# Search for Dark Matter @ the LHC: SUSY and other Searches

Albert De Roeck CERN, Geneva, Switzerland Antwerp University Belgium UG-Davis California USA IPPP, Durham UK BU, Cairo, Egypt

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The Abdus Salam International Centre for Theoretical Physics 50th Anniversary 1964–2014

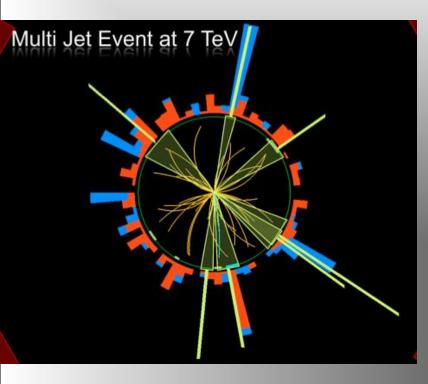
ICTP-NCP School on LHC Physics 17 – 28 November 2014 (Islamabad, Pakistan)



#### **Lecture Plan**

Overview of the 3 lectures in the next days

- Lecture 1: Searching for Physics Beyond the Standard Model: exotic signatures
- Lecture 2: The next ultimate challenge: identifying Dark Matter in the Universe, and its connection to Supersymmetry
- Lecture 3: The future program at the LHC and the studies/ideas for 'beyond the LHC'





- Introduction: Dark Matter and the WIMP miracle
- The LHC & Experiments
- The Higgs and dark matter
- Supersymmetry searches
- Generic searches via missing E<sub>T</sub>, including mono-jets, top, photons, leptons...
- New: mono-Higgs production
- Summary

#### **Dark Matter: Complementary Searches?**

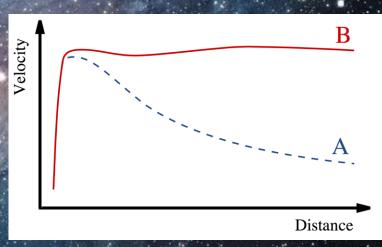
This Lecture: After the discovery of the Higgs particle @ the LHC: Dark matter is the next important physics problems to tackle for the LHC

The search is complementary to other experimental techniques used.



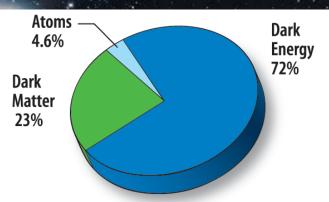
#### **Dark Matter: The Next Challenge**

Astronomers found that most of the matter in the Universe must be invisible Dark Matter



#### **'Supersymmetric' particles ?**



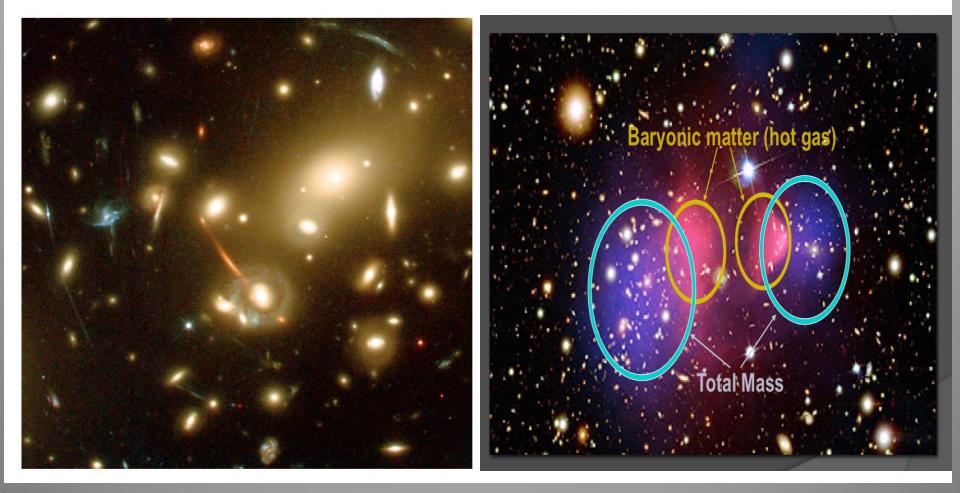


F. Zwicky 1898-1974

# **Evidence Piling-up**

- Gravitational Lensing
- much more lensing than can be explained by visible mass

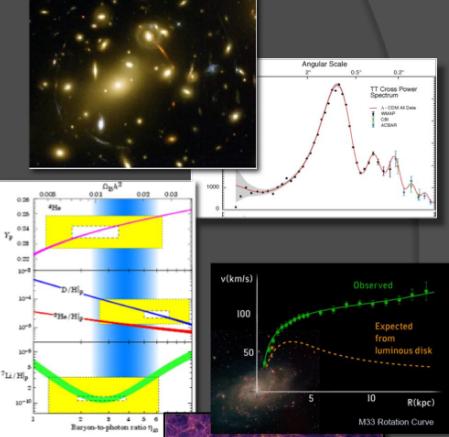
- Bullet Cluster; colliding galaxies
- Composite x-ray, visible image, 10x DM



# **Evidence Piling-up**

- There is a wide variety of evidence indicating that dark matter exists
- Each of these observations infer dark matter's presence uniquely through its gravitational influence
- To-date, no (non-controversial) observations have been made of dark matter's electroweak or other non-gravitational interactions

# Instead of dark matter, might we not understand gravity?



#### **Particle Dark Matter?**

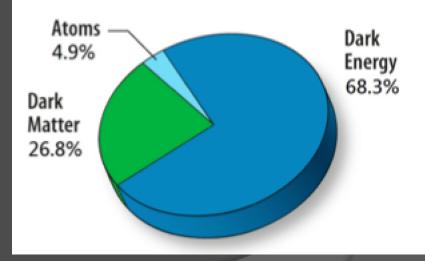
- We know only little about the nature of dark matter:
  - Cold (non-relativistic)
  - Stable
  - Dark and collisionless (no electric charge or QCD color)
- No particle contained in the Standard Model fulfills these criteria
- This leaves us with a vast range of possibilities from Planck/GUT scale "WIMPzillas" to ultra-light axions
- Dark matter candidates in the form of weakly interacting particles with masses in the GeV-TeV range (WIMPs) stand out for their
  - Testability
  - Theoretical motivation (solution to electroweak hierarchy problem)
  - The "WIMP Miracle"

The observed density of dark matter is of the magnitude expected for a thermal relic weakly-interacting massive ( $\sim$ 10-1000 GeV) particle (WIMP).

# **Particle Dark Matter?**

# The Dark Matter Candidate Zoo

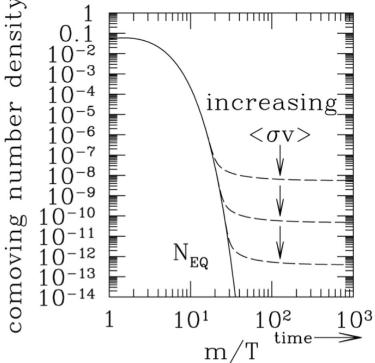
- Neutralinos (higgsino, bins, winos, singlinos)
- Axinos
- Gravitinos
- Sneutrinos
- Axions
- Sterile neutrinos
- 4<sup>th</sup> generation neutrinos
- Kaluza-Klein photons
- Kaluza-Klein gravitons
- Brane world dark matter/D-matter
- Little higgs dark matter
- Light scalars
- Superheavy states (*ie.* "WIMPzillas")
- Self-interacting dark matter
- Super-WIMPs
- Asymmetric dark matter
- Q-balls (and other topological states)
- CHAMPs (charged massive particles)
- Cryptons, mirror matter, and many, many, many others...



From D. Hooper

#### Weakly Interaction Massive Particles (WIMPs)

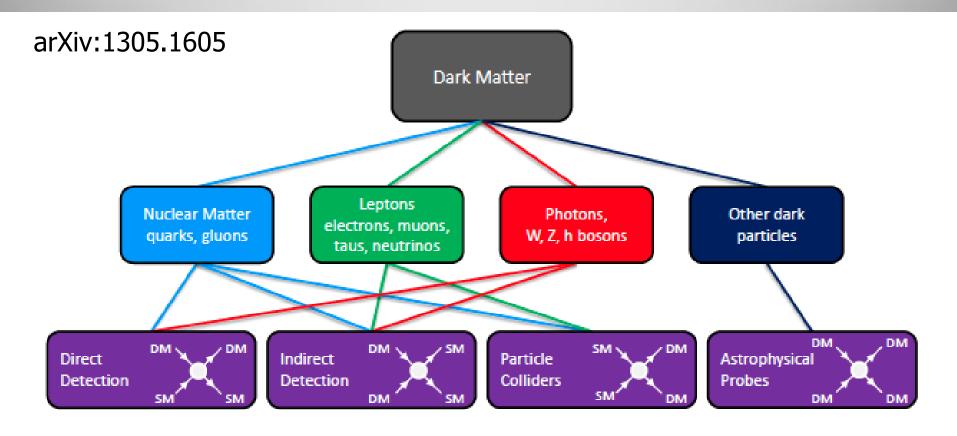
- Perhaps Dark Matter is a particle with weak-scale mass?
  - Weakly Interacting Massive Particles (WIMPs)
  - Produced in the Big Bang, interact via  $\chi + \chi \rightarrow q + q$
- As the universe expands and the temperature drops...
  - WIMPs become diluted, interact less often and 'freeze out'.
  - Higher cross-section (<σv>) yields lower relic density



Weakly-interacting massive particles naturally provide the right relic abundance - "WIMP miracle"

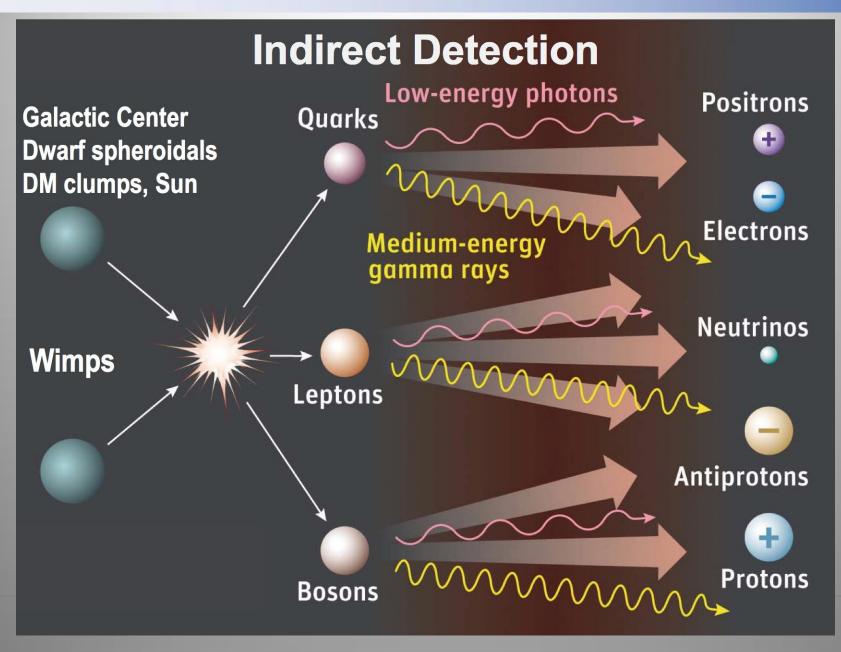
#### **Dark Matter @ LHC?**

Search for WIMP candidates in events with Missing Transverse Momentum EG: SUSY searches, monojet and mono-photon searches, W' searches...

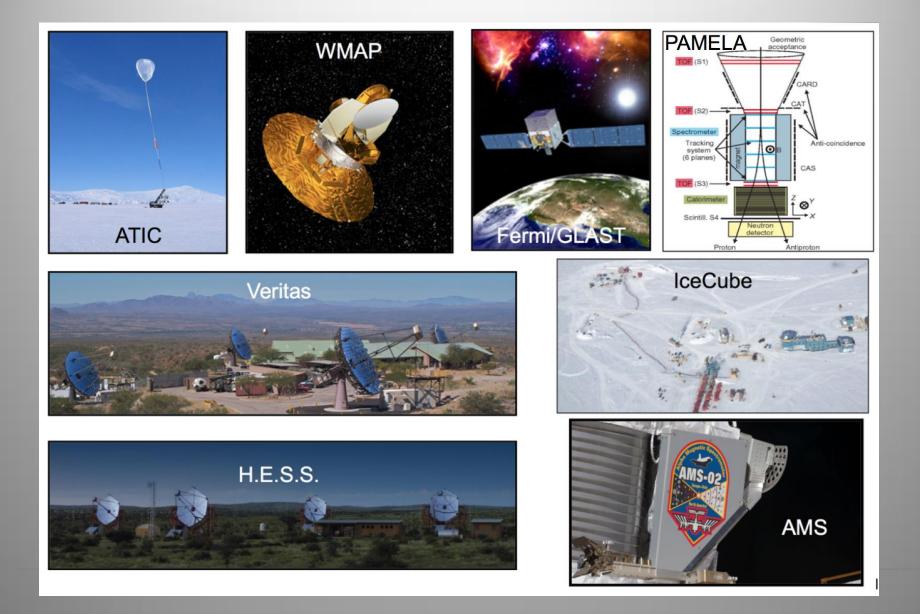


+ CAST experiment, searching for axion DM

#### **Dark Matter: Indirect Detection**



#### **Indirect Detection Experiments**



#### **Indirect Detection**

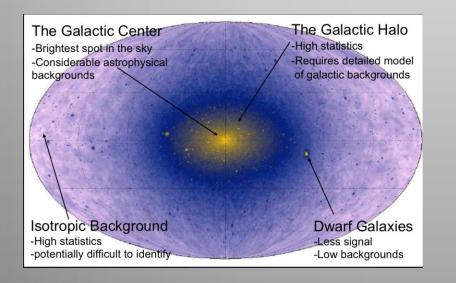
#### Some scientists are believers!!

#### arXiv:1402.6703v1

#### The Characterization of the Gamma-Ray Signal from the Central Milky Way: A Compelling Case for Annihilating Dark Matter

Tansu Daylan,<sup>1</sup> Douglas P. Finkbeiner,<sup>1,2</sup> Dan Hooper,<sup>3,4</sup> Tim Linden,<sup>5</sup> Stephen K. N. Portillo,<sup>2</sup> Nicholas L. Rodd,<sup>6</sup> and Tracy R. Slatyer<sup>6,7</sup>

# Using gamma-ray data from the FERMI satellite DM annihilation into b anti-b quarks?

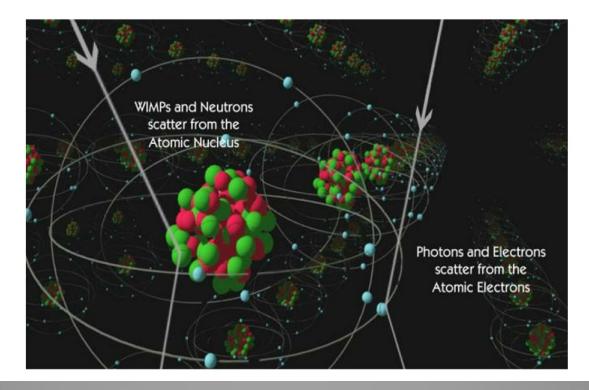




#### Also the 3.5 KeV line: light axion-like particle annihilation?

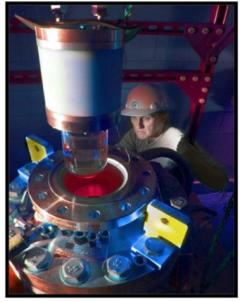
#### **Dark Matter: Direct Detection**

- Direct detection experiments: nuclear recoil from DM collision
  - Extremely sensitive, extremely difficult... extremely successful!
  - Excesses observed but not confirmed (10 GeV DM candidate?)
- Need for independent verification from non-astrophysical experiments
  - Low mass region not accessible to direct detection experiments
  - Limited by threshold effects, energy scale, bkgnds; spin-dependent couplings difficult...



#### **Direct Detection: Examples**

COUPP



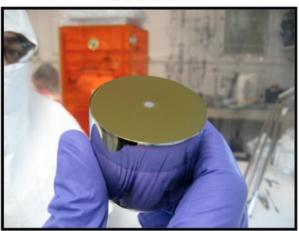
CoGeNT



CRESST



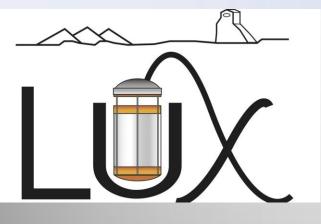
Xenon



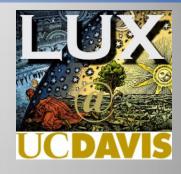
(+ EDELWEISS, DAMA, EURECA, ZEPLIN, DEAP, ArDM, WARP, LUX, SIMPLE, PICASSO, DMTPC, DRIFT, KIMS, ...)

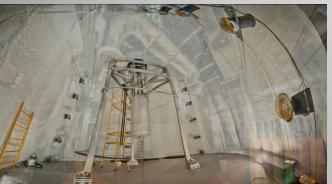


#### **Direct Searches for Dark Matter**



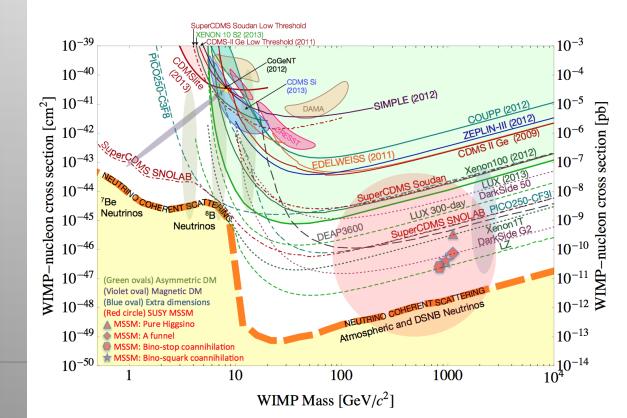
State of the art today: Driven by the results of the LUX experiment





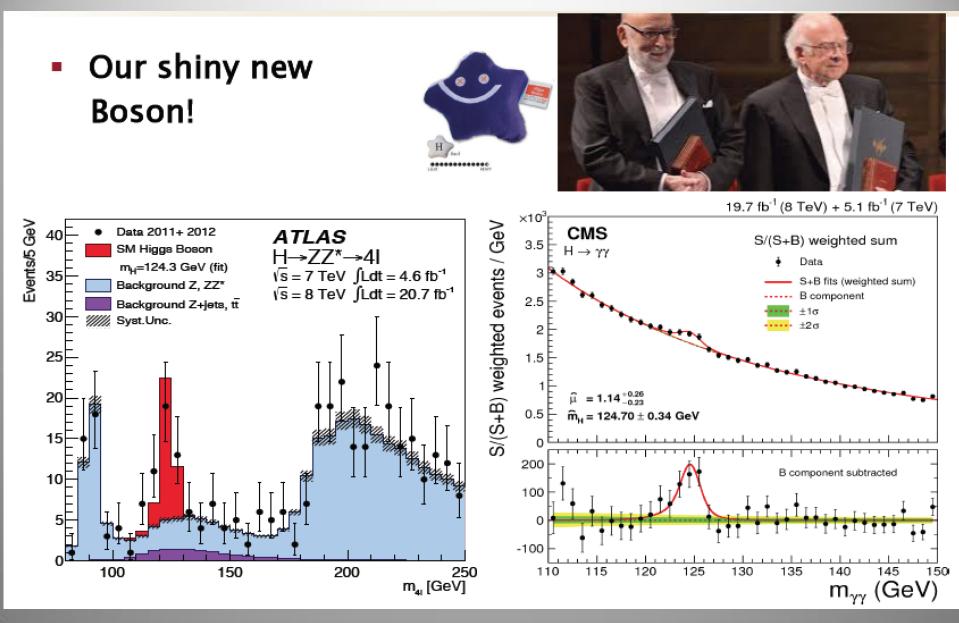
Intensive campaign of direct detection experiments since more than ~20 years

No (real) sign so far...

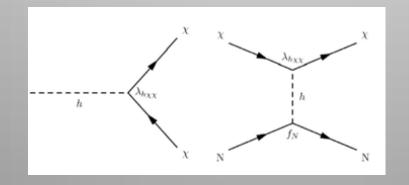


The LHC

#### **The Higgs Discovery**

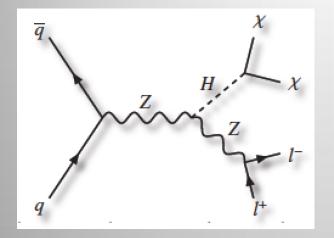


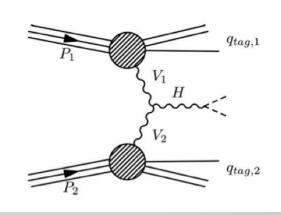
#### **Dark Matter and the Higgs**



"higgs portal models" Eg: arXiv:1205.3169

#### **Invisible Higgs Decay Channel**





#### CMS-PAS-HIG-13-005

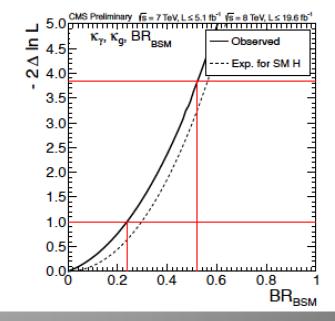
 Possible decay of the Higgs in Dark Matter particles (if M<M<sub>H</sub>/2)
 Different searches:

-Direct search

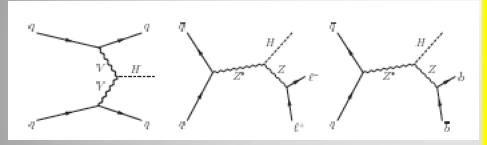
Look for the invisible decay channels

-Indirect search

Make a global fit of all production and decays (and some modest assumptions)

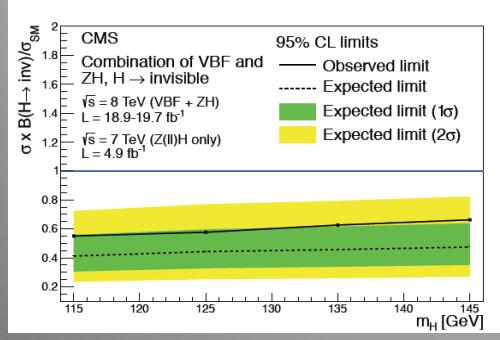


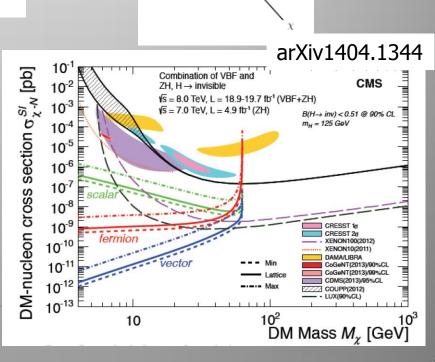
### **Invisible Higgs Decay Channel**



Search for invisible Higgs decays using  $Z+H \rightarrow 2$  leptons + missing  $E_T$ VBF H  $\rightarrow 2$  jets + missing  $E_T$ Possible decay in Dark Matter particles (if M<M<sub>H</sub>/2): Higgs Portal Models

Combined result from the three channels  $BR(H \rightarrow invisible) < 58\%(44\% exp)$  at 95% CL. for a Higgs with a mass of 125 GeV

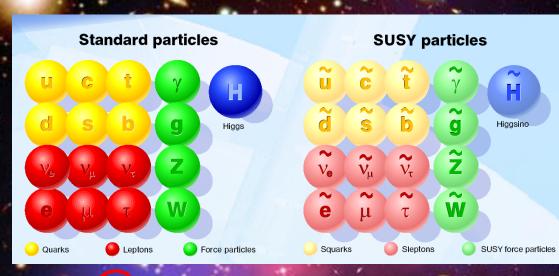


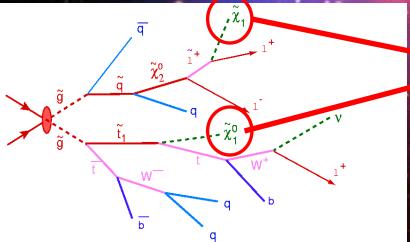


### **Searches for Supersymmetry**

Supersymmetry was not "invented" to solve the dark matter problem, but can provide a great solution! WIMP Dark Matter candidate comes for free

#### Supersymmetry: a new symmetry in Nature?



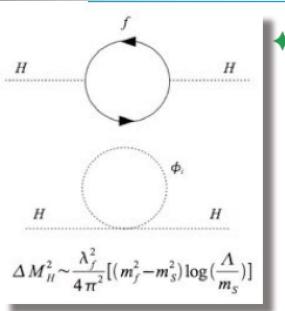


SUSY particle production at the LHC

Candidate particles for Dark Matter  $\Rightarrow$  Produce Dark Matter in the lab

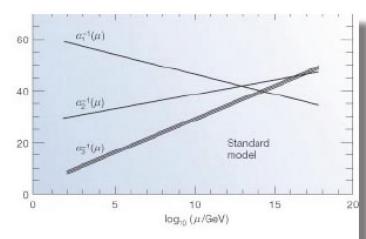
Picture from Marusa Bradac

#### Summary: Why SUSY is good for you!!

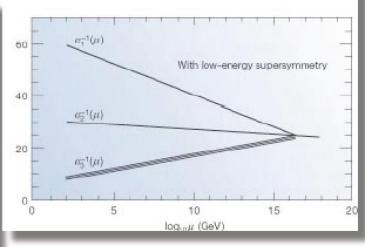


Elegant solution to the hierarchy problem (i.e., why the Higgs mass is not at the Planck scale)

#### Gauge unification

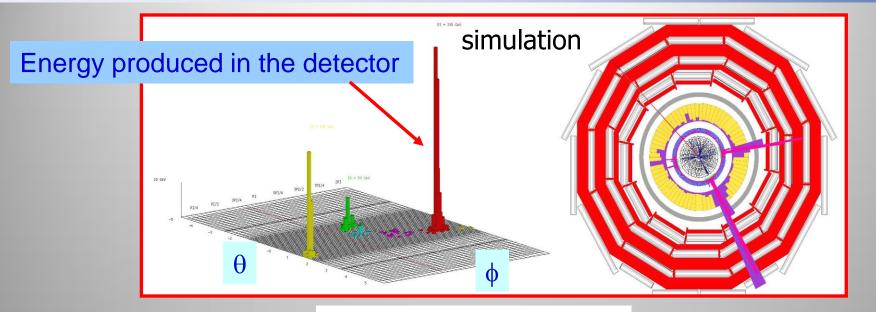


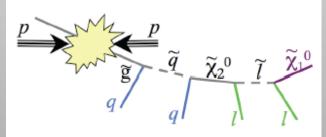




#### Dark matter candidate with the right abundance

#### **Detecting Supersymmetric Particles**





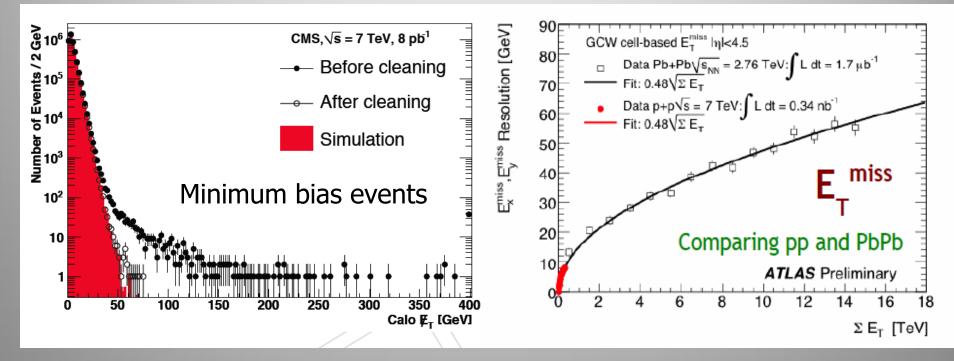
Supersymmetric particles decay and produce a cascade of jets, leptons and missing transverse energy (MET) due to escaping 'dark matter' particle candidates

Very prominent signatures in CMS and ATLAS

# **Missing Transverse Energy**

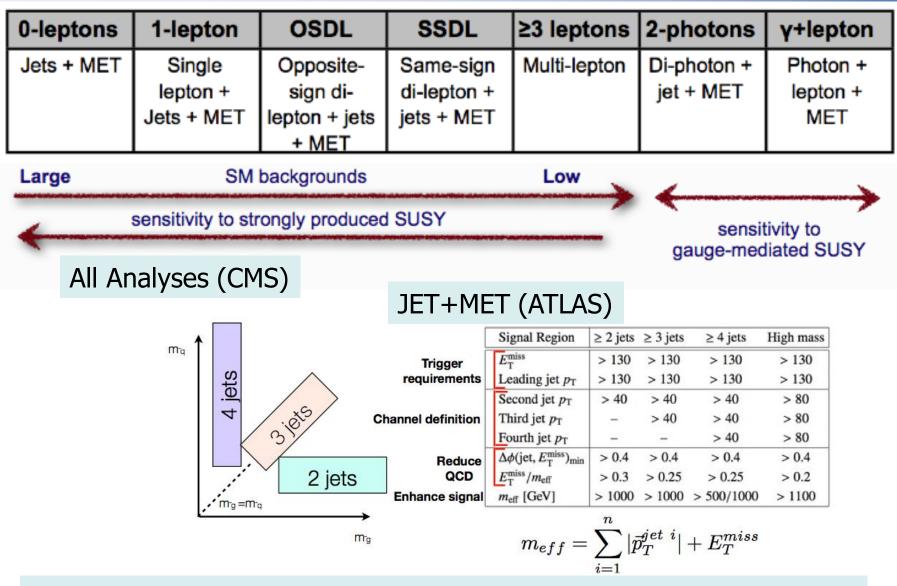
Total transverse momentum imbalance

Generally appreciated to be a difficult quantity to measure Very sensitive to fluctuations, miss-measurements, noise, backgrounds



In practice, rather well under control, from the start
Good resolution using 'particle flow' ie maximally identifying particles
More pile-up in future will NOT make this simpler

#### **SUSY Searches**



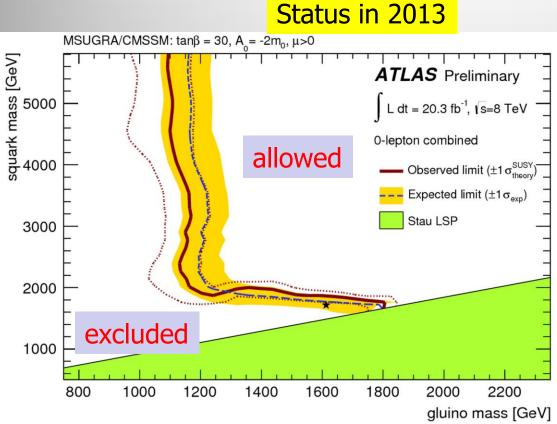
Note: Strong effort to get background (tail) estimates from data itself

#### **Example: Search for SUSY**

Take one example to show steps involved:

- Define event selection criteria
- Go through ~2.000.000.000 events triggered and stored on-line, to select candidates
- Use eg kinematical cuts to suppress background
- "Predict" backgrounds in signal region
- Determine efficiencies and systematics
- Open the data box: Excess or no excess?

#### SUSY Searches: No signal yet to date...



•So far NO clear signal of supersymmetric particles has been found

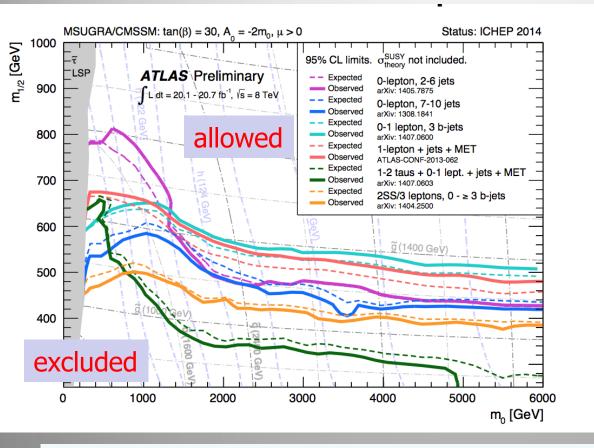
•We can exclude regions where the new particles could exist.

 Searches will continue for the higher energy in 2015

Plenty of searches ongoing: with jets, leptons, photons, W/Z, top, Higgs, with and without large missing transverse energy Also special searches for contrived model regions

#### **Constrained MSSM: Various Studies**

Status in 2014



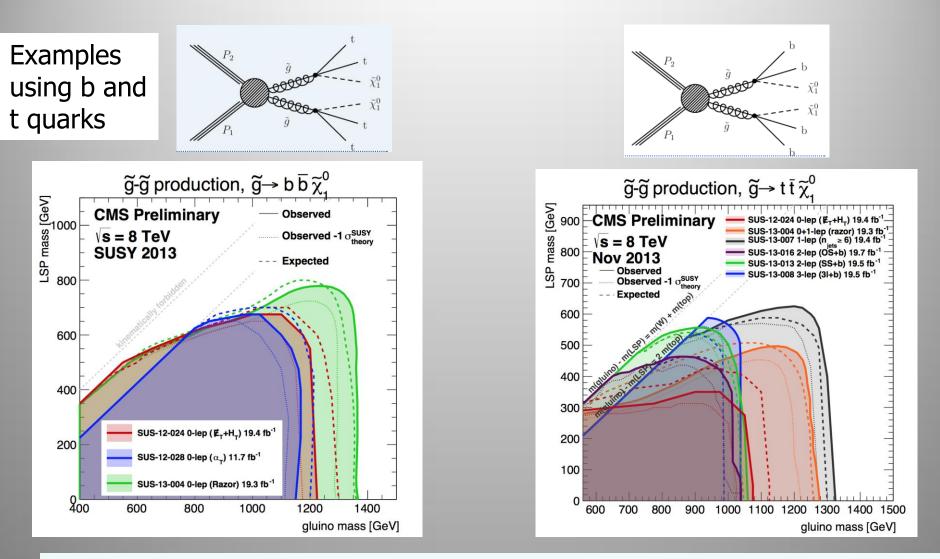
•So far NO clear signal of supersymmetric particles has been found

•We can exclude regions where the new particles could exist.

•m<sub>1/2</sub>: universal gaugino mass at GUT scale
•m<sub>0</sub>: universal scalar mass at GUT scale

Plenty of searches ongoing: with jets, leptons, photons, W/Z, top, Higgs, with and without large missing transverse energy Also special searches for contrived model regions

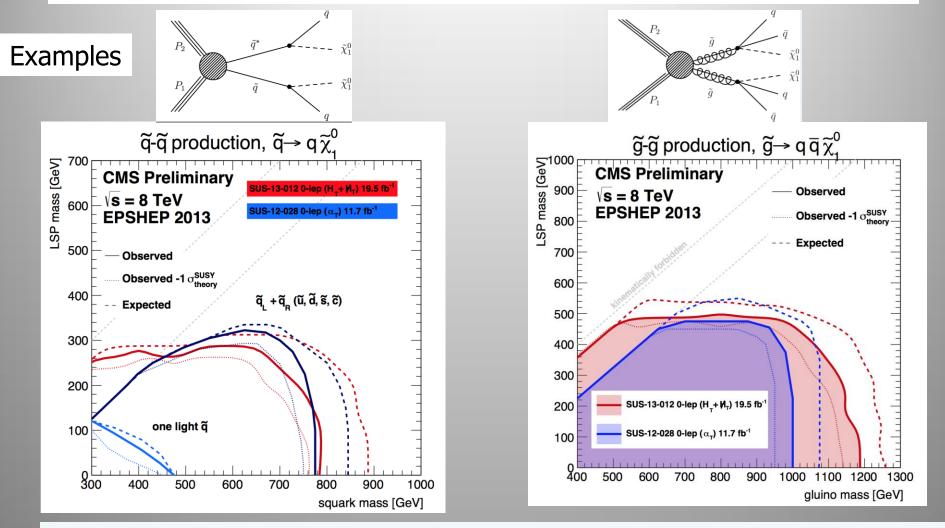
#### **Limits on Squarks and Gluinos**



#### Combined limits typically > 1-1.5 TeV on sparticle masses

#### **Limits on Squarks and Gluinos**

Results depend on the topologies studies, assumed mass of the LSP etc.



Combined limits typically > 1-1.5 TeV on sparticle masses

### What is really needed from SUSY?

#### End 2011: Revision!

N. Arkani-Ahmed CERN Nov 2011

Papucci, Ruderman, Weiler arXiv:1110.6926

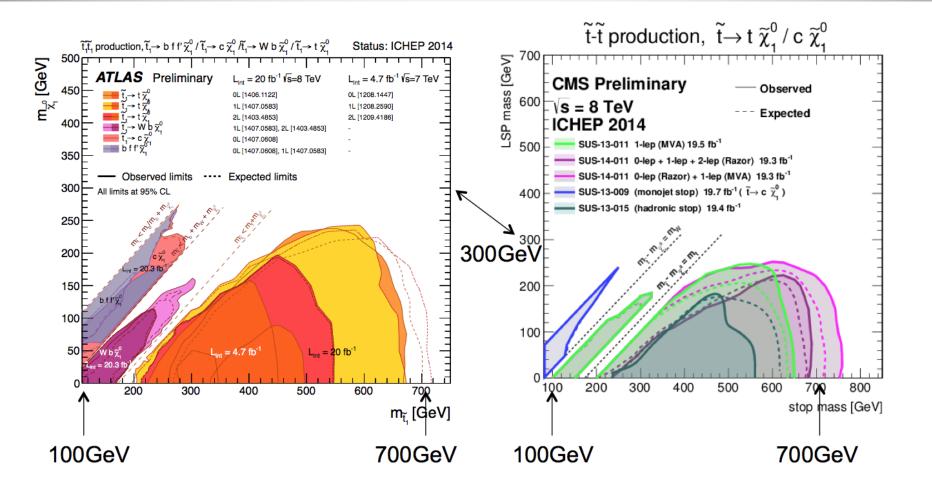
LHC data end 2011 Stops > 200-300 GeV Gluino > 600-800 GeV

Moving away from constrained SUSY models to 'natural' models

Natural SUSY survived LHC so far, but we are getting close to push it to its limits!

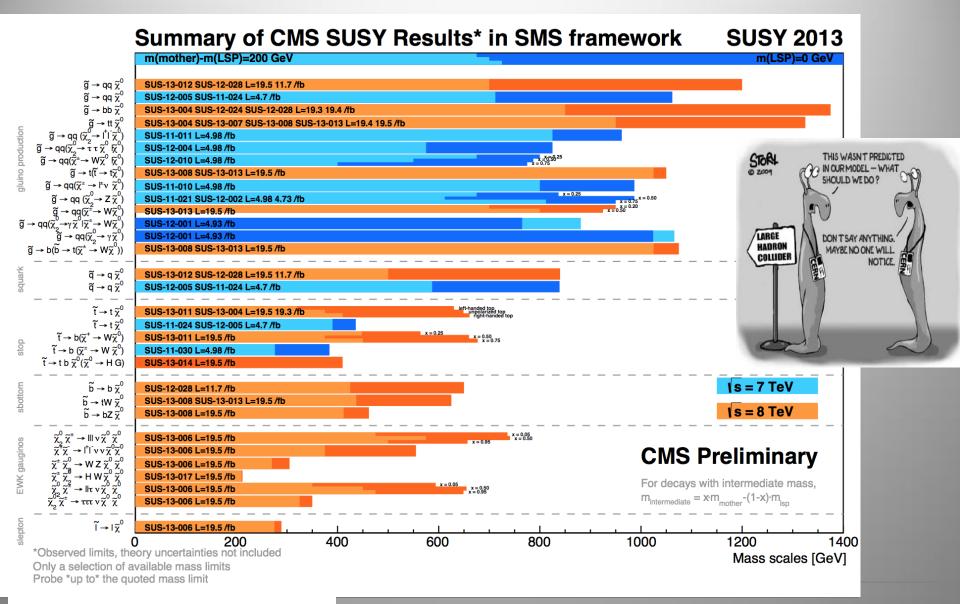
Cumpulsory Natural SUSY 1500 tL,R,b 400 120 Unavoidable tunings:  $\left(\frac{400}{m_{1}^{2}}\right)^{2}$ ,  $\left(\frac{4m_{1}^{2}}{M_{q}^{2}}\right)^{2}$ 

#### **Natural SUSY?**



Searches for stop quarks are pushing limits now to 700 GeV

# **Summary of SUSY Searches**



#### Similar table for ATLAS

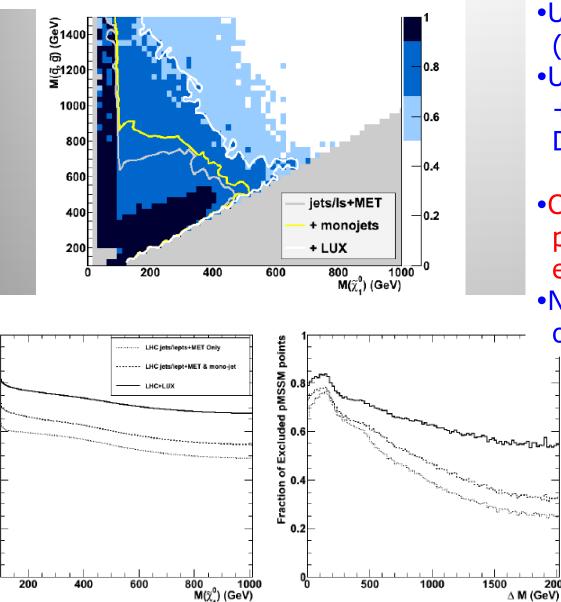
#### **Dark Matter SUSY Space Left?**

#### arXiv:1311.7641

Fraction of Excluded pMSSM points

0.4

0.2



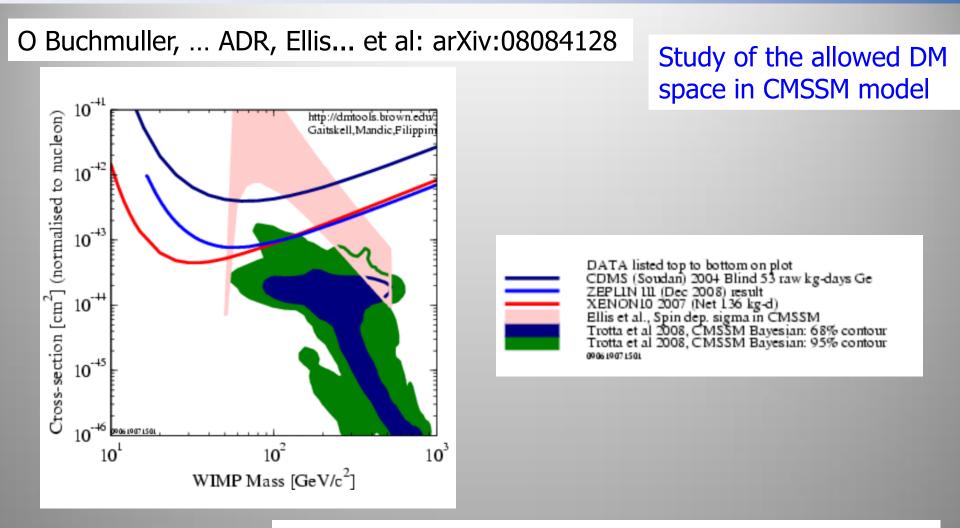
- Use the pMSSM SUSY model (19 parameters)
- Use all the ATLAS SUSY Data + mono-jet searches + LUX **DM** results
- Check what fraction of pMSSM solutions that is excluded.

2000

 No real full systematic study of the SUSY space yet

> 60-80% of the solutions excluded as function of the neutralino mass

#### **Constrained SUSY Models**

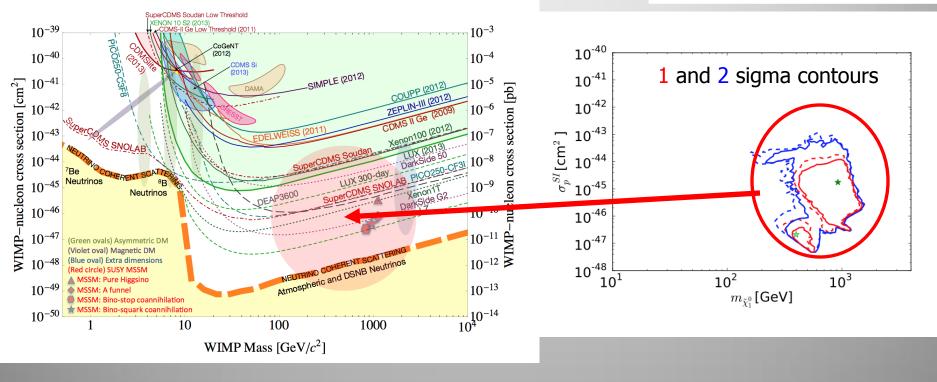


Region allowed in the CMSSM includes constraints of the Run-I LHC searches (SUSY) and precision data, g-2, cold dark matter constraints...

#### **Constrained SUSY Models**

O Buchmuller, ... ADR, Ellis... et al: arXiv:13125250

Study of the allowed DM space in CMSSM model

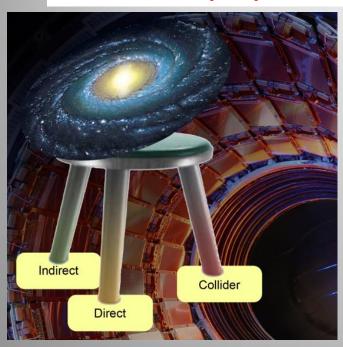


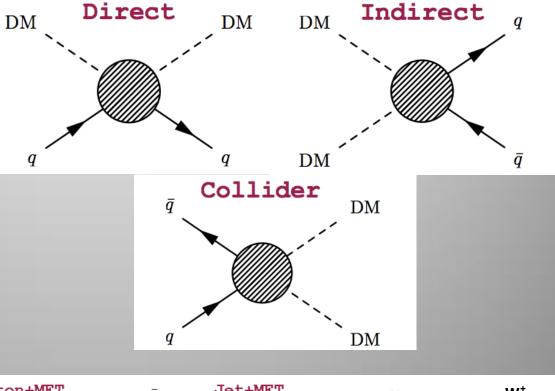
Region allowed in the CMSSM includes constraints of the Run-I LHC searches (SUSY) and precision data, g-2, cold dark matter constraints...

#### **General Searches for Dark Matter**

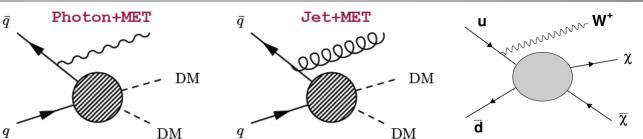
### **The Generic Dark Matter Connection**

Searches for mono-jets and mono-photons can be used to search for Dark Matter (DM)





Use effective theory or simplified models to relate measurements to Dark Matter studies



# **Mono-object Searches in CMS**

• Mono-jets: Generally the most powerful

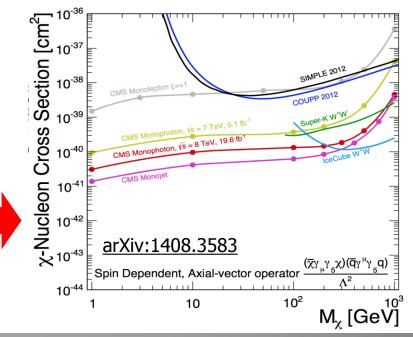
gluor

- Mono-photons: First used for dark matter Searches
- Mono-Ws: Distinguish dark matter couplings to u- and dtype of quarks
- Mono-Zs: Clean signature
- Mono-Tops: Couplings to tops
- Mono-Higgs: Higgs-portals
- Higgs Decays?

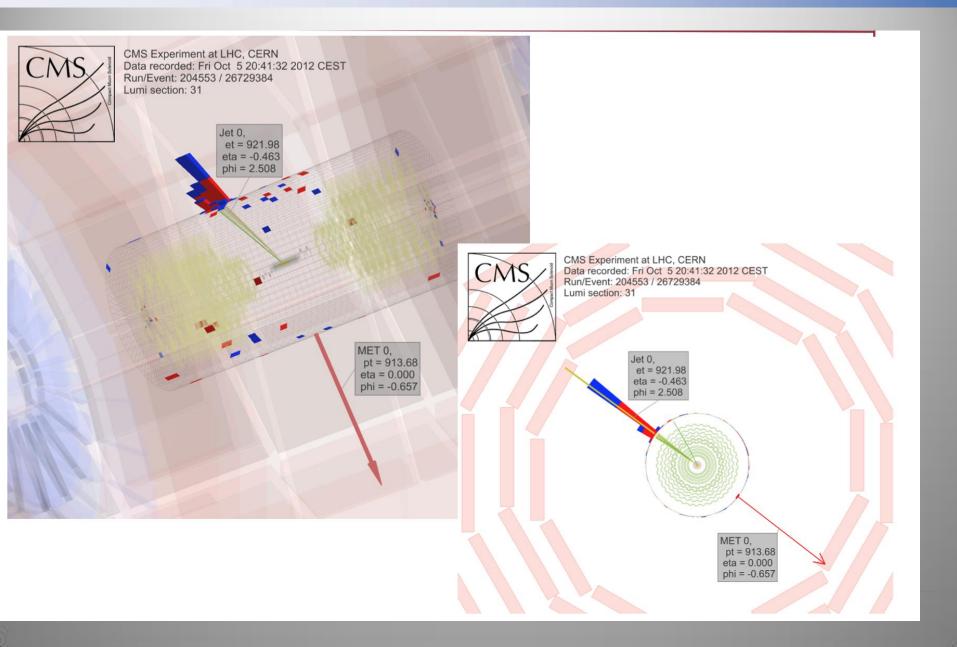
**Example Monojets** 

**Dark Matter?** 

Effective Field Theories for DM interpretation are under attack! Alternatives like SMS proposed...



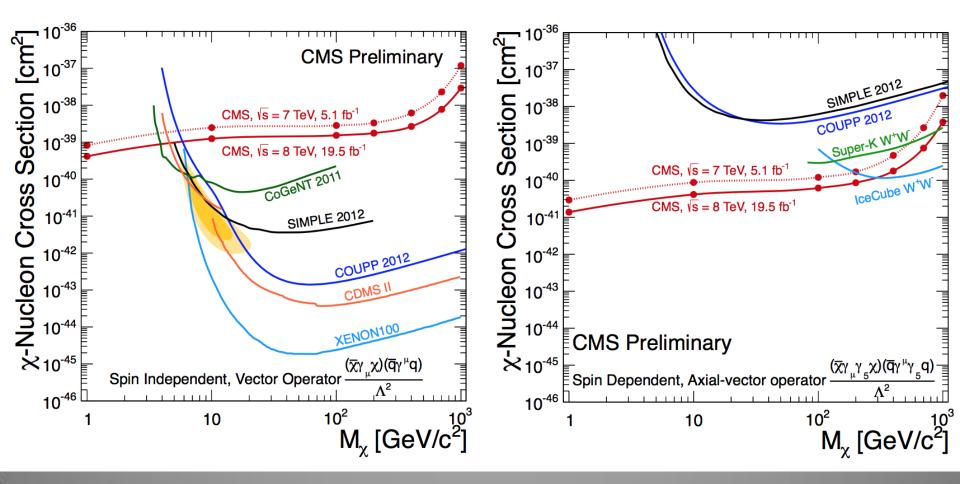
#### **Mono-Jet Event**



#### **Results for Mono-jets**

[CMS EXO-12-048]

- Derived EFT limits then compared to direct-detection experiments
- CMS results improved with 8 TeV (higher E, more data)



### Summary

- Dark Matter is an important open point in fundamental physics right now and the LHC data can contribute to the quest.
- The Higgs particle may couple to DM or may even decay into it. Invisible decays and deviations from SM Higgs couplings are explored
- Supersymmetry scenarios with R<sub>P</sub>-conservation can have a natural DM candidate. Discovery of supersymmetry will have important impact on DM
- Generic searches for DM in analyses dealing with missing E<sub>T</sub>: typically mono-object searches

WFLCOMF.

• So far exclusion limits only, but maybe soon: