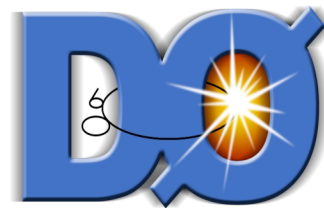


Search for Gauge-Mediated Supersymmetry

Yuri Gershtein



Tevatron Collider



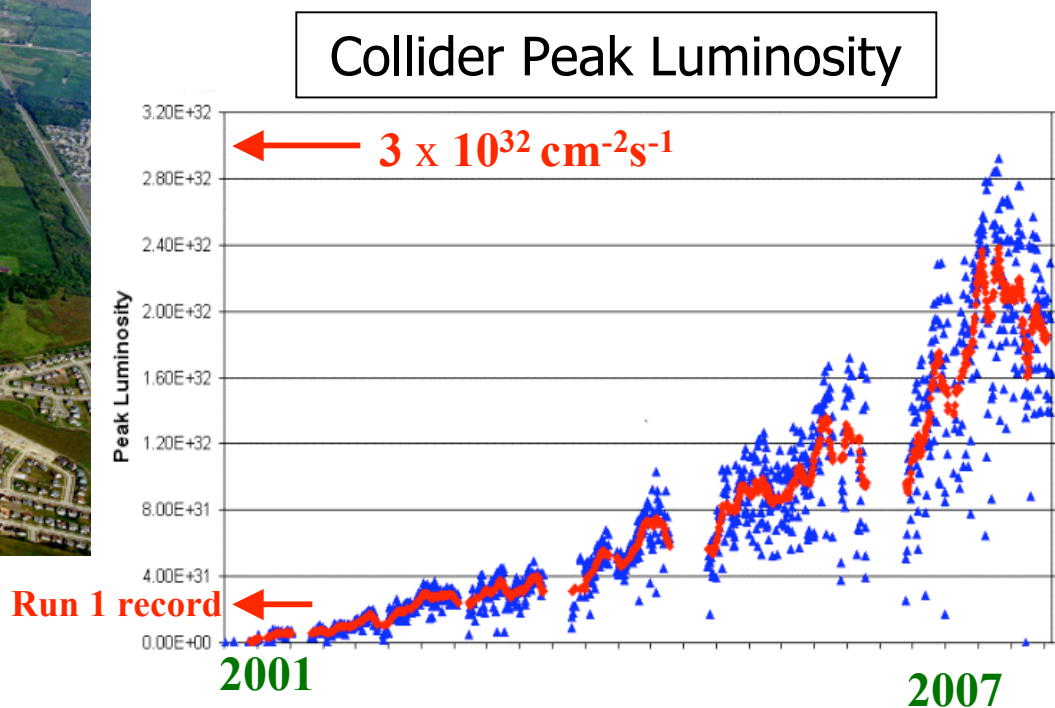
1992-95

Run 1: 100 pb^{-1} , 1.8 TeV

2001-2009 Run 2: major upgrades

higher $E_{\text{CM}} = 1.96 \text{ TeV}$

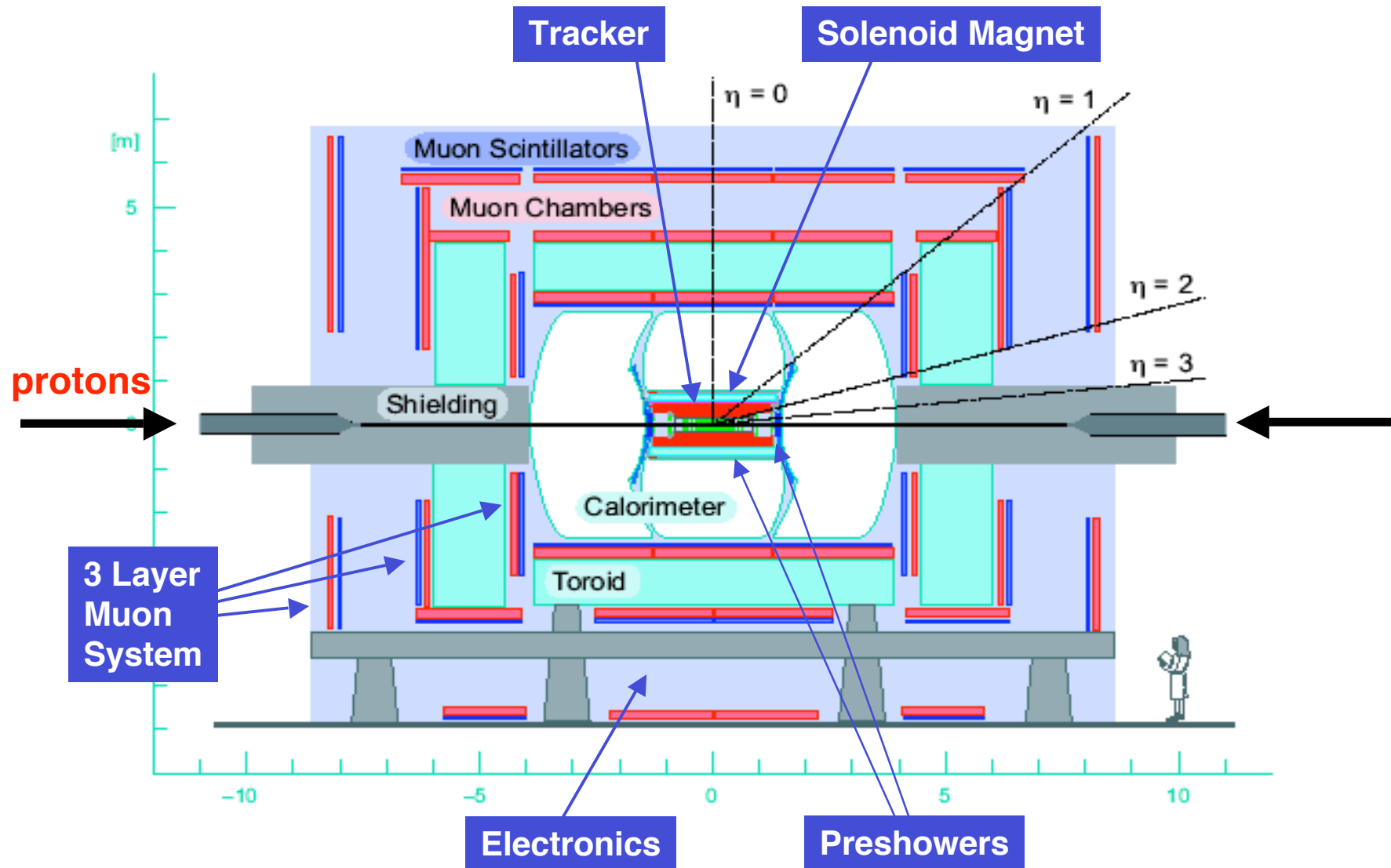
$\sim 3 \text{ fb}^{-1}$ recorded, expect $\sim 8 \text{ fb}^{-1}$ in 2009



7/28/2007

Yuri Gershtein, SUSY07

DØ Detector



7/28/2007

Yuri Gershtein, SUSY07

Gauge-Mediated SUSY Breaking

- Alternative to gravity mediated SUSY breaking
 - SUSY breaking occurs at scale Λ much lower than GUT (10 – 100 TeV)
 - Mediated by new gauge fields – “messengers”
 - Gravitino is very light ($< \text{keV}$) and is LSP
 - Dark Matter is a mix of the gravitino and the lightest “messenger”
- Lifetime of NLSP is a free parameter
 - All SUSY particles cascade to NLSP, so if R-parity is conserved all final states have two NLSPs
- NLSP can be neutralino or stau:
 $\chi_1^0 \rightarrow \gamma G$ or $\tau_1^+ \rightarrow \tau^+ G$
 - If neutralino is higgsino-like then it can also decay to hG and ZG
 - Gravitino is weakly interacting and is registered as missing transverse energy (MET)

Typical Final States

	prompt decays	inside detector	outside detector
Stau	multiple τ + MET	tracks with kinks	muon-like highly ionising slow moving particles
Neutralino	$\gamma\gamma$ + MET	non-pointing photon(s)	mSUGRA-like
Higgsino	γ bb + MET, γ Z+MET	Non-pointing b-jets, detached Z	mSUGRA-like

- Many of the final states are predicted by other models (extra dimensions, 4th generation quarks, etc...)

mGMSB

● Model Parameters:

- Λ - mass parameter (effective scale of SUSY breaking)
- M_m - messenger mass scale
- N_5 - number of messenger fields
- $\tan \beta = \langle \phi_1 \rangle / \langle \phi_2 \rangle$
- $\text{sign } \mu = \pm 1$ - sign of higgsino mass term
- C_{grav} - determines NLSP lifetime

● Gaugino masses proportional to N_5 , scalar masses proportional to $\sqrt{N_5}$

● Snowmass slope E:

- Λ - varies
 - $M_m = 2 \cdot \Lambda$
 - $N_5 = 1$
 - $\tan \beta = 15$
 - $\text{sign } \mu = +1$
- } Neutralino NLSP

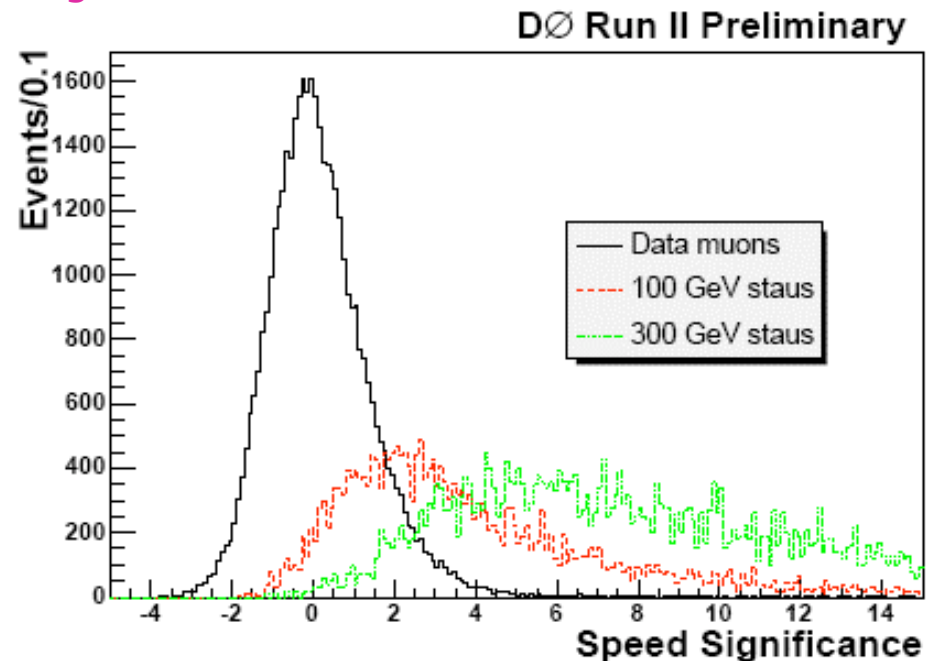
● Snowmass slope D:

- Λ - varies
 - $M_m = 2 \cdot \Lambda$
 - $N_5 = 3$
 - $\tan \beta = 15$
 - $\text{sign } \mu = +1$
- } Stau NLSP

Charged Massive “Stable” Particles

- “Old” analysis - 0.35 fb^{-1} , update is in the works...
- Exist in many models in addition to GMSB: AMSB, stable stop, R-hadrons...
- Appear in the detector as a “slow muons”
- Study exclusive pair production of CMSP’s
 - require two muons in event
 - measure muon speed with scintillator counters (counter resolution ~ 2 to 4 ns)
 - speed resolution depends on detector region and number of reconstructed hits -> construct “speed significance”
- Main background: Drell-Yan
- Use low-mass di-muons ($< 120 \text{ GeV}$) to calibrate speed significance
- Look for excess in high mass di-muons

Data agrees with SM expectation...



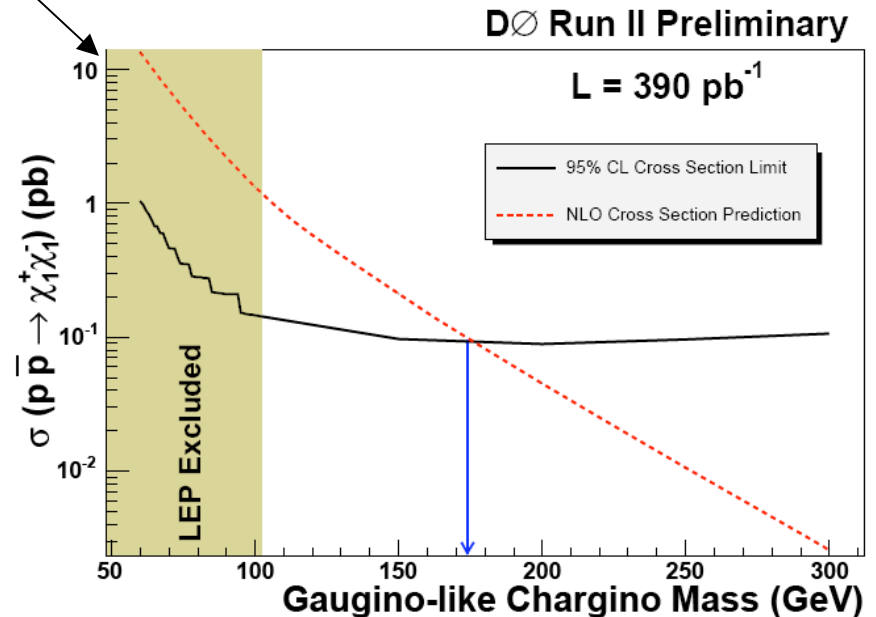
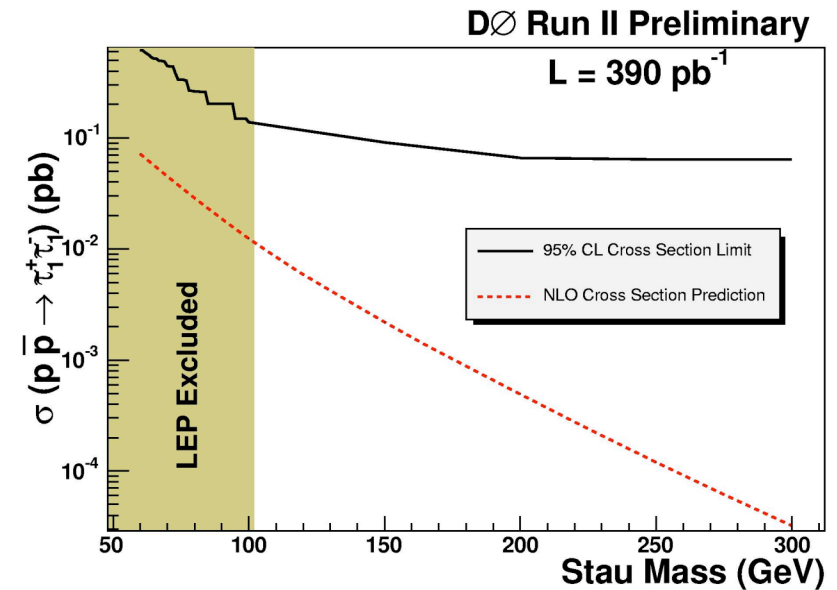
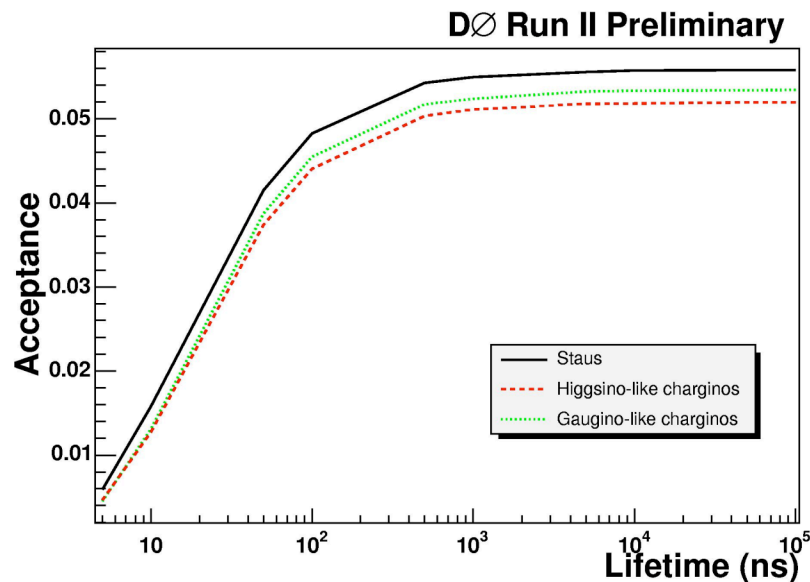
Charged Massive "Stable" Particles

● No excess, set limits

GMSB line (Snowmass slope D)
 $M=2\Lambda$, $N_5=3$, $\tan\beta=15$, $\text{sign } \mu > 0$

AMSB Gauginos

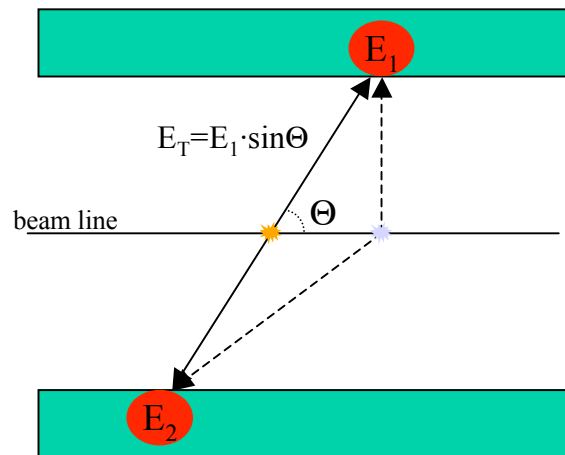
$M_1=3M_2$, $M_3=500$, $\mu=10$ TeV
 $\tan\beta = 15$, $M(\text{squark}) = 800$ GeV



Di-photon Data Analysis

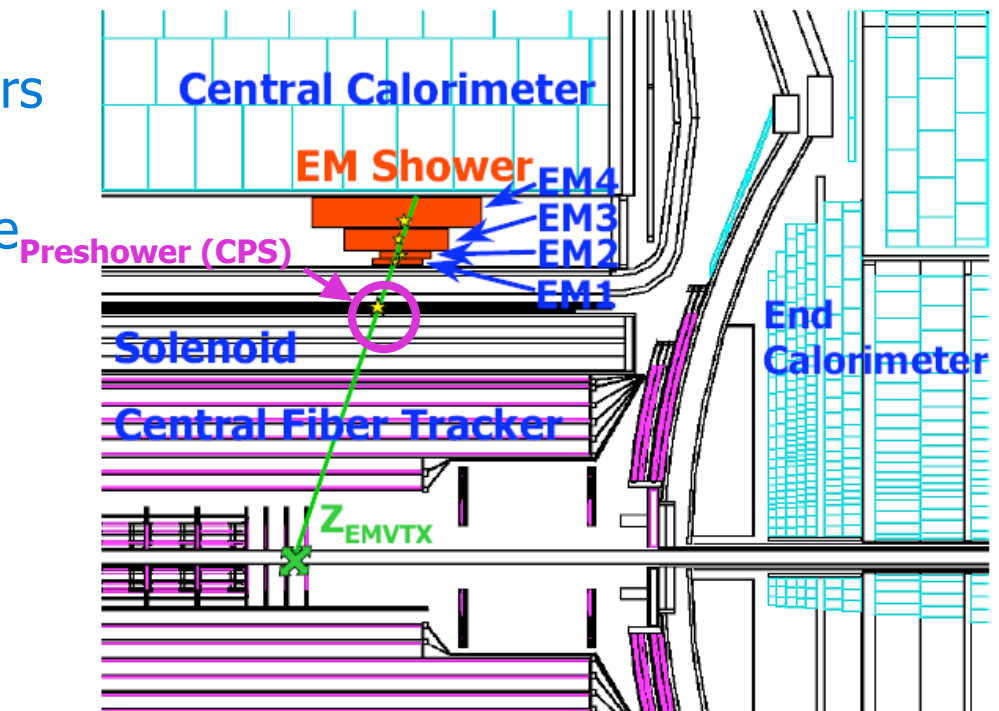
- Data collected in the Run IIa: 1.1 fb^{-1}
- Photon and electron identification:
 - Calorimeter cluster with $> 95\%$ energy in EM calorimeter
 - Isolated in calorimeter $(E_{\text{TOT}}^{R=0.4} - E_{\text{EM}}^{R=0.2})/E_{\text{EM}}^{R=0.2} < 0.07$
 - Scalar sum of track p_T in $0.05 < R < 0.4$ annulus around the direction of the cluster is less than 2 GeV
 - Shower shape is consistent with photon
 - Cluster is electron if there is a central track match (or an electron-like hit pattern in the tracker) and is a photon otherwise
- Event selection
 - Two photons, $E_T > 25 \text{ GeV}$ and $|\eta| < 1.1$
 - $\Delta\phi(\text{jet}, \text{MET}) < 2.5$ – for leading jet (if present) - to remove events with mismeasured missing E_T
 - use photon pointing to eliminate mis-vertexing

Photon Pointing & Vertex Selection



- There can be several interactions per event
- Vertex distribution RMS ~ 28 cm
- MET is significantly affected if the vertex is shifted by >10 cm
- Which vertex did the photons come from?

- DØ has four longitudinal EM layers and a preshower
- Fit a straight line through the five points (obtain resolution of each layer using $Z \rightarrow e e$ events)
- Z_{VTX} resolution ~ 2 cm (verified with $Z \rightarrow l l \gamma$ events)
- In the analysis require at least one photon to have CPS cluster



Backgrounds

- Physics backgrounds are small: $W\gamma\gamma$, $Z\gamma\gamma$ - COMPHEP MC
- All instrumental backgrounds can be determined from data

without true MET -fake MET

- QCD: $\gamma\gamma$, $\gamma+j$, $j+j$ (jet is faking γ)
- Drell-Yan (lost tracks)

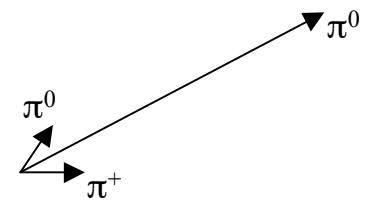
largest

with true MET -fake γ

- $W\gamma \rightarrow e\nu\gamma$ (lost track)
 $Wj \rightarrow e\nu j$ (lost track, fake γ)
- $Z \rightarrow \tau\tau \rightarrow ee + X$ (lost tracks)
- $t\bar{t} \rightarrow ee + X$ (lost tracks)
- WW, WZ, \dots (lost tracks)

Backgrounds with No Genuine MET

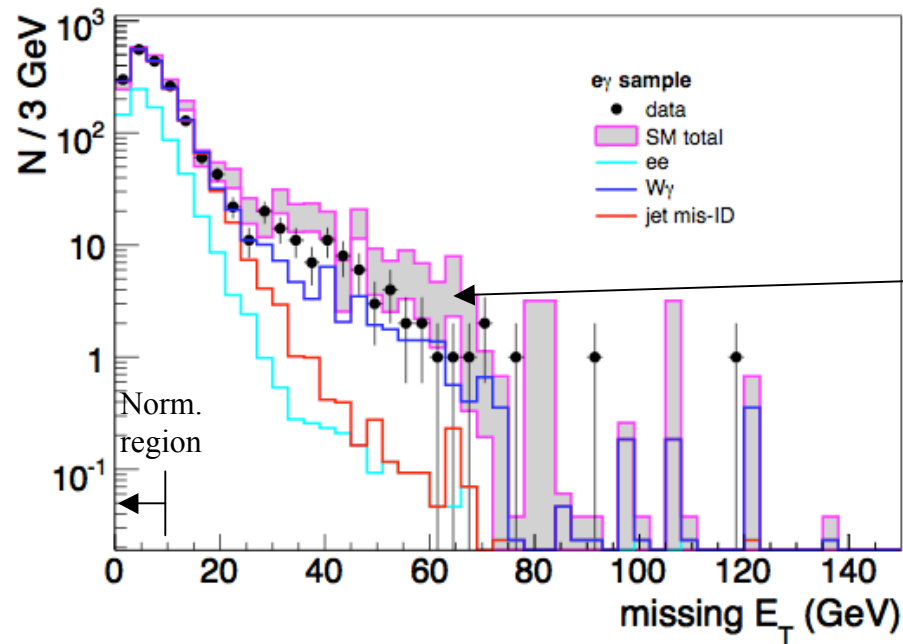
- Need to know the shape of MET distribution for them
 - and normalize to data with low MET
- MET resolution is dominated by the energy resolution of the photons
 - to first order, the shape does not depend on whether the photons are real or faked by jets
 - still, there is a small difference due to the fact that when jet fakes a photon, the photon's energy is less than original parton's energy
- Take $\gamma\gamma$ shape from $Z \rightarrow ee$
- Take fake shape from a sample that is the same as signal sample except both photons fail shower shape cut (*hh sample*)
- Take relative contribution that fits data best
 - $60 \pm 20\%$ of real $\gamma\gamma$
 - agrees with MC expectation for $\gamma\gamma$
 - purity cross-checked by looking at shower shape in the preshower (not used for photon ID)



fluctuated jet: most energy is carried by π^0

Backgrounds with Genuine MET

- Always involve electron - photon mis-identification
 - can determine using $e\gamma$ events and known mis-identification rate
- Select $e\gamma$ events using the same kinematical cuts as $\gamma\gamma$
 - Contributions from
 - $Z \rightarrow ee$ (one lost track) - get contribution from di-EM mass fit
 - QCD (jets faking electron and photon) - subtract using shape of hh sample
 - the processes we want to measure ($W\gamma$, Wj , WZ , $t\bar{t}$, etc)

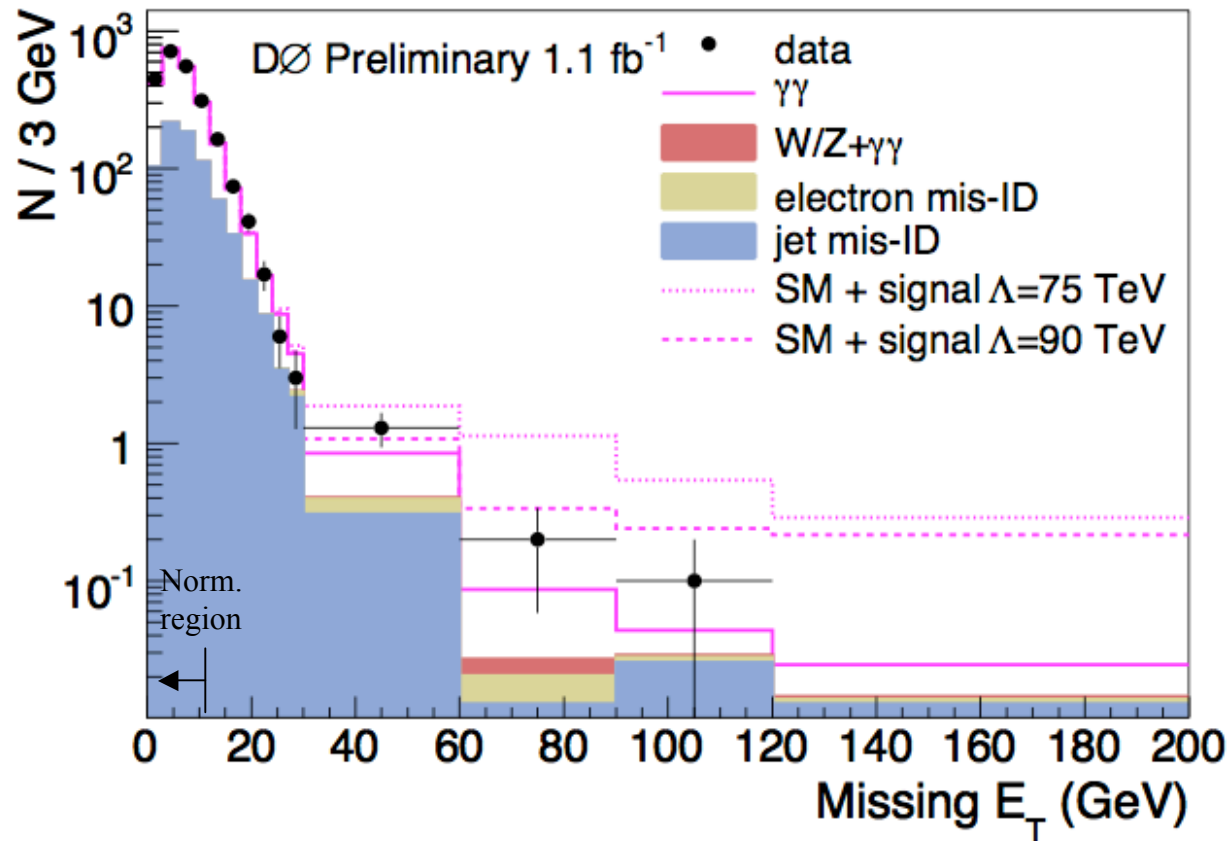


also serves as a cross-check of QCD subtraction method:

grey area is $Z+QCD+W\gamma$ MC + Wj obtained from data

error is dominated by $j \rightarrow \gamma$ fake rate (correlated between bins)

MET in $\gamma\gamma$ Sample



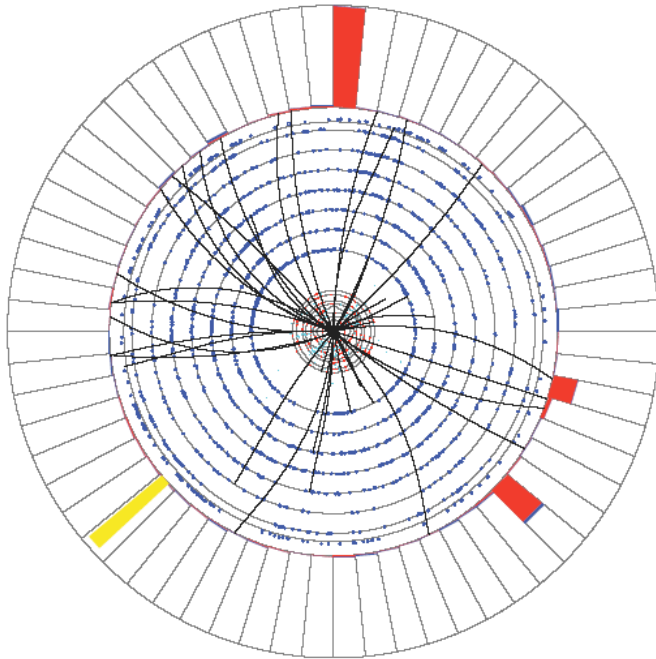
- Signal MC: ISAJET for masses and branchings, PYTHIA for event generation, full detector simulation plus real zero-bias events overlay to simulate multiple interactions per crossing

Two Highest MET Events

$ME_T = 63$ GeV, $E_T(\gamma) = 69, 27$ GeV
plus electron with $E_T = 23$ GeV

Run 187800 Evt 82968527 Sat Jan 3 16:42:02 2004

ET scale: 69 GeV

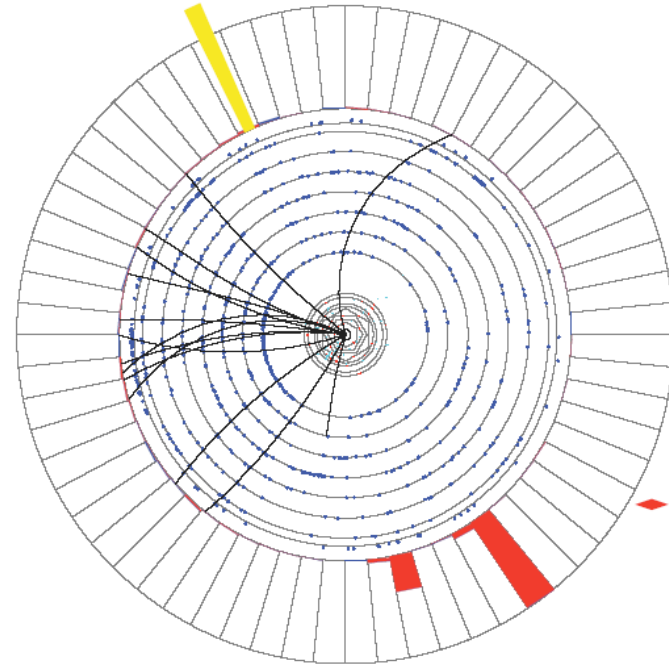


Expect ~ 0.15 events from $W\gamma\gamma$
with $E_T(\gamma) > 25$ GeV

$ME_T = 105$ GeV, $E_T(\gamma) = 82, 33$ GeV

Run 175918 Evt 28681786 Mon Apr 21 22:37:15 2003

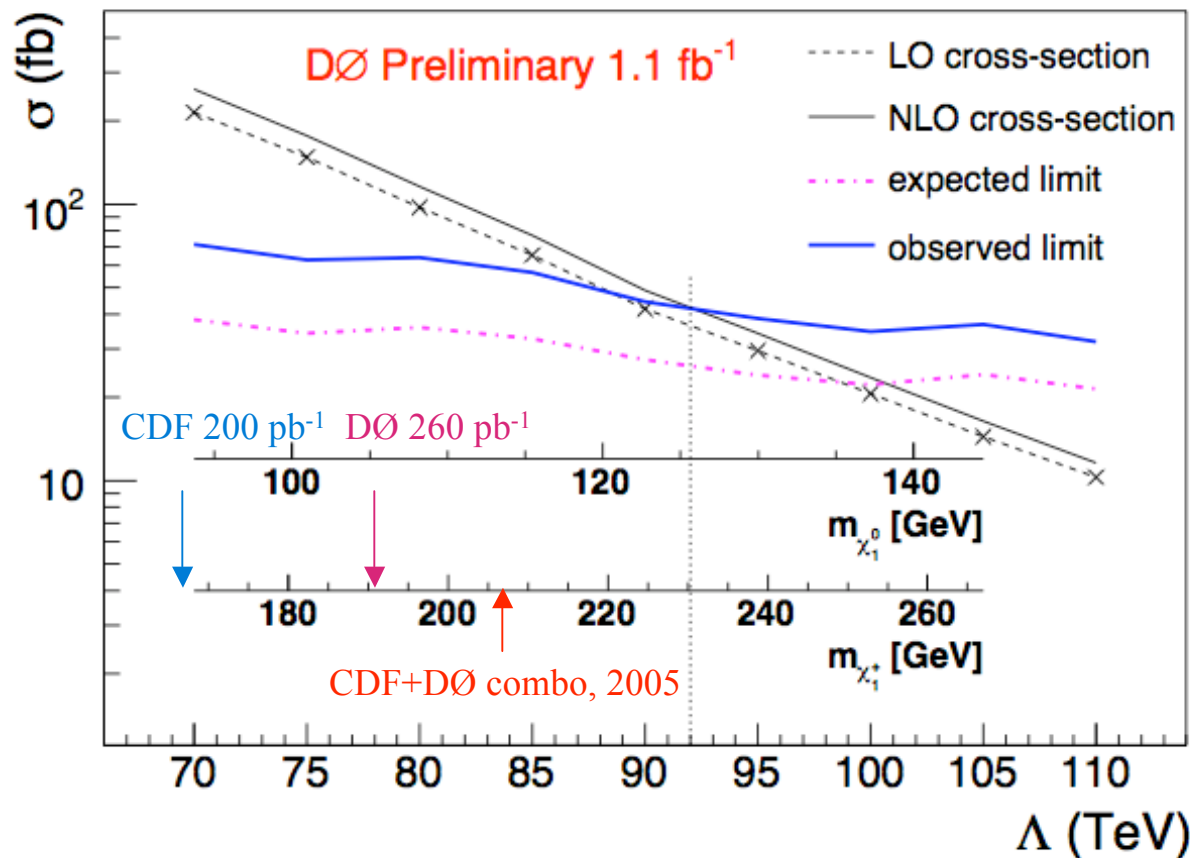
ET scale: 71 GeV



Expect ~ 0.1 event from $Z\gamma\gamma$
with $E_T(\gamma) > 25$ GeV

Limit Setting

- Since data agrees with the MC proceed to limit-setting
- Use CLS method - takes into account shape of the distributions



@ 95% CL:

$\Lambda > 92$ TeV

$m(\chi_1^0) > 126$ GeV

$m(\chi_1^+) > 230$ GeV

Summary and Outlook

- Still no sign of SUSY, although some interesting events are showing up...
- Not the end of the story - have more than twice the data on tape, will get $\sim 8 \text{ fb}^{-1}$ by the end of 2009

Stay tuned!