

The background of the slide features a complex network of blue lines and dots, resembling a particle detector or a data visualization. The dots vary in size and brightness, and the lines connect them in a web-like pattern. The overall color scheme is a deep blue.

# HEPHY

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## The Kalman Alignment Algorithm

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# Outline

- Motivation
- The Kalman Alignment Algorithm (KAA)
  - basic functionality and formulae
  - application to big systems: restricted updates
- Alignment studies
  - Monte Carlo studies with the full CMS Tracker
  - cosmic ray studies at the Tracker Integration Facility
- Current status and outlook

## Motivation

- When people started thinking about aligning the CMS Tracker it became obvious that this was not a trivial task.
- People realized that conventional approaches might face serious problems:
  - computing times
  - memory
  - stability
- The KAA was specifically designed to circumvent these problems.
- Starting from a basic idea, overcoming problems arising from conceptual details, the algorithm is now working properly and is well understood.

## Basic functionality

- The particle tracks are processed one-by-one using a specialized Kalman filter:
  - use a global track-model that parameterizes all measurements
  - track-model depends not only on the (ideal) track parameters but also the alignment constants
  - estimate and update alignment parameters and their full variance-covariance matrix at every step (Kalman filter update)
- Method accounts for all statistical correlations between the individual measurements (hits) due to multiple scattering
  - off-diagonal terms in the variance-covariance matrix of the measurements
- Method accounts for the geometrical (statistical) correlations between all detector modules
  - update is not restricted to detector modules hit by current track

## Basic formulae

- The vector of all measurements  $m$  depends via the track-model  $f$  on the track parameters  $q$  and the alignment constants  $p$ :

$$\vec{m} = \vec{f}(\vec{q}, \vec{p}) + \vec{\varepsilon}, \quad \text{cov}(\vec{\varepsilon}) = V$$

- The intrinsic detector resolution and multiple-scattering effects are accounted for in  $V$ , energy-loss effects (for high momentum muons) can be included into  $f$ .
- Using the Kalman filter formalism, update equations at step  $k$  for the estimate on the alignment parameters  $p_{k+1}$  can be derived:

$$\vec{p}_{k+1} = \vec{p}_k + K \left( \vec{m}_k - D_p \vec{p}_k - D_q \vec{q}_k + \vec{c} \right)$$

$$D_p = \partial \vec{f} / \partial \vec{p}, \quad D_q = \partial \vec{f} / \partial \vec{q}$$

- The matrix  $K$  is the *gain matrix* of the Kalman filter.
- Formula for the variance-covariance matrix of  $p_{k+1}$  follows from error propagation.
- By design, no inversions of large matrices required!

## Introduce new concept: restricted updates

- Problems for large systems (like the CMS Tracker):
  - Non negligible IO overhead due to reading and writing of the alignment parameters and their full variance-covariance matrix at every step.
  - Potentially large amount of virtual memory needed for storing the geometrical (statistical) correlations
- Proposed solution:
  - Update only those alignment parameters at each step that are significantly correlated
- Difficulties:
  - Need a-priori knowledge of which alignables should be included
  - Incomplete update can corrupt the full variance-covariance matrix, i.e. it might not be positive definite anymore



## Making the concept work ...

- No general recipe available as how to deal with these two difficulties
- However, in the case of a track-based alignment it is possible to adapt the algorithm such that these issues pose no problem
  - by bookkeeping which alignables were hit by which tracks, one can make a good guess which alignables are significantly correlated
  - storing the statistical correlations instead of the covariance entries avoids inconsistencies
  - some more minor *tricks* ...
- Needs tuning of a few *internal parameters* for a given geometrical setup
  - trade-off between precision and used memory and computing time

## Monte Carlo Alignment Studies

- Monte Carlo studies have been performed to prove that the algorithm is able to align the full CMS Tracker
- Realistic startup scenario:
  - realistic start-up conditions (initial misalignment, calibration, etc.)
  - skimmed data-samples from the *CMS Computing, Software and Analysis Challenge 2008 (CSA08)*
  - using the resources from the *CERN Analysis Facility (CAF)*
    - limited amount of RAM (2GB per core)
    - used max. 10 cores parallel (*fair use*)
  - automated using a specialized production system (modified version of the *CMS MillePede Production System*)
- Results were computed within reasonable amounts of time and look good



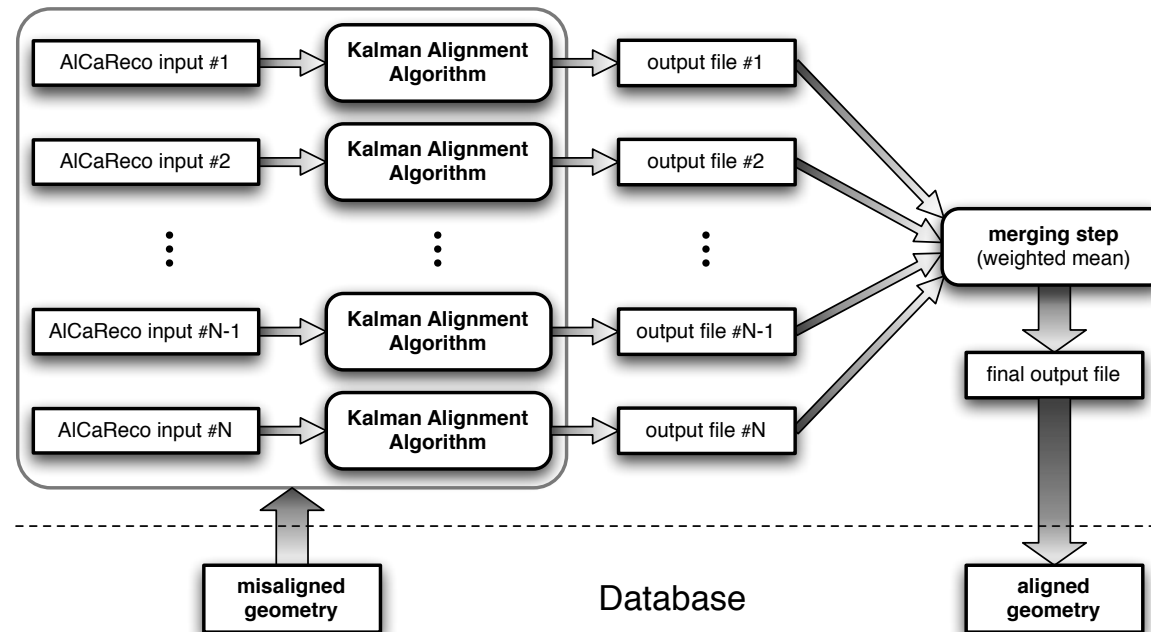
## Alignment Strategy

- Due to the constraint on virtual memory the CMS Tracker was aligned in three steps:
  - start with the outer silicon micro-strip tracker (*Tracker Outer Barrel, Tracker Endcaps*)
  - continue with the inner silicon micro-strip tracker (*Tracker Inner Barrel, Tracker Inner Disks*)
  - finish with the pixel tracker
- The first two steps were done using collision data and cosmic ray data
- The final step was done using additional  $Z \rightarrow \mu^+ \mu^-$  events
- Tuning of *internal parameters* was based on constraints regarding memory and computing time only
  - no feedback from comparison between the results and the MC-truth (which will obviously not be possible with real data)

# Workflow for a single alignment step

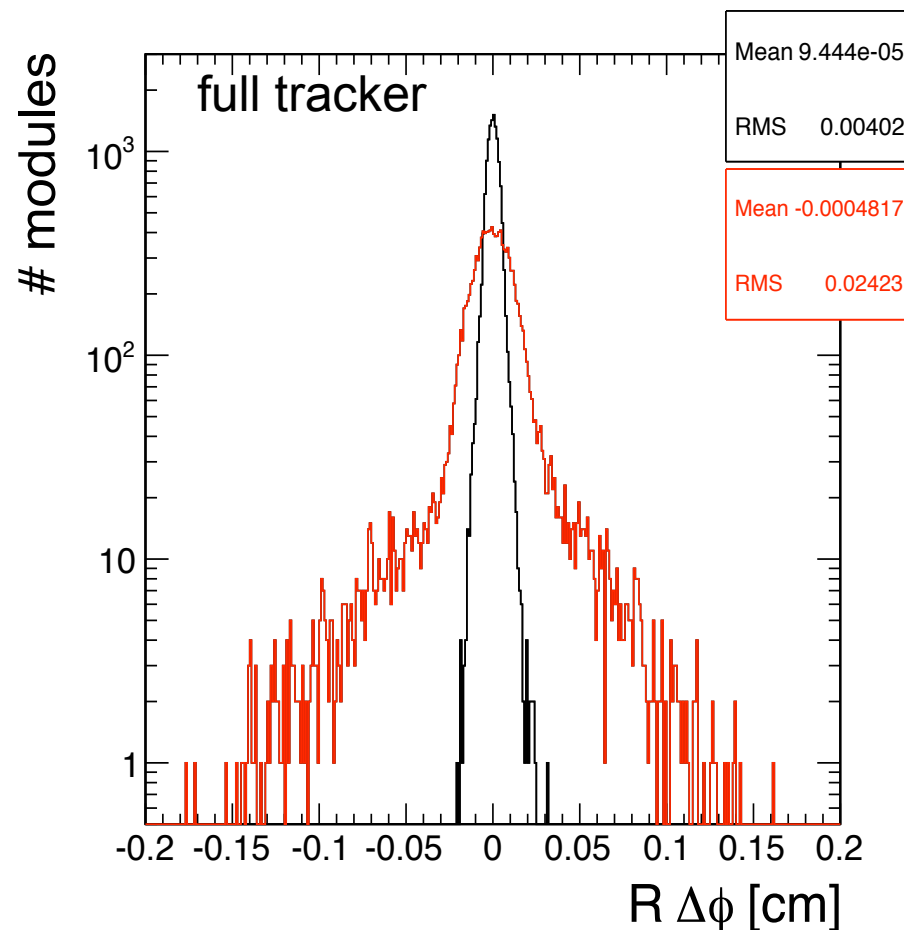
**Step 1: alignment runs on subsamples of the full data**

**Step 2: merge output of all alignment runs**



- KAA is completely implemented within the reconstruction and analysis software framework of CMS (*CMSSW*)
- Use CMS-specific data-skims for alignment and calibration tasks (*AlCaReco* format)

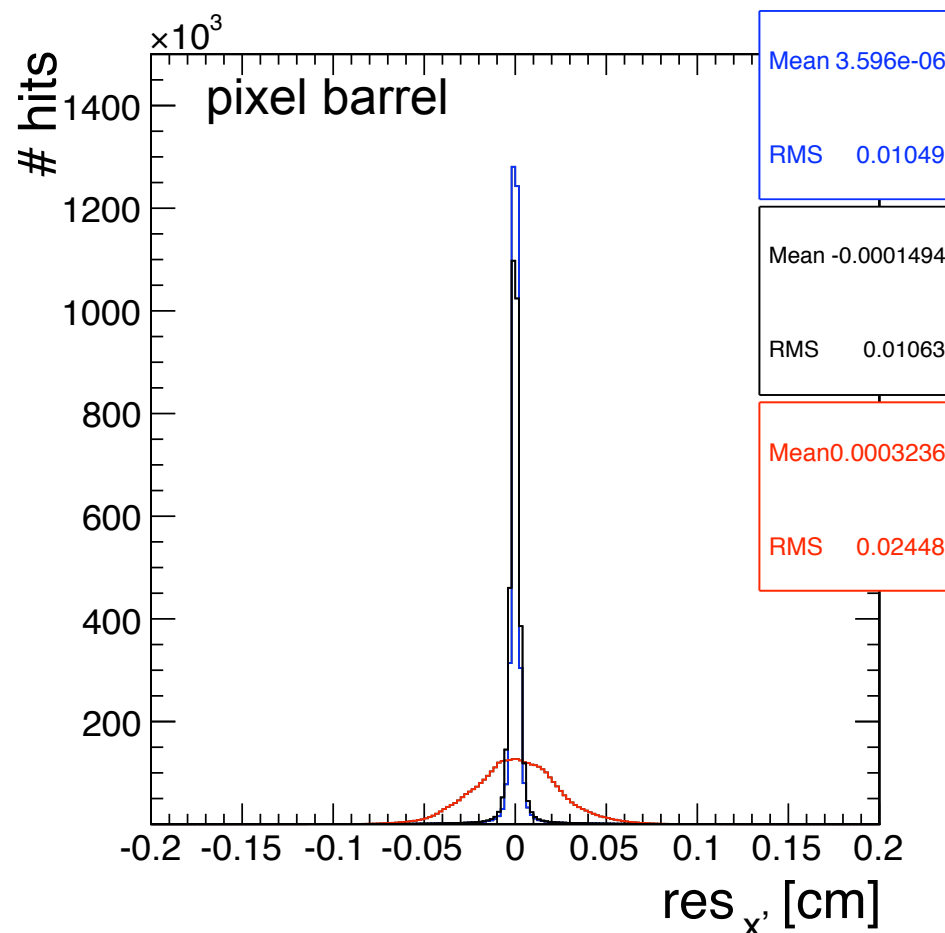
## Alignment with CMS Monte Carlo collision data (corresponding to the first 1 pb<sup>-1</sup>)



Red: misaligned geometry (startup)  
Black: aligned geometry (1 pb<sup>-1</sup>)

Overall precision of all detector modules in global  $R\Delta\phi$  is 40  $\mu\text{m}$ !

## Alignment with CMS Monte Carlo collision data (corresponding to the first 1 pb<sup>-1</sup>)



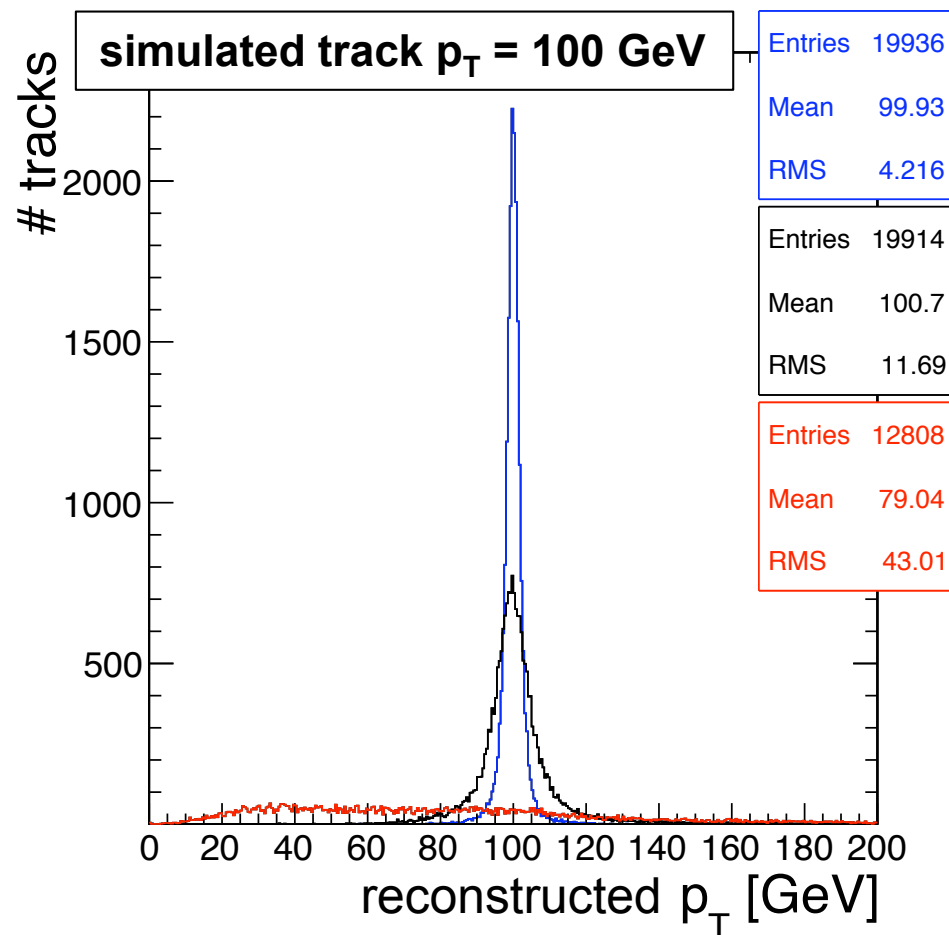
Red: misaligned geometry (startup)

Black: aligned geometry (1 pb<sup>-1</sup>)

Blue: ideal geometry

Distinct improvement in the tracking residuals.

## Impact on track reconstruction



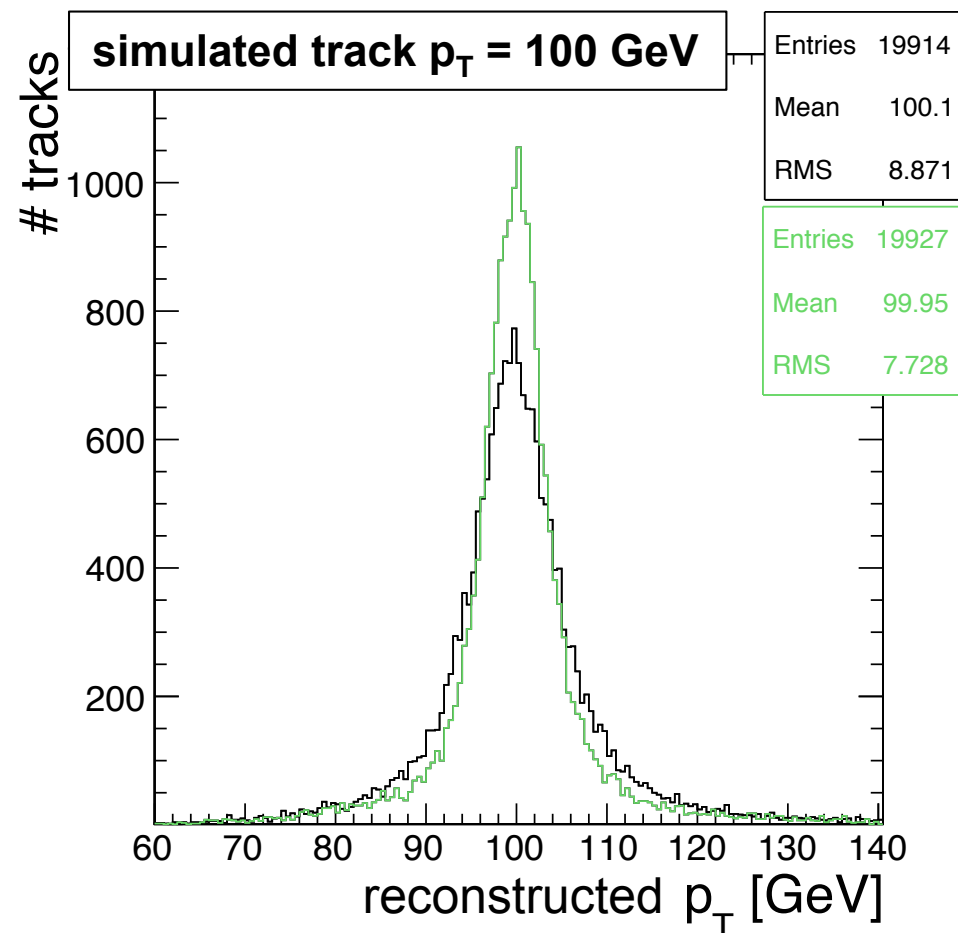
Red: misaligned geometry (startup)

Black: aligned geometry ( $1 \text{ pb}^{-1}$ )

Blue: ideal geometry

Recovers the reconstruction performance to a great extent.

## Alignment with CMS Monte Carlo collision data (corresponding to the first 10 pb<sup>-1</sup>)



Black: aligned geometry (1 pb<sup>-1</sup>)

Green: aligned geometry (10 pb<sup>-1</sup>)

Processing more high momentum tracks gives further improvements.



## Tracker Integration Facility

- The final assembly of the CMS silicon micro-strip tracker took place at a dedicated facility, referred to as the *Tracker Integration Facility*.
- The tracker has been operated at different temperatures, powering and reading out about 15% of all modules.
- Cosmic ray data has been recorded and used for track-based alignment.
- No magnetic field present, therefore no momentum estimate for individual tracks (and hence for effects due to multiple scattering)
- The results for the three different alignment algorithms used for the CMS tracker have been compared (HIP, MillePede, KAA).
- The quality assessment of the alignment objects was done by analyzing tracking residuals.

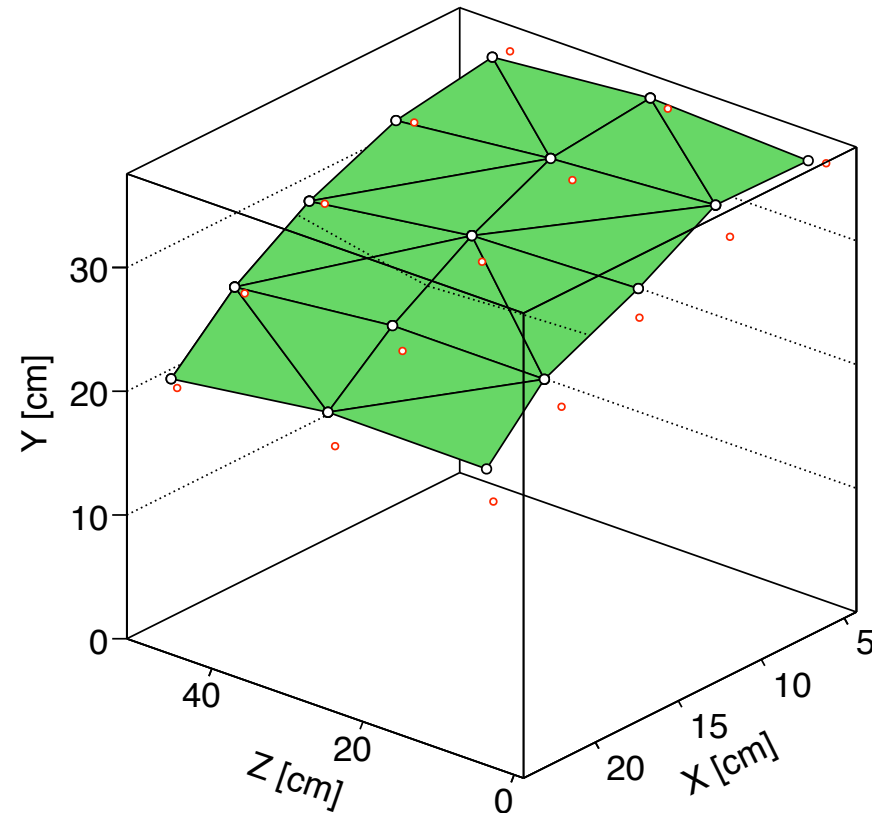
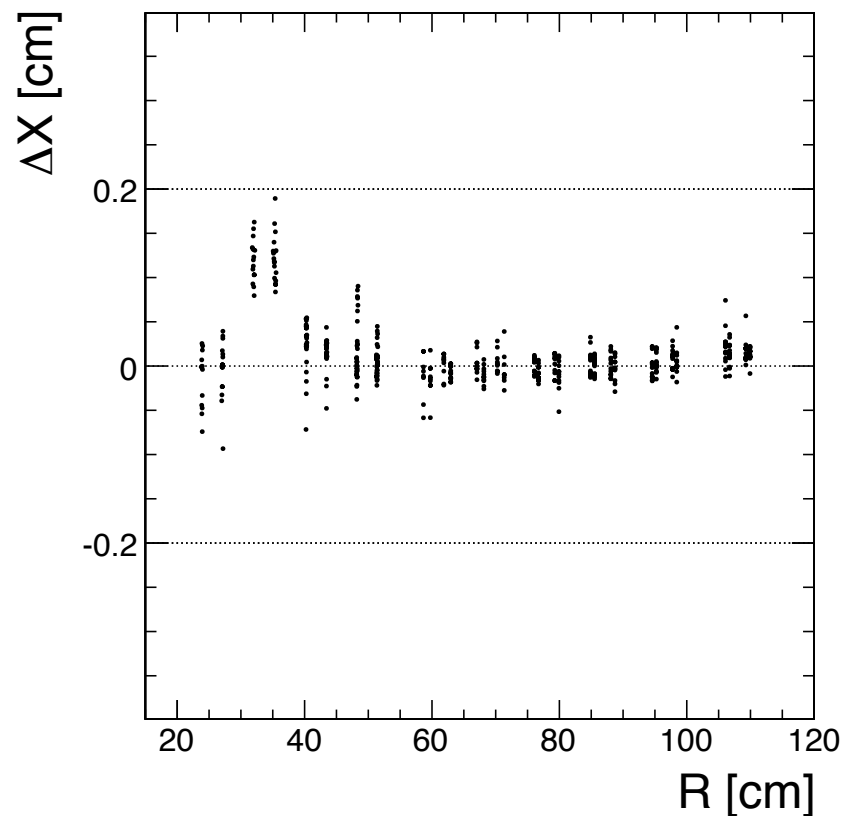
## TIF Setup



## Alignment Strategy

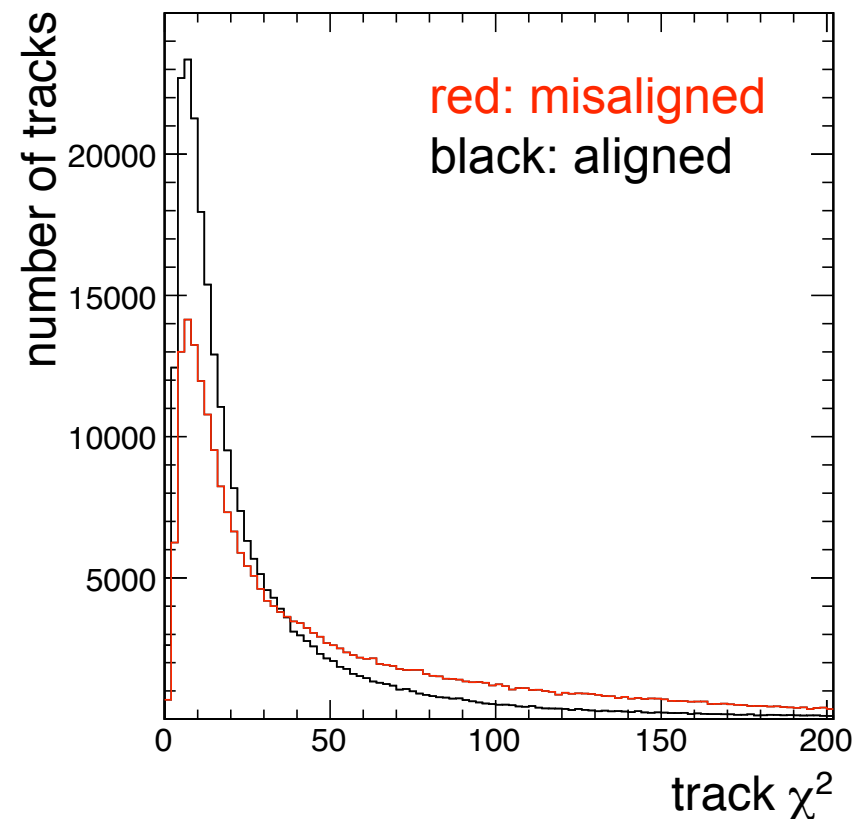
- A set of 430 detector modules in the barrel region has been chosen for alignment
  - relatively small number, hence the concept of *restricted updates* was not applied
  - same set for all three algorithms (comparability)
- Biggest data sample recorded at a single temperature has been processed
  - rules out possible effects due to different thermal expansion at different temperatures
  - approximately 70.000 well isolated cosmic particle tracks
  - same data sample for all three algorithms (comparability)

## Computed constants



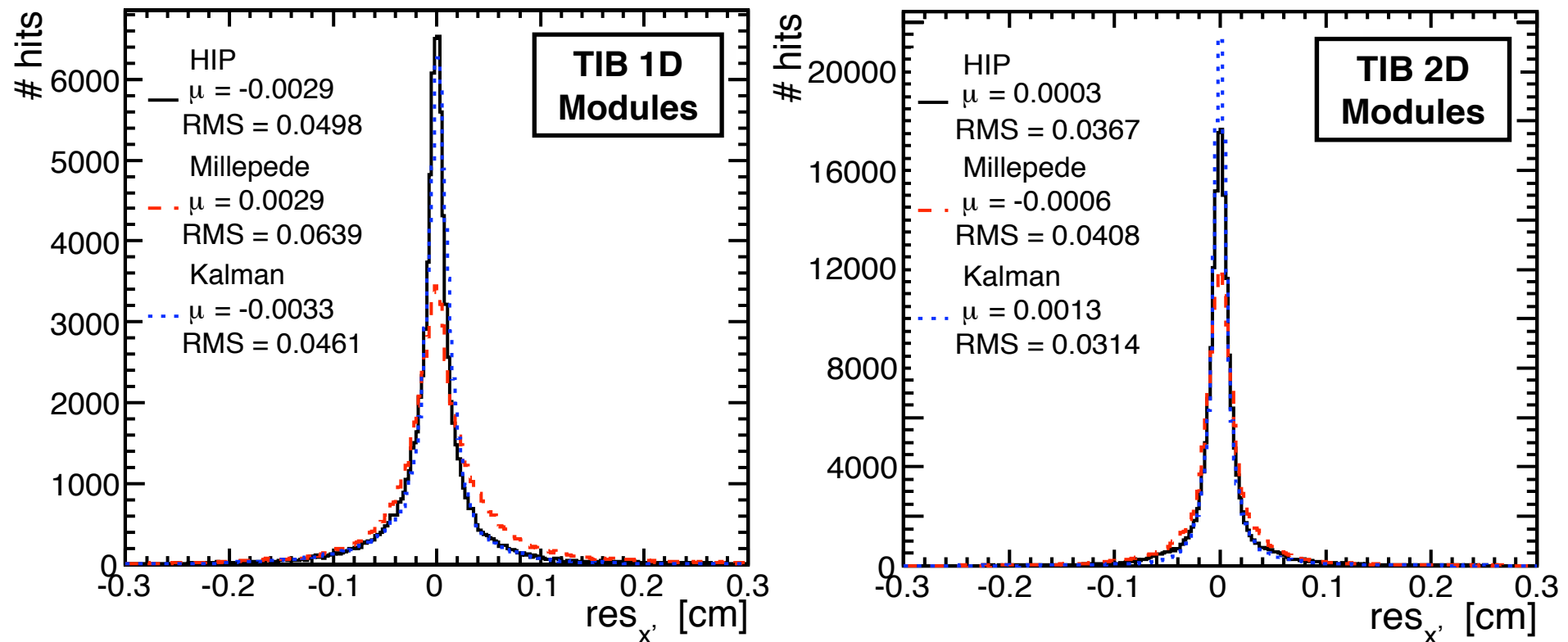
- Observed large deviations from the ideal geometry
- Due to deformations of the mechanical support structure on a larger scale

## Improvement of the tracking residuals



- Improved distribution of the track  $\chi^2$ 
  - sum of the squared residuals, each normalized by the detector resolution (no uncertainties due to misalignment included)

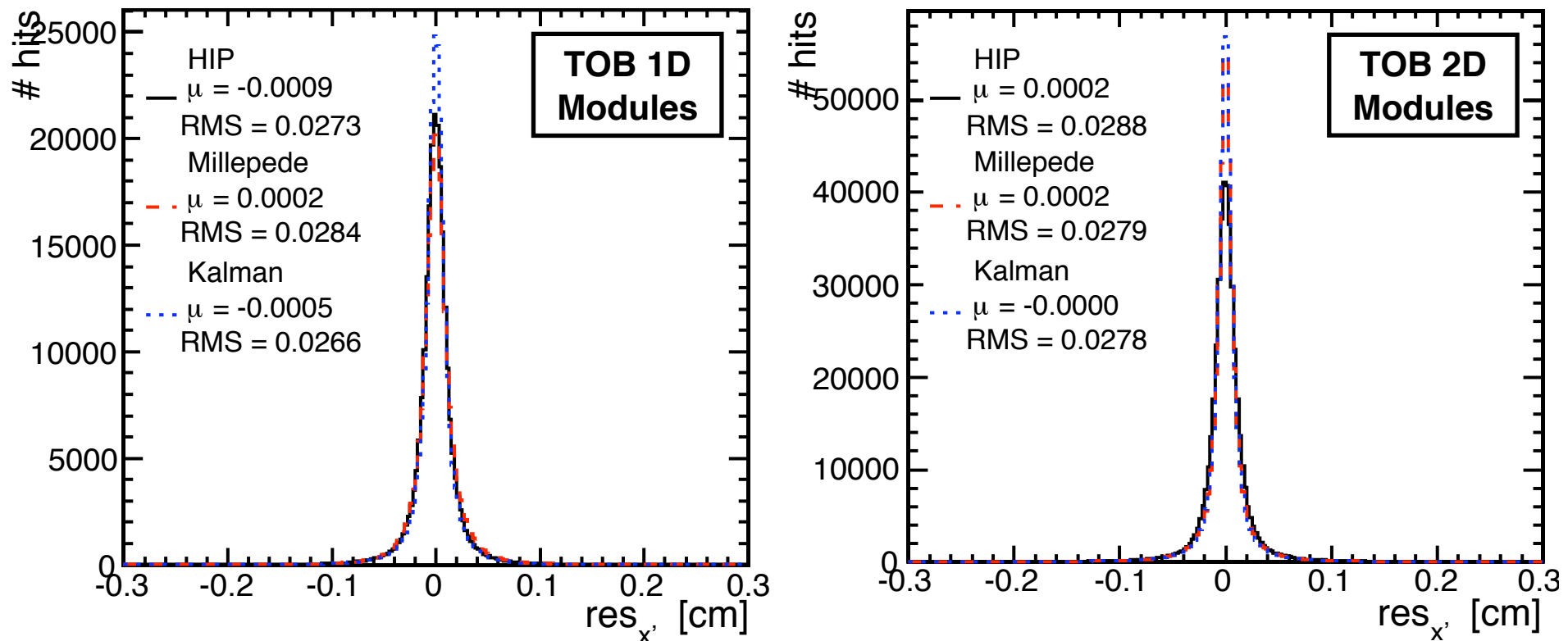
## Comparison of the tracking residuals (Tracker Inner Barrel)



Results of the three algorithms in good agreement



## Comparison of the tracking residuals (Tracker Outer Barrel)



Results of the three algorithms in good agreement

## Current status and outlook

- The KAA is a functional and well understood method.
- Results from detailed simulation studies show that the KAA is able to align the CMS Tracker – even under conditions expected during the LHC start-up phase.
- The associated computational effort can be kept at a reasonable level.
- For production, a dedicated production system is available.
- An analysis of the first experimental data from cosmic particle tracks, recorded at the Tracker Integration Facility, shows that the KAA is competitive to existing algorithms.
- The KAA will be used to analyze collision data in the future.