

# Lepton Flavour Violation in SUSY-seesaw: an update

María J. Herrero

Dpt. Física Teórica/IFT, Universidad Autónoma, Madrid

From works

E.Arganda and M.H. PRD73,055003(2006)

S.Antusch, E.Arganda, M.H. and A.Teixeira JHEP 0611(2006)090

E.Arganda, M.H. and A.Teixeira hep-ph/0707.2955

E.Arganda, M.H. and J.Portolés IFT-UAM/CSIC-07-27

SUSY07, Karlsruhe, 27-July-2007

# Motivation

- ★ Lepton Flavour Violation (LFV) occurs in Nature,  $\nu_i - \nu_j$  oscill.
- ★ In SM: no LFV if  $m_\nu = 0$ ; very suppressed if  $m_\nu \neq 0$
- ★ Many exp. bounds (present/future sensitivities):  
MEGA, SINDRUM, BABAR, Belle / MEG,...**PRISM/PRIME**

$$\text{BR}(\mu \rightarrow e\gamma) < 1.2 \times 10^{-11}/10^{-13}$$

$$\text{BR}(\tau \rightarrow \mu\gamma) < 4.5 \times 10^{-8}/10^{-8}$$

$$\text{BR}(\tau \rightarrow e\gamma) < 1.2 \times 10^{-7}/10^{-8}$$

$$\text{BR}(\mu \rightarrow 3e) < 1.0 \times 10^{-12}/10^{-13}$$

$$\text{BR}(\tau \rightarrow 3\mu) < 1.9 \times 10^{-7}/10^{-8}$$

$$\text{BR}(\tau \rightarrow 3e) < 2.0 \times 10^{-7}/10^{-8}$$

.....

$$\text{BR}(\tau \rightarrow \mu\eta) < 6.5 \times 10^{-8}$$

$$\text{BR}(\tau \rightarrow \mu\eta') < 1.3 \times 10^{-7}$$

$$\text{BR}(\tau \rightarrow \mu\pi) < 1.2 \times 10^{-7}$$

$$\text{BR}(\tau \rightarrow \mu\rho) < 2 \times 10^{-7}$$

$$\text{CR}(\mu - e, \text{Au}) < 7 \times 10^{-13}$$

$$\text{CR}(\mu - e, \text{Ti}) < 4.3 \times 10^{-12}/\mathbf{10^{-18}}$$

....

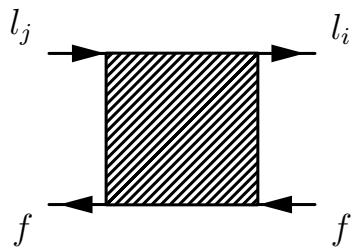
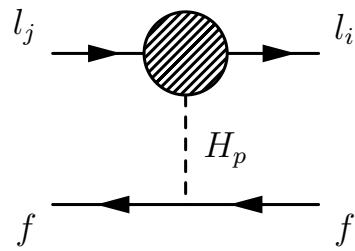
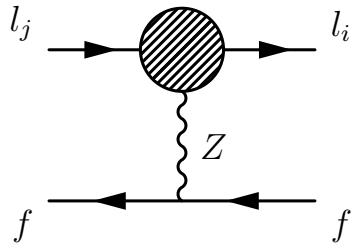
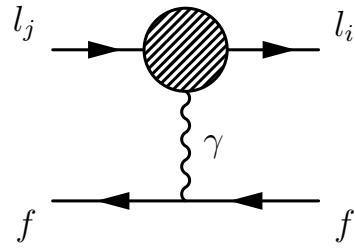
- ★ Measurement of LFV  $\Rightarrow$  Window for new physics
- ★ Very sensitive to SUSY: if Majorana  $\nu$ ,  $Y_\nu$  can be  $\mathcal{O}(1)$   
Large  $Y_\nu$  induce, via SUSY loops, large LFV rates
- ★ If no LFV found  $\Rightarrow$

**Restrictions on SUSY and/or seesaw parameters**

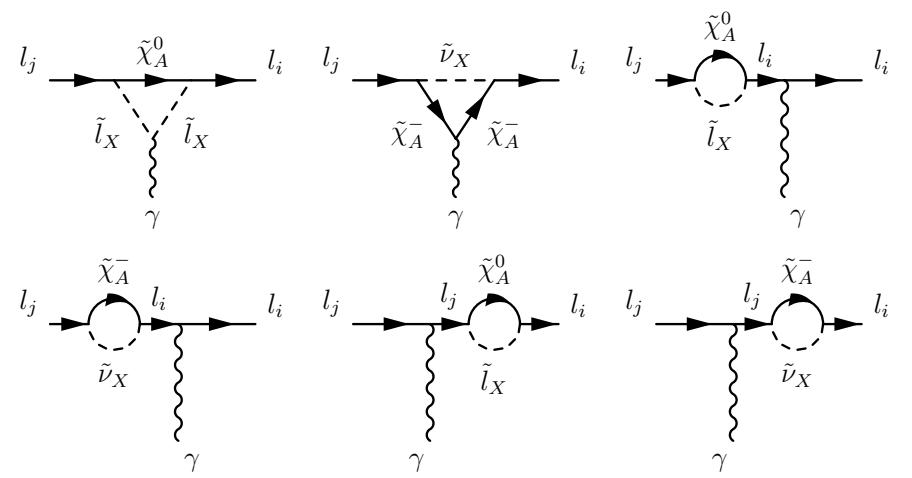
# Our Work

- **Prediction of LFV rates within SUSY-seesaw:**
  - ★ all  $l_j \rightarrow l_i \gamma$
  - ★ all  $l_j \rightarrow 3l_i$
  - ★ some semileptonic tau decays:  $\tau \rightarrow \mu\eta$ ,  $\tau \rightarrow \mu\pi$ ,  $\tau \rightarrow \mu\rho\dots$
  - ★  $\mu - e$  conversion in nuclei: Ti, Au,...
- **Full one-loop computation of LFV BRs**
- **Require compatibility with  $\nu$  data**
- **Compare with present/future LFV bounds/sensitivities**
- **Explore sensitivity to SUSY and seesaw parameters**
- **Study impact of  $\theta_{13}$  on LFV, specially  $\mu - e$**
- **Study of correlated/un-correlated processes: competing future sensitivities?**

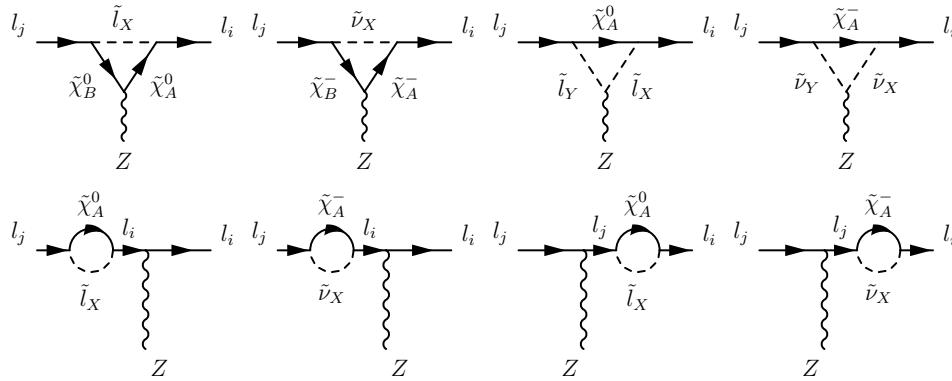
# 1-loop diagrams in $l_j \rightarrow 3l_i$ , $\tau \rightarrow \mu$ hadrons, $\mu - e$ conversion



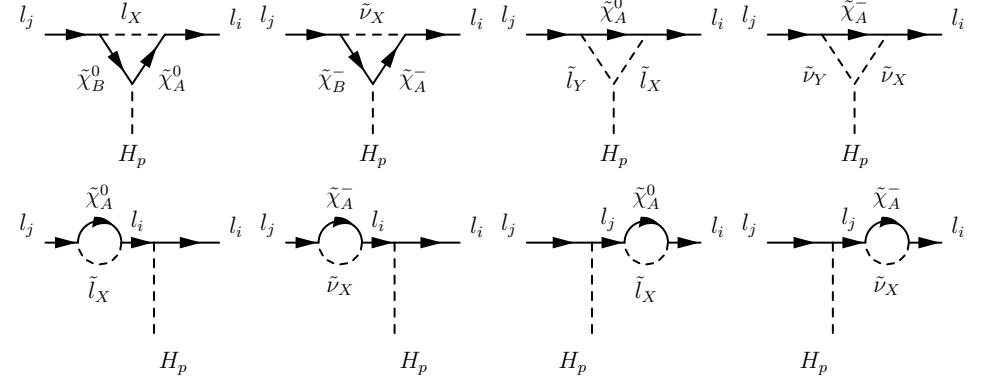
Generic



Photon-mediated



Z boson-mediated



H boson-mediated

# Framework

- Use seesaw (Type I) for  $\nu$  mass generation
- Work within CMSSM +  $3\nu_R$  (Majorana) +  $3\tilde{\nu}_R$

Two scenarios for soft parameters at  $M_X = 2 \times 10^{16}$  GeV:

- ★ Universal soft Higgs masses: CMSSM-seesaw  
 $(M_0, M_{1/2}, A_0, \tan \beta, \text{sign}(\mu))$
- ★ Non-universal soft Higgs masses: NUHM-seesaw  
 $(M_0, M_{1/2}, M_{H_1}, M_{H_2}, A_0, \tan \beta, \text{sign}(\mu))$
- LFV generated by 1-loop running from  $M_X$  to  $M_Z$   
Full RGEs including  $\nu$  and  $\tilde{\nu}$  sectors (No Llog approx)
- Mass eigenstates for all SUSY and Higgs particles (No MIA approx)
- Numerical estimates:
  - ★ SPheno 2.2.2 (W.Porod) for int. of RGEs and SUSY spectrum
  - ★ Additional subroutines for all LFV processes (by us)  
Also subroutines for checks of BAU, EDM and  $(g - 2)_\mu$

# Seesaw parameters versus neutrino data

SeeSaw equation:  $m_\nu = -m_D^T m_N^{-1} m_D$

**Solution:**  $m_D = i \sqrt{m_N^{diag}} \textcolor{red}{R} \sqrt{m_\nu^{diag}} U_{MNS}^\dagger$  [Casas, Ibarra ('01)]

$R$  is a  $3 \times 3$  complex matrix and orthogonal

$$R = \begin{pmatrix} c_2 c_3 & -c_1 s_3 - s_1 s_2 c_3 & s_1 s_3 - c_1 s_2 c_3 \\ c_2 s_3 & c_1 c_3 - s_1 s_2 s_3 & -s_1 c_3 - c_1 s_2 s_3 \\ s_2 & s_1 c_2 & c_1 c_2 \end{pmatrix}, c_i = \cos \theta_i, s_i = \sin \theta_i, \theta_{1,2,3} \text{ complex}$$

**Parameters:**  $\theta_{ij}, \delta, \alpha, \beta, m_{\nu_i}, m_{N_i}, \theta_i$  (18) ;  $m_{N_i}, \theta_i$  drive the size of  $Y_\nu$

Hierarchical  $\nu$ 's :  $m_{\nu_1}^2 \ll m_{\nu_2}^2 = \Delta m_{\text{sol}}^2 + m_{\nu_1}^2 \ll m_{\nu_3}^2 = \Delta m_{\text{atm}}^2 + m_{\nu_1}^2$

2 Scenarios

- Degenerate  $N$ 's  
 $m_{N_1} = m_{N_2} = m_{N_3} = m_N$
- Hierarchical  $N$ 's  
 $m_{N_1} \ll m_{N_2} \ll m_{N_3}$

# Our choice of input parameters

Constrained MSSM + $3\nu_R$  +  $3\tilde{\nu}_R$  + seesaw

- CMSSM:

$$\left\{ \begin{array}{l} M_0, M_{1/2}, A_0 \text{ (at } M_X \sim 2 \times 10^{16} \text{ GeV)} \\ \tan \beta = \langle H_2 \rangle / \langle H_1 \rangle \text{ (at EW scale)} \\ \text{sign}(\mu) \text{ (\mu derived from EW breaking)} \end{array} \right\} \text{Choose SPS points}$$

- NUHM:  $(M_0, M_{1/2}, M_{H_1}, M_{H_2}, A_0, \tan \beta, \text{sign}(\mu))$

**Choose**  $M_0 = M_{1/2}$ ,  $M_{H_1}^2 = M_0^2(1 + \delta_1)$ ,  $M_{H_2}^2 = M_0^2(1 + \delta_2)$

- Seesaw parameters  $\left\{ \begin{array}{l} m_{\nu_{1,2,3}} \text{ (set by data)} \\ m_{N_{1,2,3}} \text{ (input)} \\ U_{MNS} \text{ (set by data)} \\ R(\theta_1, \theta_2, \theta_3) \text{ (input)} \end{array} \right\}$

- For numerical estimates:

$$(\Delta m^2)_{12} = \Delta m_{\text{sol}}^2 = 8 \times 10^{-5} \text{ eV}^2$$

$$(\Delta m^2)_{23} = \Delta m_{\text{atm}}^2 = 2.5 \times 10^{-3} \text{ eV}^2$$

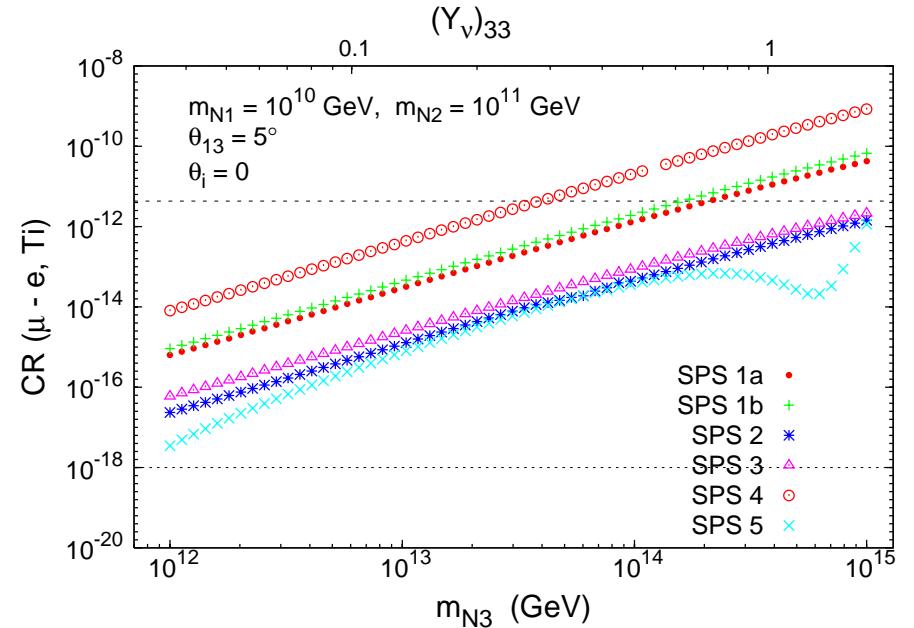
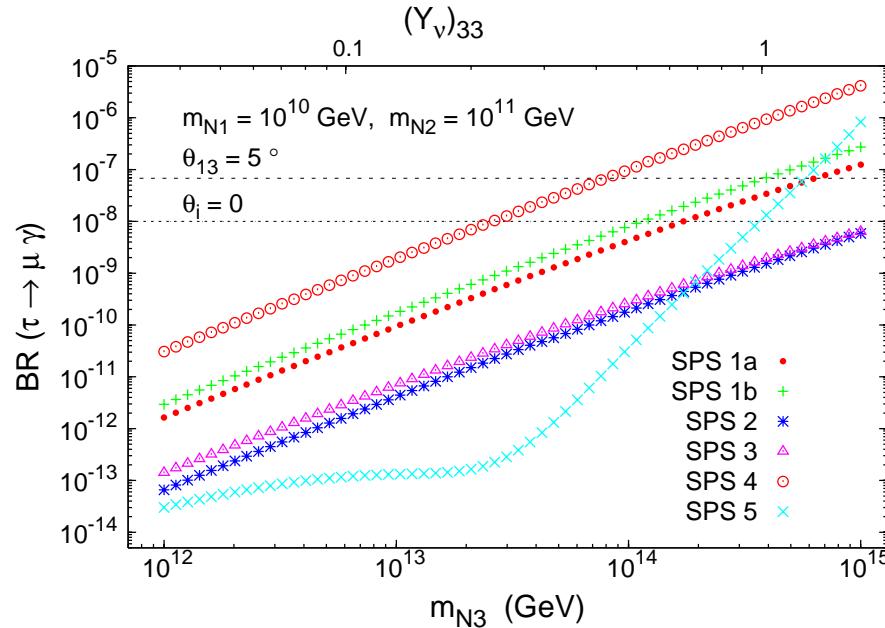
$$\theta_{12} = 30^\circ; \theta_{23} = 45^\circ; \delta = \alpha = \beta = 0; 0 \leq \theta_{13} \leq 10^\circ$$

$$250 \text{ GeV} < M_0, M_{1/2} < 1000 \text{ GeV}, -500 \text{ GeV} < A_0 < 500 \text{ GeV}$$

$$5 < \tan \beta < 50, -2 < \delta_{1,2} < 2$$

# **Results for CMSSM-seesaw**

# The most relevant parameter: Hierarchical: $m_{N_3}$ / Degenerate: $m_N$

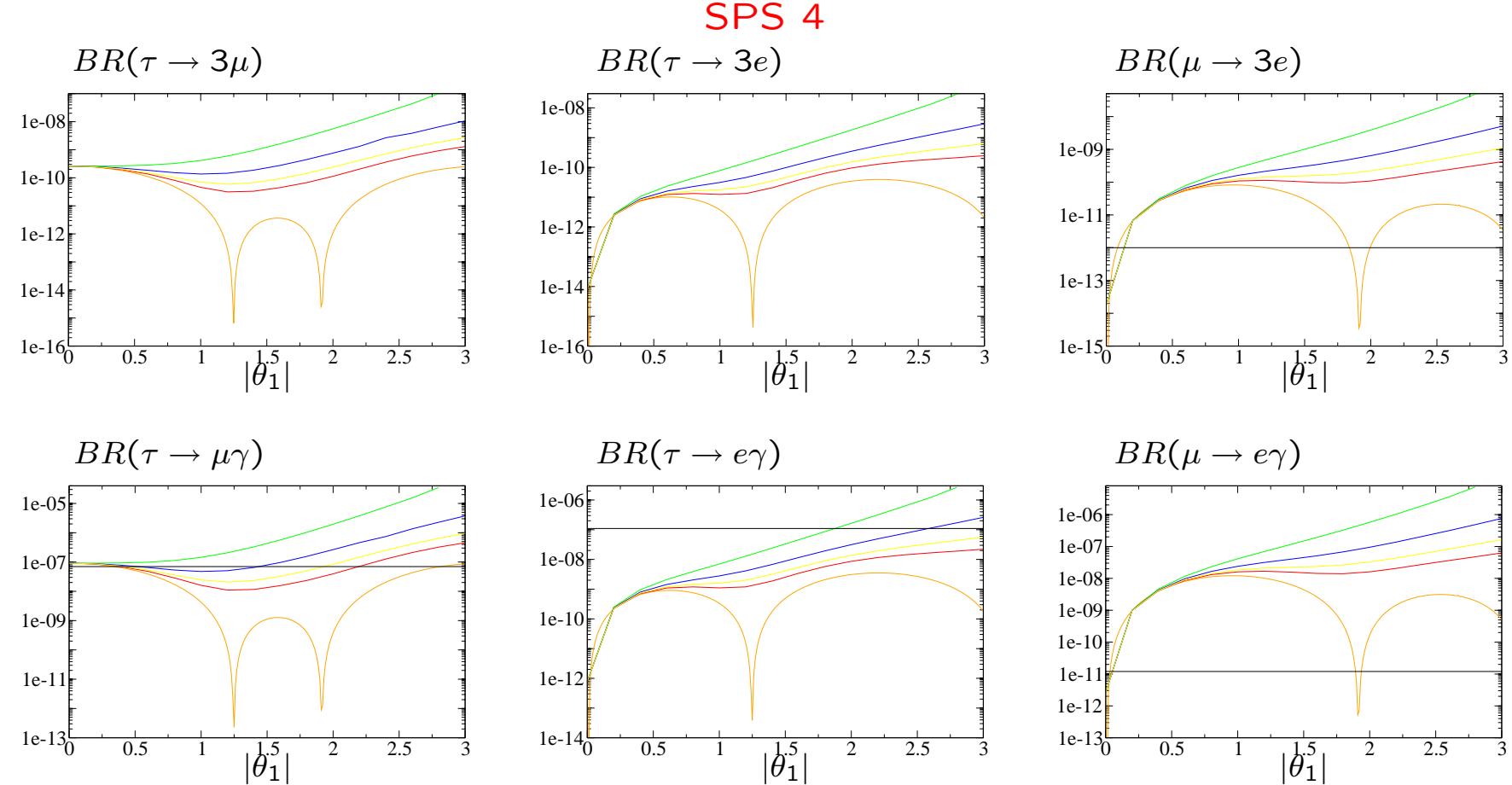


- ★ Most obs. reach exp. lim. at  $m_{N_3} \in [10^{13}, 10^{15}]$  GeV,  $(Y_\nu)_{33,32} \sim 0.1 - 1$
- ★  $BR_4 (\tan \beta = 50) > BR_{1b} \gtrsim BR_{1a} > BR_3 \gtrsim BR_2 > BR_5$
- ★  $BR \sim |m_{N_3} \log m_{N_3}|^2$  except for SPS5: Llog fails in  $\sim 10^4$ !!
- ★ Present: the most restrictive one is  $\mu \rightarrow e\gamma$  (if  $\theta_{13} \neq 0$ )  
Bounds for SPS1a  $m_{N_3} < 10^{13} - 10^{14}$  GeV  
Next are  $\mu - e$ ,  $\mu \rightarrow 3e$ ;  $\tau \rightarrow \mu\gamma$  competitive if  $\theta_{13} \simeq 0$  and N's hier.
- ★ Future:  $\mu - e$  conversion the best: sensitive to  $m_{N_3} > 10^{12}$  GeV

# Large Yukawa couplings: role of $\theta_i$

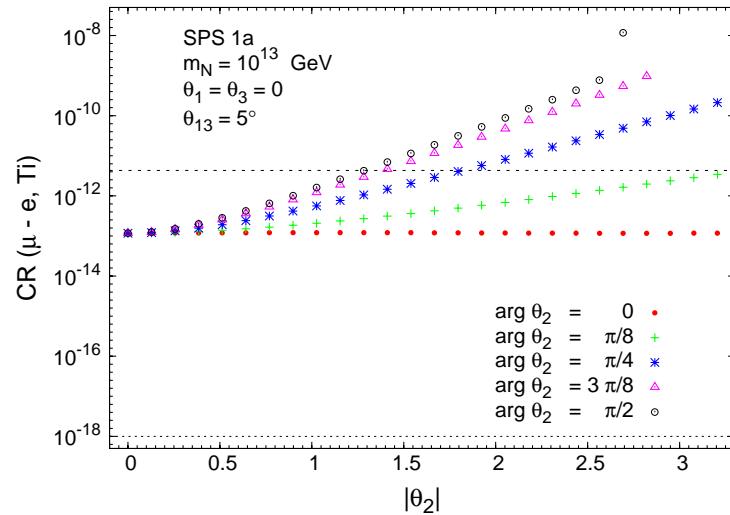
Hierarchical  $m_{N_i}$  and complex  $\theta_i$

$(m_{N_1}, m_{N_2}, m_{N_3}) = (10^8, 2 \times 10^8, 10^{14})$  GeV,  $\arg(\theta_1) = 0, \pi/10, \pi/8, \pi/6, \pi/4$  ( $\theta_2 = \theta_3 = 0$ )

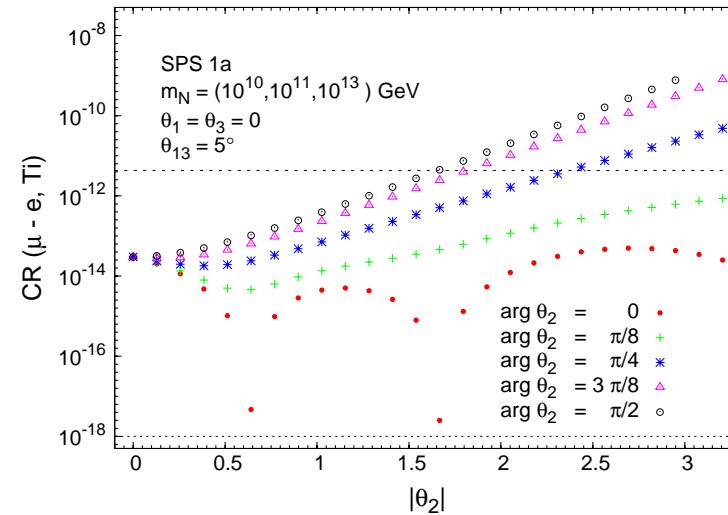


- ★ BRs for  $0 < |\theta_i| < \pi$ ,  $0 < \arg \theta_i < \pi/2$  can increase up to  $10^2 - 10^4$  respect to  $\theta_i = 0$
- ★ **BRs above present experimental bounds:** mainly  $\mu \rightarrow e\gamma$ ,  $\mu \rightarrow 3e$  and  $\tau \rightarrow \mu\gamma$
- ★ Similar results for  $\theta_2$ . BRs nearly constant with  $\theta_3$  in the case of hier. N's

# Role of $\theta_i$ (cont.)

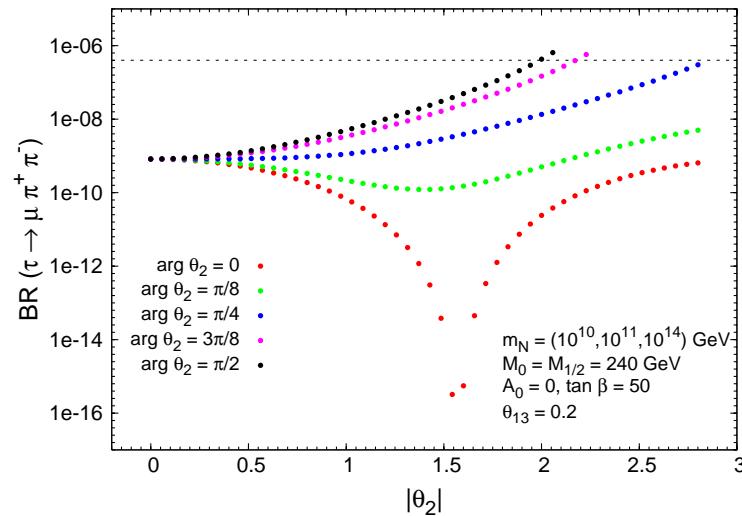


N's Deg. Eq. dep. all  $\theta_i$



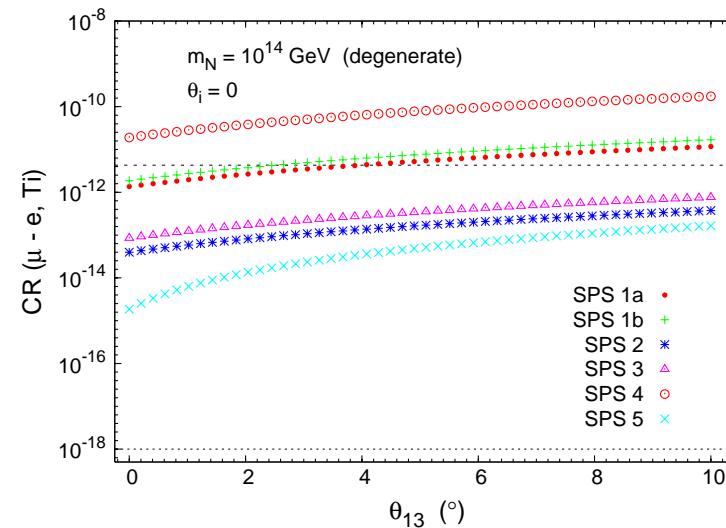
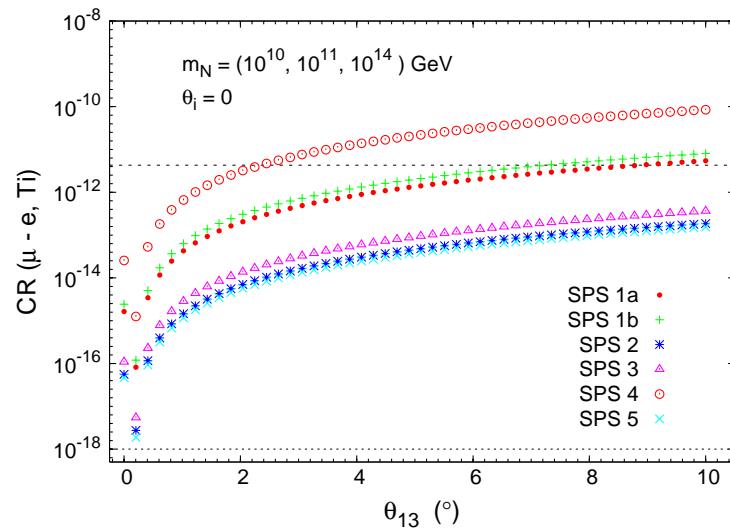
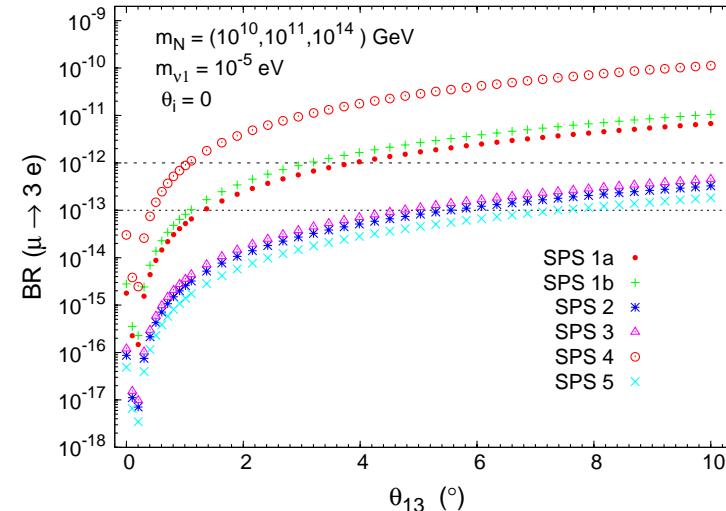
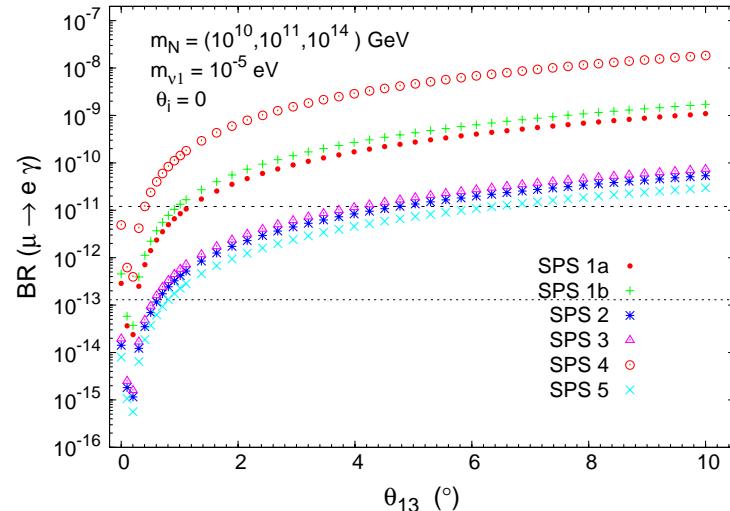
N's Hier. relevant:  $\theta_{1,2}$

- ★  $CR(\mu - e)$  reach exp. bound even for  $m_N = 10^{13}$  GeV if complex  $\theta_i \neq 0$



- ★  $BR(\tau \rightarrow \mu \pi^+ \pi^-)$  reach exp. bound at high  $m_{N_3} \sim 10^{14}$  GeV and large complex  $\theta_{1,2}$

# High sensitivity to $\theta_{13}$ : the case $\theta_i = 0$

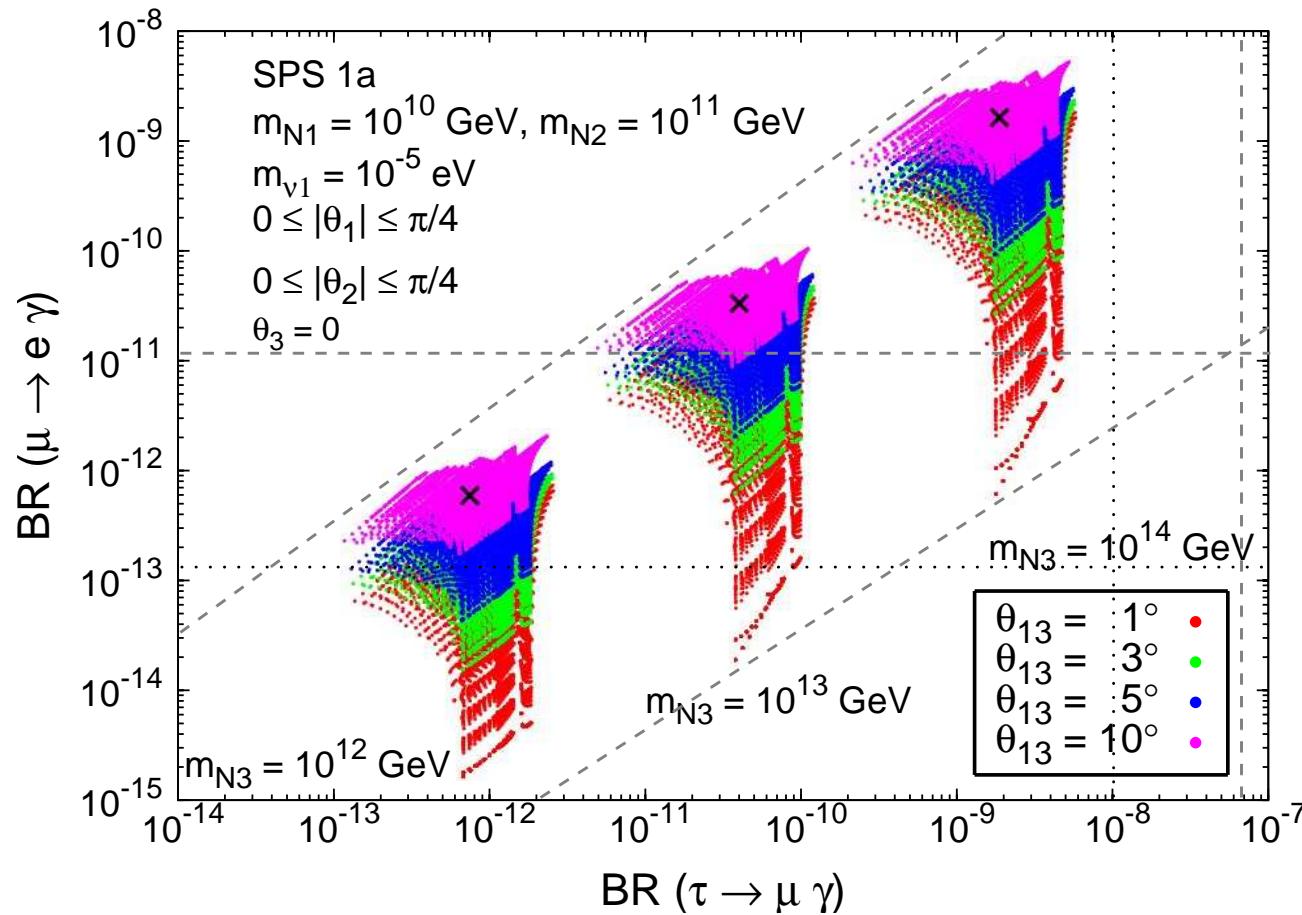


- ★  $\mu \rightarrow e\gamma, \mu \rightarrow 3e, \mu - e$  extremely sensitive to  $\theta_{13}$  if Hier.N's: BRs  $\times 10^5$  in  $0^\circ < \theta_{13} < 10^\circ$
- ★  $\tau \rightarrow e\gamma, \tau \rightarrow 3e$  also, but not within exp.reach ( $\tau \rightarrow \mu\gamma, \tau \rightarrow 3\mu$  are not!!)
- ★ Sensitivity of  $\mu \rightarrow e\gamma$  clearly **within exp. reach**:  $\text{BR}_{\text{all SPS}} > \text{BR}_{\text{exp}}^{\text{present}}$  for  $\theta_{13} \gtrsim 5^\circ$  !!

# Impact of $\theta_{13}$ on LFV processes

(All plotted points lead to 'viable BAU', respect EDM bounds, OK with  $(g - 2)_\mu$ )

$$(-\pi/4 \lesssim \arg\theta_1 \lesssim \pi/4, 0 \lesssim \arg\theta_2 \lesssim \pi/4)$$

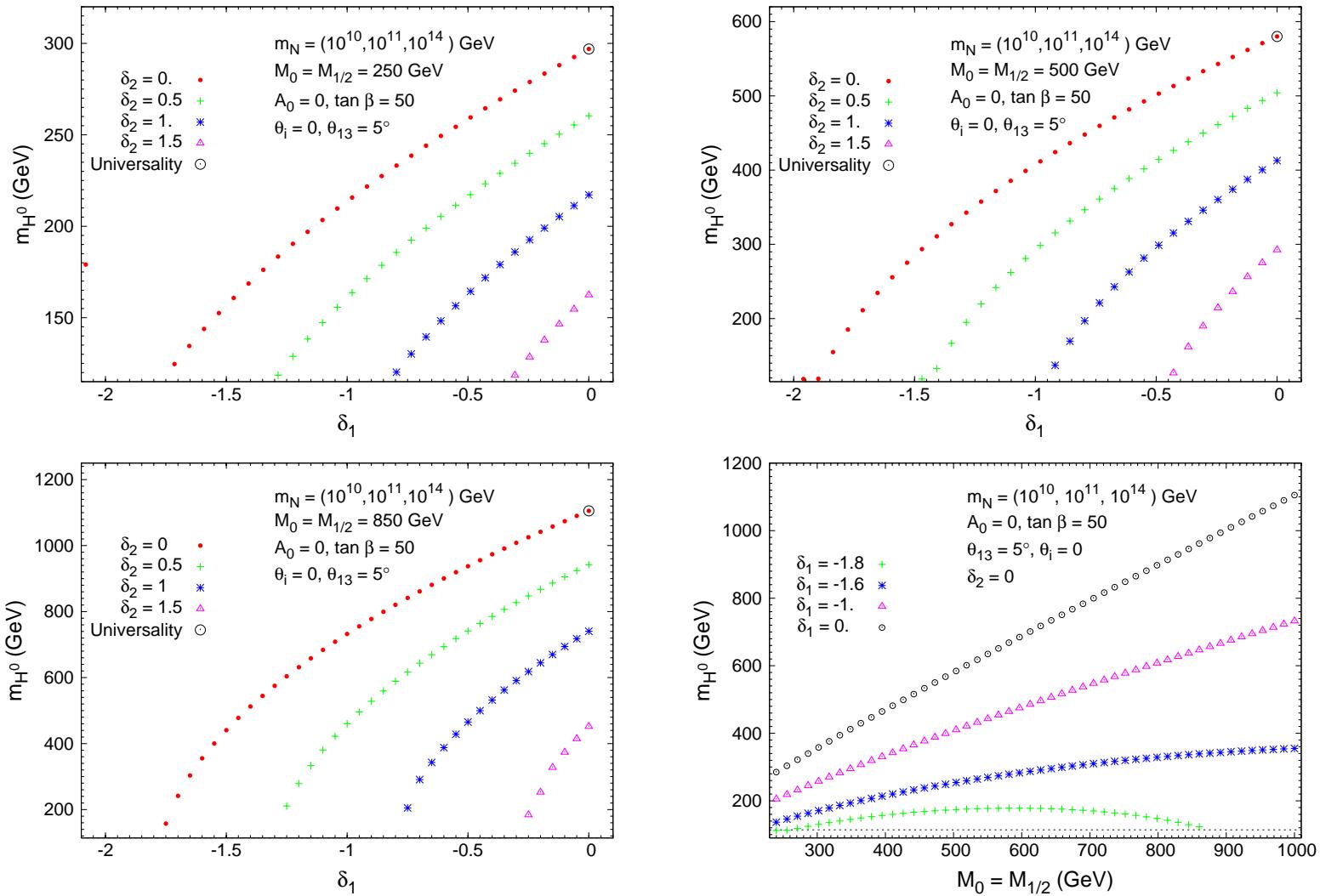


Correlations in LFV BRs help: similar dependence on  $m_{N_3}$  and  $\tan\beta$   
MEGA bound already disfavours  $\theta_{13} \gtrsim 10^\circ (2^\circ)$  for  $m_{N_3} \gtrsim 10^{13} (10^{14})$  GeV  
A measurement of BRs and  $\theta_{13}$  will provide some insight into  $m_{N_3}$  !!

# Results for NUHM-seesaw

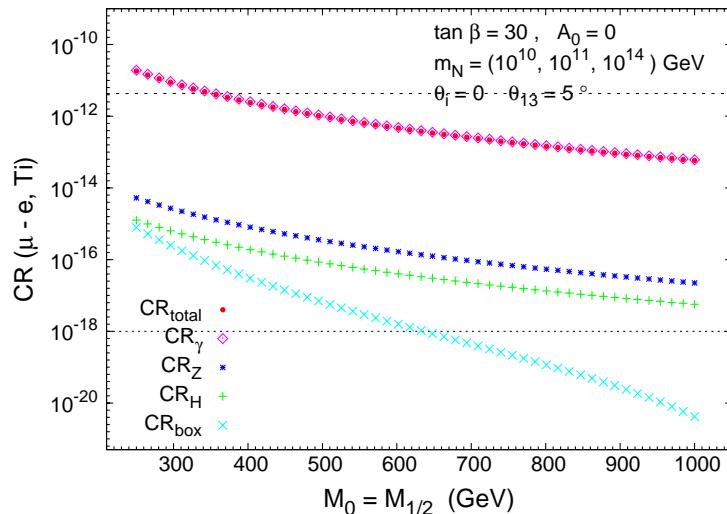
- Looking for solutions with light Higgs sector as functions of non-universal soft parameters:  $M_{H_{1,2}}^2 = M_0^2(1 + \delta_{1,2})$ ,  $-2 < \delta_{1,2} < 2$
- Study of correlation loss between related LFV observables due to Higgs-mediated contributions that can dominate the photon-mediated ones  
Relevant for:
  - ★  $\mu - e$  conversion in nuclei versus  $\mu \rightarrow e\gamma$
  - ★  $\tau \rightarrow \mu\eta$  versus  $\tau \rightarrow \mu\gamma$
  - ★  $\tau \rightarrow \mu\eta$  versus  $\tau \rightarrow \mu\rho$due to large Higgs couplings to strange quarks (not relevant for  $l_j \rightarrow 3l_i$ )
- Explore ratios of observables looking for enhancements respect to the universal case

# Light Higgs sector in NUHM-seesaw

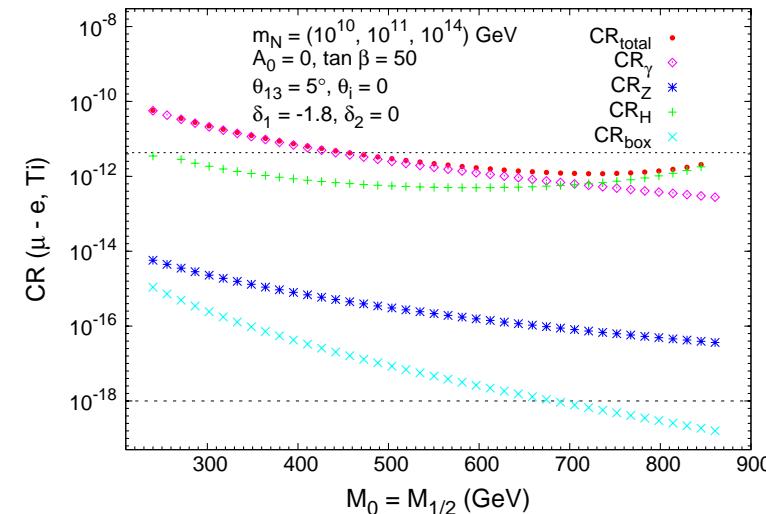


- ★ We find solutions with **light Higgs particles** even for large  $M_0 = M_{1/2} = M_{\text{SUSY}}$
- ★ Ex.: for  $M_{\text{SUSY}} = 850$  GeV,  $\tan \beta = 50$ ,  $A_0 = 0$ ,  $\delta_1 = -1.8$ ,  $\delta_2 = 0$ , we find:  
light Higgs:  $m_{H^0} = 127$  GeV,  $m_{h^0} = 123$  GeV,  $m_{A^0} = 127$  GeV,  $m_{H^+} = 155$  GeV  
heavy SUSY:  $m_{\tilde{l}_1} = 734$  GeV,  $m_{\tilde{\nu}_1} = 971$  GeV,  $m_{\tilde{\chi}_1^-} = 687$  GeV,  $m_{\tilde{\chi}_1^0} = 362$  GeV

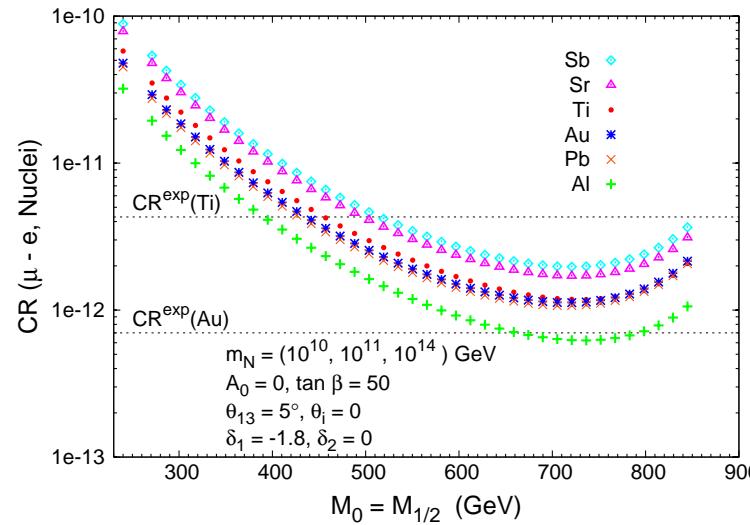
# $\mu - e$ conversion in nuclei: universality versus non-universality



universality:  $\gamma$  dominance

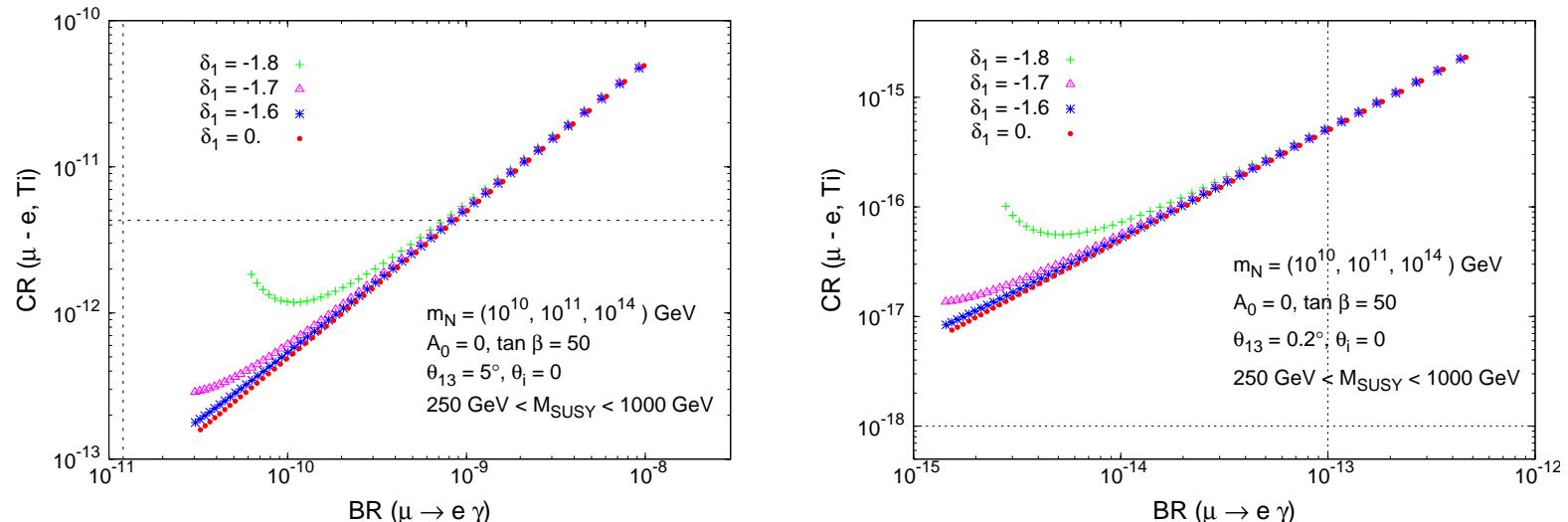


non-universality:  $H^0$  dominance

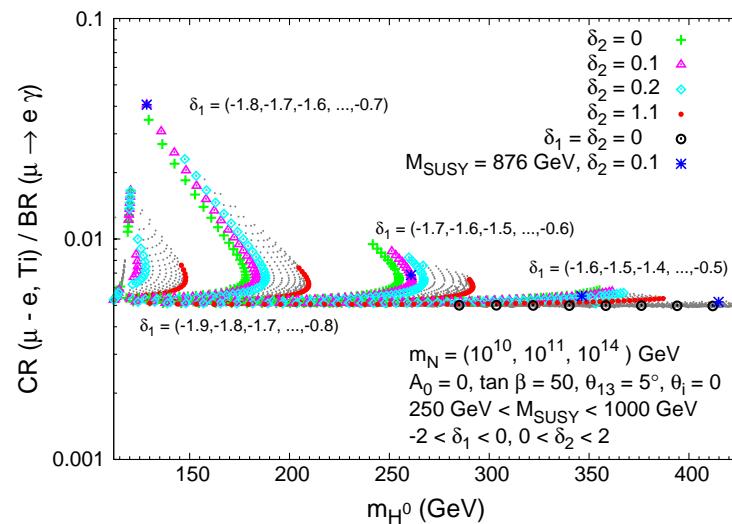


- ★  $H^0$  dominance  $\Rightarrow$  heavy SUSY spectra do not decouple in  $\mu - e$  conversion
- $CR(\mu - e, Au)$  above present experimental bound even for heavy SUSY

# Correlation loss between $\mu - e$ conv. and $\mu \rightarrow e\gamma$

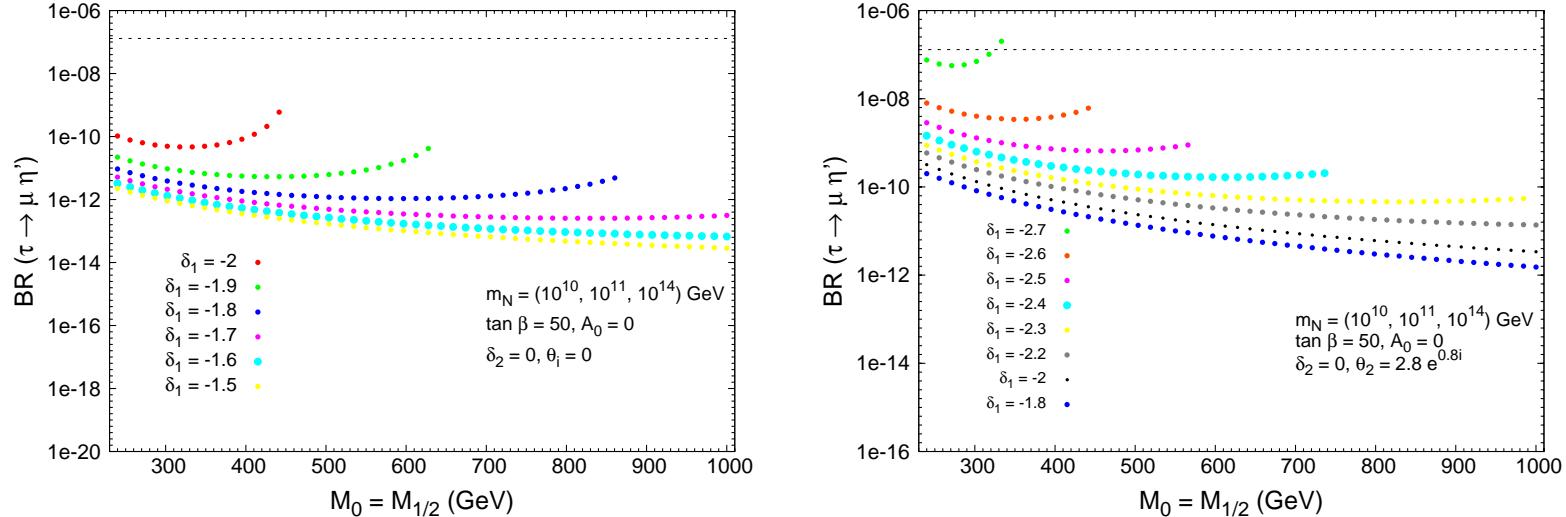


- ★ Deviation from linearity due to Higgs-dominated rates. Visible at future?

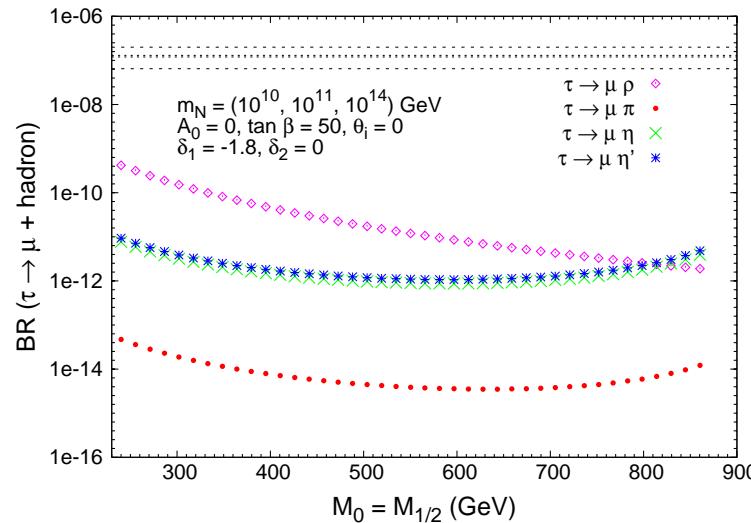


- ★ Ratio of  $\mu - e$  to  $\mu \rightarrow e\gamma$  can be a factor 10 larger than in photon-dominated scenarios like CMSSM (where this ratio is  $\mathcal{O}(\alpha)$ )

# Higgs-dominance in semileptonic tau decays



- ★ Heavy SUSY do not decouple in  $\tau \rightarrow \mu \eta'$  for scenarios with Higgs-dominance
- $\tau \rightarrow \mu \eta'$  decay rates close to exp. bound if large  $\theta_i$



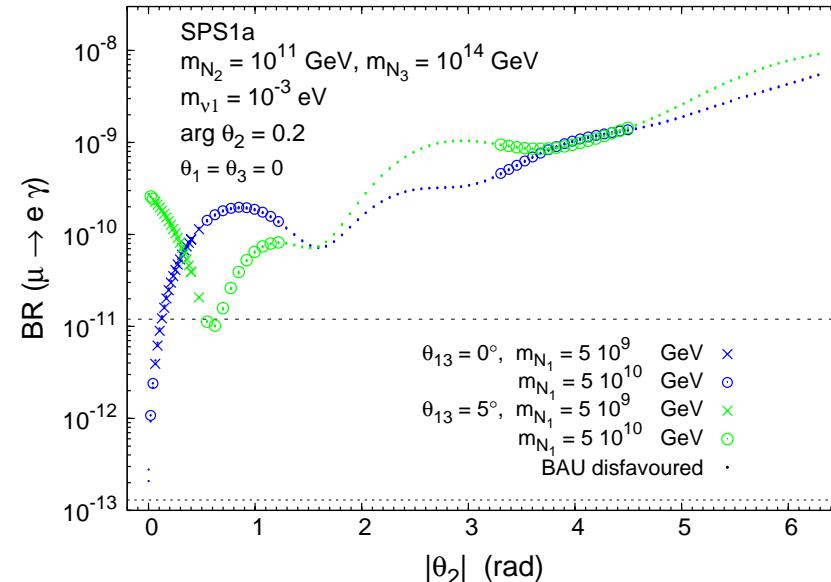
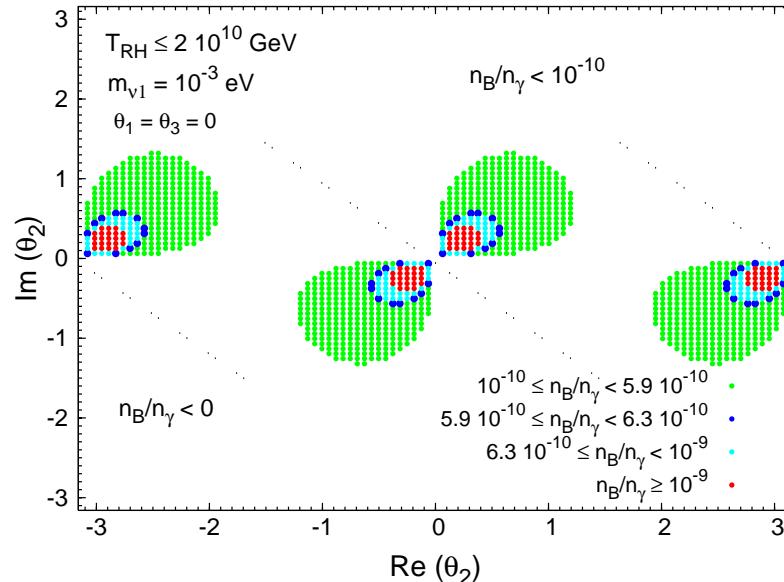
- ★ Ratio of  $\tau \rightarrow \mu \eta$  to  $\tau \rightarrow \mu \rho$  can be a factor 10 larger than in CMSSM

**If SUSY exists:**  
**LFV observables constitute an interesting lab that together with low-energy  $\nu$ -data can provide some insight into the heavy neutrino sector and SeeSaw parameters. Also into Higgs sector if NUHM like scenarios**

# **Additional transparencies**

# Constraints from 'viable' BAU

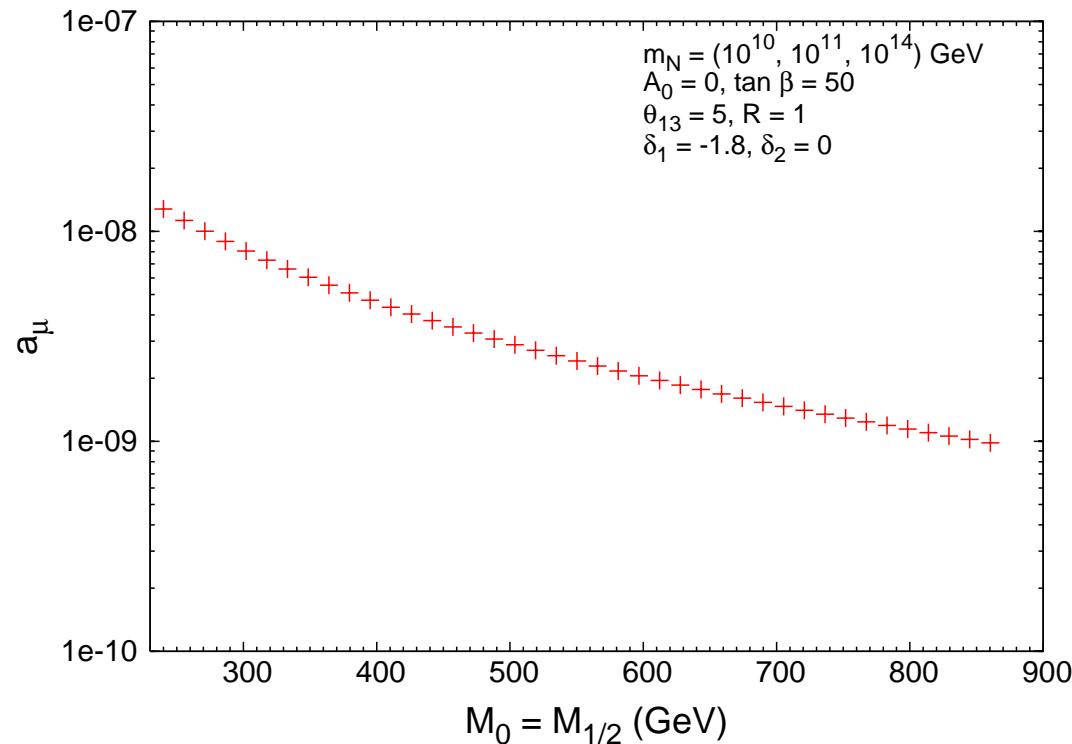
BAU requires complex  $R \neq 1 \Rightarrow$  complex  $\theta_i \neq 0$ . Most relevantly  $\theta_2$   $n_B/n_\gamma \in$  interval  $\Rightarrow (\text{Re}(\theta_2), \text{Im}(\theta_2)) \in$  area ('ring') (WMAP in darkest ring)



## Implications for LFV

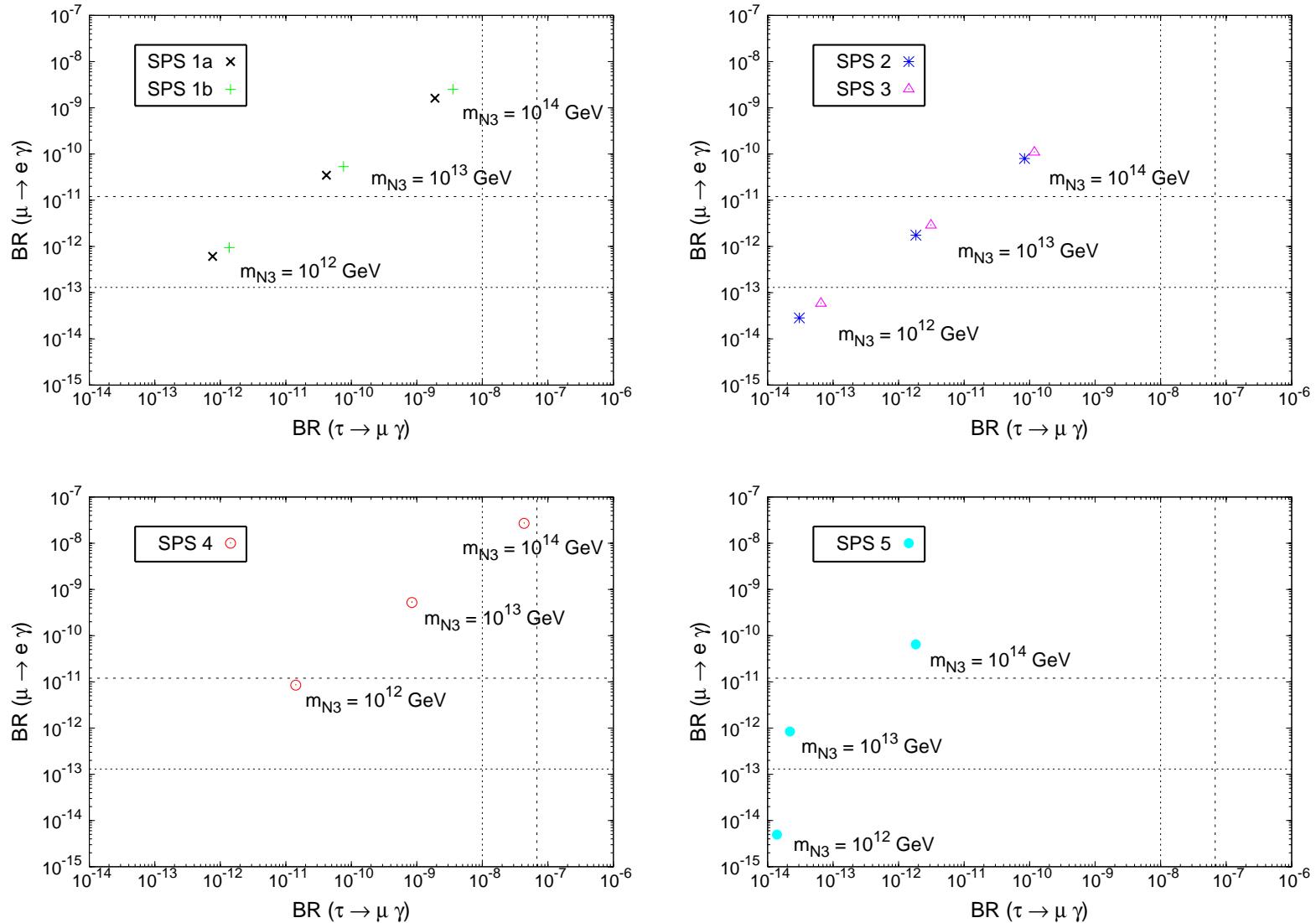
- \* **'viable' BAU**  $\leftrightarrow n_b/n_\gamma \in [10^{-10}, 10^{-9}]$  (WMAP  $\sim 6.1 \times 10^{-10}$ , '06)
  - BAU [disfav]-[fav]-[disfav]-[fav]-[disfav] pattern in  $0 < |\theta_2| < 3$
  - The BAU [fav] windows occur at small ( $\neq 0$ )  $|\theta_2| \lesssim 1.5$
- \* **smaller  $|\theta_2| \Rightarrow$  smaller LFV rates**
- \* The existence, location and size of the windows depend on  $m_{N_1}$ 
  - $m_{N_1} \sim O(10^{10})$  GeV BAU [fav] windows at  $|\theta_2| \sim O(1)$  and  $|\theta_2| \sim O(10^{-2})$
  - $m_{N_1} \sim O(10^9)$  GeV only one window at  $|\theta_2| \sim O(5 \times 10^{-1})$

# Contributions to $\Delta a_\mu^{\text{SUSY}}$



$\Delta a_\mu^{\text{SUSY}} \in [10^{-8}, 10^{-9}]$ : compatible with  $a_\mu^{\text{EXP}} - a_\mu^{\text{SM}} = 3.32 \times 10^{-9} (3.8\sigma)$

# Impact of $\theta_{13}$ on LFV: Other SPS points

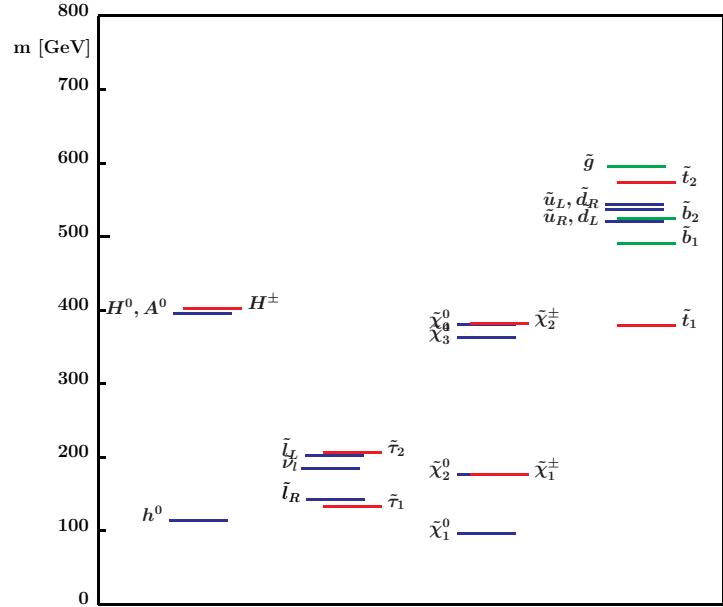


Similar for SPS1a,1b. Slightly worse prospects for SPS2,3. SPS5 the worst.

SPS4 the most restrictive one (due to  $\tan \beta = 50$ ):

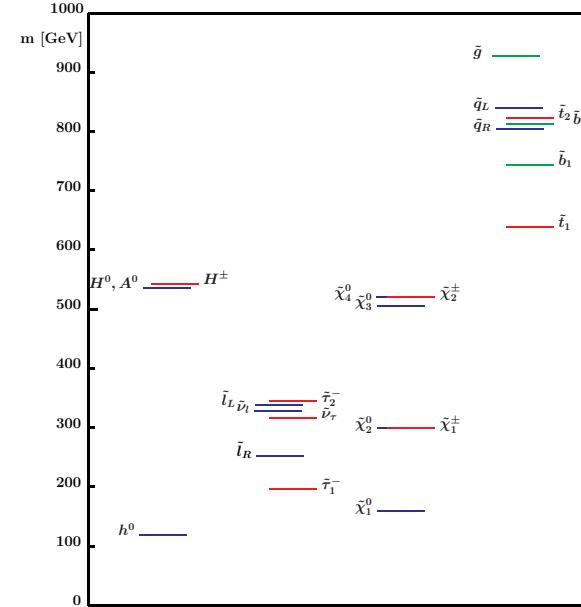
Present bounds from  $\mu \rightarrow e\gamma$  and  $\tau \rightarrow \mu\gamma$  already exclude  $m_{N_3} \gtrsim 10^{14}$  GeV!!

# SUSY SPS points (I)



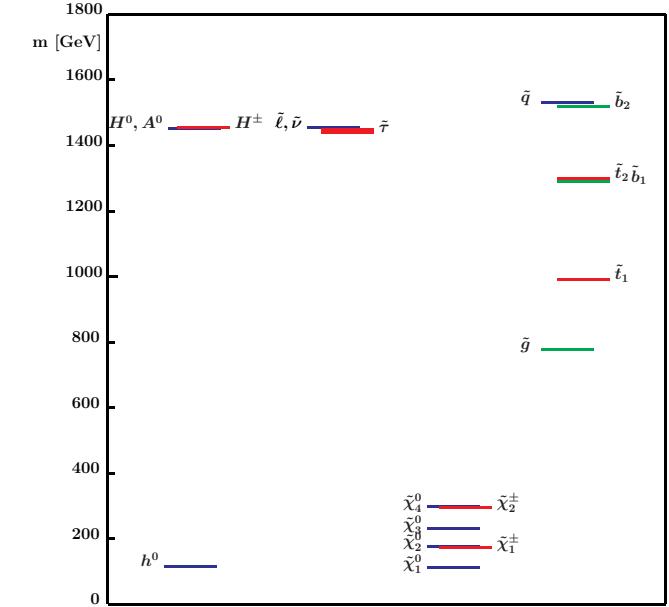
SPS1a

$M_0 = 100 \text{ GeV}$   
 $M_{1/2} = 250 \text{ GeV}$   
 $A_0 = -100 \text{ GeV}$   
 $\tan \beta = 10$   
 $\mu > 0$



SPS1b

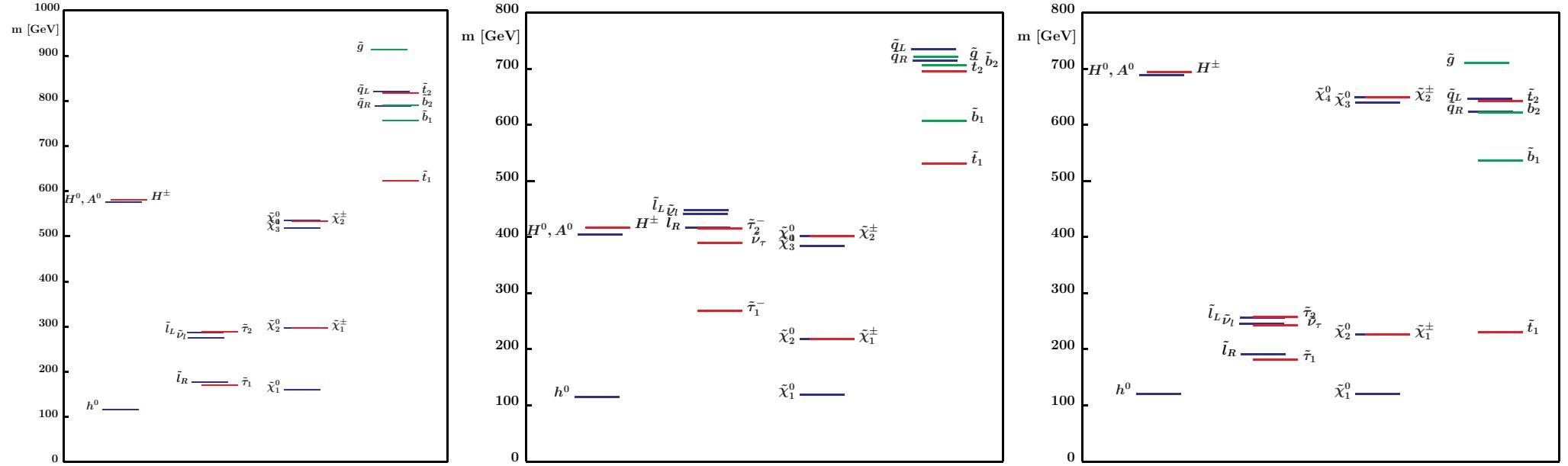
$M_0 = 200 \text{ GeV}$   
 $M_{1/2} = 400 \text{ GeV}$   
 $A_0 = 0 \text{ GeV}$   
 $\tan \beta = 30$   
 $\mu > 0$



SPS2

$M_0 = 1450 \text{ GeV}$   
 $M_{1/2} = 300 \text{ GeV}$   
 $A_0 = 0 \text{ GeV}$   
 $\tan \beta = 10$   
 $\mu > 0$

# SUSY SPS points (II)



**SPS3**

$M_0 = 90 \text{ GeV}$   
 $M_{1/2} = 300 \text{ GeV}$   
 $A_0 = 0 \text{ GeV}$   
 $\tan \beta = 10$   
 $\mu > 0$

**SPS4**

$M_0 = 400 \text{ GeV}$   
 $M_{1/2} = 300 \text{ GeV}$   
 $A_0 = 0 \text{ GeV}$   
 $\tan \beta = 50$   
 $\mu > 0$

**SPS5**

$M_0 = 150 \text{ GeV}$   
 $M_{1/2} = 300 \text{ GeV}$   
 $A_0 = -1000 \text{ GeV}$   
 $\tan \beta = 5$   
 $\mu > 0$