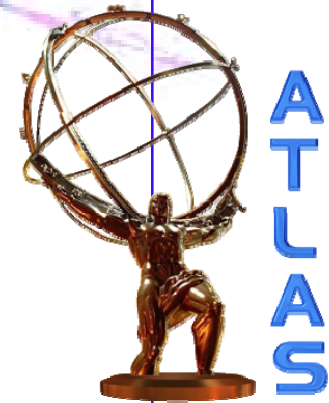


# The reach of the ATLAS experiment in the SUSY parameter space

International Conference on Particle Physics  
-in memoriam Engin Arik and her colleagues -  
Istanbul, 28<sup>th</sup> of October 2008



**Janet Dietrich**  
University of Freiburg  
on behalf of the ATLAS Collaboration



# Supersymmetry at LHC

## Susy is one of the main topics at LHC:

- extension of SM: adds a **boson**  $\leftrightarrow$  **fermion** symmetry:
  - solves the **hierarchy** problem of the SM
  - can unify gauge coupling at GUT scale
- can provide **dark matter** candidate:
  - existence of a **lightest stable SUSY particle (LSP)**

## This talk covers only direct **SUSY searches** (results for $1\text{fb}^{-1}$ ):

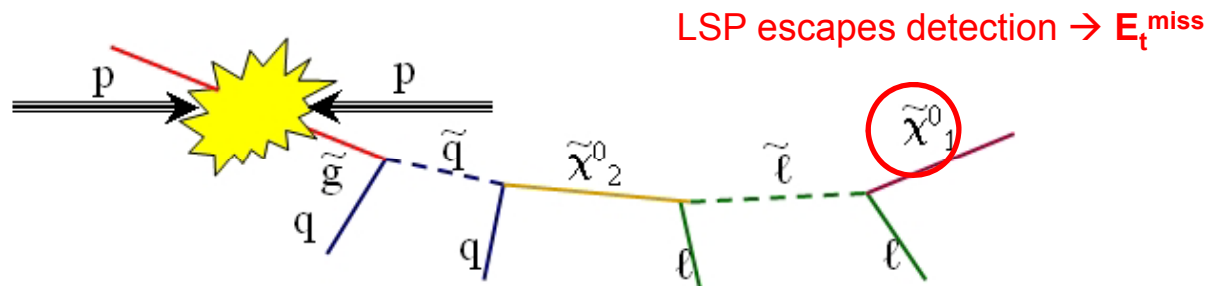
- R-parity conserving models
  - pair production of SUSY particles
  - cascade decay to lighter SUSY particles and finally to the LSP
  - LSP is stable  $\rightarrow$  escapes detection:  $E_{\text{t}}^{\text{miss}}$

**Inclusive searches for new physics at TeV scale:**

**Look for deviations between 'data' and SM expectations  
in tails of  $M_{\text{eff}}$  distributions!**

# Characteristic SUSY signatures

example for SUSY signature :



- strongly interacting SUSY particles (squarks and gluinos) dominate SUSY production
- $\rightarrow$  many **high  $p_T$  objects** (model dependent): leptons, jets, b-jets, photons
- LSP is stable:
- $\rightarrow$  clear experimental signatures with **large  $E_t^{\text{miss}}$**

**generic signature =  $E_T^{\text{miss}}$  + multi-jets (+ multi-leptons + photons)**

- $\rightarrow$  build various channels depending on number/type of objects (inclusive channels)

**signals studied at ATLAS (CSC note 2008):**

**2, 3, 4 jets + 0, 1, 2, 3 leptons (e, $\mu$ ) , tau or photon +  $E_t^{\text{miss}}$**

# SUSY searches

## SUSY signatures are quite model dependent:

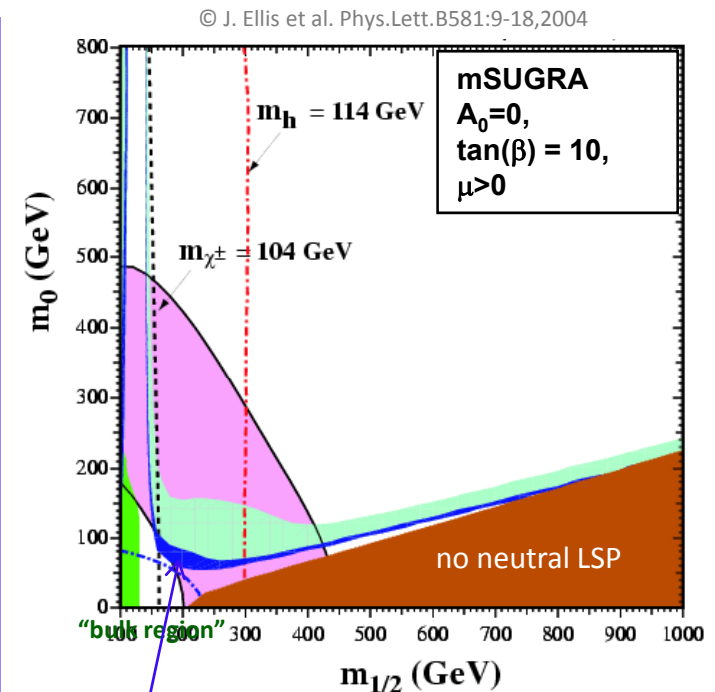
mass scale of SUSY particles, SUSY breaking scenario,.. are unknown

general “strategy”:

- define exclusive channels
- study selection cuts on benchmark points

e.g.: SU3:  $m_0=100$ ,  $m_{1/2}=300$ ,  $\tan\beta=6$ ,  $A_0=-300$

- test procedure with different scenarios
  - take **mSUGRA model** (5 parameter) as “baseline” paradigm for R-parity conserving SUSY model :
  - use also other scenarios:  
**NUHM, GMSB, AMSB**



for a particular model the available parameter space is already limited from experimental results (Dark matter constraints)

# Inclusive Searches

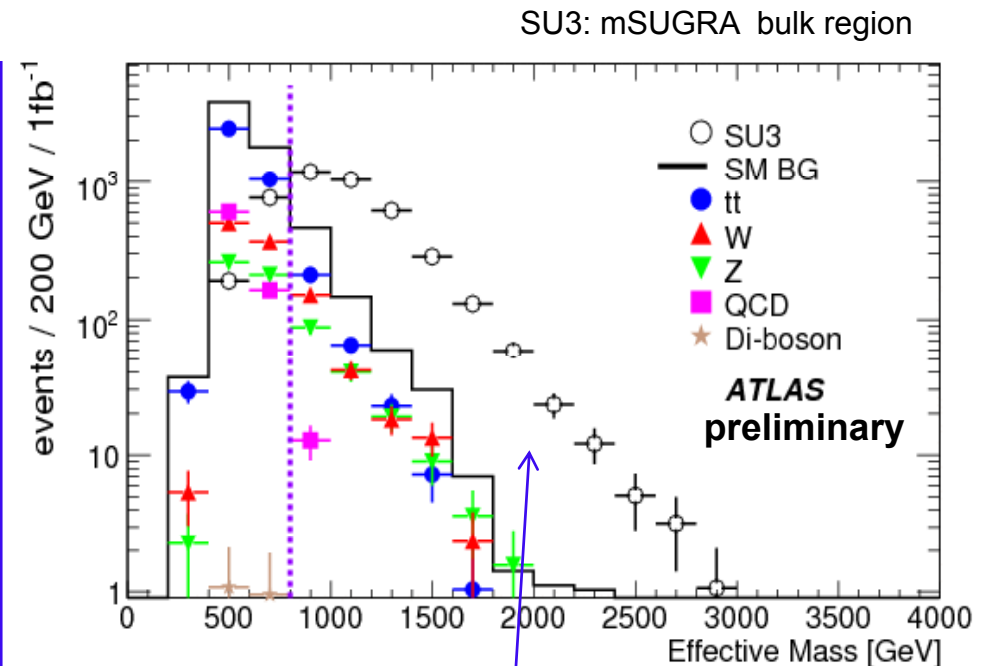
## Example: 4jet + 0 lepton channel

sensitive variable to detect SUSY: “effective Mass”  $M_{\text{eff}} = \sum |p_{\text{T}}^i| + E_{\text{T}}^{\text{miss}}$

### 0-lepton channel

#### event selection:

- Trigger: combined trigger  
jet -  $p_{\text{T}} > 70 \text{ GeV}/c + E_{\text{T}}^{\text{miss}} > 70 \text{ GeV}/c$
- 4 jets with  $p_{\text{T}} > 100, 50, 50, 50 \text{ GeV}/c$
- $E_{\text{T}}^{\text{miss}} > 100 \text{ GeV}/c$  and  $> 0.2 M_{\text{eff}}$
- transverse Sphericity  $> 0.2$
- veto isolated leptons
- $\min \Delta\phi(\text{jet}^i, E_{\text{T}}^{\text{miss}}) > 0.2$   
(QCD removal)



excess at large  $M_{\text{eff}}$

$M_{\text{eff}} > 800 \text{ GeV}$ :

$S/B \approx 6430 / 1210 \text{ for } L=1\text{fb}^{-1}$

# Inclusive Searches

## Example: 4jet + 0 lepton channel

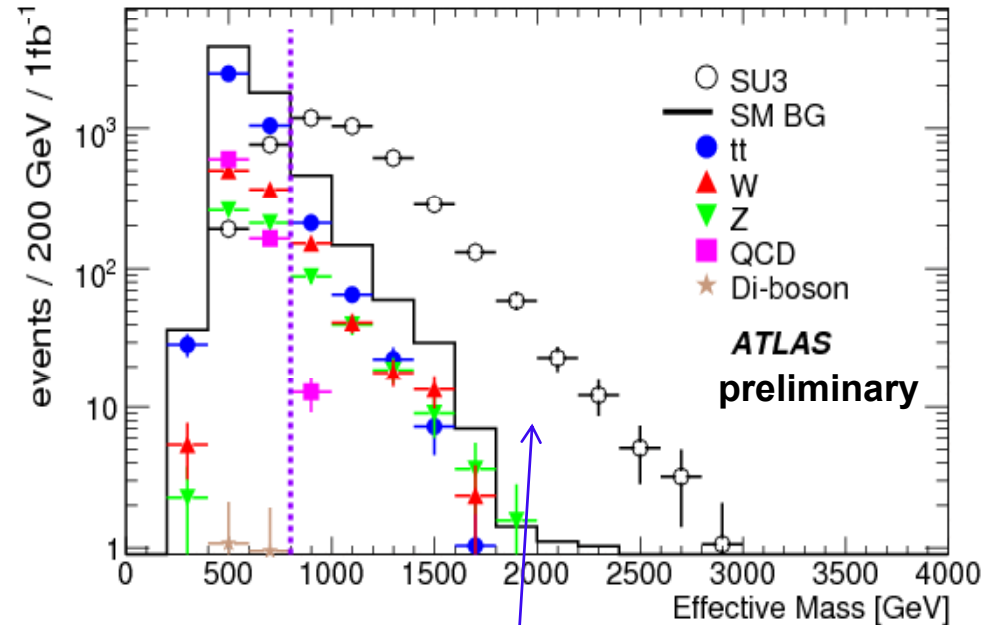
### 0-lepton channel

#### Main SM backgrounds :

- t-t bar, W(lv)+jets, Z(vv)+jets, QCD jets
- many (difficult) background contributions to control :
  - large theoretical uncertainties on backgrounds (multi-jet production)
  - try to rely on data-driven determination of backgrounds :

*additional problem:*

poor understanding of detector ( $E_t^{\text{miss}}$ , JES,...) with early data



excess at large  $M_{\text{eff}}$

$M_{\text{eff}} > 800 \text{ GeV}$ :

**S/B  $\approx 6430 / 1210$  for  $L=1\text{fb}^{-1}$**

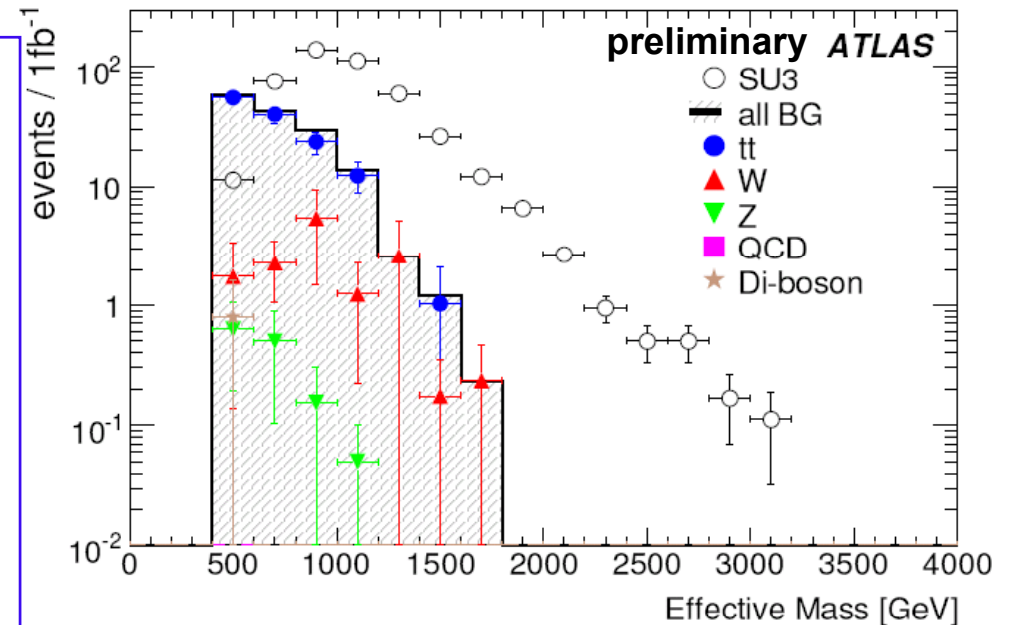
# Inclusive Searches

## 1-lepton channel

### event selection:

- Trigger: combined trigger  
jet -  $p_T > 70 \text{ GeV}/c + E_t^{\text{miss}} > 70 \text{ GeV}/c$
- 4 jets with  $p_T > 100, 50, 50, 50 \text{ GeV}/c$
- $E_T^{\text{miss}} > 100 \text{ GeV}/c$  and  $> 0.2 M_{\text{eff}}$
- transverse Sphericity  $> 0.2$
- **exactly 1 isolated lepton (e or  $\mu$ )  
with  $p_T > 20 \text{ GeV}/c$**
- **no other e or  $\mu$  ( $p_T > 10 \text{ GeV}/c$ )**
- **transverse Mass:  
 $M_T(e/\mu, E_T^{\text{miss}}) > 100 \text{ GeV}/c$**

$$M_{\text{eff}} = \Sigma |p_T^i| + E_t^{\text{miss}} + |p_T(\text{lepton})|$$



### Main SM background:

- top pairs,  $W(l\nu)$ +jets
- lower, but more clear signal
- less QCD background

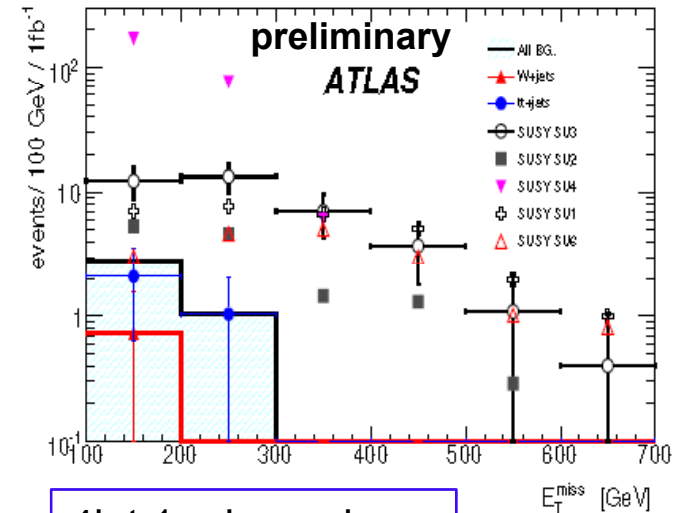
$$S/B \approx 230 / 40 \text{ for } L=1\text{fb}^{-1}$$

# Inclusive Searches

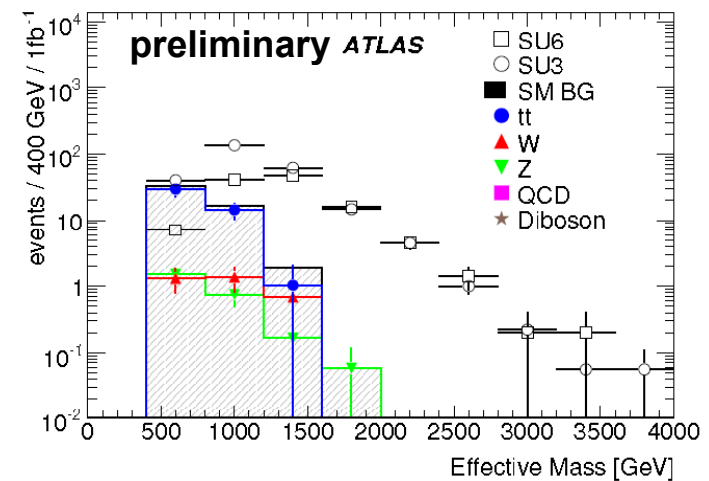
## Other inclusive Search channels considered:

- relax N jets from 4 to 3 or 2 ...
- 2 lepton same sign, opposite sign
- 3 leptons + jet (no  $E_T^{\text{miss}}$  !)
- 3 leptons +  $E_T^{\text{miss}}$  (no jets !)
- tau mode:
  - e,  $\mu$  replaced by hadronically decaying  $\tau$
  - can dominate for high  $\tan\beta$
- b-jet mode: exploit SUSY's richness in b-jets

## 4jet same sign 2 lepton channel



## 4jet 1 $\tau$ channel

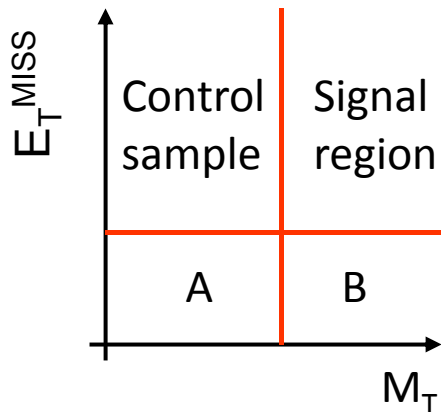




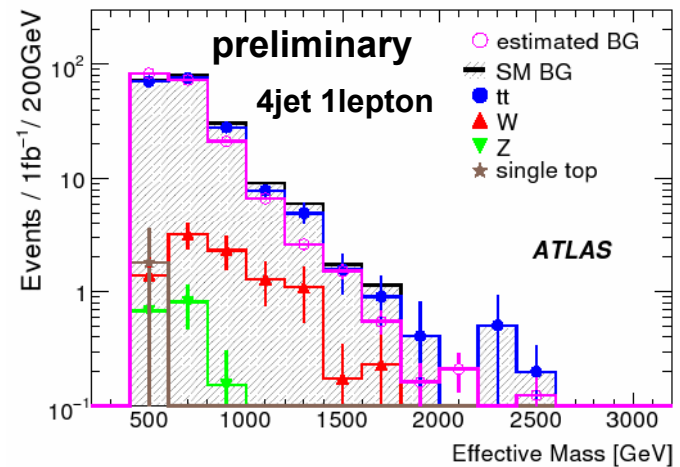
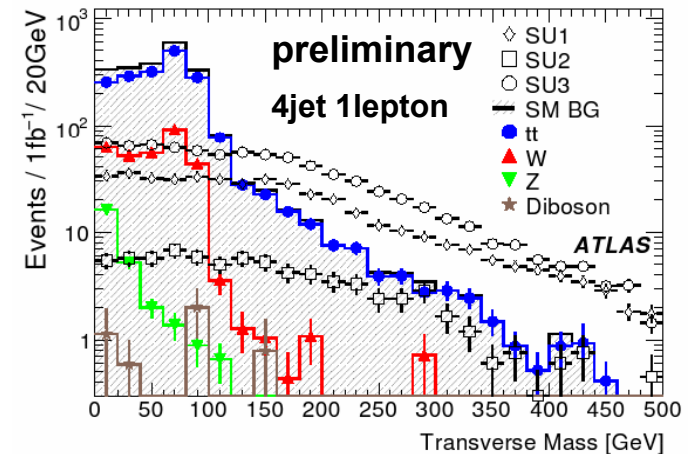
# Background estimation

before claiming any discovery we need to understand our background expectations (MC, detector)  
 → validate background estimation with data in defined control regions

estimate top / W background from data



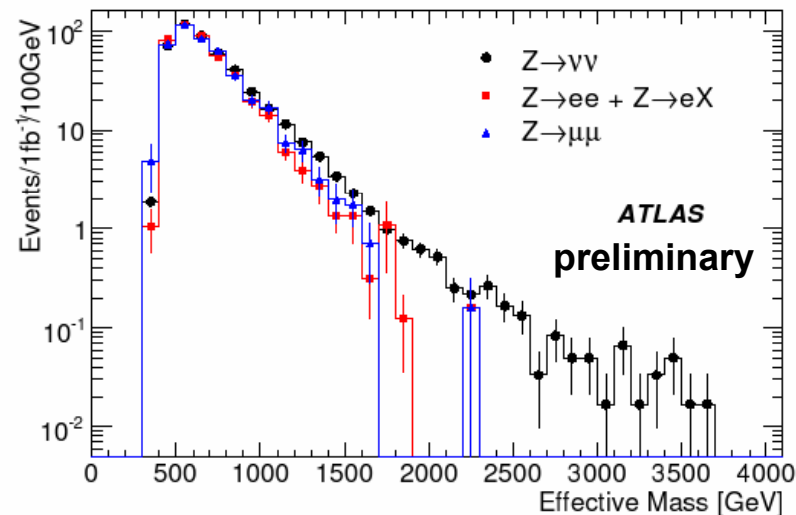
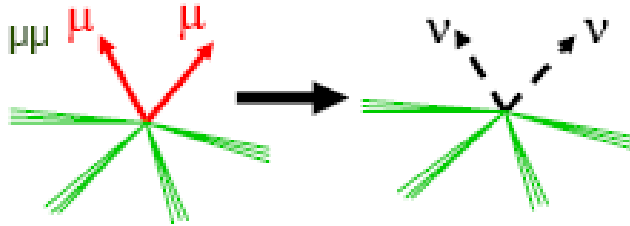
- Top background can be estimated with many methods
- use discriminating variable  $M_T < 100 \text{ GeV}/c$  to select control sample region
- $M_{\text{eff}}$  distribution in control region can be used to predict distribution in signal region ( $M_T > 100 \text{ GeV}$ )
- **uncertainty for Top:**  
about 30 % (20 % for 1 lepton mode) for  $1 \text{ fb}^{-1}$



# Background estimation

## estimate Z background from data

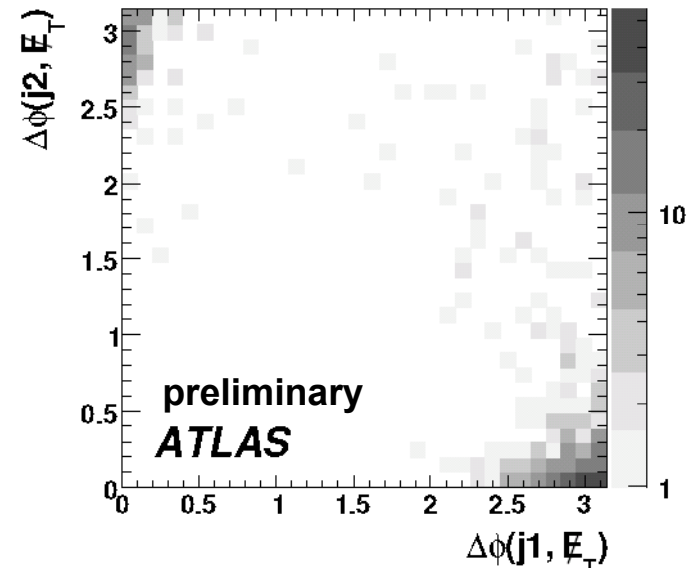
- $Z \rightarrow \nu\nu + \text{jets}$  is an irreducible background in 0-lepton channel
- estimate background from data with  $Z \rightarrow l^+ l^- + \text{jets}$



uncertainty for Z : about 20% (15 % for 0lepton mode) for 1fb<sup>-1</sup>

## estimate QCD background from data

- fake missing  $p_T$  due to jet mis-measurements and jet resolutions
- reducible with clean up cuts  $\rightarrow \Delta\phi(\text{jet} - E_T^{\text{MISS}})$  cut
- to estimate remaining QCD background after cleaning :
  - measure smearing function in events with large MET associated to a single jet
  - select seed events with low MET-significance and smear each jet

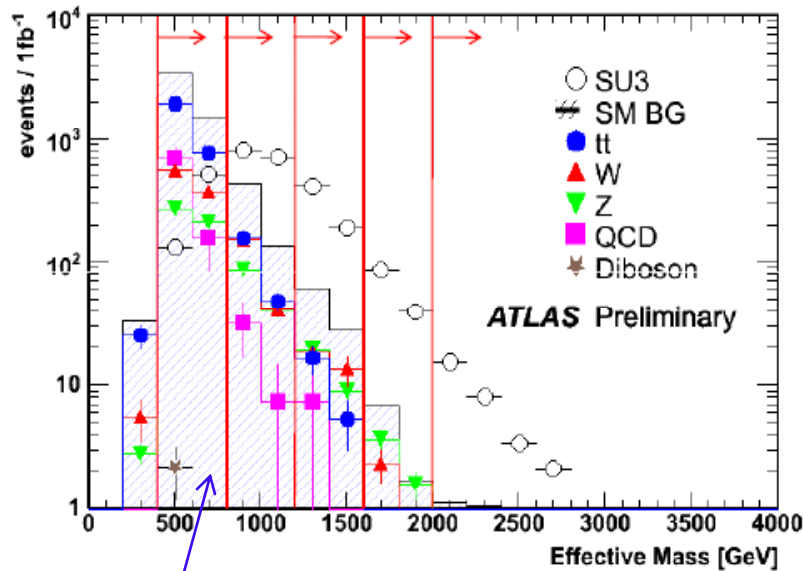


uncertainty for QCD: about 50% for 1fb<sup>-1</sup>

# Discovery reach - procedure

## Scan in mSUGRA parameter space:

repeat searches using many different mSUGRA signals



Meff scan step size: 400 GeV/c

\*prob: equation from Highland/Cousins  
= convolution of the poisson probability with a  
Gaussian probability density function

### Aim:

Scan  $M_{\text{eff}}$  distribution for deviation

### Procedure:

- each point has a different  $M_{\text{eff}}$  distribution
- scan each distribution (try n cuts in  $M_{\text{eff}}$ ):  
optimize  $M_{\text{eff}}$  cut to get best signal significance
- looking for  $M_{\text{eff}}$  cut with highest significance  
(lowest probability):

1. calculate probability prob\*:

to find D data events if B SM events are expected

2. choose region with lowest probability prob\*

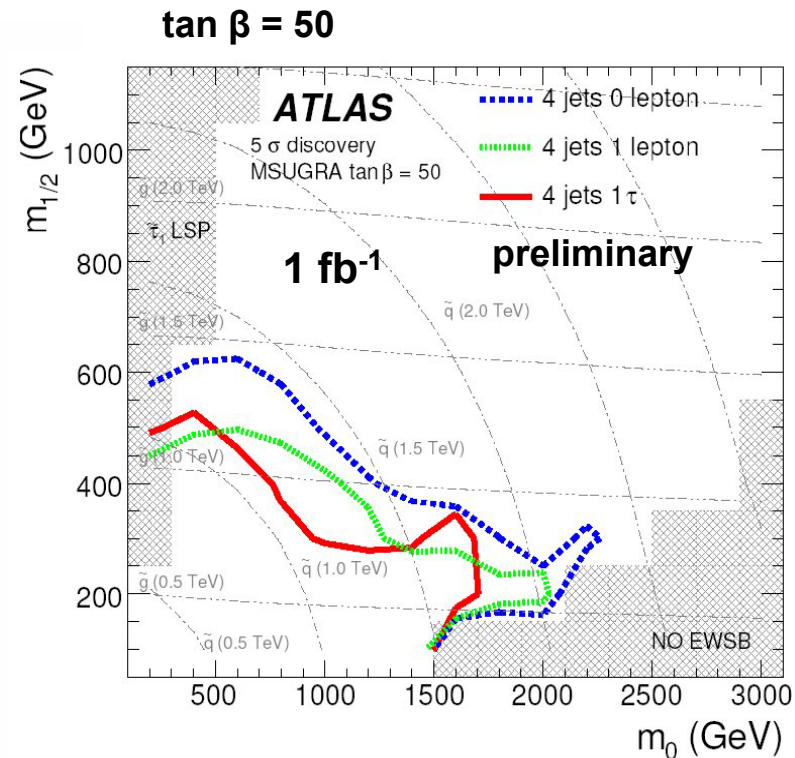
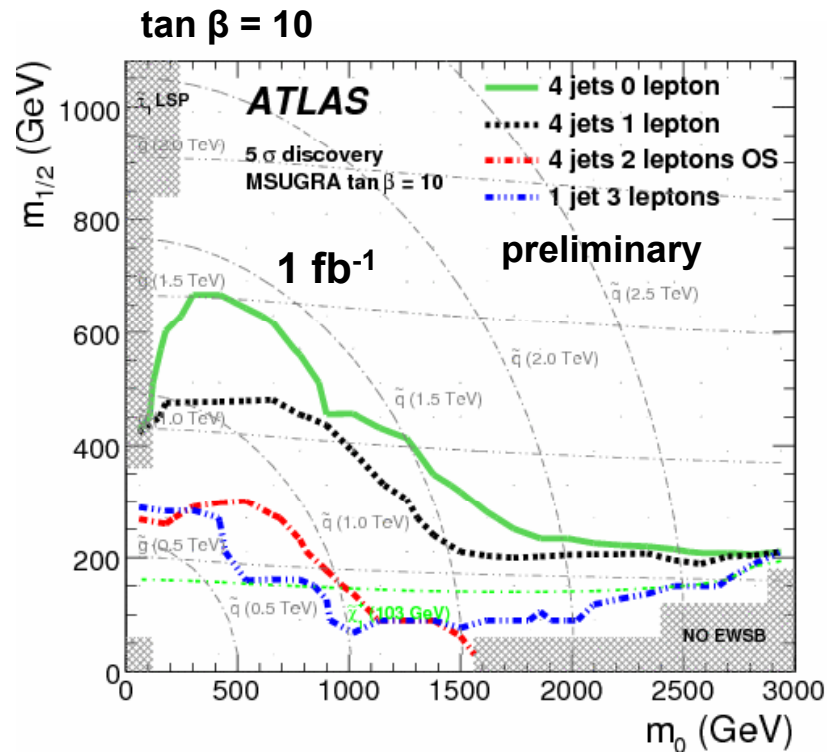
3. correct probability values for 'multiple testing'  
via a Monte Carlo method

$$\text{significance} = \text{erf}^{-1}(1 - 2 \cdot \text{prob}) \cdot \sqrt{2}$$

# Discovery reach for mSUGRA

## Scan in mSUGRA parameter space:

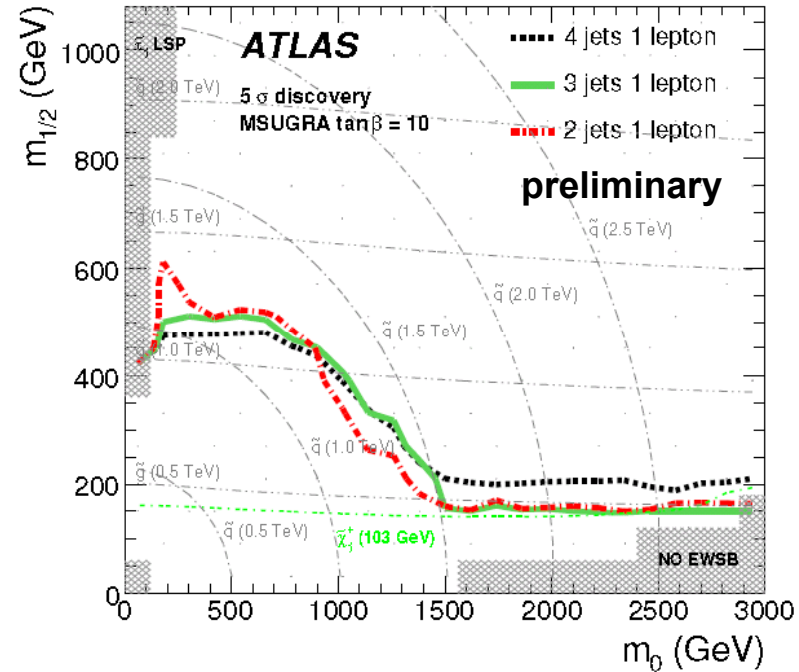
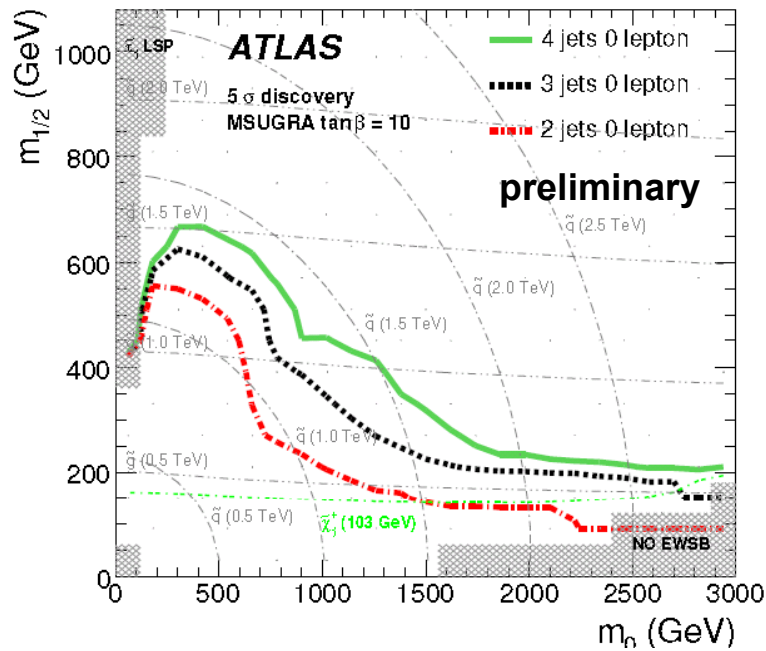
- LO signal cross sections were used
- significant (data-driven) uncertainties on background (QCD : 50%, other background 20 %)



- for 1 fb<sup>-1</sup> : cover a large part of mSUGRA phase space favored by electroweak, heavy-flavour physics and low energy precision data!
- tau channel can help for large tan $\beta$

# Discovery reach for mSUGRA

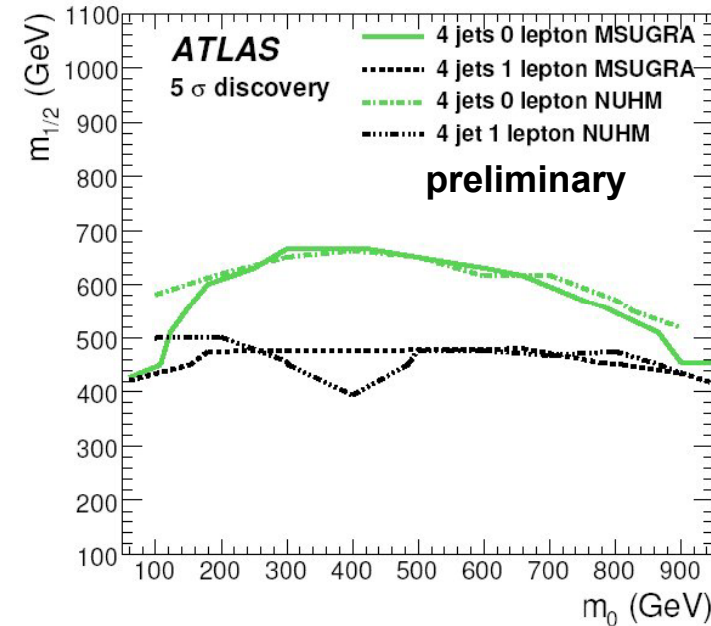
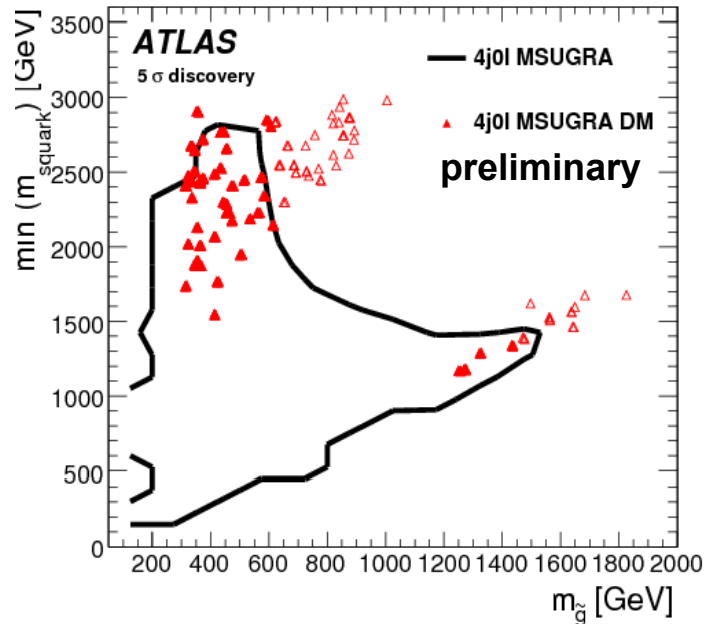
## Comparison of 2,3,4 jets + 0,1 lepton



sometimes better discovery potential with lower jet multiplicities!

# Discovery reach for other models

discovery reach for other models ( $1 \text{ fb}^{-1}$ ):



## mSUGRA $\leftrightarrow$ DM constraints ( $\mu > 0$ )

random mSUGRA points, but satisfying WMAP (dark matter constraints)

$g_\mu - 2$ ,  $M_h$ ,  $b \rightarrow s\gamma$ ,  $B_s \rightarrow \mu\mu$

(closed triangle :  $> 5\sigma$  discovery)

## mSUGRA $\leftrightarrow$ NUHM

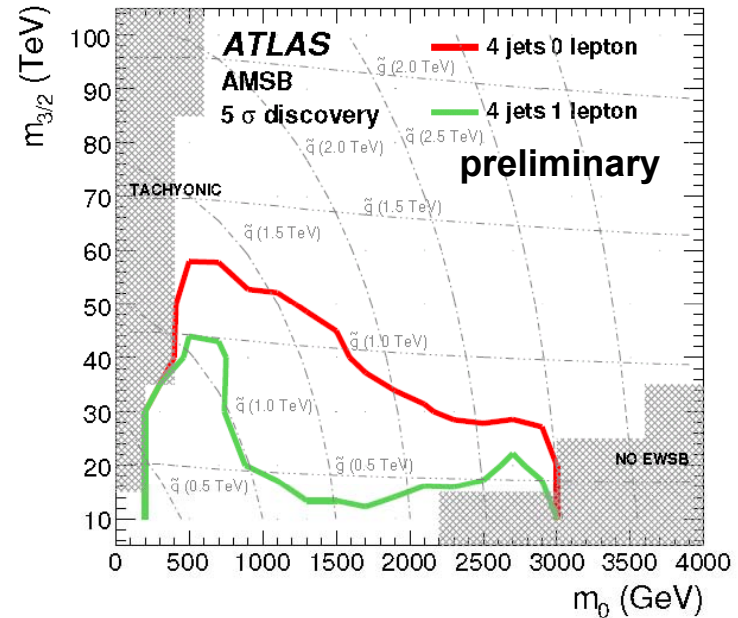
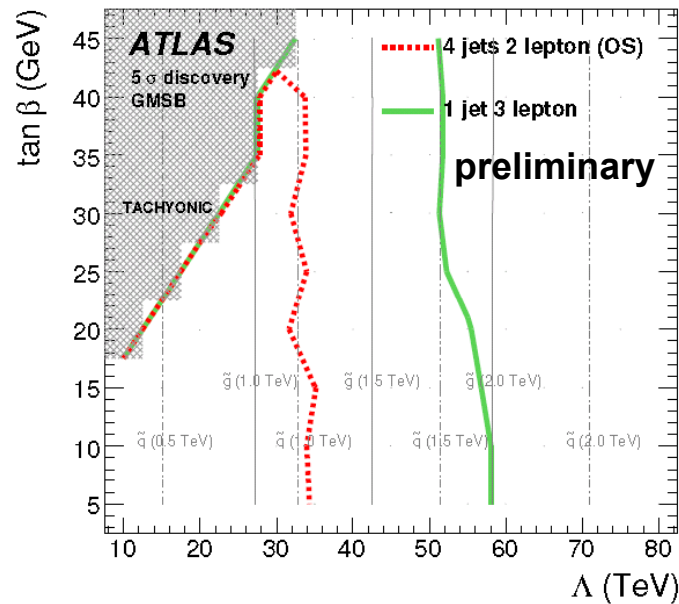
NUHM :

fix  $M_A$  and  $\tan \beta$  to values compatible with WMAP constraints

similar results: discovery reach does not differ so much

# Discovery reach for other models

discovery reach for other models (1 fb<sup>-1</sup>):



## GMSB :

- production: 2 leptons or  $\tau$  on generator level
- 1 jet 3 lepton analysis shows highest discovery reach : less Standard Model background

## AMSB:

- similar phenomenology as NUHM, but different masses and decay modes
- reach for 0 lepton channel similar to the reach for mSUGRA
- difference for 1 lepton channel: chargino decays in invisible leptons

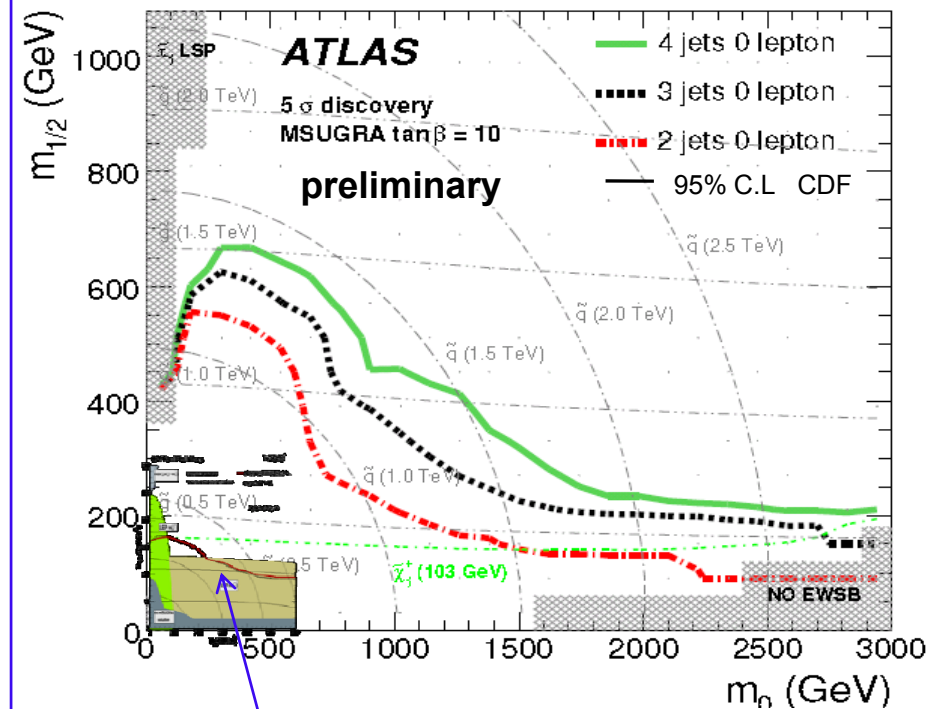
# Comparison with Tevatron

for very early data taking:

discovery reach for  $L = 100 \text{ pb}^{-1}$   
 higher than CDF 95% confidence  
 level for  $L = 2 \text{ fb}^{-1}$  !

- ATLAS is sensitive mainly to a region of squark and gluino masses up to 600 GeV for 10 TeV with  $20 \text{ pb}^{-1}$
- D0/CDF 2- $\sigma$  limits are at  $m_{\text{squark}} / m_{\text{gluino}}$  of 350-400 GeV now

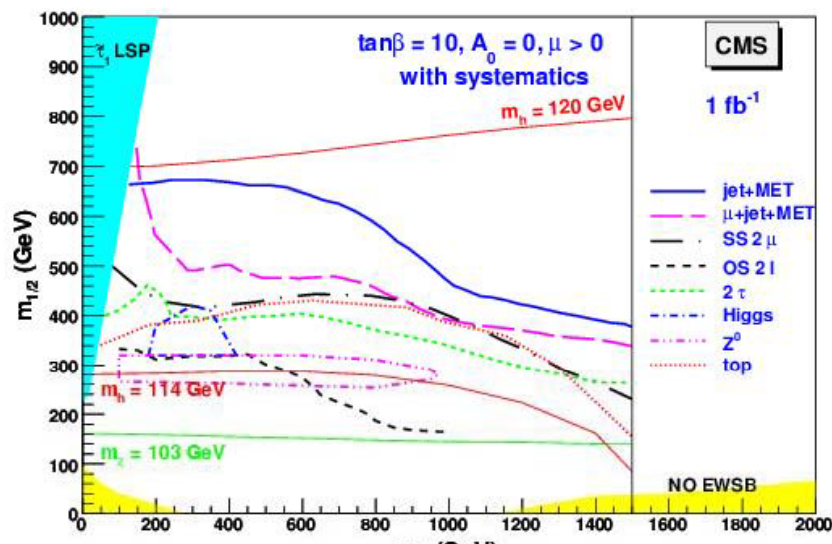
plots for 14 TeV ( $\cong 1 \text{ fb}^{-1}$ )



95 % confidence level : D0/CDF

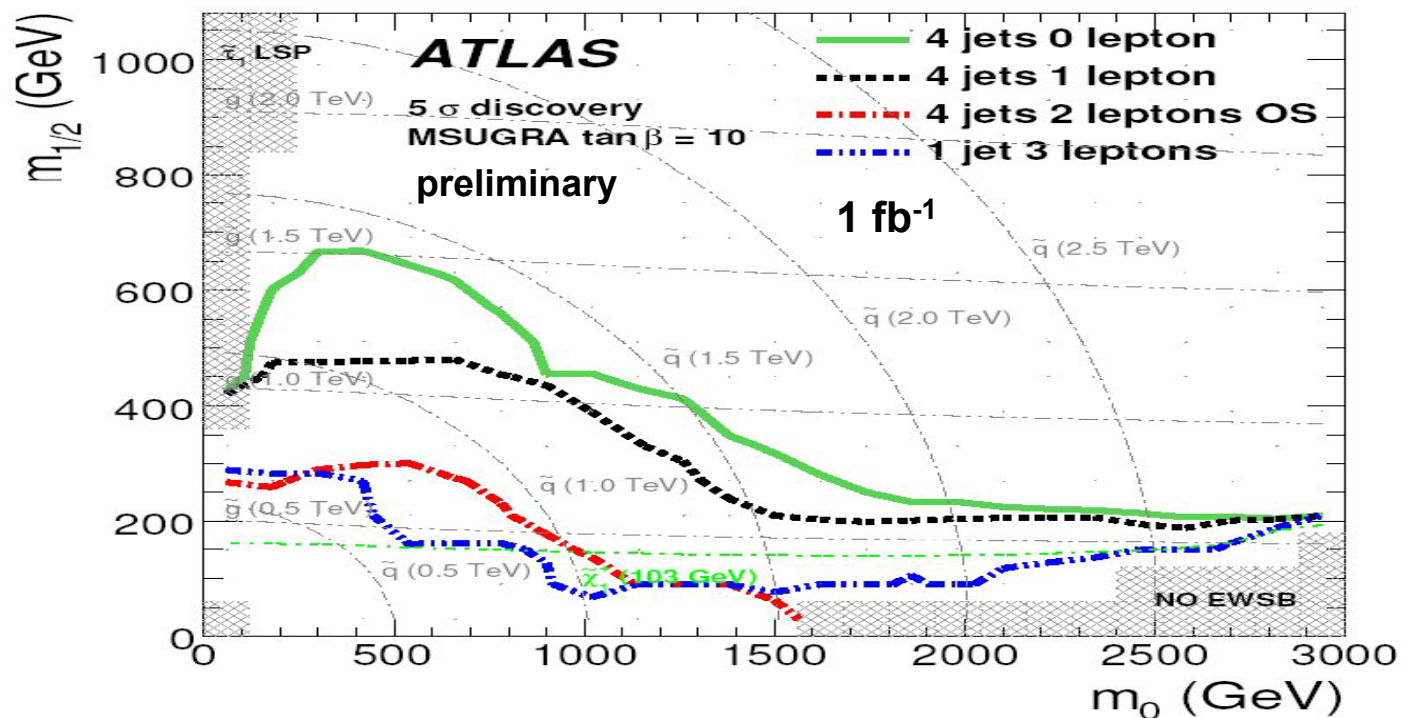


# Comparison ATLAS $\leftrightarrow$ CMS



The potentials of CMS and ATLAS are very similar.

Largely determined by signal and background cross sections!



# Conclusion

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- recent studies for various SUSY channels and various background estimation strategies have been presented
- ATLAS shows a high potential to discover R-parity conserving SUSY signatures with gluino and squark masses up to 1 TeV for 1 fb<sup>-1</sup>...

....**but** :

we have to estimate all backgrounds and understand our detectors, triggers, etc ... first!

What we need NOW is real data! (hopefully spring 2009 )

**We are at the threshold of very exciting times!**

# THANKS!

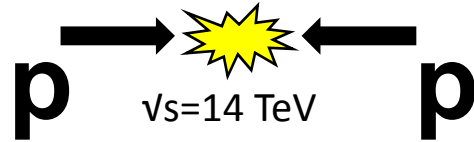
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**Many thanks to the organizers!**

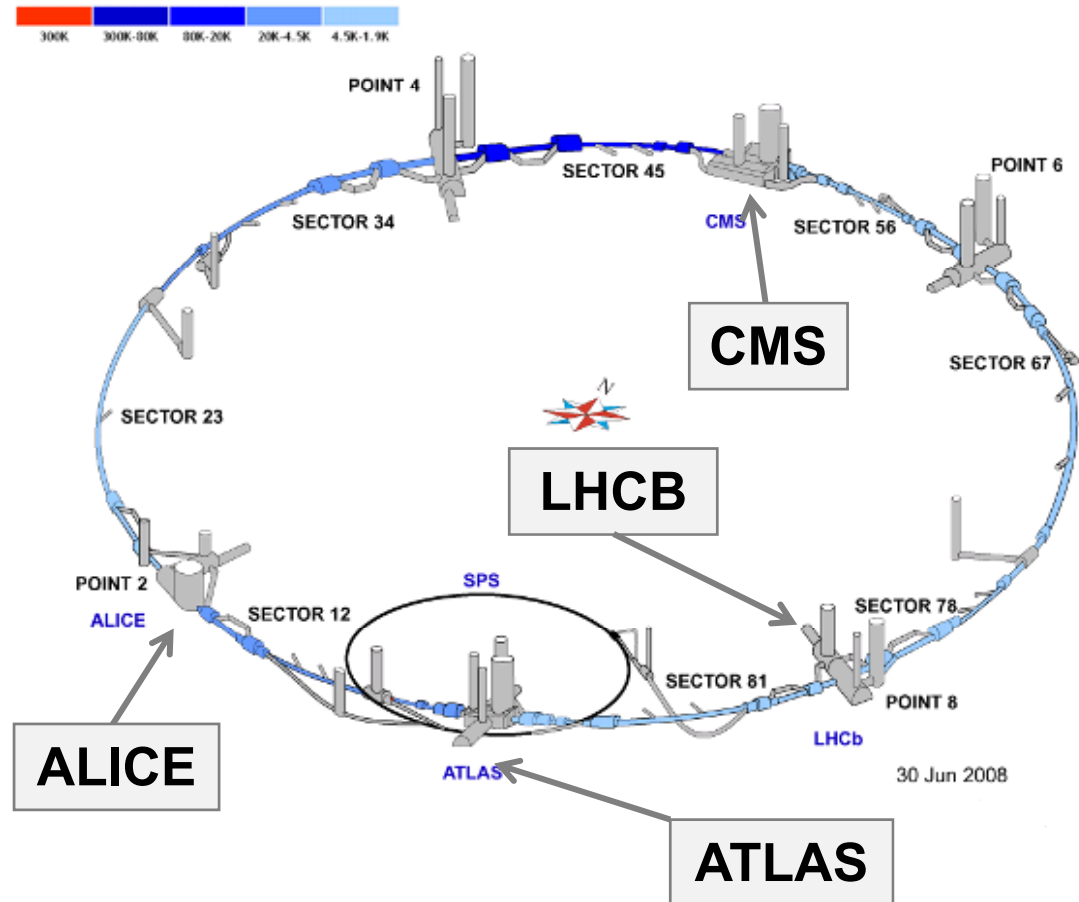
# BackUps

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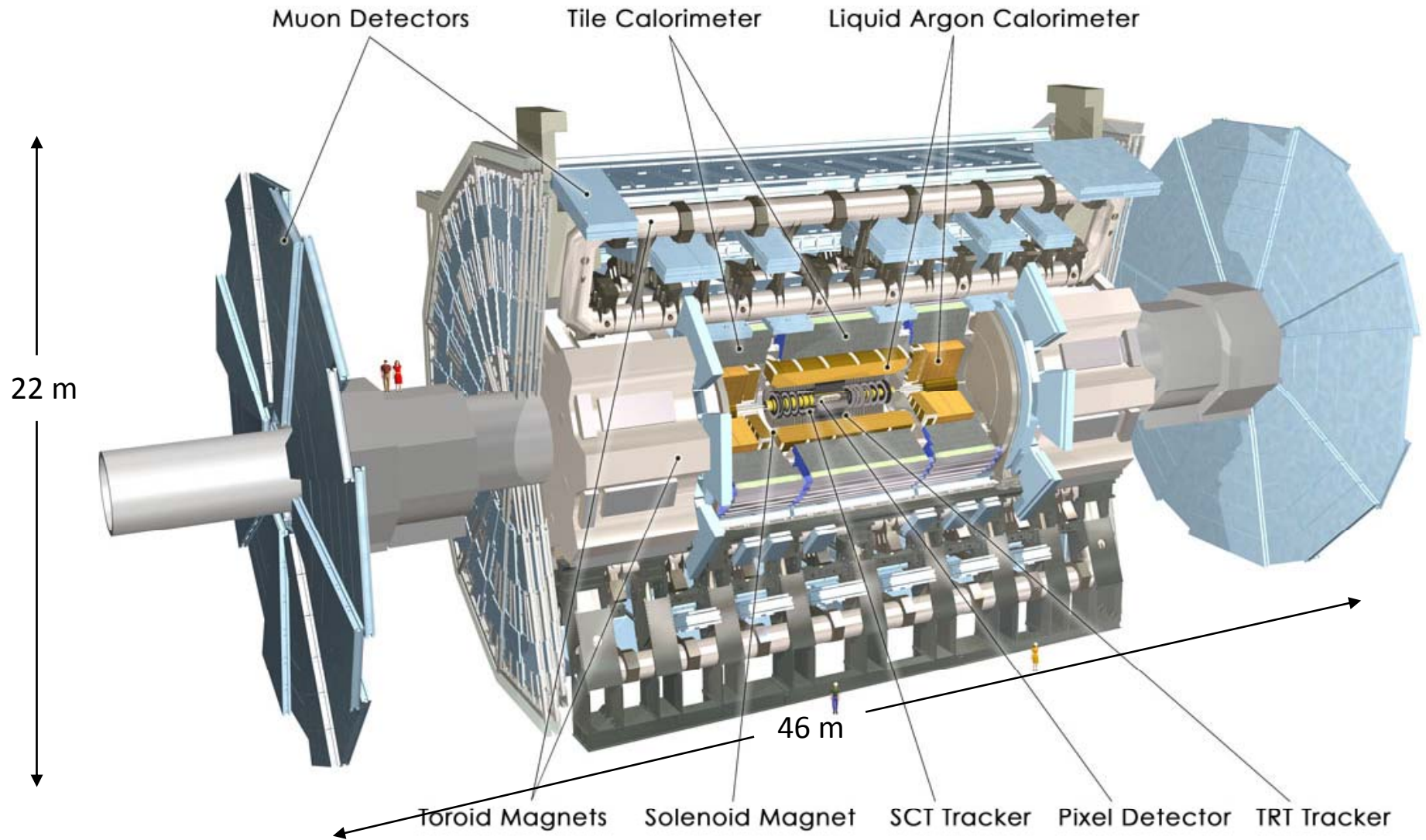
# Large Hadron Collider @ CERN



- LHC collides protons –protons (or Pb ions)
- $\sqrt{s}=14 \text{ TeV}$  (10 TeV early next year?)
- Design Luminosity  $10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
- Bunch crossing rate 40 MHz
- Events per bunch crossing  $\sim 20$  (“pile-up”)
- 4 main experiments
  - CMS, ATLAS (multi purpose)
  - ALICE (heavy ion physics)
  - LHCb (low- $p_T$  B physics)



# The ATLAS experiment, 1900 scientists, 165 institutes, 35 countries



# Supersymmetry MSUGRA - $\tan(\beta) = 10$

© J. Ellis et al. Phys.Lett.B581:9-18,2004

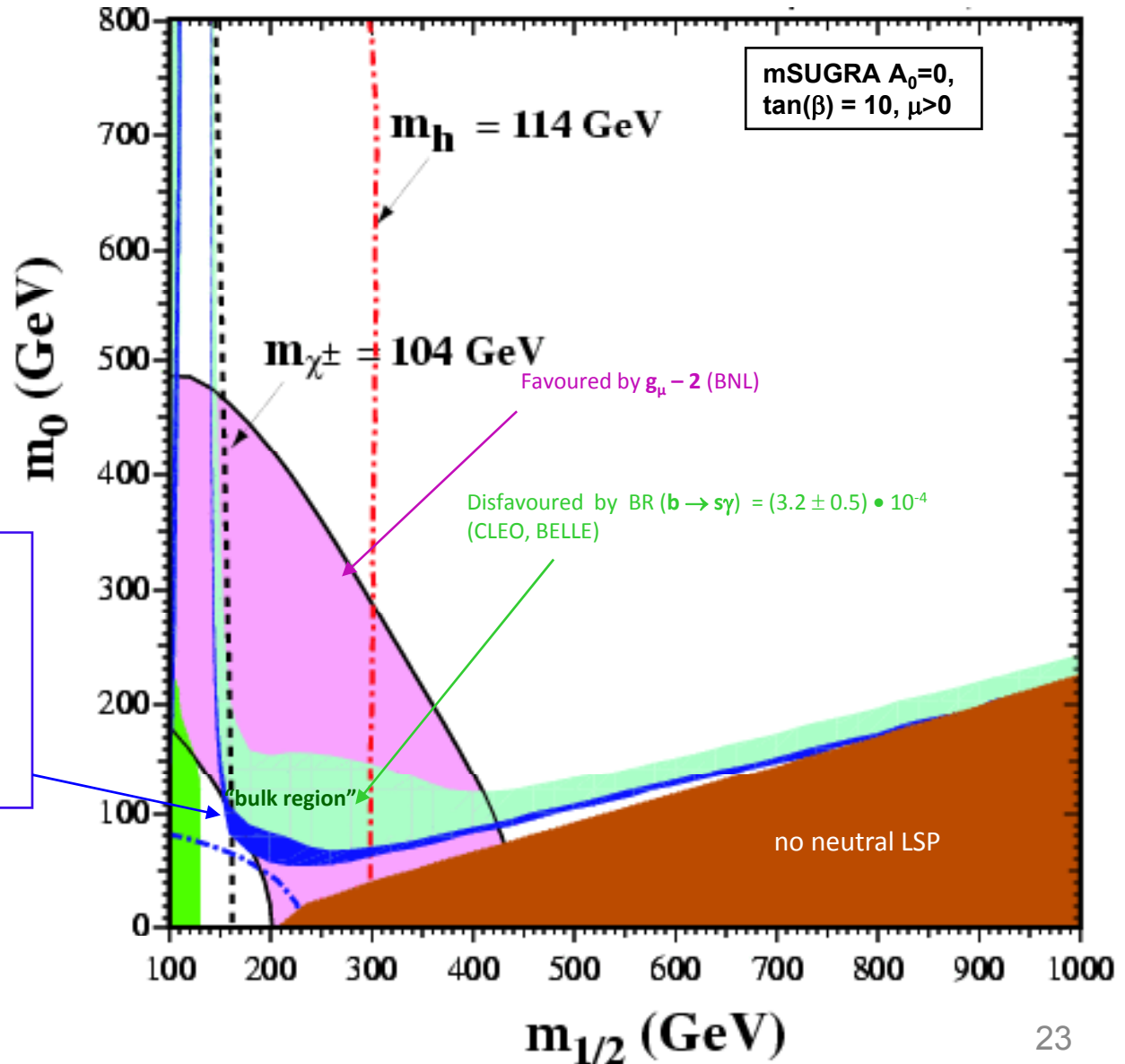
## mSUGRA model:

- minimal gravity mediated SUSY breaking
- 5 parameters ( $m_0, m_{1/2}, \dots$ )
- LSP = lightest neutralino

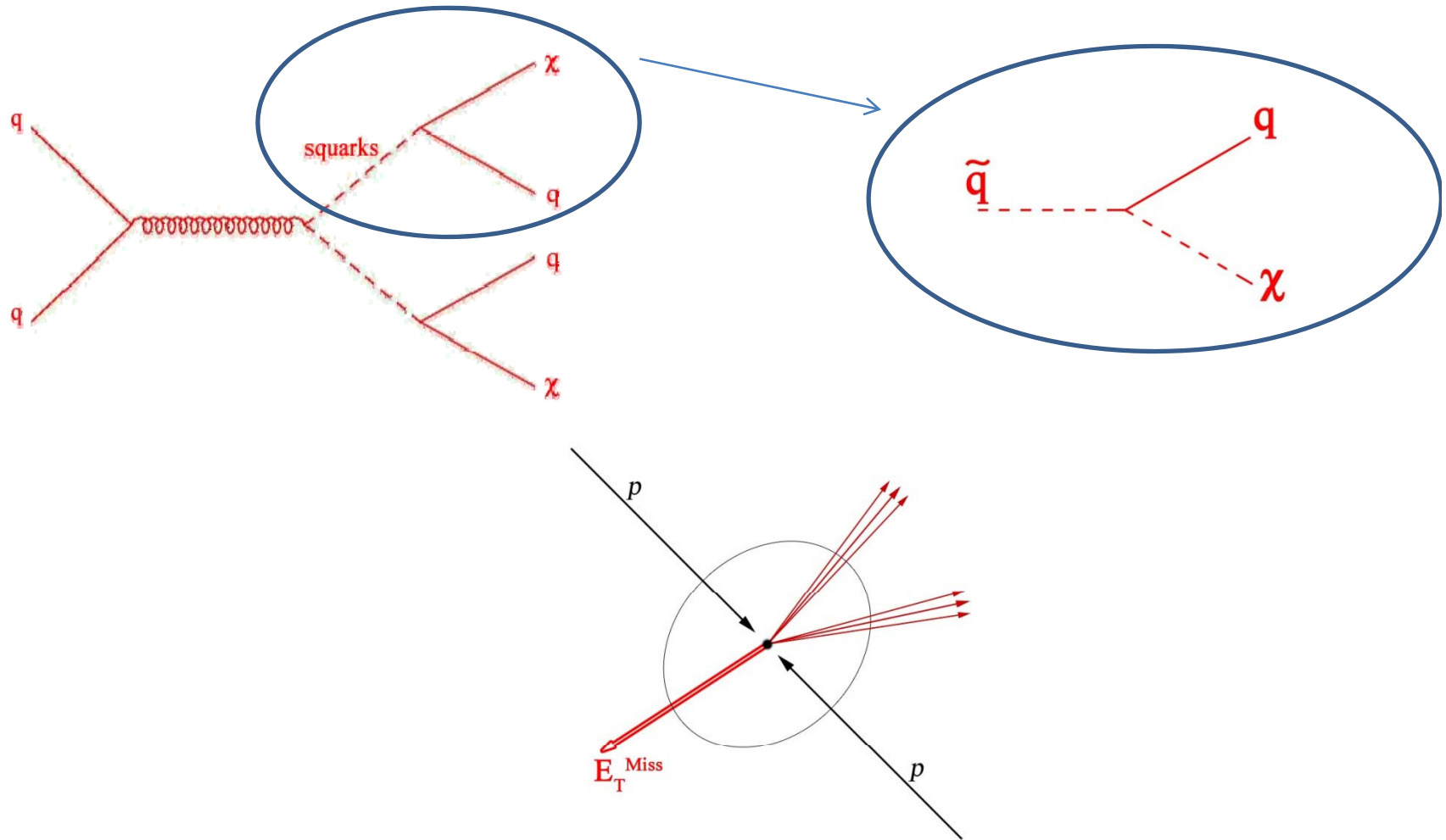
Large LSP annihilation cross-section required by WMAP data:

CMB measurement:

$0.094 < \Omega_{\text{DM}} h^2 < 0.129$  strongly bounds SUSY parameter space



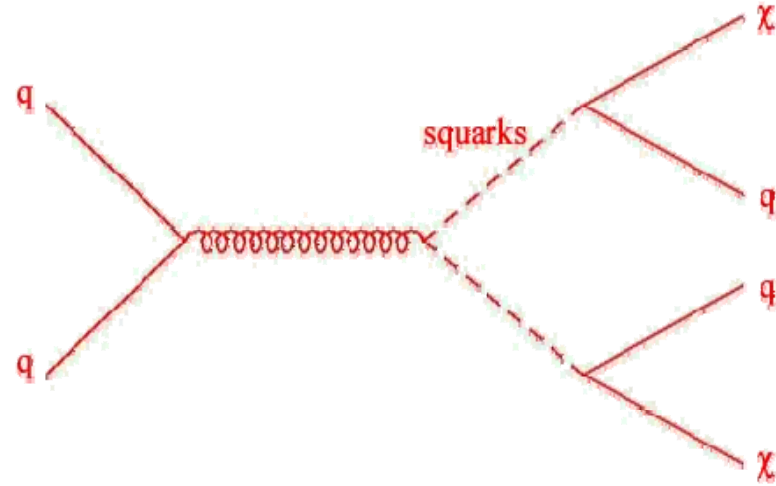
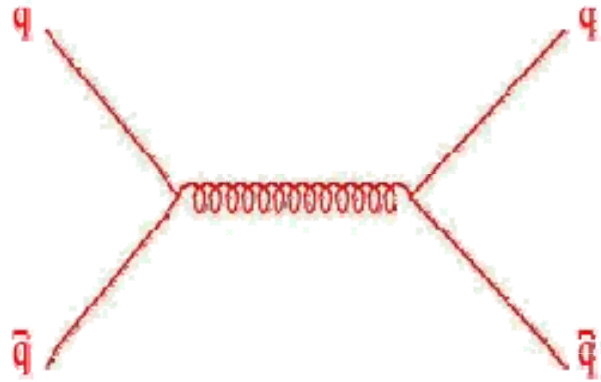
# SUSY signatures - an example





# QCD background $\leftrightarrow$ SUSY signature

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# SUSY searches

- ‘baseline model’ paradigm: **RP-conserving mSUGRA**
- gluino/squark production and cascade decay
- energetic jets, missing  $E_T$ , (leptons)

- also using :

**NUHM:** non-universal Higgs masses

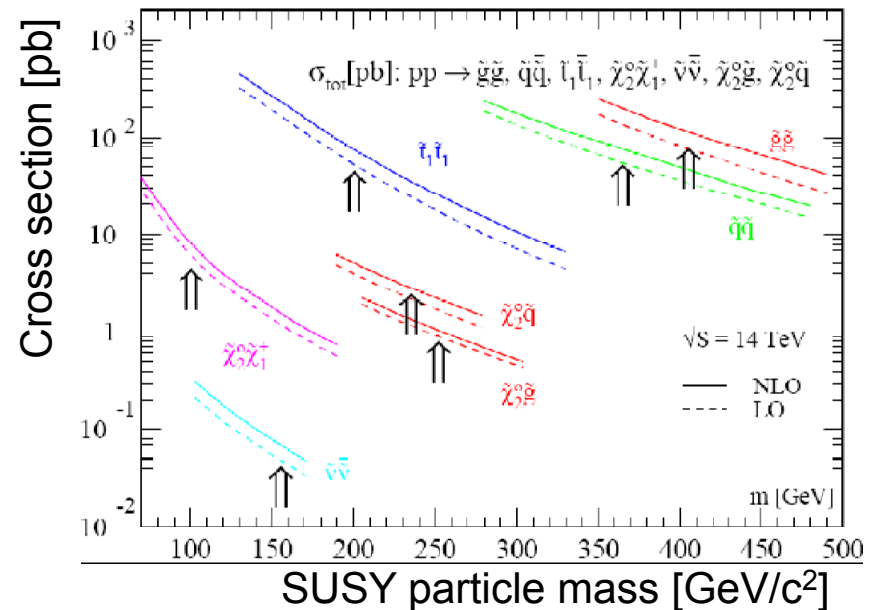
**GMSB:** gravitino LSP

NLSP: neutralino  $\rightarrow$  gravitino +  $\gamma$

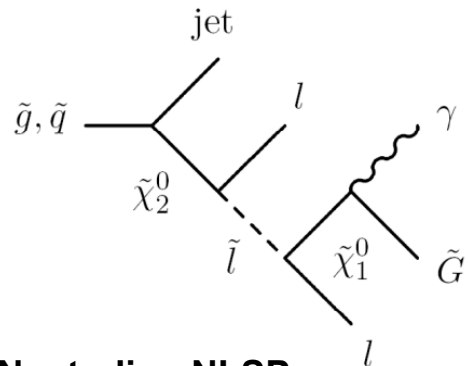
NLSP: slepton (stau)  $\rightarrow$  gravitino + lepton  
(tau)

**AMSB**

- cross-sections mainly depend on SUSY masses (masses are model dependent)

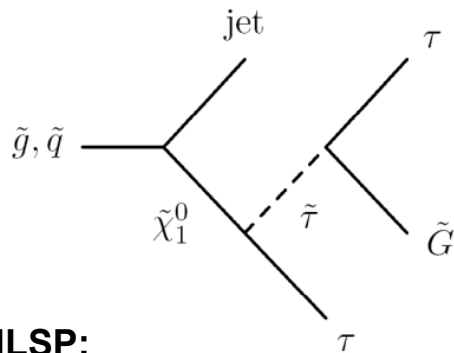


# GMSB Model



## Neutralino NLSP:

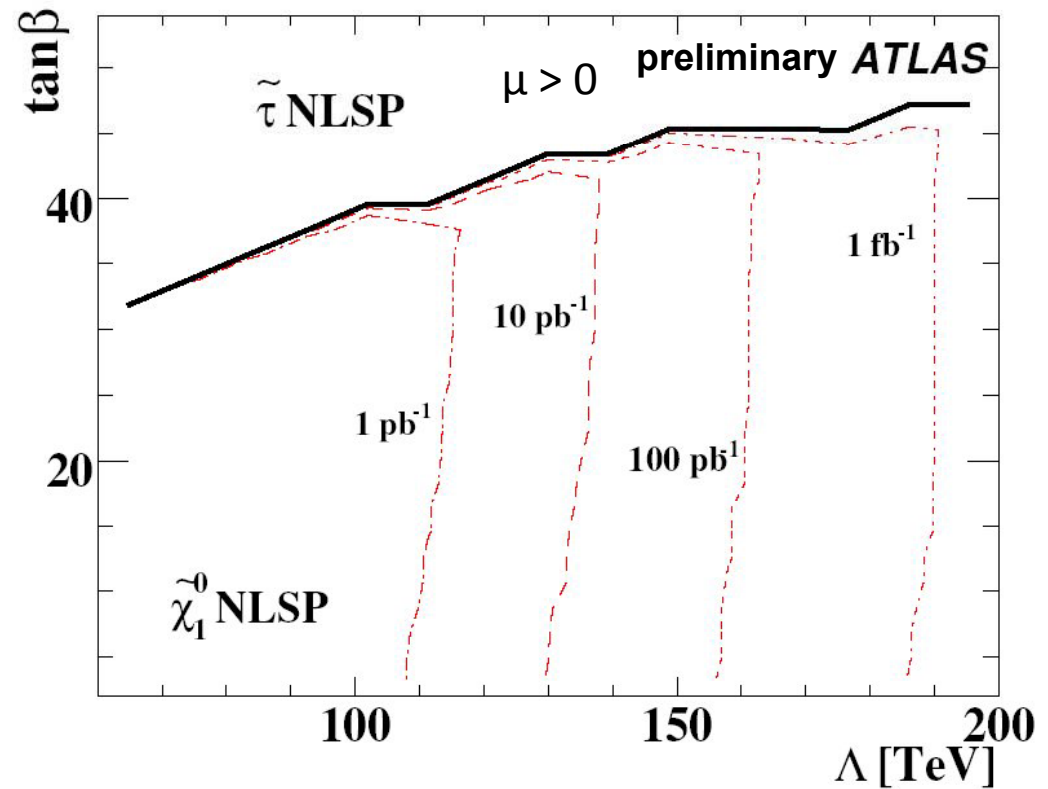
prompt photons  
non-pointing photons



## Slepton NLSP:

prompt leptons  
long living sleptons:  
semi-stable massive particle.

reach in  $\Lambda$  and  $\tan \beta$  ( $N_5 = 1$ ,  $M_m = 500$  TeV)



# Results of Optimization

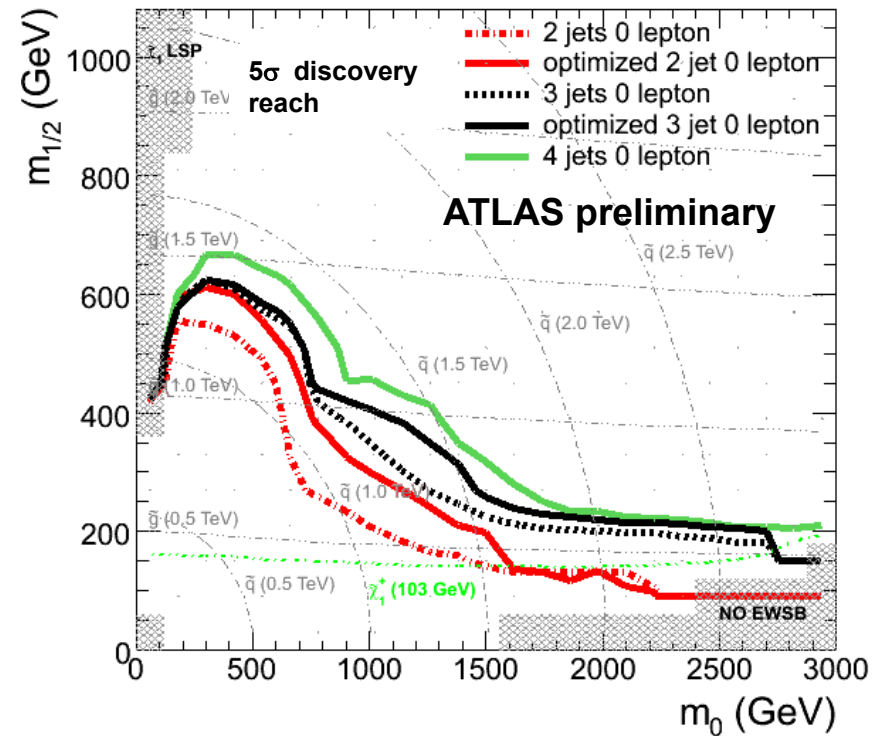
5  $\sigma$  contour line plots with optimum selection for 1 fb<sup>-1</sup>

systematic errors:

QCD = 50 %

other bg = 20 %

The cut optimization could significantly improve the discovery reach !



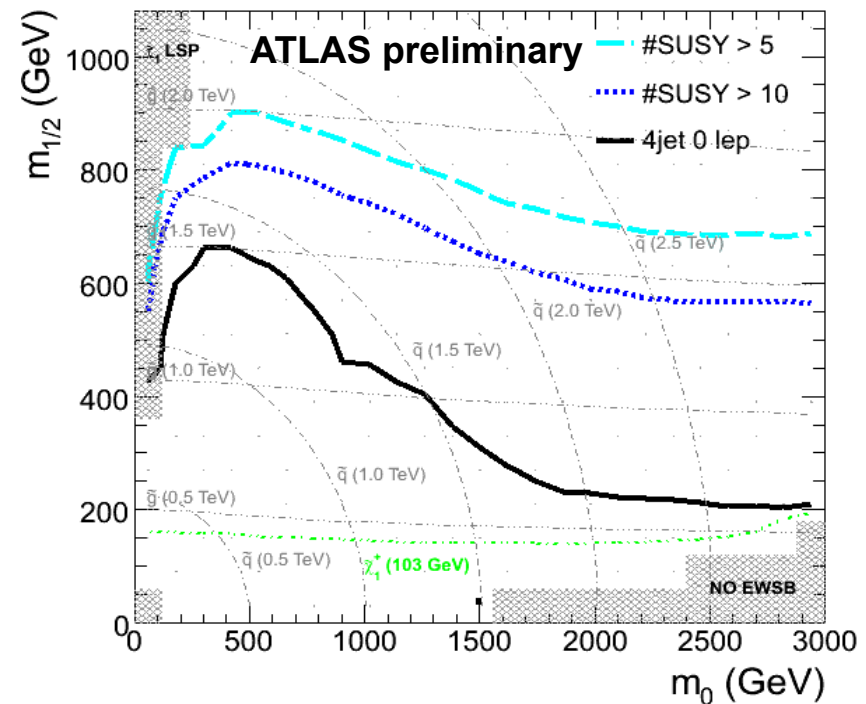
# How can we improve the discovery reach?

4 jet 0 lepton channel at  $1\text{fb}^{-1}$

CSC 5: we only claim a discovery with  $\geq 6$  data events (Susy+ bg)

blue line shows max. discovery reach we could have at all

→ improvement possible at high  $m_0$

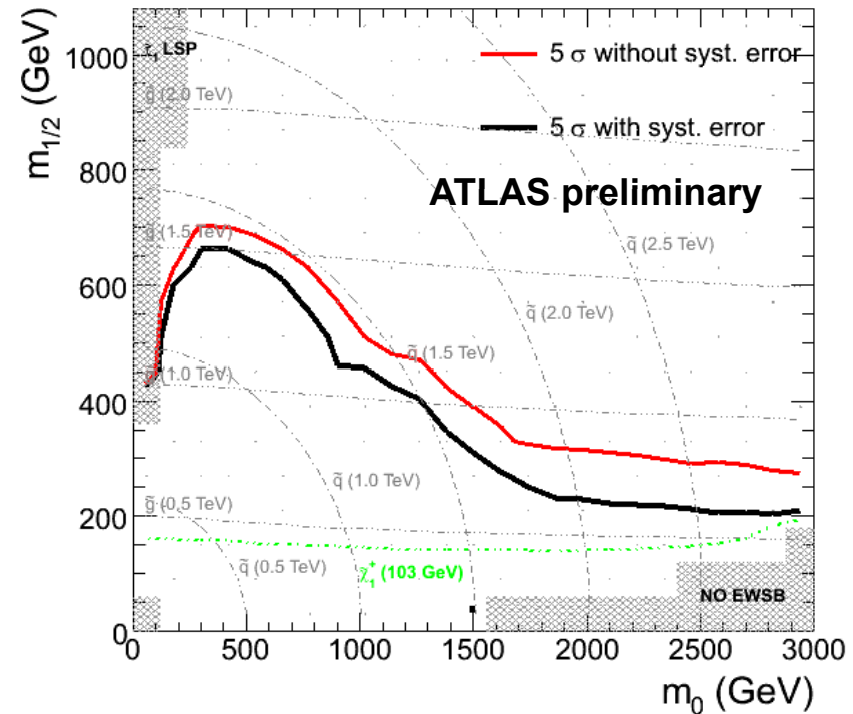


plot: M. Rammensee

# How big is the effect of systematic uncertainties?

5  $\sigma$  contour lines for 4 jet 0 lepton channel at  $1\text{fb}^{-1}$  with and without uncertainties

uncertainties have less influence on discovery reach than expected!



plot: M. Rammensee

# Systematic uncertainties

Assuming uncertainties for different samples are gaussian distributed and uncorrelated

**total systematic uncertainty for SM background:**

for each bin in the  $M_{\text{eff}}$  distribution:

$$\delta SM = \sqrt{\delta W_{e\nu}^2 + \delta W_{\mu\nu}^2 + \dots + \delta Z_{ee}^2 + \dots + \delta QCD^2}$$

and:

$$\delta \text{ sample} = \sqrt{(\text{systematic error})^2 + (\text{statistical error})^2}$$

Note:

the statistical error gives the const. systematic error a bin dependence in the  $M_{\text{eff}}$  distribution

# Uncertainties for backgrounds

---

	method	stat. unc.	syst. unc.
QCD	combine MC and data-driven jet smearing	1 %	50 %
$Z \rightarrow \nu\nu$	data (or MC) from $Z \rightarrow ll$ ( $Z \rightarrow ee$ / $Z \rightarrow \mu\mu$ )	8-13 %	10-15 %
top+W	reverse cut on $M_T$	4-8 %	15 %
top $\rightarrow \tau$	select sl $\tau$ decay	6 %	15 %
tt $\rightarrow l\nu b\bar{b}jj$	top sample using top mass	5 %	22 %
tt $\rightarrow l\nu l\nu b\bar{b}$	control sample using $H_{T2}$ kinematic reconstruction top MC redecay method	10 %	20 %



# Calculation of the probability

- the background can never be known exactly
- uncertainties on the background are incorporated in the significance by convoluting the Poisson probability that the background fluctuates to the observed signal with a Gaussian background probability density function with mean  $N_b$  and standard deviation  $\delta N_b$
- given these assumptions, the probability  $p$  that the background fluctuates by chance to the measured value  $N_{\text{data}}$  or above is given by:

$$p = A \int_0^{\infty} db G(b; N_b, \delta N_b) \sum_{i=N_{\text{data}}}^{\infty} \frac{e^{-b} b^i}{i!},$$

- where  $G(b; N_b, \delta N_b)$  is a Gaussian and the factor  $A$  ensures that the function is normalized to unity:

$$A = \left[ \int_0^{\infty} db G(b; N_b, \delta N_b) \sum_{i=0}^{\infty} e^{-b} b^i / i! \right]^{-1}$$

# Correction of probability

## looking at many different data selections :

- statistical fluctuations can be misinterpreted as new phenomena if the number of selections is not considered : 'multiple comparisons'.

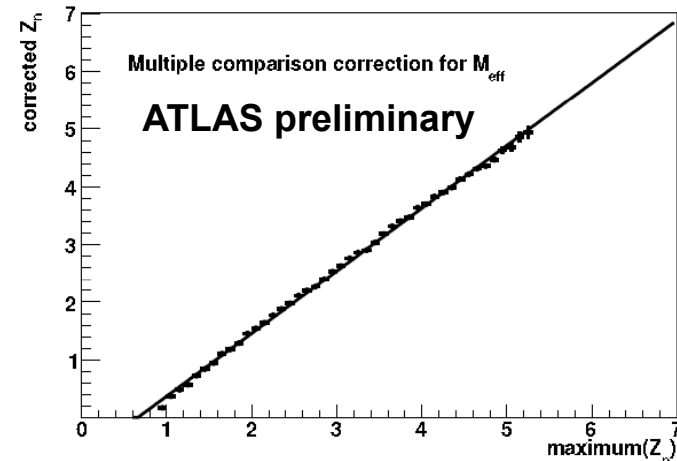
→ correct probability values for 'multiple comparisons' via a Monte Carlo method (,toy experiment')

1. generate 50 Mio.  $M_{\text{eff}}$  histograms by dicing in each bin a random number of events (poisson distributions)

- each histogram = possible outcome of measurement
- no SUSY signal is involved in this procedure

2. for each histogram apply  $M_{\text{eff}}$  scan algorithm

- select cut with largest significance  $Z_n$



3. fraction  $F$  of histograms with a larger significance than  $Z_n$  is calculated

$F$  = corrected probability for finding a deviation with significance  $\max(Z_n)$  in data  
 $F$  is comparable to the probability observed in a single counting experiment!

4. convert  $F$  into corrected  $Z_n$

# Files

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**studies are for Susy discovery potential at  $1\text{fb}^{-1}$**

**background files:**

full simulation  
for 4jet studies:

W/Z : Alpgen  
QCD: : Pythia  
DiBoson : Alpgen  
top: : MCNLO

for 2/3 jet studies: W/Z Phythia samples

**Susy signals:**

**benchmark points:**

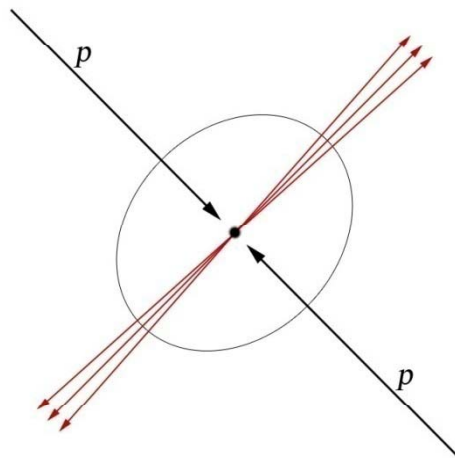
Isajet 7.74 + herwig/Jimmy + full simulation

**grids:**

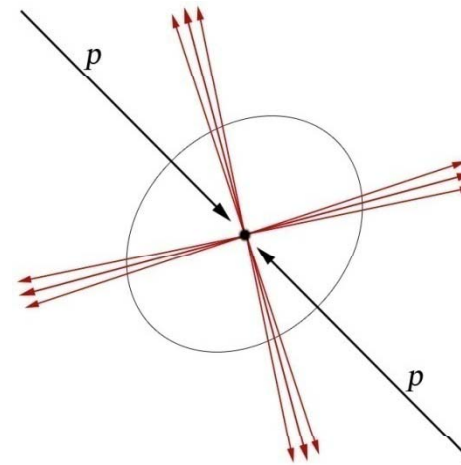
Isajet 7.75 + Herwig/Jimmy 6.5.10 + fast simulation  
+ correction

# Transverse sphericity

- measure of the isotropy of the event perpendicular to beam axis.
- $0 \leq S_T \leq 1$



$$S_T = 0$$



$$S_T = 1$$

# Files

## Grids:

### Msugra grid:

$$\tan\beta = 10$$

- 600 points a 20000 events
- $\tan\beta=10$ ,  $\text{sgn } \mu=+1$ ,  $A=0$ ,  $m_t = 175 \text{ GeV}/c^2$

$$m_0 = \{60, 180, \dots, 2940\}$$

$$m_{1/2} = \{30, 90, \dots, 1470\}$$

$$\tan\beta = 50$$

- 625 points with 1000/2000 events
- $\tan\beta=50$ ,  $\text{sgn } \mu=-1$ ,  $A=0$ ,  $m_t = 175 \text{ GeV}/c^2$

$$m_0 = \{200, 400, \dots, 3000\}$$

$$m_{1/2} = \{100, 200, \dots, 1500\}$$

### *random DM constraints grid*

- about 25 points with  $\text{sgn } \mu=1$

# Files

## Grids:

## NUHM:

step size of 100 GeV in both  $m_0$  and  $m_{1/2}$   
for each point the values of  $\mu$  and  $MA$  at the weak scale are adjusted to give acceptable cold dark matter

## GMSB

described by 5+1 parameters:

$M_{\text{mess}}$  ;  $\Lambda$  ;  $N_{\text{mess}}$  ;  $C_{\text{grav}}$  ;  $\tan \beta$  ;  $\text{sign } \mu$ :

$M_{\text{mess}}$  is a mass scale of the messenger fields

$\Lambda$  is the scale of SUSY breaking

$N_{\text{mess}}$  is the number of SU(5) messenger multiplets

$\tan \beta$  is the usual ratio of vacuum expectation values  
 $C_{\text{grav}}$  is the gravitino mass factor in  $m_{3/2}$

# Files

Grids:

**GMSB**

$\Lambda$	$\tan\beta$	$M_{\text{mess}}$	$N_{\text{mess}}$	$C_{\text{grav}}$	sign $\mu$
10-80 by step 5TeV	5-40 by step 5 TeV	500 TeV	5	1	>0

**AMSB**

- expressed by 3+1 parameters:

$m_0, m_{3/2}, \tan\beta, \text{sign } \mu$

- $m_{3/2} = \{10, 20, \dots, 150\}$  TeV
- $m_0 = \{100, 300, \dots, 3700\}$  GeV

# ATLAS benchmark points

- SU1  $m_0 = 70$  GeV,  $m_{1/2} = 350$  GeV,  $A_0 = 0$ ,  $\tan\beta = 10$ ,  $\mu > 0$ . Coannihilation region where  $\tilde{\chi}_1^0$  annihilate with near-degenerate  $\tilde{\ell}$ .
- SU2  $m_0 = 3550$  GeV,  $m_{1/2} = 300$  GeV,  $A_0 = 0$ ,  $\tan\beta = 10$ ,  $\mu > 0$ . Focus point region near the boundary where  $\mu^2 < 0$ . This is the only region in mSUGRA where the  $\tilde{\chi}_1^0$  has a high higgsino component, thereby enhancing the annihilation cross-section for processes such as  $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow WW$ .
- SU3  $m_0 = 100$  GeV,  $m_{1/2} = 300$  GeV,  $A_0 = -300$  GeV,  $\tan\beta = 6$ ,  $\mu > 0$ . Bulk region: LSP annihilation happens through the exchange of light sleptons.
- SU4  $m_0 = 200$  GeV,  $m_{1/2} = 160$  GeV,  $A_0 = -400$  GeV,  $\tan\beta = 10$ ,  $\mu > 0$ . Low mass point close to Tevatron bound.
- SU6  $m_0 = 320$  GeV,  $m_{1/2} = 375$  GeV,  $A_0 = 0$ ,  $\tan\beta = 50$ ,  $\mu > 0$ . The funnel region where  $2m_{\tilde{\chi}_1^0} \approx m_A$ . Since  $\tan\beta \gg 1$ , the width of the pseudoscalar Higgs boson  $A$  is large and  $\tau$  decays dominate.
- SU8.1  $m_0 = 210$  GeV,  $m_{1/2} = 360$  GeV,  $A_0 = 0$ ,  $\tan\beta = 40$ ,  $\mu > 0$ . Variant of coannihilation region with  $\tan\beta \gg 1$ , so that only  $m_{\tilde{\tau}_1} - m_{\tilde{\chi}_1^0}$  is small.
- SU9  $m_0 = 300$  GeV,  $m_{1/2} = 425$  GeV,  $A_0 = 20$ ,  $\tan\beta = 20$ ,  $\mu > 0$ . Point in the bulk region with enhanced Higgs production

→ SU3:  $m(\text{gluino})=720$  GeV/c<sup>2</sup>,  $m(\text{squark})=630$  GeV/c<sup>2</sup>,  $\sigma=28$  pb for 1fb<sup>-1</sup>