

Introduction to Particle Physics

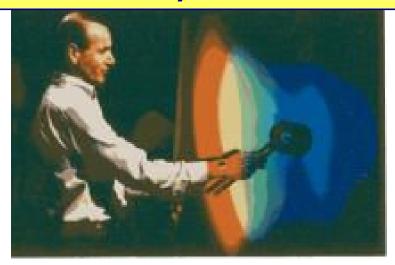
(for non physics students)

4. UNIFIED UNIVERSE

(no strings attached)

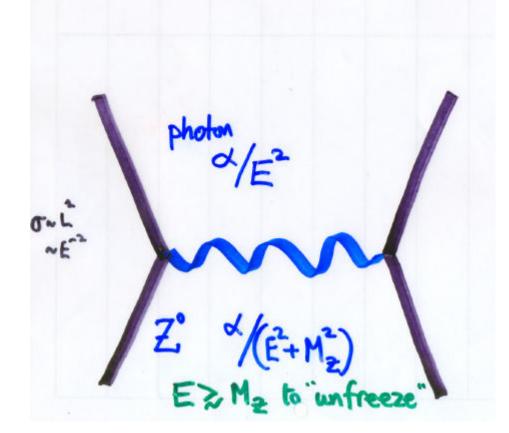


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EXETER COLLEGE
UNIVERSITY OF OXFORD



FORCES Summary (remember that waves -> particles)

NAME	action	CARRIER
Gravity	keeps us on	graviton?
Electromo	solids stops us falling to centre of Garth	byoton (A)
Weak	β-radioactivity P→He in Sun	W+ W- Z0
Strong	quatks glued inside p.n P.n in nuclei	gluons (g) 8 different
Only the	week force carriers	have MASSes
Mw ~	80 GeV/c2	
M_~	91 GeV/c2	



Feynman rules:

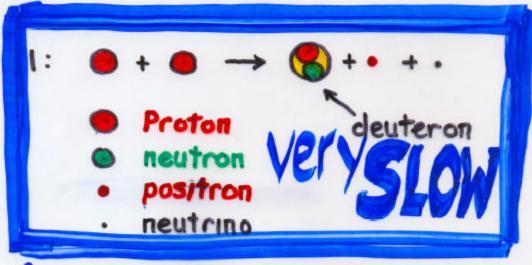
If energy E flows through the transmitted "virtual" particle (photon; Z) it costs 1/(E^2+M^2)



If E >> M the cost is 1/E^2....like the case of the photon

Only appears weak at low energy. Unified at high energy

At the heart of the Sun:

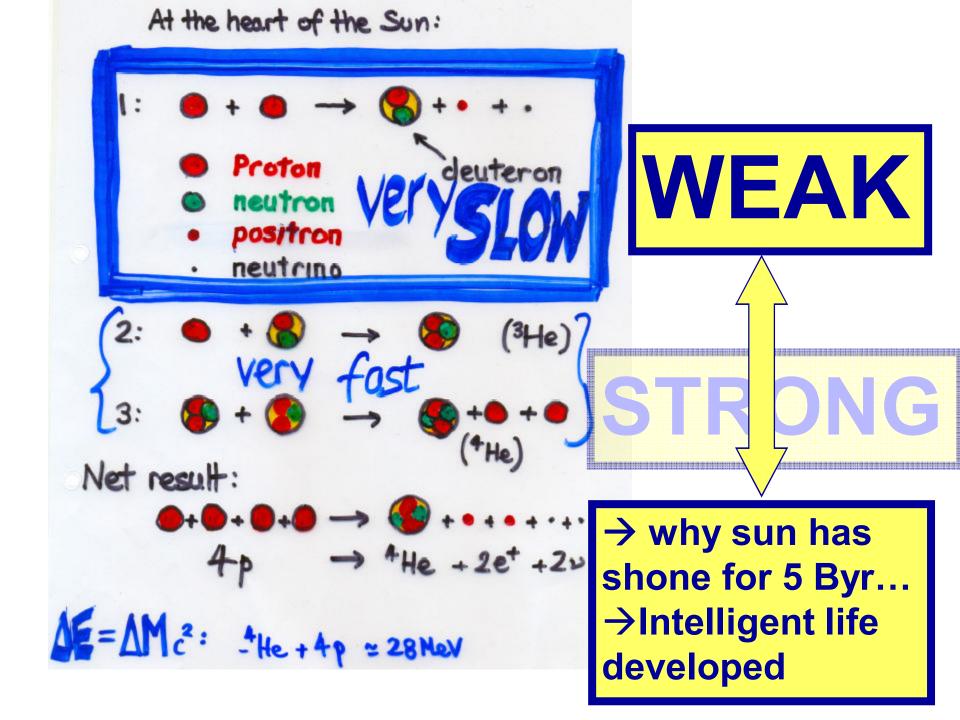


WEAK



STRONG

Net result:

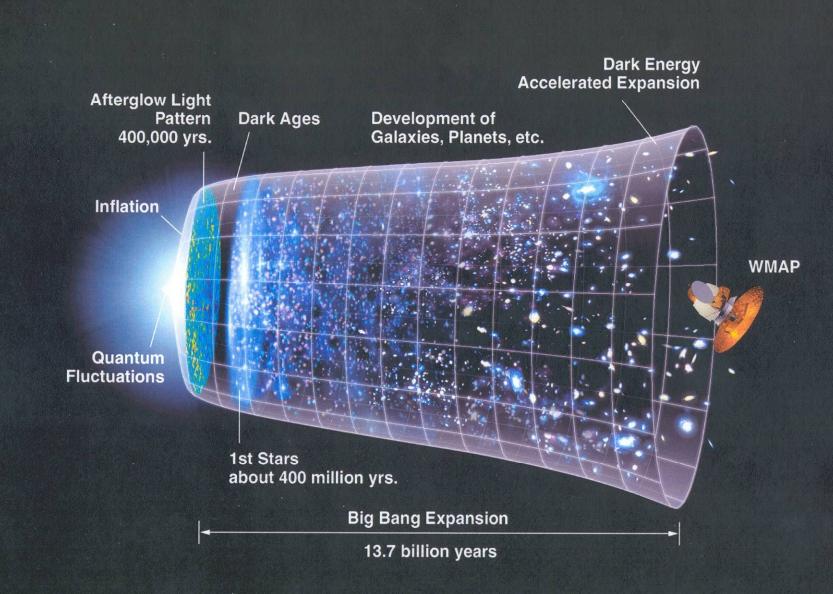


The weak force is feeble in the Sun ...

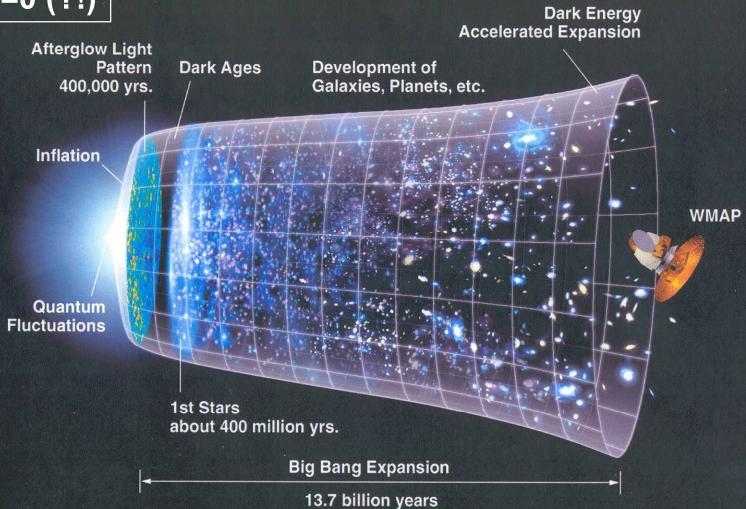
..because 10,000,000K ~ 1 keV << 80 GeV

...this is why the sun has stayed active long enough for us to have evolved and be having this conversation.

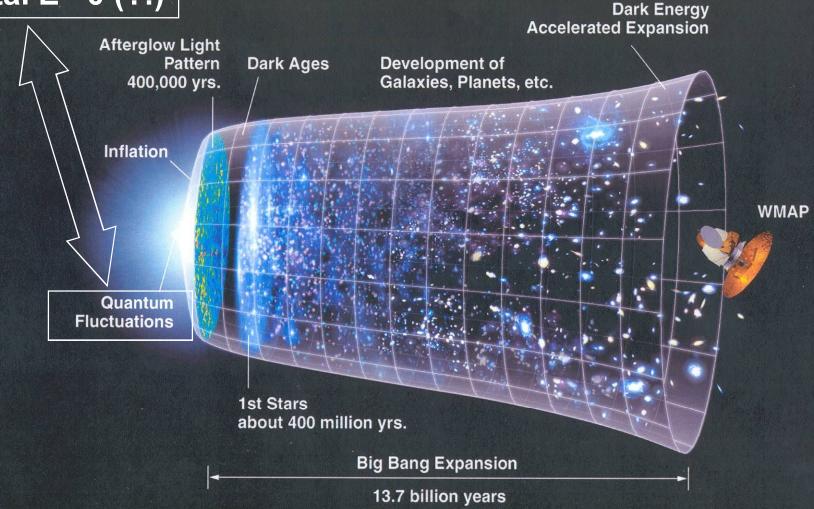
- →We exist because m(W) is not zero
- → Mass matters



Mc^2 >0 Gravity PE <0 Total E =0 (?!)



Mc^2 >0 Gravity PE <0 Total E =0 (?!)



 $Mc^2 > 0$ **Gravity PE < 0** Total E = 0 (?!)

Dark Energy

Accelerated Expansion Afterglow Light Pattern **Dark Ages Development of** Galaxies, Planets, etc. 400,000 yrs. Inflation WMAP Heisenberg: Quantum **Fluctuations** 1st Stars about 400 million yrs.

"Universe erupted as quantum fluctuation out of nothing"

E ->0

t ->infinity

Big Bang Expansion

13.7 billion years

Mc^2 >0 Gravity PE <0 Total E =0 (?!)

Dark Energy Accelerated Expansion

Afterglow Light Dark Ages Development of Pattern 400,000 yrs. Galaxies, Planets, etc. Inflation WMAP Quantum **Fluctuations** 1st Stars about 400 million yrs.

Heisenberg:

E ->0

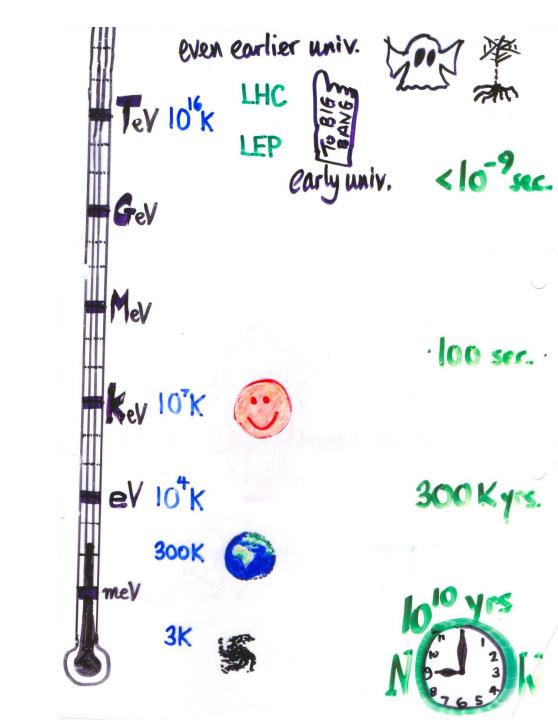
t ->infinity

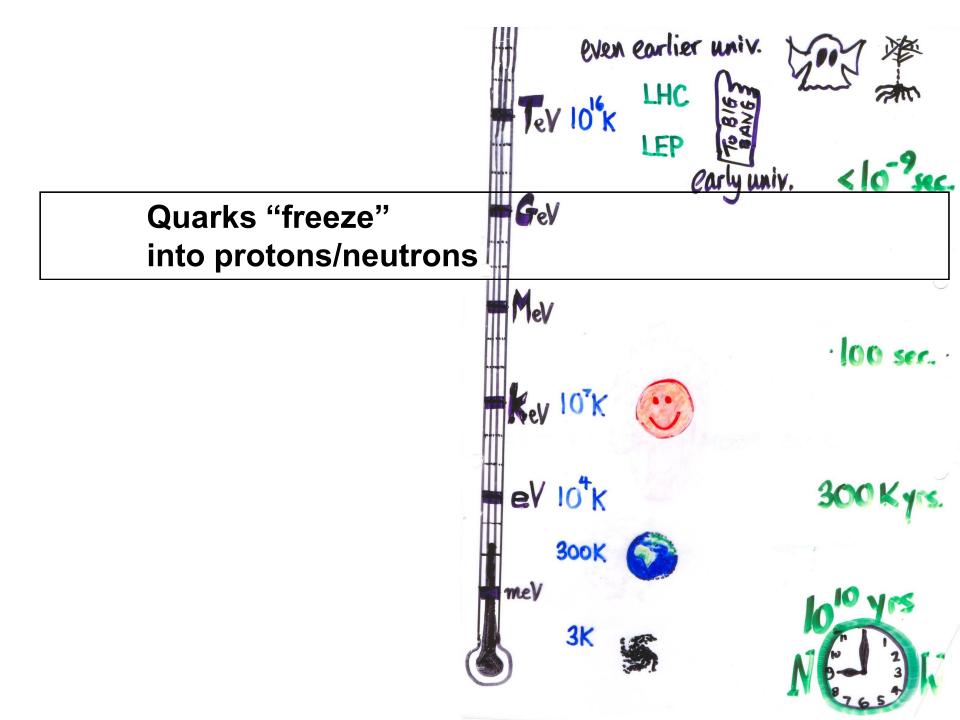
"Universe erupted as quantum fluctuation out of nothing"

(what/where/how encoded Heisenberg in "nothing"?

Big Bang Expansion

13.7 billion years





Big Bang + P

Thermal Equ	ilibr	muir	
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Temperature (energy) drops >
After 1 usec > one way only:





But at the same time:

then like processes 2 and 3 in the Sunxuntil all the neutrons have gone

* MAKING

9

particles so far apart in the expanding universe that they no longer interact

T = 1 usec after BIG BANG

neutrinos are free ("the first fossils in the Universe)

move at high speed
and if they have mass they
start clustering together
-contribute to formation of galaxies

Billion > per adom

if m(v) > m(proton)/109 = lev

they will dominate mass
density of the Universe

or future of universe and its formation

Universe expands - and cools expansion rate Rate depends on pressure which depends on the temperature in the gas and

which depends on the temperature in the gas and the number of neutrinos inside the gas volume (density) and this # depends on number of neutrino species

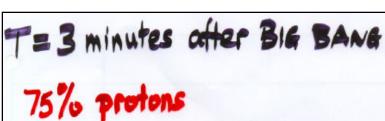
T=3 minutes after BIG BANG

75% protons
24%. Helium Nuclei
+ small amount of deuterons
+ free electrons.

Helium abundance; throngs of other depends on expansion rate of the

Universe which depends on number of neutrino species

Deuterium abundance depends on density of "ordinary malter" in the Universe.



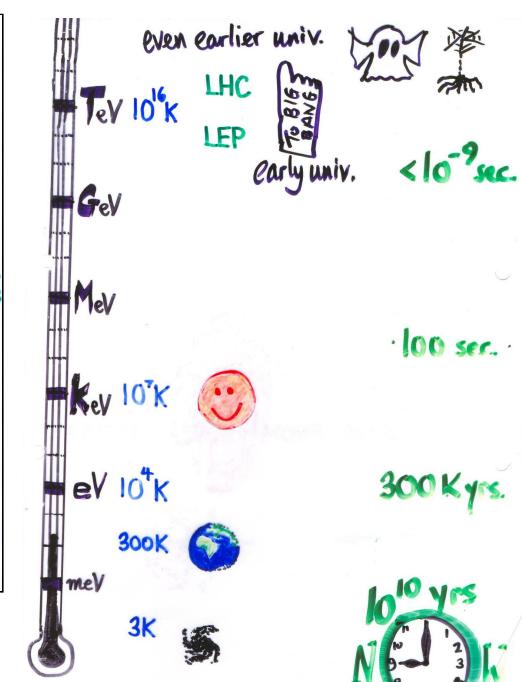
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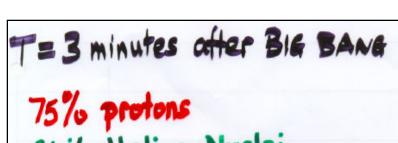
Helium abundance*; Horse elements

depends on expansion rate of the Universe which depends on number of neutrino species

Deuterium abundance depends on density of

"ordinary matter" in the Universe.





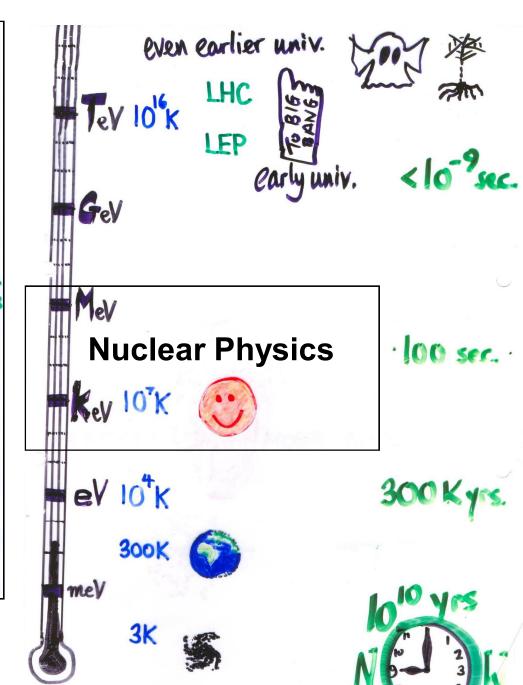
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Deuterium abundance
depends on density of
provincery malter in the Universe.

density of ordinary matter

the total in universe

part of DARK MATTER prese

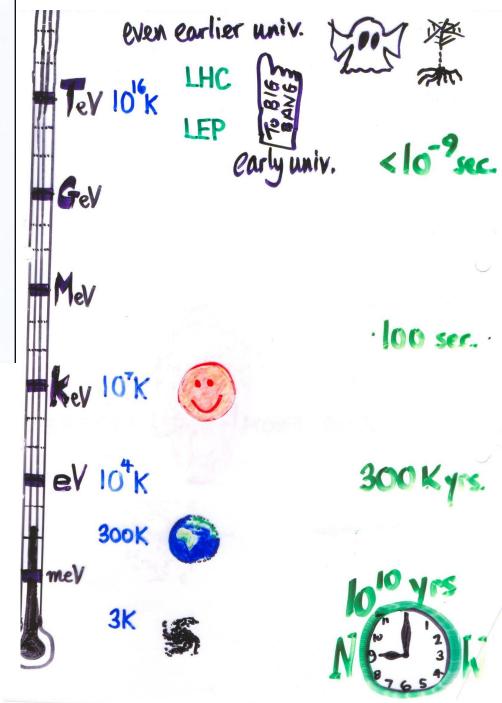
Job, 000 years later Exlock

electrons combine with nuclei
and make neutral atoms

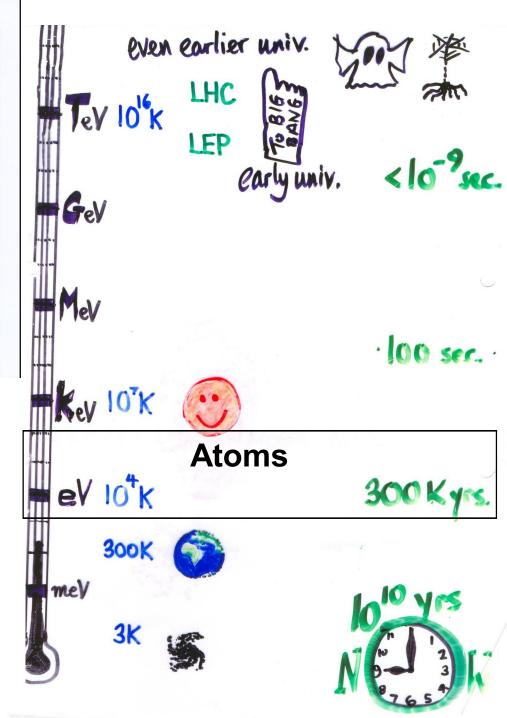
electromagnetic radiation was set free

electromagnetic radiation was set free Universe becomes transparent

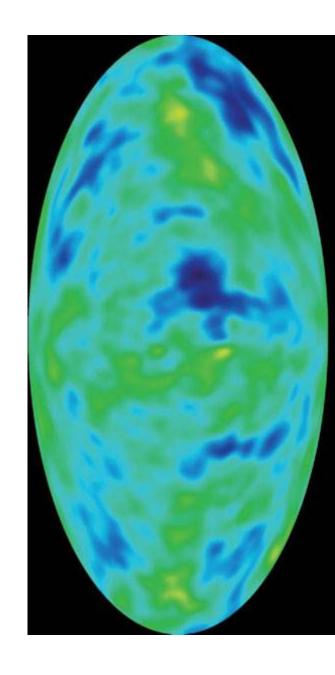
Time Passes. Temp drops 300,000 years later electrons combine with nuclei and make neutral atoms electromagnetic radiation was set free Universe becomes transparent



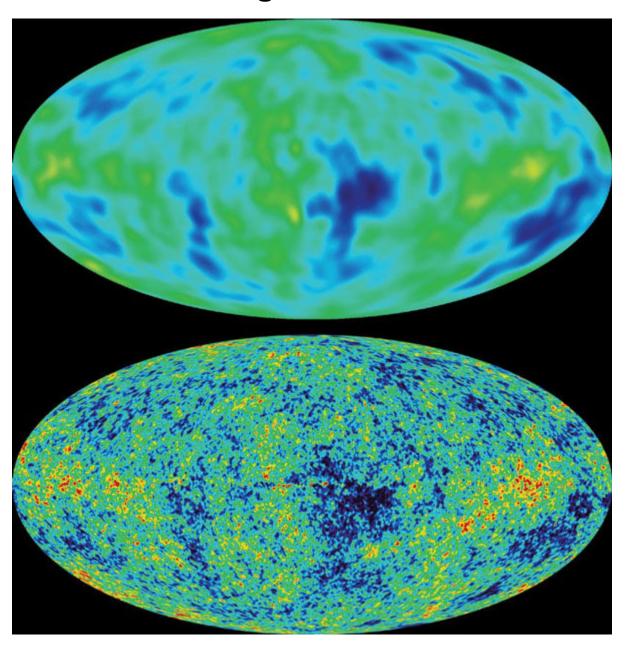
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Time Passes. Temp drops 300,000 years later Exlock T< 104K electrons combine with nuclei and make neutral atoms electromagnetic radiation was set free Universe becomes transparent 10 years later Emag 2 stretched: Micromone Rand. Black body background 3K (small fluctuations in Manuare rad = hints of proto structures, galaxier in early universe)



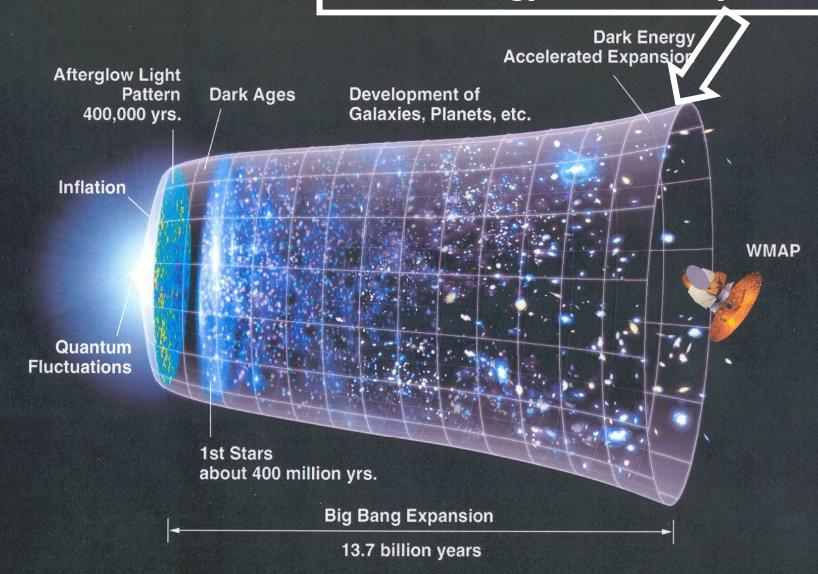
3K microwave bgnd now seen to have structure



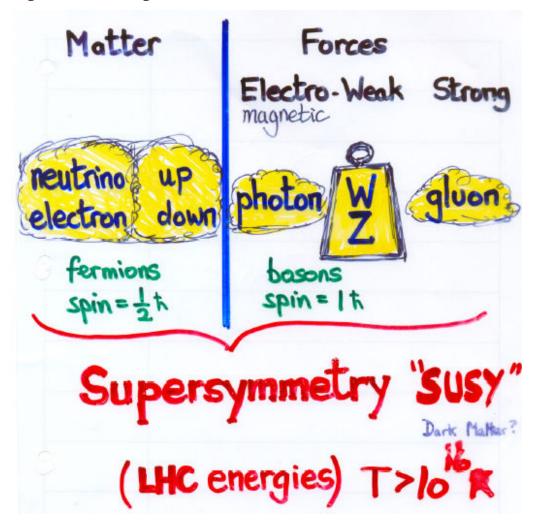
COBE 2000

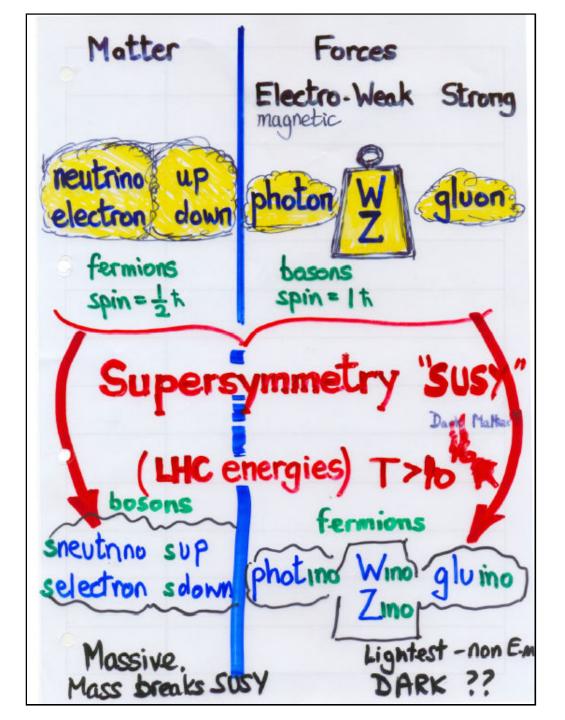
WMAP 2003

5byr ago accelerated expansion = Dark Energy. What? Why?.....



One further symmetry??





Whole new families to be found

Particle Physics @ CERN. Standard Model of PHes + Forces. · Quarks + Leptons. Spin 1/2 fermions enoulp, StW & . Spin 1 gauge basons · Higgs Spin o boson Origins of matter. Structures + pallerns at E& ITEV Symmetry revealed at E > O(TeV) Forces (and particles) unified - SUSY. Some current big pazzles. . Dark Malter, Solar v, massive v? (all the same?) . Why 3 generations What is difference between Mand M? [same? The Fifth Dimension

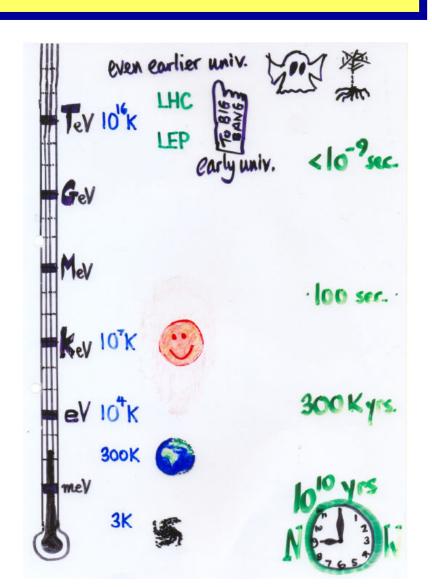
Finale: A glimpse of the future

recall from lecture 1.....

The Universe

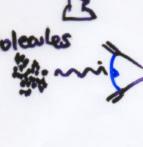
in

Temperature
Energy and
Time



even earlier univ. ...and the nature of matter LHC TeV 10 K LEP **AGPlasma** Auarks Gluons GeV neutrons & Nuclei melt Jexist Nucleus ** KeV 10 K ep Plasma H melt:plasma V.V. Alons & Lexist eV 10 K Molecules 300K

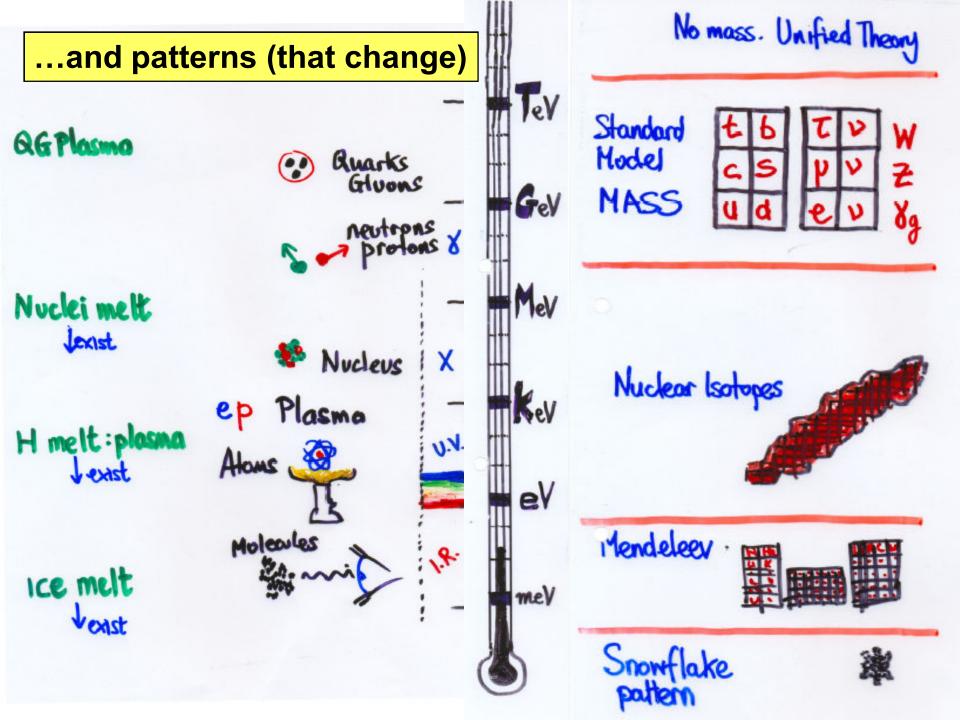
ice melt Loast



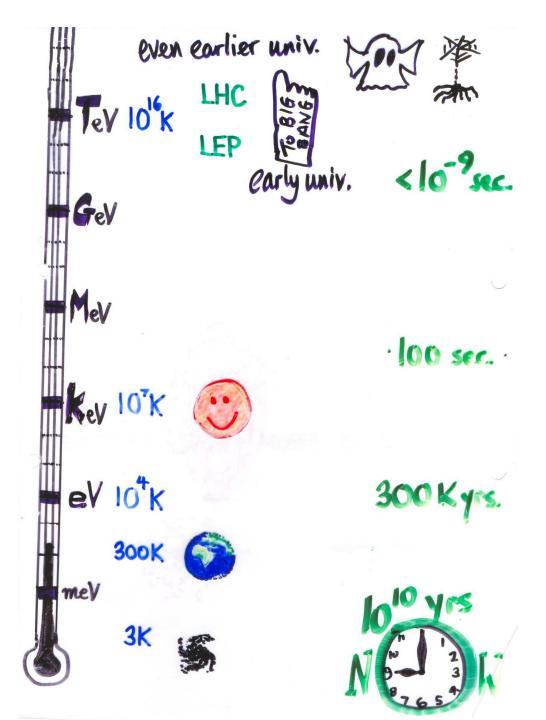




· 100 sec.



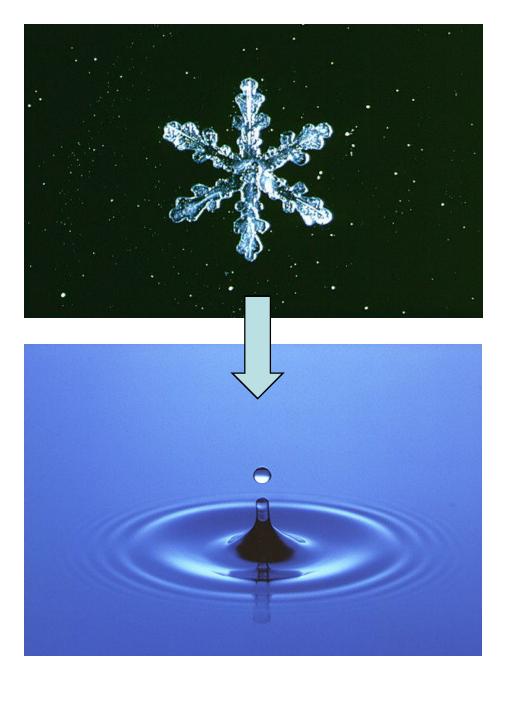
Temperature and symmetry in the universe



The Idea

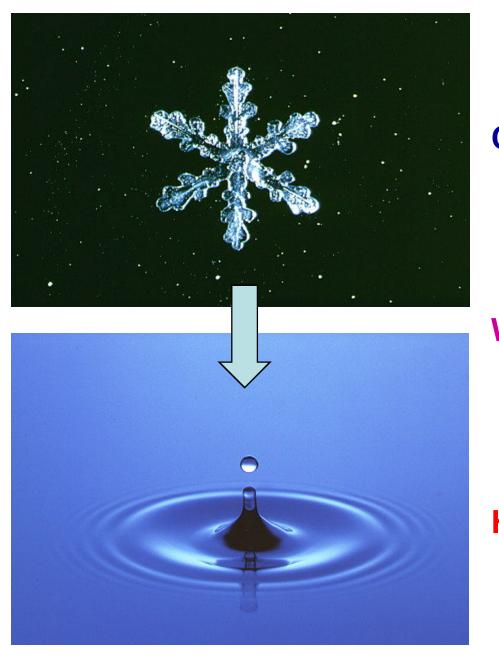


(I will tell you when to be cautious about inhaling)

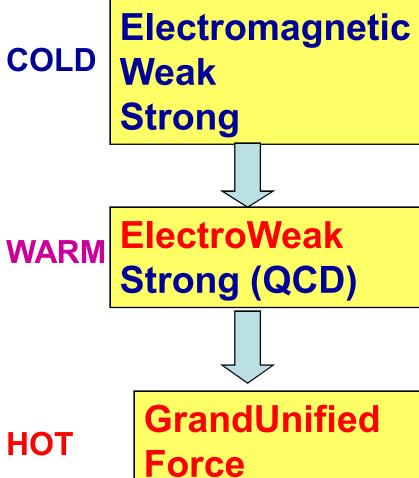


patterns and structures when cold (low energy)

Symmetry when warm (high energy)



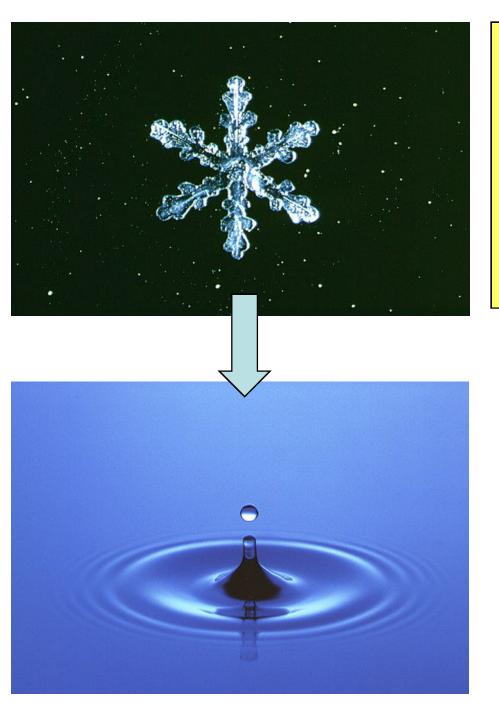
FORCES 1955-2005





Standard Model of **Quarks Leptons and forces**

- = pattern based on mass
- "cold" ="low" energy
- = below 1 TeV



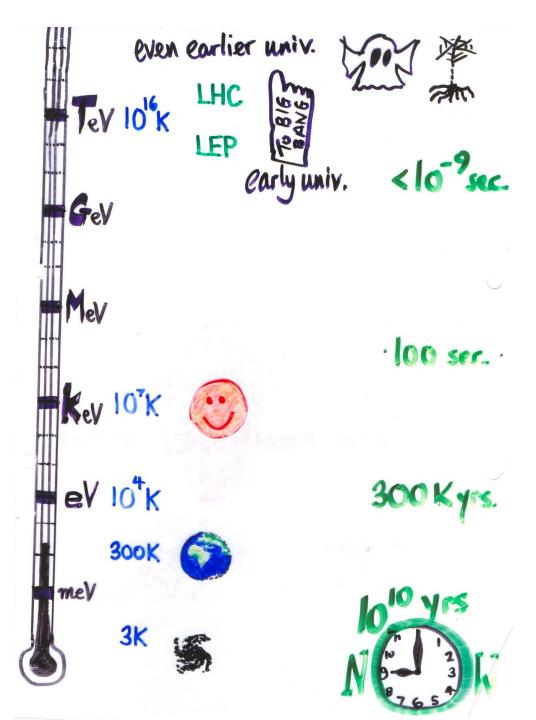
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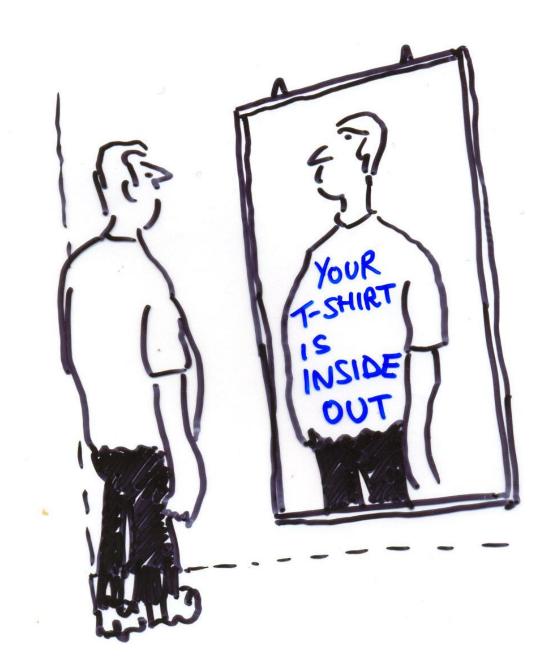
superSymmetry
when "warm"
(= high energy > 1TeV)

Higgs Boson Supersymmetry Nature of Reality

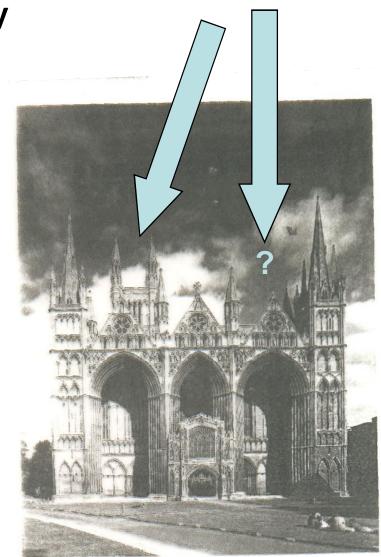




5. symmetries can disappear or change

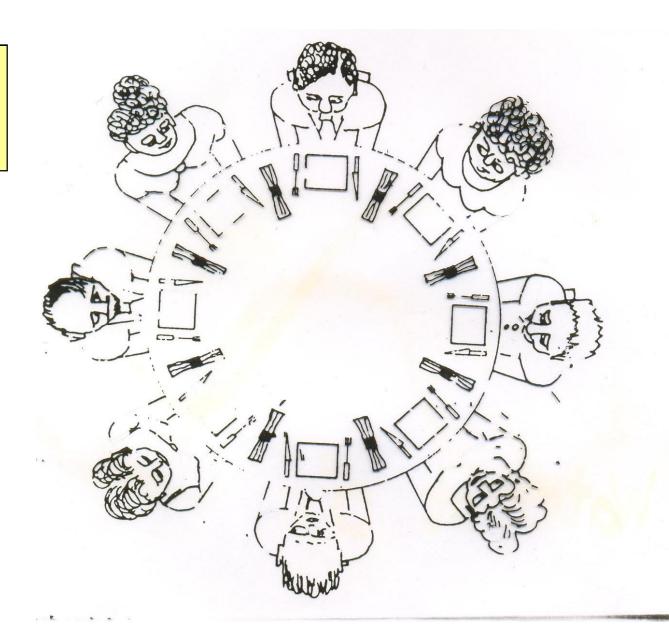


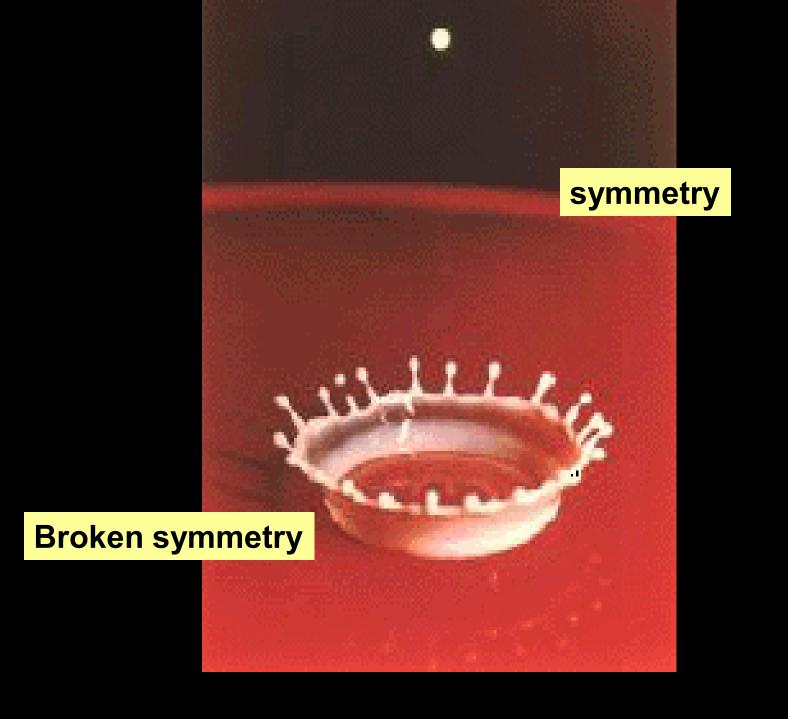
We like symmetry and when its absent we want to know why

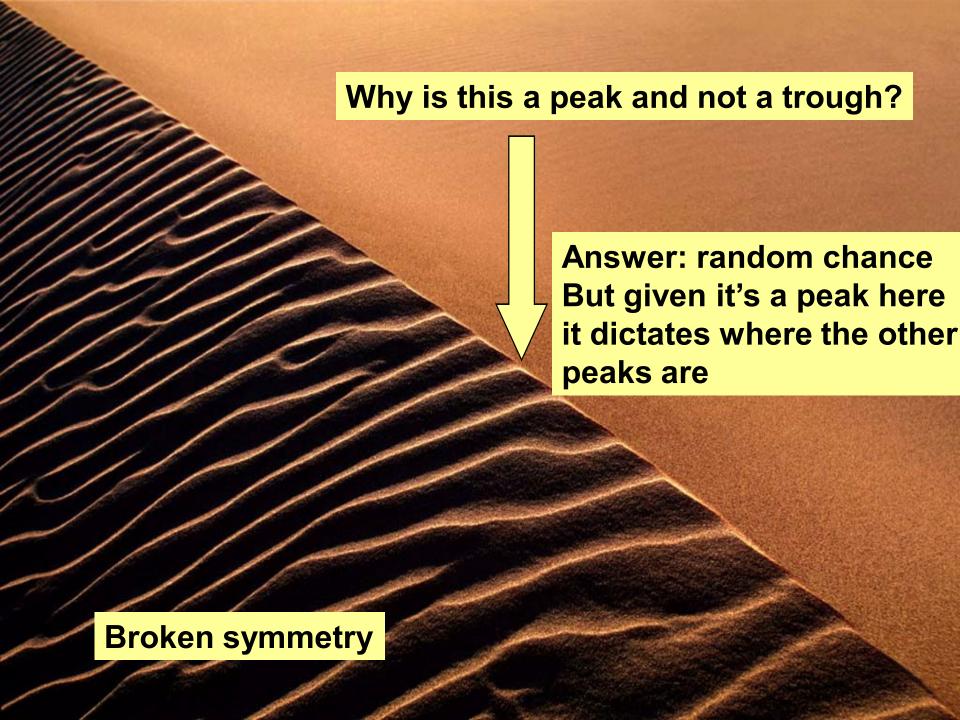


Buridan's Ass

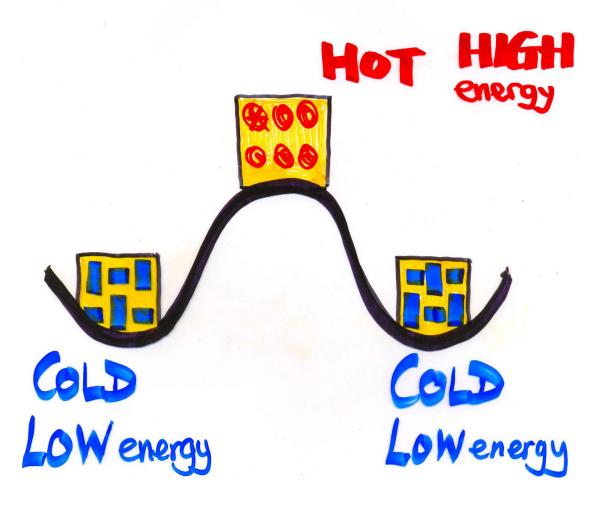
The problem of the symmetric dinner party

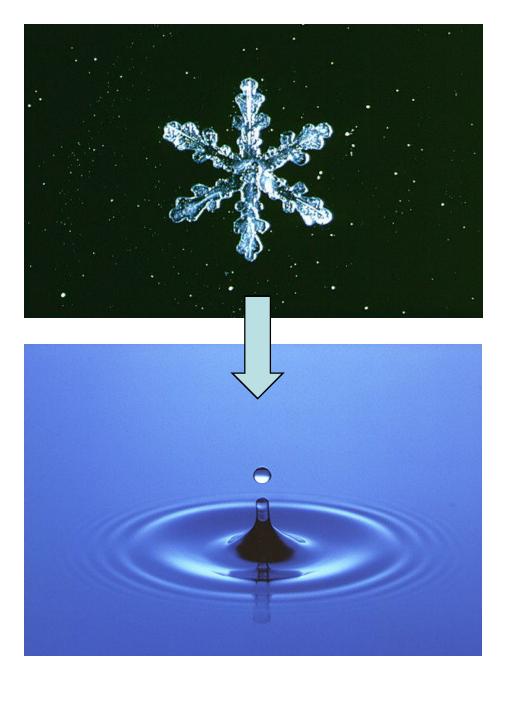






Magnets





patterns
and structures
when cold
(low energy)

Symmetry when warm (high energy)

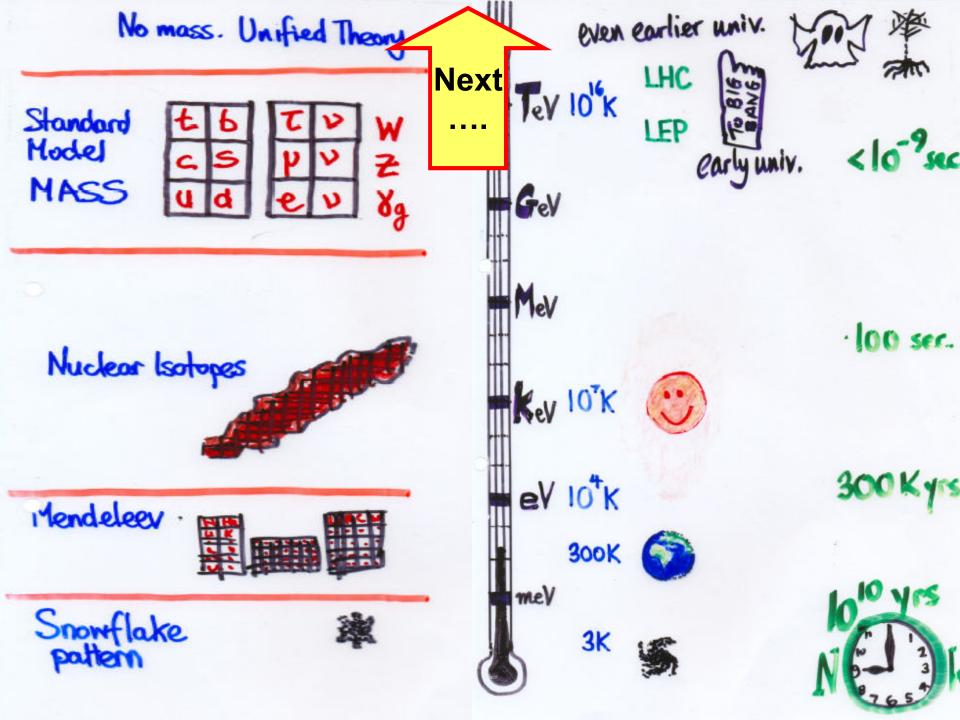
As the universe cooled after the hot big bang.....
We think that an elegant symmetry......

... "froze" into structures And patterns

Such as Atoms Mendeleev's periodic table, And particles Quarks, forces and the Standard Model

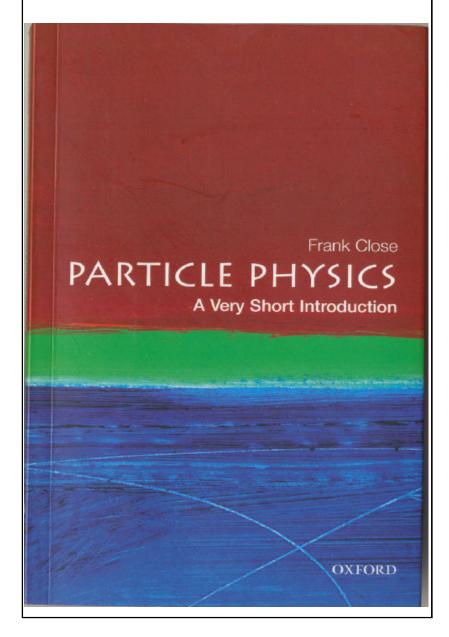
... which is a pattern based on MASS

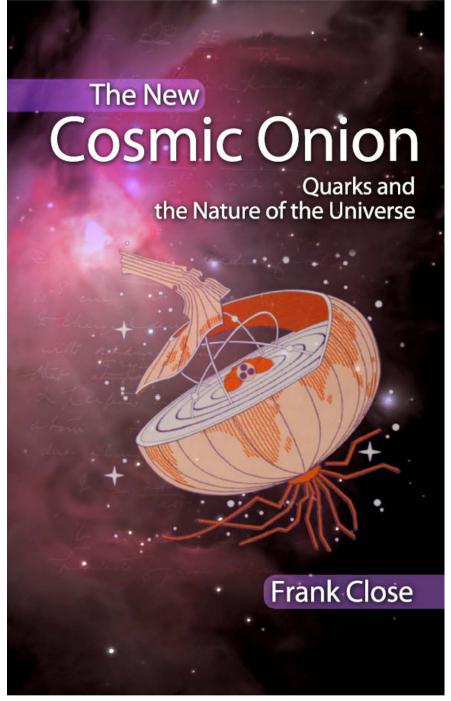
2008: heat up to energies above 1000 GeV = "1 TeV" and discover the origin of MASS (= Higgs?)





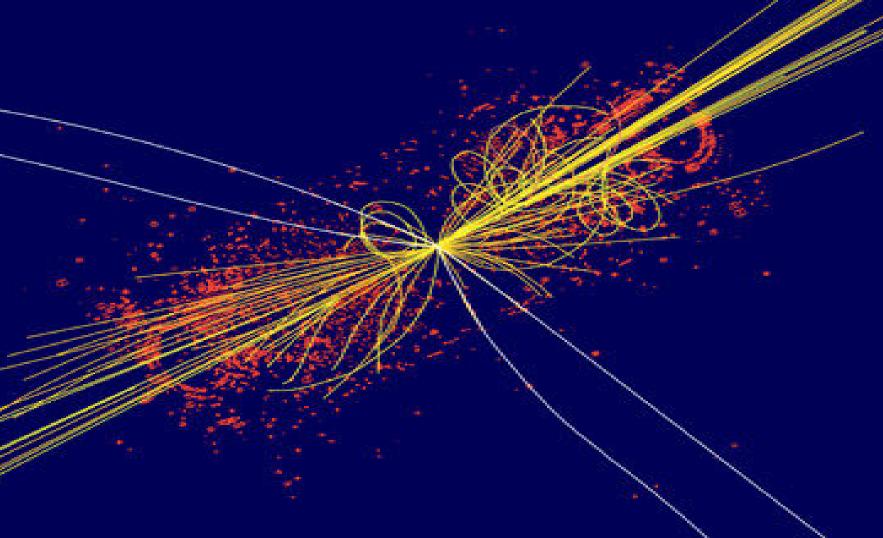
A Very Short Introduction





The Particle Odyssey

A Journey to the Heart of matter



Frank Close, Michael Marten, Christine Sutton

