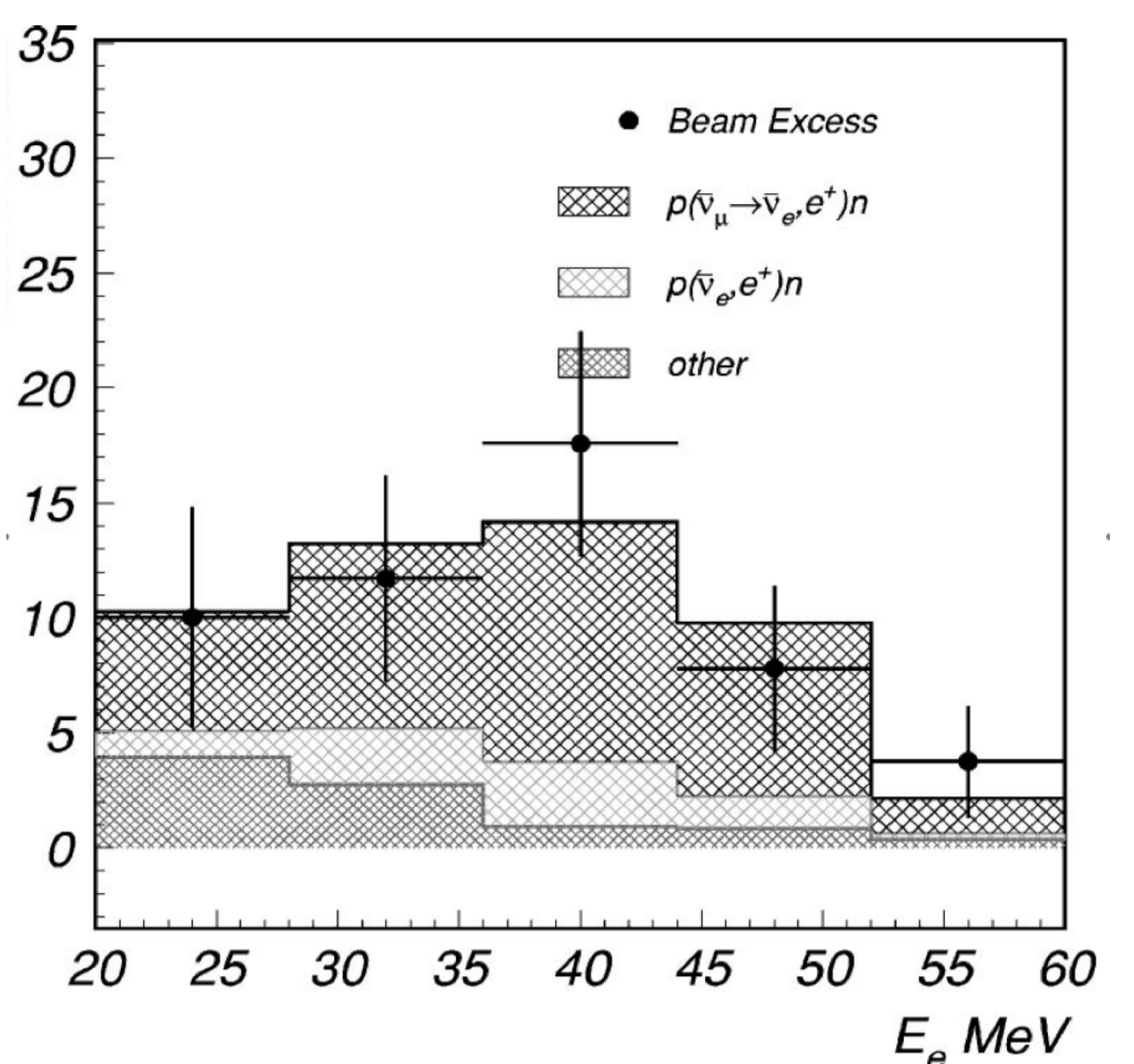


# Is there any “LSND anomaly”?

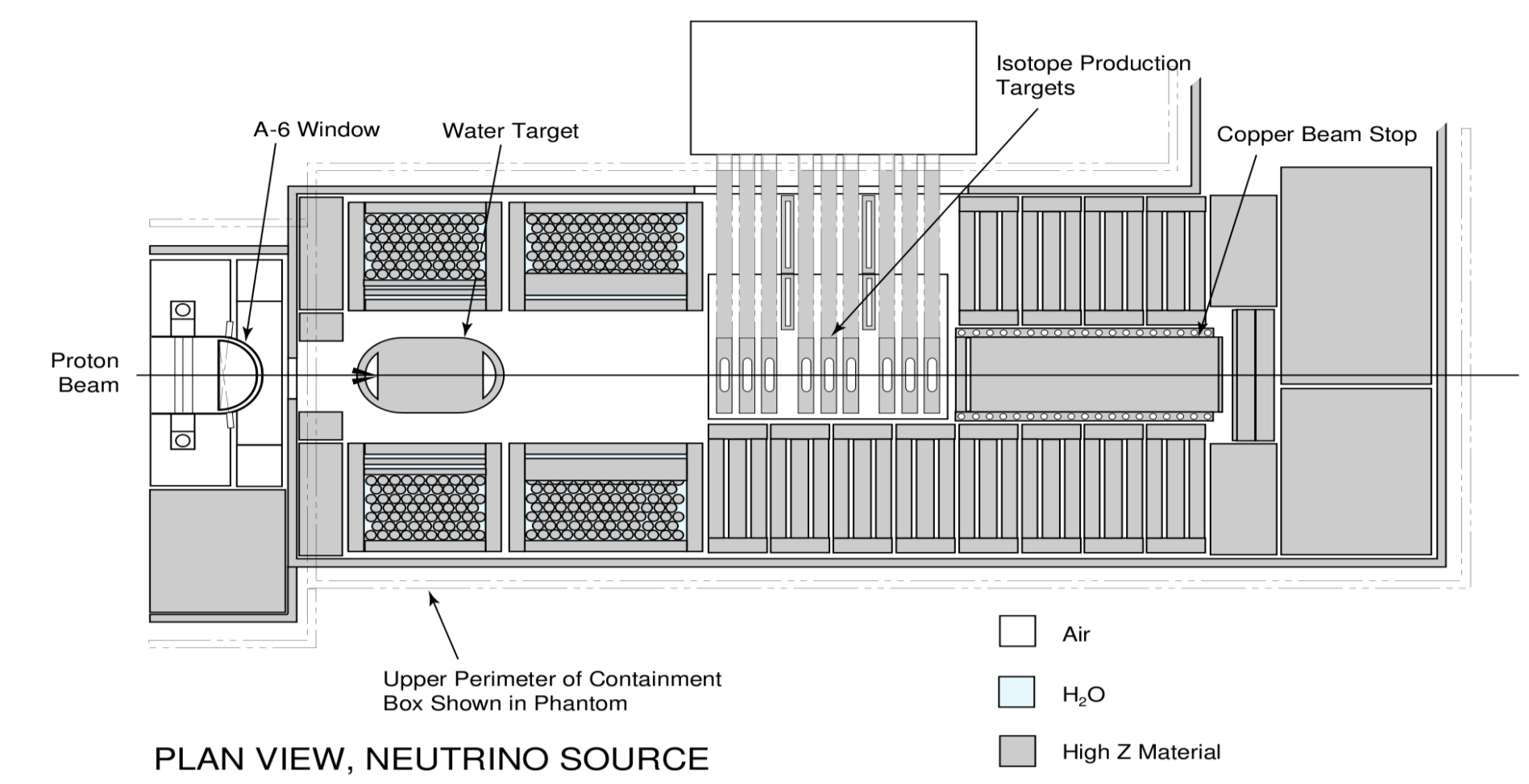
Alexey Zhemchugov for the HARP-CDP group  
 Joint Institute for Nuclear Research, Dubna, Russian Federation

## The LSND anomaly



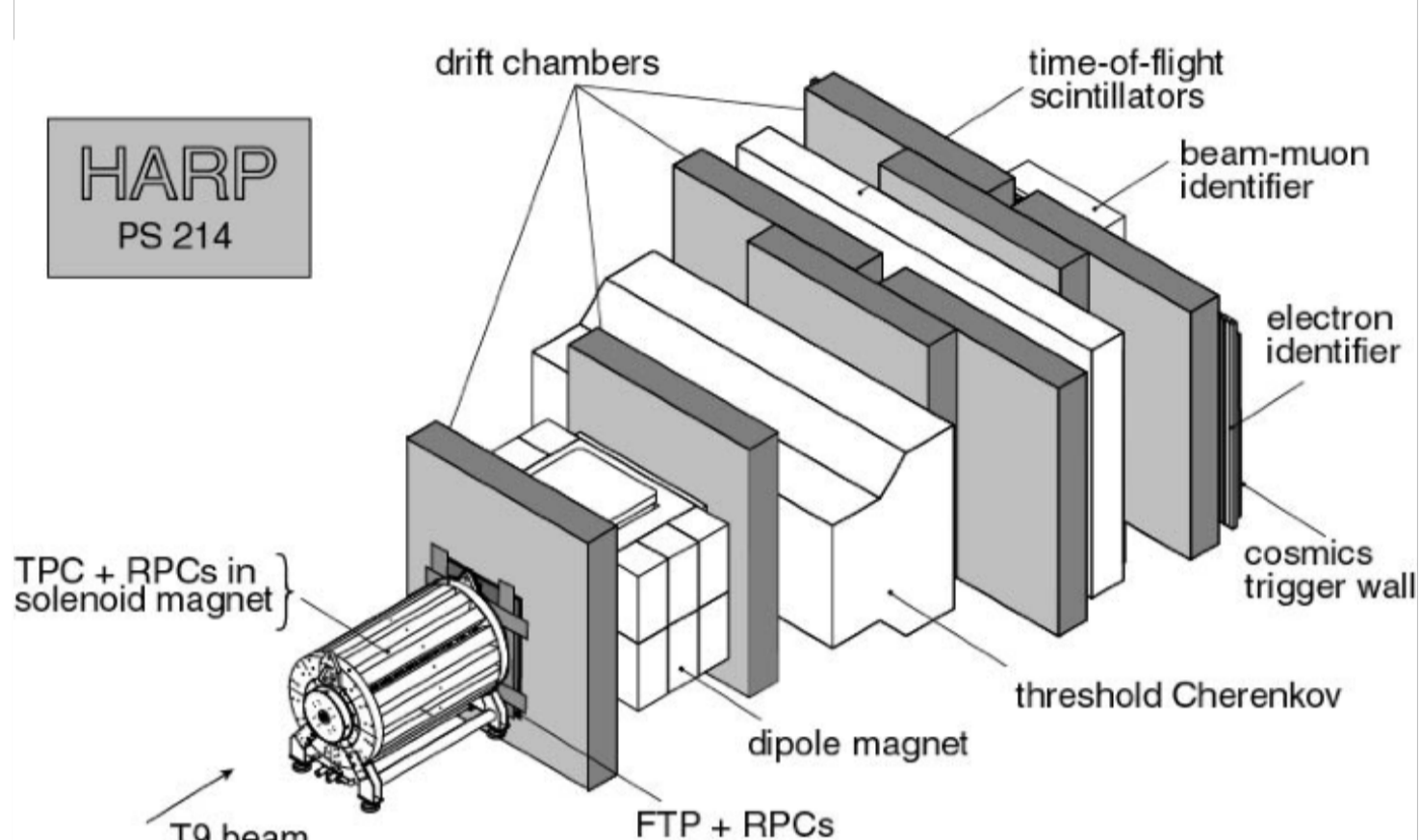
The LSND experiment reported an anomalous  $3.8 \sigma$  excess of  $\bar{\nu}_e$ , interpreted as  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation with  $\Delta m^2 \approx 1 \text{ eV}^2$ . This result has, until today, not been confirmed by other experiments. One explanation of the LSND signal is an underestimate of  $\pi^-$  production by 800 MeV protons in the calculation of the background neutrino flux in the LSND analysis. The HARP experiment measured pion production with a 1.5 GeV/c proton beam impinging on various target materials, including water and copper. The HARP measurements, together with pion production data from other experiments, were used to cross-check the calculation of the background of LSND's signal.

## The neutrino source at LANSCE



C. Athanassopoulos et al., NIM A 388 (1997) 149

## The HARP detector and its performance

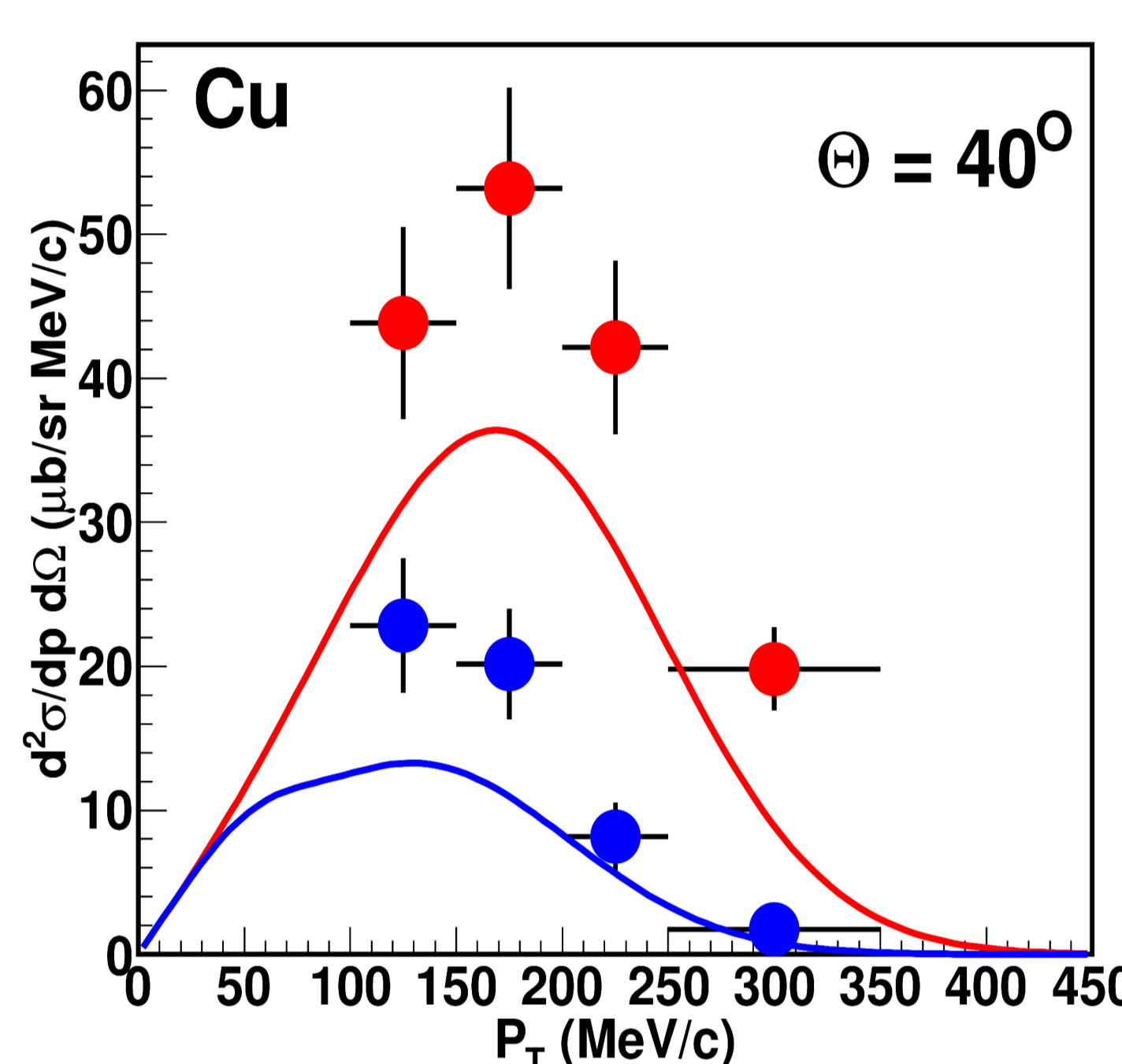
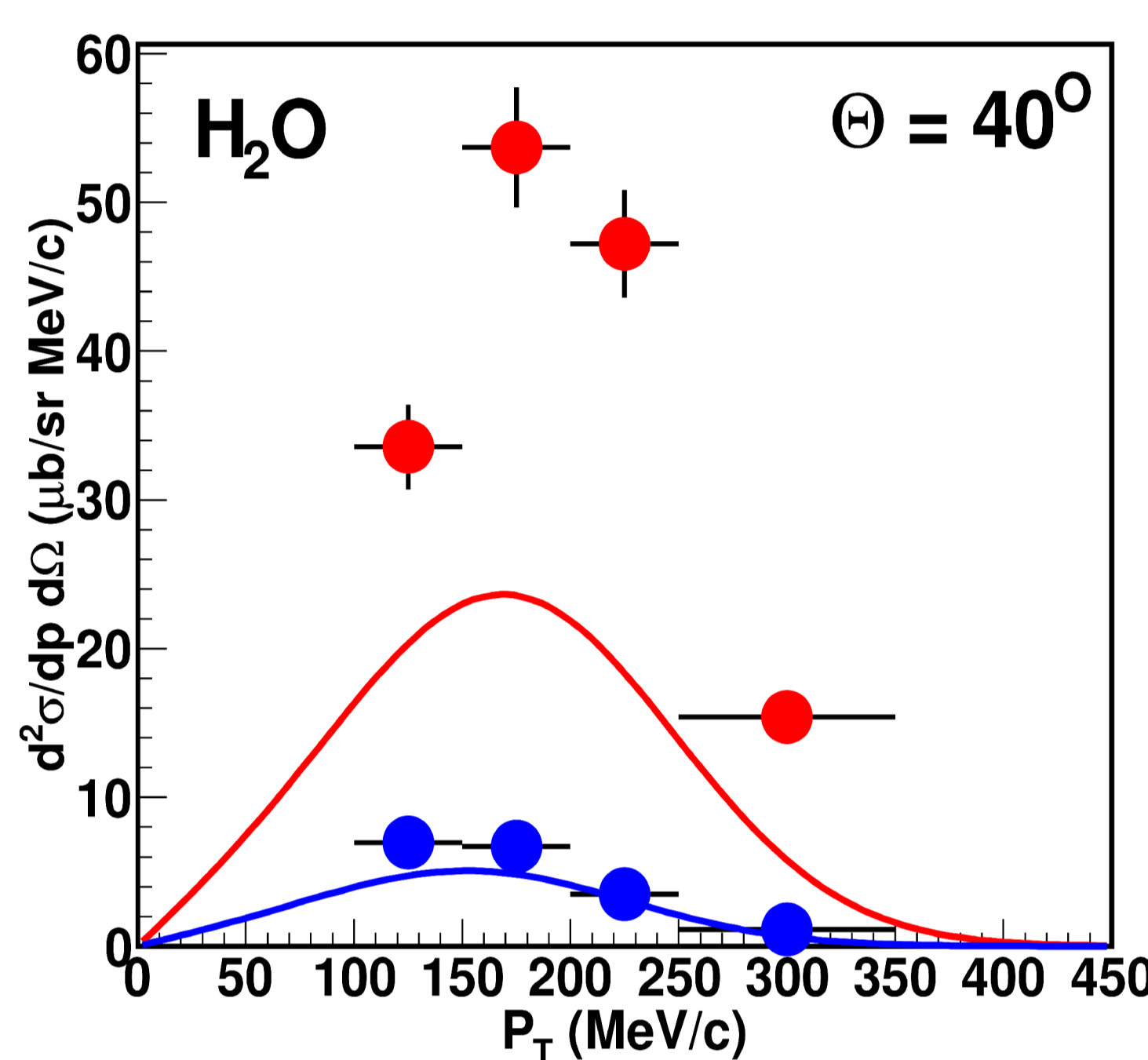


**TPC:**  
 NIM A588 (2008) 294-317  
 $\sigma(1/p_T) \approx 0.20 \text{ (GeV/c)}^{-1}$   
 $\sigma(dE/dx)/(dE/dx) \approx 0.16$   
 $\sigma(\Theta) \approx 9 \text{ mrad}$   
 Good particle identification by combining  $dE/dx$  from TPC and TOF from RPCs

**RPCs:**  
 NIM A578 (2007) 119-138  
 Efficiency  $\approx 98\%$   
 $\sigma(\text{TOF}) \approx 175 \text{ ps}$

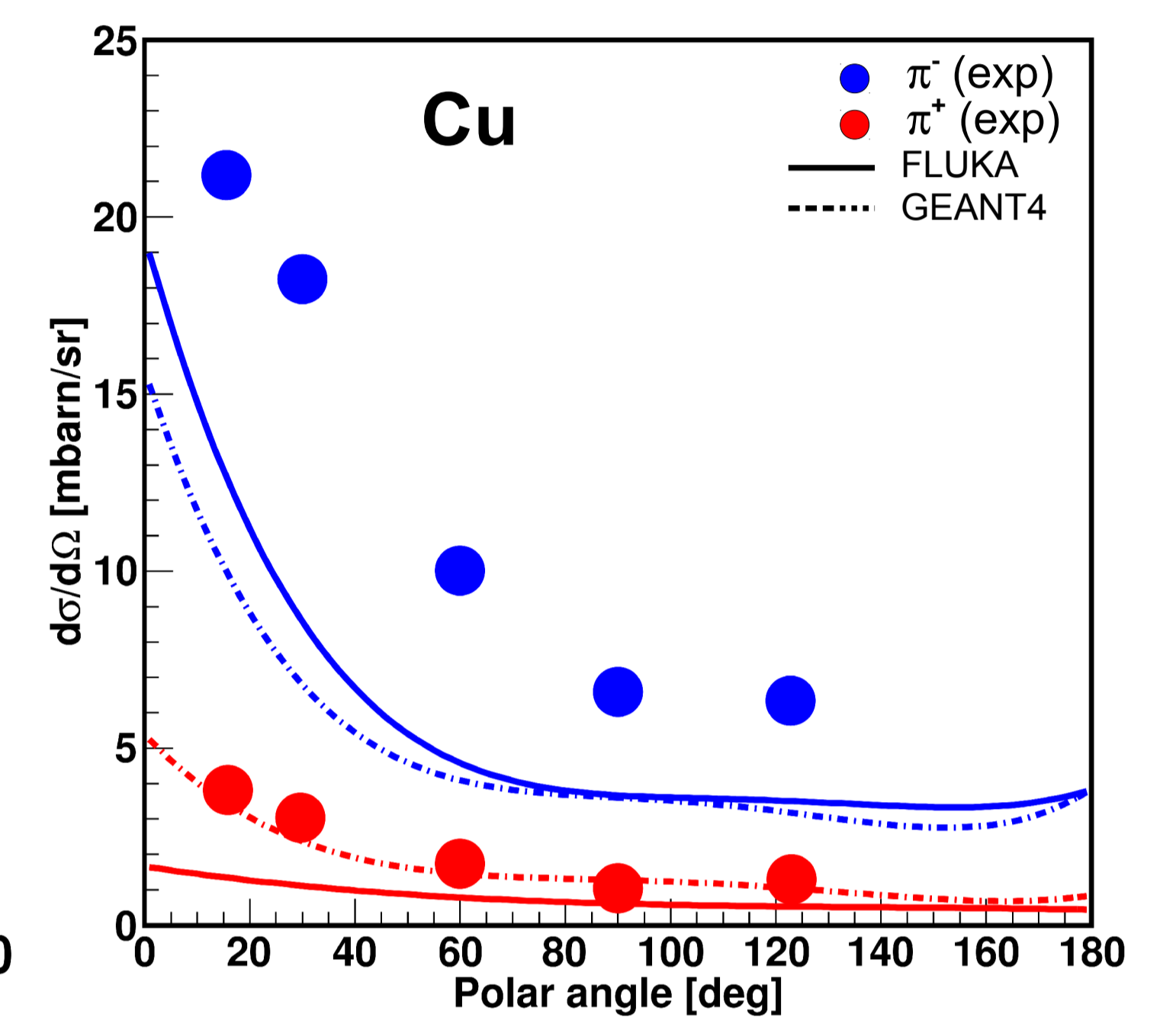
## Pion production by 1.5 GeV/c protons (HARP-CDP)

Cross-sections of the production of secondary  $\pi^+$  (red points) and  $\pi^-$  (blue points) compared with the LSND parametrization (the red solid line refers to  $\pi^+$ , and the blue one to  $\pi^-$ ).



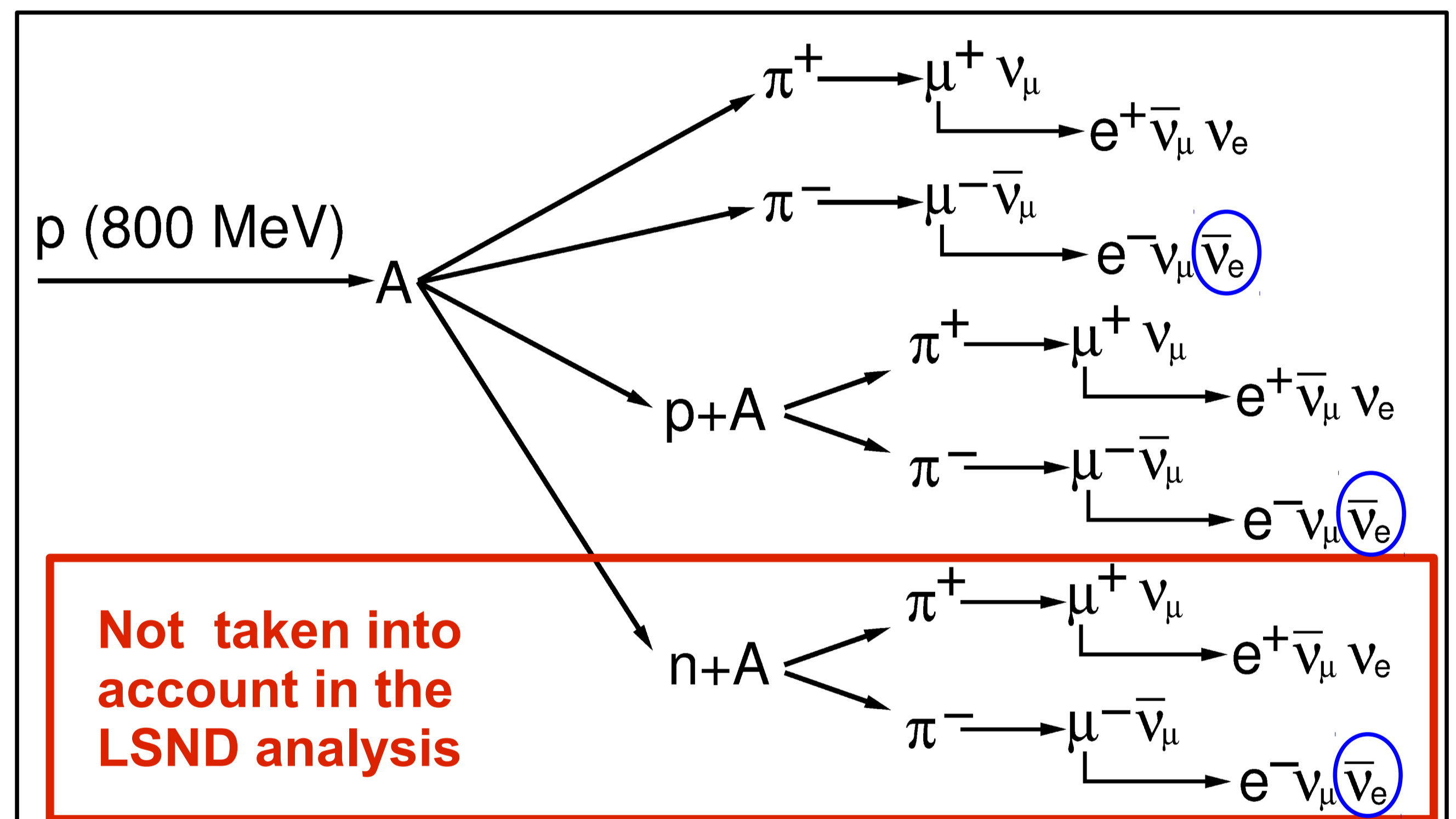
## Pion production by 600 MeV neutrons

K.O. Oganesian, JETP 54 (1968) 1273



## HARP-CDP simulation of the LSND $\bar{\nu}_e$ background:

- We “emulated” the LSND procedure of neutrino flux calculation to demonstrate understanding of the LSND geometry
- Two independent simulation programs
- Geant4 and FLUKA used to simulate hadron production
- HARP-CDP data used to tune pion production by protons
- Pion production by neutrons tuned according to *K.O. Oganesian, JETP 54 (1968) 1273*
- Neutron production by protons tuned according to *V.N. Baturin et al., JETP Lett., 30 (1979) 86*



## Pion rates and neutrino flux from muon decays at rest

	LSND published (runs 1993-1995)	Our 'emulation' of the LSND procedure	Geant4 (default)	FLUKA (default)	Geant4 $\oplus$ exp. data		FLUKA $\oplus$ exp. data	
					w/o neutrons	with neutrons	w/o neutrons	with neutrons
$\pi^+$ [ $\text{PoT}^{-1}$ ]		0.107 – 0.159	0.200	0.212	0.196	0.203	0.206	0.214
$\pi^-$ [ $\text{PoT}^{-1}$ ]		0.026 – 0.051	0.040	0.045	0.037	0.072	0.035	0.071
$\bar{\nu}_\mu, \bar{\nu}_e$ DAR flux [ $\text{PoT}^{-1} \text{cm}^{-2}$ ]	$0.65 \times 10^{-12}$	$(0.52 - 0.77) \times 10^{-12}$	$0.90 \times 10^{-12}$	$0.72 \times 10^{-12}$	$0.88 \times 10^{-12}$	$1.20 \times 10^{-12}$	$0.70 \times 10^{-12}$	$0.99 \times 10^{-12}$
$\bar{\nu}_\mu, \nu_e$ DAR flux [ $\text{PoT}^{-1} \text{cm}^{-2}$ ]	$0.80 \times 10^{-9}$	$(0.75 - 1.20) \times 10^{-9}$	$1.20 \times 10^{-9}$	$1.40 \times 10^{-9}$	$1.20 \times 10^{-9}$	$1.30 \times 10^{-9}$	$1.30 \times 10^{-9}$	$1.40 \times 10^{-9}$

**CONCLUSIONS:** the background of LSND's “anomalous  $\bar{\nu}_e$  signal” was underestimated by nearly a factor of two, and its systematic error was underestimated by at least a factor of two. The causes were too small pion production cross-sections by protons and the neglect of pion production by neutrons, which, unlike protons, predominantly produce  $\pi^-$  rather than  $\pi^+$ .

**THE PUBLISHED  $3.8 \sigma$  SIGNIFICANCE OF THE LSND SIGNAL IS LOST.**

We are grateful to M. Sung for providing us with the LSND parametrization code  
 Geant4 refers to the Geant4 software (S.Agostinelli et al., NIM A506 (2003) 250-303; J. Allison et al., IEEE Trans. Nucl. Sci. 53 (2006) 270-278)  
 FLUKA refers to the FLUKA software (G.Battistoni et al, AIP Conf. Proc. 896, 31-49, (2007); A.Fasso et al., CERN-2005-10 (2005), INFN/TC\_05/11, SLAC-R-773)