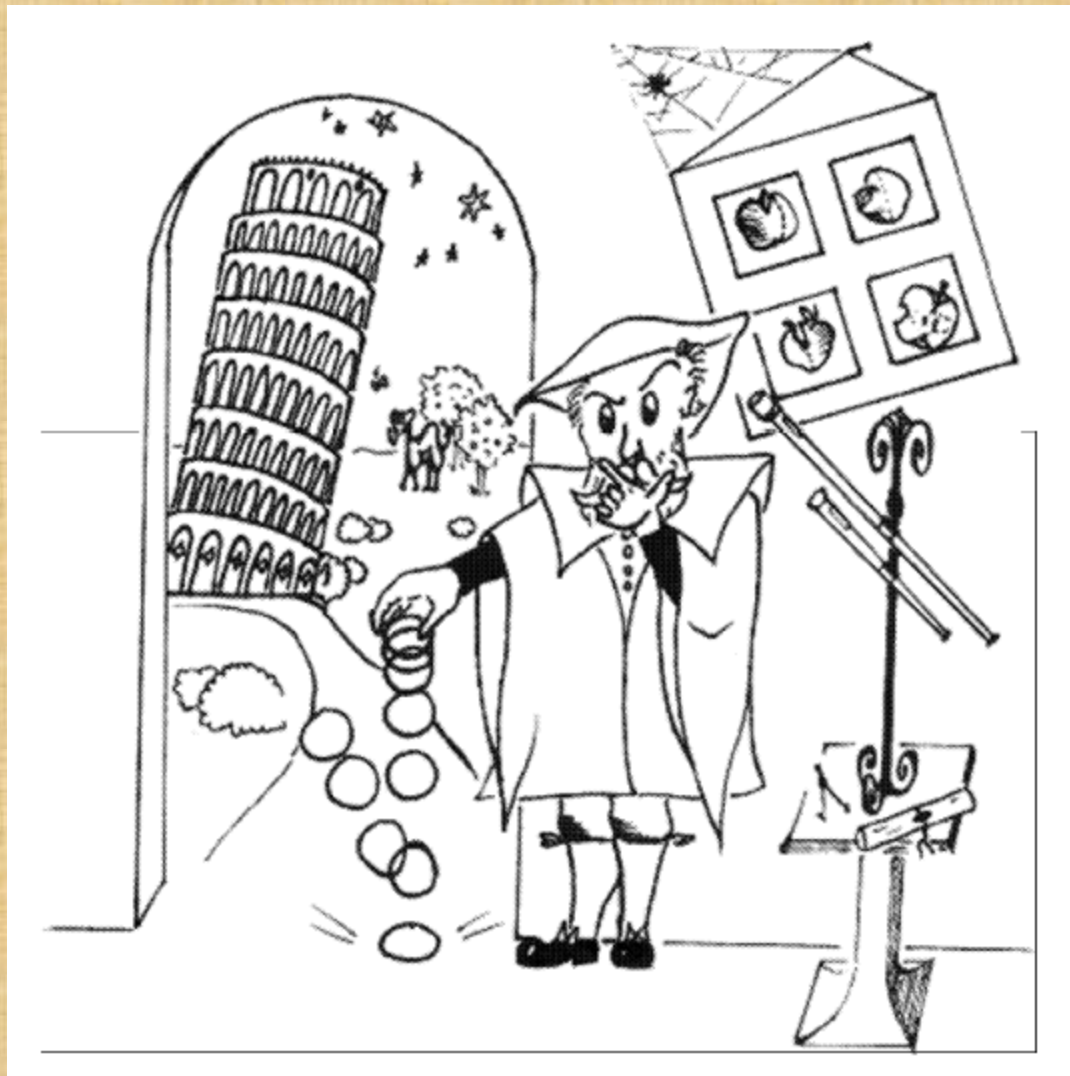
ILL reactor; precision optics; dedicated UCN source; clean room; unique detectors

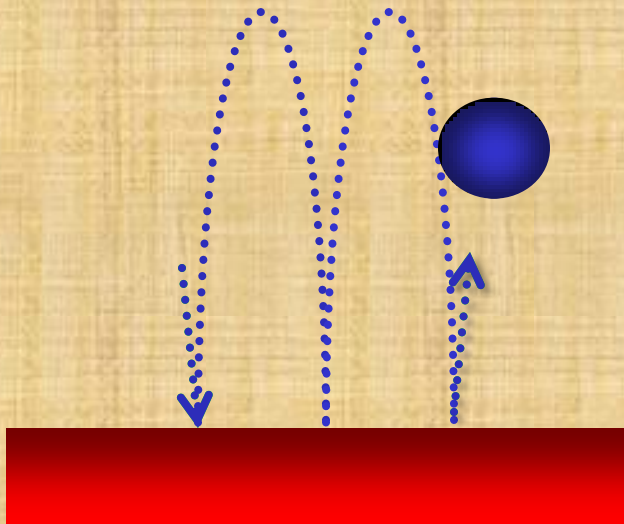


Participants: 12 countries (Europe, Asia, America, Australia)
Ones per 4 years (during Winter Olympic Games)
Previous Workshop 02-07 March 2014, Les Houches,
proceedings will be published in a special issue of
Advances in High Energy Physics « Quantum
gravitational spectroscopy »



Classical and Quantum bouncing





1) **Electrical neutrality** (usually the gravitational interaction for an object above a surface is much weaker than other interactions)

2) **Long life-time**

3) **Small mass**

$$\left(\Delta v \cdot \Delta x \approx \frac{\hbar}{m} \right) \left(\Delta E \approx \frac{\hbar}{\Delta \tau} \right)$$

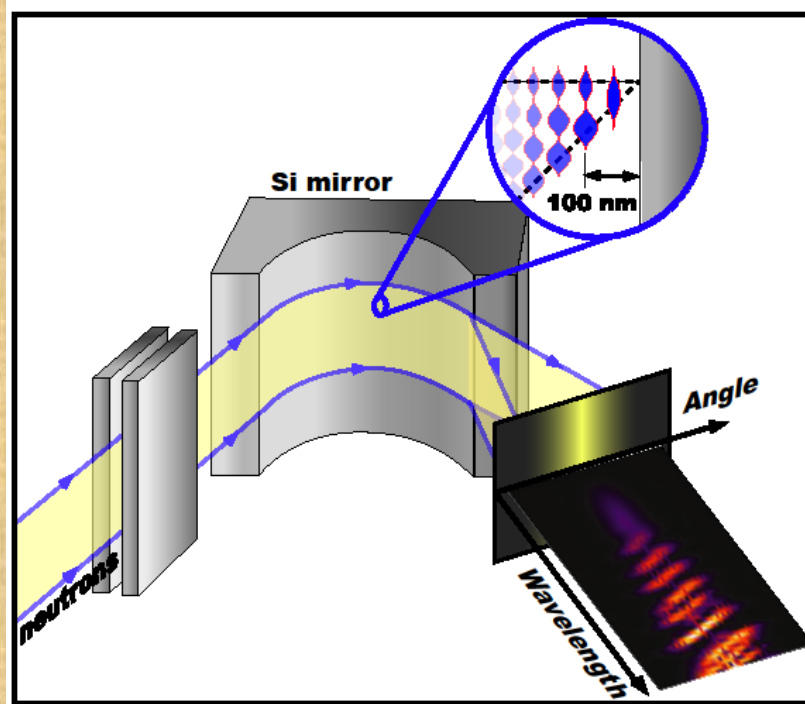
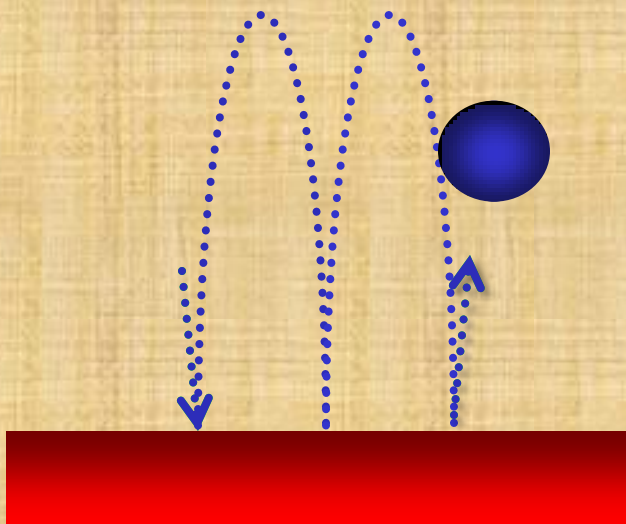
4) **Energy (effective temperature) of UCN, or an atom, is extremely low; it is not equal to the surface temperature (the effective temperature of a particle in gravitational quantum states is ~10 nK)**

A particle above a mirror in the gravity field: An ultracold neutron (V.I. Luschnikov, A.I. Frank « Quantum effects occurring when ultracold neutrons are stored on a plane », *JETP Lett.* 28 (1978) 759) and ... an anti-hydrogen atom (A.Yu. Voronin, P. Froelich, V.V. N. « Gravitational quantum states of antihydrogen », *Phys. Rev. A* 83 (2011) 032903) (a mirror represents a nearly infinitely-high and sharp potential step)

Energy of quantum states, in the Bohr-Zommerfeld approximation, equals :

$$E_n \approx \sqrt[3]{\left(\frac{9 \cdot m}{8}\right) \cdot \left(\pi \cdot \hbar \cdot g \cdot \left(n - \frac{1}{4}\right)\right)^2}$$

Observation of gravitational quantum states of neutrons (V.V. N. et al « *Quantum states of neutrons in the Earth's gravitational field* », *Nature* 415 (2002) 297



Observation of centrifugal (whispering-gallery) quantum states of neutrons (V.V. N. et al « *Neutron whispering gallery* », *Nature Physics* 6 (2010) 114

Gravity / Acceleration

$$E_n \approx \sqrt[3]{\left(\frac{9 \cdot m_n}{8}\right) \cdot \left(\pi \cdot \hbar \cdot g \cdot \left(n - \frac{1}{4}\right)\right)^2}$$

Height above
mirror



40 μm

30 μm

20 μm

10 μm



Illustration for quantum motion of an object above a mirror in the gravitational field and that in the accelerating frame. Positions of the ball correspond to the most probable heights of a neutron in the 5th quantum state.

*Yesterday's sensation is
today's calibration and
tomorrow's background*
[Richard Feynman]

Matter / Anti-matter



*Gravitational
properties of
antimatter have never
been measured
directly!*

*but aimed at, for
instance, in GBAR
project at CERN*

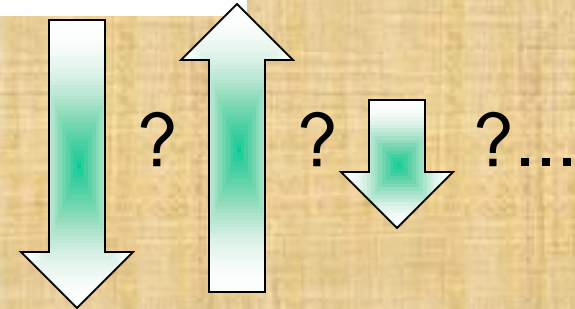


Illustration for quantum motion of a matter object and an anti-matter object above a mirror in a gravitational field.

Long storage of UCNs in gravitational quantum states

**Remember: flow-through mode;
modest energy resolution**

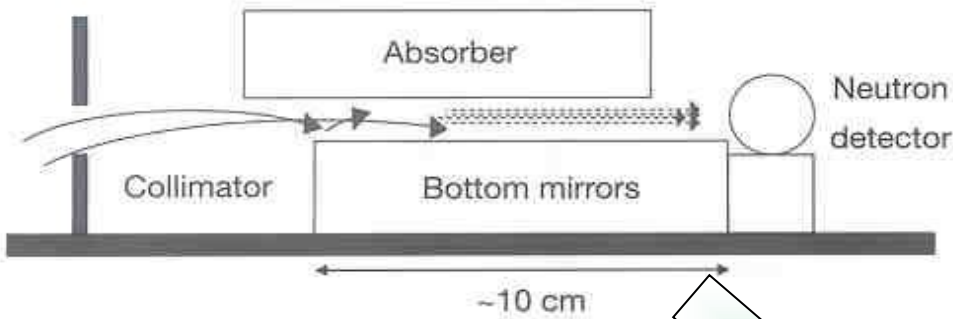


Figure 2 Layout of the experiment. The limitation of the vertical component depends on the relative position of the absorber and mirror. To limit the horizontal velocity component we use an additional entry collimator. The relative height and position of the entry collimator can be adjusted.

Transitions could be excited, for instance:

- *By periodically varying magnetic field gradient;*
- *By periodically varying local gravitational field;*
- *By oscillating mechanically the mirror*

V.V. N., K.V. Protasov, « Quantum states of neutrons in the Earth's gravitational field : state of the art, applications, perspectives », Ed. Book on Trends in Quantum Gravity Research (NOVA science publishers, New York, 2006).

Now: storage mode, long observation time and high energy resolution

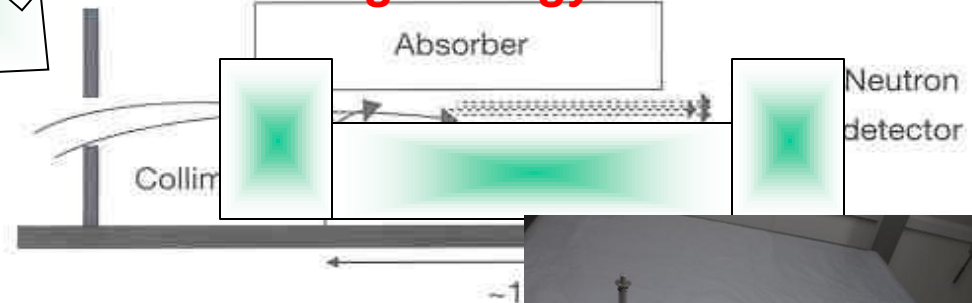
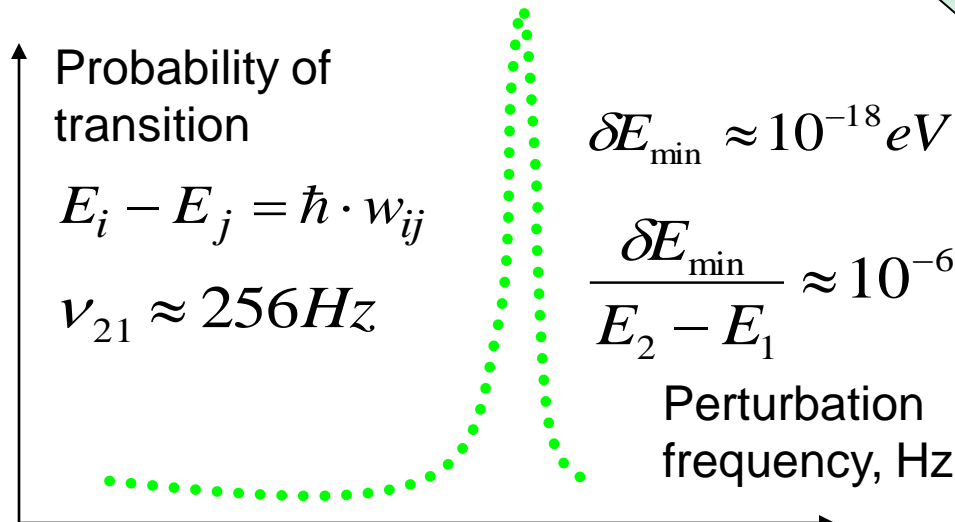
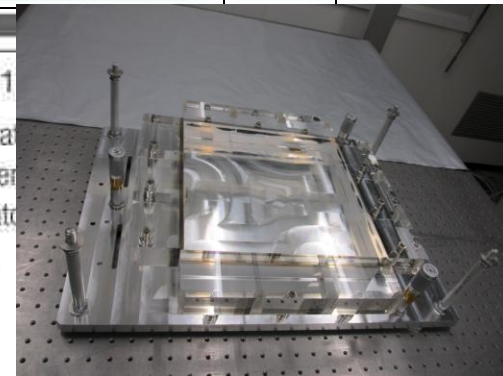
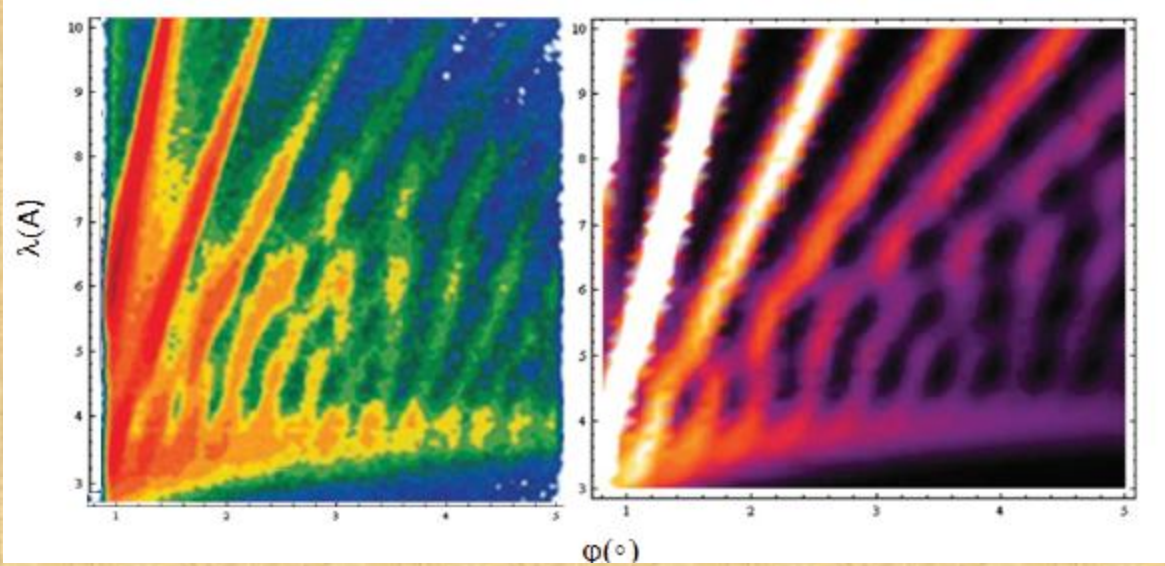


Figure 2 Layout of the experiment. The limitation depends on the relative position of the absorber component we use an additional entry collimator. The relative height and position of the entry collimator can be adjusted.



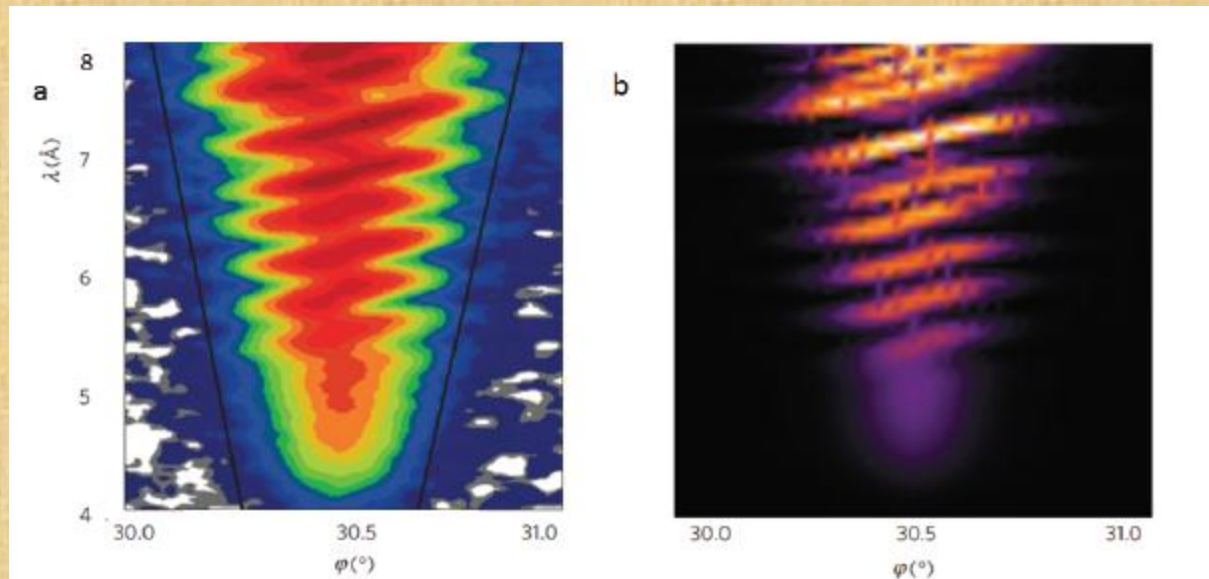
- The GRANIT facility has been constructed and is in its commissioning stage;
- First UCNs were delivered to the GRANIT detector during previous reactor cycle in 2013;
- Restart of the ILL reactor and continuation of measurement in July 2014;
- First measurements (a couple of years) in the « flow-through » mode;
- Meantime « flow-through » measurements are in progress by qBounce and Tokyo collaborations;
- Then precision measurements with GRANIT in the accumulation mode with long storage of UCNs in gravitational and whispering-gallery quantum states.

Neutron rainbow (whispering gallery states)

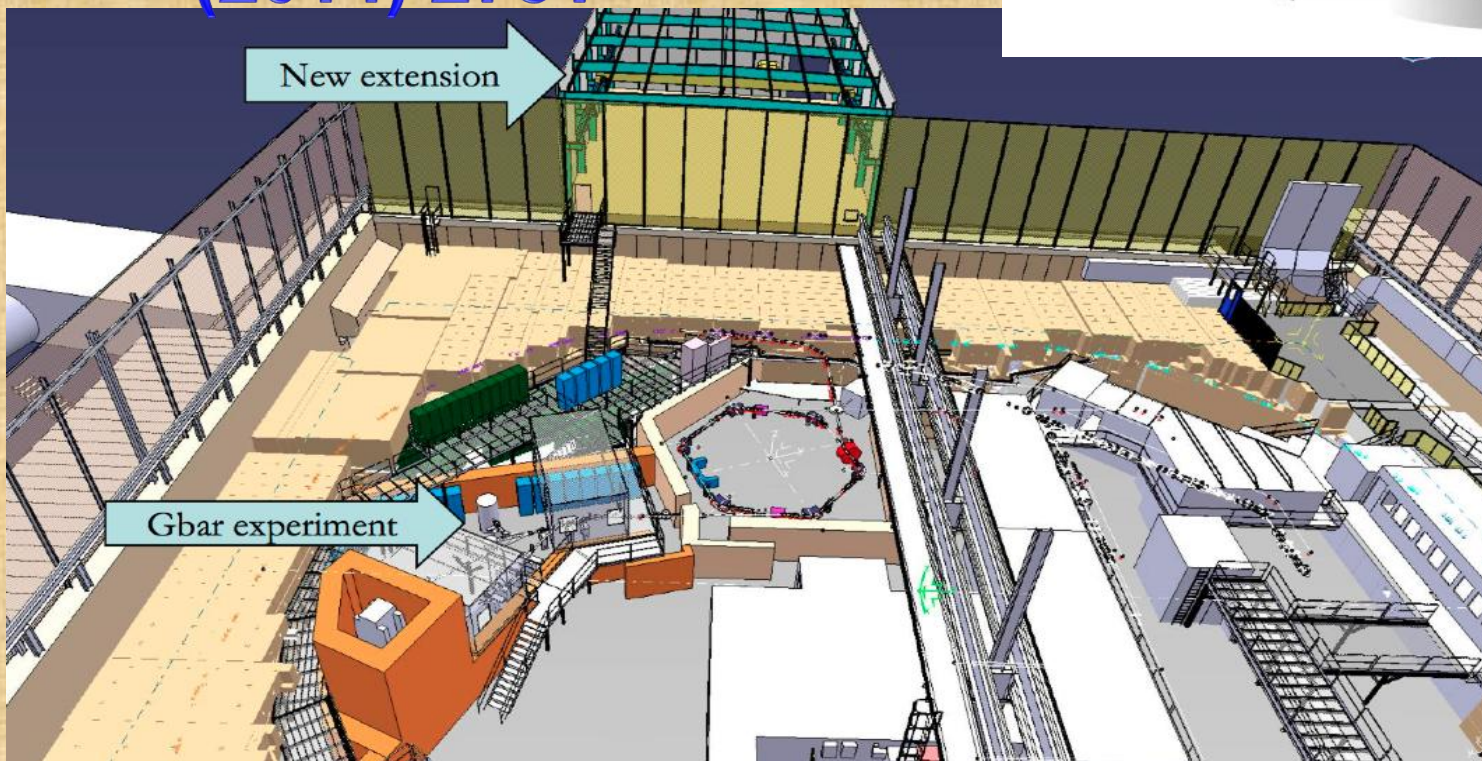
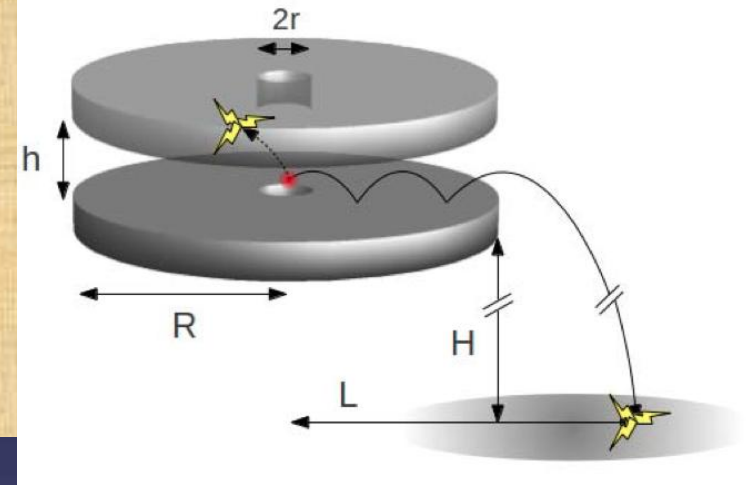


Neutrons tunneling
IN and then OUT of
quantum states

Neutrons entering
from the edge of a
truncated mirror and
then exiting from the
opposite edge

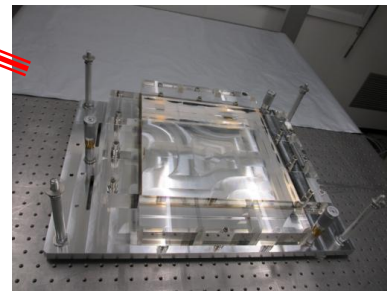
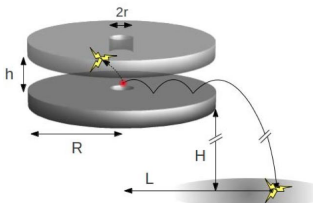
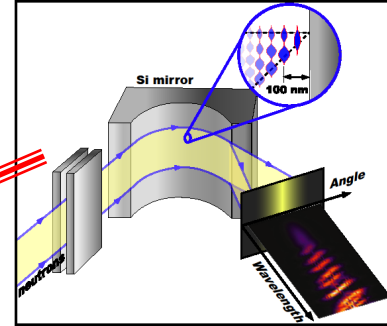
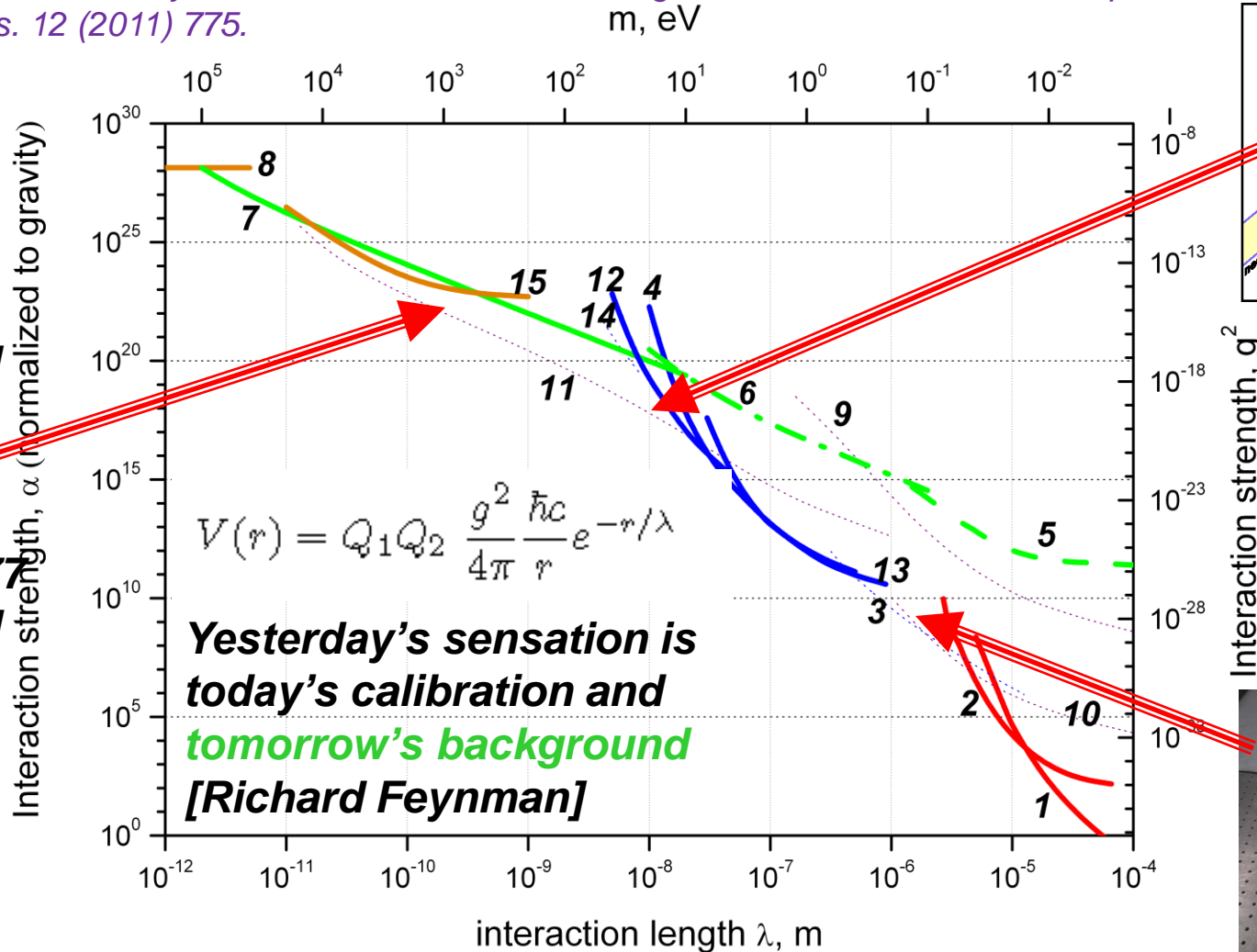


G. Dufour et al, *Shaping the distribution of vertical velocities of antihydrogen in GBAR*, *Europ. Phys. J. C* 74 (2014) 2731



Constrains for short-range forces

I. Antoniadis, S. Baessler, M. Buchner, V.V. Fedorov, S. Hoedl, V.V. N., G. Pignol, K.V. Protasov, S. Reynaud, Yu. Sobolev, « Short-range fundamental forces », *Compt. Rendu Phys.* 12 (2011) 775.



[V.V. N. et al, Neutron scattering and extra short range interactions, *Phys. Rev. D* 77 (2008) 034020]

Constraints for chameleon-like short-range forces

[Ph. Brax et al, Probing strongly coupled chameleons with slow neutrons, Phys. Rev. D 88 (2013) 083004]

instance a Ratra-Peebles model [6]

$$V(\varphi) = \Lambda^4 + \frac{\Lambda^{4+n}}{\varphi^n}$$

where $n > 0$ is the Ratra-Peebles index.

$$V_{\text{eff}}(\varphi) = V(\varphi) + \frac{\beta}{M_{\text{Pl}}} \varphi \rho, \quad (3)$$

where ρ is the mass density of matter and β is the dimensionless coupling constant of chameleons with matter. This effective potential is drastically different from

$$\varphi(\rho) = \Lambda \left(\frac{n \Lambda^3 M_{\text{Pl}}}{\beta \rho} \right)^{1/(n+1)}$$

Yesterday's sensation is today's calibration and tomorrow's background [Richard Feynman]

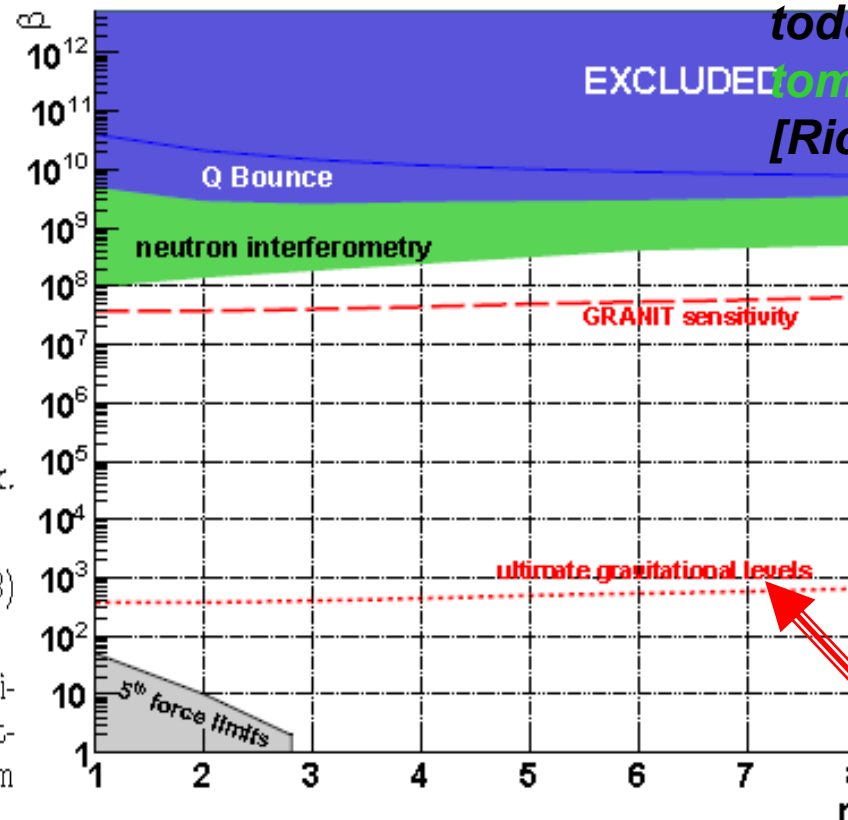
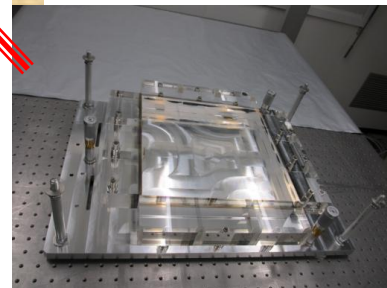
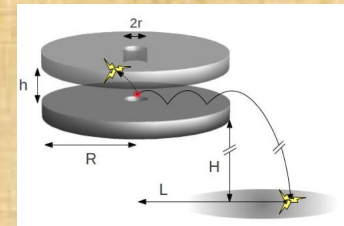
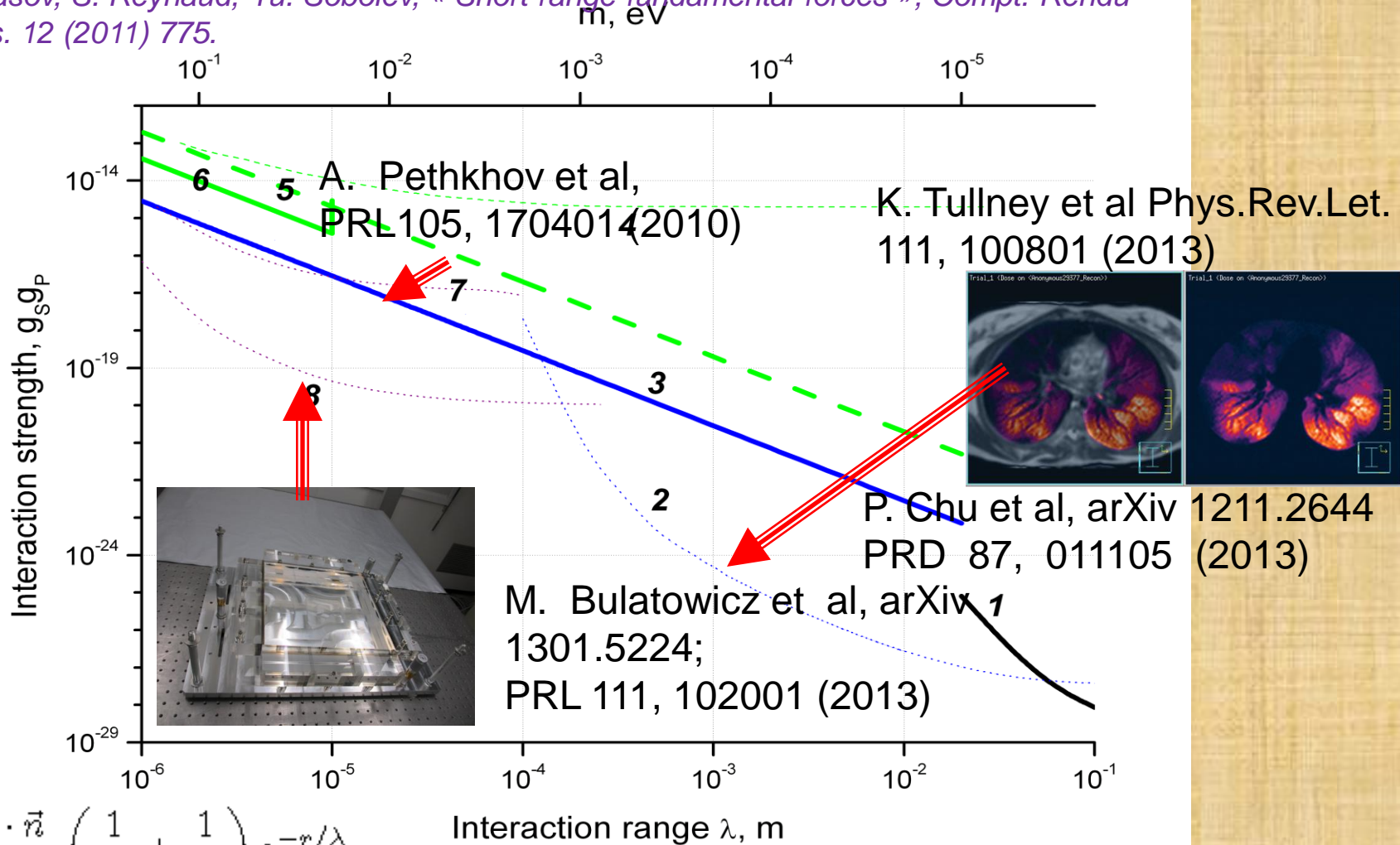


FIG. 8: The chameleon exclusion plot. The blue zone is excluded from [8]. The red lines are sensitivities calculated in [7] and confirmed in this work. The green zone is the potential reach of the neutron interferometry experiment proposed in this work.



Constrains for short-range forces

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$$V(\vec{r}) = \hbar g_P g_S \frac{\vec{\sigma} \cdot \vec{n}}{8\pi m c} \left(\frac{1}{\lambda r} + \frac{1}{r^2} \right) e^{-r/\lambda}$$

- I do not show new solid constraints, because we are analyzing new data of other and our groups at the moment;**
- For more results on spin-dependent forces at larger (than GRANIT) distances, see for instance talks at the GRANIT-2014 workshop by Mike Snow or Philipp Schmidt-Wellenburg;**
- Some of new results are expected to be summarized soon in S. Baessler et al (Advances in High Energy Physics, special issue « Quantum gravitational spectroscopy », 2014)**

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