



The dark connection between Canis Major, Monoceros Stream, gas flaring, the rotation curve and the EGRET excess



From EGRET excess of diffuse Galactic gamma rays

- Determination of WIMP mass
- Determination of WIMP halo (= standard halo + DM ring)

Confirmation:

- Rotation curve
- Canis Major/Monoceros stream
- Gas flaring

PREDICTIONS

- for LHC (if SUSY)
- for direct searches
- for solar neutrinos

Ingredients to this analysis

Astronomers

Rotation curve
Tidal streams
Gas flaring



Astrophysics

Cosmics
Gamma rays



Cosmology

23%DM, thermal history of WIMPs
Annihilation cross section
Tidal disruption of dwarf



Particle Physics

Gamma ray spectra for BG + DMA

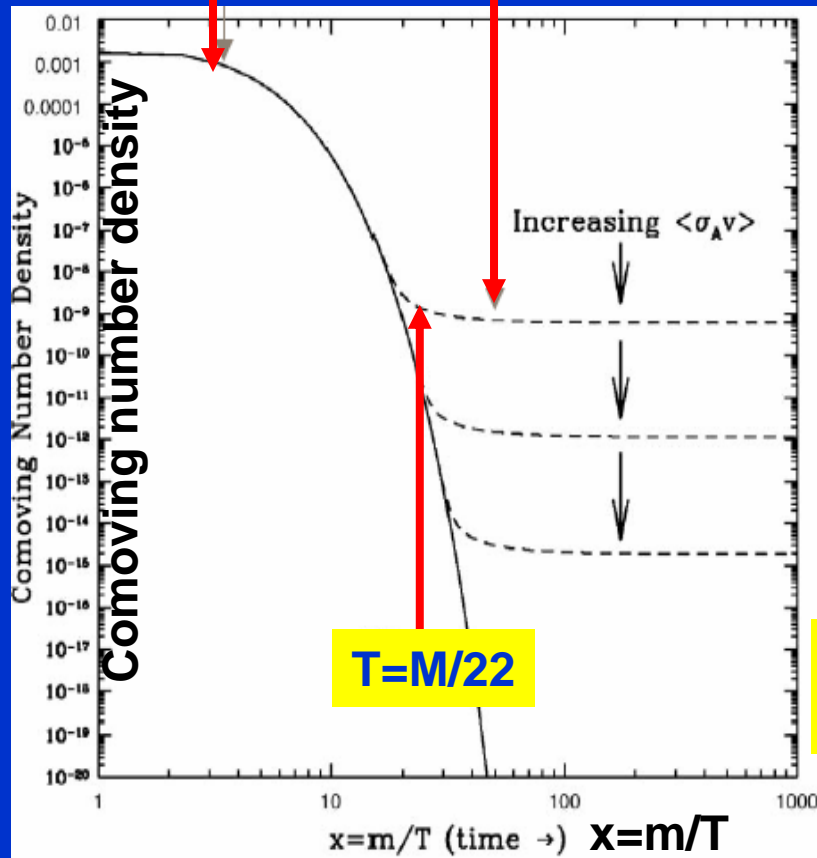


Expansion rate of universe determines WIMP annihilation cross section



Thermal equilibrium abundance

Actual abundance



$T \gg M$: $f + \bar{f} \rightarrow M + \bar{M}$; $M + \bar{M} \rightarrow f + \bar{f}$
 $T < M$: $M + \bar{M} \rightarrow f + \bar{f}$
 $T = M/22$: M decoupled, stable density
 (wenn Annihilationrate \cong Expansionsrate, i.e. $\Gamma = \langle \sigma v \rangle n_\chi(x_{fr}) \cong H(x_{fr})$!)

WMAP $\rightarrow \Omega h^2 = 0.113 \pm 0.009 \rightarrow$
 $\langle \sigma v \rangle = 2.10^{-26} \text{ cm}^3/\text{s}$

DM increases in Galaxies:
 ≈ 1 WIMP/coffee cup $\approx 10^5 \langle \rho \rangle$.
 DMA ($\propto \rho^2$) restarts again..

Annihilation into lighter particles, like
 quarks and leptons $\rightarrow \pi_0$'s \rightarrow Gammas!

Only assumption in this analysis:
 WIMP = THERMAL RELIC!

Gary Steigmann (1979)



What we get from cosmology



IF DM particles are thermal relics from early universe, they can annihilate with a cross section as large as

$$\langle \sigma v \rangle = 2 \cdot 10^{-26} \text{ cm}^3/\text{s}$$

which implies an enormous rate of gamma rays from π_0 decays (produced in quark fragmentation)
(Galaxy = 10^{40} higher rate than any accelerator)

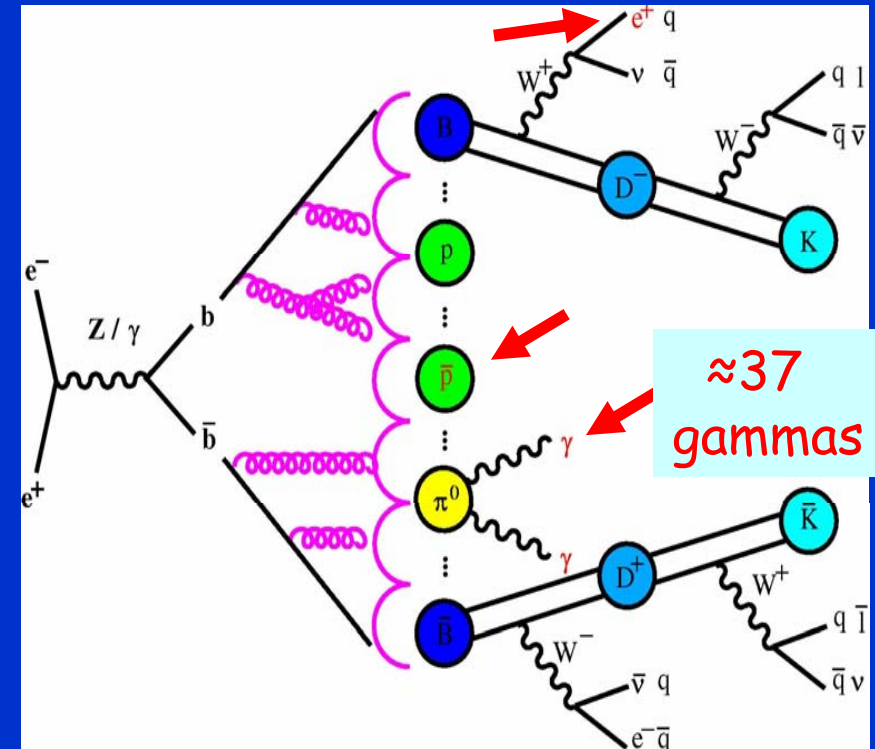
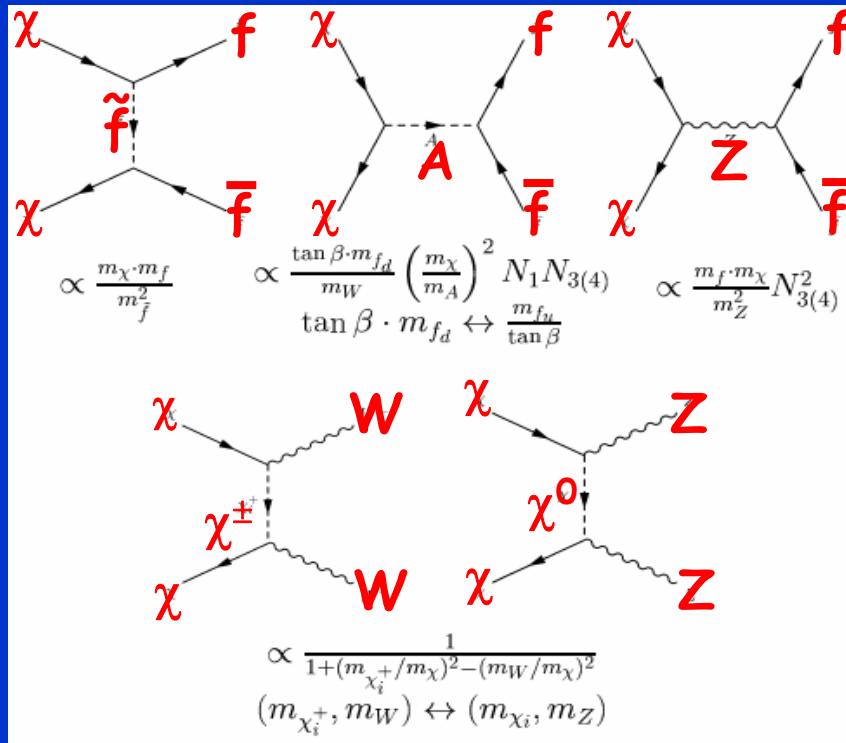
Expect large fraction of energetic Galactic gamma rays to come from DMA in this case.

Remaining ones from $p_{CR} + p_{GAS} \rightarrow \pi_0 + X$, $\pi_0 \rightarrow 2\gamma$
(+some IC+brems)

This means: Galactic gamma rays have 2 components with a shape KNOWN from the 2 BEST studied reactions in accelerators: background known from fixed target exp.
DMA known from e^+e^- annihilation (LEP)



Example of DM annihilation (SUSY)



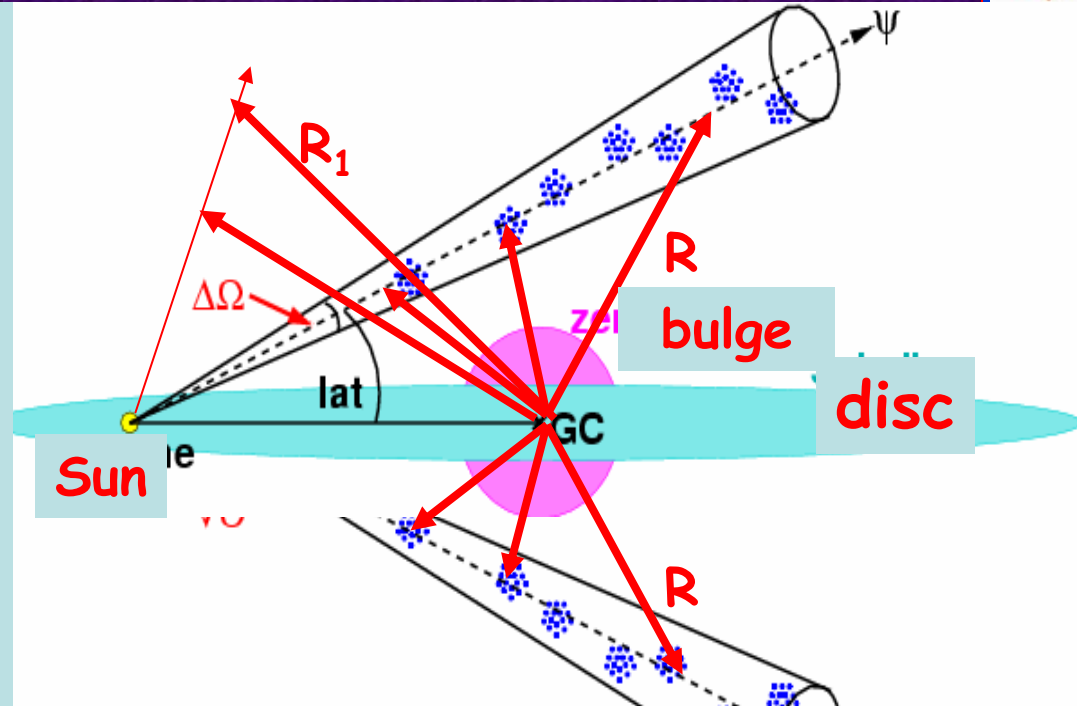
Dominant

$\chi + \chi \Rightarrow A \Rightarrow b \bar{b}$ quark pair
 Sum of diagrams should yield
 $\langle \sigma v \rangle = 2.10^{-26} \text{ cm}^3/\text{s}$ to get
 correct relic density

Quark fragmentation known!

Hence spectra of positrons,
 gammas and antiprotons known!
 Relative amount of γ, p, e^+ known
 as well.

Divergent for $r=0$?
NFW profile $\propto 1/r$
Isotherm profile const.

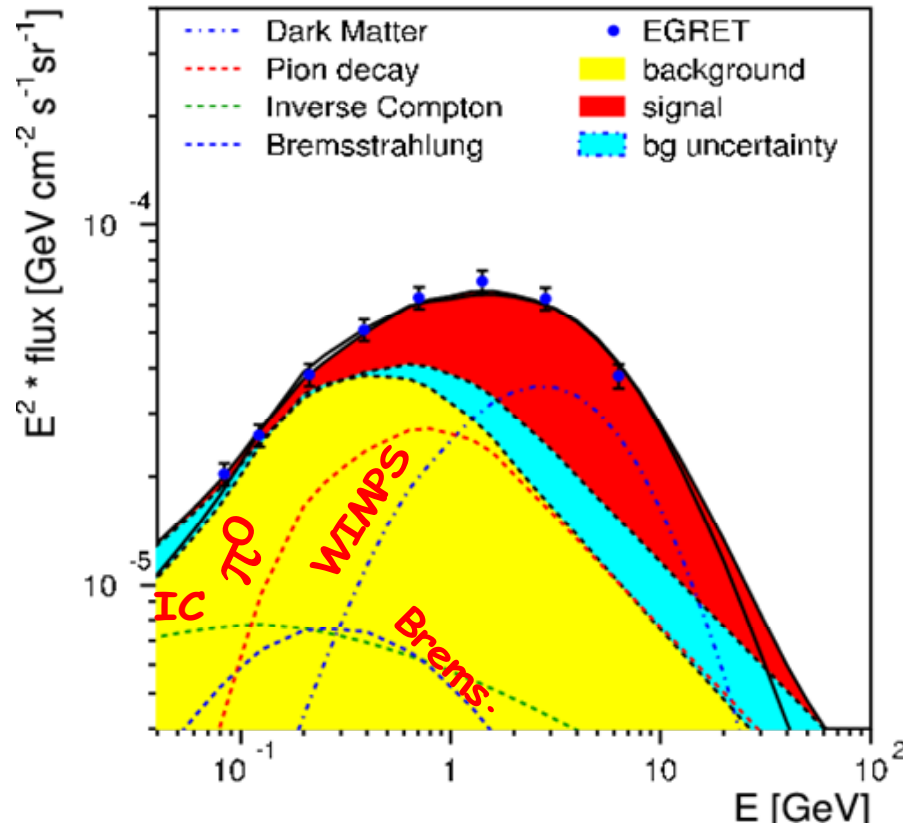


**IF FLUX AND SHAPE MEASURED IN
ONE DIRECTION, THEN FLUX AND
SHAPE FIXED IN ALL (=180) SKY
DIRECTIONS!!**

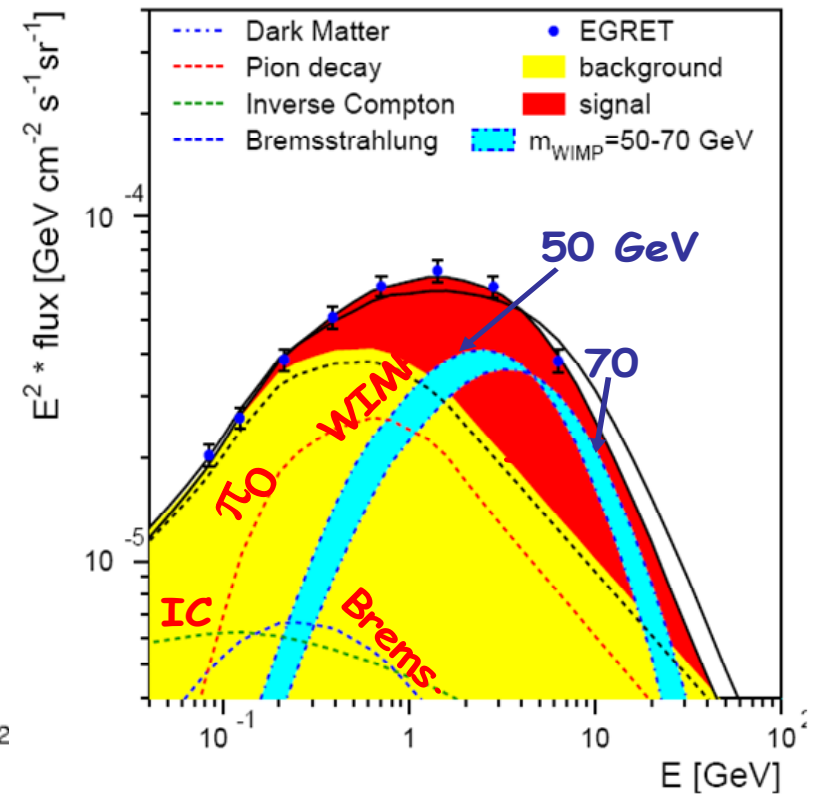
THIS IS AN INCREDIBLE CONSTRAINT, LIKE SAYING I VERIFY THE EXCESS AND WIMP MASS WITH 180 INDEPENDENT MEAS.



Background + signal describe EGRET data!



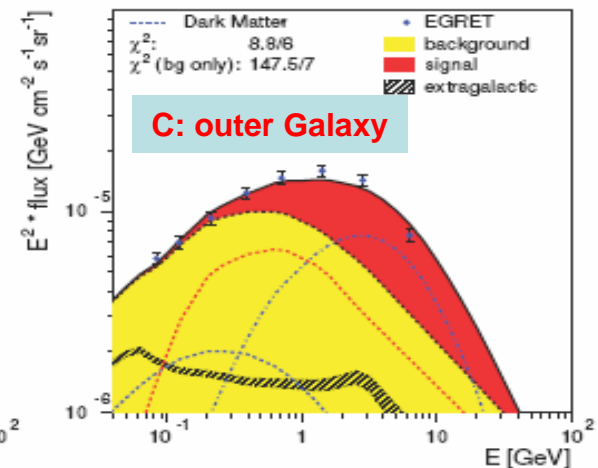
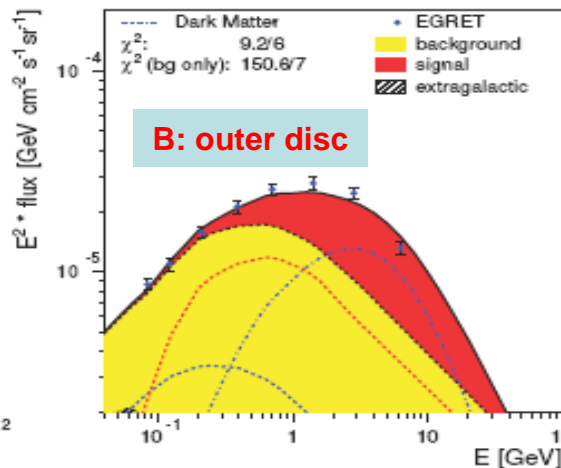
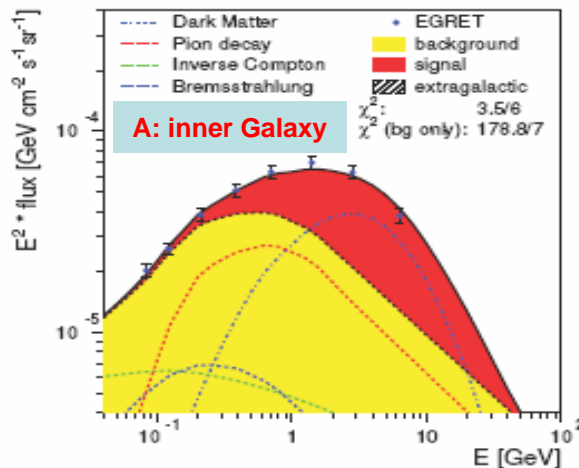
Blue: background uncertainty



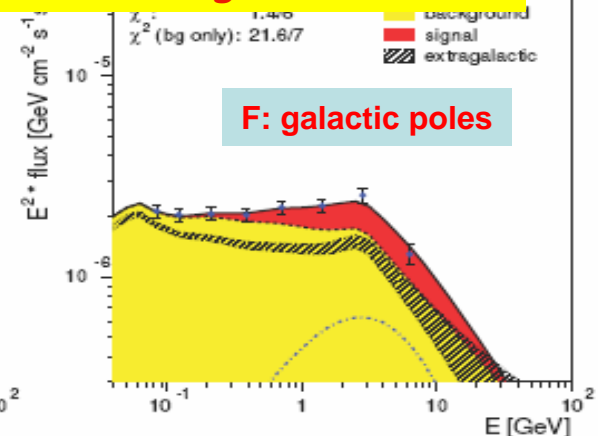
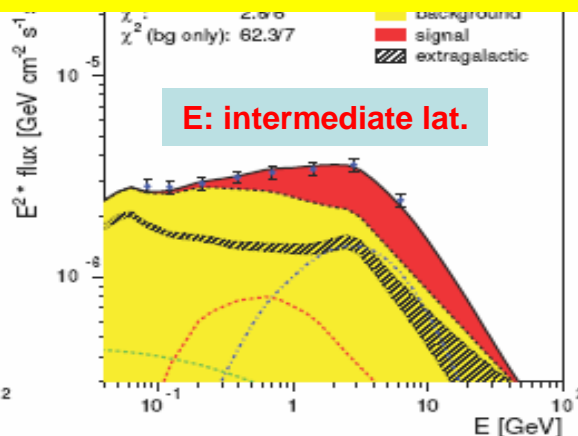
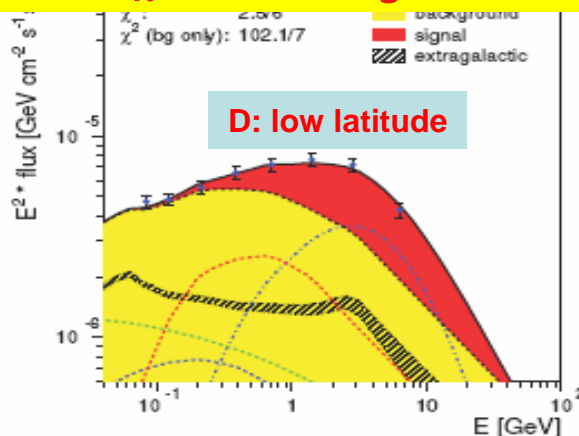
Blue: WIMP mass uncertainty



Analysis of EGRET Data in 6 sky directions

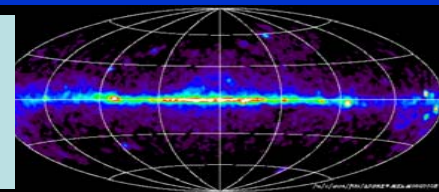


Total χ^2 for all regions :28/36 \Rightarrow Prob.= 0.8 Excess above background $> 10\sigma$.



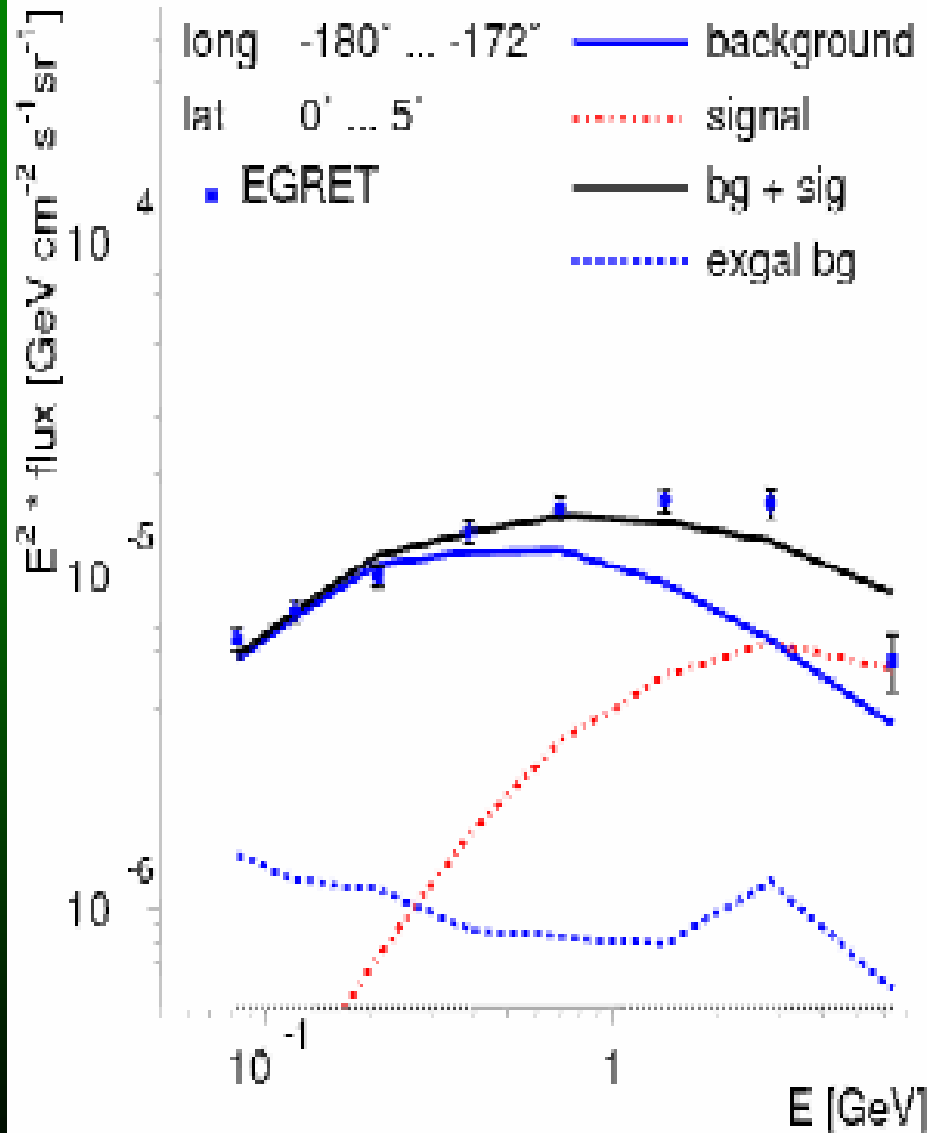
A: inner Galaxy ($l=\pm 30^\circ$, $|b|<5^\circ$)
B: Galactic plane avoiding A
C: Outer Galaxy

D: low latitude ($10-20^\circ$)
E: intermediate lat. ($20-60^\circ$)
F: Galactic poles ($60-90^\circ$)





Fits for 180 instead of 6 regions



180 regions:

8° in longitude \Rightarrow 45 bins

4 bins in latitude $\Rightarrow 0^\circ < |b| < 5^\circ$

$5^\circ < |b| < 10^\circ$

$10^\circ < |b| < 20^\circ$

$20^\circ < |b| < 90^\circ \Rightarrow$

$4 \times 45 = 180$ bins \Rightarrow

>1400 data points.

Reduced $\chi^2 \approx 1$ with 7% errors

BUT NEEDED IN ADDITION to $1/r^2$ profile, substructure in the form of 2 doughnut-like rings in the Galactic disc!

ONE RING COINCIDES WITH ORBIT FROM CANIS MAJOR DWARF GALAXY which loses mass along orbit by tidal forces

OTHER RING coincides with H_2 ring

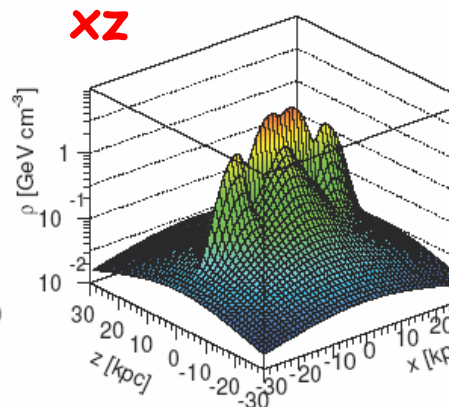
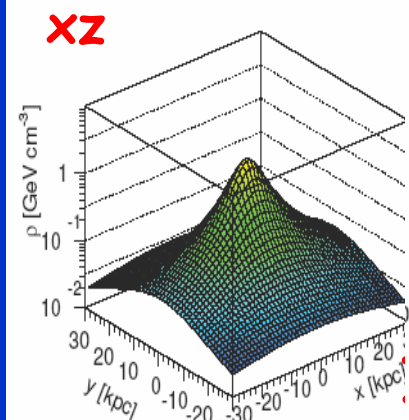
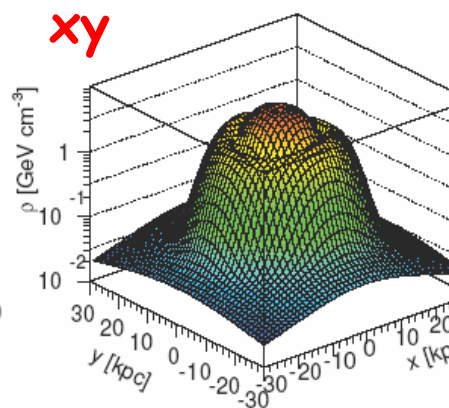
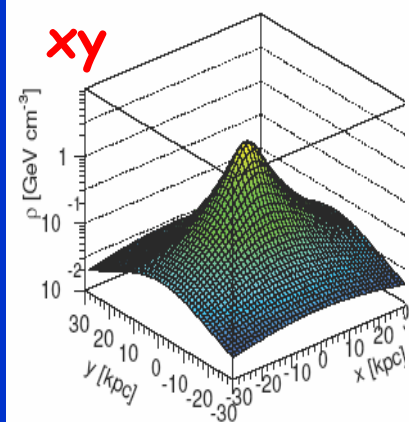


Dark Matter distribution

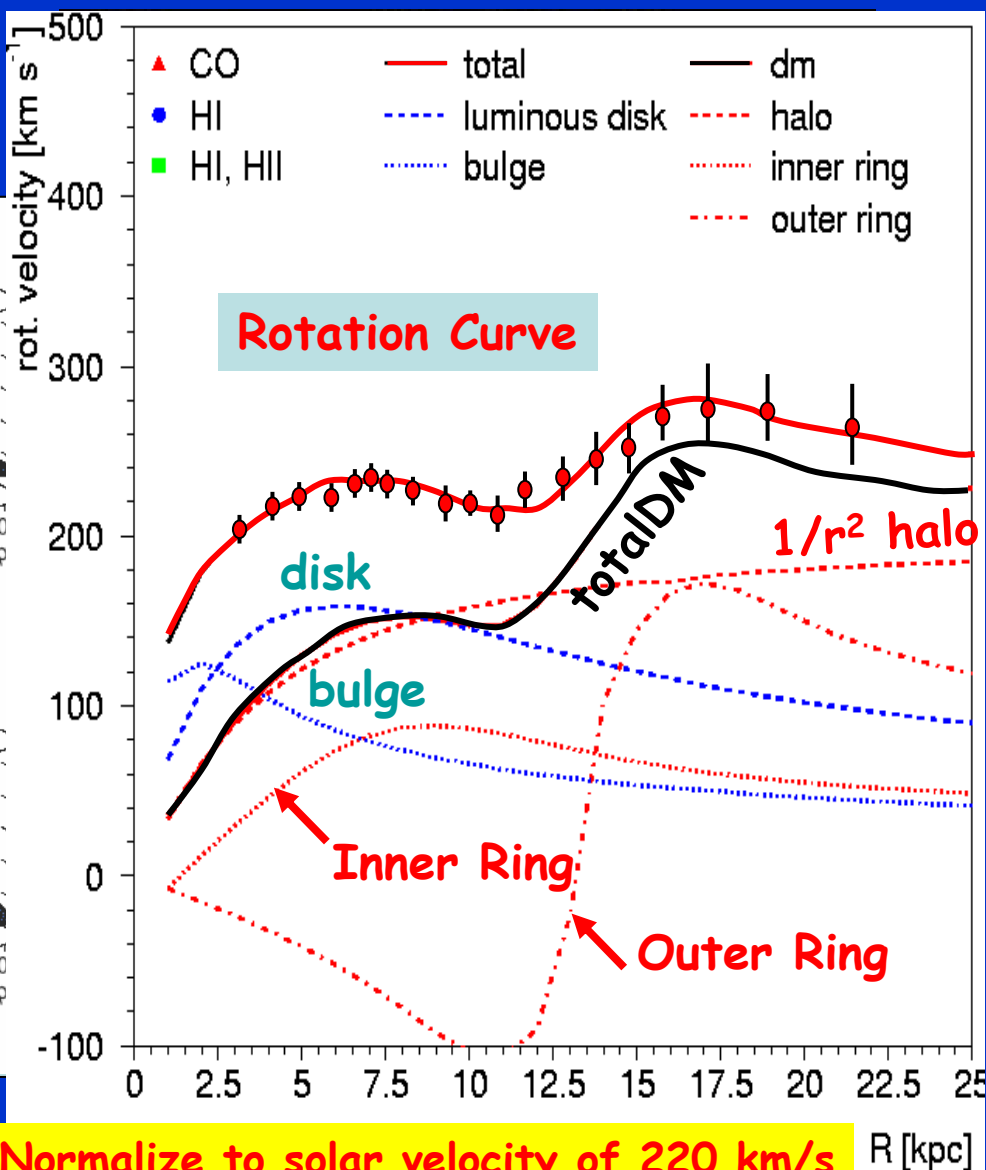


Expected
Profile

Observed
Profile



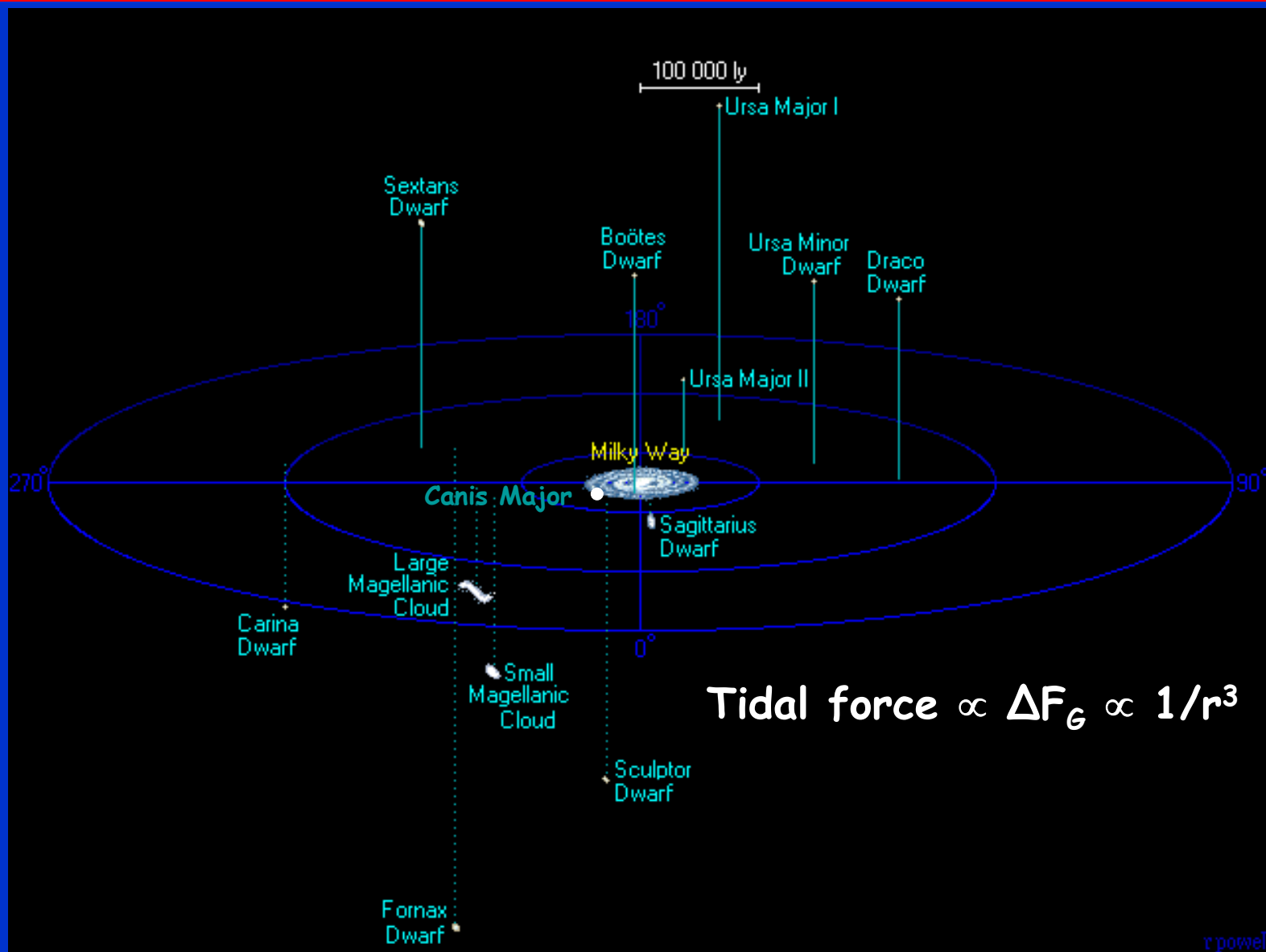
Halo profile



Normalize to solar velocity of 220 km/s

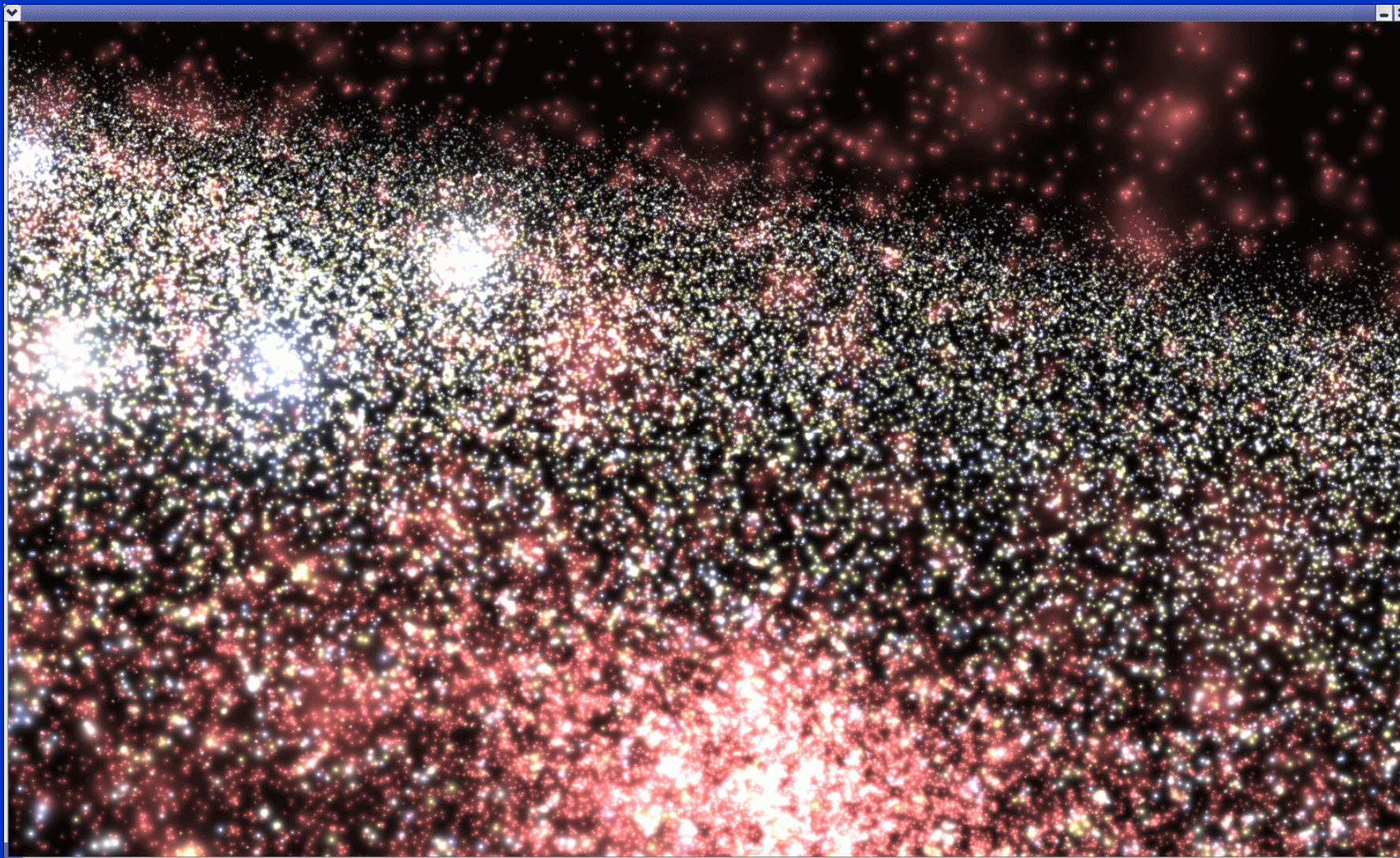


The Milky Way and its 13 satellite galaxies



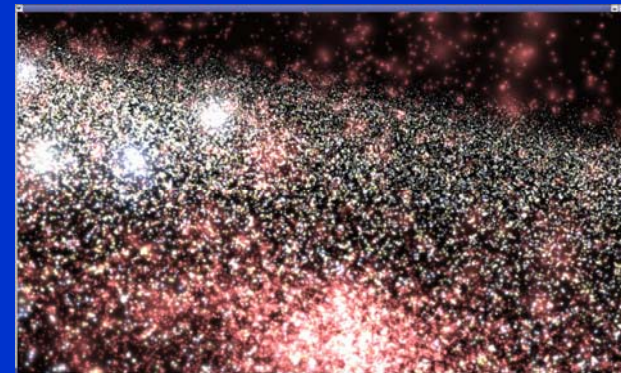


Artistic view of Canis Major Dwarf just below Galactic disc





Canis Major Dwarf orbits from N-body simulations to fit visible ring of stars at 13 and 18 kpc

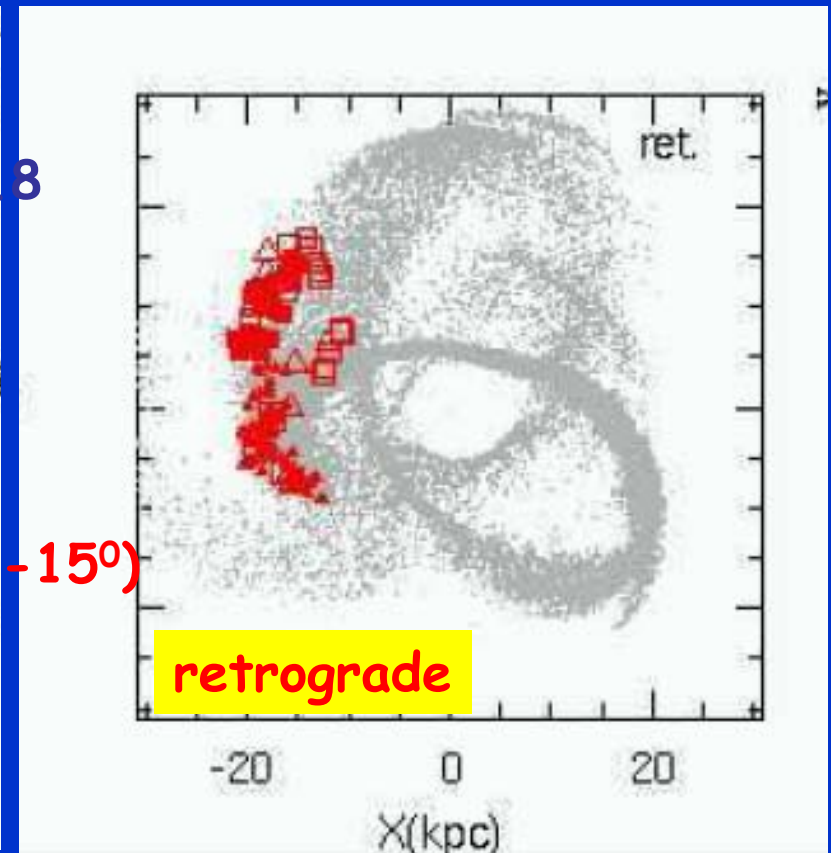
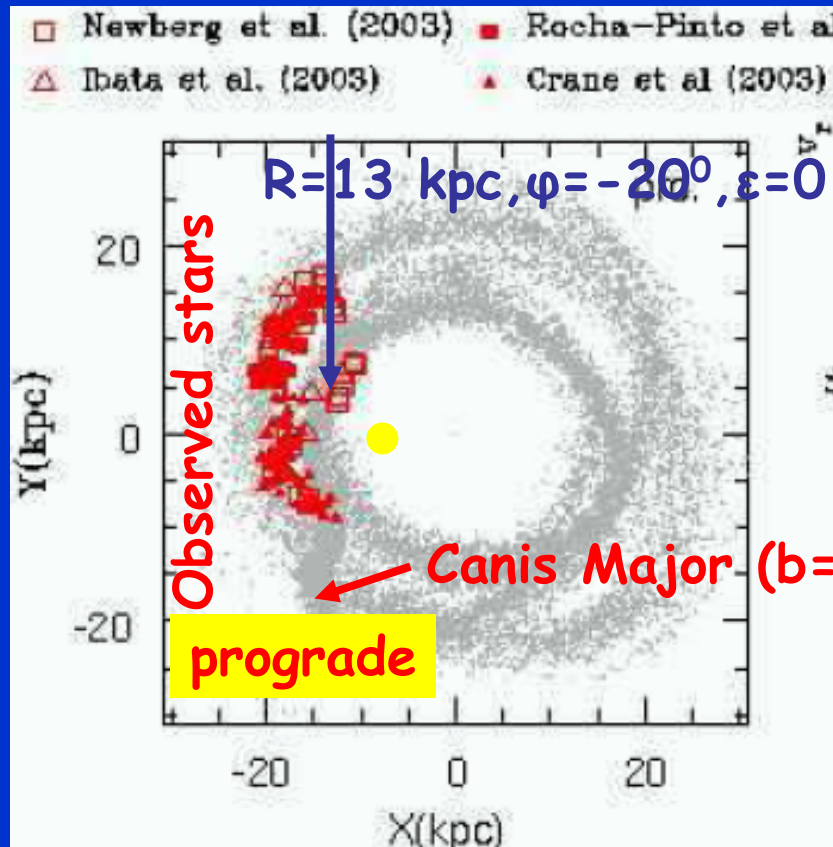


Movie from Nicolas Martin, Rodrigo Ibata
http://astro.u-strasbg.fr/images_ri/canm-e.html

Canis Major leaves at 13 kpc tidal stream of
gas($10^6 M_{\odot}$ from 21 cm line), stars ($10^8 M_{\odot}$, visible),
dark matter ($10^{10} M_{\odot}$, EGRET)



N-body simulation from Canis-Major dwarf galaxy

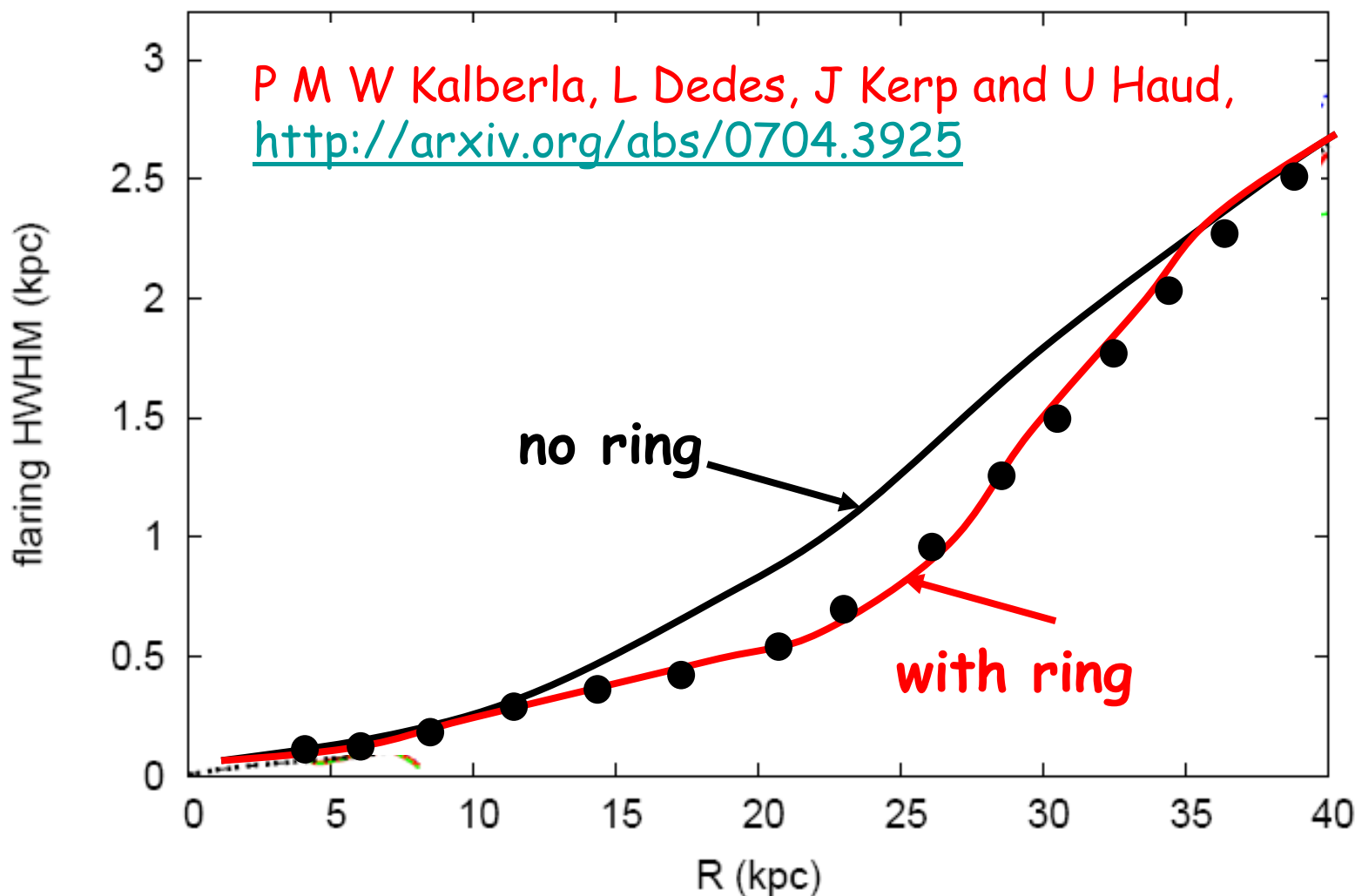


A comprehensive model for the Monoceros tidal stream

J. Peñarrubia¹, D. Martínez-Delgado¹, H.W. Rix¹, M.A Gómez-Flechoso², J. Munn³, H. Newberg⁴, E.F. Bell¹, B. Yanny⁵, D. Zucker¹, E. K. Grebel⁶



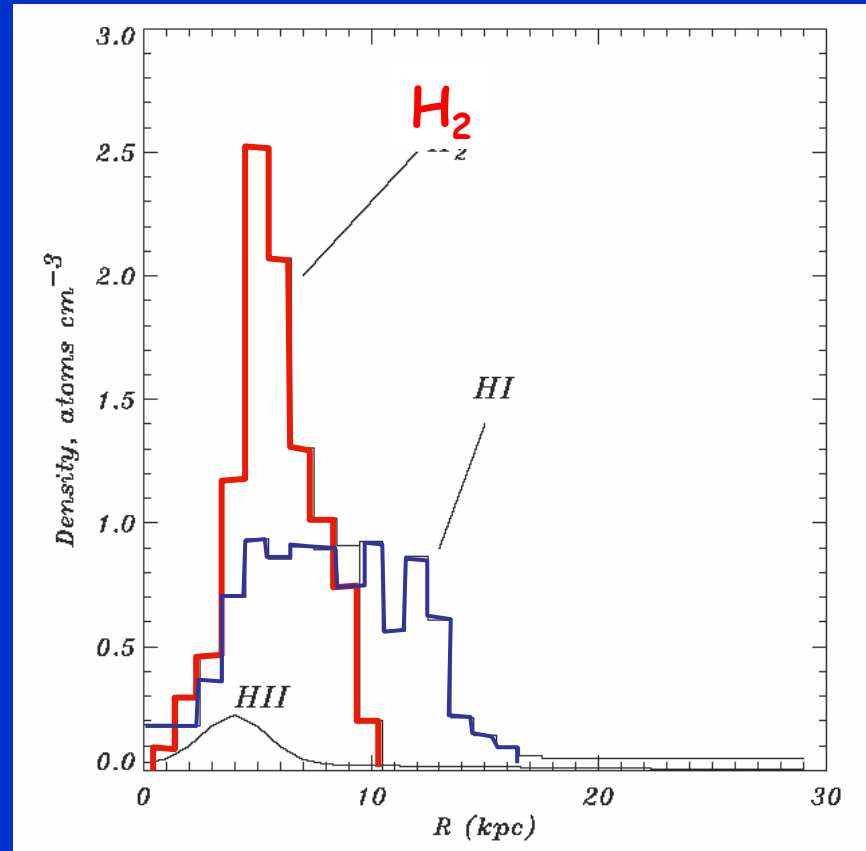
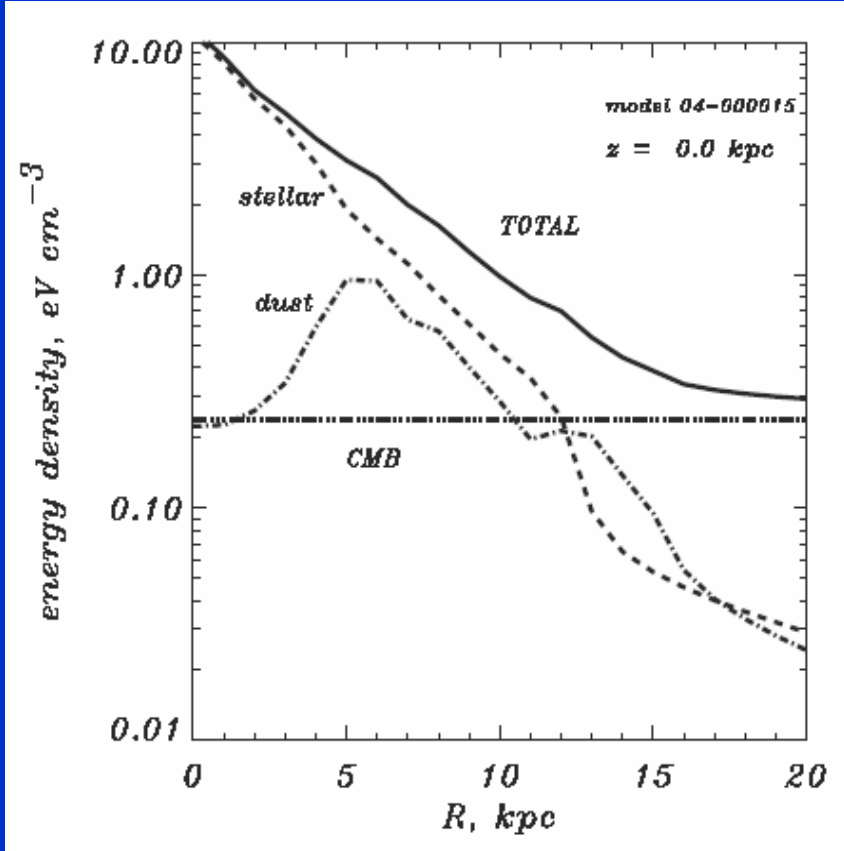
Gas flaring in the Milky Way



Gas flaring needs EGRET ring with mass of $2 \cdot 10^{10} M_{\odot}$!



Inner Ring coincides with ring of dust and H_2 -> gravitational potential well!



Enhancement of inner (outer) ring
over $1/r^2$ profile 6 (8).
Mass in rings 0.3 (3)% of total DM

4 kpc coincides with ring of
neutral hydrogen molecules!
 $H+H \rightarrow H_2$ in presence of dust ->
grav. potential well at 4-5 kpc.



8 physics questions answered SIMULTANEOUSLY if WIMP = thermal relic



- **Astrophysicists:**

What is the origin of "GeV excess" of diffuse Galactic Gamma Rays? A: DM annihilation

- **Astronomers:**

Why a change of slope in the galactic rotation curve at $R_0 \approx 11$ kpc? A: DM substructure

Why ring of stars at 13 kpc?

Why ring of molecular hydrogen at 4 kpc?

Why S-shape in gas flaring?

- **Cosmologists:** How is DM annihilating? A: into quark pairs

How is Cold Dark Matter distributed? A: standard profile + substructure

- **Particle physicists:**

Is DM annihilating as expected in Supersymmetry?

A: Cross sections perfectly consistent with mSUGRA for light gauginos, heavy squarks/sleptons



What about Supersymmetry?

Perfectly consistent
with mSUGRA, see
Talk by Christian Sander



What if DM is clumpy?

See talk by Martin Niegel?



Do antiproton data exclude interpretation of EGRET data?



Bergstrom et al. astro-ph/0603632, Abstract:

we investigate the viability of the model using the DarkSUSY package to compute the gamma-ray and antiproton fluxes. We are able to show that their (=WdB et al) model is excluded by a wide margin from the measured flux of antiprotons.

Answer: option that DarkSusy is wrong
much more likely
See talk by I. Gebauer



Summary



>>10 σ EGRET excess shows intriguing hint that:

WIMP is thermal relic with expected annihilation into quark pairs

DM becomes visible by gamma rays from fragmentation
(30-40 gamma rays of few GeV pro annihilation from π_0 decays)

Results rather model independent, since only KNOWN spectral shapes of signal and background used, NO model dependent calculations of abs.fluxes. Different shapes or unknown experimental problems may change the gamma ray flux and/or WIMP mass, BUT NOT the distribution in the sky.

SPATIAL DISTRIBUTION of annihilation signal is signature for DMA which clearly shows that EGRET excess is tracer of DM by fact that one can construct rotation curve and tidal streams from gamma rays.

DM interpretation strongly supported independently by gas flaring