
Breakdown Triggers and Breakdown Rate

Motivation of breakdown studies

- Breakdown rate is currently seen as the main obstacle in achieving the maximum possible gradient in CLIC accelerating structures
- Understanding the origin of breakdowns, experimentally and theoretically, may help in improving operating gradient
- Current strategy:
 - RF tests: breakdown rate experiments with new diagnostic tools for identifying the ion species produced from the emitted light, and searching for triggering mechanism
 - DC test: rapid turnover of samples, study of the properties of old/new candidate materials, measurements of breakdown rate and comparison with RF, diagnostics for studying breakdown physics
 - Modelling of breakdown initiation phenomena, both approximate and by atomistic simulations, and provide input parameters and ideas for RF design which could help suppressing or at least reducing breakdowns
 - ... and over again.

Workshop Timetable

09:00 Welcome (05') Jean-Pierre Delahaye (CERN)

09:05 Introduction (10') Sergio Calatroni (CERN)

09:15->10:35 Review of Current Experimental Work I (Convener: Sergio Calatroni (CERN))

09:15 RF Breakdowns: Experiments and Goals (30') (Slide) Steffen Doebert (CERN)

09:45 Breakdown in RF Structures: Material Studies (30') Gonzalo Arnau Izquierdo (CERN)

10:15 Ion Current from Breakdowns in RF Structures (20') Magnus Johnson (Department of Nuclear and Particle Physics, Uppsala university)

10:35 Coffee break

10:50->12:00 Review of Current Experimental Work II (Convener: Volker Ziemann)

10:50 Spectroscopic Analysis of DC and RF Breakdowns (30') (Slide) Jan Kovermann (III. Physikalisches Institut (B) - Rheinisch-Westfaelische Tech.)

11:20 DC Breakdown: Comparison of Different Materials (20') (Slide) Antoine Descoedres (CERN)

11:40 Outgassing Measurements in DC Breakdowns (20') Yngve Inntjore Levinsen (Norges Teknisk-Naturvitens. Univ. (NTNU))

12:00 Lunch Break

13:30->14:30 Modelling I (Convener: Kai Nordlund (University of Helsinki))

13:30 Atomistic Simulations of Breakdown Triggers: Progress Report (20') Flyura Djurabekova (University of Helsinki)

13:50 Delay Times in Breakdown Triggering (20') (Slide) Antoine Descoedres (CERN)

14:10 Breakdown Triggers and Breakdown Rate (20') Sergio Calatroni (CERN)

14:30->16:00 Modelling II (Convener: Walter WUENSCH (CERN))

14:30 Scaling Laws in RF Accelerating Structures (30') (Slide) Alexej Grudiev (CERN)

15:00 Some Observations During High Gradient Linac Testing at 17GHz using a Power Amplifying Dual Resonant Ring (15') Jake Haimson

15:15 Power Circulation in a CLIC Accelerating Structure (15') (Slide) Igor Syratcev (CERN)

15:30 RF Design Options for Quenching Breakdowns (30') Riccardo Zennaro (CERN)

16:00 Coffee break

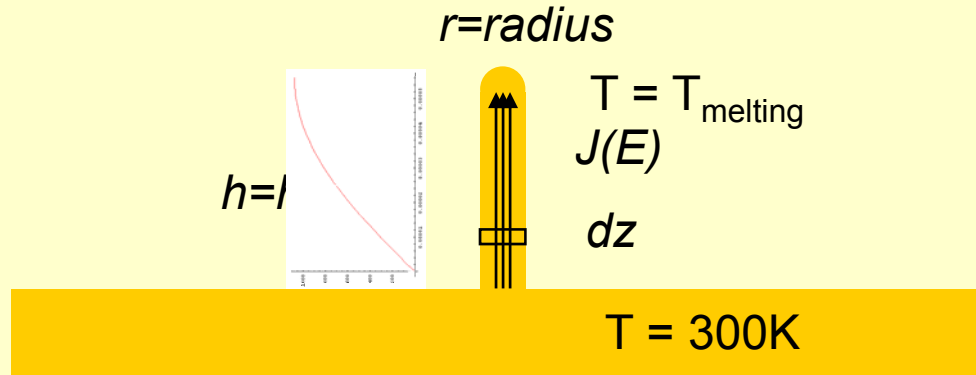
16:15 Discussion and Perspectives (45')

17:00 End of the meeting (20')

Triggering a breakdown

- A breakdown is an ionisation cascade, fuelled by some e.m. power (RF fields in a cavity, or the energy stored in a capacitor in DC testing)
- This cascade must be triggered by “something”.
- There are strong indications that it is initiated by electron field emission:
 - Evidence for cathode-initiation in DC sparks
 - Surface conditioning and associated changes in β (field enhancement factor) both in DC and RF
 - Existence of dark currents
- The following simple model suggests that some conditions must be fulfilled for initiating a breakdown. In particular:
 - Duration of RF pulse
 - Local power density
- The basic idea is that field emission currents, emitted by “tips”, heat the emitting sites by Joule effect

Heating of tips by field emission - I



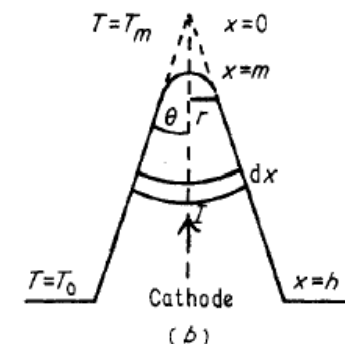
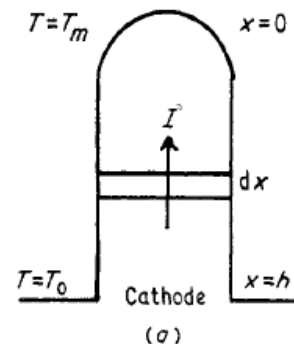
- The tip has a $\beta \approx \text{height/radius}$ (field enhancement factor)
- For a given value of applied E the Fowler-Nordheim law gives a current density $J(E) = A * (\beta E)^2 * \exp(-B/\beta E)$
- This current produces a power dissipation by Joule effect in each element dz of the tip, equal to $dP = (J \pi r^2)^2 \rho(z) dz / \pi r^2$
- The total dissipated power results in a temperature increase of the tip (the base is assumed fixed at 300 K). The resistivity itself is temperature dependent
- Using the equations we can, for example, find out for a given β what is the field that brings the “tip of the tip” up to the melting point, and in what time.

Heating of tips by field emission - II

- If the resistivity is considered temperature-independent, a stable temperature is always achieved [Chatterton Proc. Roy. Soc. 88 (1966) 231]
- If the resistivity (other material parameters play a lesser role) is temperature dependent, then its increase produces a larger power dissipation, resulting in a further temperature increase and so on [Williams & Williams J. Appl. Phys. D 5 (1972) 280]
- Below a certain current threshold, a stable regime is reached
- Above the threshold, a runaway regime is demonstrated
- The time dependence of the temperature can be calculated.

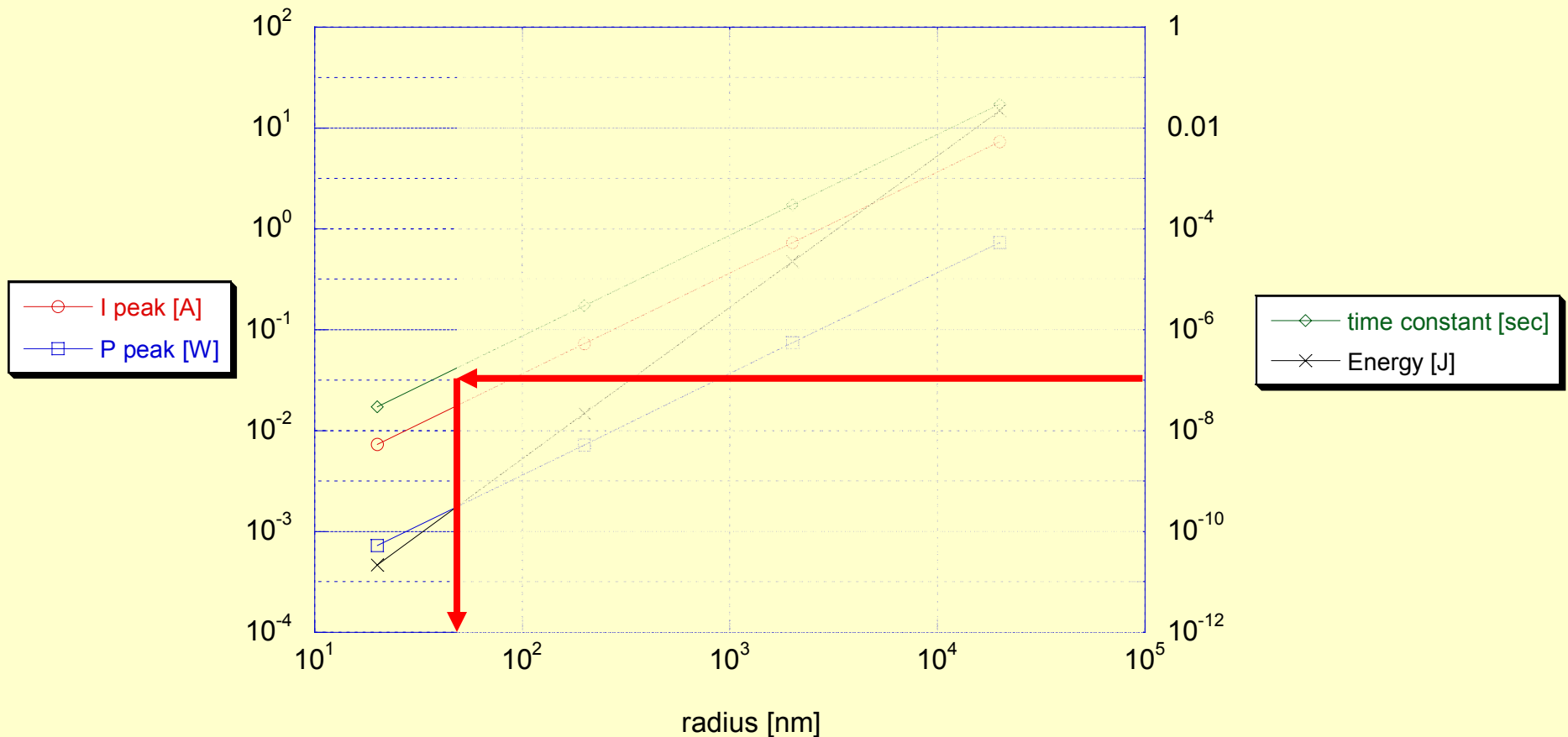
$$\frac{\partial T}{\partial t} = \frac{K}{\sigma S} \frac{\partial^2 T}{\partial x^2} + \frac{I^2 \rho_0}{\pi r^2 \sigma S} \{1 + \alpha(T - T_0)\},$$

$$\frac{\partial T}{\partial t} = \frac{K}{\sigma S} \left(\frac{\partial^2 T}{\partial x^2} + \frac{2}{x} \frac{\partial T}{\partial x} \right) + \frac{\rho}{\sigma S} \left(\frac{I}{2\pi x^2 (1 - \cos \theta)} \right)^2.$$



Time constant to reach the copper melting point (cylinders, $\beta=30$)

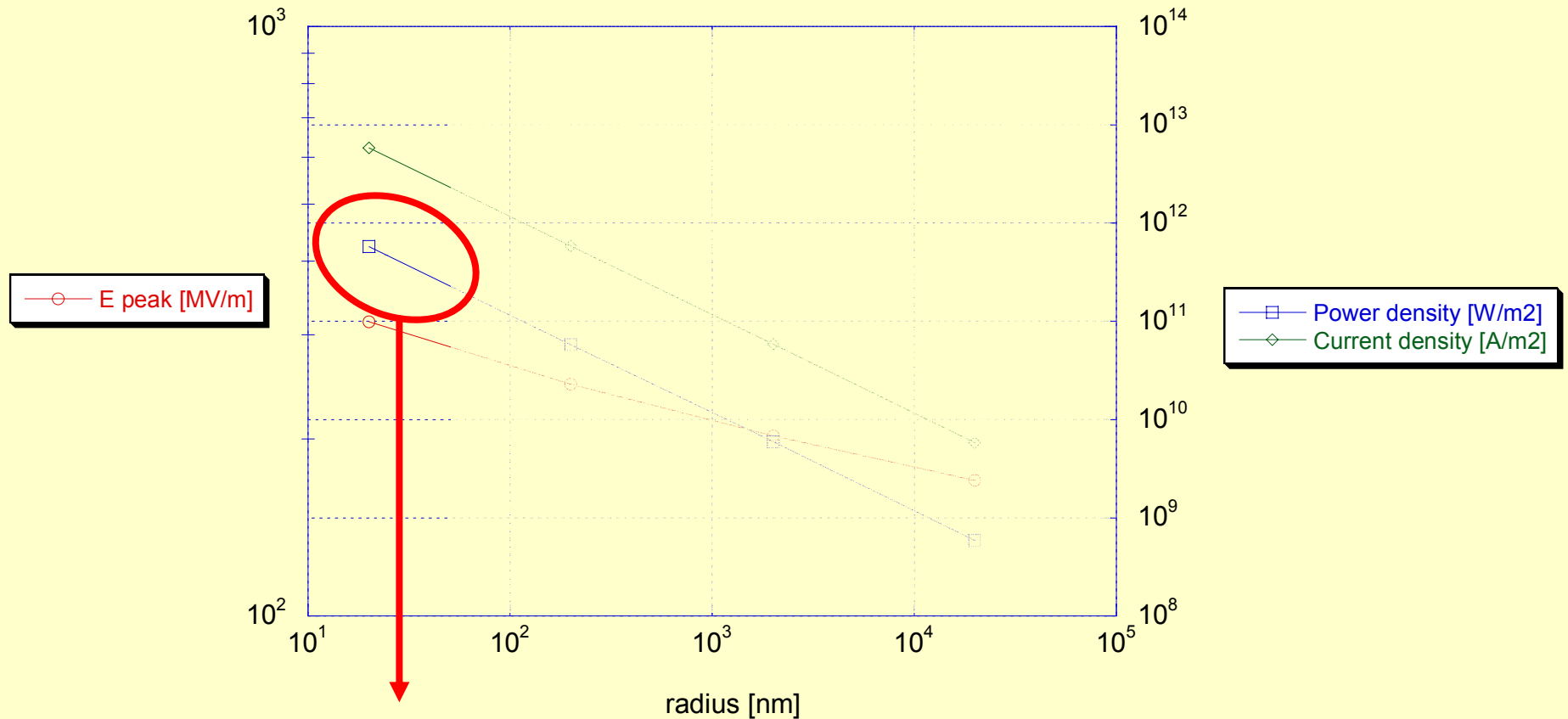
Parameters to attain the melting point of the tip of a Cu cylinder of given radius and $\beta=30$



The tips which are of interest for us are extremely tiny, <100 nm (i.e. almost invisible even with an electron microscope)

Power density at the copper melting point (cylinders, $\beta=30$)

Parameters to attain the melting point of the tip of a Cu cylinder of given radius and $\beta=30$



Power density (power flow) during the pulse is a key issue.
See talk by A. Grudiev for RF structures scaling based on Poynting vector

Breakdown: ionisation probability

- Emitting sites can get (very) hot because of Joule heating, and emit gas (metal vapours or outgassing) which gets ionised by field-emission current
- Field emission → heating → gas → ionised by electrons → ionisation cascade
- The breakdown probability: $P(x_1, x_2, \dots, x_n) = f(x_1)f(x_2)\dots f(x_n)$
- Where x_i might be E , τ or a even a combination of these or other physical quantities.
- The probability of igniting a cascade depends linearly on the amount of gas available and on the primary electron current

$$P_{breakdown} \propto I_{electrons} pressure_{gas}$$

- Where do the electrons and the gas come from?

Electron current and gas sources

- The electron current is given by the standard Fowler-Nordheim equation:

$$I_{electrons} = Const * (\beta E)^2 \exp(-B/\beta E)$$

- The constant includes the emitter area, the work function. $I_{electrons}$ is “exponential” with E (steep slope like in measured breakdown probabilities)
- β can be material dependent, and the tips might as well be “dynamical” tips, thus depending on field, number of pulses etc.
- The gas molecules can be either metallic vapours or other gas released by thermal outgassing (or both!), due to Joule heating
- T depend on pulse duration (simulations suggest that T is proportional to [current density J]² and grows with exp[time τ]).

Metal vapour pressure
(H_0 is the heat of vaporisation)

$$p = p_0 \exp\left(\frac{-H_0}{RT}\right)$$

Outgassing
(sign depends on endo/exothermic)

$$p = p_0 c_H^2 \exp\left(\frac{\pm E}{RT}\right)$$

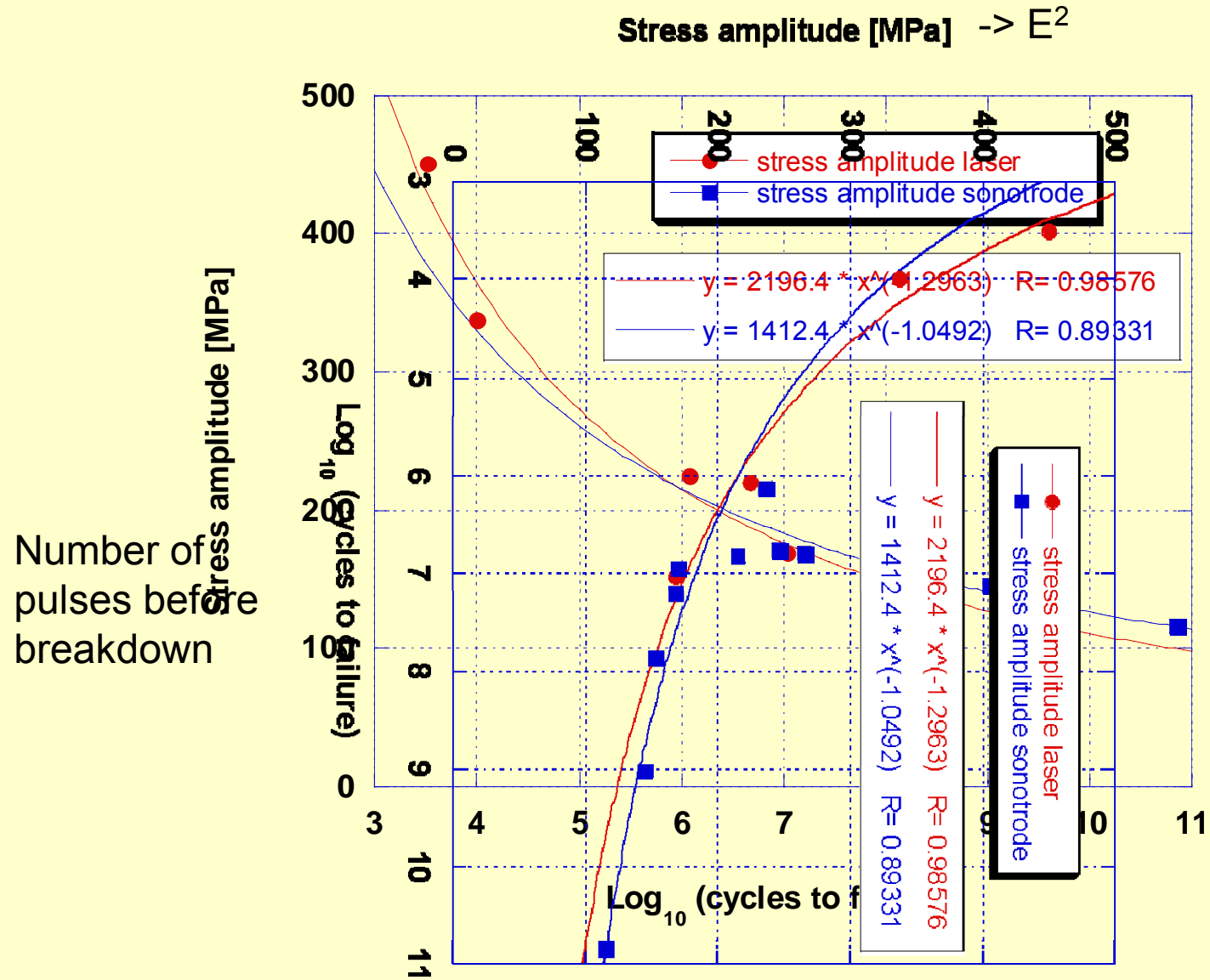
Pressures and number of molecules

- Order-of-magnitude estimate (1):
 - Vapour pressure of Cu at $T_{\text{melting}} \sim 10^{-3}$ mbar.
 - Tip radius = 25 nm, $\sim 10^6$ copper atoms emitted / second (0.1 atoms / 100 nsec: impossible to start a breakdown) ✓
 - At $T_{\text{sublimation}}$ (vapour pressure 1 bar) we have 6 orders of magnitude more atoms emitted ✓
- Order-of-magnitude estimate (2):
 - At T_{melting} , the equilibrium pressure for 10 ppm of H in Cu is ~ 1 bar ✓
 - Electron stimulated desorption. If we have currents in the mA range before breakdown, this may results in ESD $\sim 5 \cdot 10^{13}$ H₂ / sec (yield 10^{-2}) ✓
- Order-of-magnitude estimate (3):
 - Tip $\beta=30$ radius=25 nm $\Rightarrow \sim 10^8$ copper atoms ✓
 - 10 ppm of hydrogen in such a tip $\Rightarrow \sim 10^3$ hydrogen atoms (seems small) ✓

Breakdown probability as result of fatigue

- Two possible schemes for breakdown by fatigue:
 - Sites can get (very) hot because of Joule heating
 - Field emission → heating → mechanical stress → fatigue → material break-up
- Or:
 - Tips are simply pulled by electric field
 - Tip with field enhancement → pulling with force $\sim E^2$ → mechanical stress → fatigue → material break-up
- In the first case, the stress should go with temperature, which is proportional to J^2 (in turn related “exponentially” to E field) and exponentially with pulse duration. However the stress profile is unclear to me (there are no constrained surfaces)
- In the second case, the stress goes with E^2 and should not depend on the pulse duration, unless there is some change of mechanical properties due to temperature (and in this case we fall back partially in the FN-dominated mode).

Fatigue by pulling



7 orders of magnitude in # of cycles, with 1/4 stress (1/2 gradient?)

Breakdown by fatigue

- In case breakdown is due simply to pulling, some numbers are very easy
- The electrostatic pulling stress is: $\sigma \text{ [Pa]} = 0.5\epsilon_0 E^2$
- $E = 200 \text{ MV/m} \Rightarrow \sigma = 1.8 \cdot 10^5 \text{ Pa}$ (should be divided by 2 to have RMS field values).
- However: for $\beta = 50$ the stress increases by a factor 2500 $\Rightarrow \sigma = 2.5 \cdot 10^8 \text{ Pa}$. This is 250 MPa, 125 MPa amplitude.
- For info, pulsed heating target for Cu at 10^{11} cycles is 80 MPa stress amplitude for copper. 125 MPa correspond to 10^6 cycles lifetime.
- But still no pulse length dependance...

Conclusions

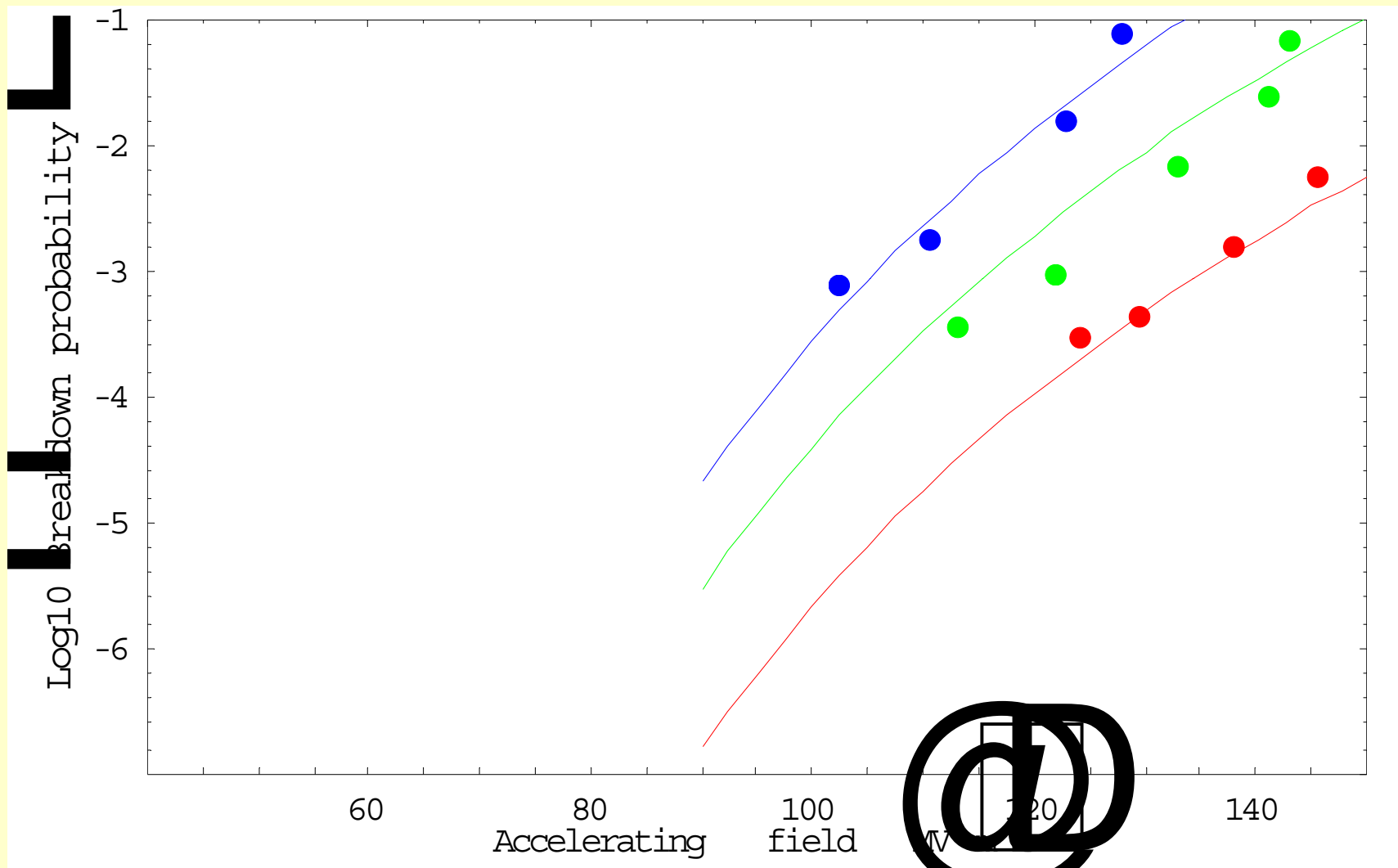
- Modelling of breakdown precursors as tips heated by Joule effect due to FN currents seems globally reasonable.
- This naturally may lead to breakdowns (by simple ionisation or by explosive process in extreme conditions)
- However understanding of why there should be a breakdown rate needs still efforts. Simple pictures as illustrated before do not satisfy all known dependences on field or pulse length
- Other possible paths (atomistic simulations):
 - Development of tips under E-field
 - Coupled with vaporization
 - And outgassing
 - And change of mechanical properties with temperature
 - And with fatigue.

Acknowledgements

- Mauro Taborelli
- Walter Wuensch
- Alexej Grudiev
- Antoine Descoeurdes
- Yngve Lenvinsen
- Trond Ramsvik
- Igor Syratchev
- Riccardo Zennaro

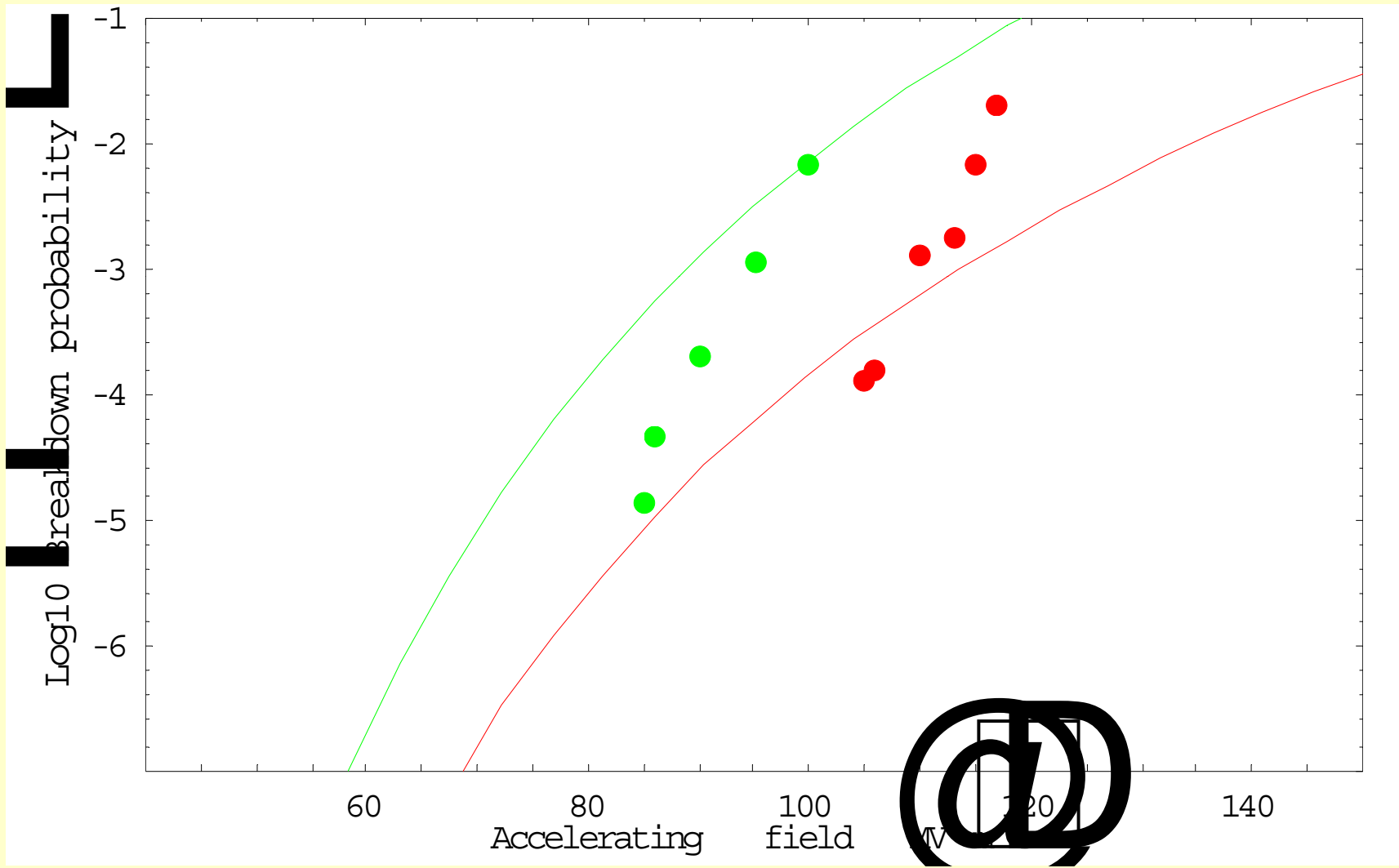
Fit to Mo data, 30 GHz circular iris

- $\beta = 30$, $k = 138 \text{ Wm}^{-1}\text{K}^{-1}$, $p_0 = 10^{14.5} \text{ mbar}$, $H_0 = 598 \text{ kJ/mol}$



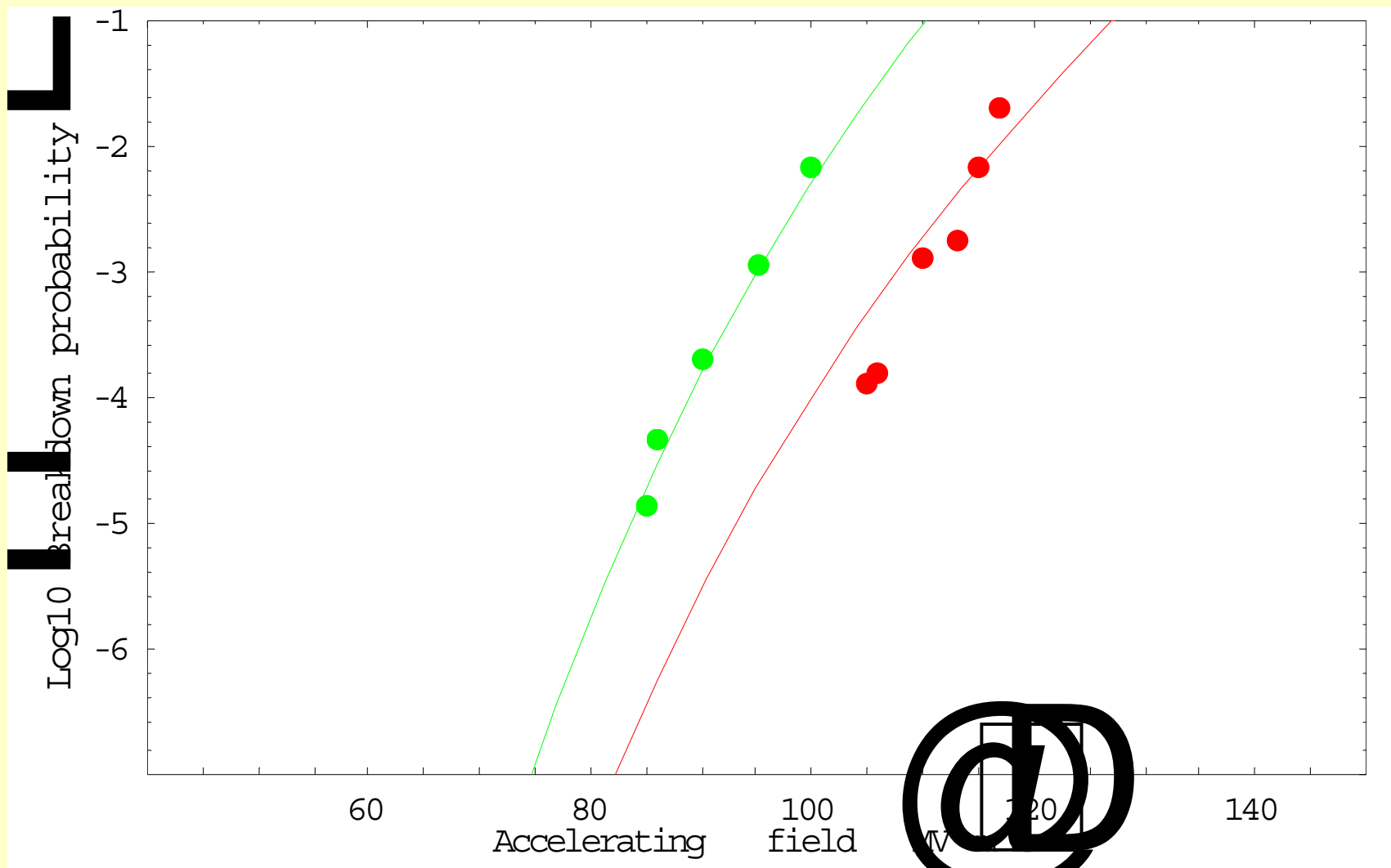
Keeping the same fit parameters and comparing to Cu data, 30 GHz

- $\beta = 45$, $k = 400 \text{ Wm}^{-1}\text{K}^{-1}$, $p_0 = 10^{12} \text{ mbar}$, $H_0 = 300 \text{ kJ/mol}$.



Letting free the F-N fit parameters and comparing to Cu data, 30 GHz

- B doubles and A increases of 6 units



Temperature rise calculations

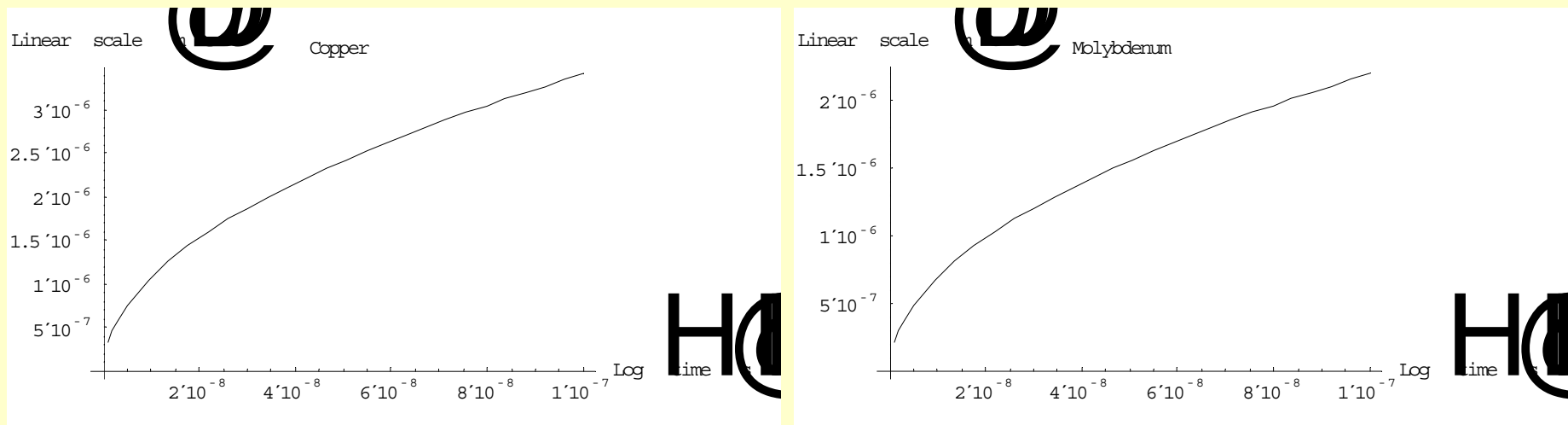
- Here starts the main part of the talk
- 1D, 2D, 3D time dependent heating
 - Relevant for the discussion on breakdown limit
- Heating of tips by field emission currents
 - Relevant for the discussion on breakdown probability

Time-dependent heating

- The breakdown limit of materials in RF tests is observed to follow the dependence: $P\tau^a$ with $a=1/3$ for copper and $a=2/3$ for molybdenum
- Is there any intrinsic material dependence? Heat flow equation:

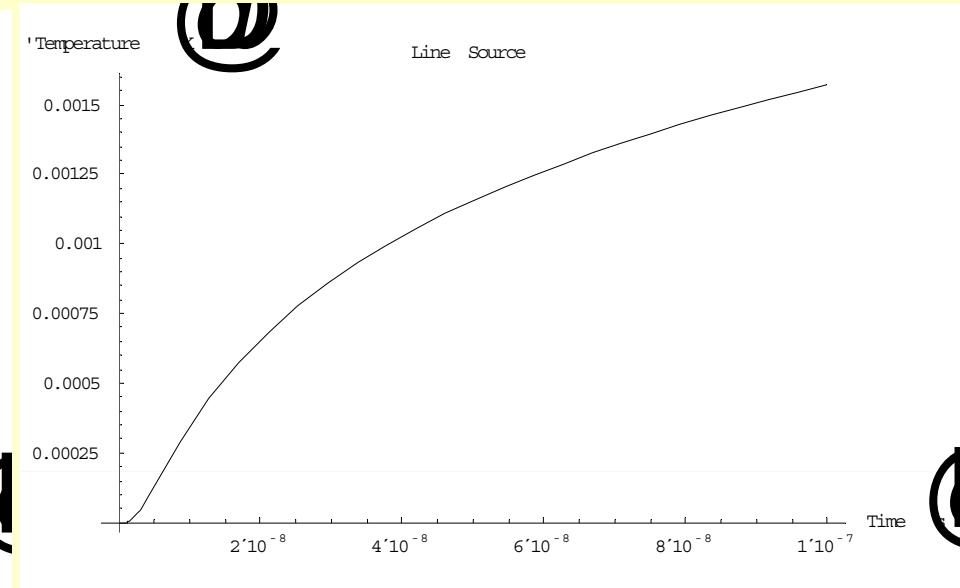
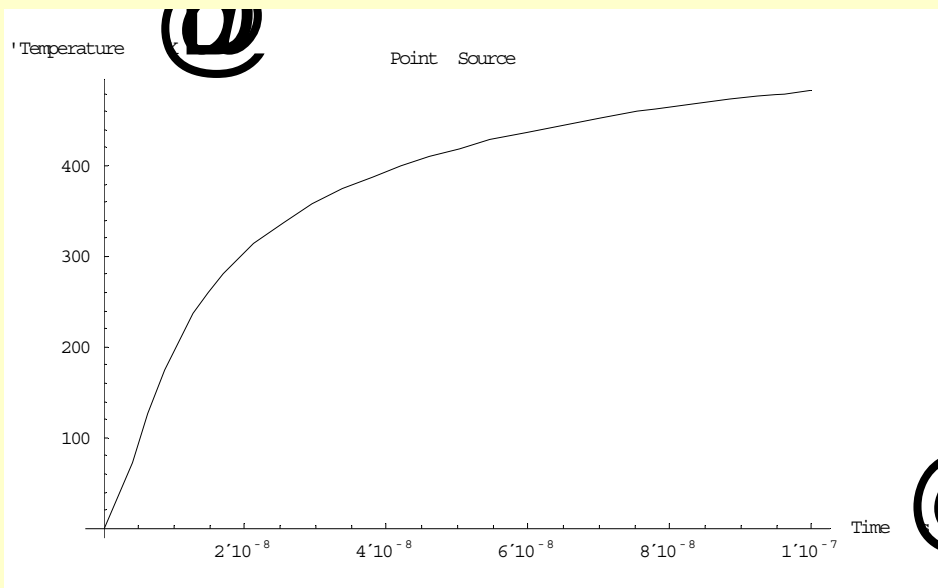
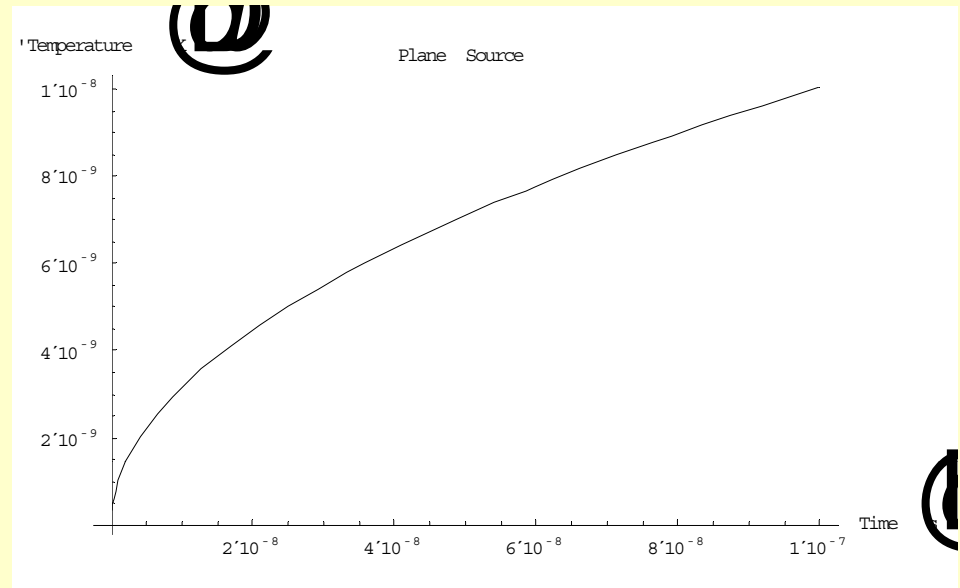
$$\nabla^2 T + \frac{\dot{q}}{k} = \frac{1}{\alpha} \frac{\partial T}{\partial t}$$

- With: k = thermal conductivity, $\alpha = k/(c*\rho)$, c = specific heat, ρ = density
- In-time dependent calculations the distinction between a “fast” and “slow” regime is based on the diffusivity time $\tau_D = R^2/\alpha$. R is the linear scale of the phenomena that are under consideration

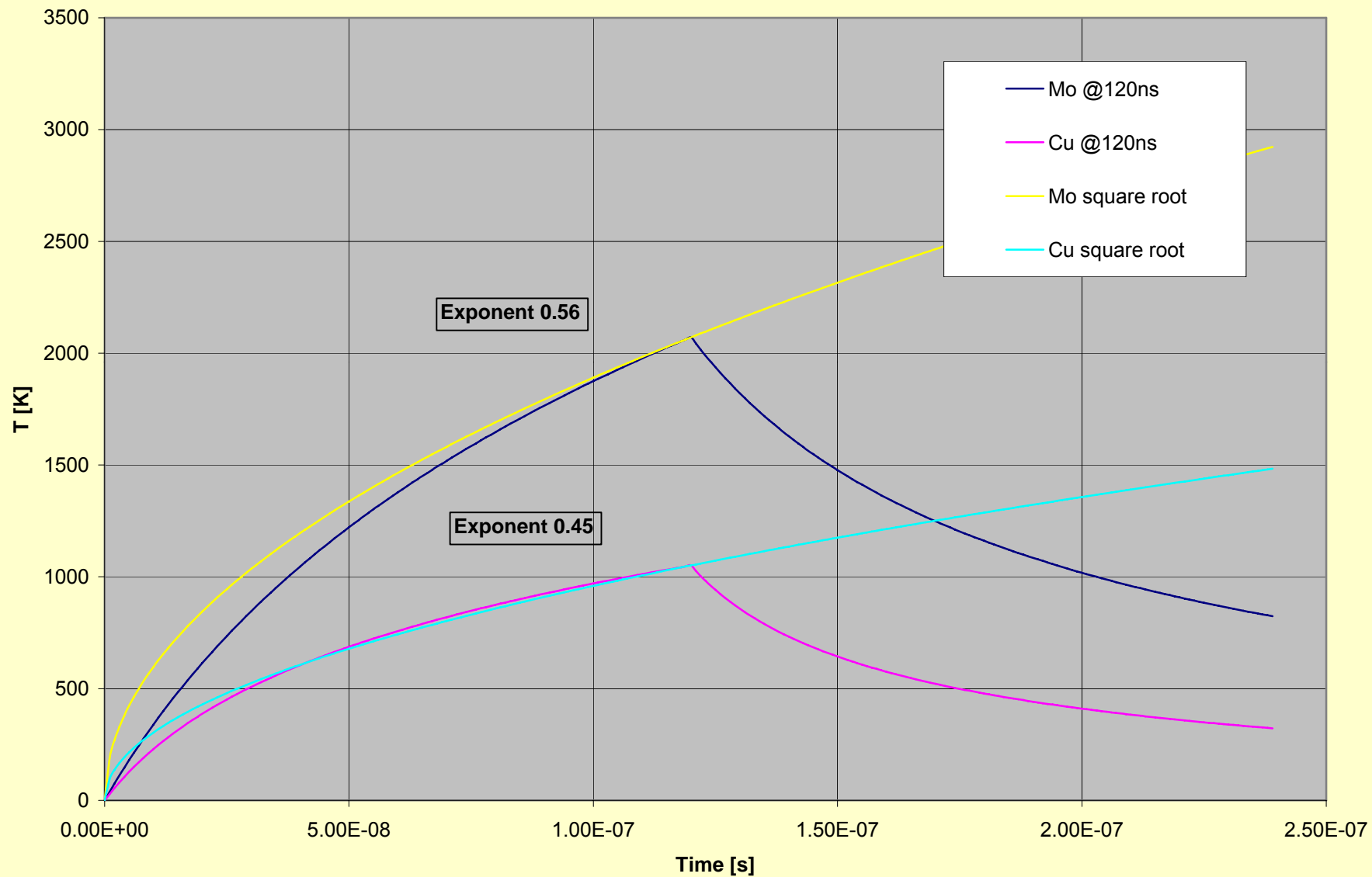


1D, 2D, 3D heating profiles inside a solid, or over a semi-infinite solid

- Clockwise:
- 1D heat flow → plane source gives square-root time dependence
- 2D heat flow → line source
- 3D heat flow → point source



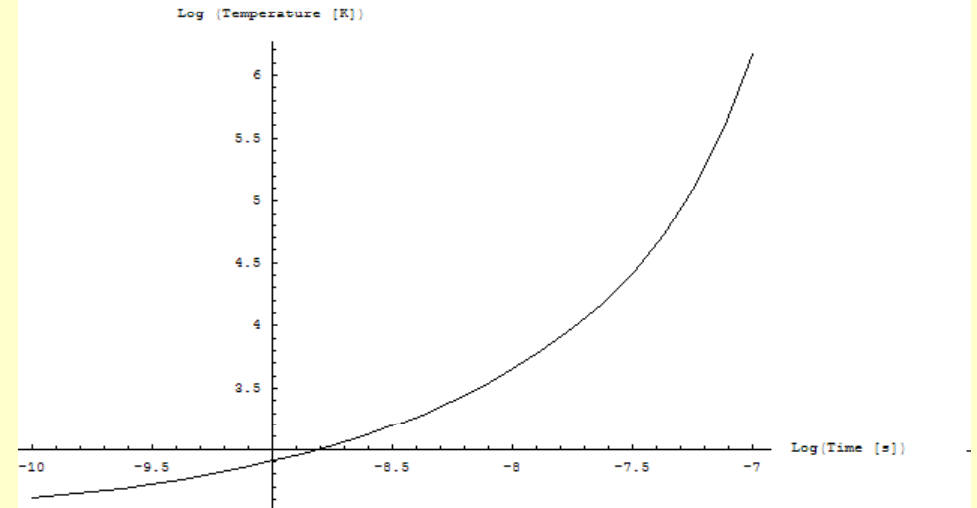
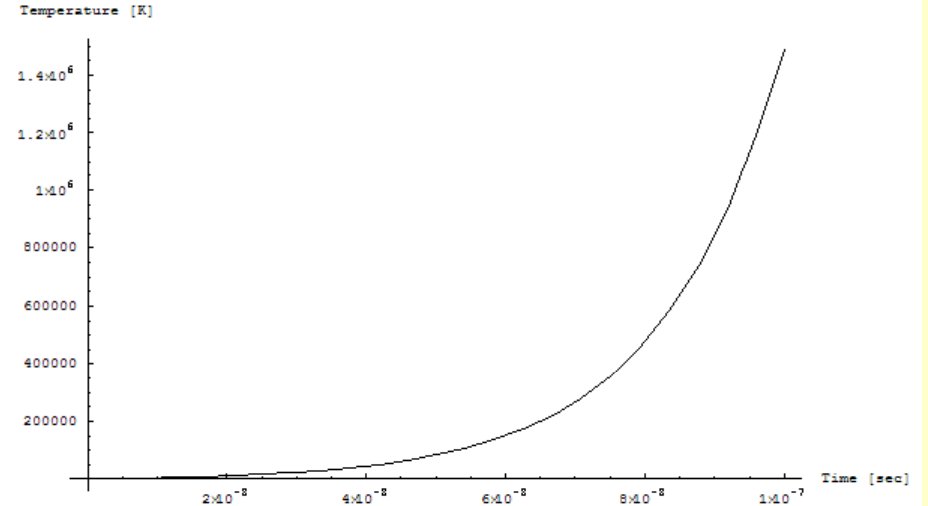
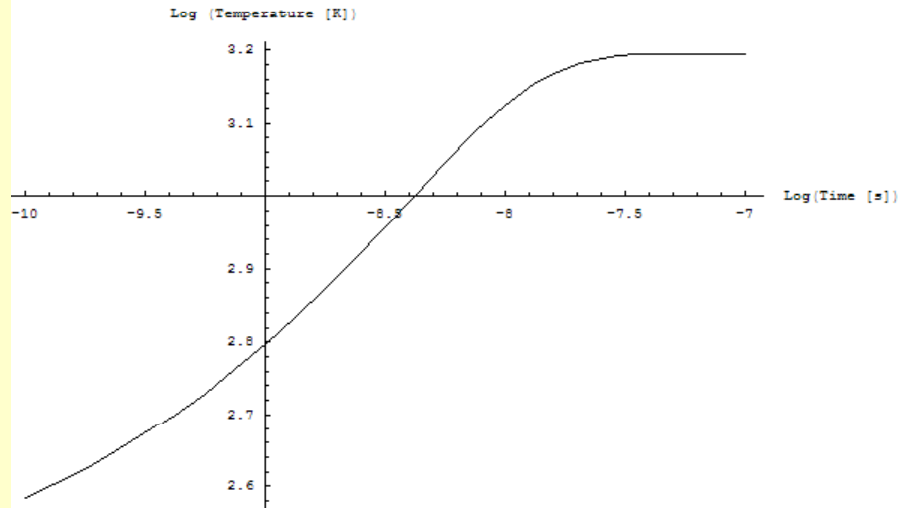
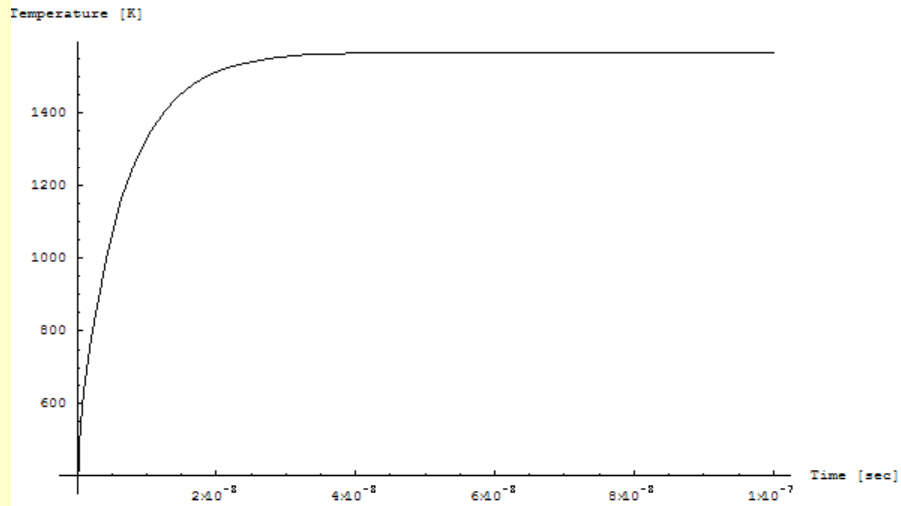
From Alessandro Bertarelli: 2 μ m x 2 μ m heat source



Simulation for Mo cone: diameter 20 nm, beta = 30

E=374 MV/m

E=378 MV/m



Simulation for Mo cone: beta = 30

**Diameter 20 nm, E=374 MV/m,
current = 0.028 A**

**Diameter 2000 nm, E=226 MV/m,
current = 2.8 A**

