

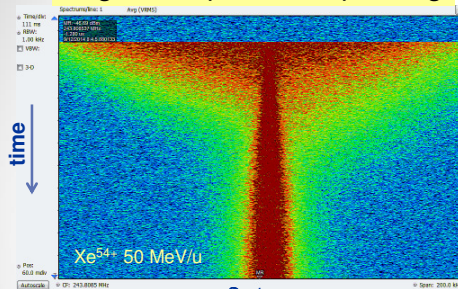
Beam Cooling

M. Steck, GSI, Darmstadt

CAS, Warsaw,
27 September – 9 October, 2015

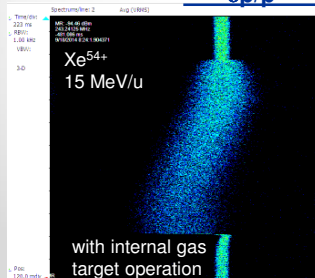
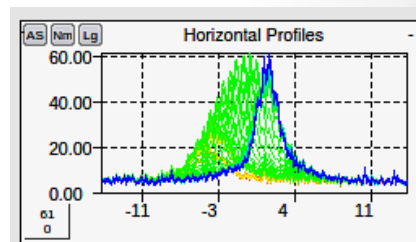
Cooling

longitudinal (momentum) cooling



injection into the storage ring

transverse cooling



cooling **off**

heating (spread) and
energy loss (shift)

cooling **on**

cooling:
good energy definition
small beam size
⇒ highest precision

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Beam Cooling

Introduction

1. Electron Cooling
2. Ionization Cooling
3. Laser Cooling
4. Stochastic Cooling

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Beam Cooling

Beam cooling is synonymous for a reduction of beam temperature. Temperature is equivalent to terms as phase space volume, emittance and momentum spread.

Beam Cooling processes are not following Liouville's Theorem:
`in a system where the particle motion is controlled by external conservative forces the phase space density is conserved'
(This neglects interactions between beam particles.)

Beam cooling techniques are non-Liouvillian processes which violate the assumption of a conservative force.

e.g. interaction of the beam particles with other particles
(electrons, photons, matter)

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Cooling Force

Generic (simplest case of a) cooling force:

$$F_{x,y,s} = -\alpha_{x,y,s} v_{x,y,s}$$

$v_{x,y,s}$ velocity in the rest frame of the beam

non conservative, cannot be described by a Hamiltonian

For a 2D subspace distribution function $f(z, z', t)$

$$F_z = -\alpha_z v_z \quad z = x, y, s \quad v_z = v_0 \cdot z'$$

$$\frac{df(z, z', t)}{dt} = -\lambda_z f(z, z', t) \quad \lambda_z \text{ cooling (damping) rate}$$

in a circular accelerator:

Transverse (emittance) cooling rate

$$\epsilon_{x,y}(t_0 + t) = \epsilon_{x,y}(t_0) e^{-\lambda_{x,y} t}$$

Longitudinal (momentum spread) cooling rate

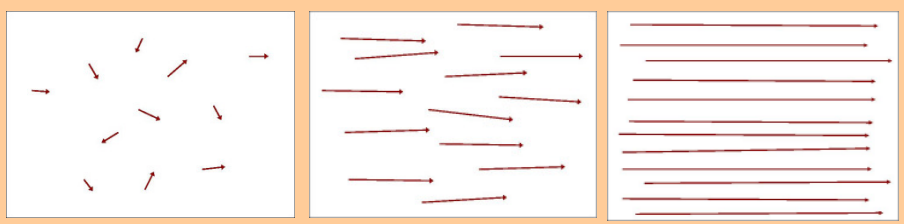
$$\frac{\delta p_{\parallel}}{p_0}(t_0 + t) = \frac{\delta p_{\parallel}}{p_0}(t_0) e^{-\lambda_{\parallel} t}$$

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Beam Temperature

Where does the beam temperature originate from?

The beam particles are generated in a 'hot' source



at rest (source)

at low energy

at high energy

In a standard accelerator the beam temperature is not reduced
(thermal motion is superimposed the average motion after acceleration)

but: many processes can heat up the beam

e.g. heating by mismatch, space charge, intrabeam scattering,
internal targets, residual gas, external noise

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Beam Temperature Definition

Longitudinal beam temperature

$$\frac{1}{2}k_B T_{\parallel} = \frac{1}{2}mv_{\parallel}^2 = \frac{1}{2}mc^2\beta^2\left(\frac{\delta p_{\parallel}}{p}\right)^2$$

Transverse beam temperature

$$\frac{1}{2}k_B T_{\perp} = \frac{1}{2}mv_{\perp}^2 = \frac{1}{2}mc^2\beta^2\gamma^2\theta_{\perp}^2 \quad \theta_{\perp} = \frac{v_{\perp}}{\beta c}, \quad \theta_{\perp}(s) = \sqrt{\frac{\epsilon}{\beta_{\perp}(s)}}$$

Distribution function

$$f(v_{\perp}, v_{\parallel}) \propto \exp\left(-\frac{mv_{\perp}^2}{2k_B T_{\perp}} - \frac{mv_{\parallel}^2}{2k_B T_{\parallel}}\right)$$

Particle beams can be anisotropic: $k_B T_{\parallel} \neq k_B T_{\perp}$

e.g. due to laser cooling or the distribution of the electron beam

Don't confuse: beam energy \leftrightarrow beam temperature

(e.g. a beam of energy 100 GeV can have a temperature of 1 eV)

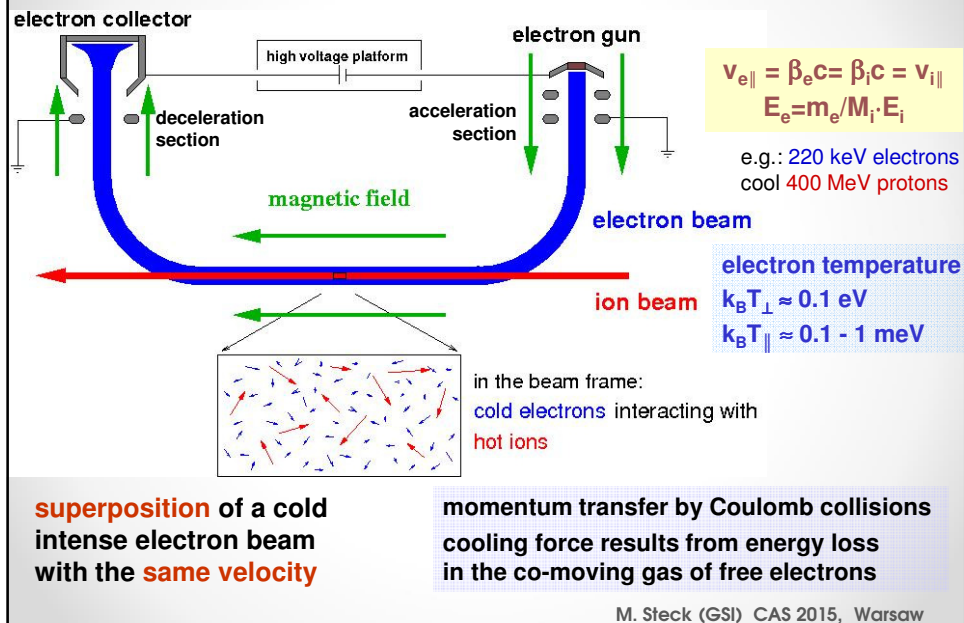
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Benefits of Beam Cooling

- Improved beam quality
 - Precision experiments
 - Luminosity increase
- Compensation of heating
 - Experiments with internal target
 - Colliding beams
- Intensity increase by accumulation
 - Weak beams from the source can be enhanced
 - Secondary beams (antiprotons, rare isotopes)

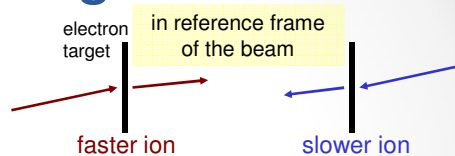
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1. Electron Cooling



Simple Derivation of the Electron Cooling Force

Analogy: energy loss in matter (electrons in the shell)



Rutherford scattering: $2 \tan\left(\frac{\theta}{2}\right) = \frac{2Z_1 Z_2 e^2}{4\pi\epsilon_0 \Delta p v b}$ $Z_1 = Q$ (ion), $Z_2 = -1$ (electron)

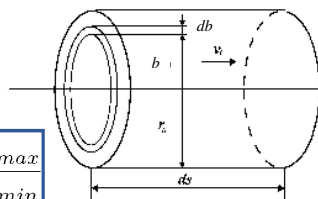
Energy transfer: $\Delta E(b) = \frac{(\Delta p)^2}{2m_e} \simeq \frac{2Q^2 e^4}{(4\pi\epsilon_0)^2 m_e v^2} \frac{1}{b^2}$ (for $b \gg b_{min}$)

Minimum impact parameter: $b_{min} = \frac{Qe^2}{(4\pi\epsilon_0)^2 m_e v^2}$
 from: $\Delta E(b_{min}) = \Delta E_{max} \simeq 2m_e v^2$

Energy loss:

$$-\frac{dE}{dx} = 2\pi \int_{b_{min}}^{b_{max}} b n_e \Delta E db = \frac{4\pi Q^2 e^4}{(4\pi\epsilon_0)^2 m_e v^2} n_e \ln \frac{b_{max}}{b_{min}}$$

Coulomb logarithm $L_C = \ln(b_{max}/b_{min}) \approx 10$ (typical value)



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Characteristics of the Electron Cooling Force

$$\vec{F}(\vec{v}_i) = -\frac{4\pi Q^2 e^4 n_e}{(4\pi\epsilon_0)^2 m_e} \int L_C(\vec{v}_{rel}) f(\vec{v}_e) \frac{\vec{v}_{rel}}{v_{rel}^3} d^3 v_e$$

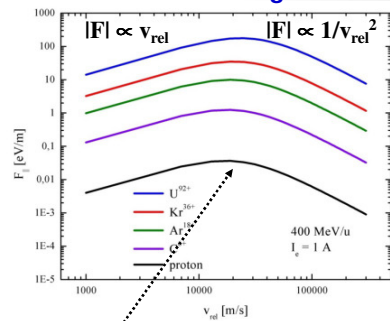
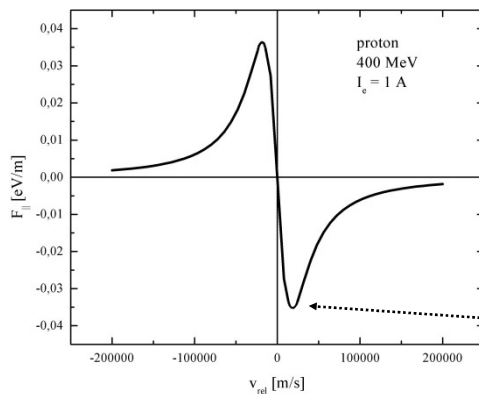
$$\vec{v}_{rel} = \vec{v}_i - \vec{v}_e$$

cooling force F

for small relative velocity: $\propto v_{rel}$

for large relative velocity: $\propto v_{rel}^{-2}$

increases with charge: $\propto Q^2$



maximum of cooling force
at effective electron temperature

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Electron Cooling Time

first estimate:
(Budker 1967)

$$\tau = \frac{3}{8\sqrt{2}\pi n_e Q^2 r_e r_i c L_C} \left(\frac{k_B T_e}{m_e c^2} + \frac{k_B T_i}{m_i c^2} \right)^{3/2}$$

for large relative velocities

$$\text{cooling time } \tau_z \propto \frac{A}{Q^2} \frac{1}{n_e \eta} \beta^3 \gamma^5 \theta_z^3 \begin{cases} \theta_{x,y} = \frac{v_{x,y}}{\gamma \beta c} \\ \theta_{\parallel} = \frac{v_{\parallel}}{\gamma \beta c} \end{cases}$$

cooling rate (τ^{-1}):

- slow for hot beams $\propto \theta^3$
- decreases with energy $\propto \gamma^2$ ($\beta \cdot \gamma \cdot \theta$ is conserved)
- linear dependence on electron beam intensity n_e and cooler length $\eta = L_{ec}/C$
- favorable for highly charged ions Q^2/A
- independent of hadron beam intensity

for small relative velocities

cooling rate is constant and maximum at small relative velocity

$$F \propto v_{rel} \Rightarrow \tau = \Delta t = p_{rel}/F = \text{constant}$$

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Models of the Electron Cooling Force

- **binary collision model**

description of the cooling process by successive collisions of two particles and integration over all interactions
analytic expressions become very involved, various regimes (multitude of Coulomb logarithms)

- **dielectric model**

interaction of the ion with a continuous electron plasma
(scattering off of plasma waves)
fails for small relative velocities and high ion charge

- **an empiric formula (Parkhomchuk) derived from experiments:**

$$\vec{F} = -4 \frac{n_e}{m_e} \frac{(Qe^2)^2}{(4\pi\epsilon_0)^2} \ln\left(\frac{b_{max} + b_{min} + r_c}{b_{min} + r_c}\right) \frac{\vec{v}_{ion}}{(v_{ion}^2 + v_{eff}^2)^{3/2}}$$

$$b_{min} = \frac{Qe^2/4\pi\epsilon_0}{m_e v_{ion}^2}; \quad b_{max} = \frac{v_{ion}}{\min(\omega_{pe}, 1/T_{cool})}, \quad v_{eff}^2 = v_{e,\parallel}^2 + v_{e,\perp}^2$$

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Electron Beam Properties

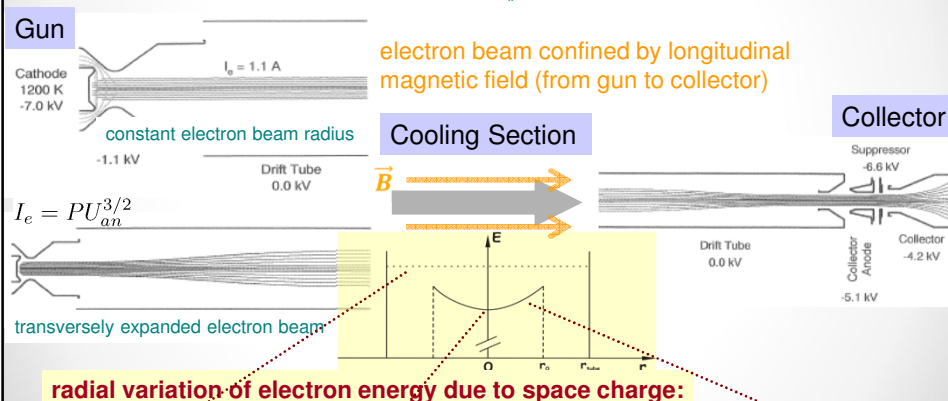
electron beam temperature

transverse $k_B T_{\perp} = k_B T_{cat}$, with transverse expansion ($\propto B_c/B_{gun}$)

longitudinal $k_B T_{\parallel} = (k_B T_{cat})^2/4E_0 \ll k_B T_{\perp}$

lower limit : $k_B T_{\parallel} \geq 2e \frac{n_c^{1/3}}{4\pi\epsilon_0}$

typical values: $k_B T_{\perp} \approx 0.1$ eV (1100 K), $k_B T_{\parallel} \approx 0.1 - 1$ meV



$$E(r) = eU_{cat} - \frac{n_e \pi r_0^2}{2} m_e c^2 [1 + 2 \ln(r_{tube}/r_0)] + \frac{n_e \pi r_e}{2} m_e c^2 r^2$$

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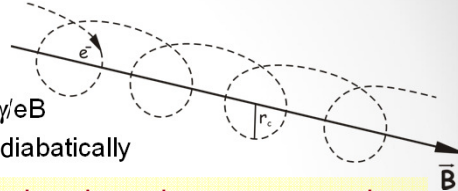
Electron Motion in Longitudinal Magnetic Field

single particle cyclotron motion

cyclotron frequency $\omega_c = eB/\gamma m_e$

cyclotron radius $r_c = v_\perp/\omega_c = (k_B T_\perp m_e)^{1/2} \gamma / eB$

electrons follow the magnetic field line adiabatically



important consequence: for interaction times long compared to the cyclotron period the ion does not sense the transverse electron temperature \Rightarrow magnetized cooling ($T_{\text{eff}} \approx T_\parallel \ll T_\perp$)

electron beam space charge:

transverse electric field + B-field \Rightarrow azimuthal drift

$$v_{azi} = r \omega_{azi} = r \frac{2\pi r_c n_e c^2}{\gamma \omega_c}$$

\Rightarrow electron and ion beam should be well centered

Favorable for optimum cooling (small transverse relative velocity):

- high parallelism of magnetic field lines $\Delta B_\perp/B_0$
- large beta function (small divergence) in cooling section

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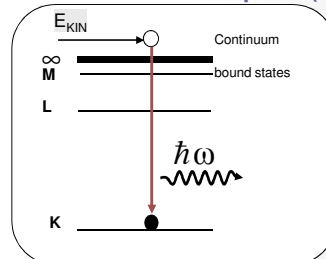
Imperfections and Limiting Effects in Electron Cooling

technical issues:

- ripple of accelerating voltage
- magnetic field imperfections
- beam misalignment
- space charge of electron beam and compensation

physical limitation:

Radiative Electron Capture (REC)



losses by recombination (REC)

loss rate $\tau^{-1} = \gamma^{-2} \alpha_{REC} n_e \eta$

$$\alpha_{REC} = \frac{1.92 \times 10^{-13} Q^2}{\sqrt{k_B T}} \left(\ln \frac{5.66 Q}{\sqrt{k_B T}} + 0.196 \left(\frac{k_B T}{Q^2} \right)^{1/3} \right) [cm^3 s^{-1}]$$

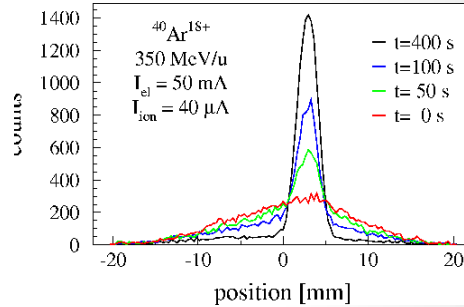
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Examples of Electron Cooling

fast transverse cooling at TSR, Heidelberg

measured with residual gas ionization beam profile monitor

transverse cooling at ESR, Darmstadt



cooling of **350 MeV/u Ar¹⁸⁺** ions
0.05 A, 192 keV electron beam
 $n_e = 0.8 \times 10^6 \text{ cm}^{-3}$

note! time scale, cooling time varies strongly with beam parameters

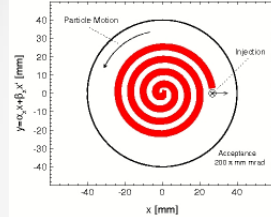
profile every 0.1 s. x [mm]

cooling of **6.1 MeV/u C⁶⁺** ions
0.24 A, 3.4 keV electron beam
 $n_e = 1.56 \times 10^7 \text{ cm}^{-3}$

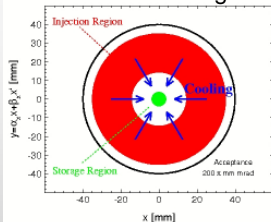
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Accumulation of Heavy Ions by Electron Cooling

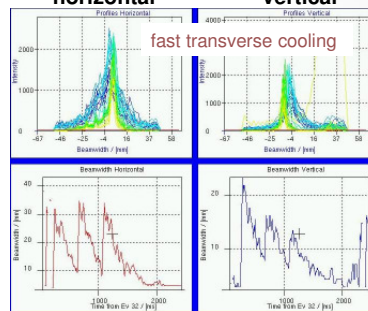
standard multiturn injection



fast accumulation by repeated multiturn injection with electron cooling

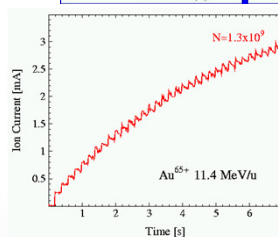


horizontal vertical



profile

beam size



intensity increase in 5 s by a factor of ≈ 10

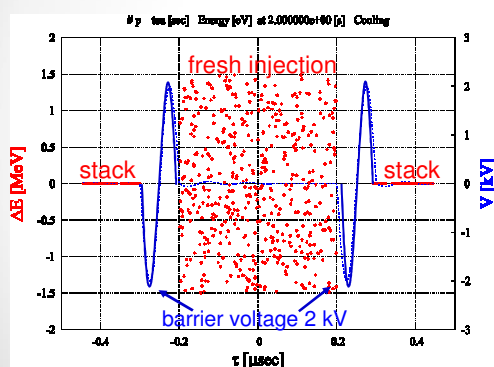
limitations:
 space charge tune shift,
 recombination (REC)

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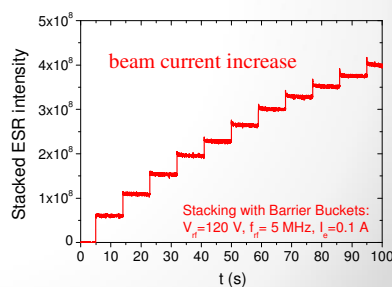
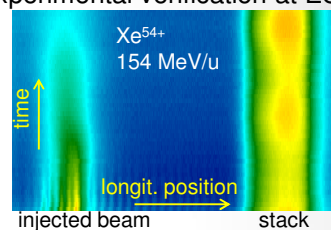
Accumulation of Secondary Particles

basic idea: confine stored beam to a fraction of the circumference, inject into gap and apply cooling to merge the two beam components
 \Rightarrow fast increase of intensity (for secondary beams)

experimental verification at ESR



simulation of longitudinal stacking with barrier buckets and electron cooling

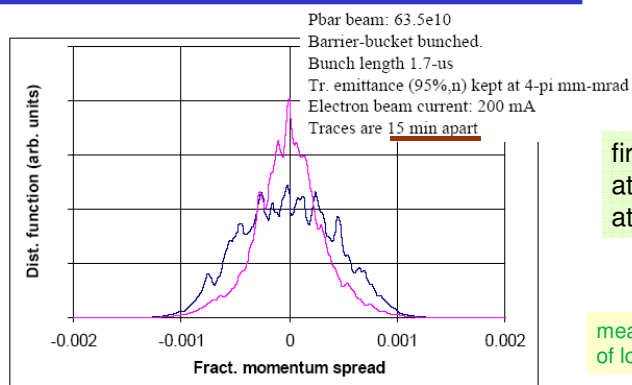


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Examples of Electron Cooling

high energy electron cooling of 8 GeV antiprotons
 longitudinal cooling with 0.2 A, 4.4 MeV electron beam

First e-cooling demonstration - 07/15/05



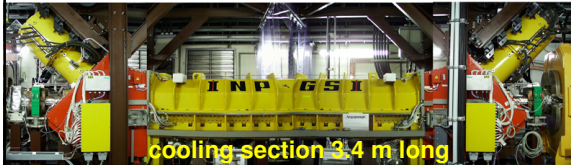
first electron cooling at relativistic energy at Recycler, FNAL

measured by detection of longitudinal Schottky noise

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Electron Cooling Systems

Low Energy: 35 keV SIS/GSI



Medium Energy:
300 keV
ESR/GSI



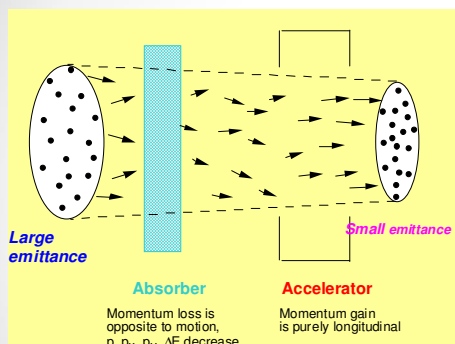
High Energy:
4.3 MeV Recycler/FNAL



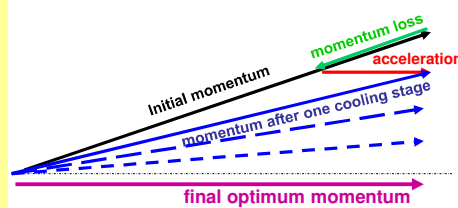
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2. Ionization Cooling

energy loss in solid matter



proposed for muon cooling



not useful for heavy particles
due to strong interaction with matter

transverse cooling

$$\frac{d\epsilon_N}{ds} = -\frac{1}{\beta^2 E} \frac{dE}{ds} \epsilon_N + \frac{\beta \gamma \beta_{\perp}}{2} \frac{\langle \theta_{rms}^2 \rangle}{ds}$$

$$= -\frac{1}{\beta^2 E} \frac{dE}{ds} \epsilon_N + \frac{\beta_{\perp} E_s^2}{2 \beta^3 m_{\mu} c^2 L_R E}$$

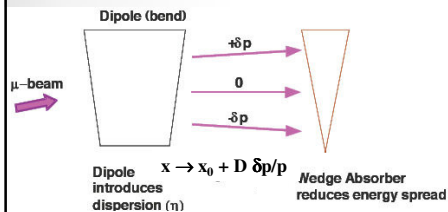
small β_{\perp} at absorber in order
to minimize multiple scattering

large $L_R, (dE/ds) \Rightarrow$ light absorbers (H_2)

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Ionization Cooling

increased longitudinal cooling
by longitudinal-transverse emittance exchange



$$\frac{d\sigma_E^2}{ds} = -2 \frac{\partial(dE/ds)}{\partial E} \sigma_E^2 + \frac{d(\Delta E_{rms}^2)}{ds}$$

cooling term heating term

cooling, if $\frac{\partial(dE/ds)}{\partial E} > 0$

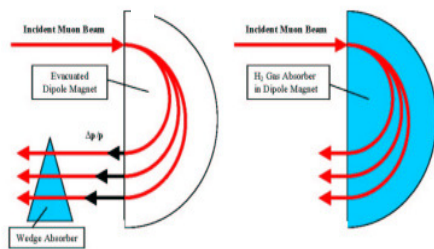
emittance exchange

increased longitudinal cooling

$$\frac{\partial \frac{dE}{ds}}{\partial E} \Rightarrow \frac{\partial \frac{dE}{ds}}{\partial E} \Big|_0 + \frac{dE}{ds} \frac{D\rho'}{\beta c p \rho_0}$$

reduced transverse cooling

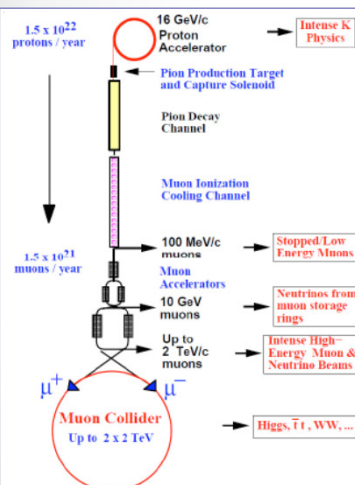
$$\frac{d\epsilon_N}{ds} = -\frac{1}{\beta^2 E} \frac{dE}{ds} \left(1 - \frac{D\rho'}{\rho_0}\right) \epsilon_N$$



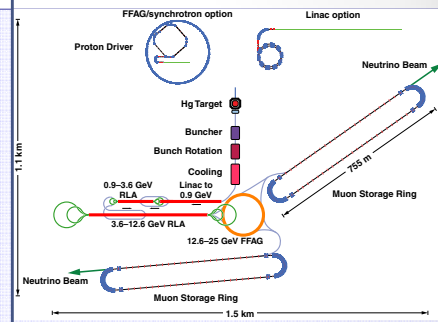
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Scenarios with Ionization Cooling

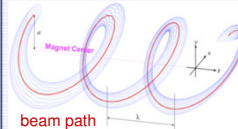
Muon Collider



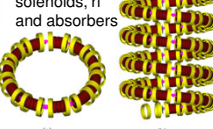
Neutrino Factory



helical cooling channel



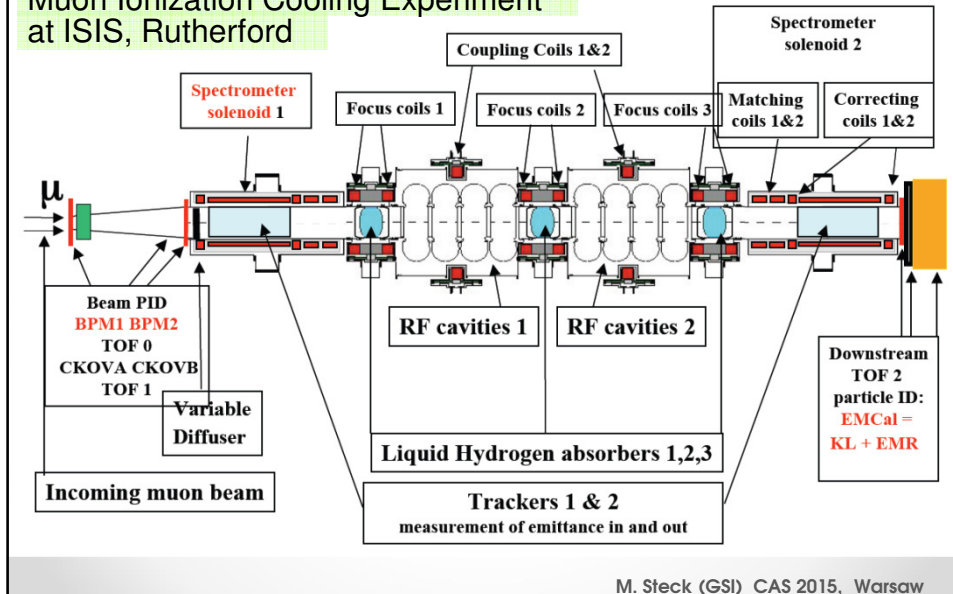
sequence of solenoids, rf and absorbers



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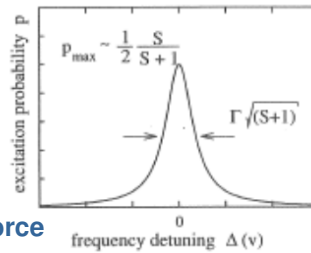
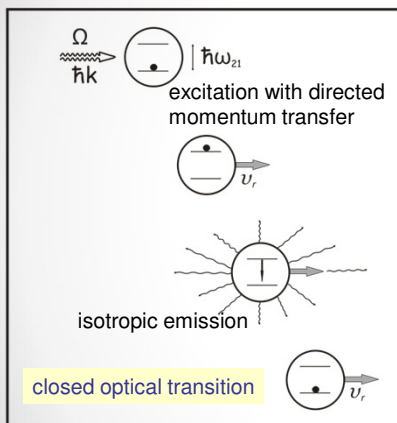
MICE

Muon Ionization Cooling Experiment at ISIS, Rutherford



3. Laser Cooling

$$\Omega = \gamma \omega_{21} (1 - \beta \cos \theta)$$



$$\vec{F}(\vec{v}, \vec{k}) = \frac{\hbar \vec{k}}{2} S \Gamma \frac{(\Gamma/2)^2}{(\omega - \omega_{21} - \vec{v} \cdot \vec{k}) + (\Gamma/2)^2 (1 + S)}$$

Lorentzian with width $\Gamma/k \sim 10$ m/s

minimum temperature $T_D = \frac{\hbar \Gamma}{2k_B}$ (Doppler limit)
typical $10^{-5} - 10^{-4}$ K

typical cooling time ~ 10 μ s

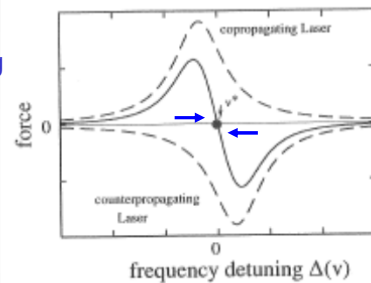
drawback: only longitudinal cooling

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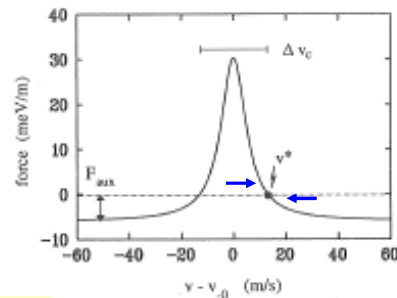
Laser Cooling

a single laser does not provide cooling (only acceleration or deceleration)

schemes
for cooling



two counter-propagating lasers
(matched to beam velocity, but slightly detuned)



auxiliary force
(betatron core, rf)

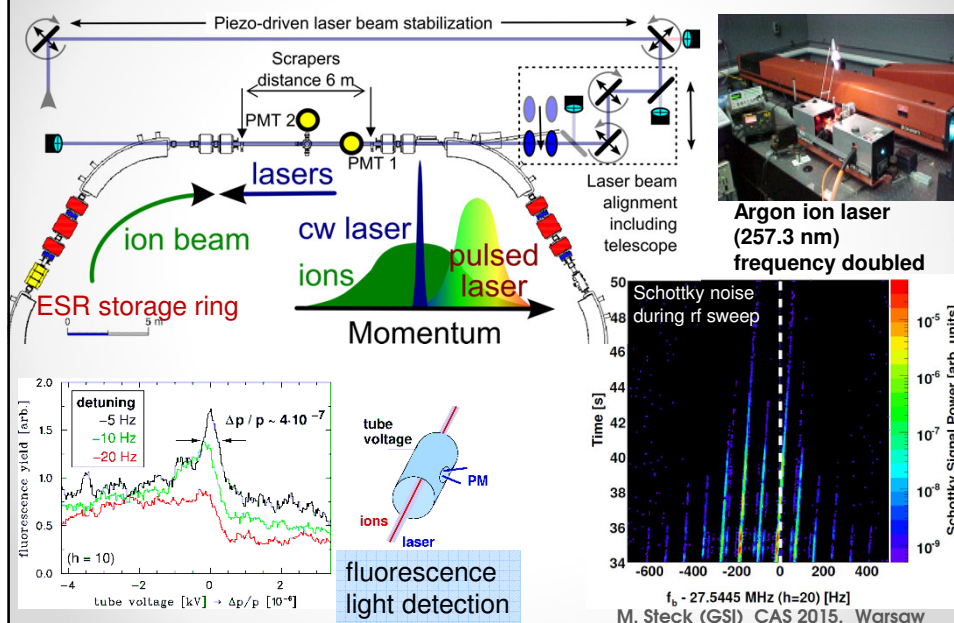
capture range of laser is limited \Rightarrow frequency sweep (snowplow)

ions studies so far: $^7\text{Li}^+$, $^9\text{Be}^{1+}$, $^{24}\text{Mg}^{1+}$, $^{12}\text{C}^{3+}$

in future: Li-like heavy ions at relativistic energies
large relativistic energy \Rightarrow large excitation energy in PRF

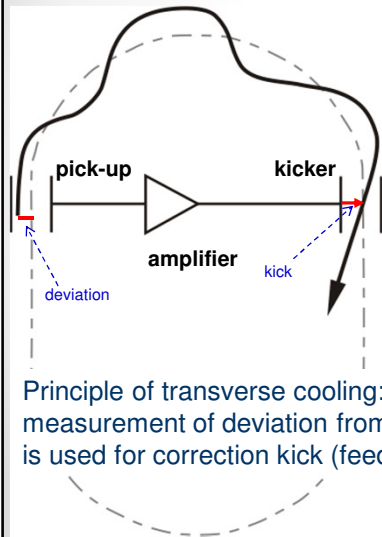
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Laser Cooling of C^{3+}



4. Stochastic Cooling

First cooling method which was successfully used for beam preparation



S. van der Meer, D. Möhl, L. Thorndahl et al.
(1925 – 2011) (1936-2012)

Conditions:

Betatron motion phase advance
(pick-up to kicker): $(n + \frac{1}{2}) \pi$

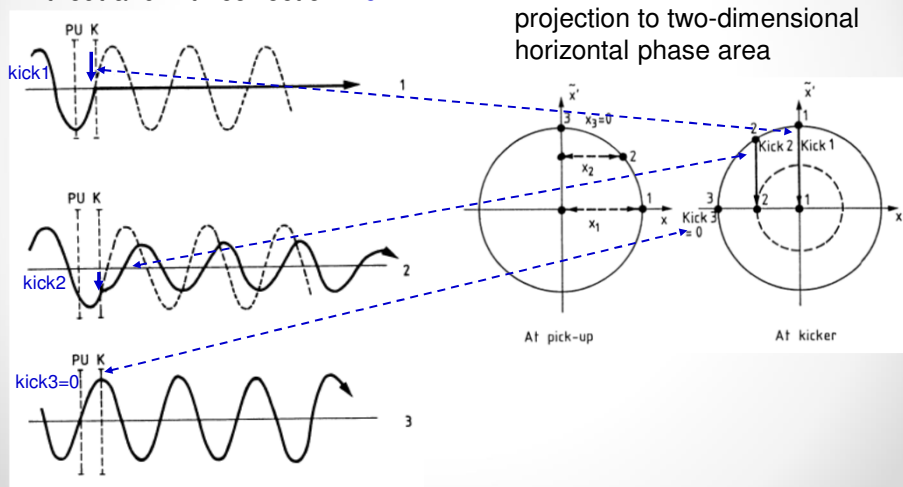
Signal travel time = time of flight of particle
(between pick-up and kicker)

Sampling of sub-ensemble of total beam

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Stochastic Cooling

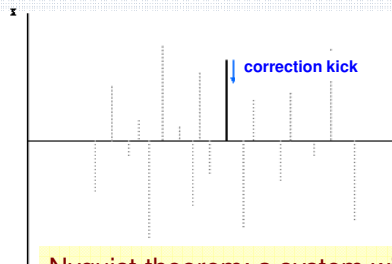
single particle betatron motion
along storage ring
without and with correction kick



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Stochastic Cooling

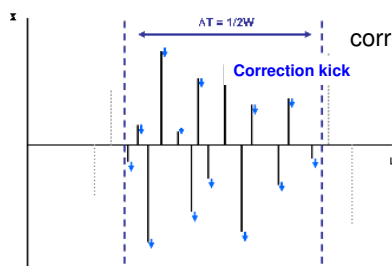
in time domain



correction kick
(unlimited resolution)

$$\Delta x = g \times x$$

Nyquist theorem: a system with a band-width $\Delta f = W$ in frequency domain can resolve a minimum time duration $\Delta T = 1/(2W)$



$$\text{correction kick } \Delta x = \frac{g}{N_s} \times \sum_{i=1..N_s} x_i, \quad N_s = N \frac{\Delta T}{T_0} = \frac{N}{2WT_0}$$

For exponential damping ($x(t) = x(t_0) \cdot \exp(-(t-t_0)/\tau)$):

$$\tau^{-1} = T_0^{-1} \times \frac{\Delta x}{x} = \frac{g2W}{N}, \quad \text{if } \sum_{i=1..N_s} x_i = x$$

cooling rate $\tau^{-1} \leq \frac{2W}{N} \quad \text{if } g \leq 1$

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Stochastic Cooling

some refinements of cooling rate formula

noise: thermal or electronic noise adds to the beam signal

mixing: change of relative longitudinal position of particles due to momentum spread

cooling rate $\lambda = \tau^{-1} = \frac{2W}{N} (2\underbrace{g}_{\text{cooling}} - \underbrace{g^2(M+U)}_{\text{heating}})$ M mixing factor
U noise to signal ratio

maximum of cooling rate

$$\lambda_{max} = \frac{2W}{N} \frac{1}{M+U}$$

$$\frac{d\lambda}{dg} = 0 \Rightarrow g = \frac{1}{M+U}$$

further refinement (wanted ↔ unwanted mixing):

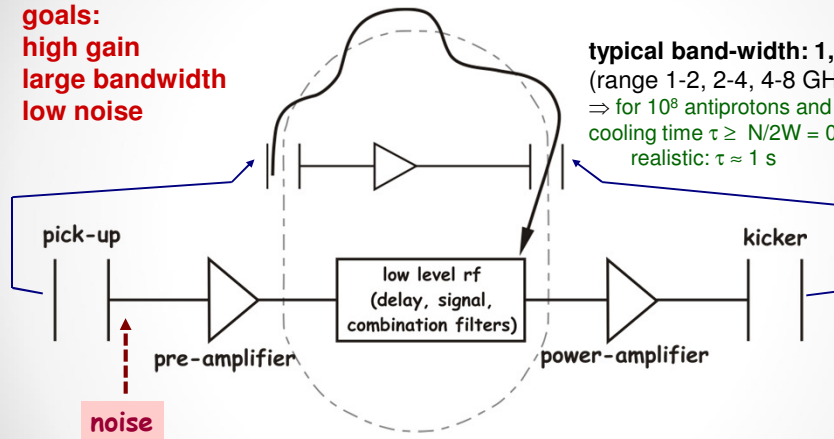
with wanted mixing M (kicker to pick-up) and unwanted mixing \tilde{M} (pick-up to kicker) $\lambda = \tau^{-1} = \frac{2W}{N} (2g(1 - \tilde{M}^2) - g^2(M+U))$

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Stochastic Cooling Circuit

goals:
high gain
large bandwidth
low noise

typical band-width: 1, 2, 4 GHz
(range 1-2, 2-4, 4-8 GHz)
⇒ for 10^8 antiprotons and $W = 1$ GHz
cooling time $\tau \geq N/2W = 0.05$ s
realistic: $\tau \approx 1$ s



Transfer Function:

$$Z_{pick-up} \cdot G_{pick-up}(E) \cdot H(t_{delay}) \cdot F(E) \cdot G \cdot G_{kicker}(E) \cdot Z_{kicker}$$

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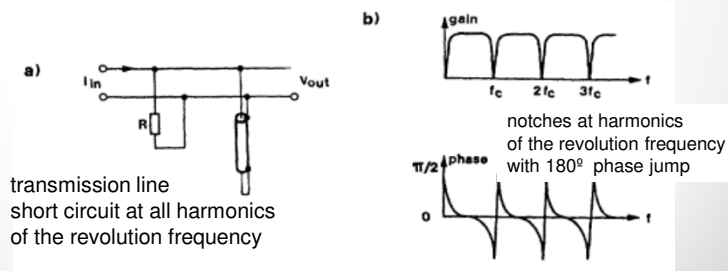
Longitudinal Stochastic Cooling

1) Palmer cooling

pick-up in dispersive section detects horizontal position
⇒ acceleration/deceleration kick corrects momentum deviation

2) Notch filter cooling

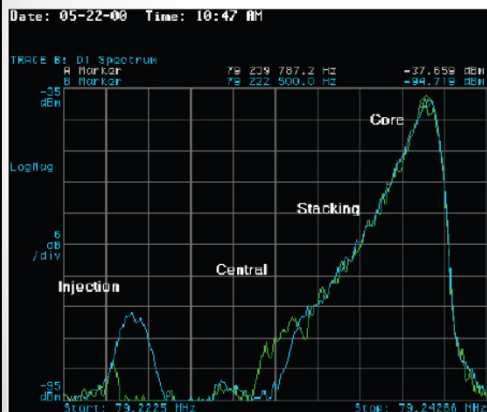
filter creates notches at the harmonics of the nominal revolution frequency
⇒ particles are forced to circulate at the nominal frequency



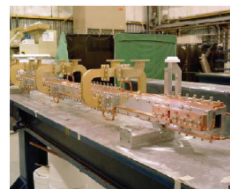
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Antiproton Accumulation by Stochastic Cooling

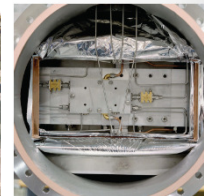
accumulation of 8 GeV antiprotons at accumulator ring, FNAL, shut down 09/2011
a similar facility AC/AA at CERN was shut down 11/1996



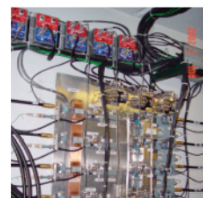
momentum distribution of accumulated antiproton beam



kicker array



cryogenic microwave amplifier



microwave electronics



power amplifiers (TWTs)

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Stochastic Cooling of Rare Isotopes at GSI

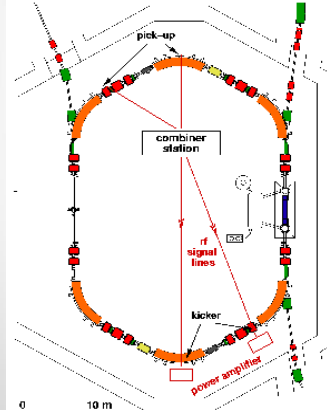
fast pre-cooling of hot fragment beams

energy 400 (-550) MeV/u

bandwidth 0.8 GHz (range 0.9-1.7 GHz)

$\delta p/p = \pm 0.35\%$ \rightarrow $\delta p/p = \pm 0.01\%$

$\epsilon = 10 \times 10^{-6} \text{ m}$ \rightarrow $\epsilon = 2 \times 10^{-6} \text{ m}$



electrodes
installed
inside magnets



combination of
signals from
electrodes



power amplifiers
for generation of
correction kicks

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Comparison of Cooling Methods

Stochastic Cooling

Useful for: low intensity beams
hot (secondary) beams
high charge
full 3 D control

Limitations: high intensity beams
/problems beam quality limited
bunched beams

Electron Cooling

low energy
all intensities
warm beams (pre-cooled)
high charge
bunched beams

space charge effects
recombination losses
high energy

laser cooling (of incompletely ionized ions)
and ionization cooling (of muons) are quite particular
and not general cooling methods

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Trends in Beam Cooling

Stochastic cooling was mainly developed for the production of high intensity antiproton beams for colliders (CERN, FNAL, 1972 – 2011). It is still in operation at AD (CERN), COSY (FZJ) and ESR (GSI).

It will also be used in the FAIR project (Germany) for cooling of antiprotons and rare isotope beams.

First demonstration of **bunched beam stochastic cooling** (2008) with ions (BNL) made it also attractive for ion colliders.

Now it is proposed for the collider of the Russian NICA project.

Electron cooling was and still is used in low energy storage rings for protons, ions, secondary beams (antiprotons, rare isotopes).

Electron cooling is interesting for low energy storage rings, but also application at higher energies (**MeV electron energies**) is envisaged after the successful demonstration of the 4 MeV electron cooler at FNAL.

Other cooling methods, like **muon (ionization) cooling** or **coherent electron cooling** are under investigation, but are still far from implementation in a full scale machine.

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Biannual Workshops on Beam Cooling: e. g. COOL'15, Jefferson Lab, USA

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