



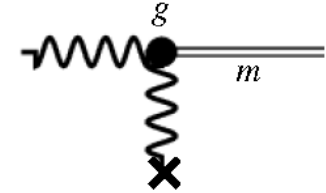
# Ultra-light Particles beyond the Standard Model:

## Laboratory Experiments



# A Landscape

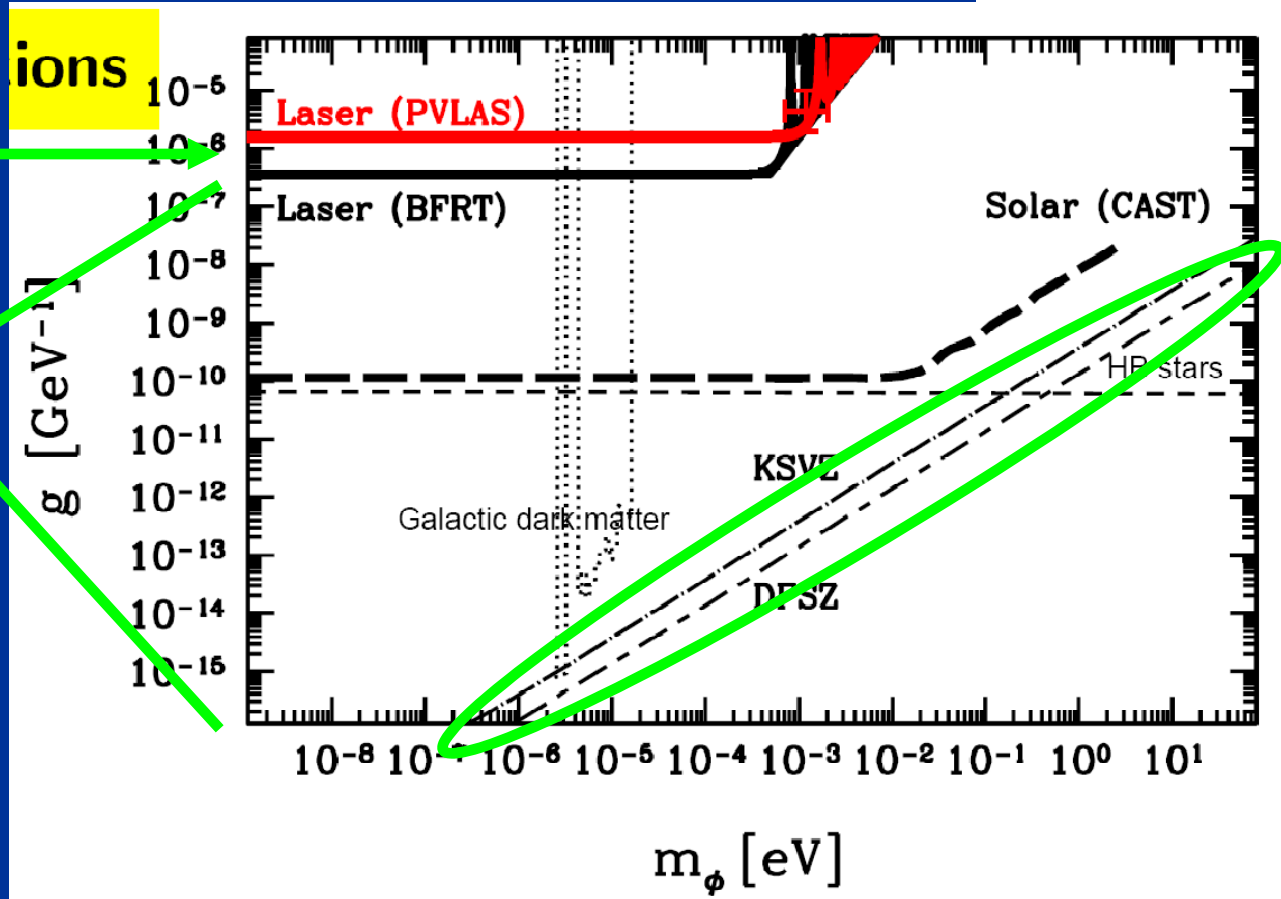
related to scalar and pseudoscalar WISPs



today's lab. limits

terra incognita

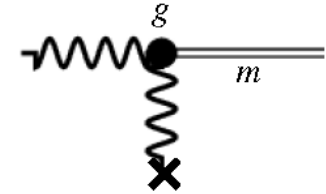
a promised continent:  
the QCD axion





# A Landscape

related to scalar and pseudoscalar WISPs

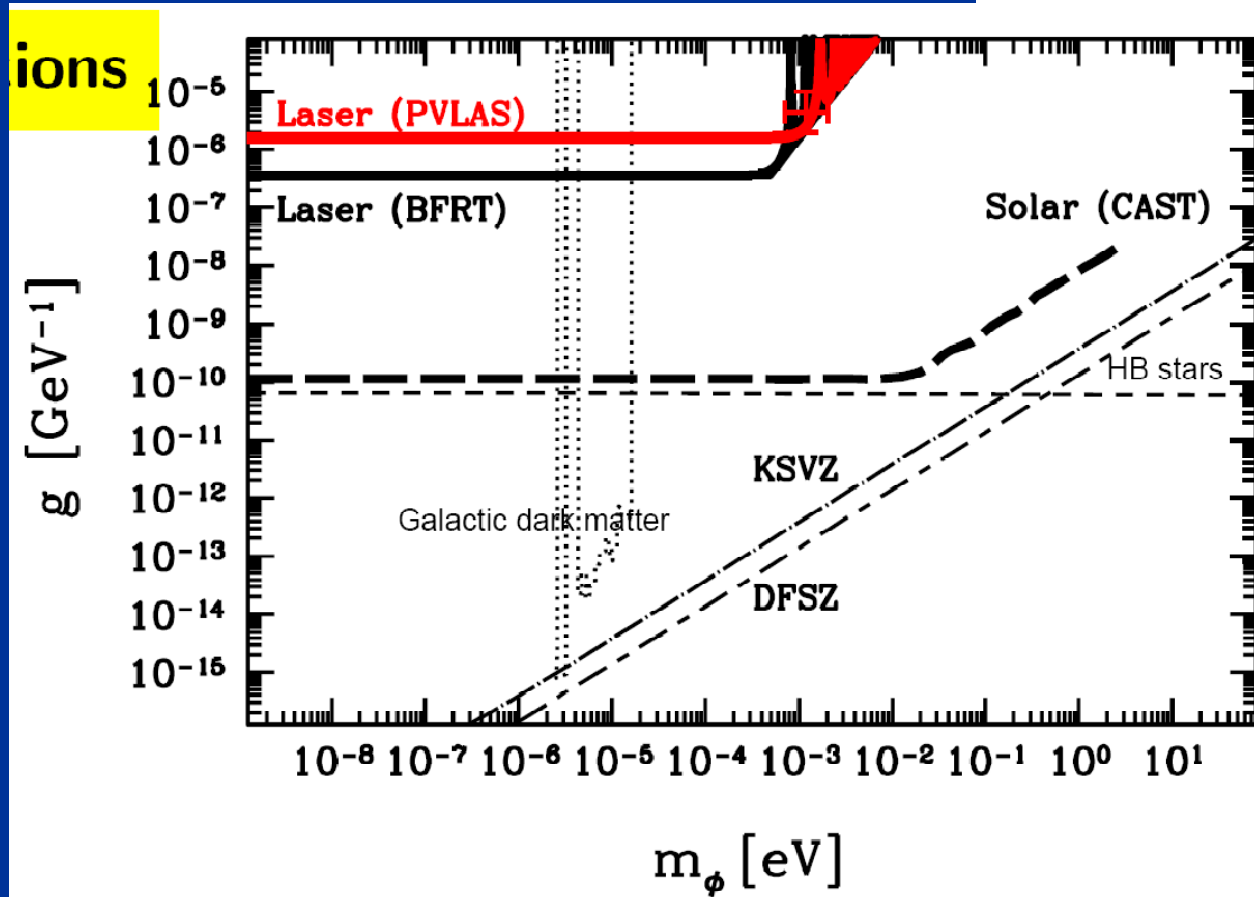


Scales of new physics

TeV

PeV

How to explore this landscape?





# Low Energy Photon Colliders

Interactions involving multiple photons:

- Interaction of intense laser light with strong electromagnetic fields.
- Schwinger pair production.
- Oscillation effects.



# Low Energy Photon Colliders

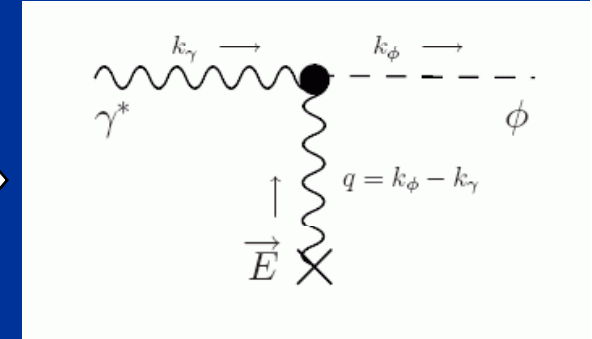
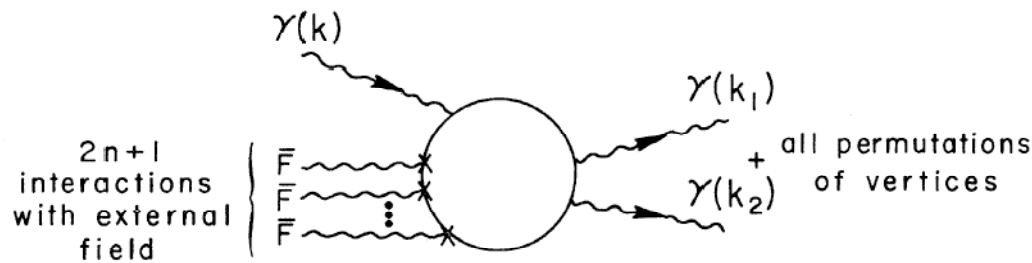
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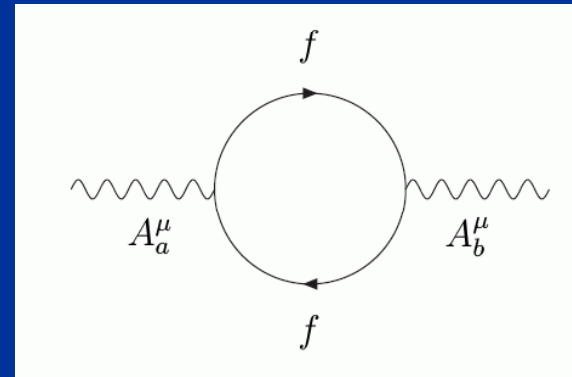
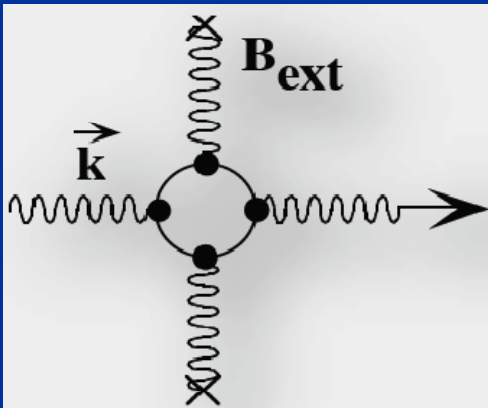


# New Physics may interfere with QED!

## Rotation:



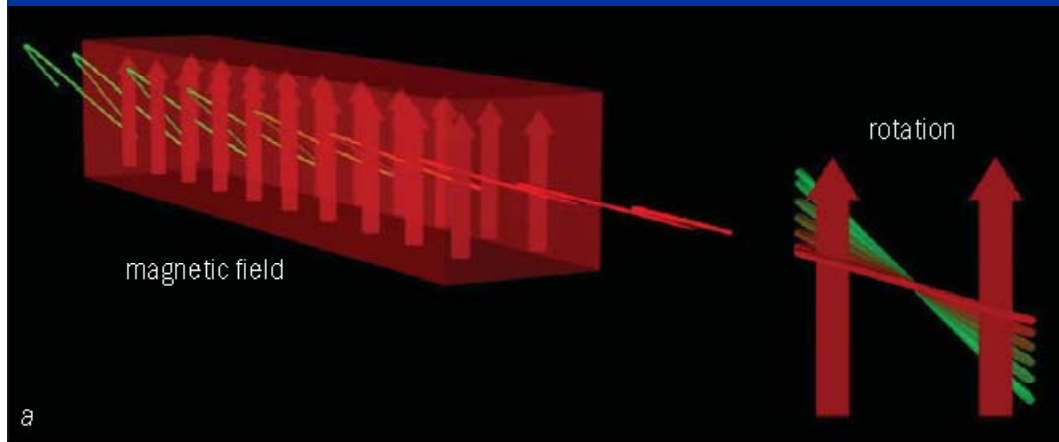
## Ellipticity:



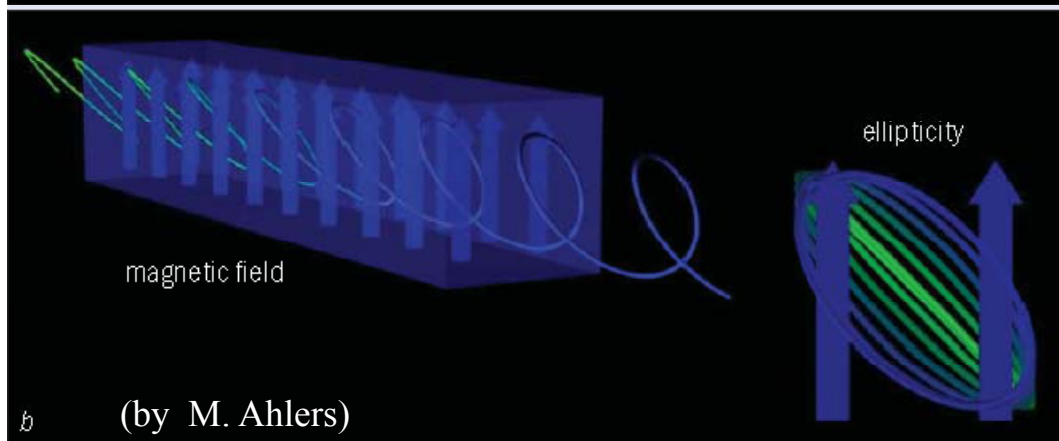


# Indirect WISP Search

## Change of laser light polarization passing a magnetic field



rotation:  
dichroism due to  
real WISP production



ellipticity:  
birefringence due to  
virtual WISP production



# Indirect WISP Search Experiments

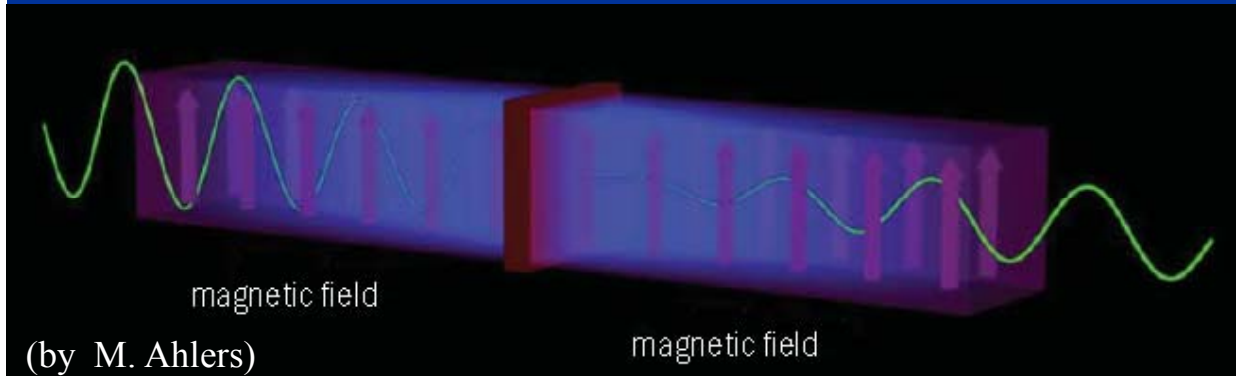
- Q&A in Taiwan
  - ongoing, sensitivity comparable to PVLAS
- BMV at Toulouse (France)
  - starting, aiming for QED prediction of ellipticity due to  $e^+e^-$  loop
  - also direct search experiments
- OSQAR at CERN (using two LHC dipoles)
  - starting, aiming also for QED-ellipticity
  - also direct search experiments





# Direct WISP Search

“Light shining through a wall” (LSW) or  
“photon regeneration” experiments.



- cross-check of indirect searches,
- more simple determination of properties of new particles,
- access to WISPs not detectable in indirect searches.



# LSW Experiments

- **BFRT at Brookhaven National Laboratory, finished in 1993**
  - limits on existence of WISPs
- **BMV at Toulouse (France)**
  - ongoing, limits published
- **GammeV at Fermilab**
  - ongoing, limits published
  - most sensitive limits on WISPs
- **LIPSS at Jefferson Lab. (USA)**
  - ongoing, preliminary results



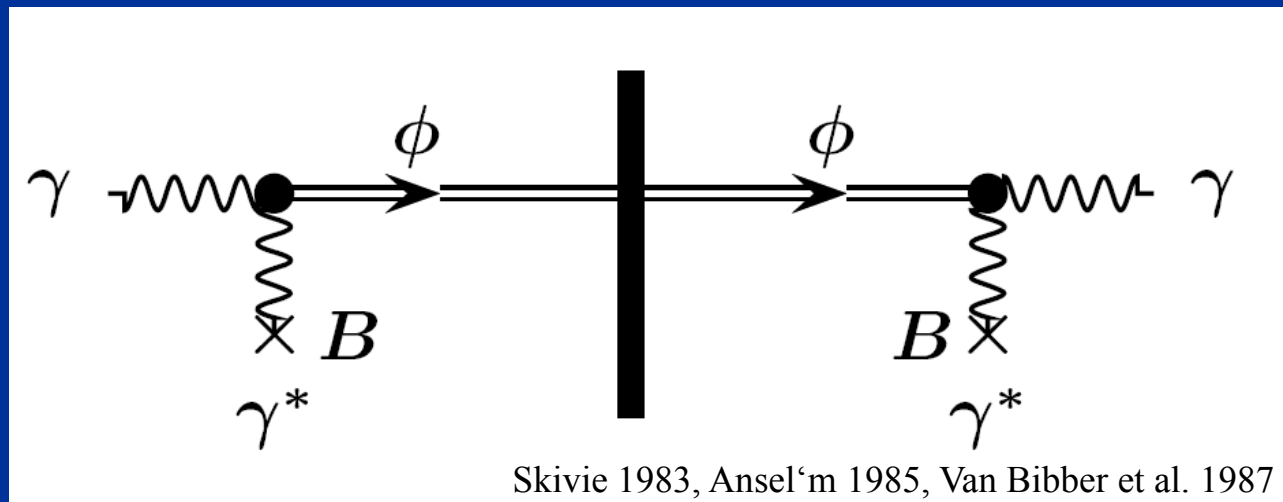
# LSW Experiments

- PVLAS at INFN Legnaro (Italy) ?
- OSQAR at CERN (using two LHC dipoles)
  - ongoing, preliminary results
- ALPS at DESY



# Brief Introduction to Photon Regeneration Experiments

*Light shining through a wall*





# The ALPS Project

## Axion-Like Particle Search @ DESY



A photon regeneration experiment





# The ALPS Project

## Axion-Like Particle Search @ DESY

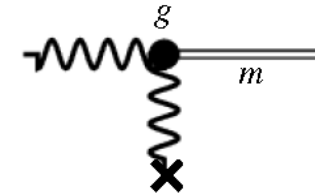
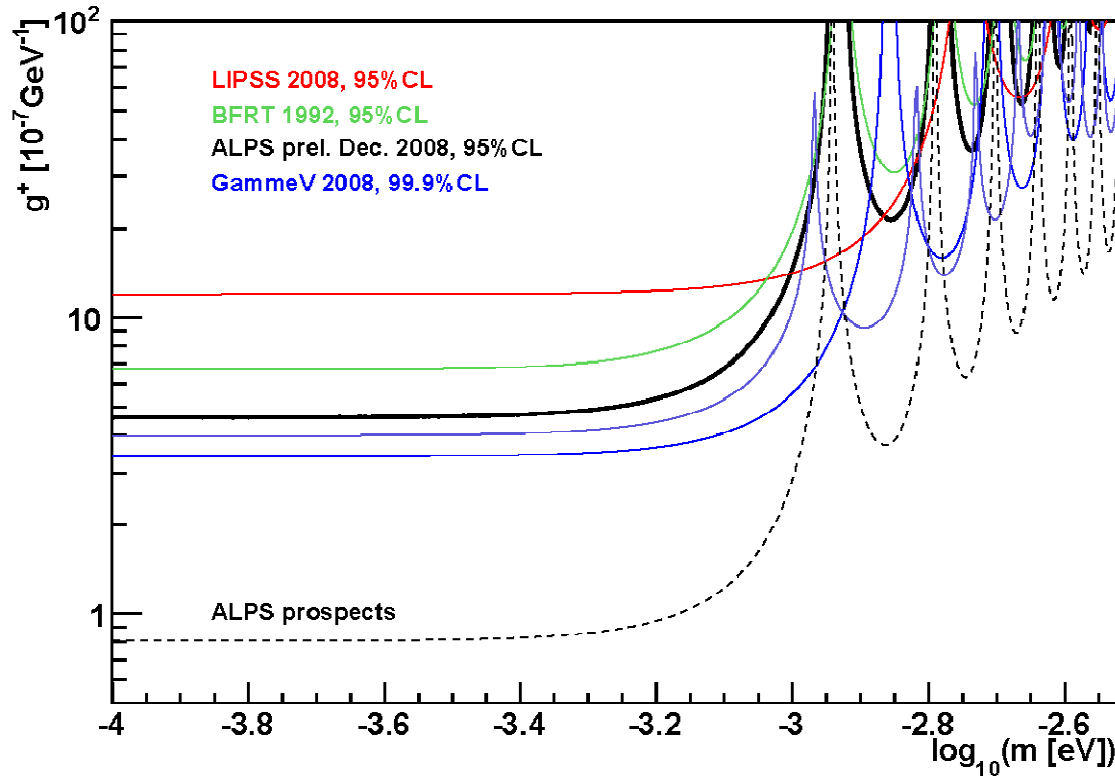
- DESY
- Max Planck Institute for Gravitational Physics (Albert Einstein Institute), and Institute for Gravitational Physics, Leibniz University Hannover
- Laserzentrum Hannover
- Hamburger Sternwarte

A photon regeneration experiment



# ALPS Sensitivity

Limit on Coupling for scalar ALPs



Successful and stable operation of a cavity in the generation part!



# Future Prospects of direct WISP Searches

- Increase sensitivity for lower couplings
- Extend mass range to higher values.





# Towards lower Couplings

## “Brute force” approach:

- Laser (power + optical cavity)
- Magnet (field strength + length)
- Detector sensitivity



# Towards lower Couplings

## Laser (flux of incoming photons) in tight magnet bore:

- ALPS at present: 0.7 W 532 nm,  
cavity with power built up of 40 0.03 kW
- ALPS prospects: 10 W 532 nm (enhanced LIGO),  
cavity with power built up of 500 5 kW
- ALPS “dream”: 100 W 532 nm (advanced LIGO),  
cavity with power built up of 1000 100 kW
- OSQAR proposal: 1kW 1064 nm (Nd-YAG),  
cavity with power built up of 10.000 10,000 kW  
([http://axion-wimp2007.desy.de/e30/e126/talk\\_Siemko.pdf](http://axion-wimp2007.desy.de/e30/e126/talk_Siemko.pdf))

Laser: 100 kW seems to be possible, MW real challenge!



# Towards lower Couplings

## Magnet (interaction probability):

- ALPS at present:  $\frac{1}{2} + \frac{1}{2}$  HERA dipole,  
B=5.2 T, l=4.2 m 22 Tm
- OSQAR proposal: 1+1 LHC, dipoles  
B=9.7 T, l=14.3 m 139 Tm
- “dream”:  
2+2 dLHC dipoles (or 4+4 LHC dipoles),  
B=20 T, l=30m 600 Tm

Magnet:  $B \cdot l$  up to  $\approx 25 \cdot (B \cdot l)_{\text{ALPS}}$  possible?



# Towards lower Couplings

## Detector sensitivity:

- ALPS at present: SBIG-402 40 mHz
- ALPS near future: PIXIS 1024:B 4 mHz
- “kind-of-limit”: radioactivity, CR in  $20 \cdot 20 \mu\text{m}^2$  signal region at ALPS about 0.02 mHz
  - may be reached with TES (single photon counting) for example.

Detector sensitivity: increase by factor of 200 in future?



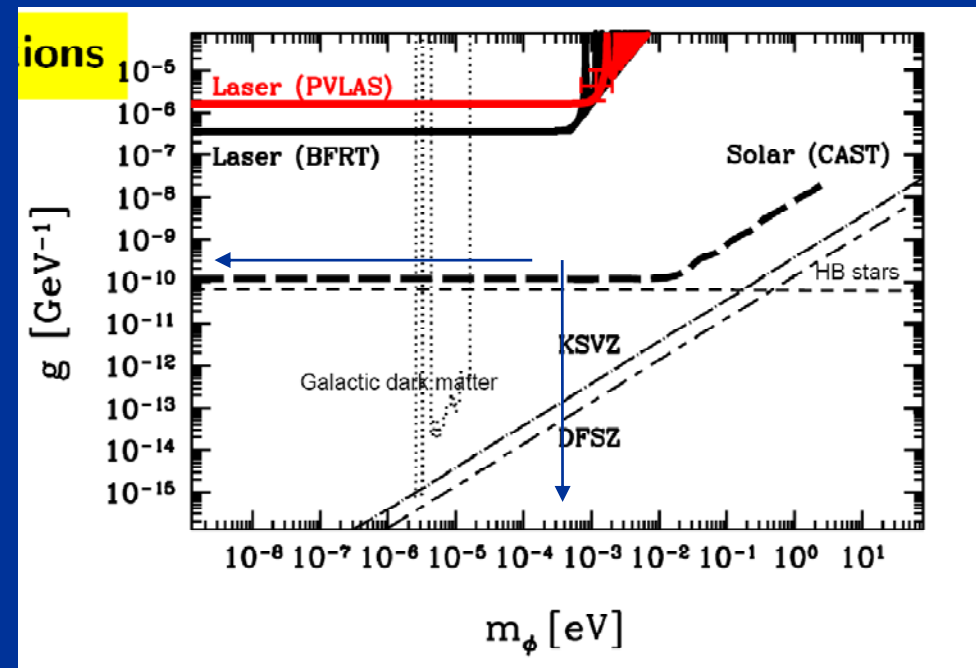
# Towards lower Couplings

Relative to ALPS in summer '09 aiming for  $g=10^{-7}\text{GeV}^{-1}$ :

- Laser (power + optical cavity): 100
- Magnet (field strength + length): 25
- Detector sensitivity: 200

## Physics:

$$\begin{aligned}
 g &= 2 \cdot (Bl)^{-1} \cdot (P_{\gamma \rightarrow \phi \rightarrow \gamma})^{-4} \\
 &= g_{\text{ALPS}} \cdot (25)^{-1} \cdot (100 \cdot 200)^{-4} \\
 &= g_{\text{ALPS}} / 300 \\
 &\approx 3 \cdot 10^{-10} \text{GeV}^{-1}
 \end{aligned}$$





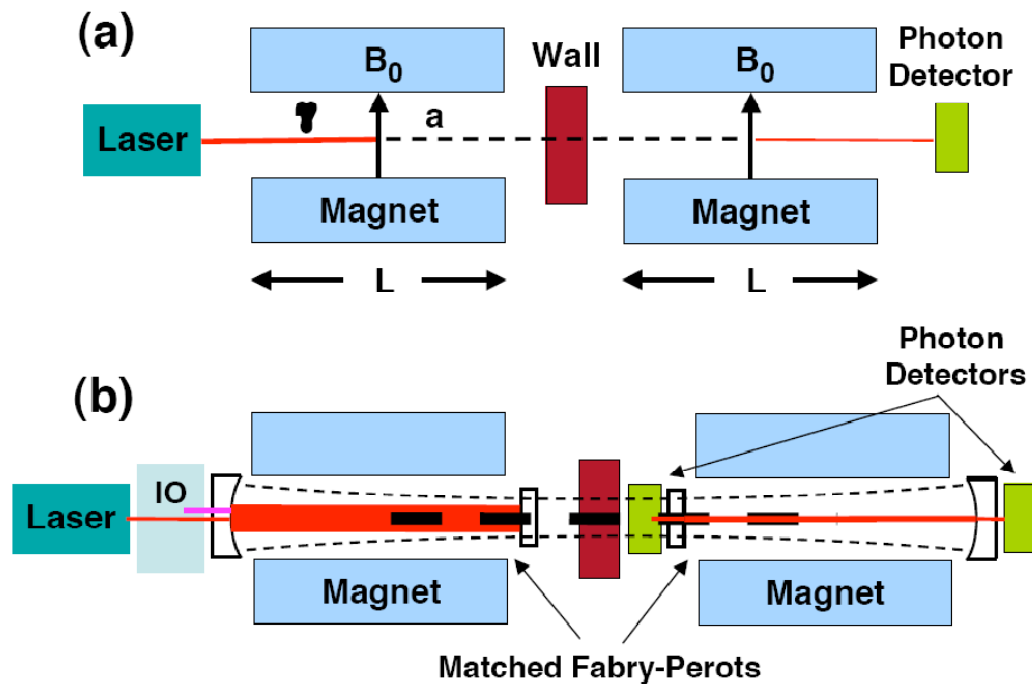
# Towards lower Couplings

Ingenuity in addition to “brute force”:

“Resonantly enhanced Axion-Photon Regeneration”

*P. Sikivie, D.B. Tanner, Karl van Bibber. Phys.Rev.Lett.98:172002,2007.*

*(also F. Hoogeveen, T. Ziegenhagen, DESY-90-165, Nucl.Phys.B358)*



Optical cavity also for the regeneration of photons from WISPs!

Increase power output by finesse of cavity:  
 $10^4$  seems to be possible.



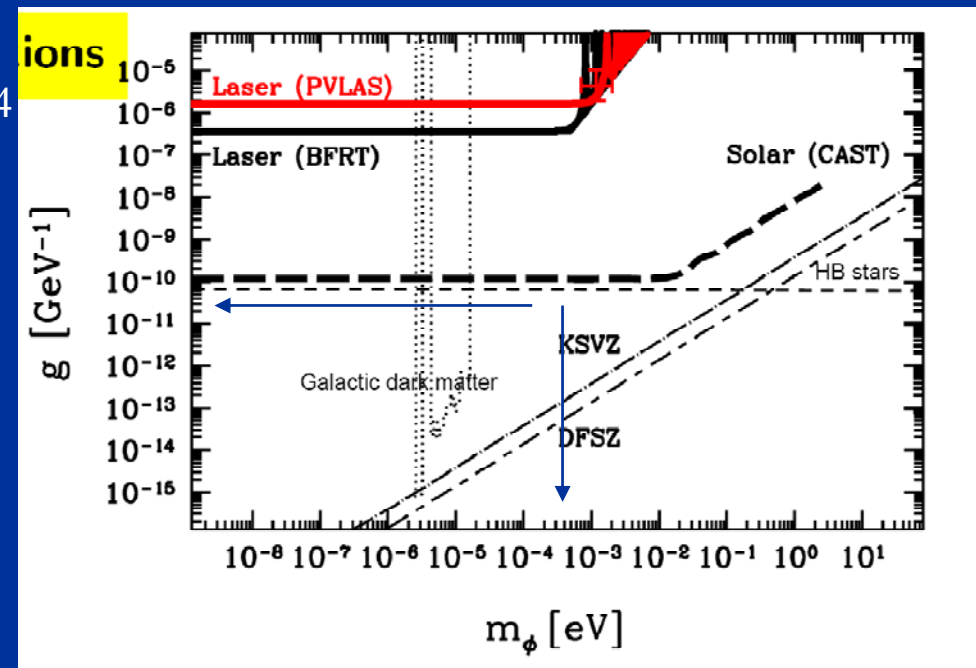
# Towards lower Couplings

With resonantly enhanced photon regeneration using a  $10^4$  optical cavity:

Physics:

$$\begin{aligned}
 g &= 2 \cdot (B1)^{-1} \cdot (P_{\gamma \rightarrow \phi \rightarrow \gamma})^{-4} \\
 &= g_{\text{ALPS}} \cdot (25)^{-1} \cdot (20,000 \cdot 10^4)^{-4} \\
 &= g_{\text{ALPS}} / 3,000 \\
 &\approx 3 \cdot 10^{-11} \text{GeV}^{-1}
 \end{aligned}$$

Within a decade limits from astrophysics might be surpassed.



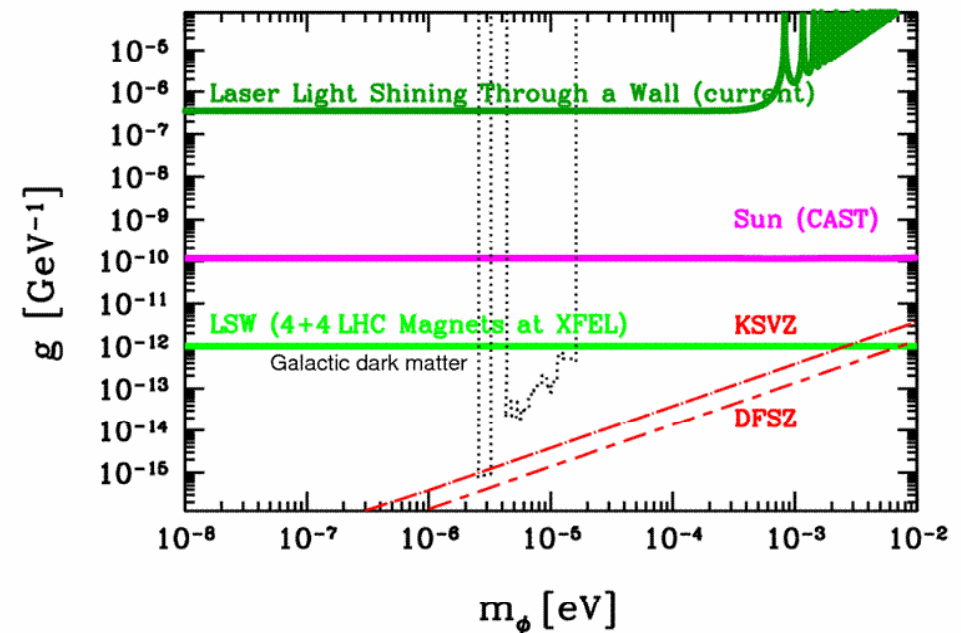


# Towards higher Masses

Big science (again):

With about one year of beam time at the XFEL and an installation of 4+4 LHC dipoles axions with masses above 1 meV could be probed.

- Advantage:  
detection of keV  $\gamma$  “easy”.
- Challenge:  
how to convince the  
internat. XFEL-GmbH?







# Low Energy Photon Colliders

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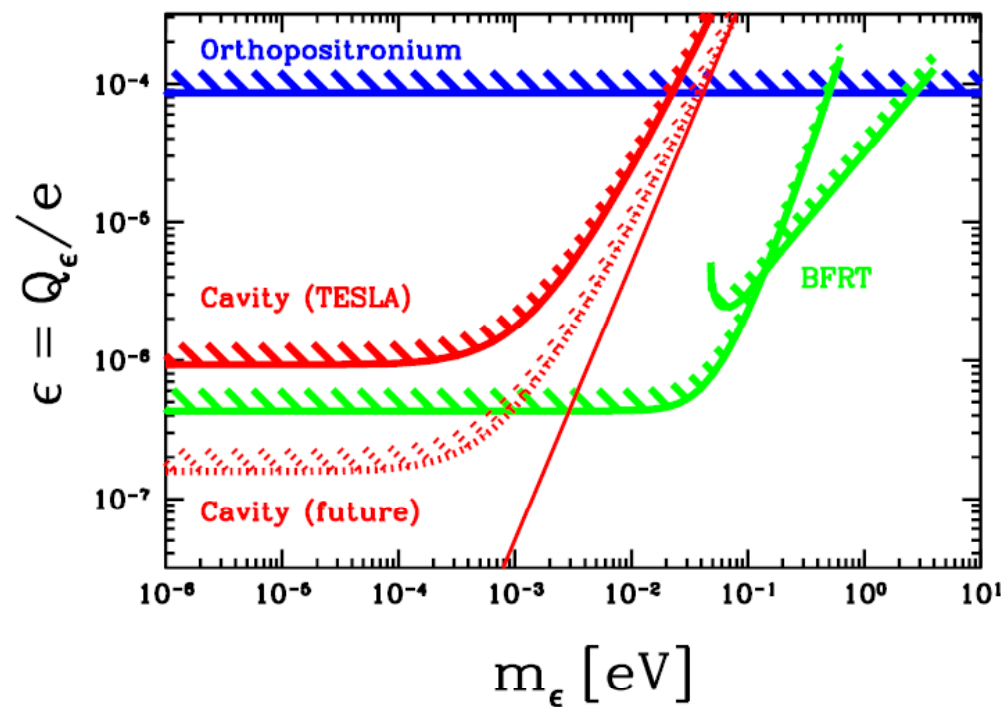


# Schwinger Pair Production

If field strength  $E > m^2/\epsilon e$ : spontaneous pair production!

- for electron/positron:  $E > 10^{18}$  V/m
- for MCP with  $\epsilon=10^{-6}$ ,  $m=1$ meV:  $E > 5$  MV/m

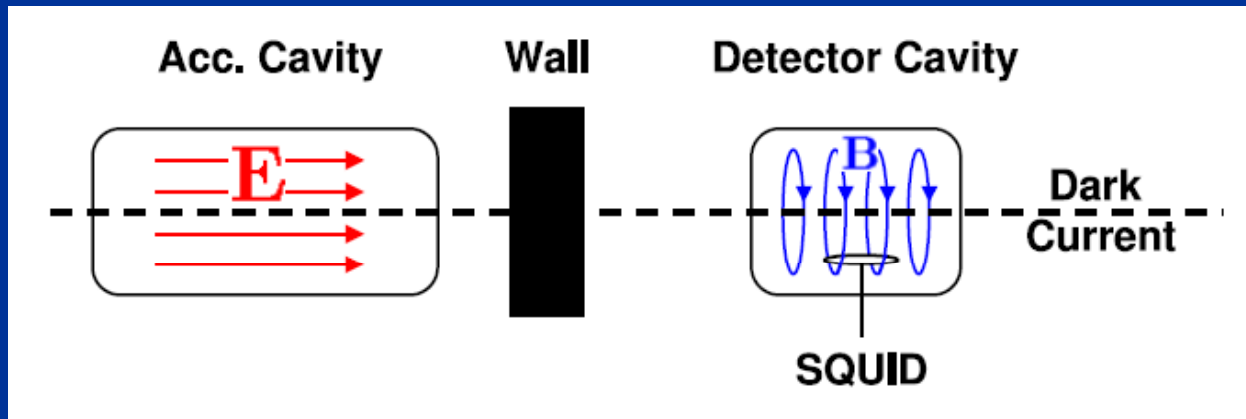
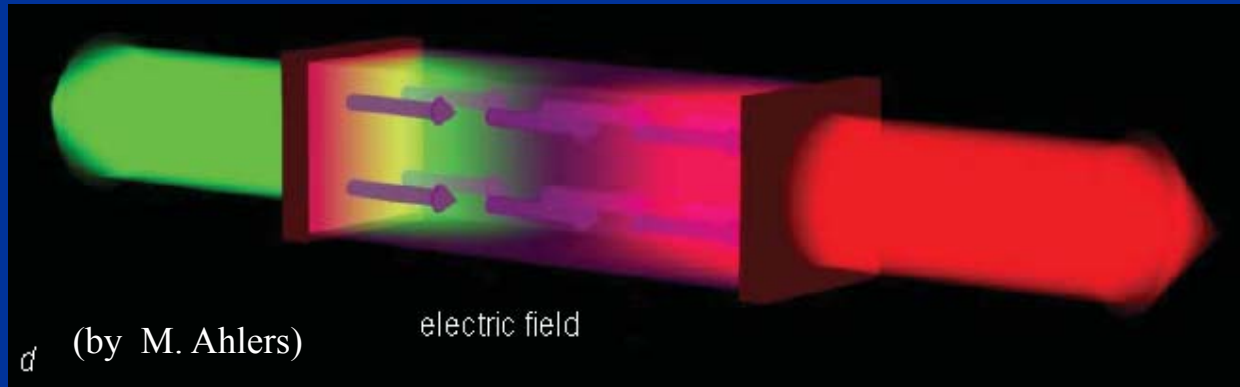
*Accelerator Cavities as a probe  
of millicharged particles,  
H. Gies, J. Jaeckel, A. Ringwald,  
Europhys. Lett. 76(5), 794 (2006)*





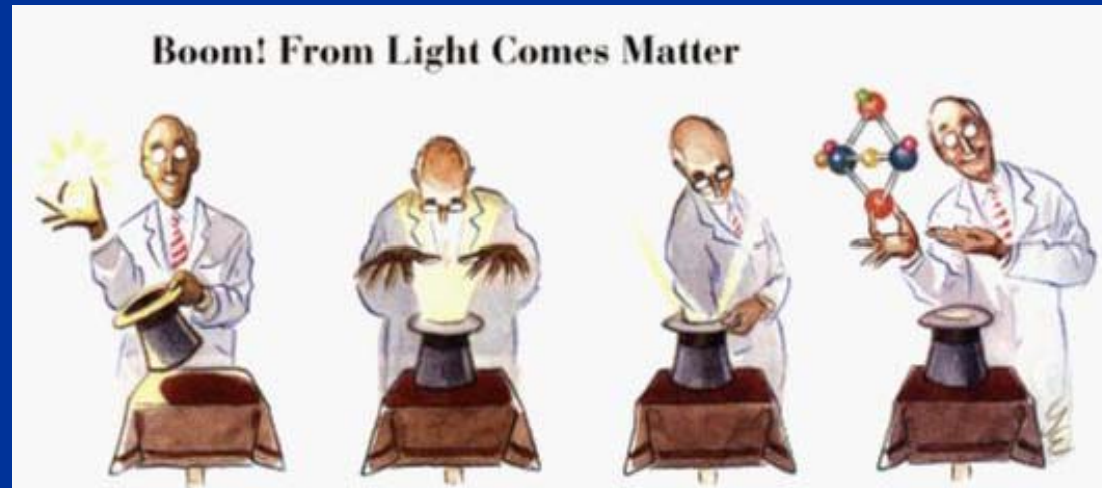
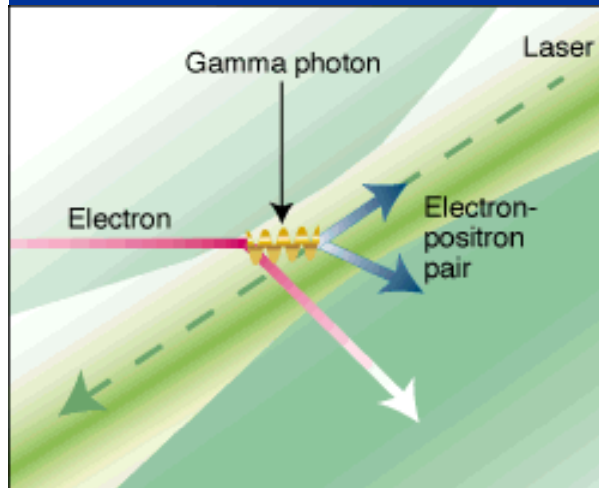
# Schwinger Pair Production

“Current flowing through a wall” experiments.





# Schwinger Pair Production



<http://www.slac.stanford.edu/exp/e144/e144.html>

Repeat the SLAC E-144 experiments at XFEL to probe QED in the non-linear regime?

First discussions with G. Carugno and G. Ruoso have started.



# Low Energy Photon Colliders

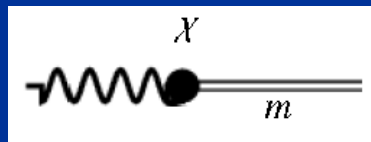
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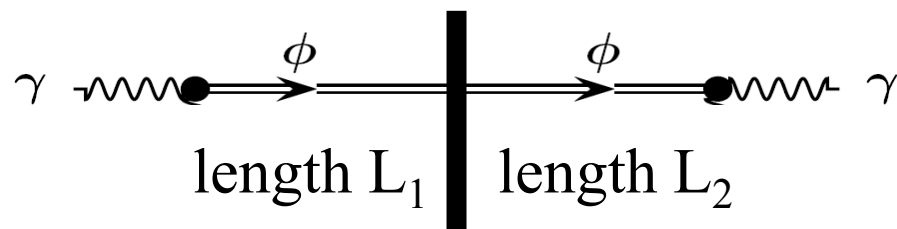
# Search for massive Photons

“Light shining through a wall” without external fields:



(like neutrino oscillations).

Principle of an experiment:



Experimental requirements very similar to searches for axion-likes, but:

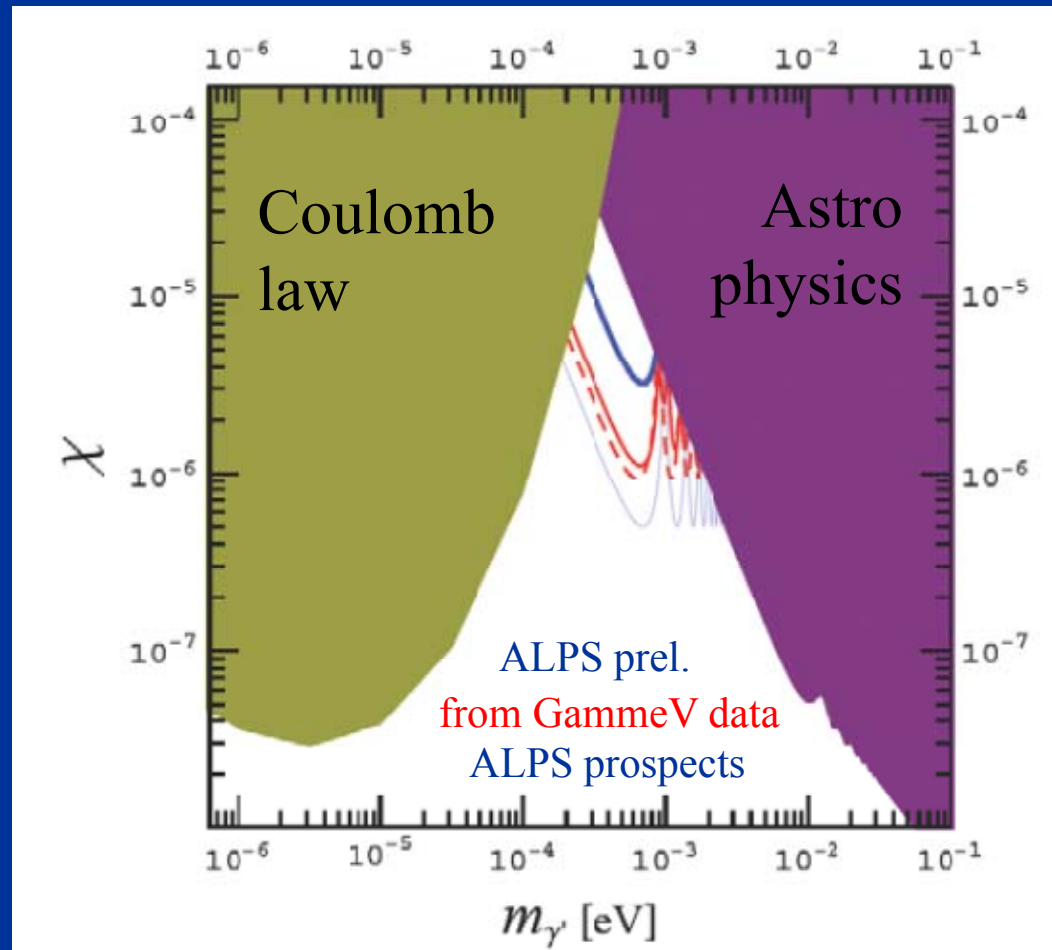
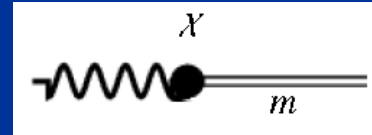
- no magnetic field,
- UHV conditions.

$$P_{\text{reconv.}} = 16\chi^4 \cdot [\sin(qL_1/2) \cdot \sin(qL_2/2)]^2 \quad (\text{kinetic mixing})$$



# Preliminary ALPS Sensitivity

95% CL limits for massive hidden sector  $\gamma$



Only laboratory experiments searching for massive hidden sector  $\gamma$  might close the gap in the meV mass region!



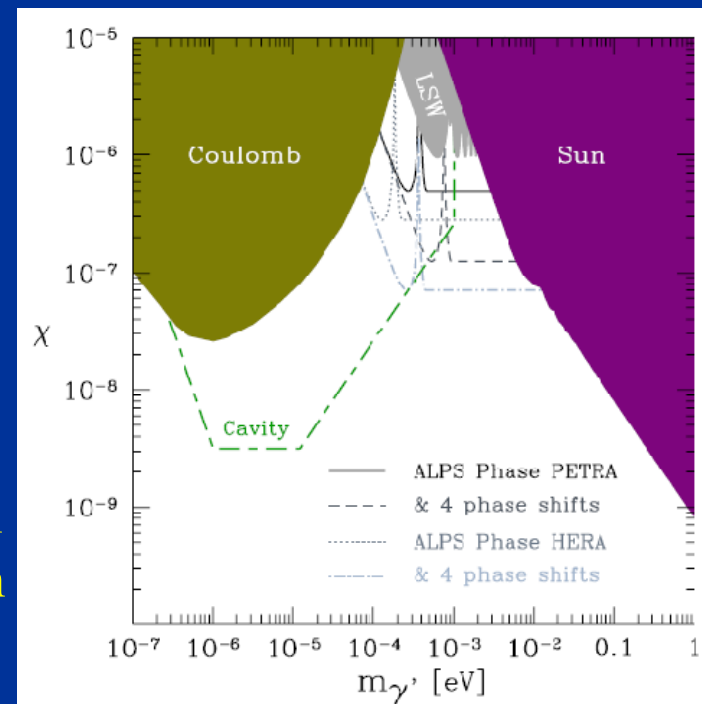
# Future Prospects

Searches for massive hidden sector photons benefit from the laser and detector developments sketched above.

Different mass regions may be probed by different lengths ( $L_1$  and  $L_2$ ) of the vacuum tubes.

*Laser experiments explore the hidden sector,  
M. Ahlers, H. Gies, J. Jaeckel,  
J. Redondo, A. Ringwald,  
Phys. Rev. D 77 (2008) 095001*

Phase “PETRA”:  $L_1 = L_2 = 40\text{m}$   
Phase “HERA”:  $L_1 = L_2 = 170\text{m}$





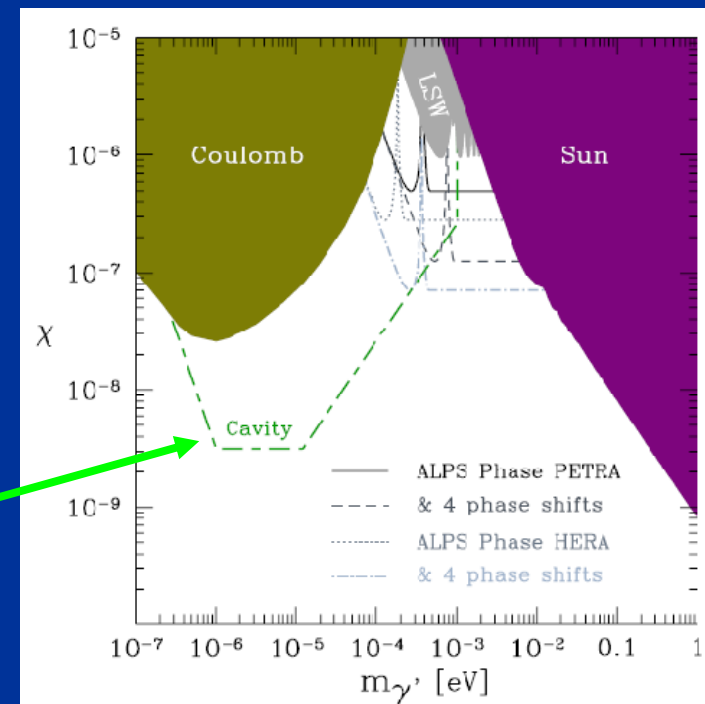
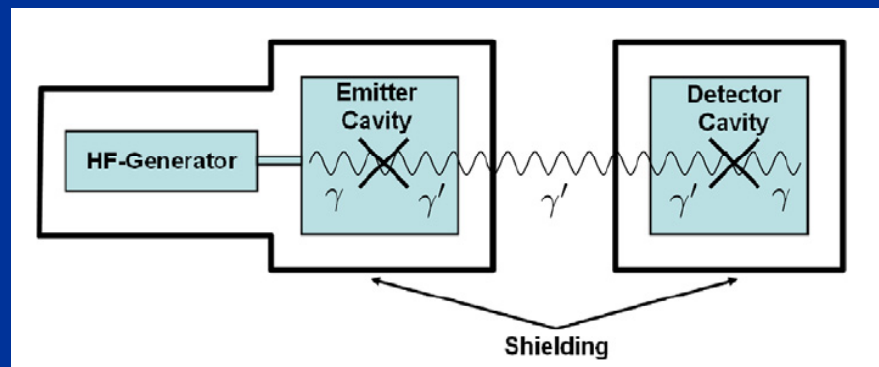


# Future Prospects

Searches for massive hidden sector photons benefit from the laser and detector developments sketched above.

Different mass regions may be probed also by different technologies.

*A Cavity Experiment to Search for Hidden Sector Photons,*  
*J. Jaeckel, A. Ringwald,*  
*Phys. Lett. B659 (2008) 509*





# Two Cavity Test Stands @ DESY



- Two adjacent test stands,
- well shielded,
- perfectly matched for WISP searches.





# Summary

- Searching for WISPs in the laboratory is necessary to complement astrophysics experiments.
- There is a wealth of different experimental approaches.
- Microwave cavities might be minicharged particle factories.
- The sensitivity of future experiments will likely surpass present day limits from astrophysics.
- Finding the QCD axion remains a challenging target.
- The typical size of WISP direct search experiments are perfectly matched to laboratories like DESY / CERN.



# A WISP Future at DESY?

## The laboratory's interest:

- small scale particle physics experiments on site
- exploit possibilities of new light sources (PETRA III, FLASH, XFEL) for particle physics.
- help developing the research field further (theory).



# A WISP Future at DESY?

## Resources:

- Rather limited at present, therefore participation in external experiments unlikely at present.
- The situation will be reviewed in autumn 2009 depending
  - on results achieved at ALPS,
  - on the outcome of the strategic Helmholtz review in spring 2009 for the funding period 2010 to 2014.

Collaboration very welcome!



# It might look challenging ...

THIS IS THE MOUNTAIN  
WHICH I MUST CLIMB



ITS UPPERMOST PEAKS STRETCH  
INTO THE DISTANCE, AND I  
AM BEGINNING TO DOUBT THAT  
I WILL EVER MAKE IT



.. but surprises might be close!

