

The SuperB Project

(Super Flavor Factory)



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November 28, 2008

Plenary ECFA

CERN



SuperB: a $10^{36} \text{ cm}^{-2} \text{ s}^{-1}$ accelerator

SuperB is a project sustained by an international collaboration aiming at a regional project:

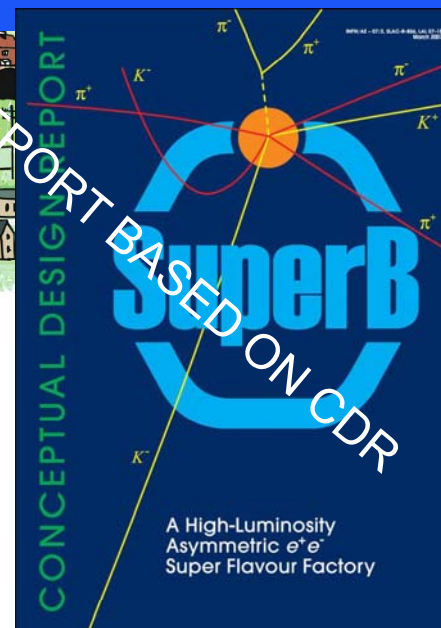
- the construction of a very high luminosity ($10^{36} \text{ cm}^{-2} \text{ s}^{-1}$) asymmetric e^+e^- **Super Flavor Factory**, with location at the campus of the University of Rome Tor Vergata, near the INFN Frascati National Laboratory
- A Super Flavor Factory such as *SuperB* will open a window on NEW PHYSICS **complimentary** to LHC and ILC
- The physics studies possible at such a machine will provide a uniquely important source of deeper understanding of the NP found at LHC, and if not found, will bring a sensitivity to seeing signs of NP at even higher energies than LHC to help set the scale of NP
- A **Conceptual Design Report**, signed by 85 Institutions was published in March 2007 ([arXiv:0709.0451 \[hep-ex\]](https://arxiv.org/abs/0709.0451))

Previous presentations of the Project

Manchester ('07) :



Lisbon (March'08)



320
Signatures
About 85
institutions
174 Babar
members
65 non
Babar.

Some Highlight on Physics Program

Quick update on Detector

**Accelerator : preliminary results from test on
SuperB concepts in DaΦne upgrade at LNF.**



CERN (Nov''08):

Quick update on Physics Program and Detector

Accelerator test results

Update on Process and Organization for TDR

After CDR

SuperB workshops and reviews of Physics motivation and machine project.

- Topical meetings on Machine, Detector and Physics (main physics meeting in Valencia Jan.7-15,2008).
- Main General meeting in Elba may 30-june 2,2008.

IRC appointed by the President of INFN in summer 2007:

- Preliminary meeting in Rome end July 2007 (committee with INFN management and proponents)
- First review meeting in LNF Nov. 12-13 , 2007.
- Final meeting before report to President of INFN in Rome, Apr. 29-30, 2008 .

Machine advisory committee (MiniMac) appointed after Elba.

- Meeting on July 16-17,2008 in Frascati.

SuperB major meetings



Proceedings of SuperB Workshop VI

New Physics at the Super Flavour Factory

Valencia, Spain
January 7-15, 2008



May 2008 SuperB Meeting

La Biodola, Isola d'Elba
May 31st-June 3rd, 2008

Istituto Nazionale Fisica Nucleare - Sez. Pisa
Universita' di Pisa, Dipartimento di Fisica
Scuola Normale Superiore



**Attended by ECFA ad hoc
Subcommittee**

Update on physics (potential discovery of New Physics with a 75 ab⁻¹ in 5 years) for B, Charm, Tau's and new Spectroscopy. **Examine carefully the potential benefits of running at 4 GeV c.m.s. Energy and of the Polarization.**

Organize the preparation of the simulation tools to evaluate the correct experimental sensitivity to the most relevant physics channels

Reviews

INFN International Review Committee

- John Dainton – UK/Daresbury, chair
- Jacques Lefrancois – F/Orsay
- Antonio Masiero – I/Padova
- Rolf D.Heuer *– D/ Desy
- Daniel Schulte – CERN
- Abe Seiden – USA/UCSC
- Young-Kee Kim – USA/FNAL
- Hiroaki Aihara – Japan/Tokyo
- + Tatsuya Nakada (RECFA)
- + Steve Myers – accel expert

Report released at the end of May 2008
They

“recommend strongly continuation of
work for $10^{36} \text{ cm}^{-2} \text{ s}^{-1}$ asymmetric e^+e^-
collider.”

(*)R.D.Heuer attended only the last meeting

DOE Particle Physics Project Prioritization Panel (P5)

Report released May 2008

Recommend significant US
Participation in offshore flavour
factory in the intermediate funding
scenario

Mini Machine Advisory Committee

- Klaus Balewski (DESY)
- John Corlett (LBNL)
- Jonathan Dorfman (SLAC, Chair)
- Tom Himel (SLAC/ DESY)
- Claudio Pellegrini (UCLA)
- Daniel Schulte (CERN)
- Ferdi Willeke (BNL)
- Andy Wolski (Liverpool)
- Frank Zimmermann (CERN)

First meeting in July 16-17,2008

No glaring showstoppers
Form a management structure

Comments on reviews

- Link to meetings and reports:

<http://www.pi.infn.it/SuperB/reviews>

- Dainton committee →
- Mini MAC

- Very exciting project -- Committee is exhilarated by the challenge
- Physics requirement of $10^{36} \text{ cm}^{-2} \text{ sec}^{-1}$ or 75 ab-1/5yr is very demanding
- Committee considers the SINGLE MOST ESSENTIAL ingredient for moving forward **is the formation of a sanctioned management structure which formally incorporates a dedicated machine design team**. The team members must have the strong support of their home institutions to work on the design. The team needs a designated leader, who is as close to full time as is possible
- The Committee sees no glaring showstoppers wrt achieving the design performance. However, in several key areas, more work is needed before the design can be blessed

28 Nov, 2008

Marcello A. Giorgi



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John Dainton
INFN SuperB
La Biolada
June 1st 2008

IRC First report



5. Conclusion

- recommend strongly continuation of work for $10^{36} \text{ cm}^{-2} \text{ s}^{-1}$ asymmetric e^+e^- collider
- even more concerted effort to fully evaluate physics potential ↔ machine specifications
- major design program to establish credibility of machine now **critical** ← showstoppers?
- MAC now essential
- preservation of detectors

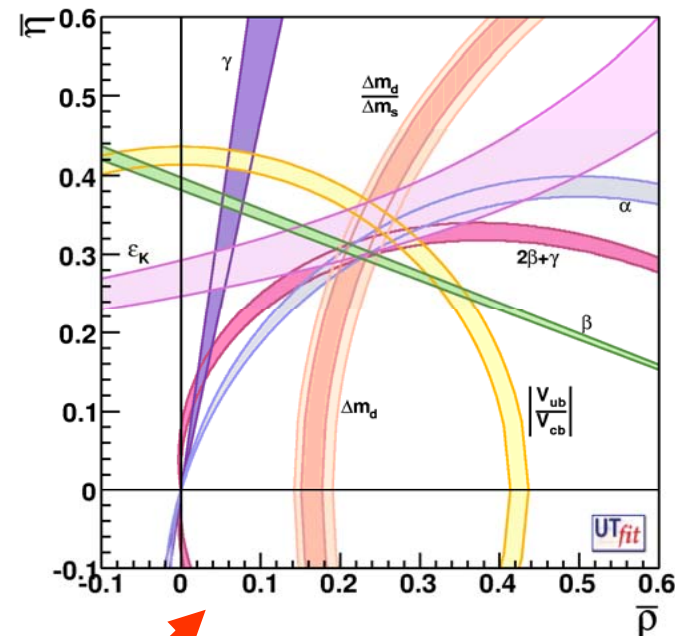
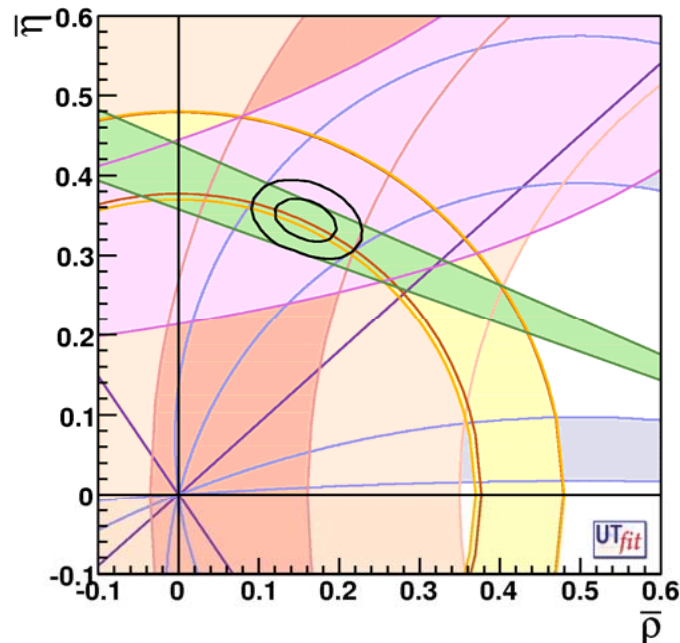
PEP2 components

→ increasing global involvement if timescale for a TDR is to be met

Still on CKM:NP sensitivity with 75 ab⁻¹

Today

SuperB+Lattice improvements



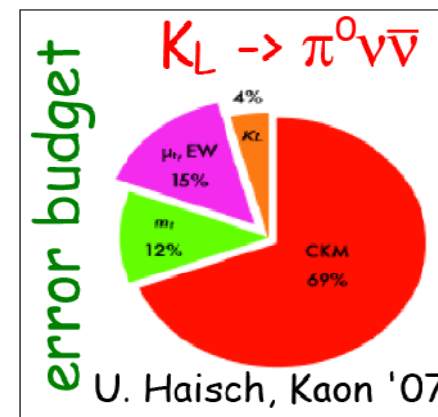
$$\rho = 0.163 \pm 0.028$$

$$\eta = 0.344 \pm 0.016$$

$$\rho = \pm 0.0028$$

$$\eta = \pm 0.0024$$

Improving CKM is
crucial to look for NP



B Physics @ Y(4S)

Observable	B factories (2 ab ⁻¹)	SuperB (75 ab ⁻¹)	Observable	B Factories (2 ab ⁻¹)	SuperB (75 ab ⁻¹)
$\sin(2\beta) (J/\psi K^0)$	0.018	0.005 (†)	$ V_{cb} $ (exclusive)	4% (*)	1.0% (*)
$\cos(2\beta) (J/\psi K^{*0})$	0.30	0.05	$ V_{cb} $ (inclusive)	1% (*)	0.5% (*)
$\sin(2\beta) (Dh^0)$	0.10	0.02	$ V_{ub} $ (exclusive)	8% (*)	3.0% (*)
$\cos(2\beta) (Dh^0)$	0.20	0.04	$ V_{ub} $ (inclusive)	8% (*)	2.0% (*)
$S(J/\psi \pi^0)$	0.10	0.02	$\mathcal{B}(B \rightarrow \tau \nu)$	20%	4% (†)
$S(D^+ D^-)$	0.20	0.03	$\mathcal{B}(B \rightarrow \mu \nu)$	visible	5%
$S(\phi K^0)$	0.13	0.02 (*)	$\mathcal{B}(B \rightarrow D \tau \nu)$	10%	2%
$S(\eta' K^0)$	0.05	0.01 (*)	$\mathcal{B}(B \rightarrow \rho \gamma)$	15%	3% (†)
$S(K_s^0 K_s^0 K_s^0)$	0.15	0.02 (*)	$\mathcal{B}(B \rightarrow \omega \gamma)$	30%	5%
$S(K_s^0 \pi^0)$	0.15	0.02 (*)	$A_{CP}(B \rightarrow K^* \gamma)$	0.007 (†)	0.004 († *)
$S(\omega K_s^0)$	0.17	0.03 (*)	$A_{CP}(B \rightarrow \rho \gamma)$	~ 0.20	0.05
$S(f_0 K_s^0)$	0.12	0.02 (*)	$A_{CP}(b \rightarrow s \gamma)$	0.012 (†)	0.004 (†)
$\gamma (B \rightarrow DK, D \rightarrow CP \text{ eigenstates})$	~ 15°	2.5°	$A_{CP}(b \rightarrow (s+d) \gamma)$	0.03	0.006 (†)
$\gamma (B \rightarrow DK, D \rightarrow \text{suppressed states})$	~ 12°	2.0°	$S(K_s^0 \pi^0 \gamma)$	0.15	0.02 (*)
$\gamma (B \rightarrow DK, D \rightarrow \text{multibody states})$	~ 9°	1.5°	$S(\rho^0 \gamma)$	possible	0.10
$\gamma (B \rightarrow DK, \text{combined})$	~ 6°	1-2°	$A_{CP}(B \rightarrow K^* \ell \ell)$	7%	1%
$\alpha (B \rightarrow \pi \pi)$	~ 16°	3°	$A^{FB}(B \rightarrow K^* \ell \ell)_{s_0}$	25%	9%
$\alpha (B \rightarrow \rho \rho)$	~ 7°	1-2° (*)	$A^{FB}(B \rightarrow X_s \ell \ell)_{s_0}$	35%	5%
$\alpha (B \rightarrow \rho \pi)$	~ 12°	2°	$\mathcal{B}(B \rightarrow K \nu \bar{\nu})$	visible	20%
$\alpha (\text{combined})$	~ 6°	1-2° (*)	$\mathcal{B}(B \rightarrow \pi \nu \bar{\nu})$	-	possible
$2\beta + \gamma (D^{(*)\pm} \pi^\mp, D^\pm K_s^0 \pi^\mp)$	20°	5°			

Charm mixing and CP

Mode	Observable	$\Upsilon(4S)$ (75 ab ⁻¹)	$\psi(3770)$ (300 fb ⁻¹)
$D^0 \rightarrow K^+ \pi^-$	x'^2	3×10^{-5}	
	y'	7×10^{-4}	
$D^0 \rightarrow K^+ K^-$	y_{CP}	5×10^{-4}	
$D^0 \rightarrow K_s^0 \pi^+ \pi^-$	x	4.9×10^{-4}	
	y	3.5×10^{-4}	
	$ q/p $	3×10^{-2}	
	ϕ	2°	
$\psi(3770) \rightarrow D^0 \bar{D}^0$	x^2		$(1-2) \times 10^{-5}$
	y		$(1-2) \times 10^{-3}$
	$\cos \delta$		$(0.01-0.02)$

Charm FCNC

Channel	Sensitivity
$D^0 \rightarrow e^+ e^-, D^0 \rightarrow \mu^+ \mu^-$	1×10^{-8}
$D^0 \rightarrow \pi^0 e^+ e^-, D^0 \rightarrow \pi^0 \mu^+ \mu^-$	2×10^{-8}
$D^0 \rightarrow \eta e^+ e^-, D^0 \rightarrow \eta \mu^+ \mu^-$	3×10^{-8}
$D^0 \rightarrow K_s^0 e^+ e^-, D^0 \rightarrow K_s^0 \mu^+ \mu^-$	3×10^{-8}
$D^+ \rightarrow \pi^+ e^+ e^-, D^+ \rightarrow \pi^+ \mu^+ \mu^-$	1×10^{-8}
$D^0 \rightarrow e^\pm \mu^\mp$	1×10^{-8}
$D^+ \rightarrow \pi^+ e^\pm \mu^\mp$	1×10^{-8}
$D^0 \rightarrow \pi^0 e^\pm \mu^\mp$	2×10^{-8}
$D^0 \rightarrow \eta e^\pm \mu^\mp$	3×10^{-8}
$D^0 \rightarrow K_s^0 e^\pm \mu^\mp$	3×10^{-8}
$D^+ \rightarrow \pi^- e^+ e^+, D^+ \rightarrow K^- e^+ e^+$	1×10^{-8}
$D^+ \rightarrow \pi^- \mu^+ \mu^+, D^+ \rightarrow K^- \mu^+ \mu^+$	1×10^{-8}
$D^+ \rightarrow \pi^- e^\pm \mu^\mp, D^+ \rightarrow K^- e^\pm \mu^\mp$	1×10^{-8}

τ Physics

Sensitivity

$\mathcal{B}(\tau \rightarrow \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow e \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow eee)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow \mu \eta)$	4×10^{-10}
$\mathcal{B}(\tau \rightarrow e \eta)$	6×10^{-10}
$\mathcal{B}(\tau \rightarrow \ell K_s^0)$	2×10^{-10}

B_s Physics @ Y(5S)

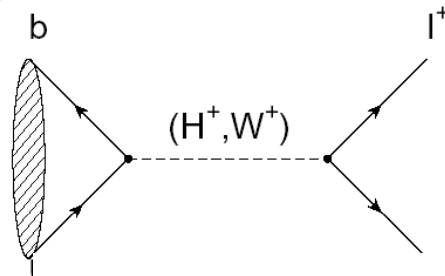
	ab ⁻¹	Error with 30 ab ⁻¹
$\Delta\Gamma$	0.16 ps ⁻¹	0.03 ps ⁻¹
Γ	0.07 ps ⁻¹	0.01 ps ⁻¹
β_s from angular analysis	20°	8°
A_{SL}^*	0.006	0.004
A_{CH}	0.004	0.004
$\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)$	-	$< 8 \times 10^{-9}$
$ V_{td}/V_{ts} $	0.08	0.017
$\mathcal{B}(B_s \rightarrow \gamma \gamma)$	38%	7%
β_s from $J/\psi \phi$	10°	3°
β_s from $B_s \rightarrow K^0 \bar{K}^0$	24°	11°

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some **EXAMPLES:**

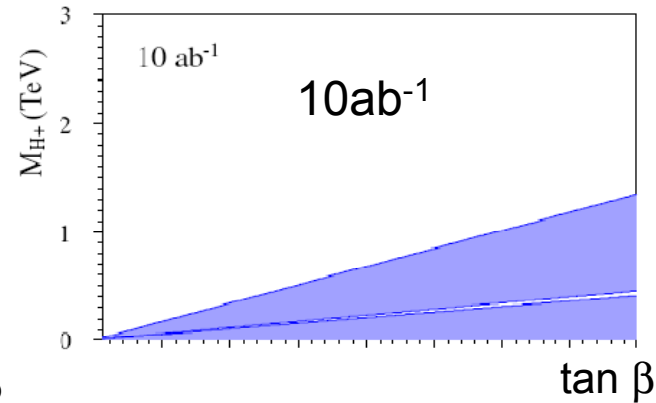
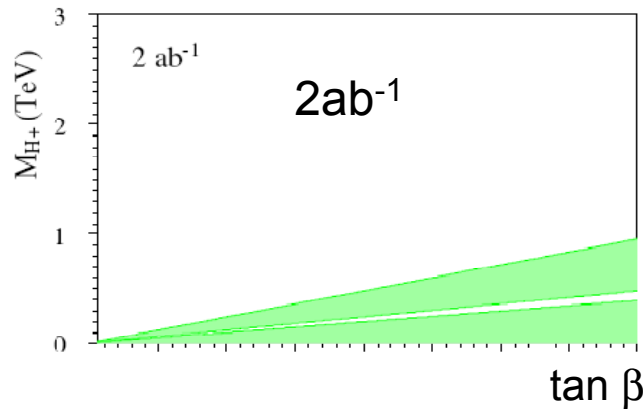
Higgs-mediated NP in MFV at large $\tan\beta$



$$\text{BR}(B \rightarrow \tau \nu) = \text{BR}_{\text{SM}}(B \rightarrow \tau \nu) \left(1 - \frac{m_B^2}{M_H^2} \tan^2 \beta \right)^2$$

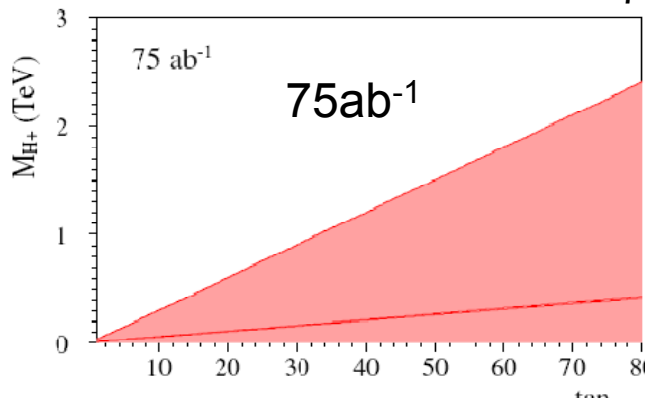
2ab⁻¹

$M_H \sim 0.4-0.8$ TeV
for $\tan\beta \sim 30-60$

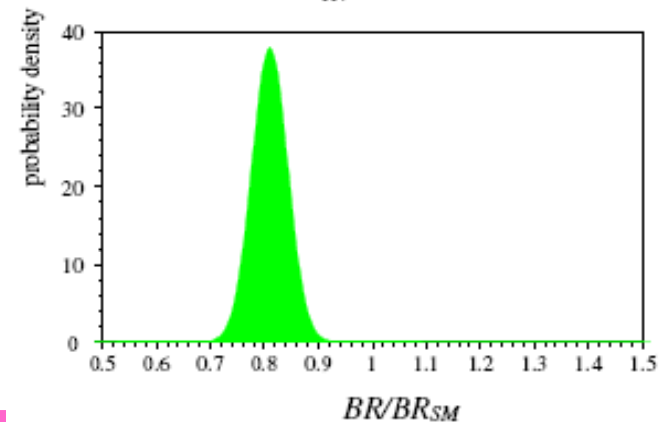


SuperB -75ab⁻¹

$M_H \sim 1.2-2.5$ TeV
for $\tan\beta \sim 30-60$



How signal would like
with $M_H=350$ GeV

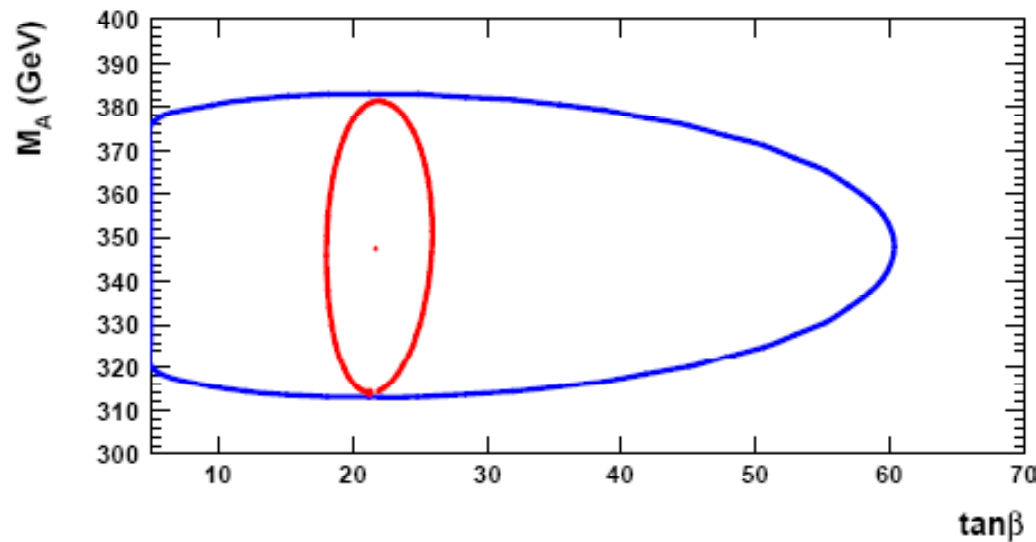


Importance of having very large sample $\geq 75\text{ab}^{-1}$

28 Nov,2008

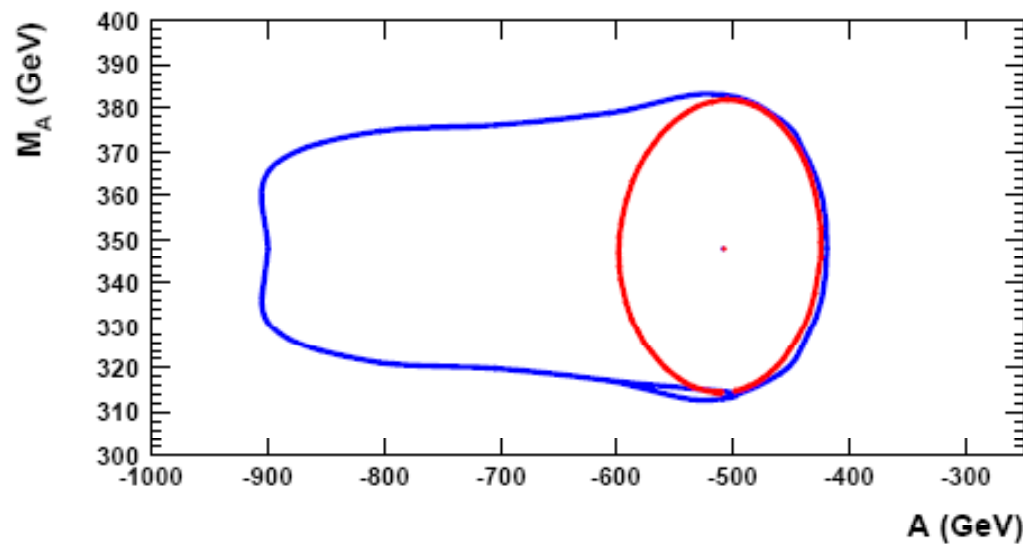
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COMPLEMENTARY: LHC and Flavour with 75 ab⁻¹



IF LHC DISCOVERS
SUPERSYMMETRY

some **EXAMPLES:**



Red are LHC+EW constraints+**SuperB**

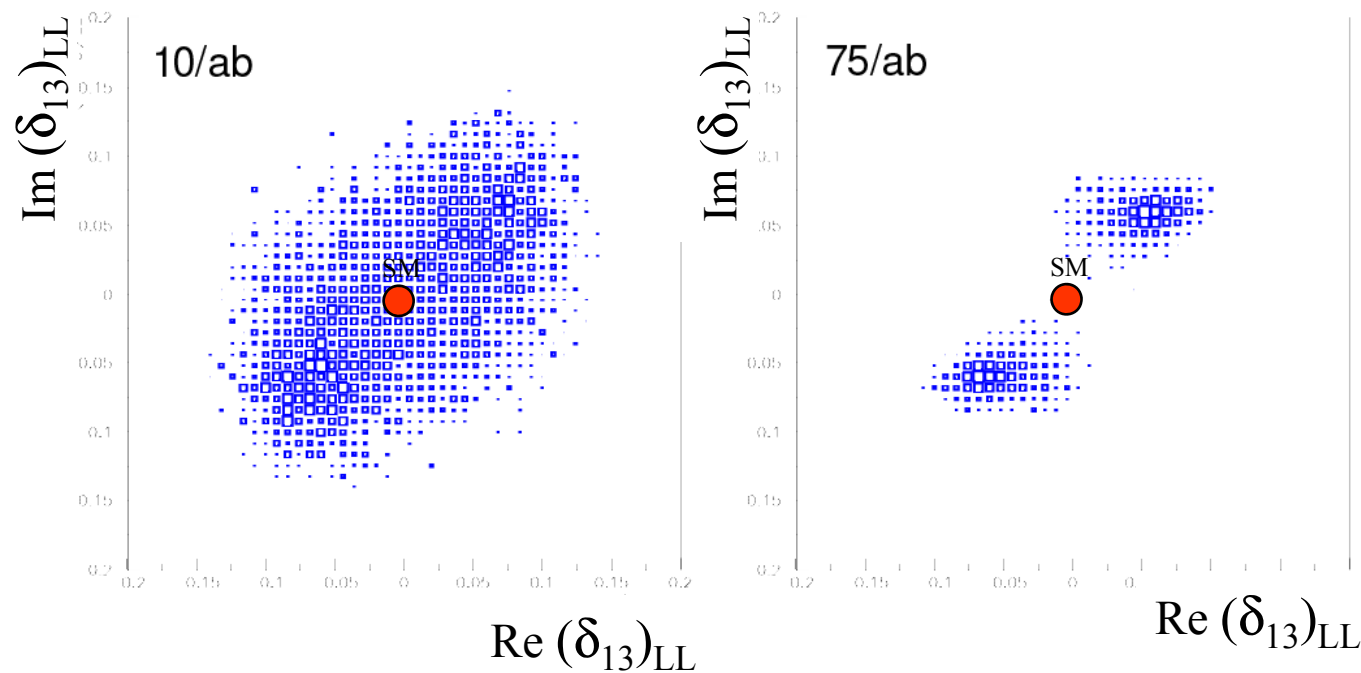
Blue is LHC alone

Importance of having very large sample $>75\text{ab}^{-1}$

Determination of coupling [in this case : $(\delta_{13})_{\text{LL}}$]

with 10ab^{-1} and 75ab^{-1}

some **EXAMPLES:**



is also a τ factory \rightarrow golden measurement LFV (Complementary to $\mu \rightarrow e \gamma$)

Process	Sensitivity
$\mathcal{B}(\tau \rightarrow \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow e \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow e e e)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow \mu \eta)$	4×10^{-10}
$\mathcal{B}(\tau \rightarrow e \eta)$	6×10^{-10}
$\mathcal{B}(\tau \rightarrow \ell K_s^0)$	2×10^{-10}

90% CL limits

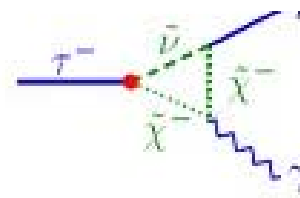
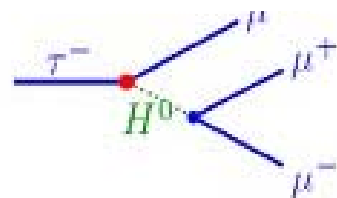
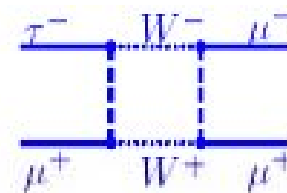
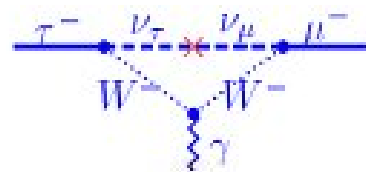
$$\text{Br}(\tau^- \rightarrow e^- \gamma) < 12 \times 10^{-8}$$

$$\text{Br}(\tau^- \rightarrow \mu^- \gamma) < 4.1 \times 10^{-8}$$

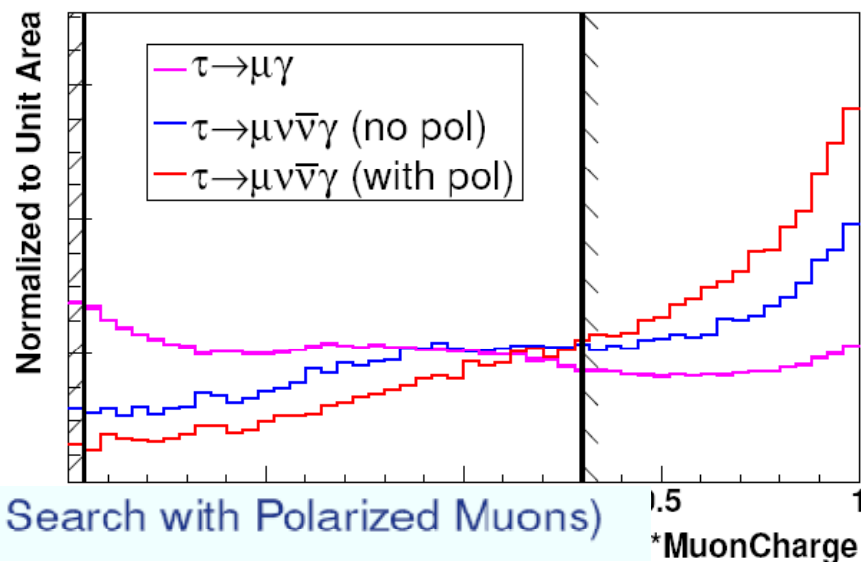


$$\text{Br}(\tau^- \rightarrow e^- \gamma) < 11 \times 10^{-8}$$

$$\text{Br}(\tau^- \rightarrow \mu^- \gamma) < 6.7 \times 10^{-8}$$



Further improvements if polarized beams.



Optimization of BKG rejection is in progress. Pol.
Helps also to discriminate models. In some model
there is a strong effect on the angular distribution
(see hep-ph/9604296, Y.Kuno, Y.Okada, $\mu \rightarrow e \gamma$ Search with Polarized Muons)

Comparison with Snowmass points on Tau using also Polarization

SuperB with 75 ab-1, evaluation assuming the most conservative scenario about syst. errors

SPS	$M_{1/2}$ (GeV)	M_0 (GeV)	A_0 (GeV)	$\tan\beta$	μ
1 a	250	100	-100	10	> 0
1 b	400	200	0	30	> 0
2	300	1450	0	10	> 0
3	400	90	0	10	> 0
4	300	400	0	50	> 0
5	300	150	-1000	5	> 0

◆ NP predictions for experimentally constrained SUSY in a number of standard scenarios
B.C.Allanach *et al.*, hep-ph/0202233

	Snowmass points predictions						SuperB	
	1 a	1 b	2	3	4	5	90% UL	4σ disc.
$\text{BF}(\tau \rightarrow \mu\gamma) \times 10^{-9}$	4.2	7.9	0.18	0.26	97	0.019	2	5
$\text{BF}(\tau \rightarrow 3\mu) \times 10^{-12}$	9.4	18	0.41	0.59	220	0.043	200	880

SuperKEKB worse by a factor 2.5 and 4.5 in $\tau \rightarrow \mu\gamma$ and >5 in $\tau \rightarrow 3\mu$

Tau g-2

Start with the expt. with μ

$$\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} \approx (3 \pm 1) \times 10^{-9}$$

assume SuperB at 75 fb^{-1} , 80% e^- beam polarization

extend to all tau decay channels

combine 2 measurement methods for $\text{Re}\{F_2\}$

studies on simulated events show no limiting syst. effects

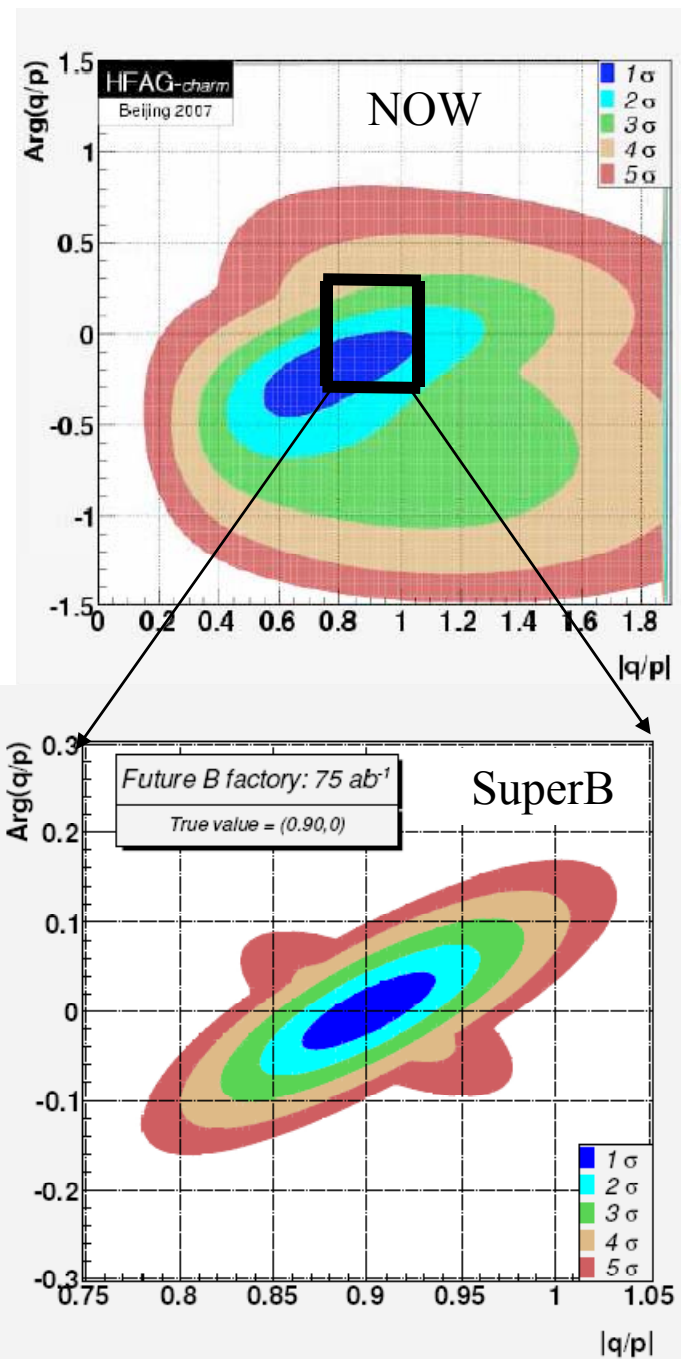
	Snowmass points predictions						SuperB
	1 a	1 b	2	3	4	5	exp. resolution
$\Delta a_\mu \times 10^{-9}$	3.1	3.2	1.6	1.4	4.8	1.1	
$\Delta a_\tau \times 10^{-6}$	0.9	0.9	0.5	0.4	1.4	0.3	<1

SuperKEKB, without beam polarization, expected worse by factor ≈ 10 , and worse systematics

Make use of all the informations (total x-section, angular distribution, f-b asymmetry.
Measure Re and Im parts

CP Violation in charm

Mode	Observable	$\Upsilon(4S)$ (75 ab ⁻¹)	$\psi(3770)$ (300 fb ⁻¹)
$D^0 \rightarrow K^+ \pi^-$	x'^2	3×10^{-5}	
	y'	7×10^{-4}	
$D^0 \rightarrow K^+ K^-$	y_{CP}	5×10^{-4}	
$D^0 \rightarrow K_S^0 \pi^+ \pi^-$	x	4.9×10^{-4}	
	y	3.5×10^{-4}	
	$ q/p $	3×10^{-2}	
	ϕ	2°	
$\psi(3770) \rightarrow D^0 \bar{D}^0$	x^2		$(1-2) \times 10^{-5}$
	y		$(1-2) \times 10^{-3}$
	$\cos \delta$		$(0.01-0.02)$

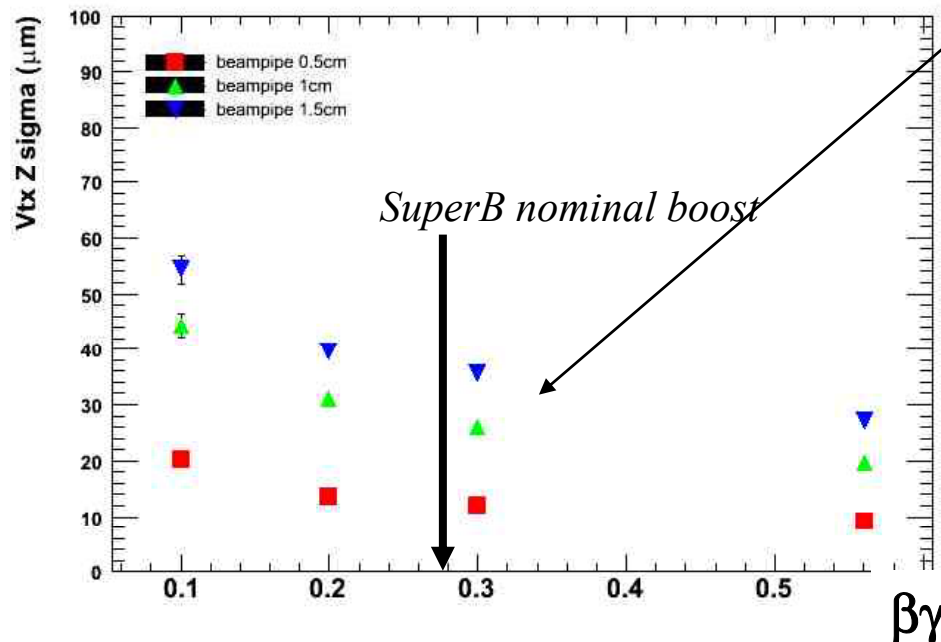


Charm at threshold

- Charm events at threshold are very clean: pure DD, no additional fragmentation
- High signal/bkg ratio: optimal for decays with neutrinos.
- Quantum Coherence: new and alternative CP violation measurement wrt to $\Upsilon(4S)$. Unique opportunity to measure D^0 - D^0 relative phase.
- Increased statistics is not an advantage running at threshold: cross-section 3x wrt 10GeV but luminosity 10x smaller.
- SuperB lumi at 4 GeV = $10^{35} \text{ cm}^{-2}\text{s}^{-1}$ produces $\sim 10^9$ DD pairs per month of running. (using Cleo-c cross-section measurement $\sigma(e^+e^- \rightarrow D^0 D^0) \sim 3.6 \text{ nb}$ + $\sigma(e^+e^- \rightarrow D^+ D^-) \sim 2.8 \text{ nb} \sim 6.4 \text{ nb}$)
- Time-dependent measurements at 4 GeV only possible at SuperB.

Time dependent measurements at DD threshold: only possible at SuperB

- Proper time resolution dominated by decay vertex resolution.
 - Production vertex precisely determined thanks to nm beamspot dimensions



SuperB lumi at 4 GeV = $10^{35} \text{ cm}^{-2}\text{s}^{-1}$
produces $\sim 10^9$ DD per month

$\beta\gamma_{ct} = 0.28 \times 120 \mu\text{m} \sim 30 \mu\text{m}$
Average flight distance similar to
vertex resolution $\rightarrow \sigma_t \sim \tau$

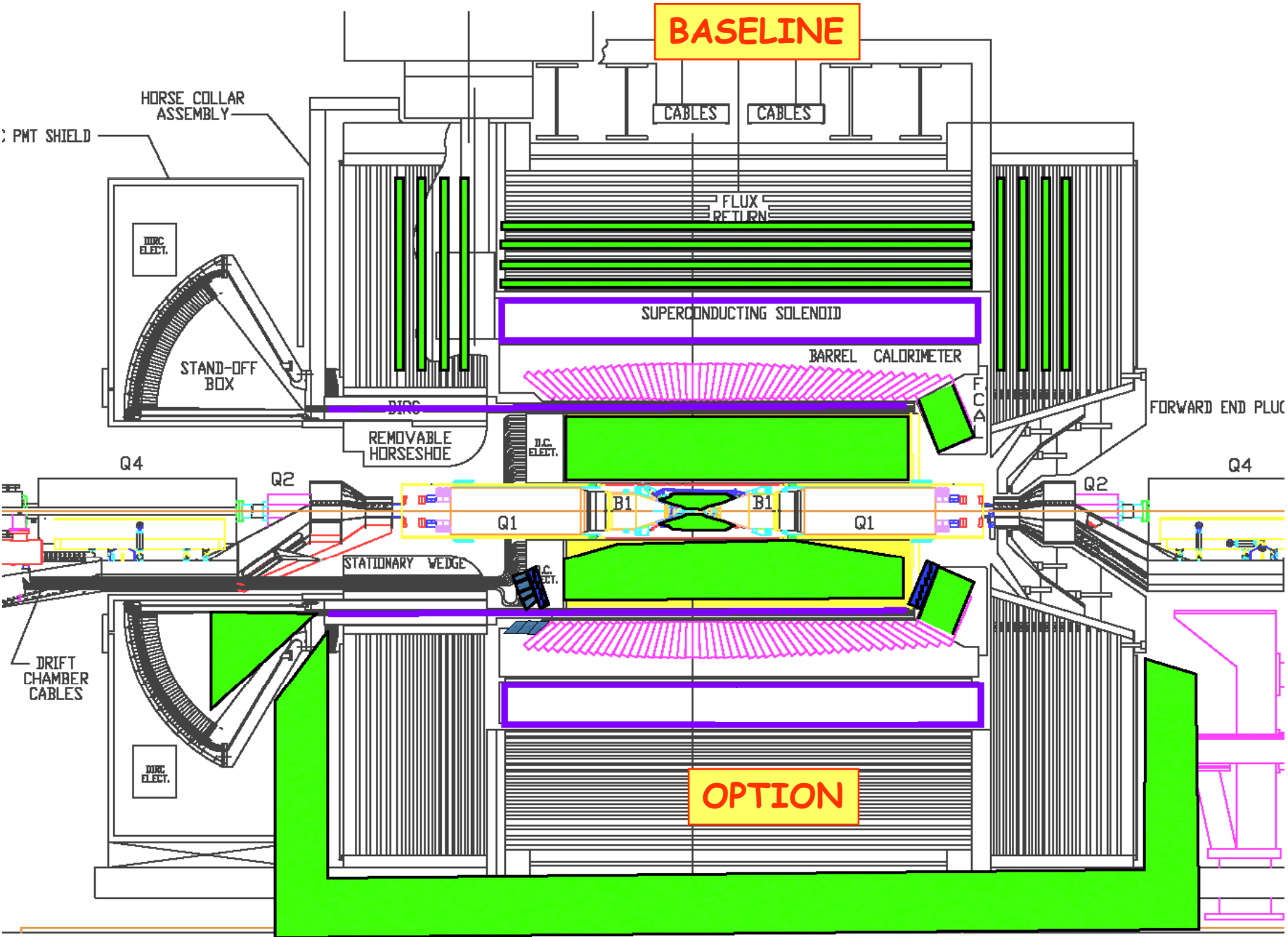
Resolution is still adequate for time dependent measurements

$$\text{Error on lifetime} \approx \sqrt{\frac{\tau^2 + \sigma_t^2}{N}} = \sqrt{2} \frac{\tau}{\sqrt{N}} \quad (\text{wrt perfect resolution})$$

SuperB Detector

- Babar and Belle designs have proven to be very effective for B-Factory physics
 - Follow the same ideas for SuperB detector
- A SuperB detector is possible with today's technology. Main issues:
 - Machine backgrounds – somewhat larger than in Babar/Belle
 - Beam energy asymmetry – a bit smaller
 - Strong interaction with machine design
- Try to reuse parts of Babar as much as possible
 - Quartz bars of the DIRC
 - Barrel EMC CsI(Tl) crystal and mechanical structure
 - Superconducting coil and flux return yoke.
- Moderate R&D and engineering required
 - Small beam pipe technology
 - Thin silicon pixel detector for first layer
 - Drift chamber CF mechanical structure, gas and cell size
 - Photon detection for DIRC quartz bars
 - Forward PID system (TOF or focusing RICH)
 - Forward calorimeter crystals (LSO)
 - Minos-style scintillator for Instrumented flux return
 - Electronics and trigger
 - Computing – large data amount

Detector Layout – Reuse parts of Babar (or Belle)



..and R&D for new components

- Beam test september '08 @ CERN (T9). Main goals:

- [REDACTED]
- [REDACTED]

ips



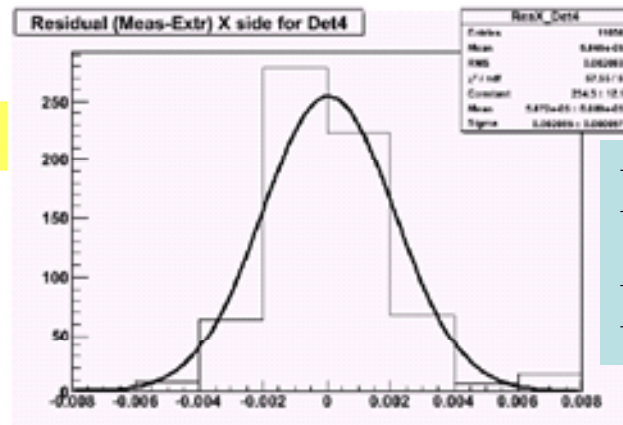
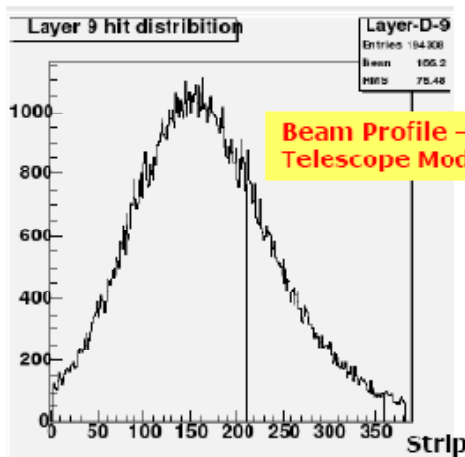
←
beam

28 Nov, 2008

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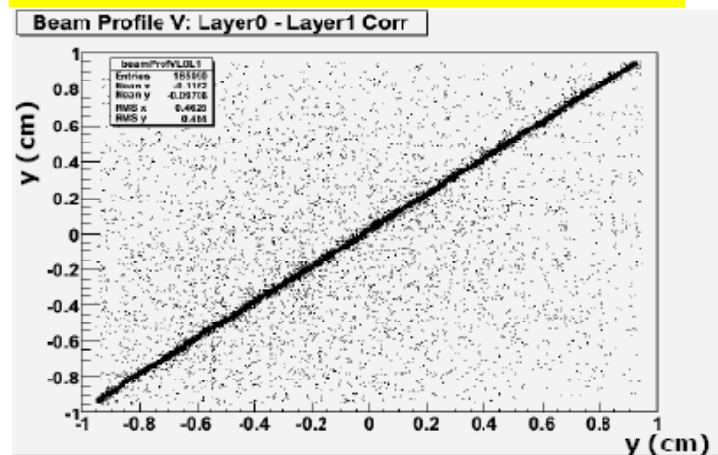
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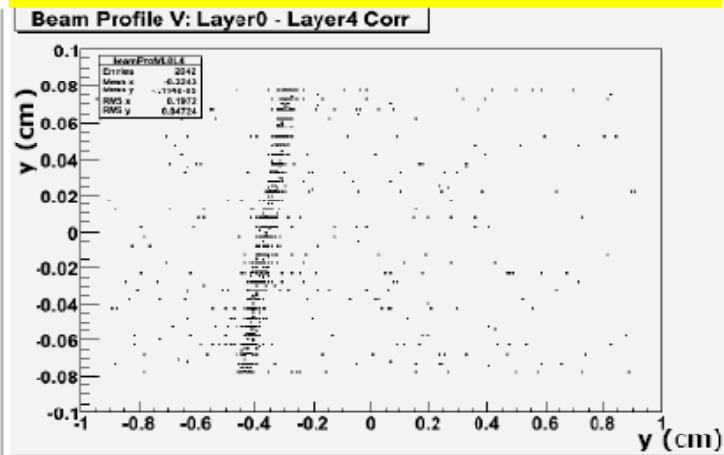
Resolution $\leq 18\mu\text{m}$

Pixel size $50\mu\text{m} \times 50\mu\text{m}$

Y hit correlation Telescope T2 vs T1



Y hit correlation MAPS vs Telescope T1



Efficiency plateau $80\div 90\%$

ACCELERATOR : High Luminosity for physics requirements

- For gaussian bunches:

$$\mathcal{L} = f_{\text{coll}} \times \frac{N_{e^+} N_{e^-}}{4\pi \sigma_x \sigma_y} \times R_l$$

geometrical
Reduction
factor

N_{e^+} (N_{e^-}) is the number of positrons (electrons) in a bunch

f_{coll} is the collision frequency

σ_x (σ_y) is the horizontal (vertical) r.m.s. size at the I.P.

R_l is the Luminosity Reduction factor by incomplete overlap: crossing angle and “hour glass” effect.

- **TRADITIONAL** (brute force): increase the numerator Currents increase: from 1 A on 2 A up to 4.1 A on 9.4 A- **Wall Plug Power**, HOM, CSR: hard to surpass $5 \cdot 10^{35} \text{ cm}^2\text{s}^{-1}$ **Crab Crossing** to increase R_l and to optimize beam dynamic

- **SuperB**: decrease the denominator (same currents as PEP-II) Bunch sizes:
from $\sigma_y = 3 \mu\text{m}$ down to $\sigma_y = 40 \text{ nm}$ Luminosity: $10^{36} \text{ cm}^2\text{s}^{-1}$ (baseline) . **Crab Waist** and large **Piinsky** angle to optimize beam dynamic

Geometric Luminosity (%)

Vertical scan

Stand.Dev.
5.56E+0

Δ Std Dev rms
2.8E-2

Max Position
9.328E-1

Δ Max Pos rms
3.5E-2

Scale Factor
6.173E+1

Δ Scale Fct rms
4.1E-1

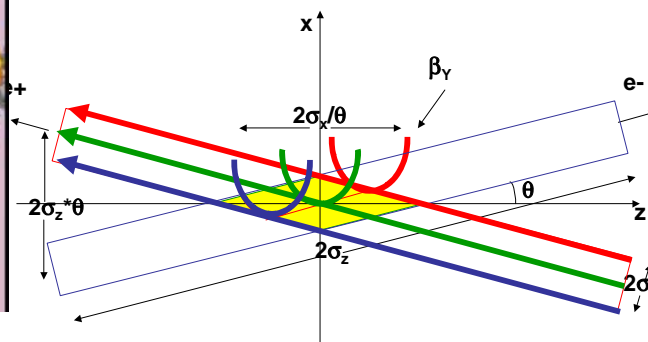
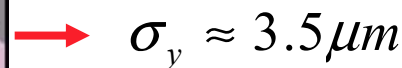
Chi square
1.77E+2

ΔL rms
5.93E-2

Step
0

L
-1.50E+1

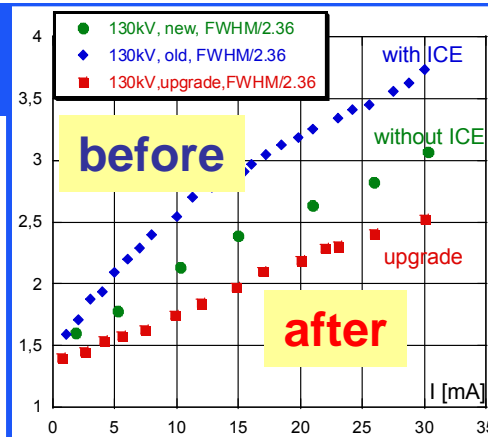
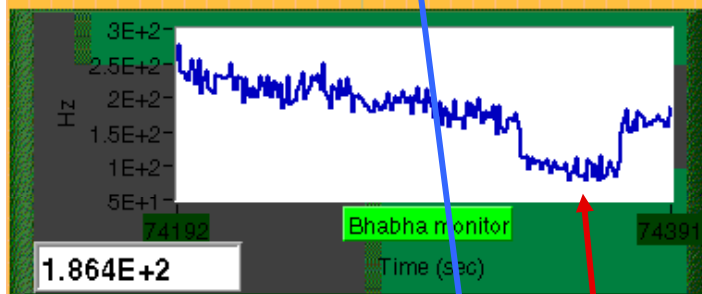
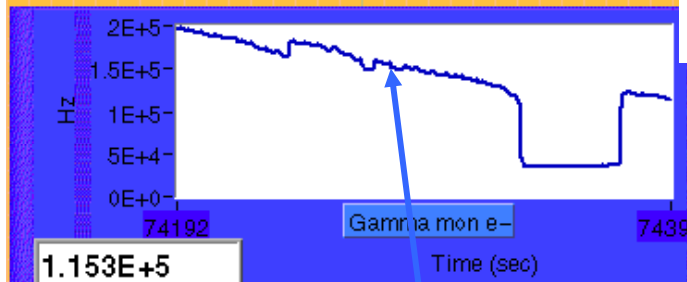
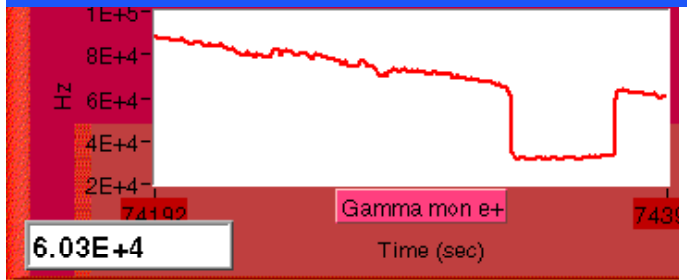
up<10%



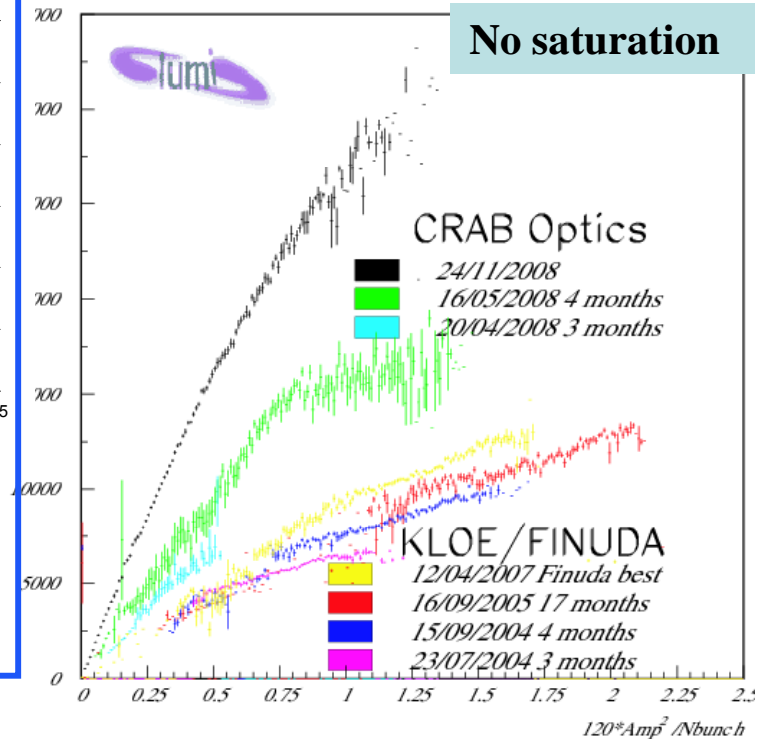
$$\Sigma_y = \sqrt{\sigma_{yp}^2 + \sigma_{ye}^2} \quad \Sigma_y = \Sigma_y^{meas} * 0.88$$

Large Piwinsky angle and “*crab waist*” with a pair of sextupoles/ring ($\Phi = tg(\theta)\sigma_z/\sigma_x$)
 Currents comparable to present Factories, lower backgrounds, less HOM and instabilities.

DaΦne test



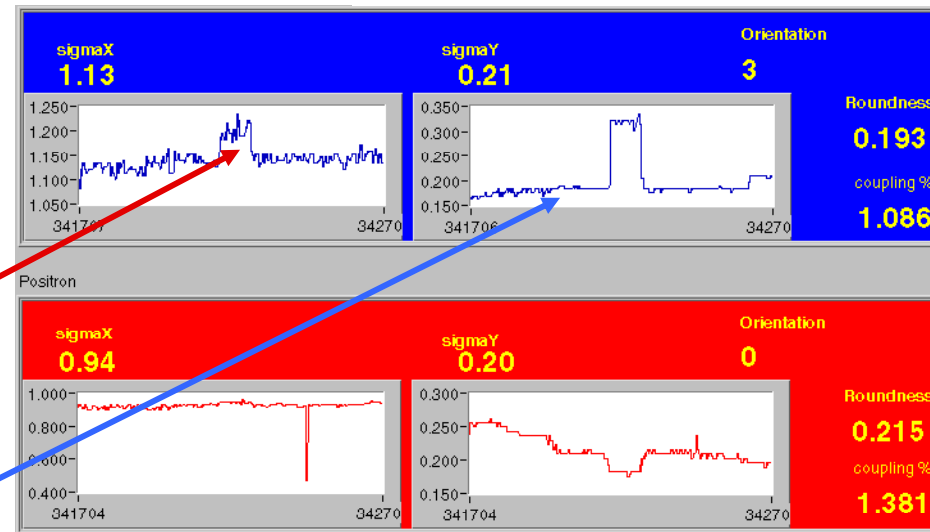
e⁻ bunch length vs
bunch current as
measured before
and after upgrade



Blow-up in beam
sizes and decrease
Bhabha rates with
crab sexts for one
ring OFF (other ring
ON) 28 Nov,2008

Crab OFF

Crab ON



Marcello A. Giorgi

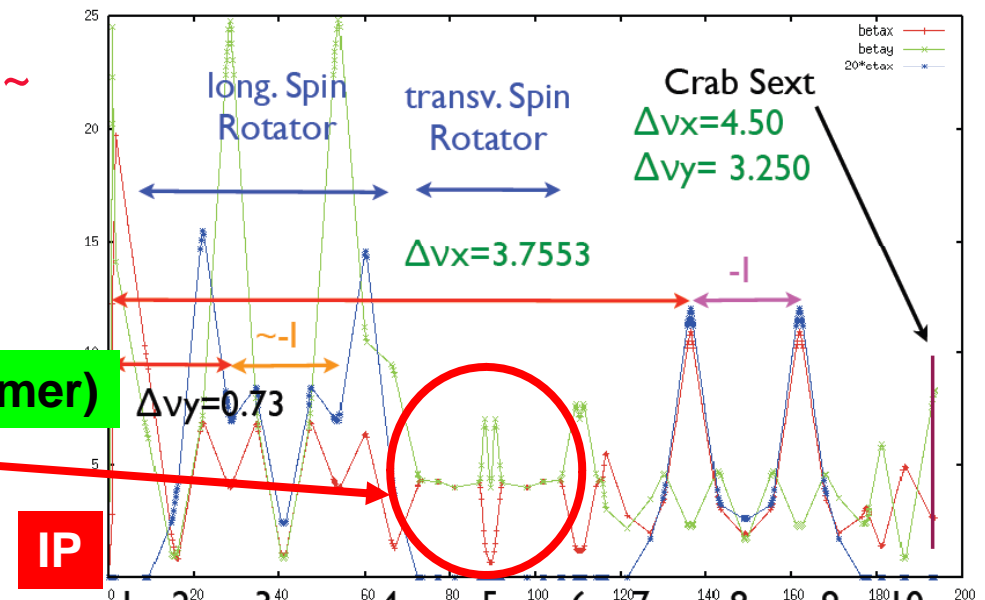


Polarization

- Polarization of one beam is included in *SuperB*
 - Either energy beam could be the polarized one
 - The LER would be less expensive, the HER easier
 - HER was chosen
- Longitudinal polarization times and short beam lifetimes indicate a need to inject vertically polarized electrons.
 - The plan is to use a polarized e^- source similar to the SLAC SLC source.
- There are several possible IP spin rotators:
 - Solenoids look better at present (vertical bends give unwanted vertical emittance growth)

- Expected longitudinal polarization at IP $\sim 87\%(\text{inj}) \times 97\%(\text{ring}) = 85\%(\text{effective})$
- Polarization section implementation in lattice is in progress

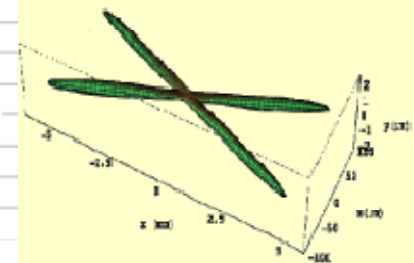
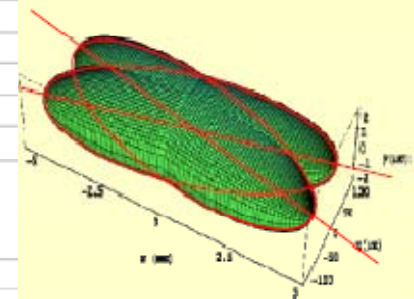
Half IR with spin rotator (Wienands, Wittmer)



SuperB parameters wrt SuperKEKB

	Nominal		Upgrade		Ultimate		SuperKEKB	
PARAMETER	LER (e+)	HER (e-)	LER (e+)	HER (e-)	LER (e+)	HER (e-)	LER (e+)	HER (e-)
Energy (GeV)	4	7	4	7	4	7	3.5	8
Luminosity $\times 10^{36}$	1.0		2.0		4.0		0.8 (0.4)	
Circumference (m)	1800	1800						
Revolution frequency (MHz)	0.167							
Eff. long. polarization (%) →	0	80						
RF frequency (MHz)	476							
Momentum spread ($\times 10^{-4}$)	7.9	5.6	9.0	8.0				
Momentum compaction ($\times 10^{-4}$)	3.2	3.8	3.2	3.8				
Rf Voltage (MV)	5	8.3	8	11.8	17.5	27		
Energy loss/turn (MeV)	1.16	1.94	1.78	2.81				
Number of bunches	1251				2502		5000	
Particles per bunch ($\times 10^{10}$)	5.52				6.78		12	5
Beam current (A) →	1.85				3.69		9.4	4.1
Beta y^* (mm)	0.22	0.39	0.16	0.27			3	
Beta x^* (mm)	35	20					200	
Emit y (pm-rad)	7	4	3.5	2			45	
Emit x (nm-rad)	2.8	1.6	1.4	0.8			9 (24)	
Sigma y^* (microns)	0.039	0.039	0.0233	0.0233			0.367	
Sigma x^* (microns)	9.9	5.66	7	4			42	
Bunch length (mm)	5		4.3				3	
Full Crossing angle (mrad)	48						30	
Wigglers (#) 20 meters each	0	0	2	2				
Damping time (trans/long)(ms)	40/20	40/20	28/14	28/14				
Luminosity lifetime (min)	6.7		3.35					
Touschek lifetime (min)	20	40	38	20				
Effective beam lifetime (min)	5.0	5.7	3.1	2.9				
Injection rate pps ($\times 10^{11}$) (100%)	2.6	2.3	5.1	4.6	10	9.1		
Tune shift y (from formula) →	0.15		0.20				0.405	
Tune shift x (from formula) →	0.0043	0.0025	0.0059	0.0034			0.209	
RF Power (MW)	17		25		58.2		83	

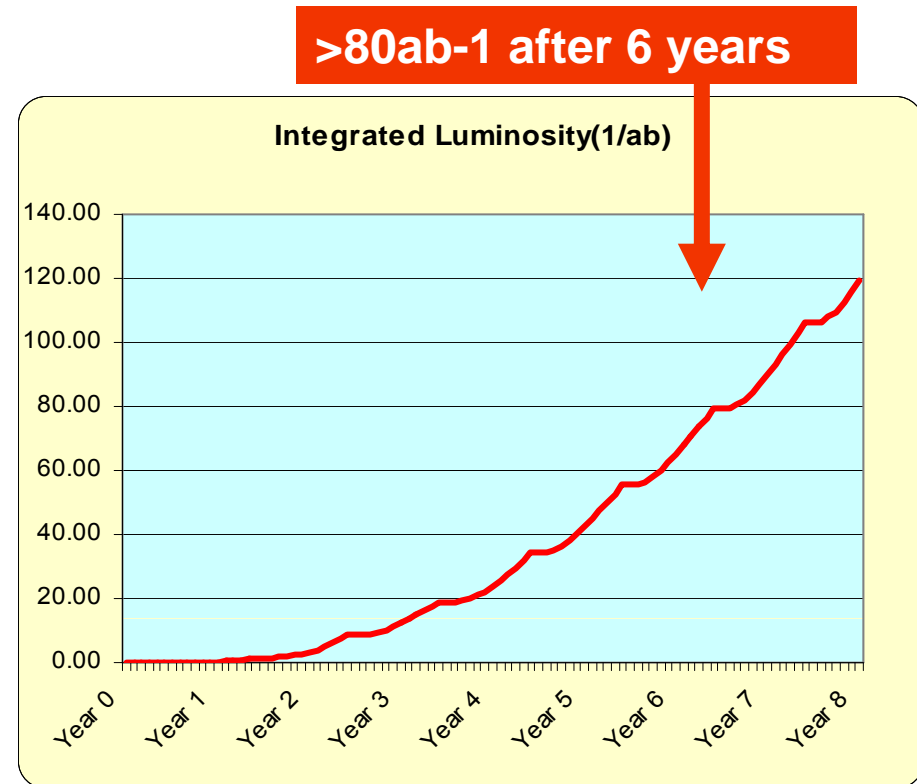
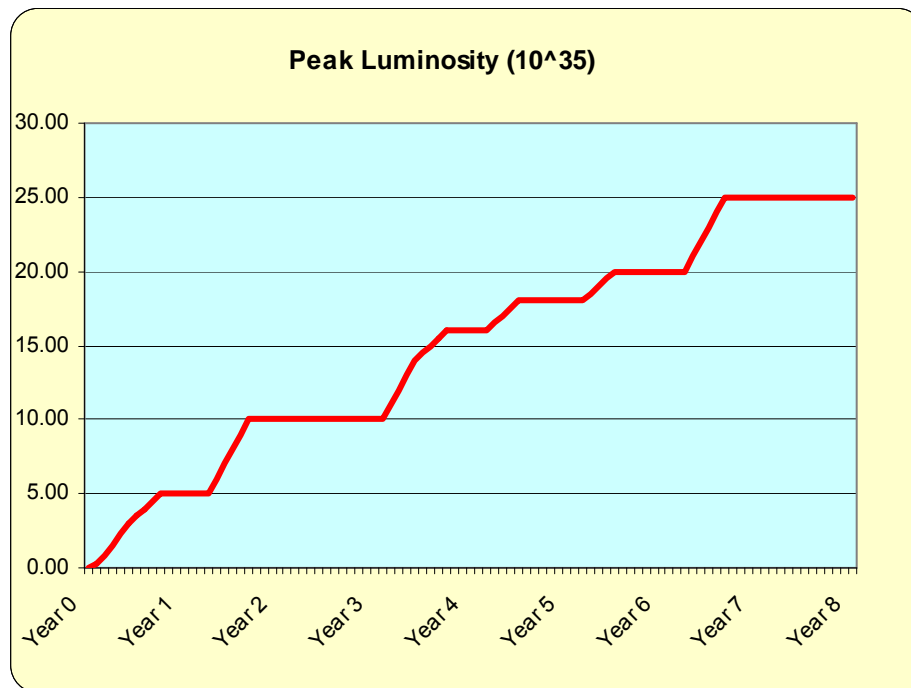
IP beam distributions for KEKB



IP beam distributions for SuperB (without transparency conditions)

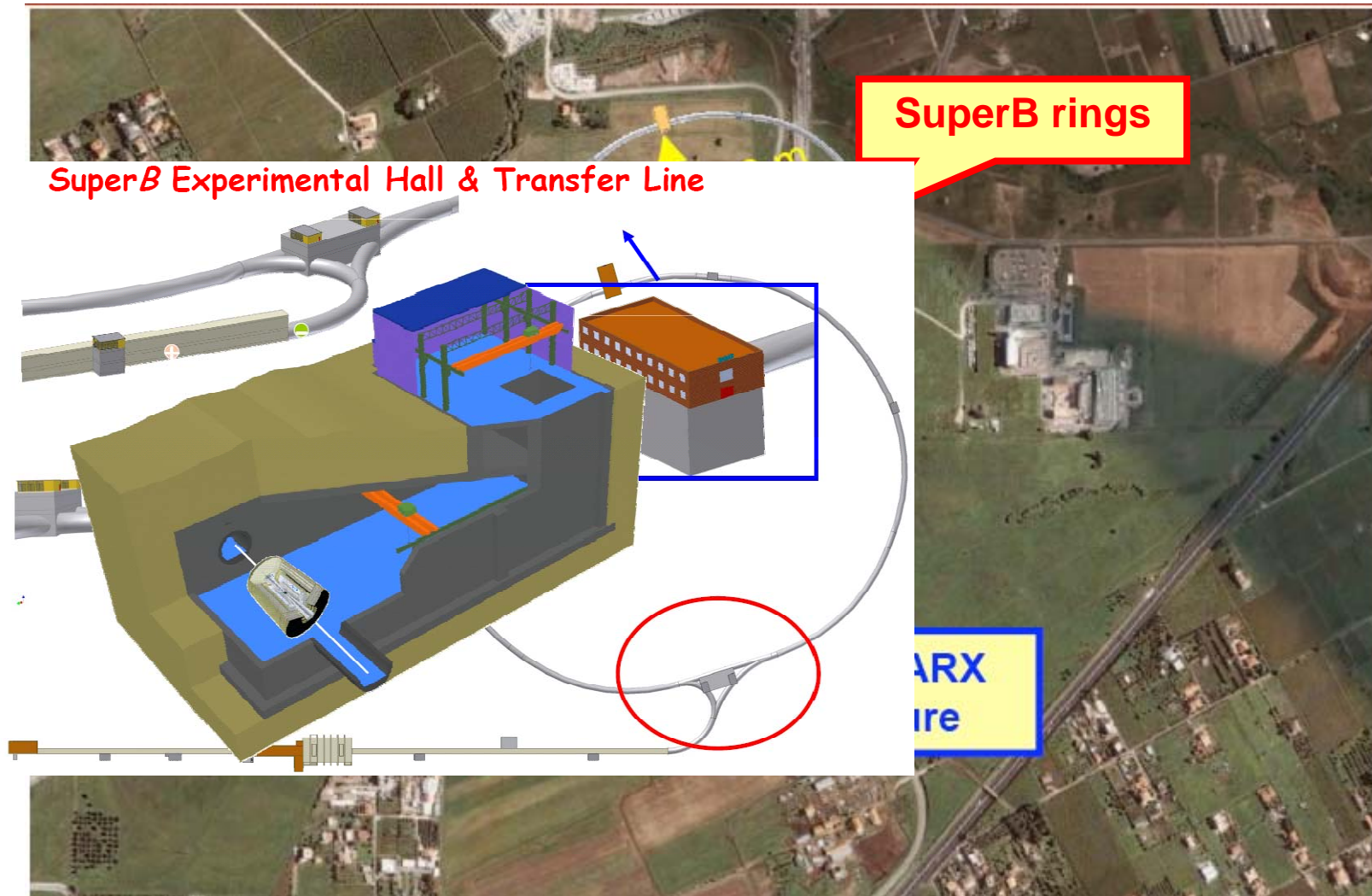
Luminosity doubled by doubling currents and wall power by a factor 2

SuperB expected LUMI



With 7th year integrated Luminosity can grow at rate of $\sim 40 \div 60 \text{ ab}^{-1}/\text{year}$

SuperB on Tor Vergata site



Cost estimate

The total estimated Project costs is ~ 500 M€.

Expected :

- in kind contribution from SLAC with a large fraction of the PEP-II components.
- From Lazio government for the infrastructure and TDR

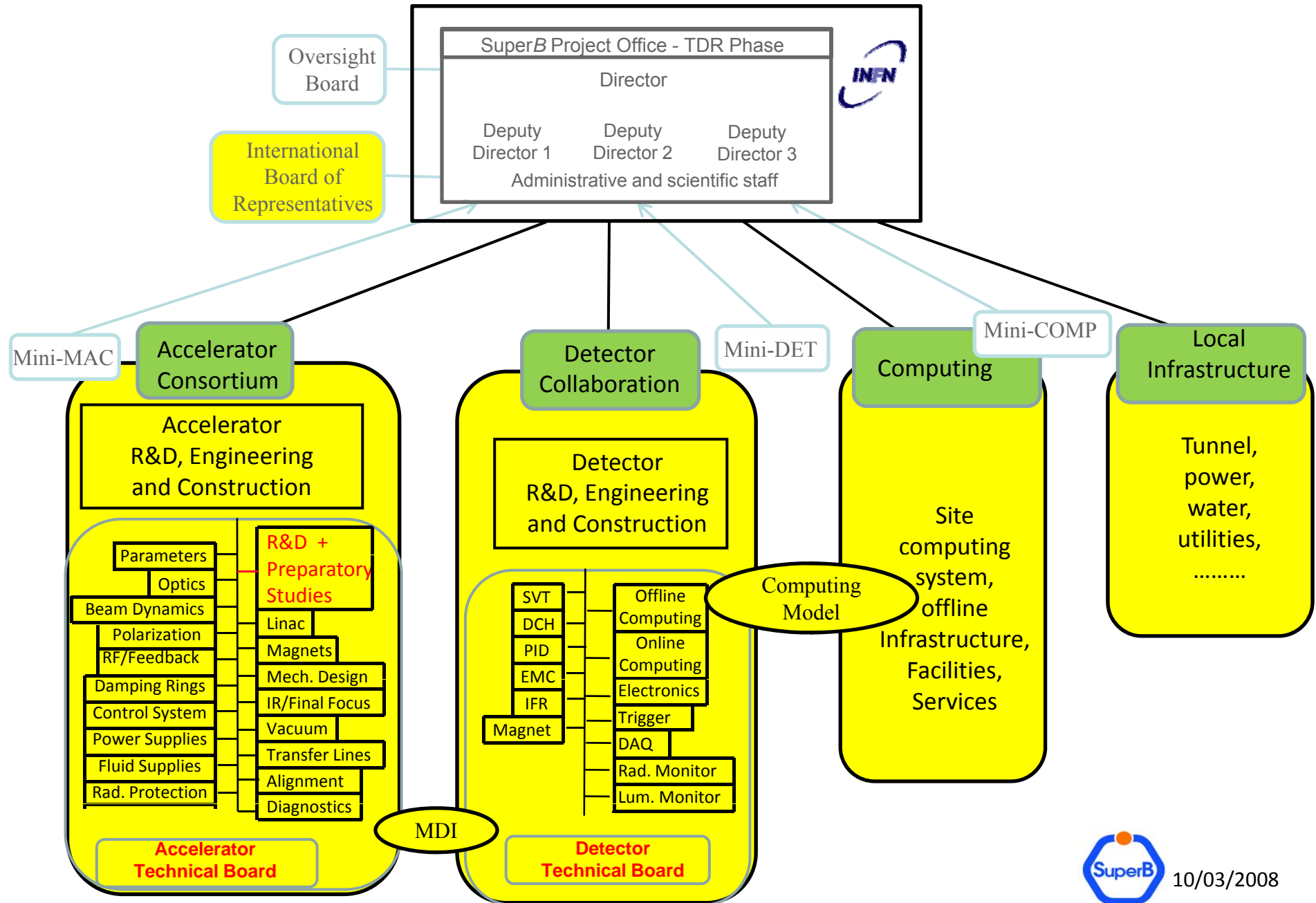
Still the remaining cost is ~ 200 M€ including detector after reuse of large part of BaBar components

An international consortium is foreseen for the construction phase and for the operation operation

A project office for TDR is being built.

We are in the phase of building a strong team for TDR

SuperB Organization Chart for TDR Phase



Conclusions

The result of crab waist tests has been satisfactory.

SuperB is entering in the TDR phase.

Project Office is being completed and project team is being assembled.

First organization meeting of collaboration to start officially the TDR phase in Paris (LAL) on Feb 15-18 2009.

MiniMAC review soon after (possibly contiguous).

Physics workshop in Warwick on April 16-19,2009.