

ICHEP 2016
Chicago, August 4-10 2016

Higgs physics at 100 TeV

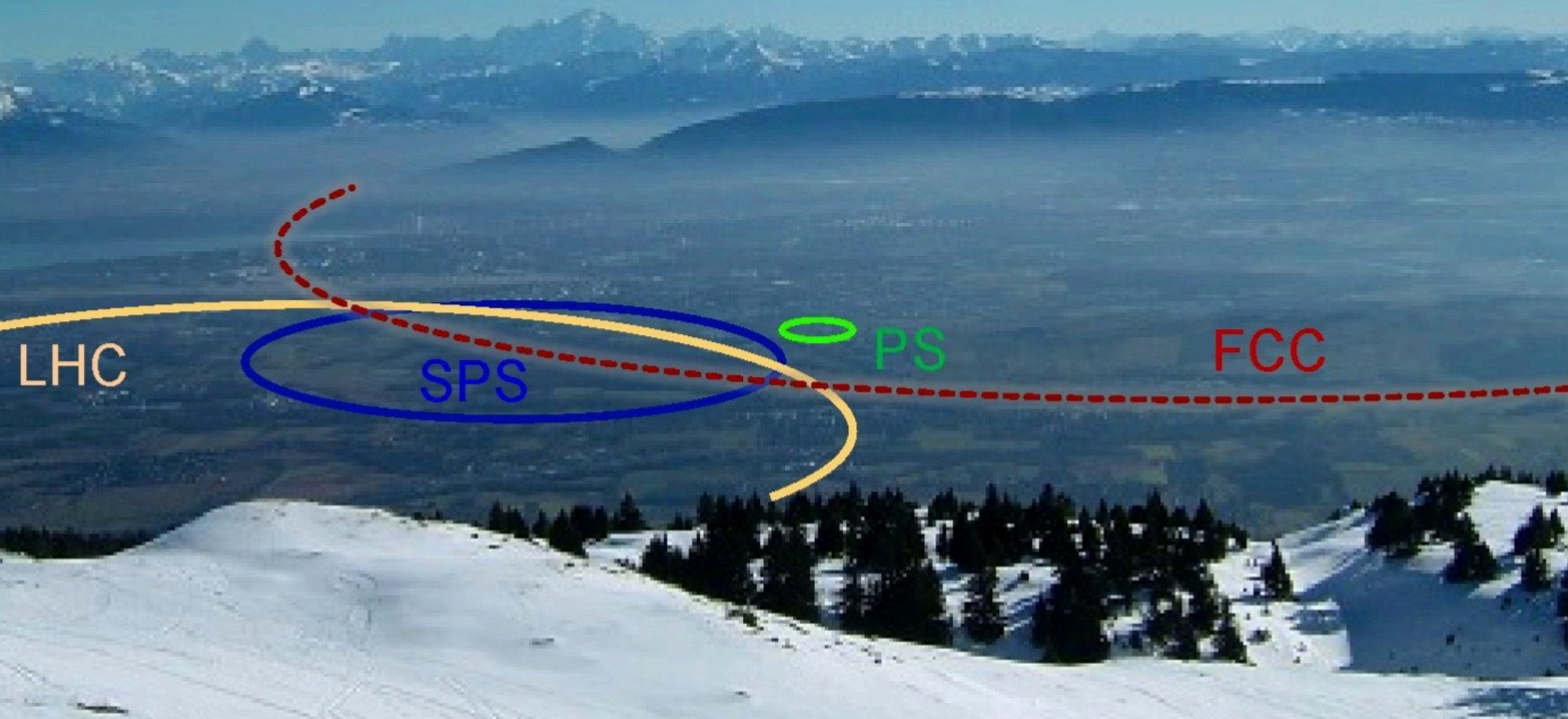
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CERN



The context

Parameter	FCC-hh	LHC
Energy [TeV]	100 c.m.	14 c.m.
Dipole field [T]	16	8.33
# IP	2 main, +2	4
Luminosity/IP _{main} [cm ⁻² s ⁻¹]	5 - 25 x 10 ³⁴	1 x 10 ³⁴
Stored energy/beam [GJ]	8.4	0.39
Synchrotron rad. [W/m/aperture]	28.4	0.17
Bunch spacing [ns]	25 (5)	25

- **Phase 1 (baseline):** $5 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$ (peak),
250 fb⁻¹/year (averaged)
2500 fb⁻¹ within 10 years (~HL LHC total luminosity)
- **Phase 2 (ultimate):** $\sim 2.5 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$ (peak),
1000 fb⁻¹/year (averaged)
→ 15,000 fb⁻¹ within 15 years
- **Yielding total luminosity O(20,000) fb⁻¹ over ~25 years of operation**



Report on Physics at a 100 TeV pp Collider

- **Volume 1: SM processes** (238 pages)
 - arXiv:1607.01831
- **Volume 2: Higgs and EW symmetry breaking studies** (175 pages)
 - arXiv:1606.09408
- **Volume 3: beyond the Standard Model phenomena** (189 pages)
 - arXiv:1606.00947
- **Volume 4: physics with heavy ions** (56 pages)
 - arXiv:1605.01389
- **Volume 5: physics opportunities with the FCC-hh injectors** (14 pages)

total ~700 pages

Higgs chapter of FCC physics report

CERN-TH-2016-113

Physics at a 100 TeV pp collider: Higgs and EW symmetry breaking studies

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Contents

1	Foreword	5
2	SM Higgs production	6
2.1	Inclusive $gg \rightarrow H$ production	6
2.2	Higgs plus jet and Higgs p_T spectrum in $gg \rightarrow H$	13
2.3	Higgs plus jets production in $gg \rightarrow H$	18
2.4	Associated VH production	28
2.5	VBF Higgs production	31
2.6	Associated $t\bar{t}H$ production	42
2.7	Rare production modes	48
3	Prospects for measurements of SM Higgs properties	50
3.1	Higgs acceptance	52
3.2	Small-BR H final states at intermediate p_T	53
3.3	Associated VH production	56
3.4	Measurement of top Yukawa coupling from the $t\bar{t}H/t\bar{t}Z$ ratio	60
3.5	Combined determination of y_t and $\Gamma(H)$ from $t\bar{t}H$ vs $t\bar{t}t\bar{t}$ production	62
3.6	Rare SM Exclusive Higgs decays	66
4	Multi-Higgs production	72
4.1	Double Higgs production from gluon fusion	77
4.2	Triple Higgs production and the quartic Higgs self-coupling	96
5	BSM aspects of Higgs physics and EWSB	105
5.1	Introduction	105
5.2	Overview	106
5.3	Electroweak Phase Transition and Baryogenesis	113
5.4	Dark Matter	122
5.5	The Origins of Neutrino Mass and Left-right symmetric model	125
5.6	Naturalness	135
5.7	BSM Higgs Sectors	145

Why Higgs at 100 TeV?

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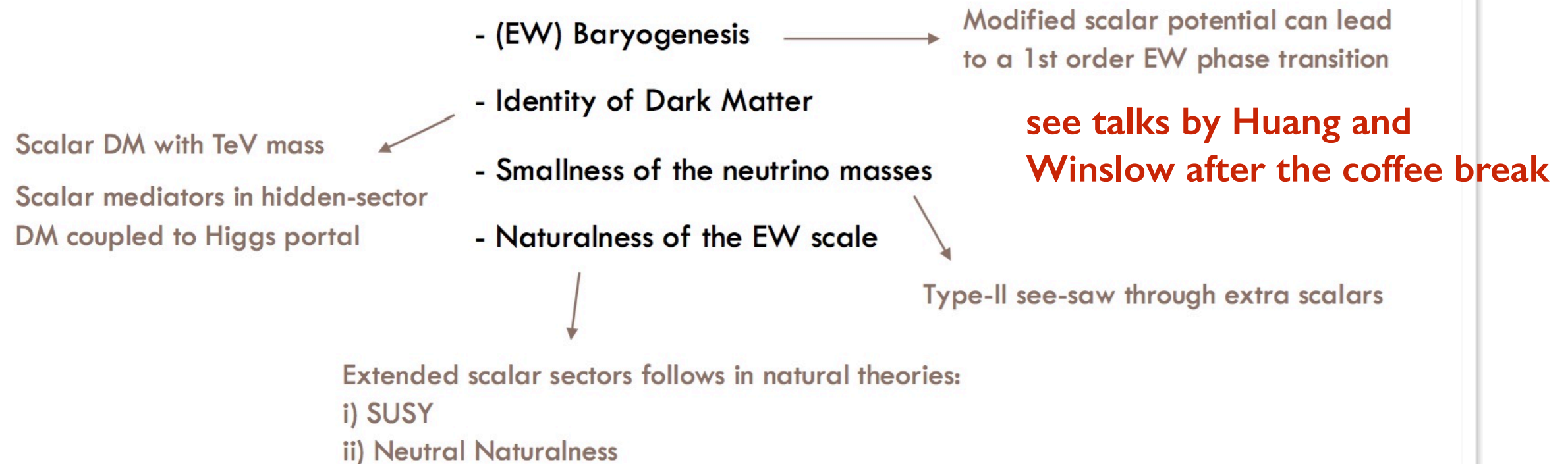
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- At the LHC, the Higgs is already an analysis tool, if not a background, in searches of new particles (like W/Z and like the top quark). This will be even more true at 100 TeV!!

Higgs and BSM

Search for Extended Higgs sectors

Extended Higgs sectors are a prediction of many BSM scenarios.
They may play a role in the following open questions:

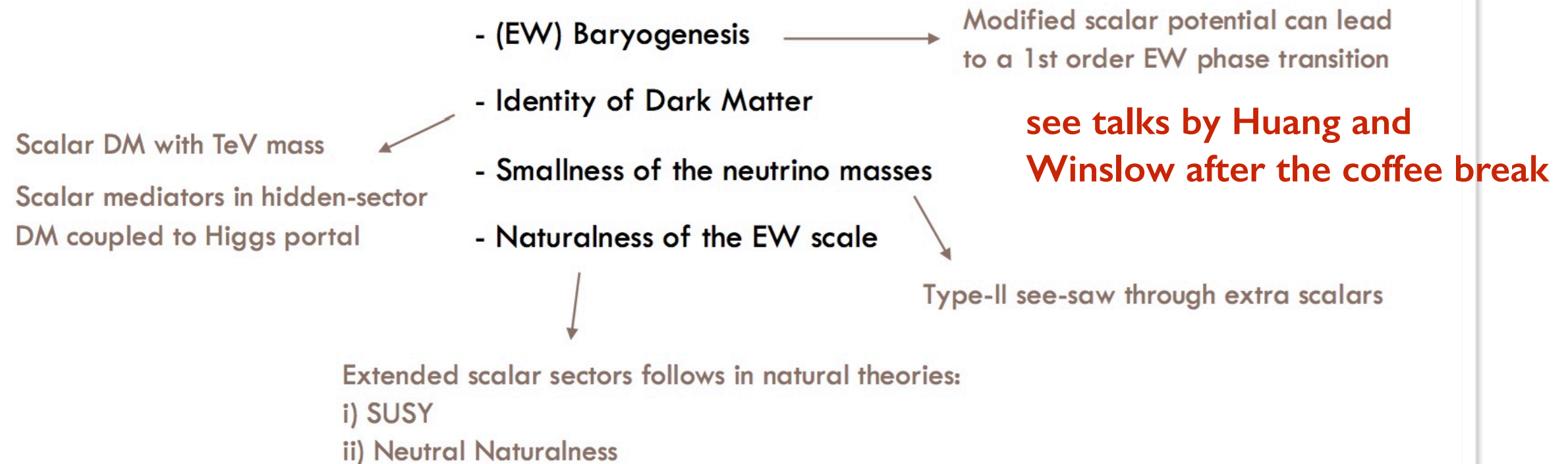


👉 A 100TeV pp collider offers the unique opportunity to discover EW-charged or SM-singlet scalars with a few TeV mass

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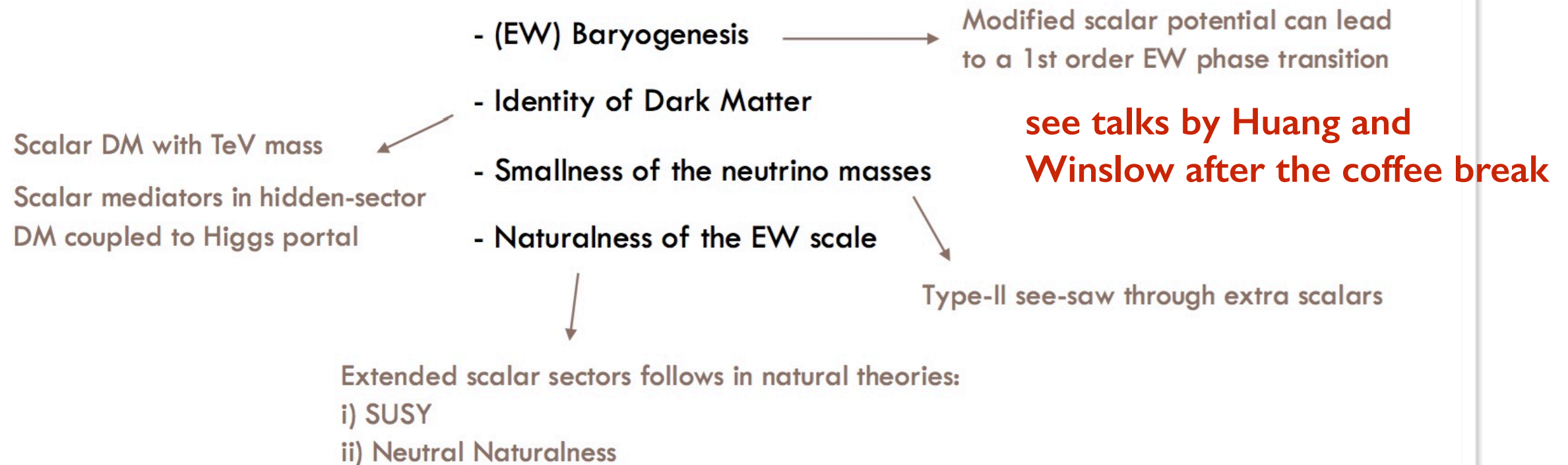
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*Unless otherwise stated, plots in the following are from the
Report, where more details and ref's can be found*

Higgs as a BSM probe: precision vs dynamic reach

$$L = L_{SM} + \frac{1}{\Lambda^2} \sum_k \mathcal{O}_k + \dots$$

$$O = | \langle f | L | i \rangle |^2 = O_{SM} [1 + O(\mu^2/\Lambda^2) + \dots]$$

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For H decays, or inclusive production, $\mu \sim O(v, m_H)$

$$\delta O \sim \left(\frac{v}{\Lambda} \right)^2 \sim 6\% \left(\frac{\text{TeV}}{\Lambda} \right)^2 \Rightarrow \text{precision probes large } \Lambda$$

e.g. $\delta O = 1\% \Rightarrow \Lambda \sim 2.5 \text{ TeV}$

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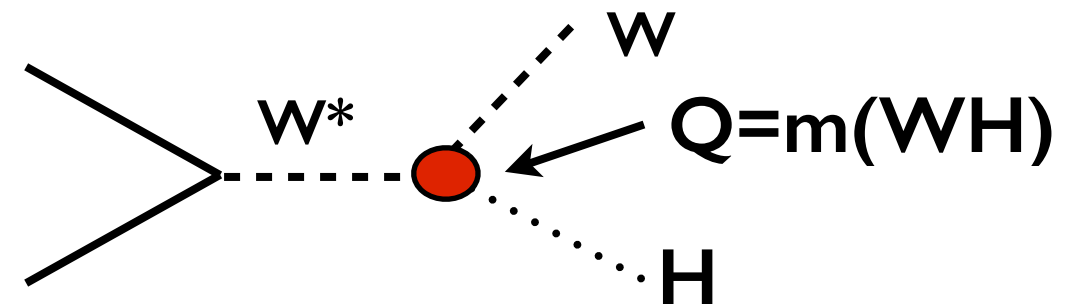
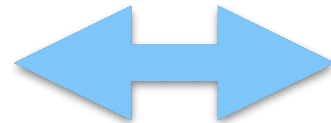
For H production off-shell or with large momentum transfer Q , $\mu \sim O(Q)$

$$\delta O \sim \left(\frac{Q}{\Lambda} \right)^2 \Rightarrow \text{kinematic reach probes large } \Lambda \text{ even if precision is low}$$

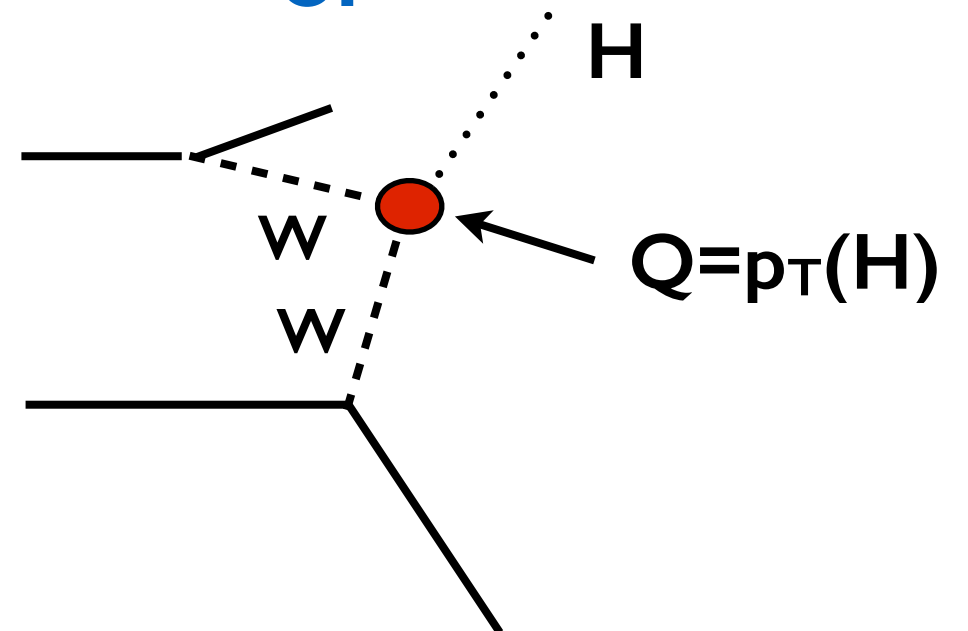
$$\text{e.g. } \delta O = 15\% \text{ at } Q = 1 \text{ TeV} \Rightarrow \Lambda \sim 2.5 \text{ TeV}$$

Examples

$\delta\text{BR}(H \rightarrow WW^*)$

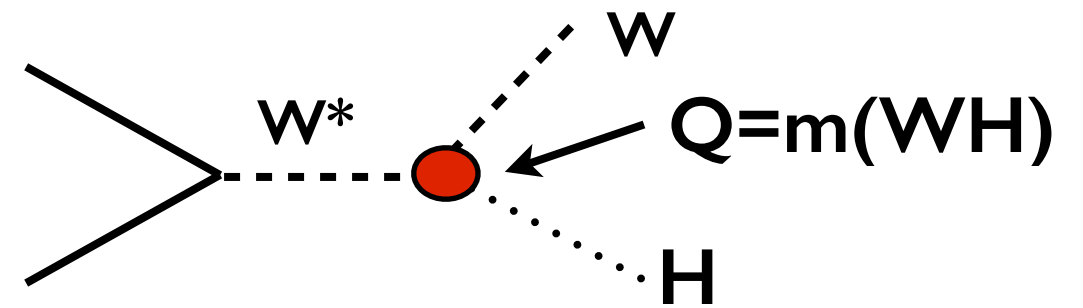
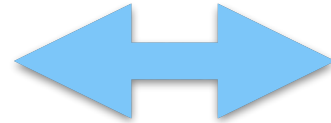


or

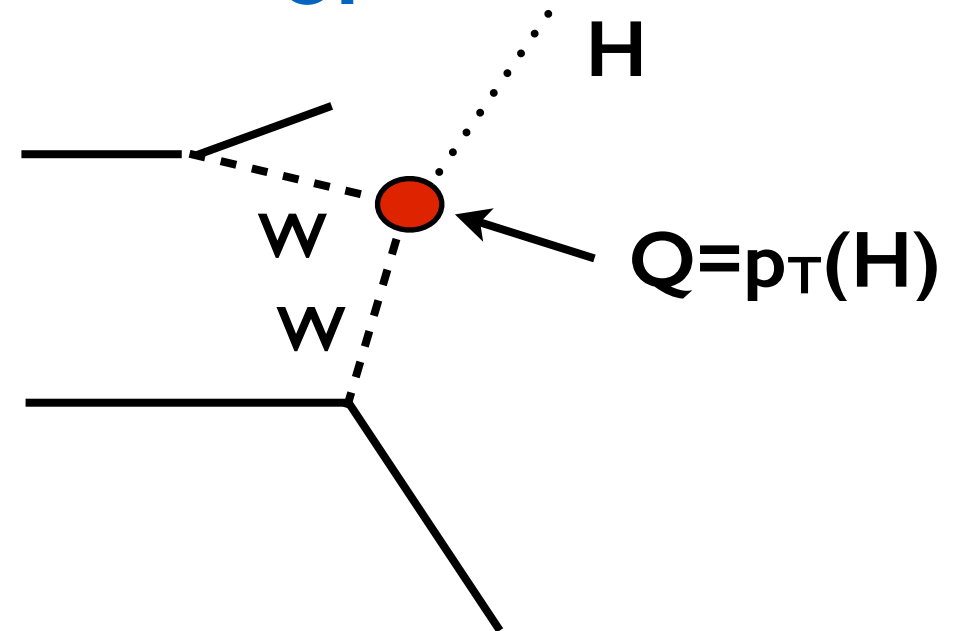


Examples

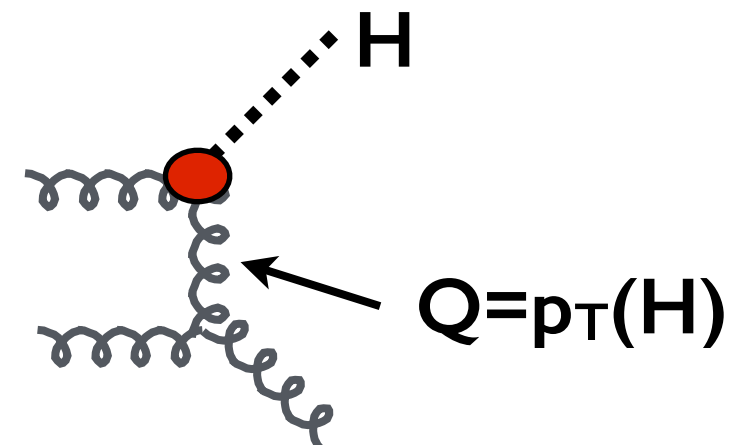
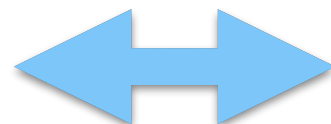
$\delta\text{BR}(H \rightarrow WW^*)$



or



$\delta\text{BR}(H \rightarrow gg)$



$$\delta O_Q \sim \left(\frac{Q}{\Lambda}\right)^2 \quad \mathbf{vs} \quad \delta O \sim \left(\frac{v}{\Lambda}\right)^2$$

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For a high-Q observable O_Q to achieve the same sensitivity of a “precision” observable O , it is sufficient, for a given Q, to reach an accuracy

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Or, for a given accuracy δO_Q , to have statistics on O_Q at a scale

$$Q \sim v \left(\frac{\delta O_Q}{\delta O}\right)^{1/2}$$

$$\delta O_Q \sim \left(\frac{Q}{\Lambda}\right)^2 \quad \mathbf{vs} \quad \delta O \sim \left(\frac{v}{\Lambda}\right)^2$$

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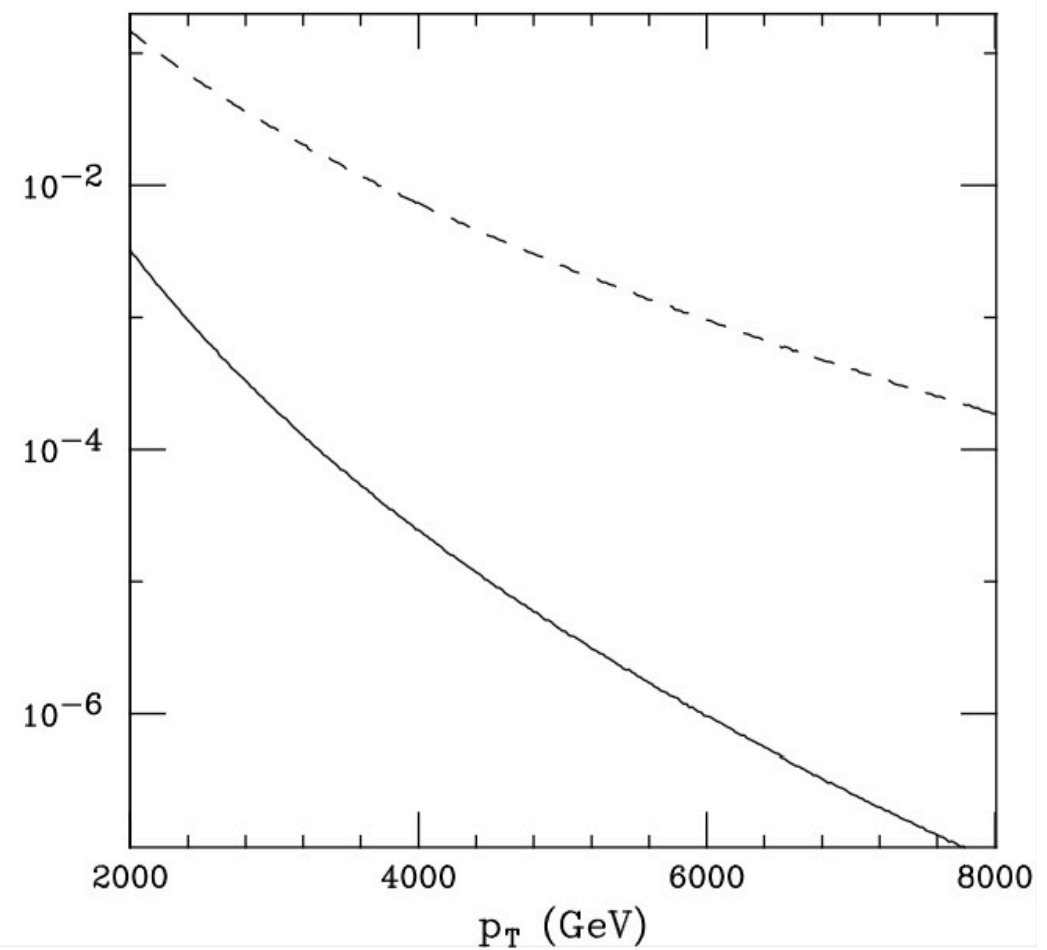
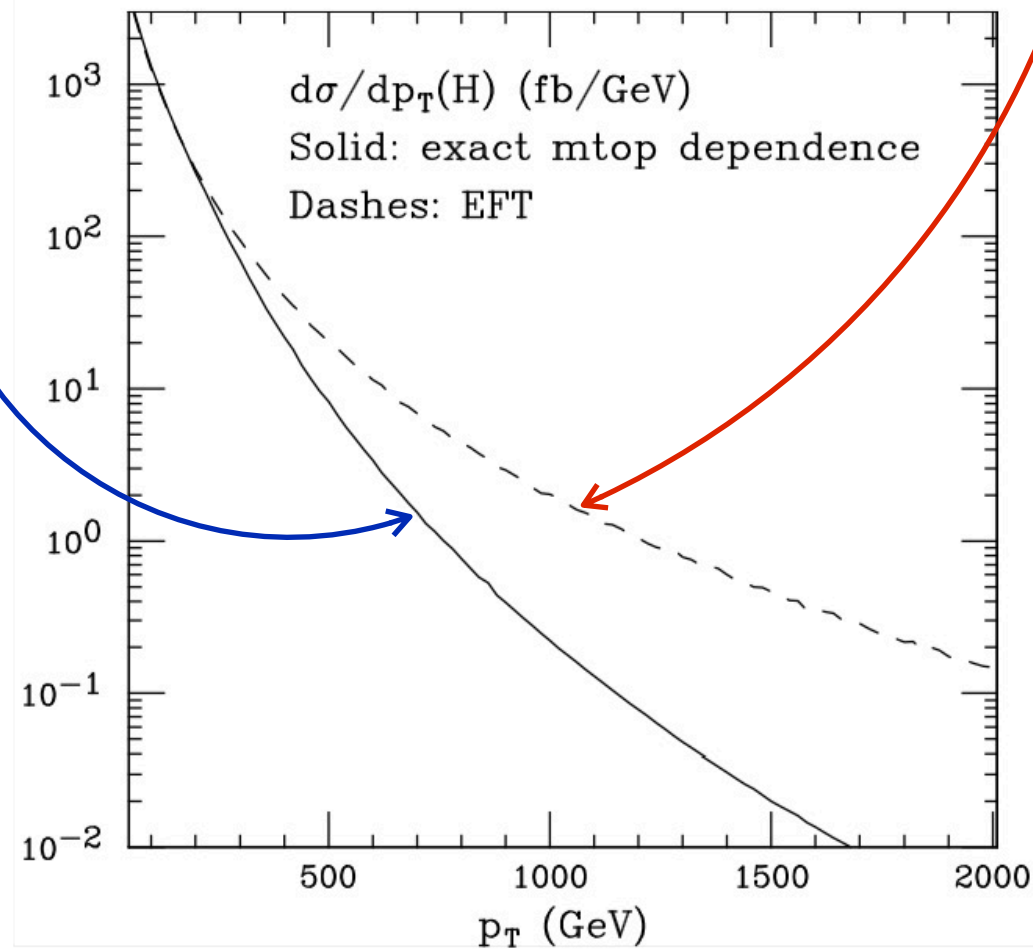
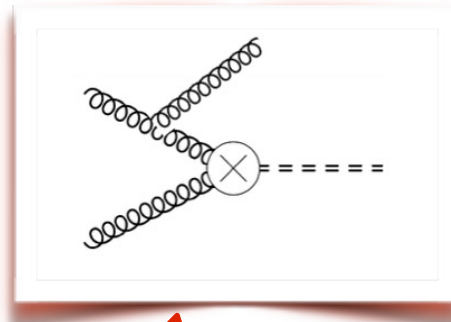
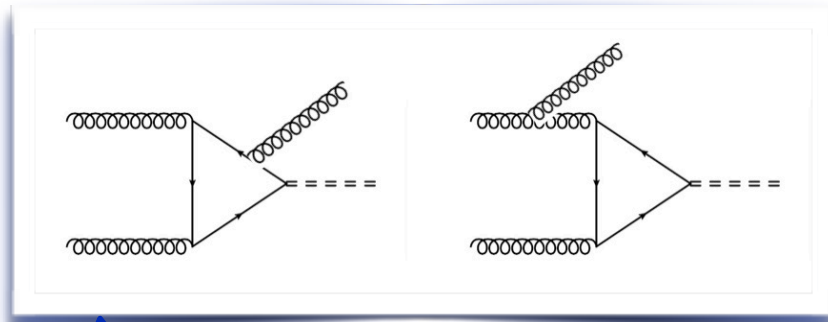
$$Q \sim v \left(\frac{\delta O_Q}{\delta O}\right)^{1/2}$$

E.g. for $\delta O \sim 10^{-3}$ (goal of precision BR measurements at FCC-ee):

$$- \delta O_Q \sim 10^{-1} \Rightarrow Q \sim 10 v \sim 2.5 \text{ TeV}$$

$$- \delta O_Q \sim 10^{-2} \Rightarrow Q \sim 3 v \sim 750 \text{ GeV}$$

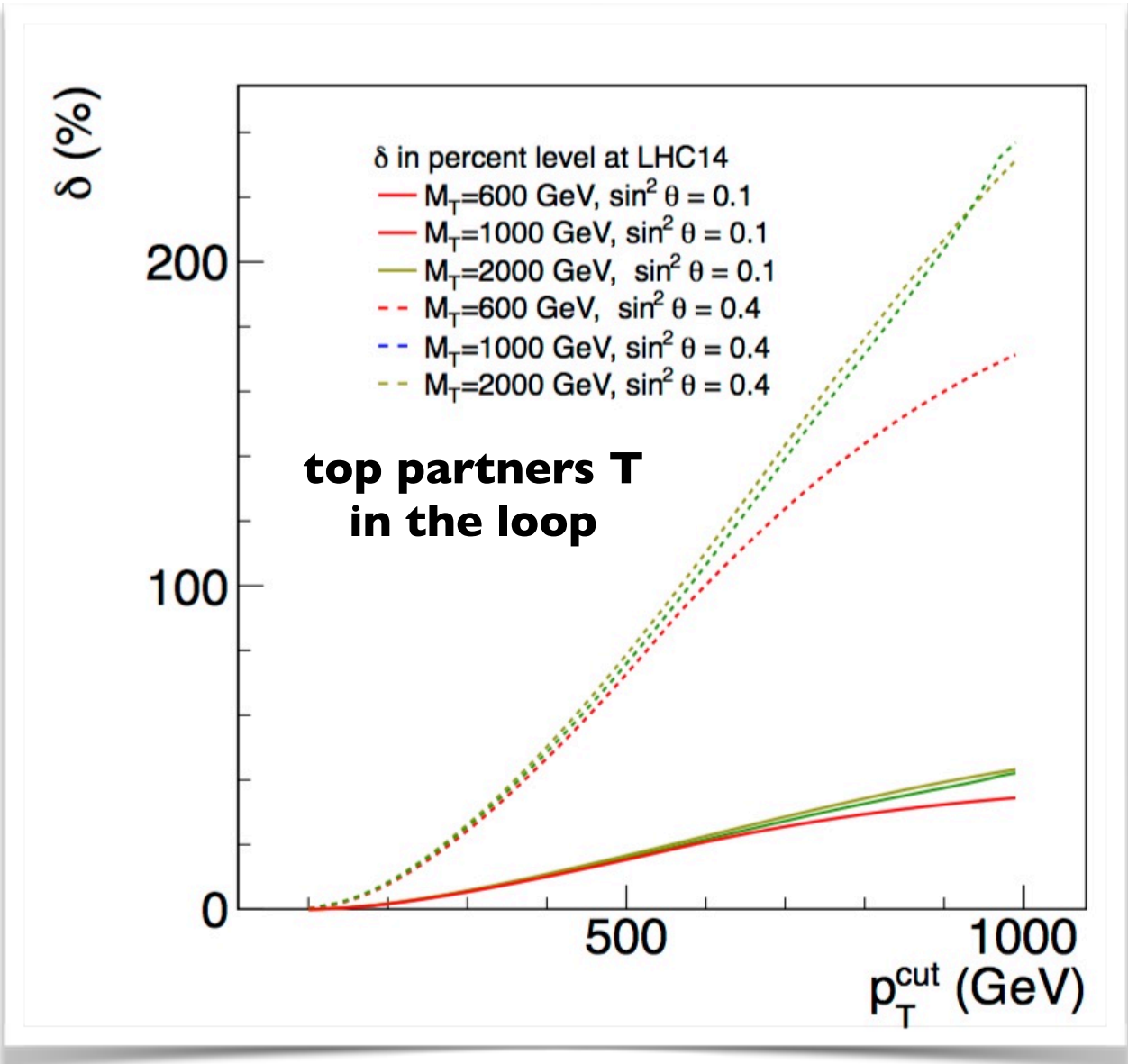
$gg \rightarrow H$ at large p_T



- At LHC, can measure only up to $p_T \sim$ few hundred GeV \Rightarrow reduced sensitivity to the inner guts of the ggH coupling
- At FCC, orders of magnitude difference between EFT and exact m_{top}

Examples

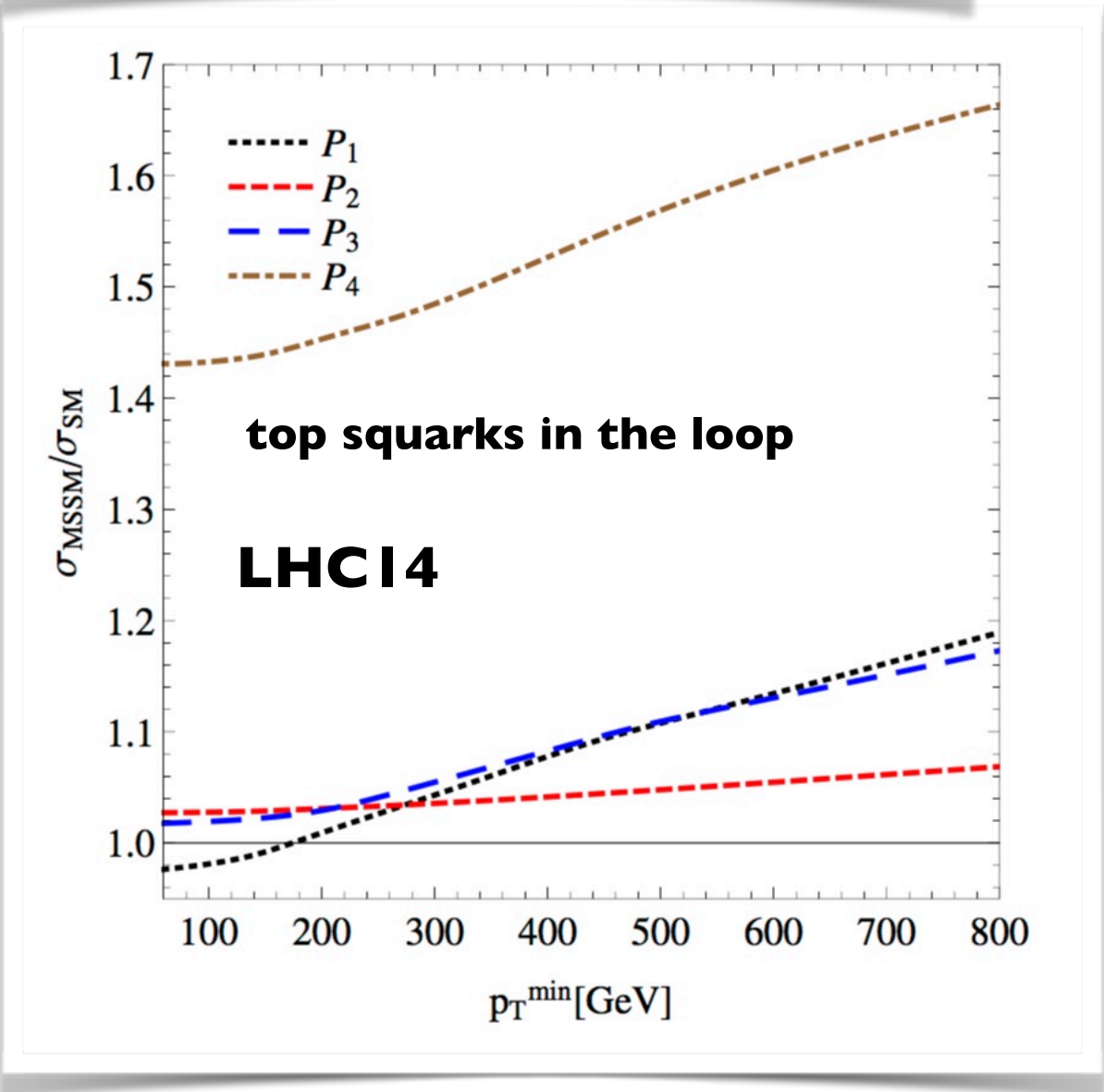
(See also
Azatov and Paul [arXiv:1309.5273v3](https://arxiv.org/abs/1309.5273v3))



Banfi Martin Sanz, [arXiv:1308.4771](https://arxiv.org/abs/1308.4771)

Table 3: The benchmark points shown in Fig. 7. We set $\tan \beta = 10$, $M_{A^0} = 500$ GeV, $M_2 = 1000$ GeV, $\mu = 200$ GeV and all trilinear couplings to a common value A_t . The remaining sfermion masses were set to 1 TeV and the mass of the lightest CP -even Higgs was set to 125 GeV.

Point	$m_{\tilde{t}_1}$ [GeV]	$m_{\tilde{t}_2}$ [GeV]	A_t [GeV]	Δ_t
P_1	171	440	490	0.0026
P_2	192	1224	1220	0.013
P_3	226	484	532	0.015
P_4	226	484	0	0.18



Grojean, Salvioni, Schlaffer, Weiler [arXiv:1312.3317](https://arxiv.org/abs/1312.3317)

SM Higgs at 100 TeV

	N_{100}	N_{100}/N_8	N_{100}/N_{14}
$gg \rightarrow H$	16×10^9	4×10^4	110
VBF	1.6×10^9	5×10^4	120
WH	3.2×10^8	2×10^4	65
ZH	2.2×10^8	3×10^4	85
$t\bar{t}H$	7.6×10^8	3×10^5	420

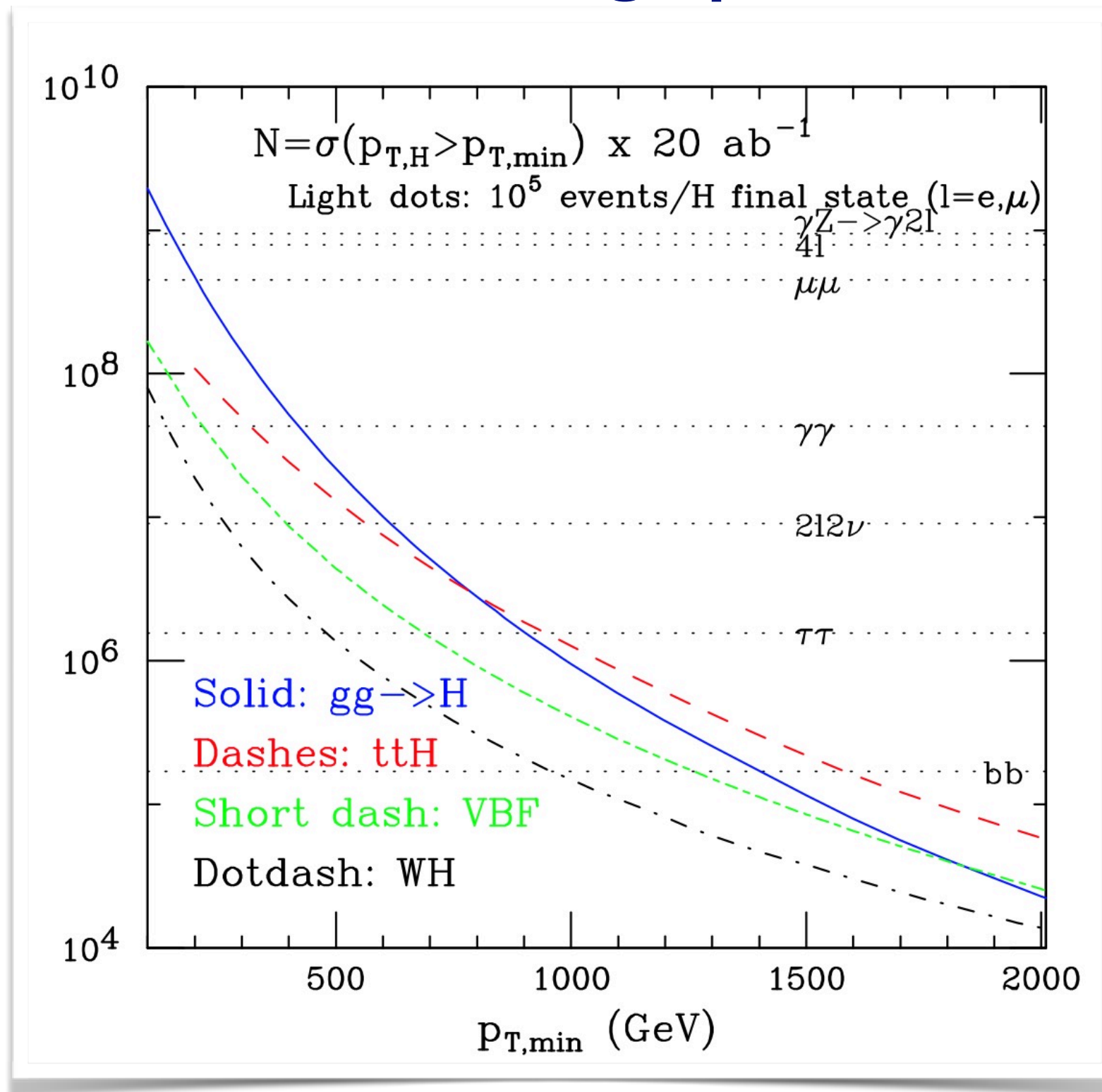
$$N_{100} = \sigma_{100\text{TeV}} \times 20 \text{ ab}^{-1}$$

$$N_8 = \sigma_{8\text{TeV}} \times 20 \text{ fb}^{-1}$$

$$N_{14} = \sigma_{14\text{TeV}} \times 3 \text{ ab}^{-1}$$

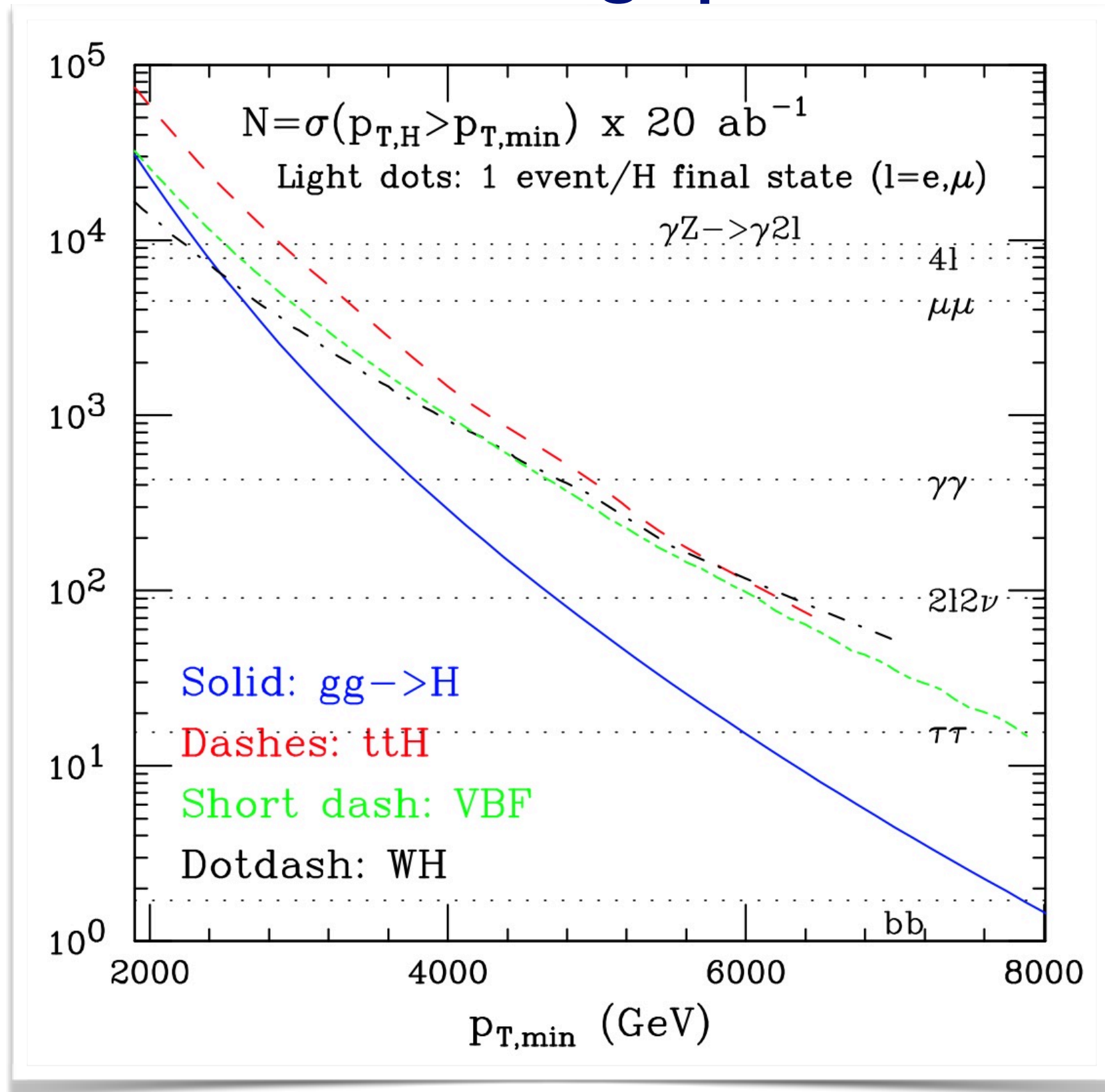
- Huge production rates imply:
 - can afford reducing statistics, with tighter kinematical cuts that reduce backgrounds and systematics
 - can explore new dynamical regimes, where new tests of the SM and EWSB can be done

H at large p_T



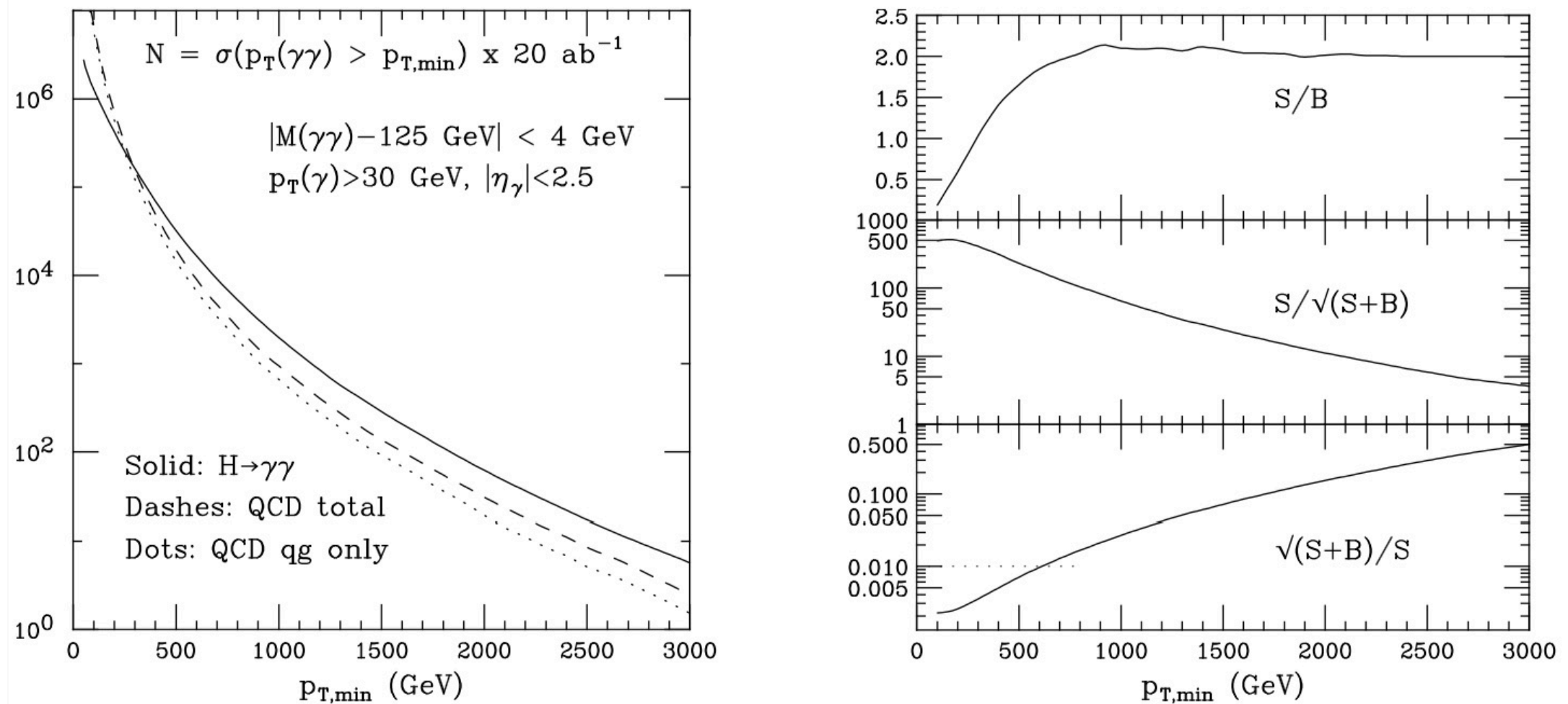
- Hierarchy of production channels changes at large $p_T(H)$:
 - $\sigma(ttH) > \sigma(gg \rightarrow H)$ above 800 GeV
 - $\sigma(VBF) > \sigma(gg \rightarrow H)$ above 1800 GeV

H at large p_T



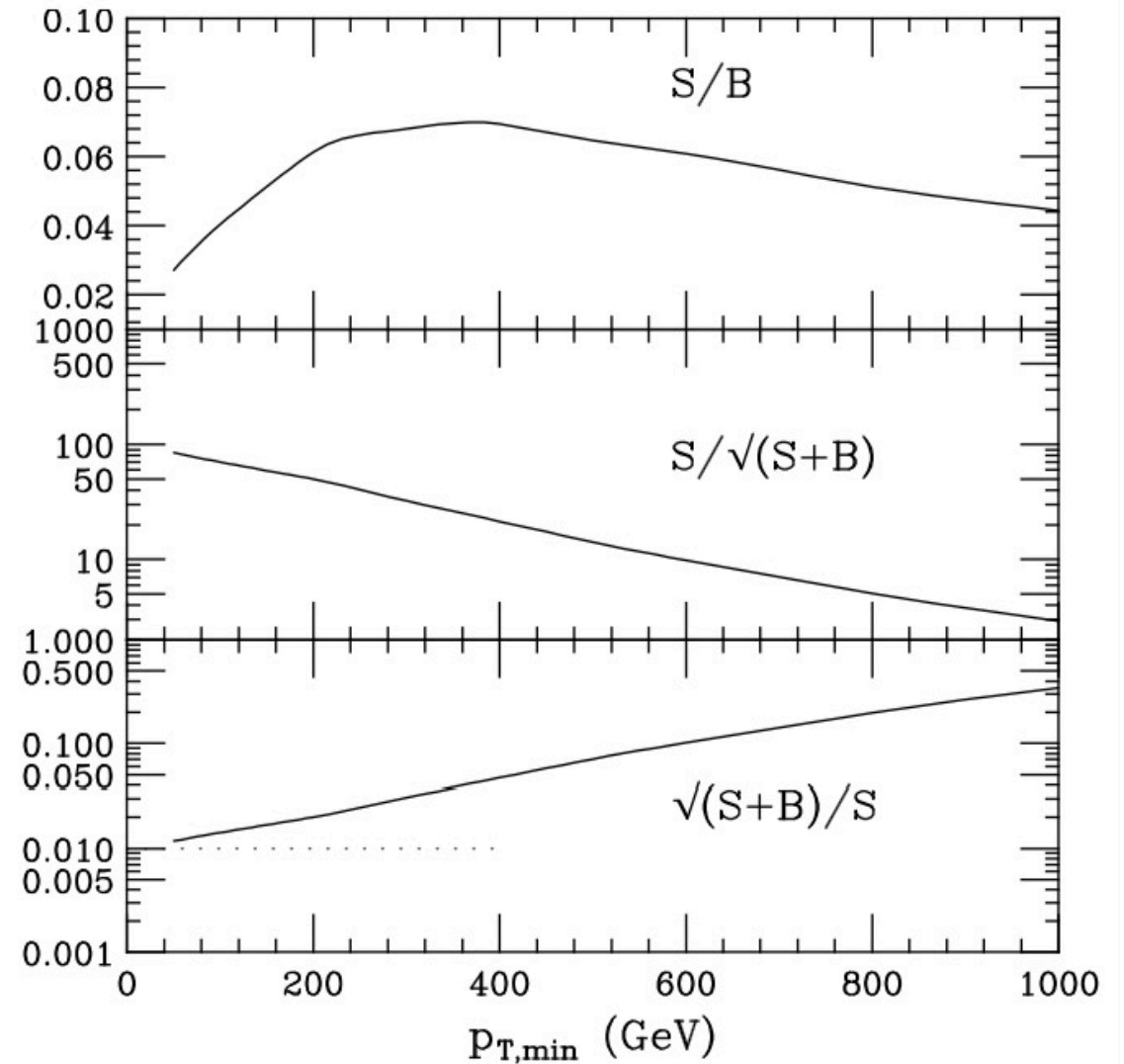
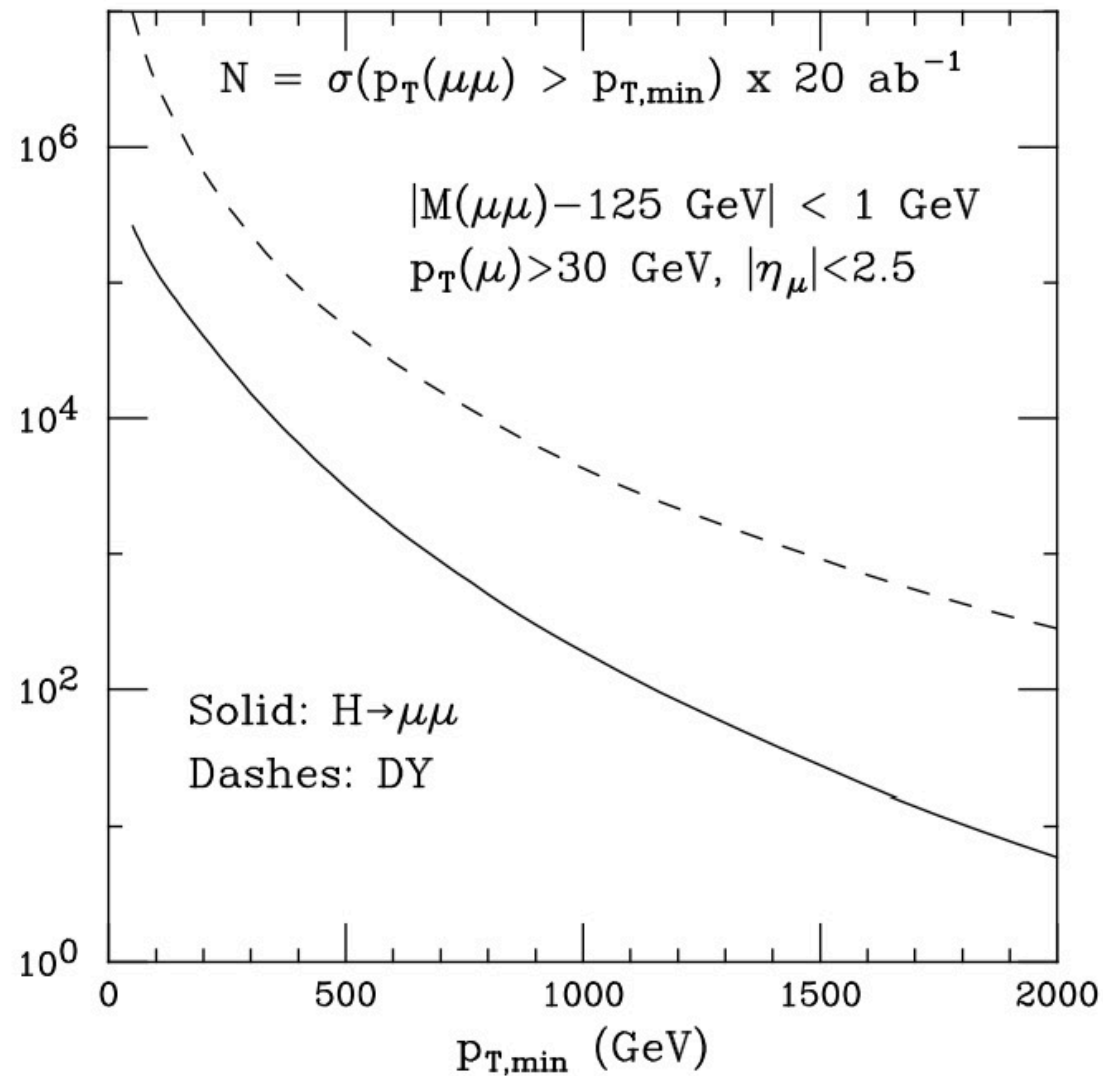
- Statistics in potentially visible final states out to several TeV

$gg \rightarrow H \rightarrow \gamma\gamma$ at large p_T



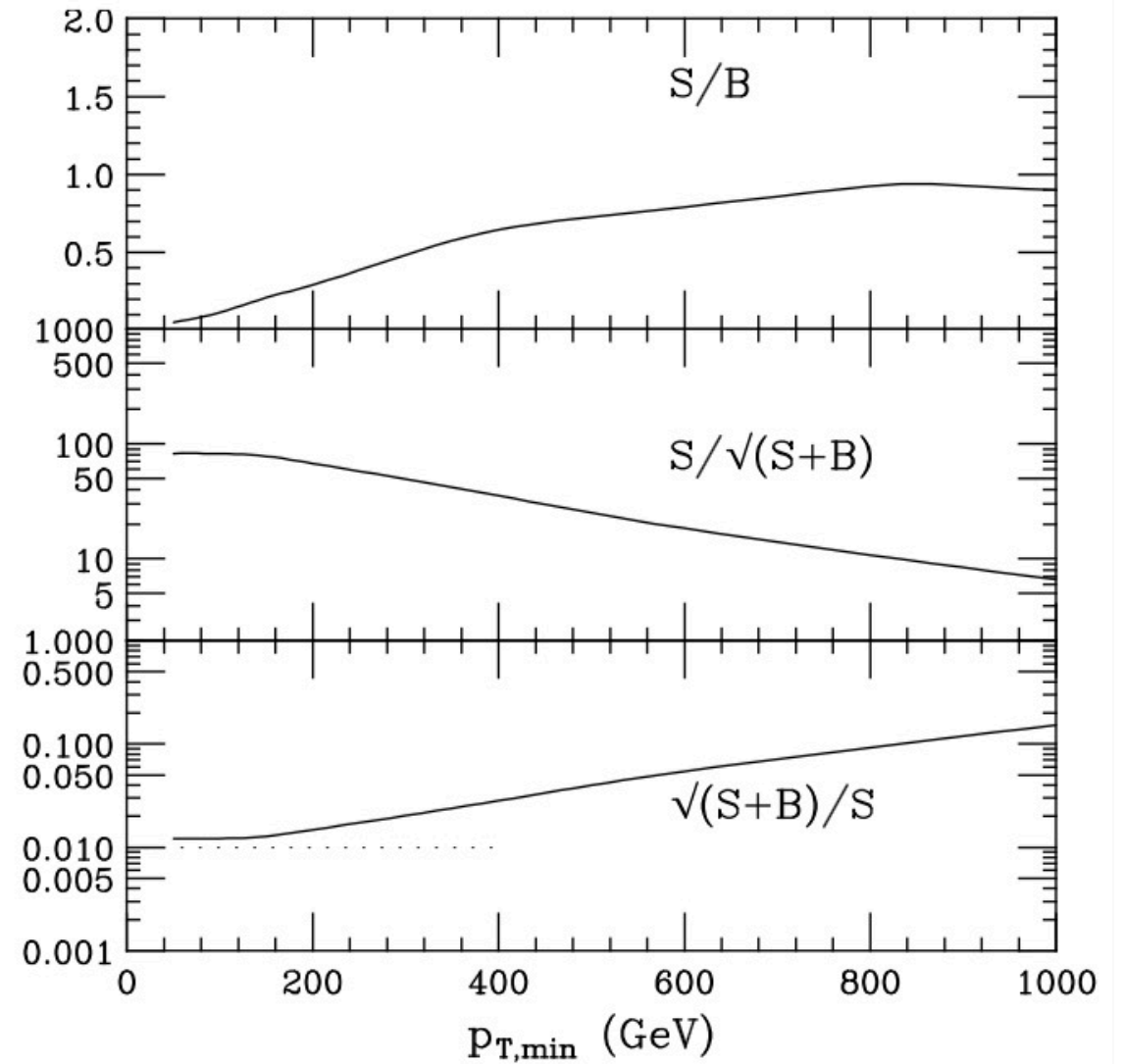
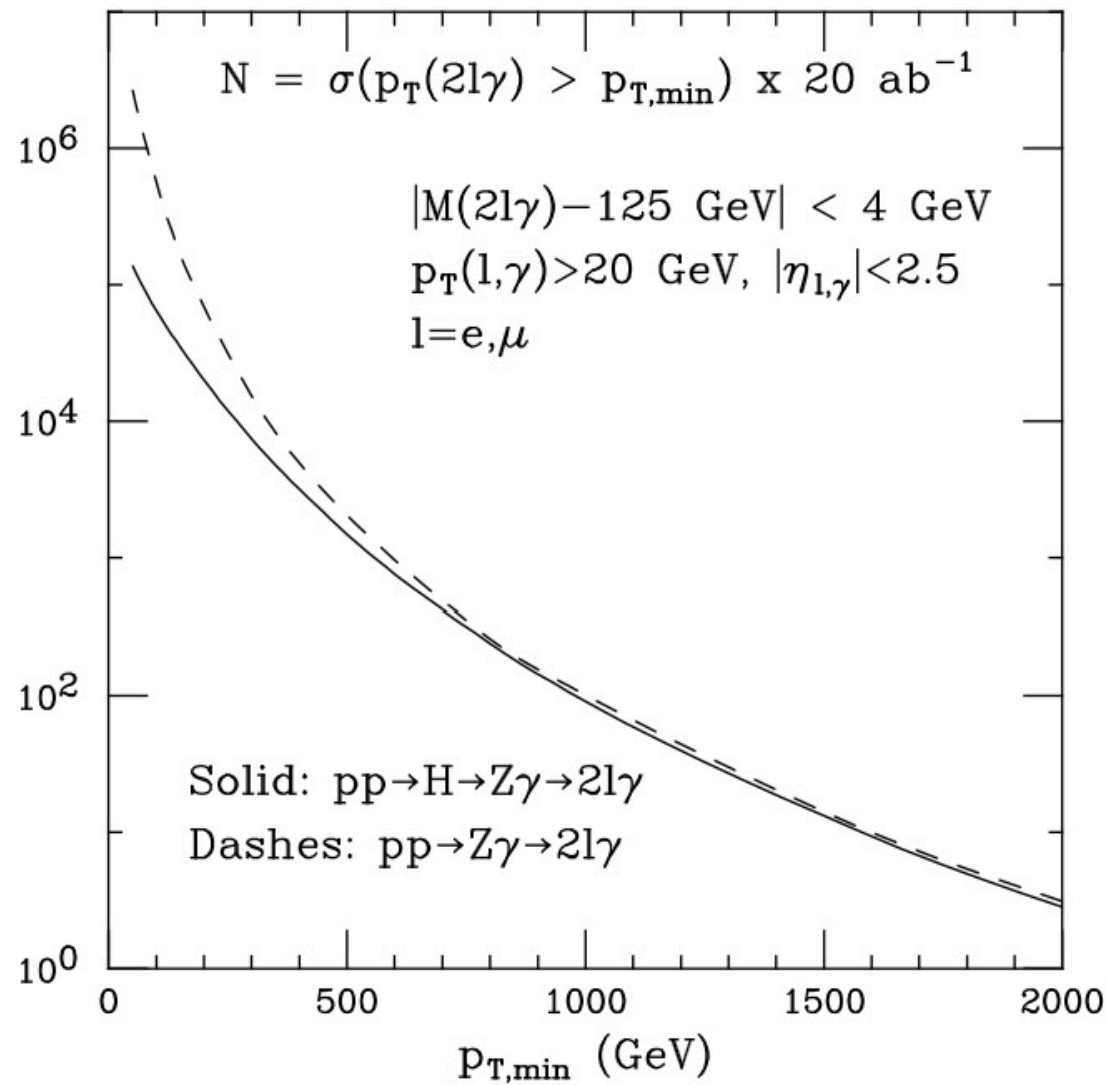
- At LHC, S/B in the $H \rightarrow \gamma\gamma$ channel is $O(\text{few } \%)$
- At FCC, for $p_T(H) > 300 \text{ GeV}$, $S/B \sim 1$
- Very clean probe of Higgs production up to large $p_T(H)$.
What's the sensitivity required to probe relevant BSM deviations from SM spectrum?
- Exptl mass resolution at large $p_T(H)$?

$gg \rightarrow H \rightarrow \mu\mu$ at large p_T



- Stat reach $\sim 1\%$ at $p_T \sim 100 \text{ GeV}$
- Exptl systematics on $BR(\mu\mu)/BR(\gamma\gamma)$? (use same fiducial selection to remove H modeling syst's)

$gg \rightarrow H \rightarrow Z\gamma \rightarrow \ell\ell\gamma$ at large p_T



- S/B improves greatly at larger p_T
- Stat reach $\sim 1\%$ at $p_T \sim 100 \text{ GeV}$
- Exptl systematics on $BR(Z\gamma)/BR(\gamma\gamma)$?

Remarks on $gg \rightarrow H$

- Reach for %-level measurement of very rare decay modes ($Z\gamma$, $\mu\mu$) (absolute, if $B(\gamma\gamma)$ or $B(ZZ^*)$ known from e^+e^- , or relative w.r.t. $B(\gamma\gamma)$ using pp-only data)
- Much larger statistics and p_T reach for modes like WW and $\tau\tau$. Needs dedicated studies to check potential precision (e.g. systematics from corrections to common fiducial regions, impact of neutrinos, ...)
- Reach for $H \rightarrow b\bar{b}$?

Open questions, for future work

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- How far can one go in setting constraints on the ggH coupling using $p_T(H)$ in the multi-TeV domain?

Open questions, for future work

- How far can one go in setting constraints on the ggH coupling using $p_T(H)$ in the multi-TeV domain?
- How do these constraints compare with
 - direct detection of possible new particles in the loop?
 - precise determination of $BR(H \rightarrow gg)$ from e^+e^- (e.g. analysis of EFT couplings)?

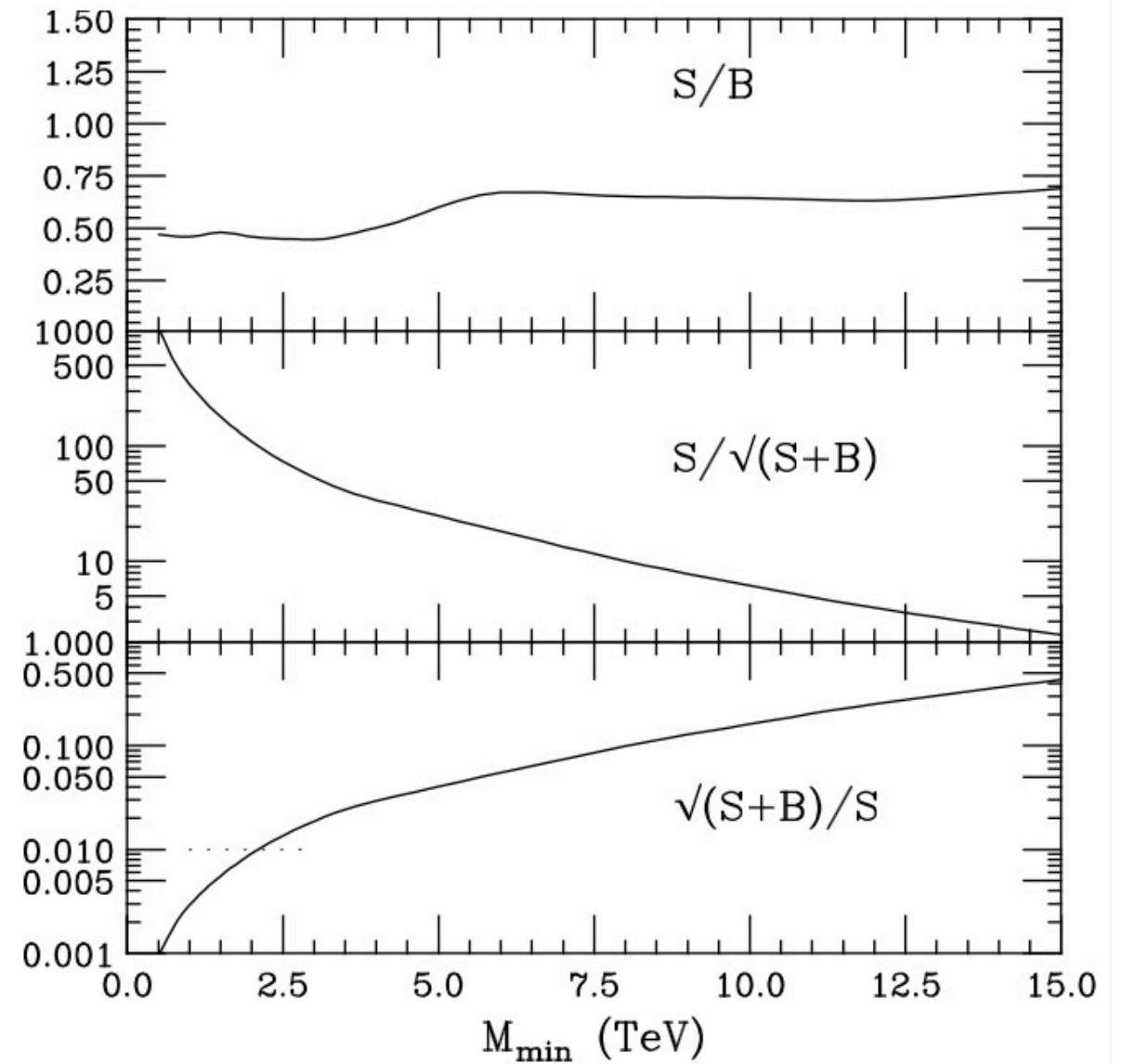
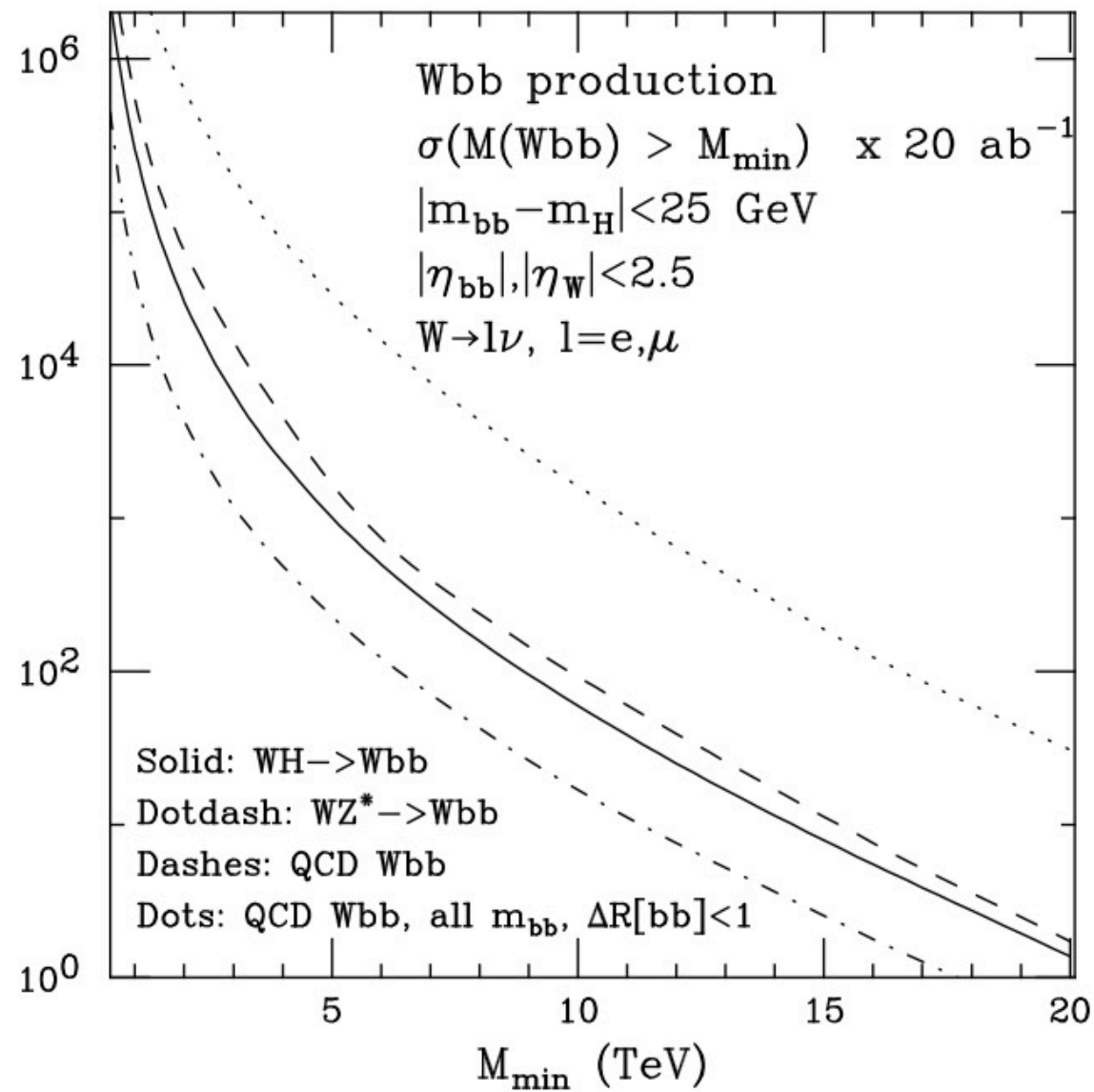
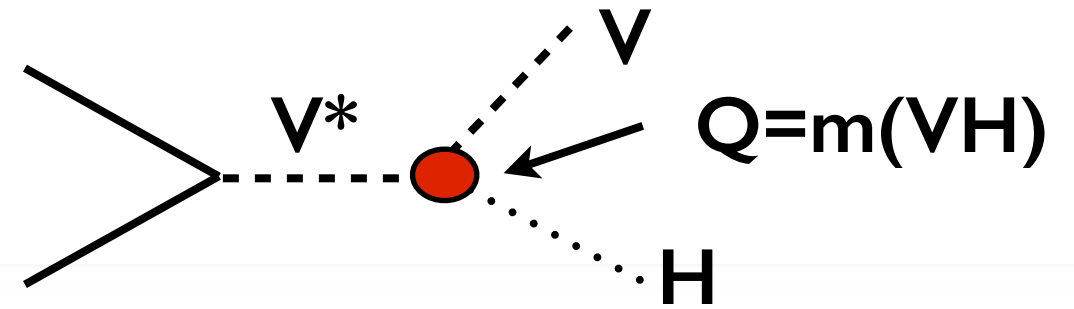
Open questions, for future work

- How far can one go in setting constraints on the ggH coupling using $p_T(H)$ in the multi-TeV domain?
- How do these constraints compare with
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 - precise determination of $BR(H \rightarrow gg)$ from e^+e^- (e.g. analysis of EFT couplings)?
- Analyses likely based on shape (e.g. $\sigma(p_T > 2 \text{ TeV})/\sigma(p_T > 1 \text{ TeV})$), to reduce dependence on absolute production rate, ttH coupling, lumi, etc:
 - ultimate TH systematics?
 - ultimate EXP systematics?
 - what are the best decay channels?

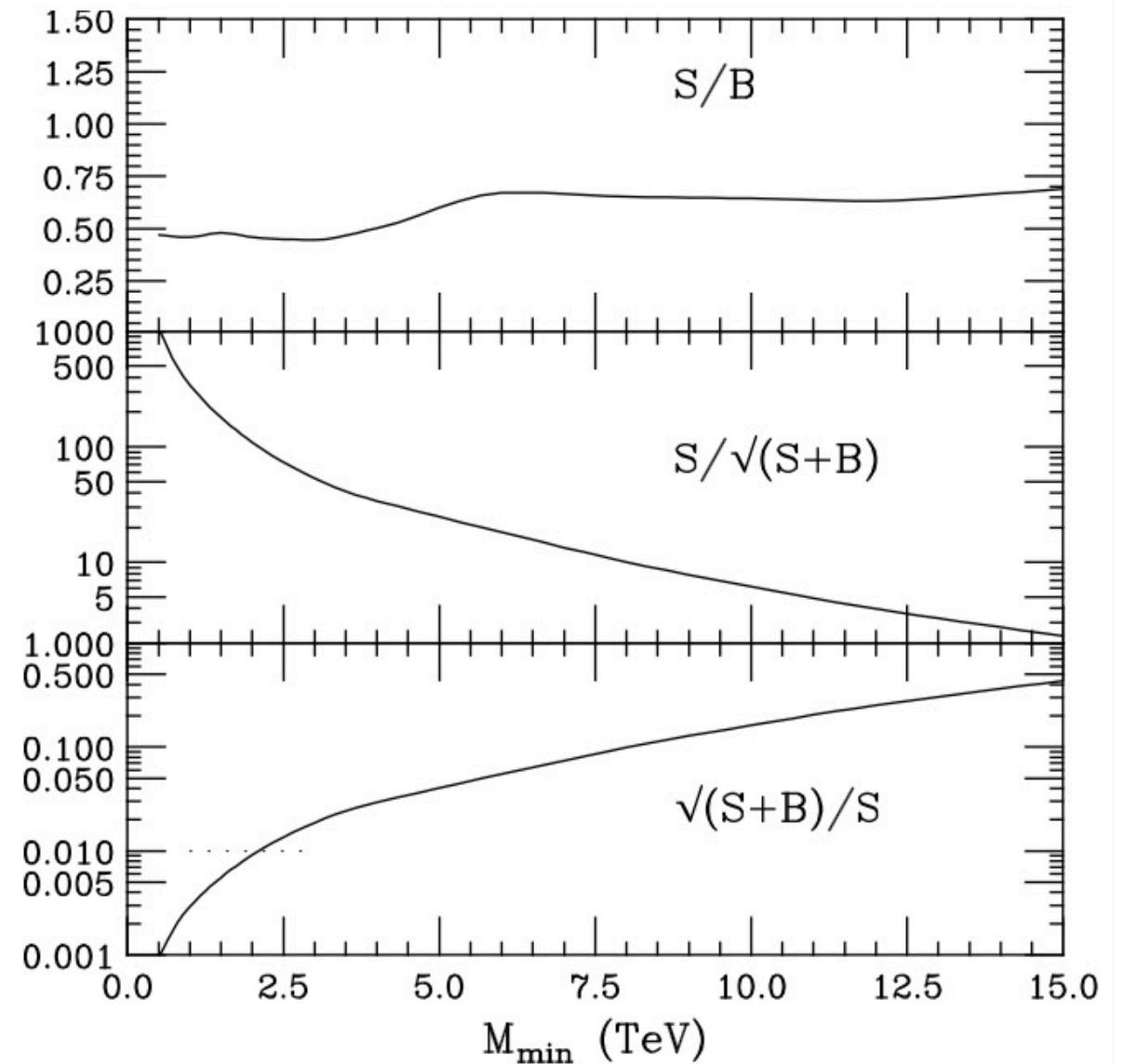
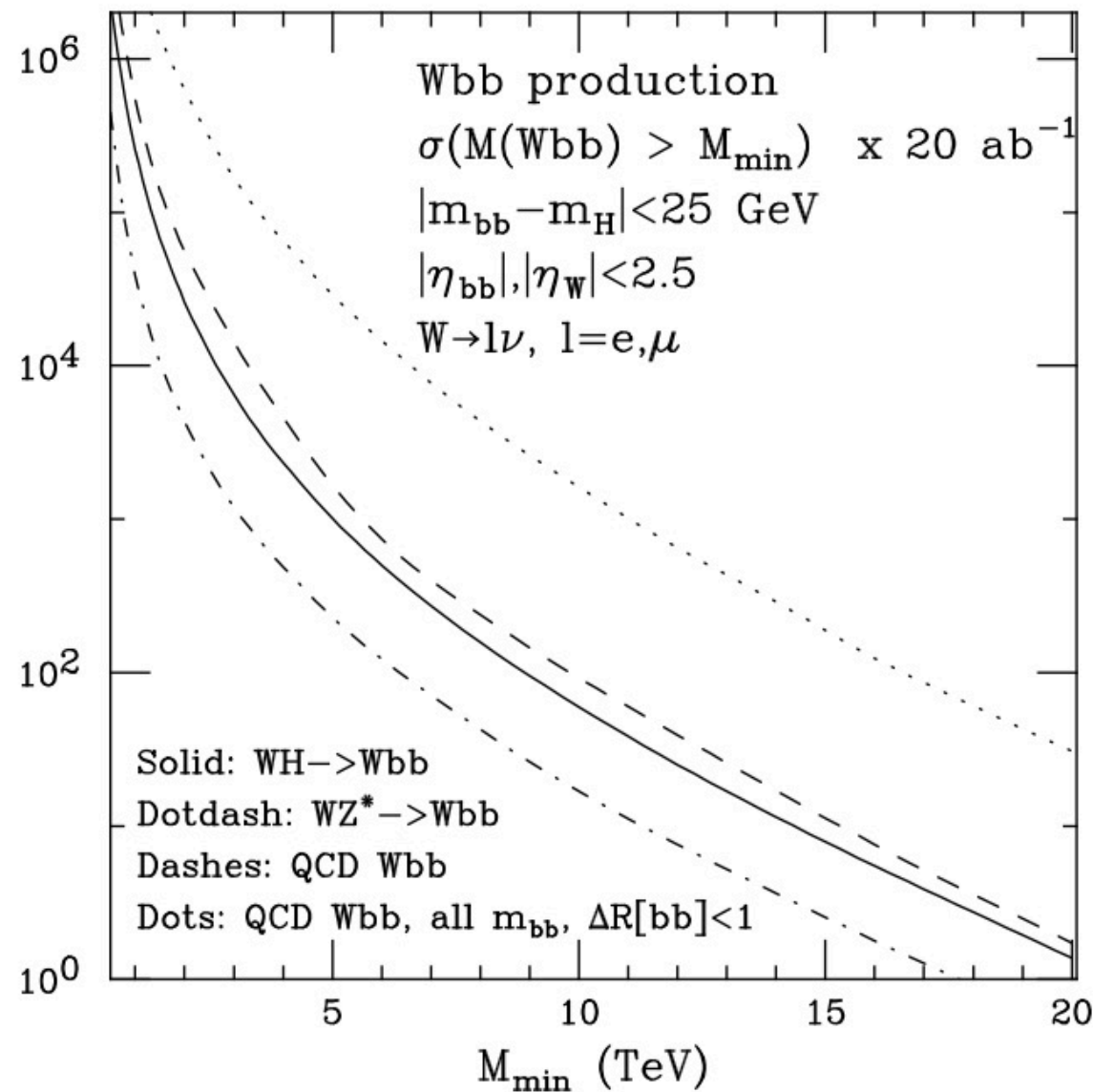
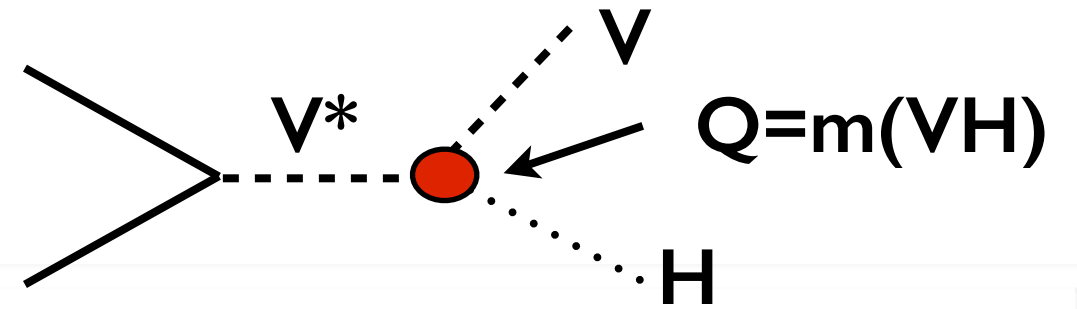
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 - ultimate TH systematics?
 - ultimate EXP systematics?
 - what are the best decay channels?
- More in general (for all production and decay channels):
 - *Can high- p_T measurements compete with precise BR's in probing EFT couplings?*

$WH \rightarrow Wbb$ at large M_{WH}

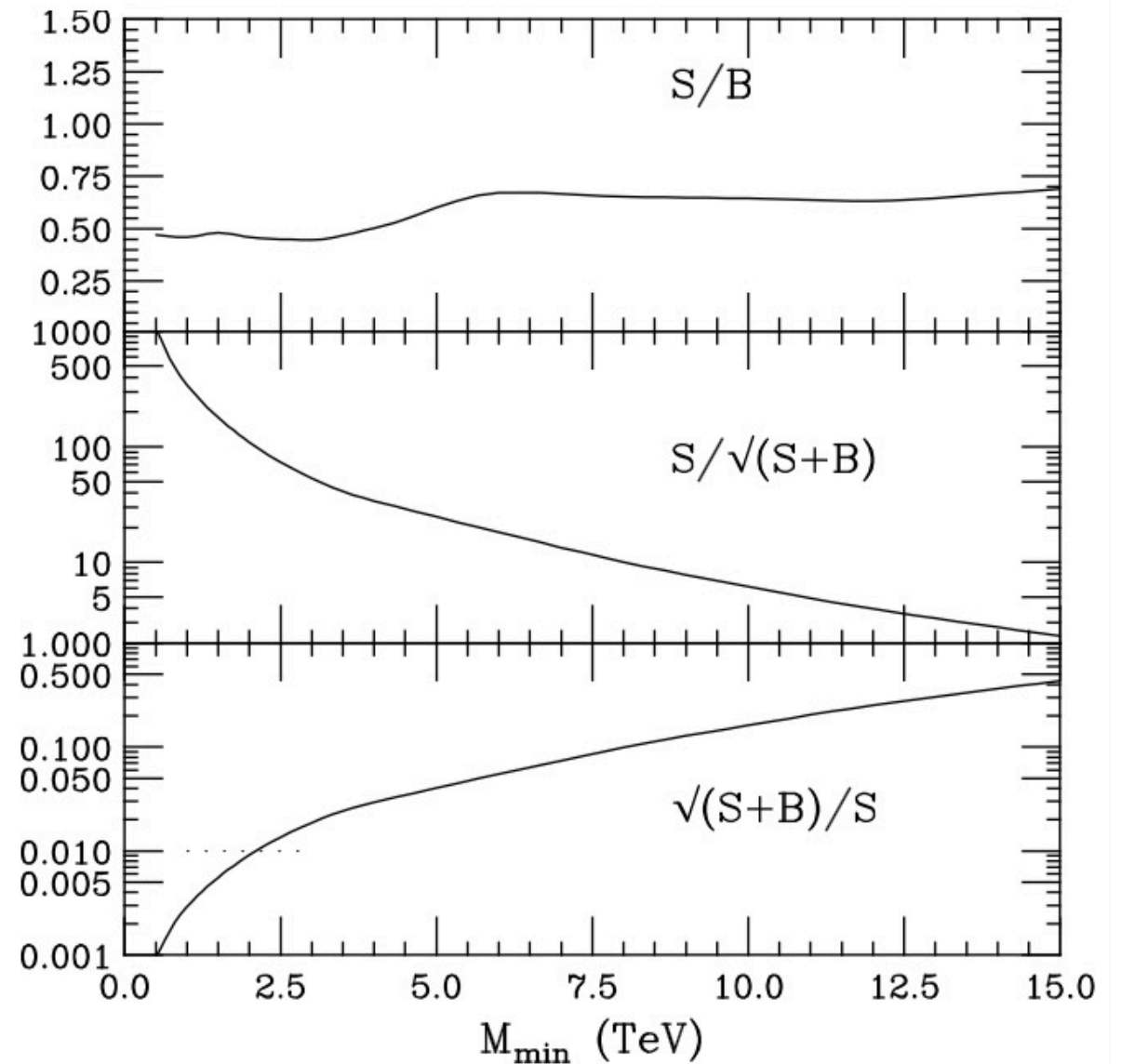
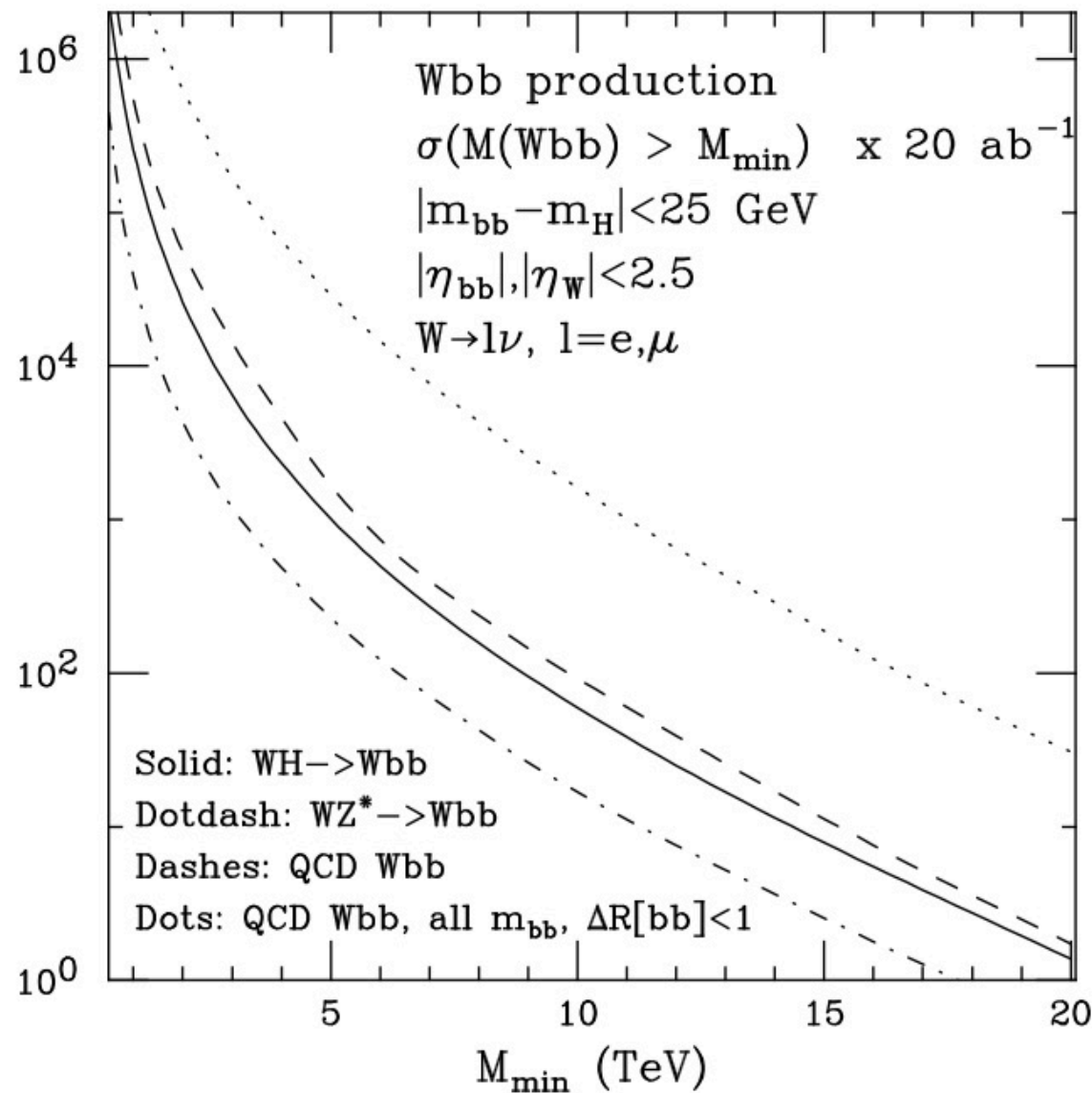
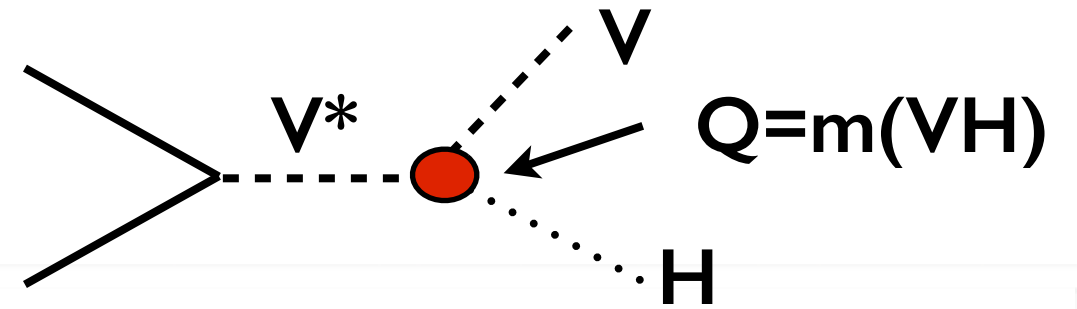


WH → Wbb at large M_{WH}



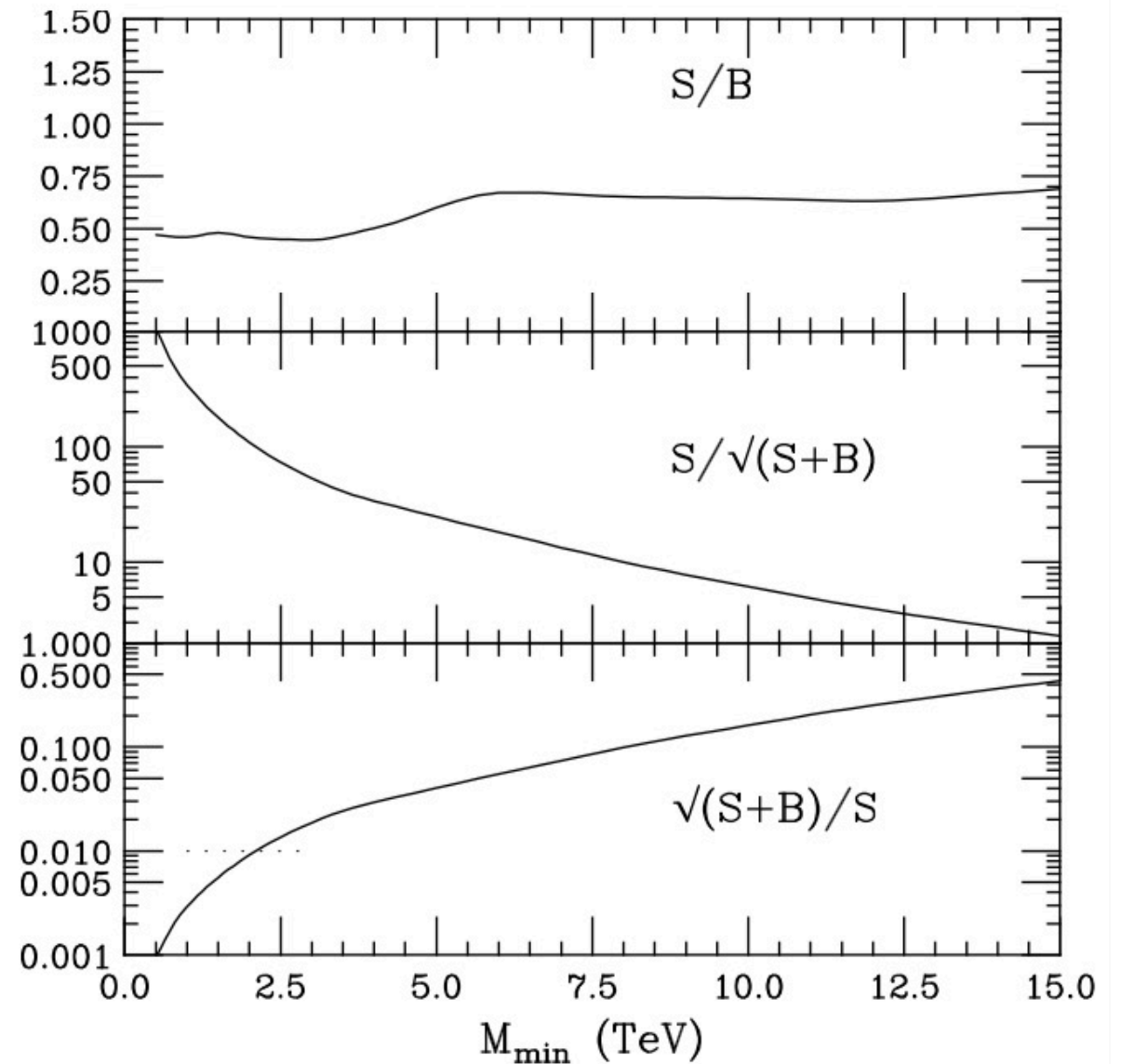
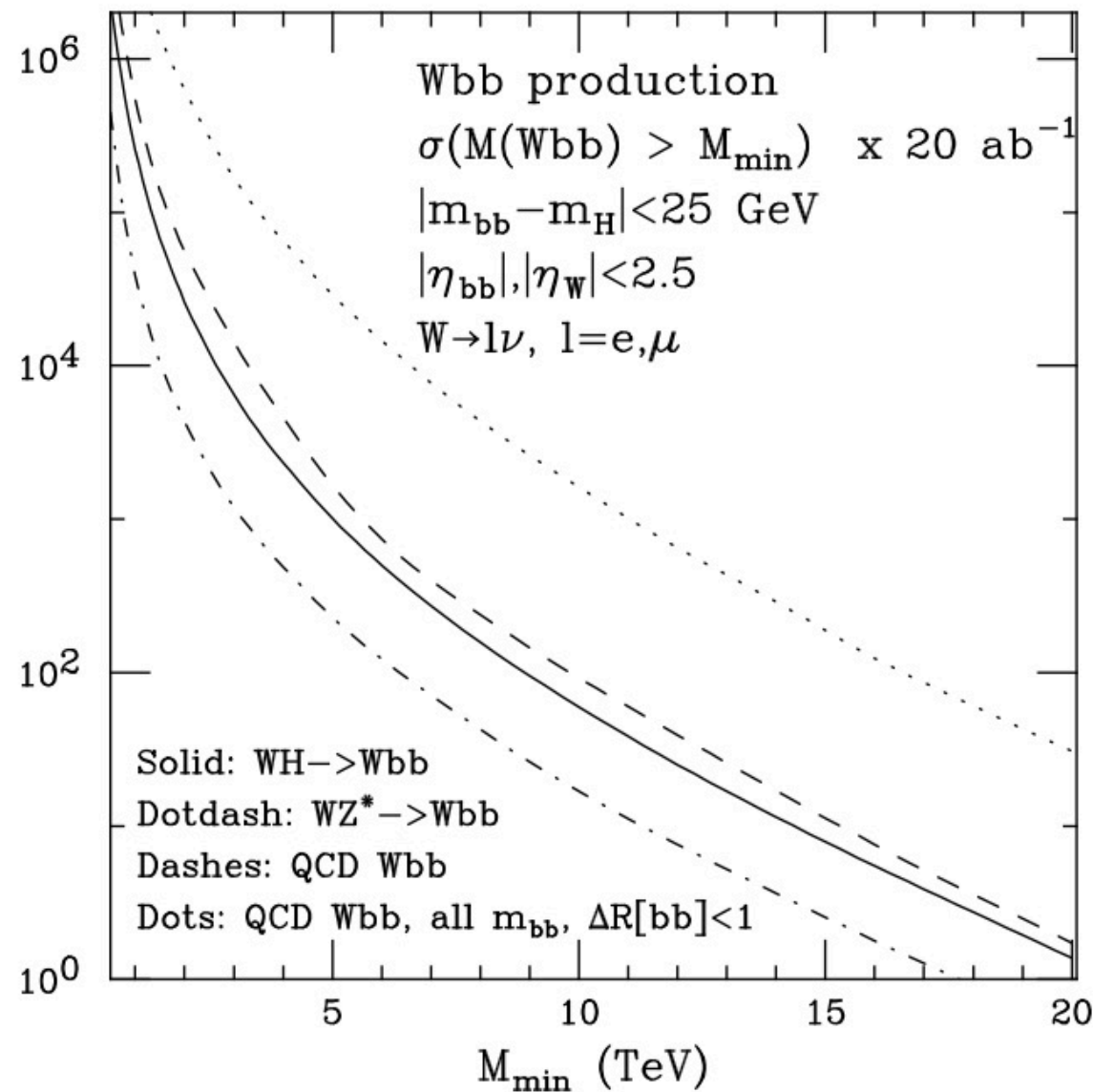
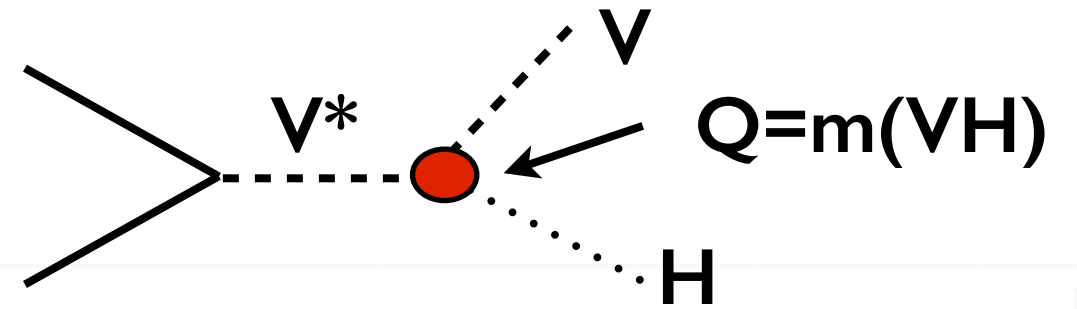
- Bg level greatly sensitive to bb mass resolution. Can be improved using jet substructure studies? => more work required

WH → Wbb at large M_{WH}

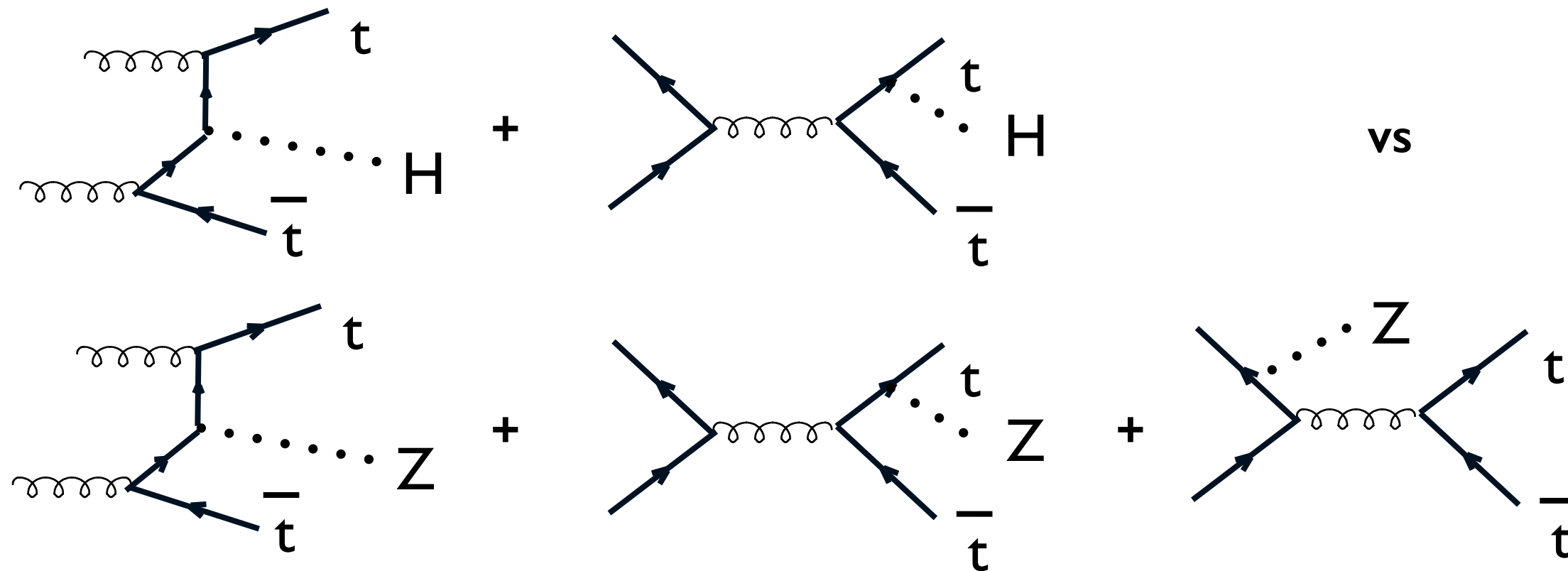


- Bg level greatly sensitive to bb mass resolution. Can be improved using jet substructure studies? \Rightarrow more work required
- Sensitivity to higher-dim ops in the VVH coupling $\Leftrightarrow B(H \rightarrow VV^*)$?

$WH \rightarrow Wbb$ at large M_{WH}



- Bg level greatly sensitive to bb mass resolution. Can be improved using jet substructure studies? \Rightarrow more work required
- Sensitivity to higher-dim ops in the VVH coupling $\Leftrightarrow B(H \rightarrow VV^*)$?
- Systematics on slope of M_{HV} ? (For EFT constraints don't need absolute rate)



To the extent that the $q\bar{q} \rightarrow tt Z/H$ contributions are subdominant:

- Identical production dynamics:

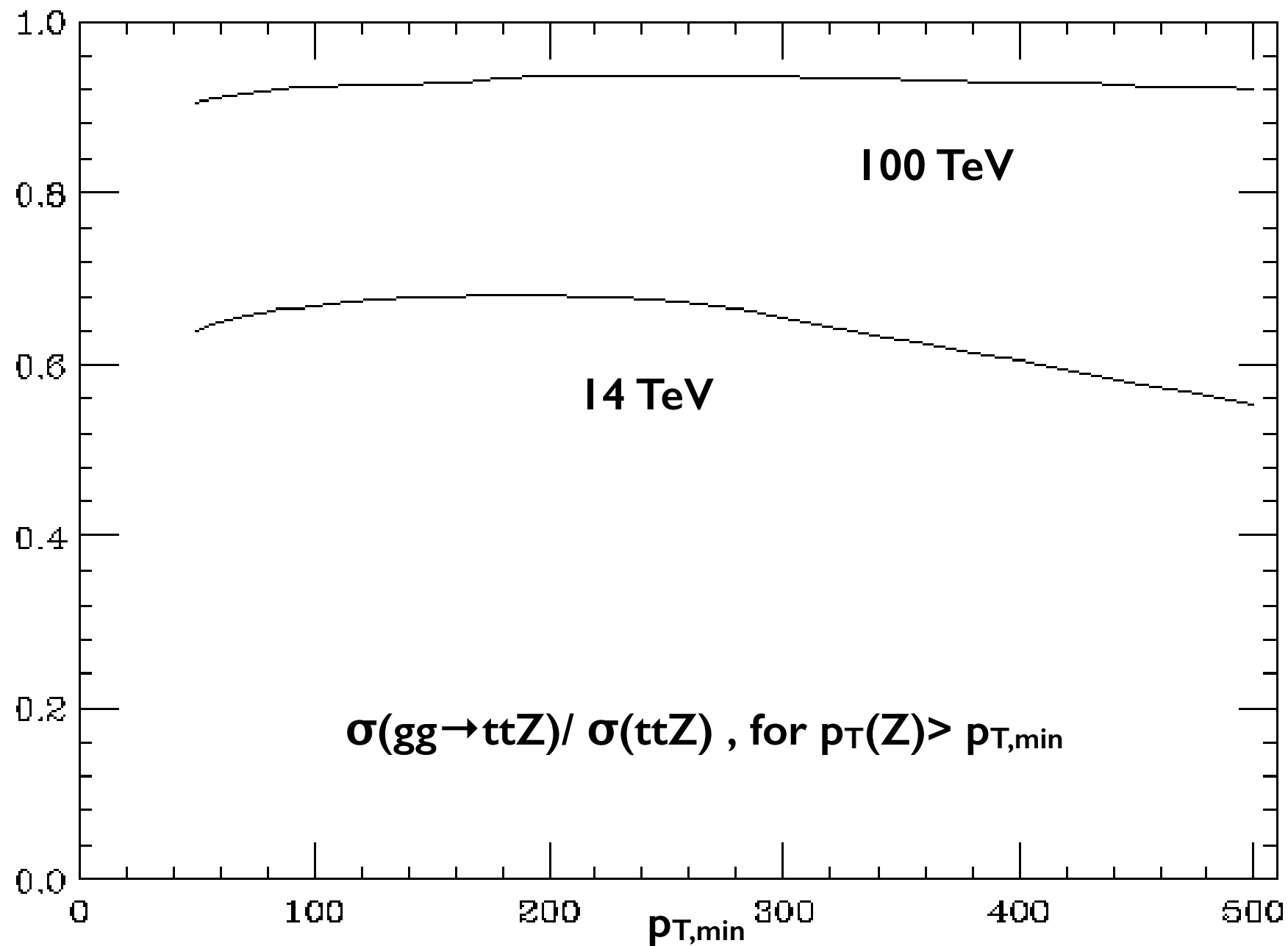
- o correlated QCD corrections, correlated scale dependence
- o correlated α_s systematics

- $m_Z \sim m_H \Rightarrow$ almost identical kinematic boundaries:

- o correlated PDF systematics
- o correlated m_{top} systematics

For a given y_{top} , we expect $\sigma(ttH)/\sigma(ttZ)$ to be predicted with great precision


At 100 TeV, $gg \rightarrow tt X$ is indeed dominant



NB: At lower p_T values, gg fraction is slightly larger for ttZ than for ttH , since $m_Z < m_H$

Cross section ratio stability

	$\sigma(t\bar{t}H)[\text{pb}]$	$\sigma(t\bar{t}Z)[\text{pb}]$	$\frac{\sigma(t\bar{t}H)}{\sigma(t\bar{t}Z)}$
13 TeV	$0.475^{+5.79\%+3.33\%}_{-9.04\%-3.08\%}$	$0.785^{+9.81\%+3.27\%}_{-11.2\%-3.12\%}$	$0.606^{+2.45\%+0.525\%}_{-3.66\%-0.319\%}$
100 TeV	$33.9^{+7.06\%+2.17\%}_{-8.29\%-2.18\%}$	$57.9^{+8.93\%+2.24\%}_{-9.46\%-2.43\%}$	$0.585^{+1.29\%+0.314\%}_{-2.02\%-0.147\%}$

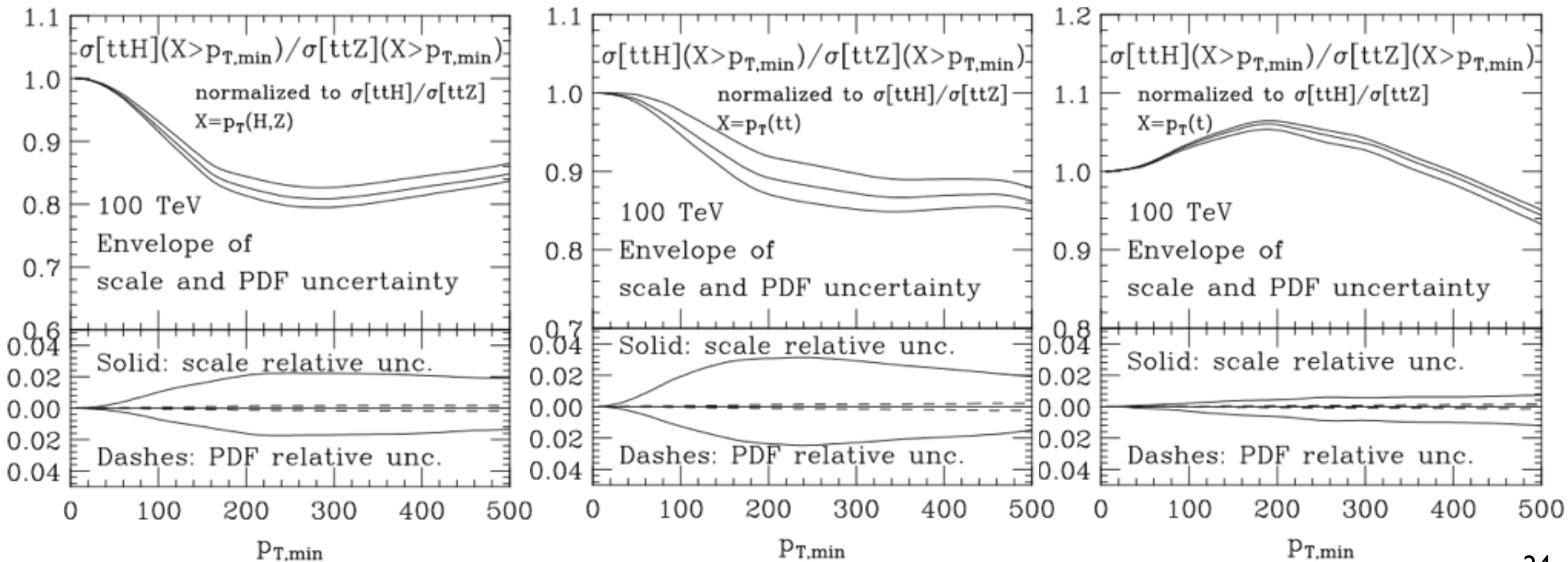

 scale PDF

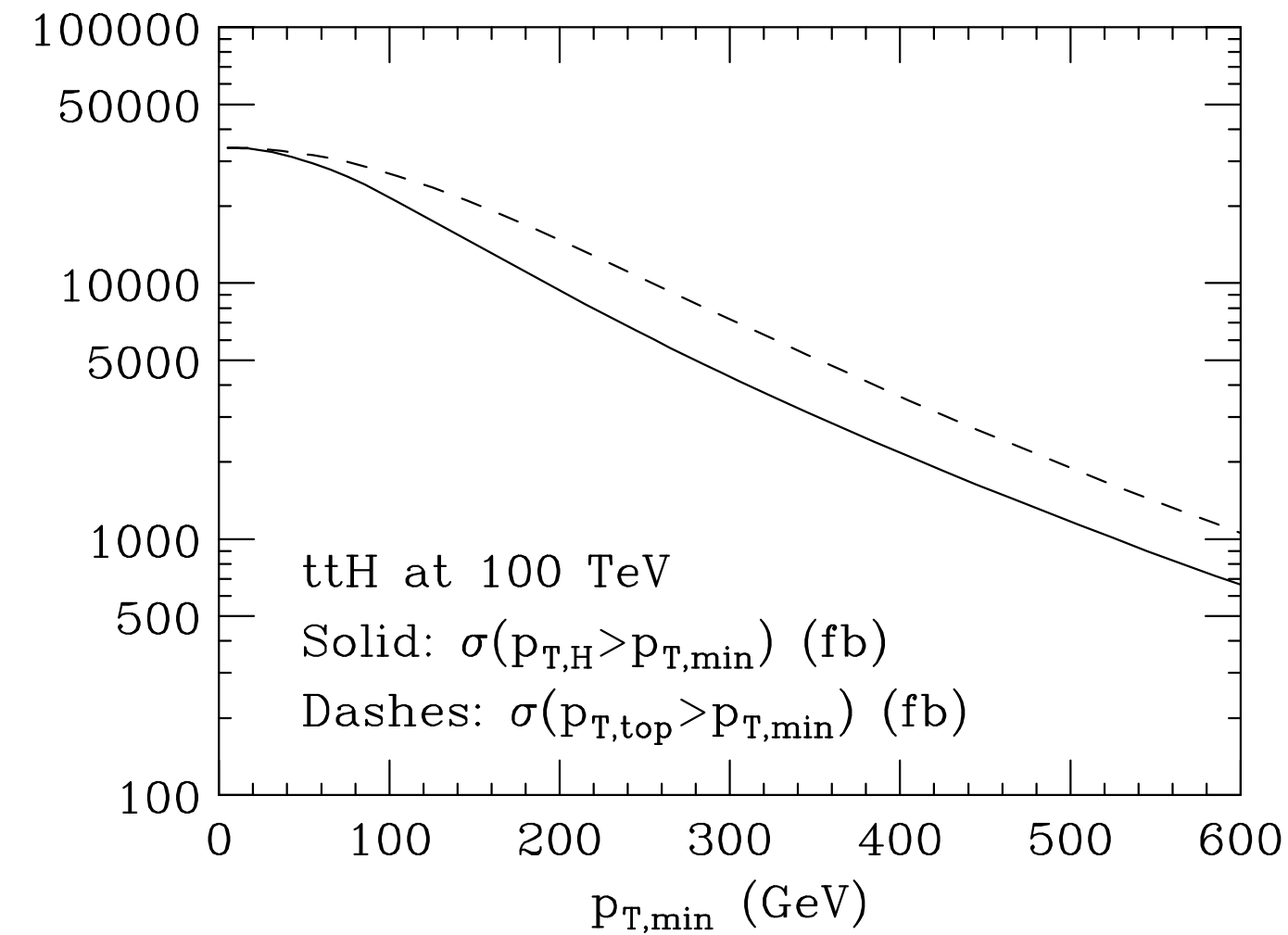
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↑ scale ↑ PDF

Production kinematics ratio stability

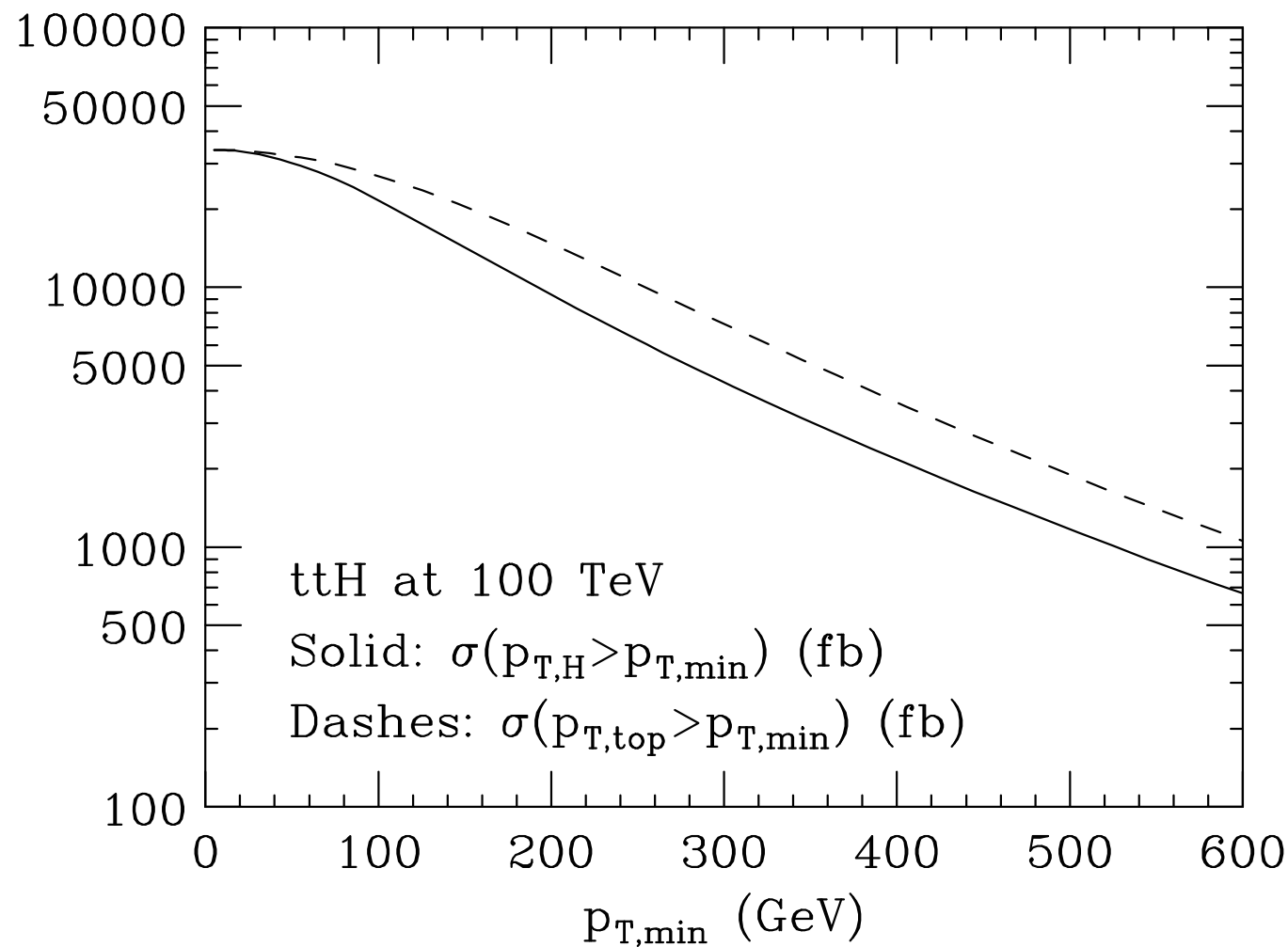




$H \rightarrow 4\ell$	$H \rightarrow \gamma\gamma$	$H \rightarrow 2\ell 2\nu$	$H \rightarrow b\bar{b}$
$2.6 \cdot 10^4$	$4.6 \cdot 10^5$	$2.0 \cdot 10^6$	$1.2 \cdot 10^8$

Events/ 20ab^{-1} , with $t\bar{t} \rightarrow \ell\nu + \text{jets}$

\Rightarrow huge rates, exploit
 boosted topologies



Top fat C/A jet(s) with $R = 1.2$, $|y| < 2.5$,
 and $p_{T,j} > 200$ GeV

- δy_t (stat + syst $_{TH}$) $\sim 1\%$

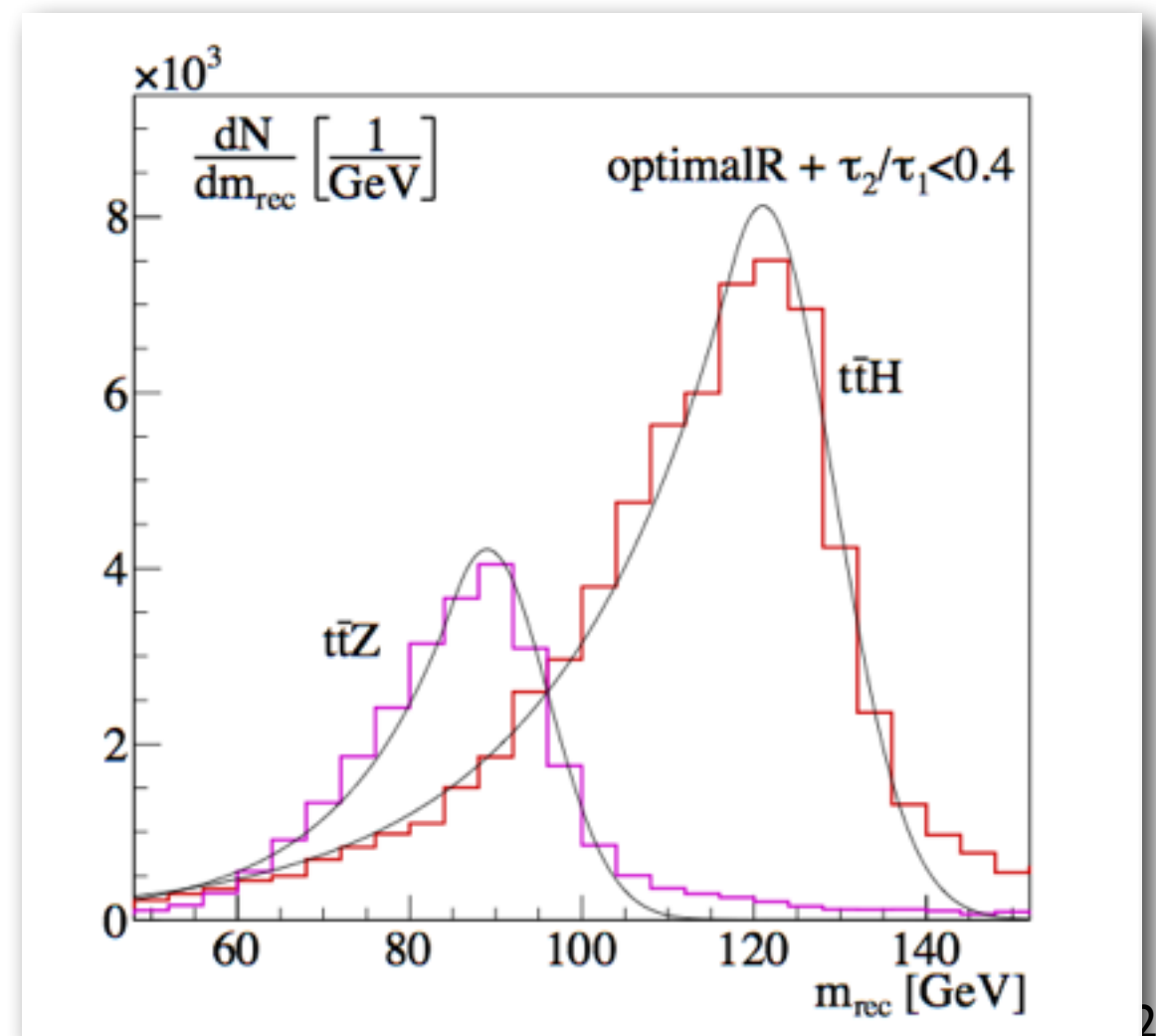
- great potential to reduce to similar
 levels δ_{exp} syst

- consider other decay modes, e.g. $2l2\nu$

$H \rightarrow 4\ell$	$H \rightarrow \gamma\gamma$	$H \rightarrow 2\ell 2\nu$	$H \rightarrow b\bar{b}$
$2.6 \cdot 10^4$	$4.6 \cdot 10^5$	$2.0 \cdot 10^6$	$1.2 \cdot 10^8$

Events/ 20ab^{-1} , with $tt \rightarrow \ell\nu + \text{jets}$

\Rightarrow huge rates, exploit
 boosted topologies



Rare production modes: any good use for them?

$pp \rightarrow HW^+W^-$ (4FS)	$4.62 \cdot 10^0$ $^{+3\%}_{-2\%}$ $^{+2\%}_{-2\%}$	$1.68 \cdot 10^2$ $^{+5\%}_{-6\%}$ $^{+2\%}_{-1\%}$	36
$pp \rightarrow HZW^\pm$	$2.17 \cdot 10^0$ $^{+4\%}_{-4\%}$ $^{+2\%}_{-2\%}$	$9.94 \cdot 10^1$ $^{+6\%}_{-7\%}$ $^{+2\%}_{-1\%}$	46
$pp \rightarrow HW^\pm\gamma$	$2.36 \cdot 10^0$ $^{+3\%}_{-3\%}$ $^{+2\%}_{-2\%}$	$7.75 \cdot 10^1$ $^{+7\%}_{-8\%}$ $^{+2\%}_{-1\%}$	33
$pp \rightarrow HZ\gamma$	$1.54 \cdot 10^0$ $^{+3\%}_{-2\%}$ $^{+2\%}_{-2\%}$	$4.29 \cdot 10^1$ $^{+5\%}_{-7\%}$ $^{+2\%}_{-2\%}$	28
$pp \rightarrow HZZ$	$1.10 \cdot 10^0$ $^{+2\%}_{-2\%}$ $^{+2\%}_{-2\%}$	$4.20 \cdot 10^1$ $^{+4\%}_{-6\%}$ $^{+2\%}_{-1\%}$	38
$pp \rightarrow HW^\pm j$	$3.18 \cdot 10^2$ $^{+4\%}_{-4\%}$ $^{+2\%}_{-1\%}$	$1.07 \cdot 10^4$ $^{+2\%}_{-7\%}$ $^{+2\%}_{-1\%}$	34
$pp \rightarrow HW^\pm jj$	$6.06 \cdot 10^1$ $^{+6\%}_{-8\%}$ $^{+1\%}_{-1\%}$	$4.90 \cdot 10^3$ $^{+2\%}_{-6\%}$ $^{+1\%}_{-1\%}$	81
$pp \rightarrow HZj$	$1.71 \cdot 10^2$ $^{+4\%}_{-4\%}$ $^{+1\%}_{-1\%}$	$6.31 \cdot 10^3$ $^{+2\%}_{-7\%}$ $^{+2\%}_{-1\%}$	37
$pp \rightarrow HZjj$	$3.50 \cdot 10^1$ $^{+7\%}_{-10\%}$ $^{+1\%}_{-1\%}$	$2.81 \cdot 10^3$ $^{+2\%}_{-5\%}$ $^{+1\%}_{-1\%}$	80

Table 1: Production of a single Higgs boson at the LHC and at a 100 TeV FCC-hh. The rightmost column reports the ratio ρ of the FCC-hh to the LHC cross sections. Theoretical uncertainties are due to scale and PDF variations, respectively. Monte-Carlo-integration error is always smaller than theoretical uncertainties, and is not shown. For $pp \rightarrow HVjj$, on top of the transverse-momentum cut of section 2, I require $m(j_1, j_2) > 100$ GeV, j_1 and j_2 being the hardest and next-to-hardest jets, respectively. Processes $pp \rightarrow Htj$ and $pp \rightarrow Hjj$ (VBF) do not feature jet cuts.

New analysis of HH production for the FCC report

R.C., C. Englert, G. Panico, A. Papaefstathiou, J. Ren, M. Selvaggi, M. Son, M. Spannowsky, W. Yao

- **Goals:**
 1. improve on previous studies and get a commonly-agreed estimate
 2. study dependence on efficiencies and systematics

Previous analyses:

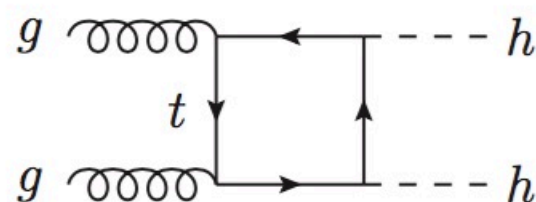
W. Yao arXiv:1308.6302 (Snowmass Summer Study 2013)

Barr, Dolan, Englert, de Lima, Spannowsky JHEP 1502 (2015) 016

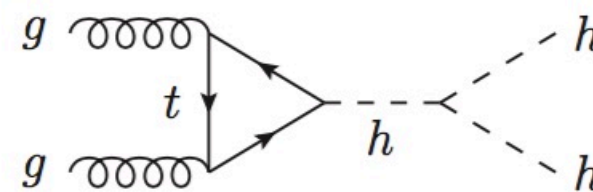
Azatov, R.C., Panico, Son PRD 92 (2015) 035001

H-J. He, J. Ren, W. Yao PRD 93 (2016) 015003

Signal: double Higgs production via gluon fusion ($gg \rightarrow hh$)



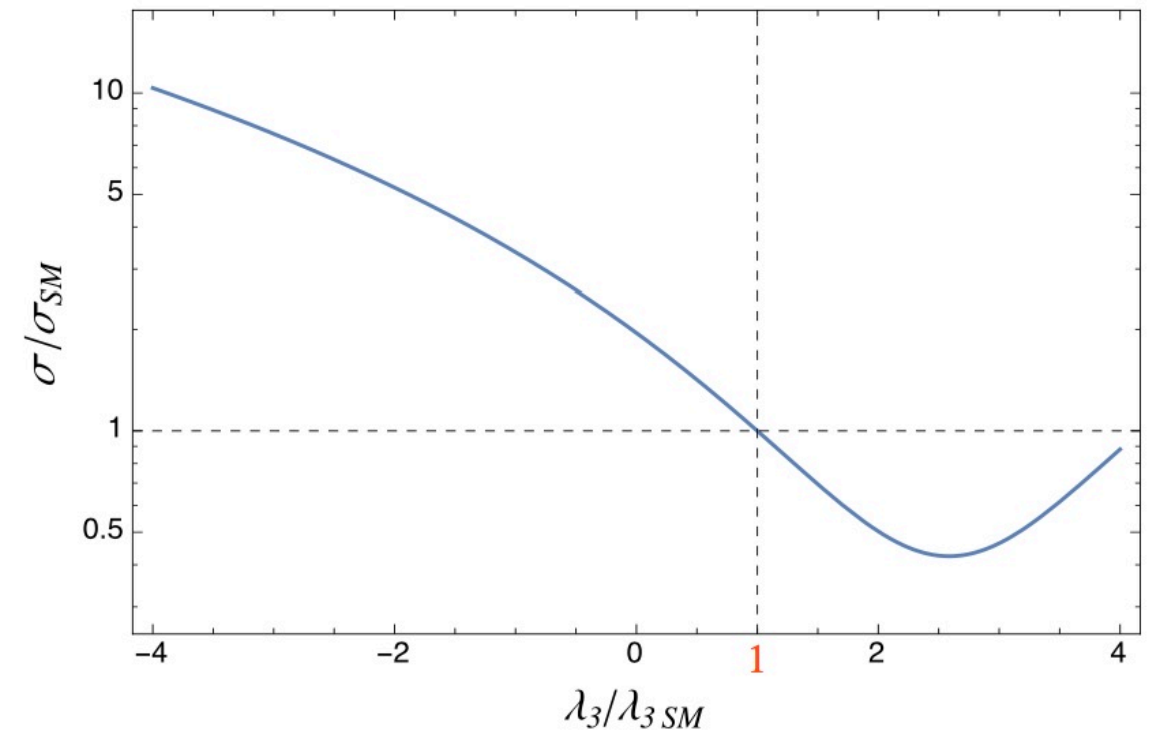
$\sim \text{const.}$



$$\sim \lambda_3 \times \frac{m_h^2}{\hat{s}} \log^2 \left(\frac{m_t^2}{\hat{s}} \right)$$

Most sensitivity on trilinear coupling comes from threshold events

	Signal cross section [fb] at NNLO+NNLL *		
14 TeV	$45.05^{+4.4\%}_{-6.0\%} \pm 3.0\% \pm 10\%$		
100 TeV	$1749^{+5.1\%}_{-6.6\%} \pm 2.7\% \pm 10\%$		
Theoretical uncertainties:		<div style="display: flex; justify-content: space-around; align-items: center;"> <div style="text-align: center;"> <div style="width: 40px; height: 10px; background-color: blue; margin: 0 auto;"></div> <p>scale</p> </div> <div style="text-align: center;"> <div style="width: 40px; height: 10px; background-color: blue; margin: 0 auto;"></div> <p>PDFs $+\alpha_s$</p> </div> <div style="text-align: center;"> <div style="width: 40px; height: 10px; background-color: blue; margin: 0 auto;"></div> <p>infinite m_t approx.</p> </div> </div>	<p style="color: red; font-size: 2em;">~ 40 × increase</p>



	# Higgs pairs to $b\bar{b}\gamma\gamma$
LHC: 14TeV 300fb ⁻¹	36
HL-LHC: 14TeV 3ab ⁻¹	360
FCC: 100TeV 20ab ⁻¹	92×10^3

← percent precision physics

Backgrounds:

$b\bar{b}\gamma\gamma$
 $t\bar{t}h(\gamma\gamma)$
 $b\bar{b}h(\gamma\gamma)$
 $jj\gamma\gamma$ (two fake b-jets)
 $b\bar{b}j\gamma$ (one fake photon)

* Results of the recent full- m_{top} NLO calculation (Borowka et al, arXiv: 1604.06447) not included here (as yet....)

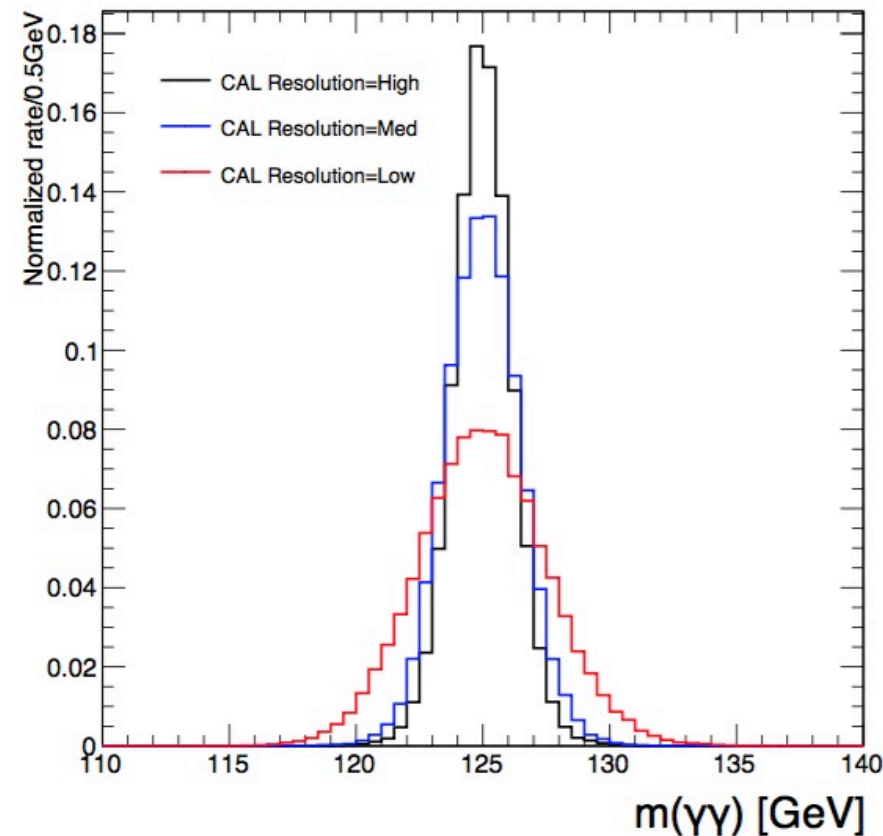
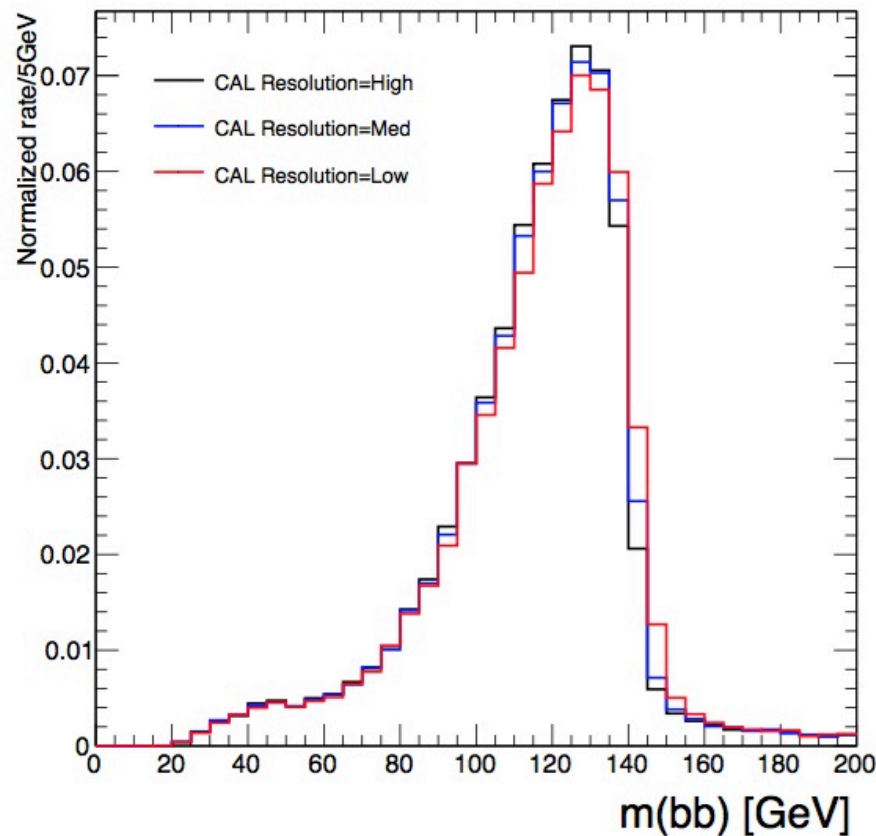
Montecarlo Simulation:

MadGraph5_aMC@NLO → Pythia 6 → Delphes (FCC card)

Three benchmark scenarios for ECAL and HCAL resolution:

$$\Delta E = \sqrt{a^2 E^2 + b^2 E}$$

	ECAL				HCAL			
	$ \eta \leq 4$		$4 < \eta \leq 6$		$ \eta \leq 4$		$4 < \eta \leq 6$	
	a	b	a	b	a	b	a	b
low	0.02	0.2	0.01	0.1	0.05	1.0	0.05	1.0
medium	0.01	0.1	0.01	0.1	0.03	0.5	0.05	1.0
high	0.007	0.06	0.01	0.1	0.01	0.3	0.03	0.5



High
 $\Delta m(\gamma\gamma) = 1.5$ GeV

Med
 $\Delta m(\gamma\gamma) = 2.0$ GeV

Low
 $\Delta m(\gamma\gamma) = 3.0$ GeV

- overall rescaling of background rate $n_B \rightarrow r_B \times n_B$

- uncertainty on signal rate $\Delta_S = \frac{\Delta\sigma(pp \rightarrow hh)}{\sigma(pp \rightarrow hh)}$

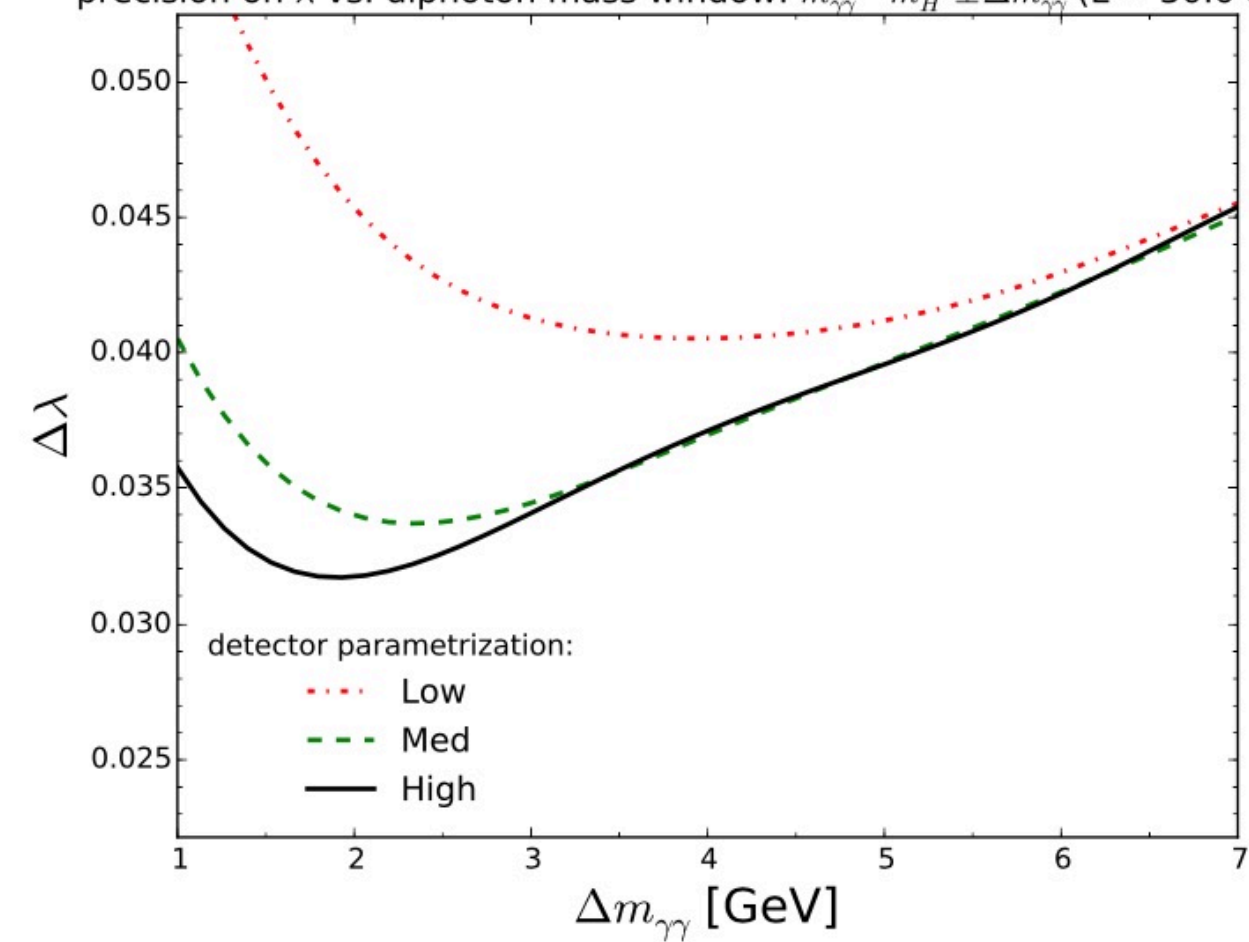
using “medium” calorimeter resolution

$\Delta\lambda_3$	$\Delta_S = 0.00$	$\Delta_S = 0.01$	$\Delta_S = 0.015$	$\Delta_S = 0.02$	$\Delta_S = 0.025$
$r_B = 0.5$	2.7%	3.4%	4.1%	4.9%	5.8%
$r_B = 1.0$	3.4%	3.9%	4.6%	5.3%	6.1%
$r_B = 1.5$	3.9%	4.4%	5.0%	5.7%	6.4%
$r_B = 2.0$	4.4%	4.8%	5.4%	6.0%	6.8%
$r_B = 3.0$	5.2%	5.6%	6.0%	6.6%	7.3%

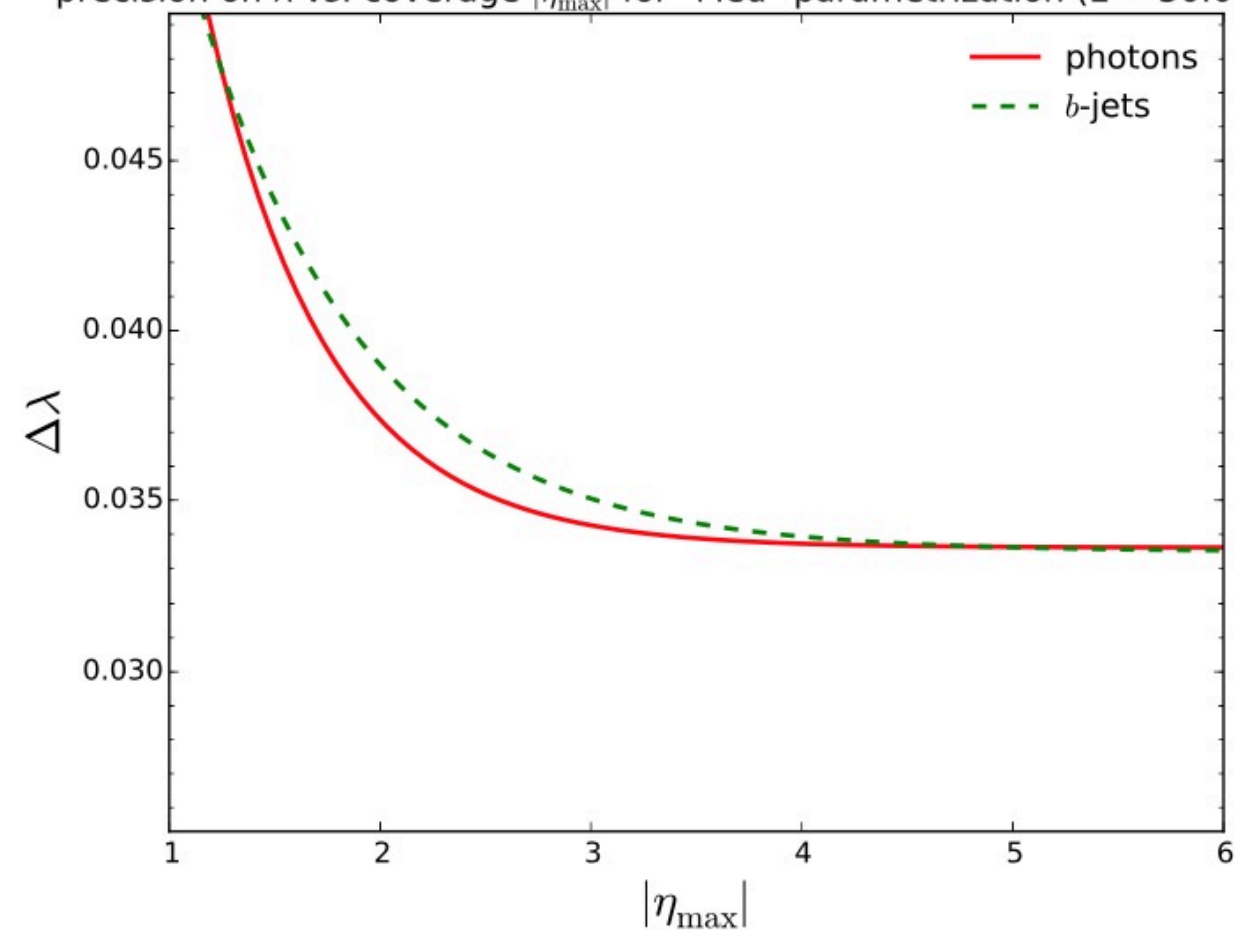
For $\Delta_S \gtrsim 2.5\%$ the precision on λ_3 is dominated
by the theory error on the signal: $\Delta\lambda_3 \simeq 2\Delta_S$

impact of detector performance, I

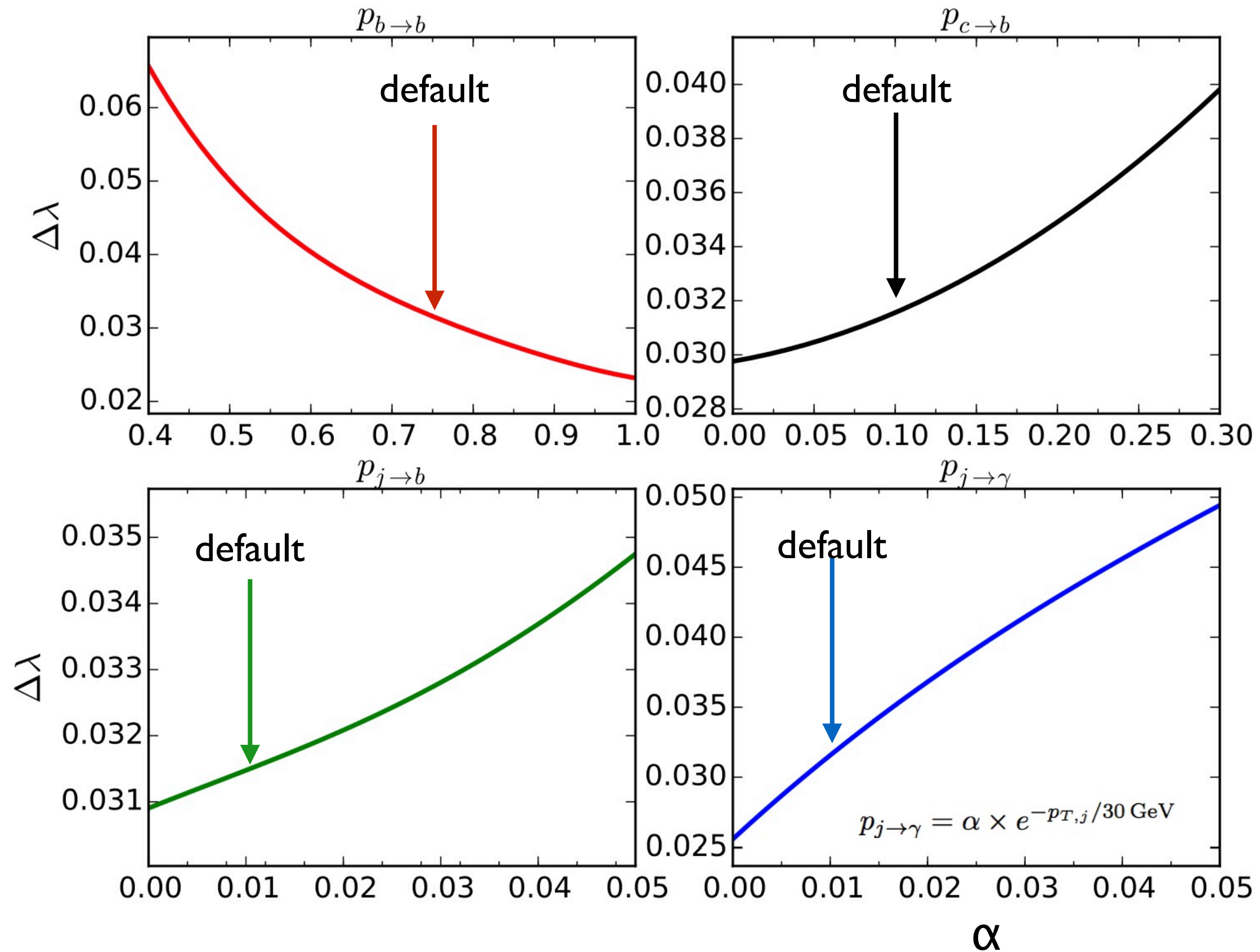
precision on λ vs. diphoton mass window: $m_{\gamma\gamma} = m_H \pm \Delta m_{\gamma\gamma}$ ($L = 30.0 \text{ ab}^{-1}$)



precision on λ vs. coverage $|\eta_{\text{max}}|$ for "Med" parametrization ($L = 30.0 \text{ ab}^{-1}$)



impact of detector performance, 2

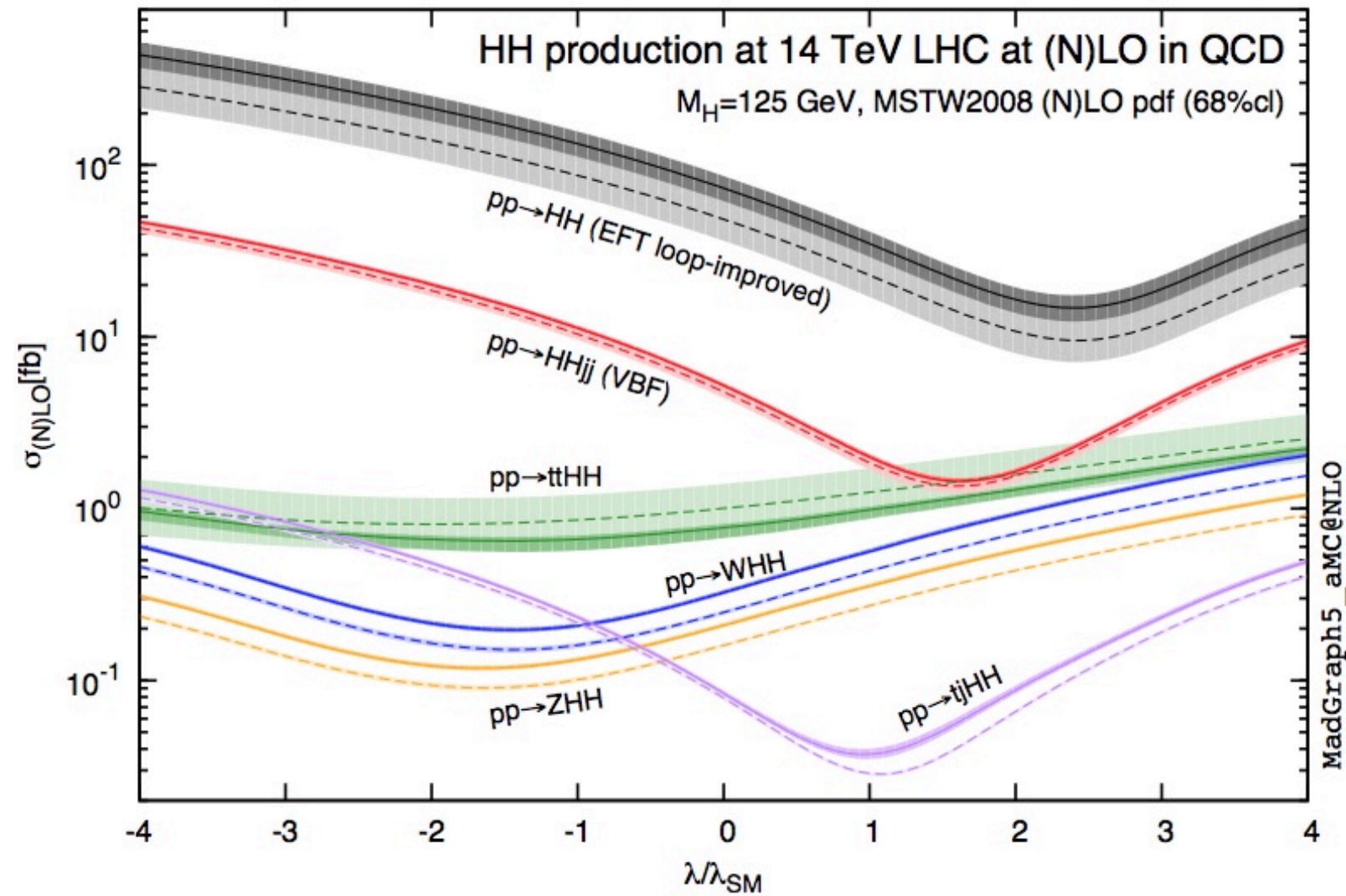


other HH+X production modes

process	$\sigma(14 \text{ TeV}) \text{ (fb)}$	$\sigma(100 \text{ TeV}) \text{ (fb)}$	accuracy
HH (ggf)	$45.05^{+4.4\%}_{-6.0\%} \pm 3.0\% \pm 10\%$	$1749^{+5.1\%}_{-6.6\%} \pm 2.7\% \pm 10\%$	NNLL matched to NNLO
$HHjj$ (VBF)	$1.94^{+2.3\%}_{-2.6\%} \pm 2.3\%$	$80.3^{+0.5\%}_{-0.4\%} \pm 1.7\%$	NLO
HHZ	$0.415^{+3.5\%}_{-2.7\%} \pm 1.8\%$	$8.23^{+5.9\%}_{-4.6\%} \pm 1.7\%$	NNLO
HHW^+	$0.269^{+0.33\%}_{-0.39\%} \pm 2.1\%$	$4.70^{+0.90\%}_{-0.96\%} \pm 1.8\%$	NNLO
HHW^-	$0.198^{+1.2\%}_{-1.3\%} \pm 2.7\%$	$3.30^{+3.5\%}_{-4.3\%} \pm 1.9\%$	NNLO
$HHt\bar{t}$	$0.949^{+1.7\%}_{-4.5\%} \pm 3.1\%$	$82.1^{+7.9\%}_{-7.4\%} \pm 1.6\%$	NLO
$HHtj$	$0.0364^{+4.2\%}_{-1.8\%} \pm 4.7\%$	$4.44^{+2.2\%}_{-2.6\%} \pm 2.4\%$	NLO
HHH	$0.0892^{+14.8\%}_{-13.6\%} \pm 3.2\%$	$4.82^{+12.3\%}_{-11.9\%} \pm 1.8\%$	NLO

Table 25: Cross sections for production of two or three SM Higgs bosons, including associated production channels, at a 14 TeV and 100 TeV hadron collider [1]. The cross sections are computed by choosing $\mu = M_{hh}/2$ ($\mu = M_{hhh}/2$ in the case of triple production). The error intervals correspond to scale variation and PDF + α_s uncertainty. In HH production in the gluon-fusion channel a conservative 10% uncertainty is included to take into account the effects of the infinite top-mass approximation (see Section 3.1.1).

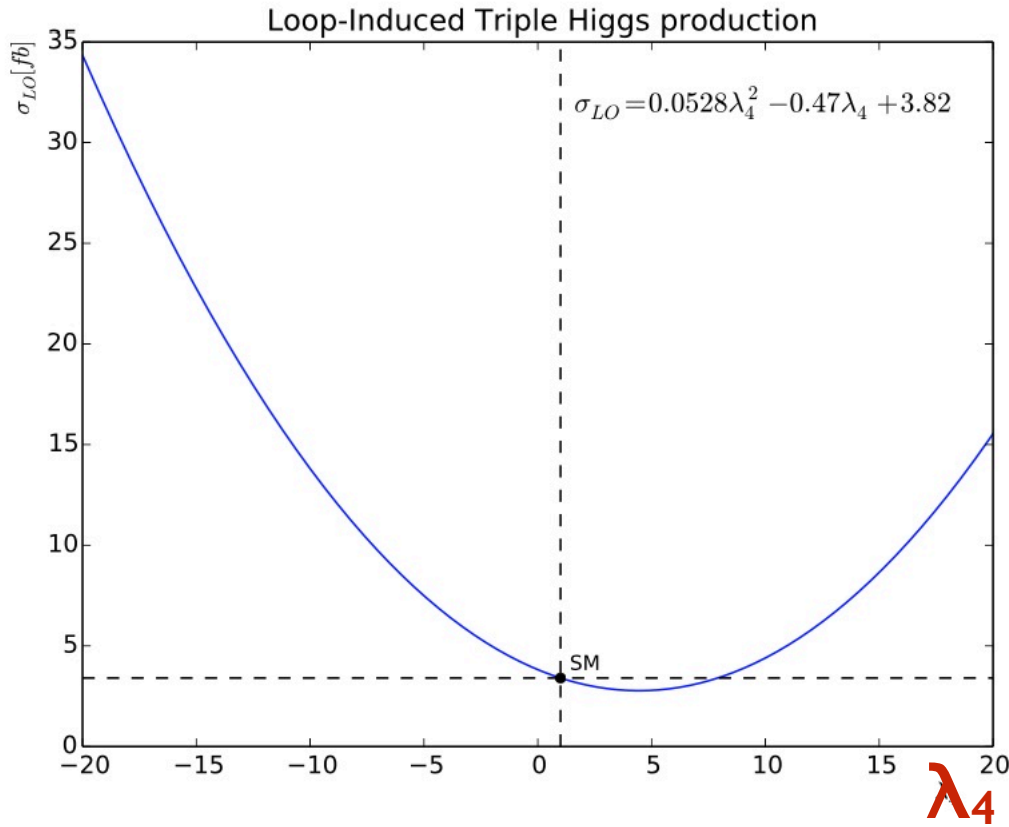
other channels, first assessments



λ dependence
at 14 and 100
TeV are similar

process	precision on σ_{SM}	68% CL interval on Higgs self-couplings
$HH \rightarrow b\bar{b}\gamma\gamma$	3%	$\lambda_3 \in [0.97, 1.03]$
$HH \rightarrow b\bar{b}b\bar{b}$	5%	$\lambda_3 \in [0.9, 1.5]$
$HH \rightarrow b\bar{b}4\ell$	$O(25\%)$	$\lambda_3 \in [0.6, 1.4]$
$HH \rightarrow b\bar{b}\ell^+\ell^-$	$O(15\%)$	$\lambda_3 \in [0.8, 1.2]$
$HH \rightarrow b\bar{b}\ell^+\ell^-\gamma$	—	—

Quartic Higgs selfcoupling



observable	selection cut
$p_{T,b\{1,2,3,4\}}$	$> \{80, 50, 40, 40\} \text{ GeV}$
$ \eta_b $	< 3.0
$m_{bb}^{\text{close},1}$	$\in [100, 160] \text{ GeV}$
$m_{bb}^{\text{close},2}$	$\in [90, 170] \text{ GeV}$
$\Delta R_{bb}^{\text{close},1}$	$\in [0.2, 1.6]$
$\Delta R_{bb}^{\text{close},2}$	no cut
$p_{T,\gamma\{1,2\}}$	$> \{70, 40\} \text{ GeV}$
$ \eta_\gamma $	< 3.5
$\Delta R_{\gamma\gamma}$	$\in [0.2, 4.0]$
$m_{\gamma\gamma}$	$\in [124, 126] \text{ GeV}$

process	$\sigma_{LO} \text{ (fb)}$	$\sigma_{NLO} \times \text{BR} \times \mathcal{P}_{\text{tag}} \text{ (ab)}$	$\epsilon_{\text{analysis}}$	$N_{30 \text{ ab}^{-1}}^{\text{cuts}}$
$hhh \rightarrow (b\bar{b})(b\bar{b})(\gamma\gamma), \text{ SM}$	2.89	5.4	0.06	9.7
$b\bar{b}b\bar{b}\gamma\gamma$	1.28	1050	2.6×10^{-4}	8.2
$hZZ, (\text{NLO}) (ZZ \rightarrow (b\bar{b})(b\bar{b}))$	0.817	0.8	0.002	$\ll 1$
$hhZ, (\text{NLO}) (Z \rightarrow (b\bar{b}))$	0.754	0.8	0.007	$\ll 1$
$hZ, (\text{NLO}) (Z \rightarrow (b\bar{b}))$	8.02×10^3	1130	$\mathcal{O}(10^{-5})$	$\ll 1$
$b\bar{b}b\bar{b}\gamma + \text{jets}$	2.95×10^3	2420	$\mathcal{O}(10^{-5})$	$\mathcal{O}(1)$
$b\bar{b}b\bar{b} + \text{jets}$	5.45×10^3	4460	$\mathcal{O}(10^{-6})$	$\ll 1$
$b\bar{b}\gamma\gamma + \text{jets}$	98.7	4.0	$\mathcal{O}(10^{-5})$	$\ll 1$
$hh + \text{jets}, \text{ SM}$	275	593	7×10^{-4}	12.4

$\Rightarrow \lambda_4 \text{ in } [-4, 16] \text{ at } 95\% \text{CL}$

Conclusions

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- It is clear from the previous slides that this is however a ballgame different from the LHC.
- There are many opportunities, exploiting the large statistics and the novel kinematical regimes
- Once a more complete assessment has been made, we can move to address questions like:

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