

EUROPEAN ORGANIZATION FOR NUCLEAR
RESEARCH

Proposal to the ISOLDE and Neutron Time-of-Flight
Committee

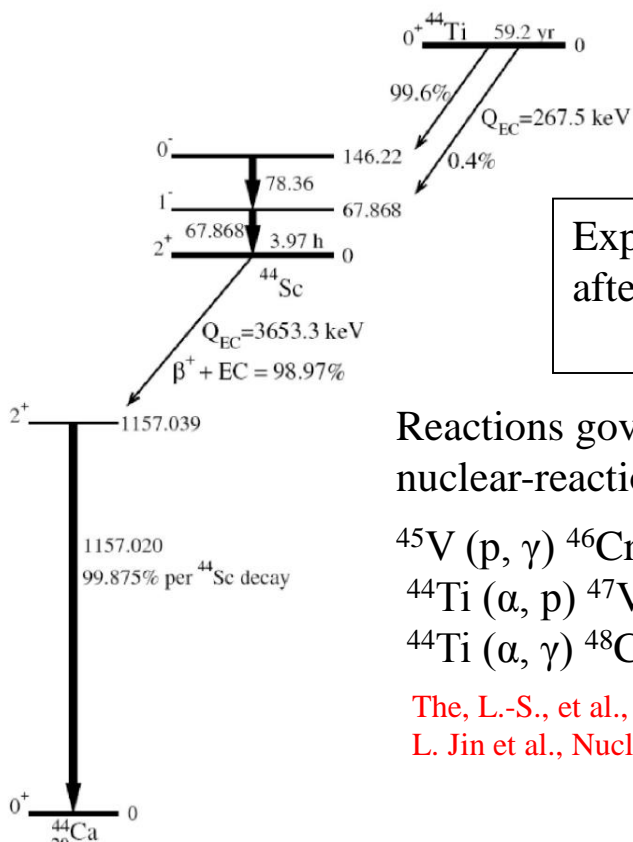
**Study of the beta delayed particle emission from
 ^{48}Mn and its relevance for explosive nucleosynthesis
(P-445)**

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Local contact: Miguel Madurga

Motivations

One of the open problems in galactic chemical evolution is the source of ^{44}Ca ; the dominant production channel is believed to be ^{44}Ti nucleosynthesis at core collapse supernovae.



The $^{44}\text{Ti} \rightarrow ^{44}\text{Sc} \rightarrow ^{44}\text{Ca}$ decay chain produces characteristic gamma lines that have been observed in young SNRs: 1157 keV from ^{44}Ca , and 78.4 keV, 67.9 keV from ^{44}Sc .

Expected synthesis of ^{44}Ti taking place at core collapse supernovae after alpha-rich freeze-out. **Main production reaction:**
 $^{40}\text{Ca} (\alpha, \gamma) ^{44}\text{Ti} \rightarrow$ Recent measurements @ (TRIUMF)

Reactions governing the abundance of ^{44}Ti have been investigated by using large nuclear-reaction network calculations.

- $^{45}\text{V} (p, \gamma) ^{46}\text{Cr}$
- $^{44}\text{Ti} (\alpha, p) ^{47}\text{V}$
- $^{44}\text{Ti} (\alpha, \gamma) ^{48}\text{Cr}$

$E \sim 3 - 7 \text{ MeV}$

The, L.-S., et al., A&A 450 (2006) 1037.
 L. Jin et al., Nuclear Physics A 621 (1997) 319c.

These reactions involve the use of low energy beam/targets of ^{45}V (T1/2 = 547(6) ms) and ^{44}Ti (T1/2 = 60.0(1) y) \rightarrow radioactive beam facilities.

Mn45	Mn46	Mn47	Mn48	Mn49	Mn50	Mn51
	41 ms	100 ms	158.1 ms	382 ms	283.88 ms	46.2 m
	ECp	ECp	ECp, EC	(α, γ)	EC	EC
Cr44	Cr45	Cr46	Cr47	Cr48	Cr49	Cr50
53 ms	50 ms	0.26 s	500 ms	21.56 h	42.3 m	1.8E+17 y
ECp	ECp	EC	EC	EC	EC	ECEC 4.345
V43	V44	V45	V46	V47	V48	V49
80 ms	547 ms	547 ms	422.3 ms	32.6 m	15.9735 d	3.30 d
(p, γ)	EC	EC	EC	EC	(α, p)	EC
Ti42	Ti43	Ti44	Ti45	Ti46	Ti47	Ti48
199 ms	509 ms	63 y	184.8 m	0+	5/2-	0+
EC	EC	EC	EC	EC	EC	EC
Ca40	Ca41	Ca42	Ca43	Ca44	Ca45	Ca46
0+	1.03E+5 y	0+	7/2-	0+	162.61 d	0+
96.941	EC	0.647	0.135	2.086	β^-	0.004

Previous studies

$^{45}\text{V} (p, \gamma) ^{46}\text{Cr}$

- dominates $(p, \gamma) - (\gamma, p)$ equilibrium with ^{45}V
- ^{45}V very difficult to produce (refractory)
- Dominated by very narrow proton resonances (\ll keV)
- Seems to be well described by statistical calculations

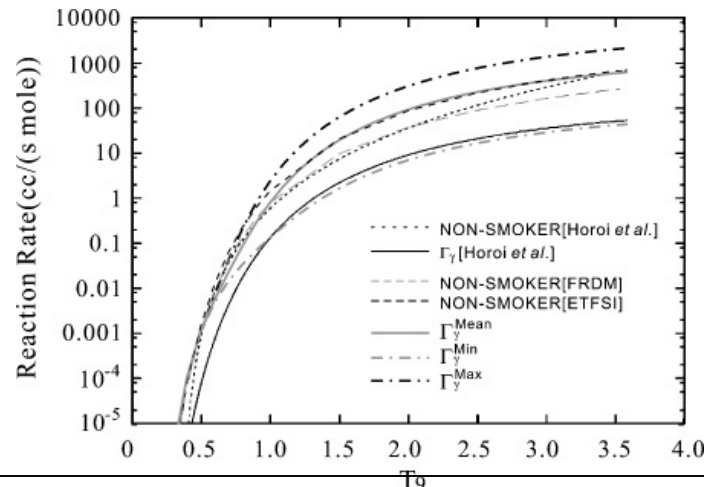
No experimental data so far

Difficult experiment:

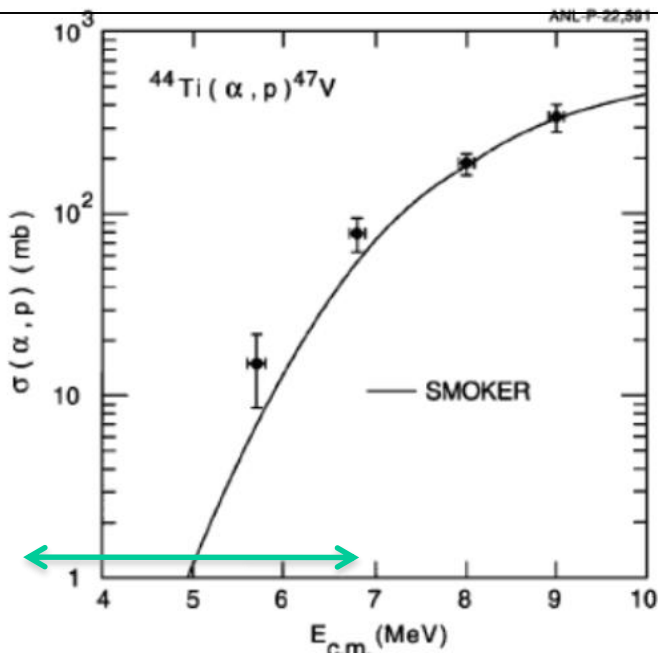
- ^{45}V not available (at present) at ISOLDE
- Proposal for production of ^{45}V at TRIUMF

Horoi, M. et al. Phys.Rev. C66 (2002) 015801

J. J. He* and A. St. J. Murphy, Phys. Rev. C75 (2007) 068801



FRDM: Finite Range Droplet Model
ETFSI: Extended Thomas-Fermi approach
with Strutinski Integral



A.A. Sonzogni et al. PRL 84 (2000)

$^{44}\text{Ti} (\alpha, p) ^{47}\text{V}$

- Dominant destruction of ^{44}Ti
- ^{44}Ti very difficult to produce (refractory)
- Previous measurements Argonne Nat. Lab. $\sim 5 \times 10^5$ pps on target
- Typical cross sections \sim few 100's μb

Difficult experiment:

- ^{44}Ti not available (at present) at ISOLDE
- Measurements using ^{44}Ti at ISODE from PSI Targets (A. St. J. Murphy et al., (2012))
- Dedicated setup using gas target & silicon detectors

$^{44}\text{Ti} (\alpha, \gamma) ^{48}\text{Cr}$

→ Destruction of ^{44}Ti

→ No experimental data so far

Difficult experiments!! → possibly at ISOLDE in near future?

Different approach → Beta delayed proton emission from ^{48}Mn

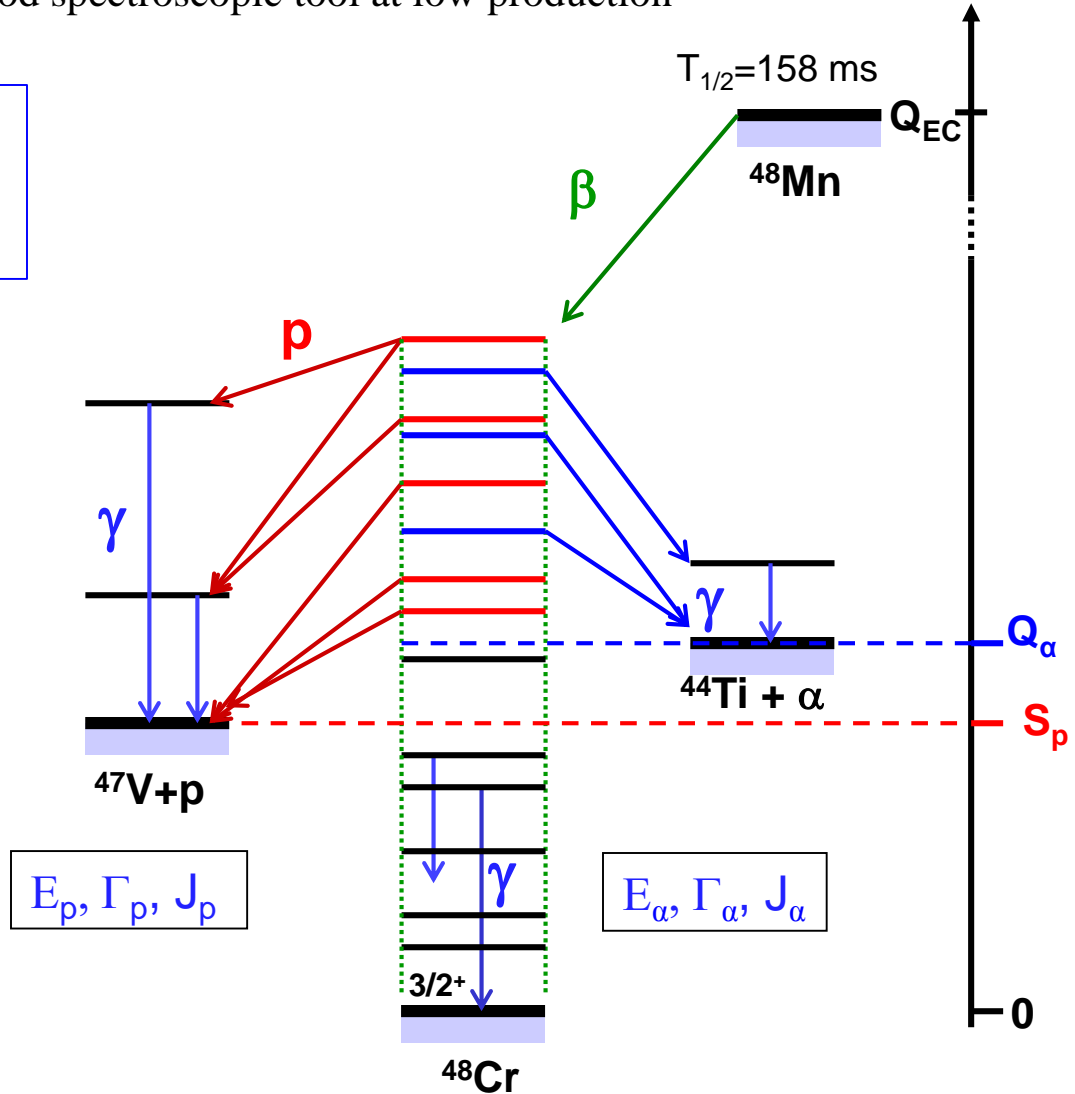
Beta delayed particle emission → a very good spectroscopic tool at low production rates (~few pps)!!

Experimental data:
 E, Γ, J, \dots
 β -decay properties

^{48}Mn
 $Q_{\beta^+} = 12480 \text{ keV}$

^{48}Cr
 $S_p = 8105 \text{ keV}$
 $Q_\alpha = 7698 \text{ keV}$

Mn45	Mn46 41 ms	Mn47 100 ms	Mn48 158.1 ms 4+	Mn49 382 ms 5/2-	Mn50 283.88 ms 0+	Mn51 46.2 m 5/2-
Cr44 53 ms 0+	Cr45 50 ms	Cr46 0.26 s 0+	Cr47 500 ms 3/2-	Cr48 21.56 h 0+	Cr49 42.3 m 5/2-	Cr50 1.8E+17 y 0+
V43 800 ms (7/2-)	V44 90 ms (2+)	V45 22.37 ms 7/2-	V46 32.6 ms 0+	V47 32.6 m 3/2-	V48 15.9735 d 4+	V49 330 d 7/2-
Ti42 199 ms 0+	Ti43 509 ms 7/2-	Ti44 63 y 0+	Ti45 184.8 m 7/2-	Ti46 0+	Ti47 5/2-	Ti48 0+
Sc41 596.3 ms 7/2-	Sc42 681.3 ms 1+	Sc43 3.891 h 7/2-	Sc44 3.927 h 2+	Sc45 7/2-	Sc46 83.79 d 4+	Sc47 3.3492 d 7/2-
Ca40 0+ 96.941	Ca41 1.03E+5 y 7/2-	Ca42 0+	Ca43 7/2-	Ca44 0+	Ca45 162.61 d 7/2-	Ca46 0+
		0.647	0.135	2.086	β	0.004



(E_p, Γ_p, J_p) & $(E_\alpha, \Gamma_\alpha, J_\alpha)$ are the critical parameters needed to evaluate the reactions rates of the reactions $^{44}\text{Ti}(\alpha, \gamma)^{48}\text{Cr}$ and $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$

The astrophysical S factors (example $\rightarrow (\alpha, p)$)

$$S(E) = \pi \frac{(2J+1)}{(2j_t+1)(2j_p+1)} \times \frac{\Gamma_\alpha \Gamma_p}{(E-E_r)^2 + \Gamma^2/4} \exp(2\pi\eta)$$

J - spin of the state in the compound nucleus

J_t - spin of the target nucleus

j_p - spin of the projectile

Γ - total decay width

Γ_α, Γ_p - partial decay widths

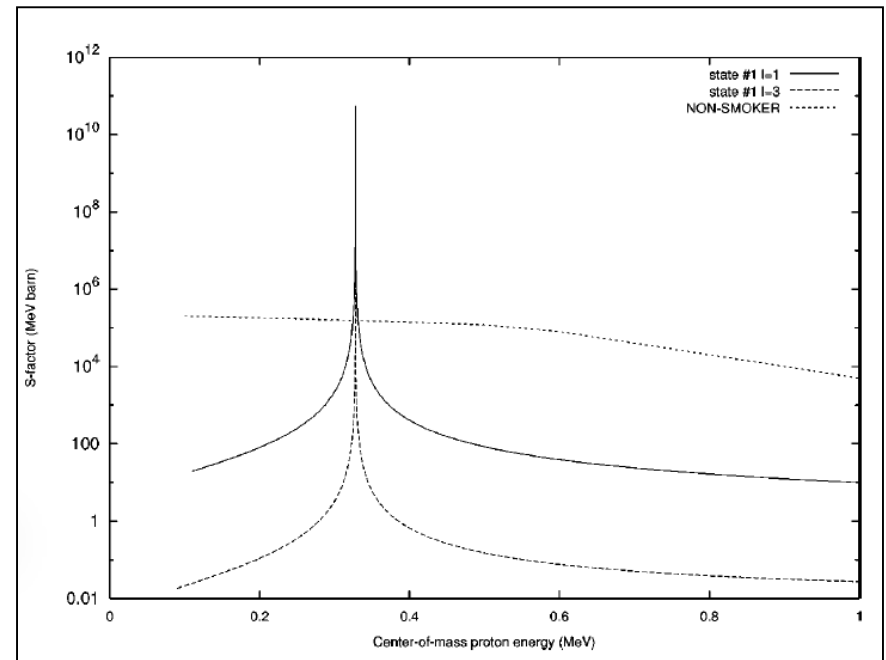
η - Sommerfeld parameter

E_r - level energy

Thermonuclear reaction rates

$$N_A \langle \sigma v \rangle = \left(\frac{8}{\pi \mu} \right)^{1/2} \frac{N_A}{(kT)^{3/2}} \int_0^\infty S(E) \exp\left[-\frac{E}{kT} - \frac{b}{E^{1/2}} \right] dE$$

Astrophysical S factor calculated using the single resonance formula as a function of proton center-of-mass energy.



Data analysis of particle decay to extract E, Γ , J

“Ideal analysis”: Shell model in the continuum (Gamow Shell Model - GSM)

→ Full shell model in the complex energy plane

N. Michel, W. Nazarewicz, M. Oloszajzak and T. Vertse, J. Phys. G.: Nucl. Part. Phys. 36 (2009) 013101// Humblet and Rosenfeld, Nucl. Phys. 26, 529 (1961); T. Berggren, Nucl. Phys. A 109 (1968) 265; R. de la Madrid, Nucl. Phys. A812, 13 (2008)

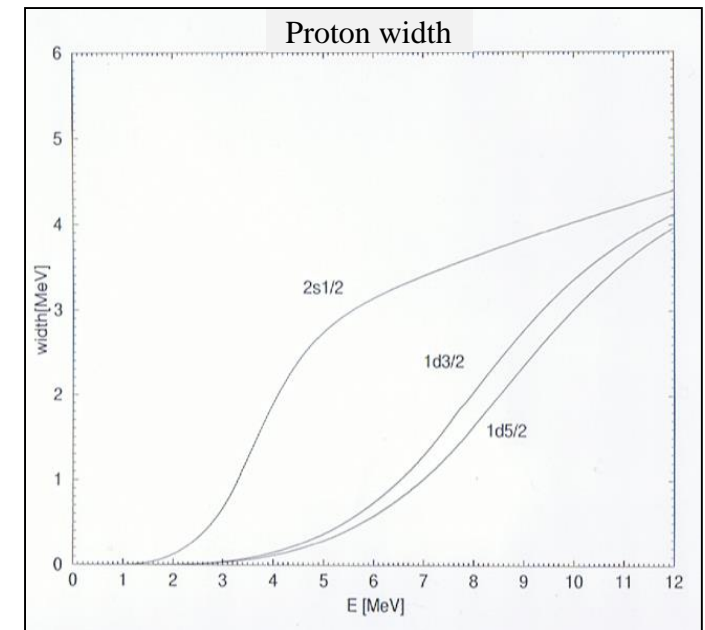
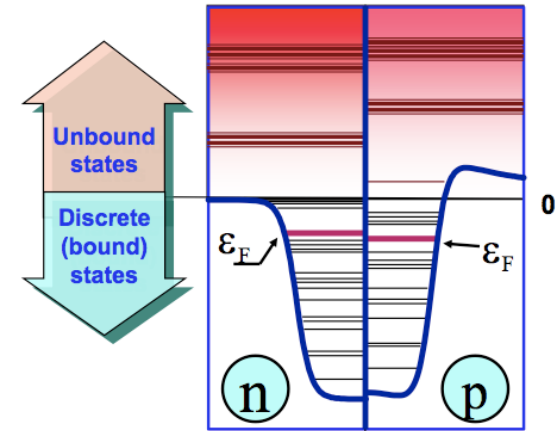
“Simple analysis”: Shell model WITH Gamow wave functions SM-G.

I. Martel et al., NPA(2001)424-436

- Shell model to calculate energy levels, spectroscopic factors, and beta decay strength
- Gamow-wave functions to evaluate particle decay widths
- Already tested in the analysis of ^{31}Ar beta delayed particle emission
- Collaboration with A. Poves (Madrid)

Total width (Γ)

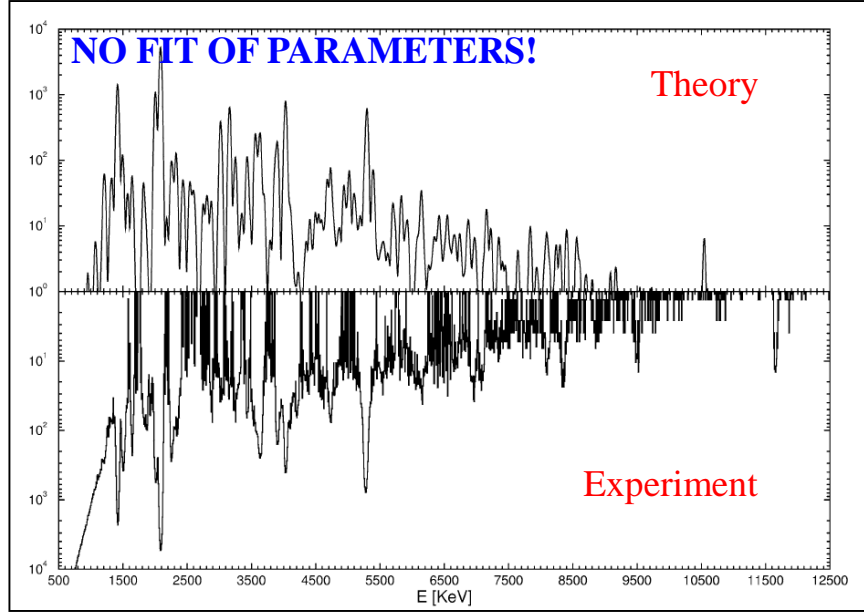
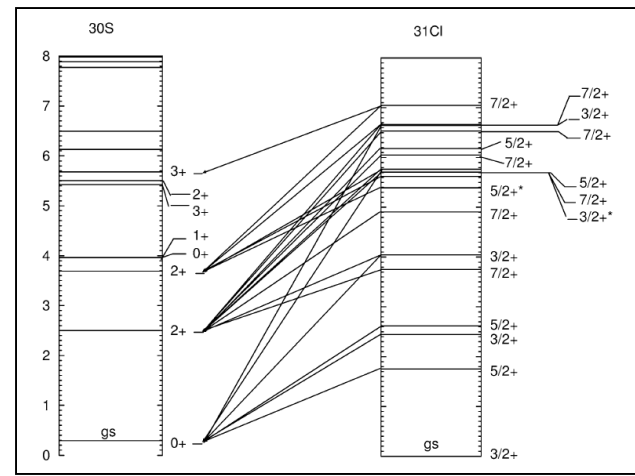
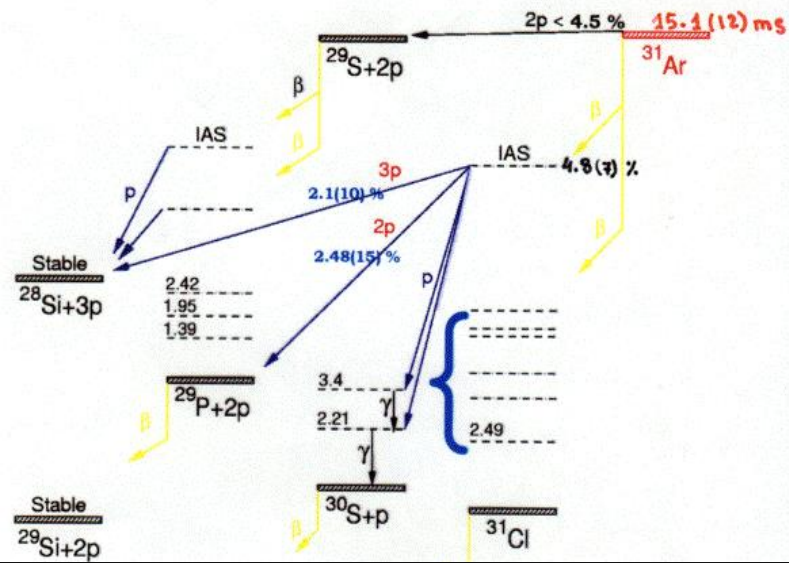
$$\Gamma(i; r; E_k) = \sum_{lj} |SPA(i; r, nlj)|^2 \gamma(nlj; E_k)$$



Gamow state calculation~ Woods-Saxon ($a=0.65\text{fm}$, $r=1.27\text{fm}$), select depth V_0 to reproduce E_k , ^{31}Ar β -delayed proton emission

Shell model WITH Gamow wave functions SM-G

I. Martel et al., NPA(2001)424-436



Results of the analysis of the beta-delayed proton energy spectrum

J_{Cl}^{π}	E_{Cl}	$B(GT)$ (th)	J_S^{π}	E_p (exp)	I_p (th)	I_p (exp)	Γ (th)
$5/2_1^+$	1753	0.053	0_1^+	1416(2)	27.1	34.0(3)	0.02
$3/2_2^+$	2443	0.243	0_1^+	2084(2)	100	100.0(6)	0.4
$5/2_2^+$	2618	0.004	0_1^+	2253(2)	1.6	4.0(3)	0.8
$7/2_1^+$	3752	0.004	2_1^+	1211(4)	1.1	1.7(5)	0.11
$3/2_3^+$	4045	0.028	0^+	3634(3)	4.4	6.1(8)	9.8
			2_1^+	1504(2)	2.3	6.2(2)	5.1
$7/2_2^+$	4905	0.014	2_1^+	2327(4)	2.4	5.1(4)	1.7
$5/2_3^+$	5390	0.011	2_2^+	1643(2)	1.1	2.88(14)	3.4
$7/2_3^+$	5621	0.057	2_1^+	3020(3)	7.4	1.08(14)	3.1
$3/2_4^+$	5658	0.010	1_1^+	1643(2)	0.8	2.88(14)	3.8
$5/2_4^+$	5767	0.360	0_1^+	5276(5)	11.5	17.6(3)	7.4
			2_1^+	3153(4)	12.2	0.44(10)	7.8
			2_2^+	2008(2)	20.9	10.0(2)	13.4
$7/2_4^+$	6047	0.022	2_1^+	3432(3)	2.1	0.89(11)	9.8
$5/2_5^+$	6180	0.047	2_1^+	3561(11)	4.5	3.6(8)	30.8
$7/2_5^+$	6533	0.044	2_1^+	3902(3)	3.4	2.22(14)	11.3
$3/2_5^+$	6640	0.023	0_1^+	6145(7)	0.5	0.51(12)	5.4
$7/2_6^+$	6665	0.186	2_1^+	4030(3)	14.7	7.0(2)	4.6
			2_2^+	2881(3)	0.4	0.99(13)	0.13
$7/2_7^+$	7050	0.050	2_2^+	3249(4)	1.9	1.17(15)	2.6
			3_2^+	1300(13)	0.9	0.7(11)	1.3

Previous measurements on beta delayed proton emission from ^{48}Mn

GSI (Darmstadt, Germany)

T. Sekine et al., Nucl. Phys. A467 (1987) 93

J. Szerypo et al., Nucl. Phys. A528 (1991) 203

UNILAC + GSI on-line mass separator

$^{12}\text{C}(^{40}\text{Ca}, p3n)$ reaction

Beam time:

50 hours,, I=200-600 pps

Contaminants:

48Cr, 48V, 48Sc and 12F (Al-F)

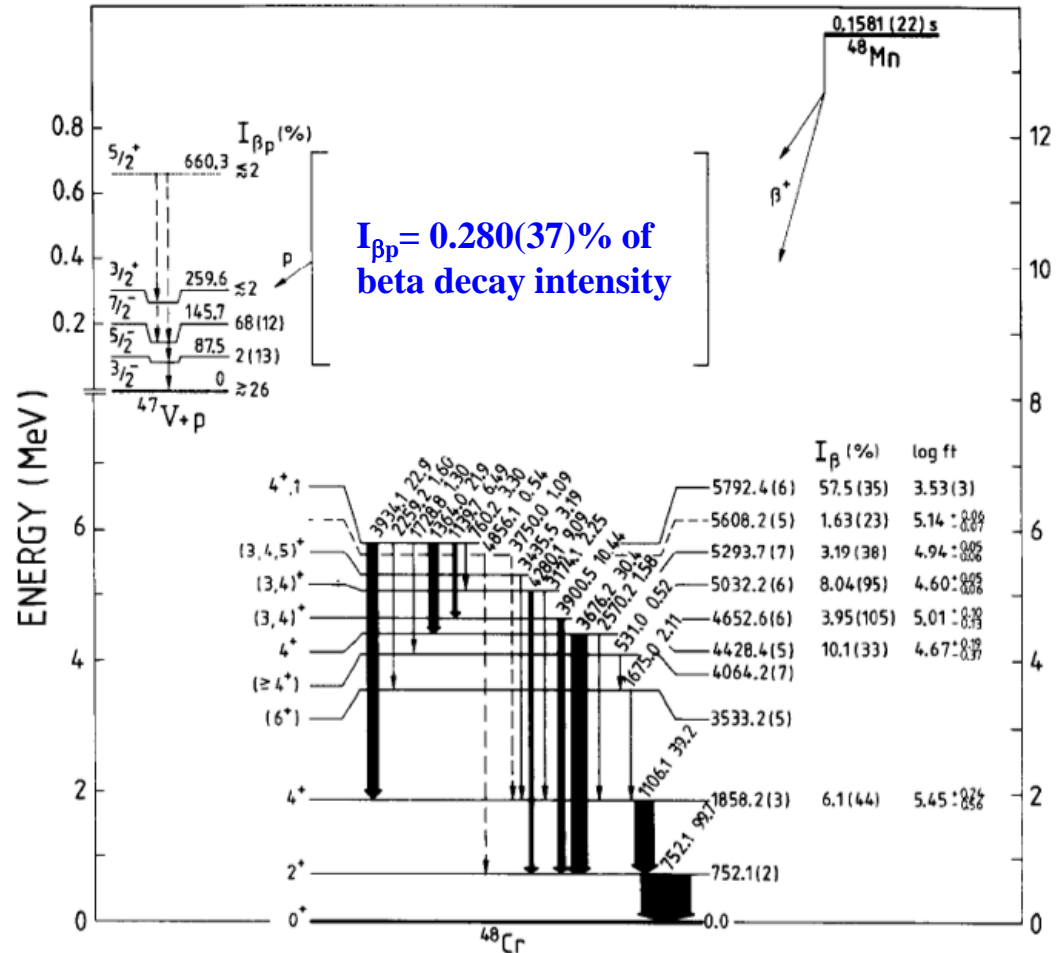
Set-up:

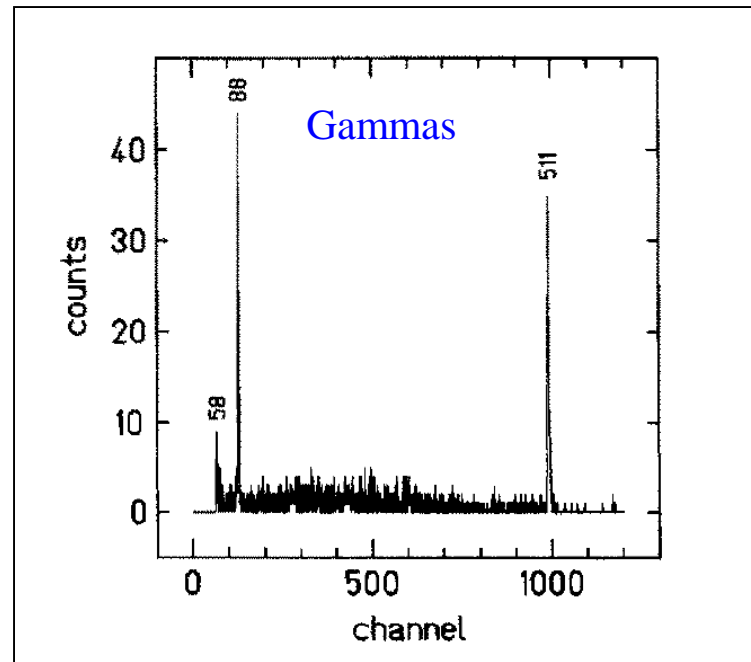
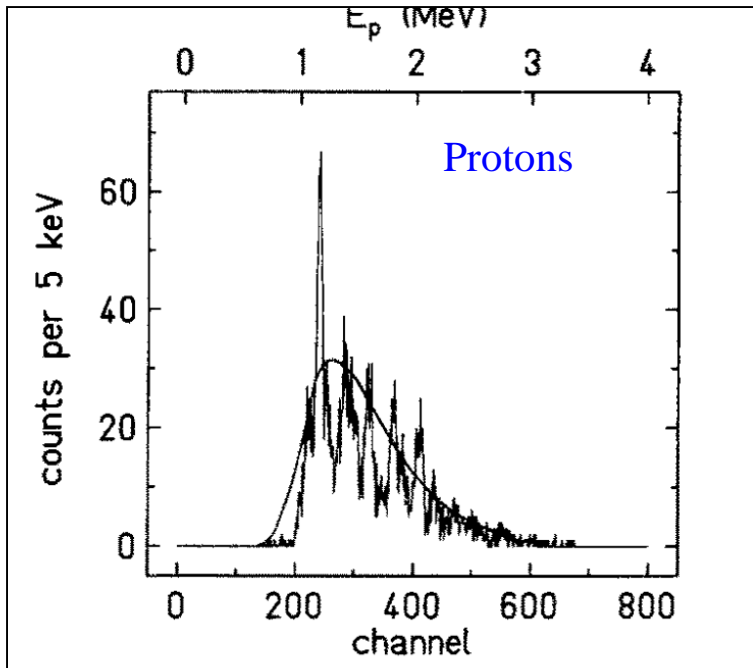
2 x particle telescopes

- Efficiency of 4%
- $E_{\text{res}} = 50 \text{ keV}$

2 x gamma detectors:

- Efficiency of 4% (1.3 MeV)





Protons: expected bell-shaped overall structure

$E_p \sim 1-3$ MeV range \rightarrow Structures @ 1.2, 1.4, 1.6, 1.8, 2.0, 2.2 MeV

Coincidence with gamma rays:

- $\sim 2\% \rightarrow ^{47}\text{V}(660.3 \text{ keV})$
- $\sim 2\% \rightarrow ^{47}\text{V}(259.6 \text{ keV})$
- 68(12)% $\rightarrow ^{47}\text{V}(145.7 \text{ keV})$
- 2(13)% $\rightarrow ^{47}\text{V}(87.5 \text{ keV})$
- $\sim 26\% \rightarrow ^{47}\text{V}(\text{gs})$

$$I_{\beta\alpha} < 6.0 \times 10^{-4} \%$$

Sekine et al. \rightarrow singles: 1 event!

However, due to statistics, it was not possible to identify the initial and final levels in ^{48}Cr involved in the process. \rightarrow Main objective of the present proposal.

Main objectives of present proposal

Detailed study of ^{48}Mn beta delayed proton emission:

- Identify proton emitter levels in ^{48}Cr
- Identify alpha emitter levels (branching ratio limits)
- Extract proton widths, angular momentum (alpha width limits)

→ Cross section evaluation

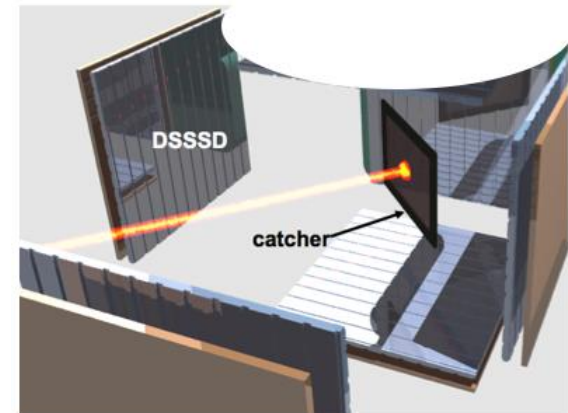
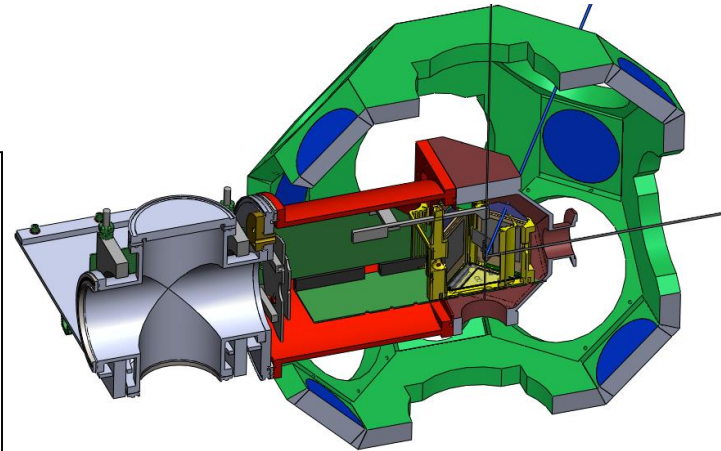
Dedicated experimental setup at the IDS

“Silicon-cube” device

- 5 **DSSSD** telescopes for protons/alphas → high granularity
 - default detection efficiency ~46%
 - energy FWHM ~ 25 keV (low noise P.A.)
 - 40um and 1mm thickness
 - back-detectors for beta suppression
 - ions deposited on a thin **carbon catcher**

Germanium detectors

- 4 Clover detectors
 - high resolution
 - efficiency ~4% (1 MeV)
- closer geometry possible: particle/gamma (60% / 10%)



Beam-time request

ISOLDE ^{48}Mn production (Nb target, 1.4 GeV) ~ 10 pps on target (Isolde Yields Data Base)

Beam request: 21 shifts

Assuming 46% efficiency (particles) & 4% efficiency (gammas)

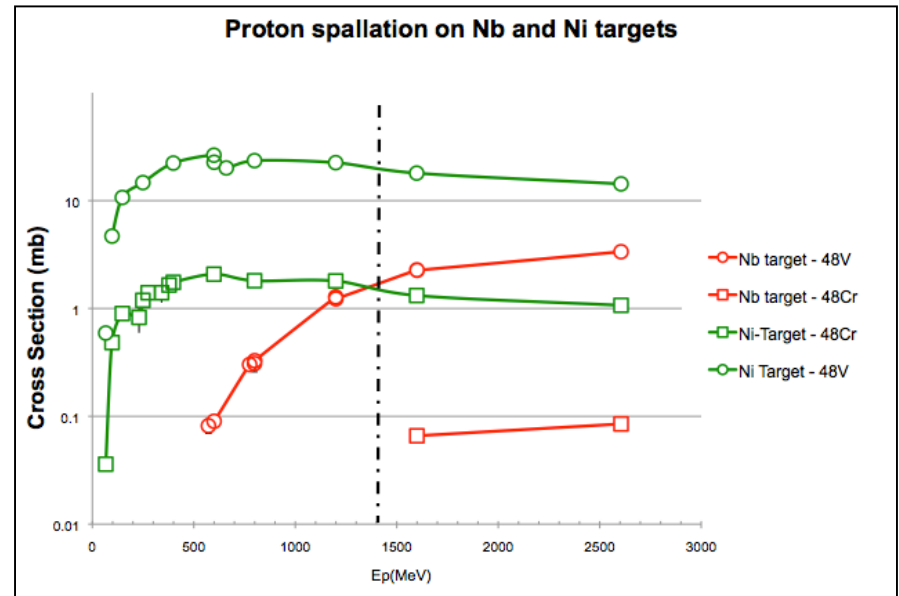
Expected beta delayed proton yield: 8300

Expected proton-gamma coincidences: 330

Remark: For doing the proposed measurements, it would be very interesting to investigate a target test with a mixed Ni-Al composition.

Estimated gain in Yield by $\sim 10 - 20$

EXFOR: Experimental Nuclear Reaction Data.
<https://www-nds.iaea.org/exfor/>



^{48}Mn collaboration

CERN – Huelva (Spain) – Madrid (Spain) – Warsaw (Poland) – Belfast (UK) –
Leuven (Belgium) – Gradignan (France) – Lund (Sweden) – Mexico (Mexico) –
Warsaw (Poland) – Aarhus (Denmark)

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Thanks for your attention!!

^{35}Ca decay beta-delayed 1- and 2-proton

spokespersons: J. Giovinazzo (CENBG), O. Tengblad (CSIC)

Germanium detectors

- **efficiency ~10-15%** (at 1 MeV)
close geometry (~8 cm)

“Silicon-cube” device

- 6 **DSSSD** for protons
- detection efficiency ~60% for 1 proton
- energy FWHM ~ 25 keV (low noise P.A.)

