

Project X: Accelerator and Experiments:

V 's

μ 's

K 's

Stephen Parke

Fermilab

Jan 22, 2010-Aspen

Physics of Project X:

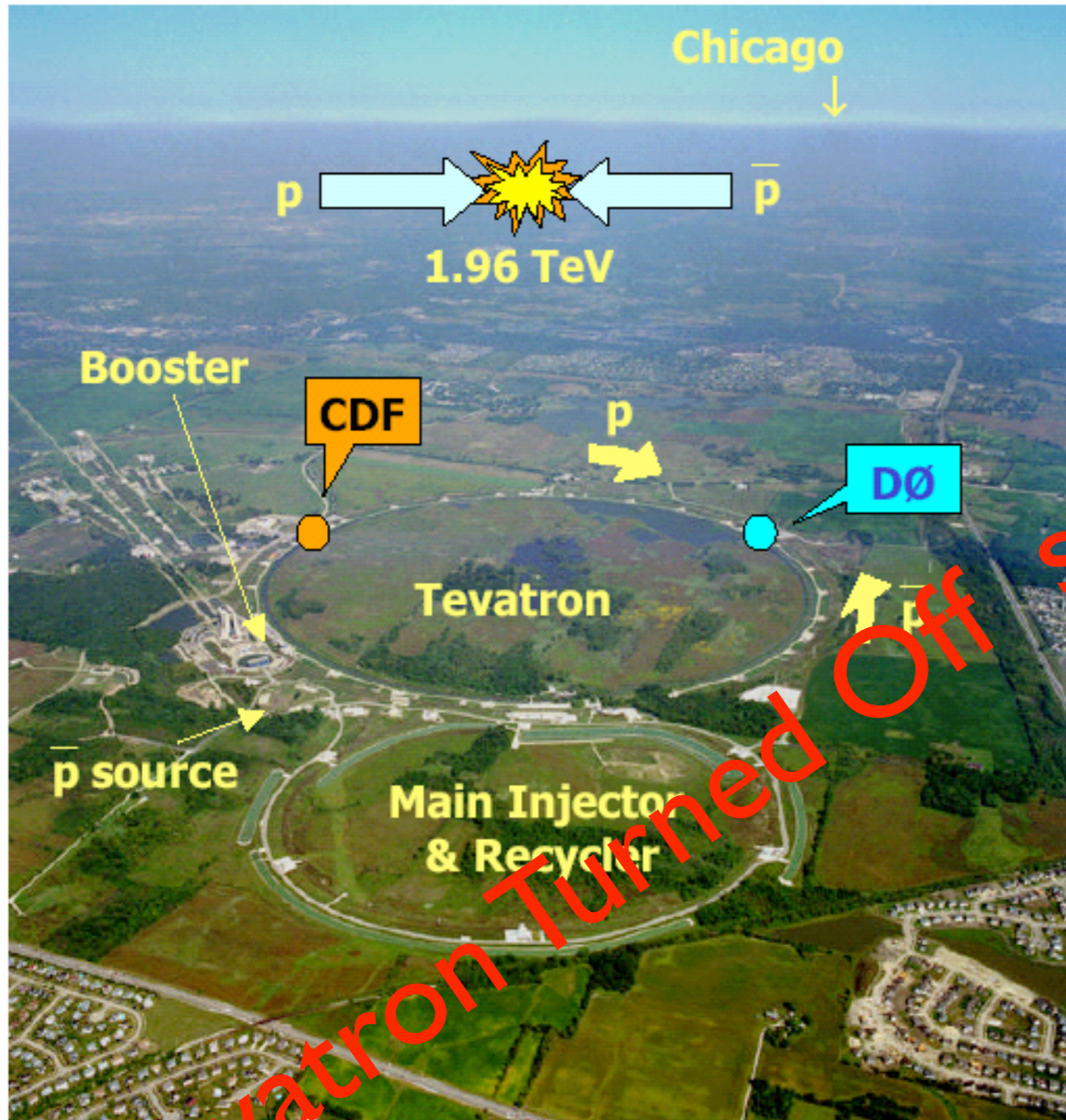
V 's

μ 's

K 's

Stephen Parke
Fermilab
Jan 22, 2010-Aspen

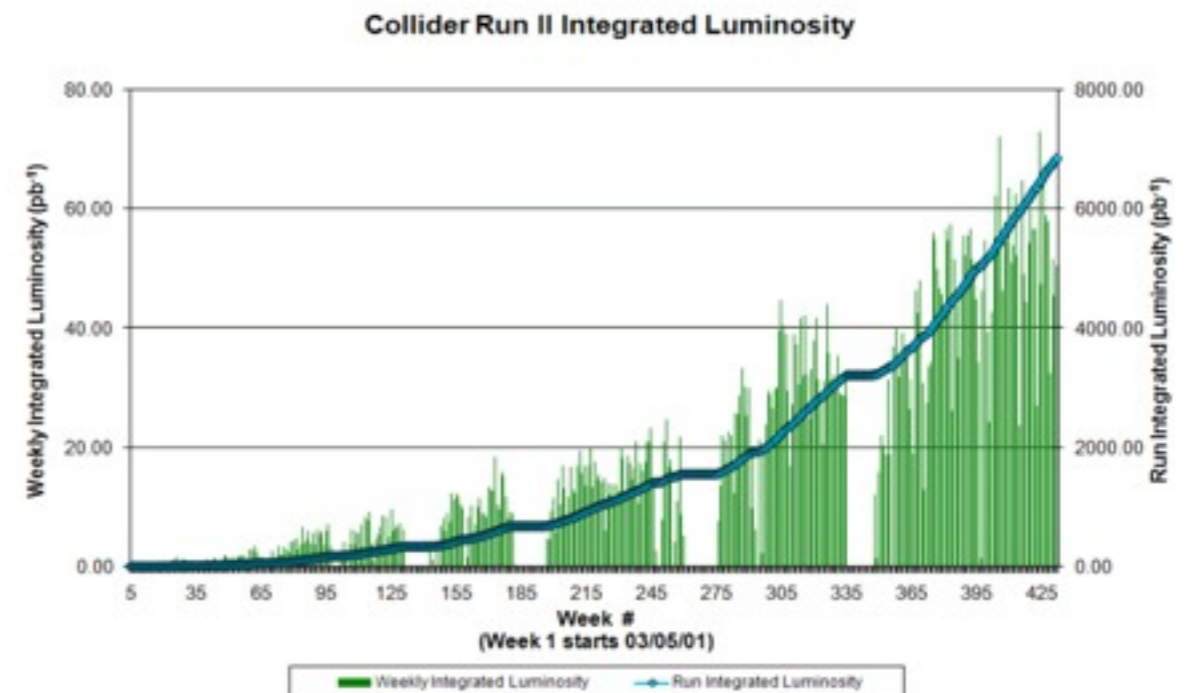
Tevatron: CDF & D0



- 36x36 bunches
- bunch crossing 396 ns
- Run II started in March 2001
- Peak Luminosity: $3.5 \times 10^{32} \text{ cm}^{-2} \text{ sec}^{-1}$
- Run II delivered: $\sim 7 \text{ fb}^{-1}$

Peak Integrated Luminosity:

$$3.5 \times 10^{32} \text{ cm}^{-2} \text{ sec}^{-1} \times 3 \times 10^7 \text{ sec} \approx 10 \text{ fb}^{-1} / \text{yr}$$

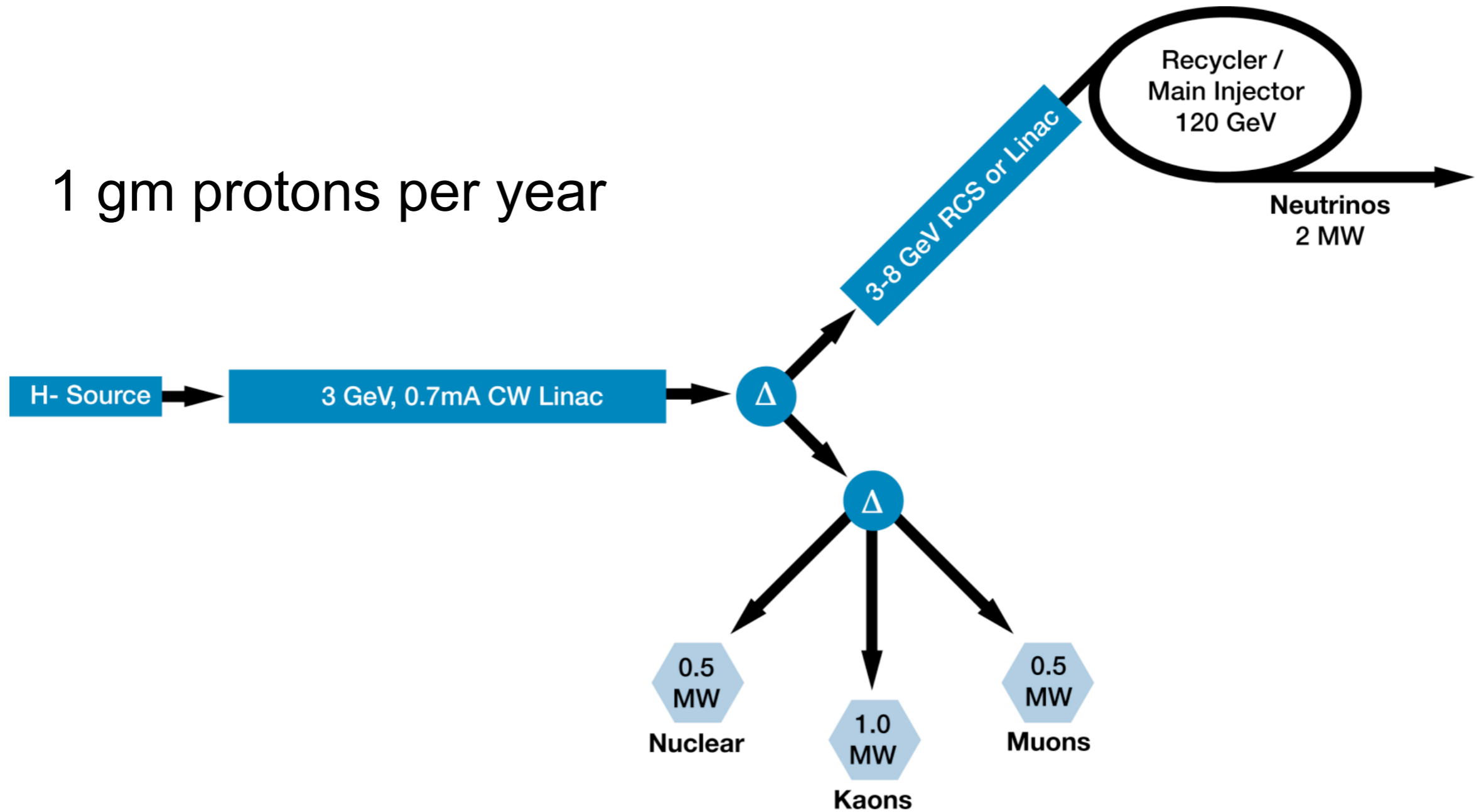


Peak \bar{p} Collection Rate:
 $7 \times 10^7 / \text{sec}$

Peak \bar{p} Burn Rate:
 $2 \times 3.5 \times 10^{32} \text{ cm}^{-2} \text{ sec}^{-1} \times 100 \text{ mb} = 7 \times 10^7 / \text{sec}$

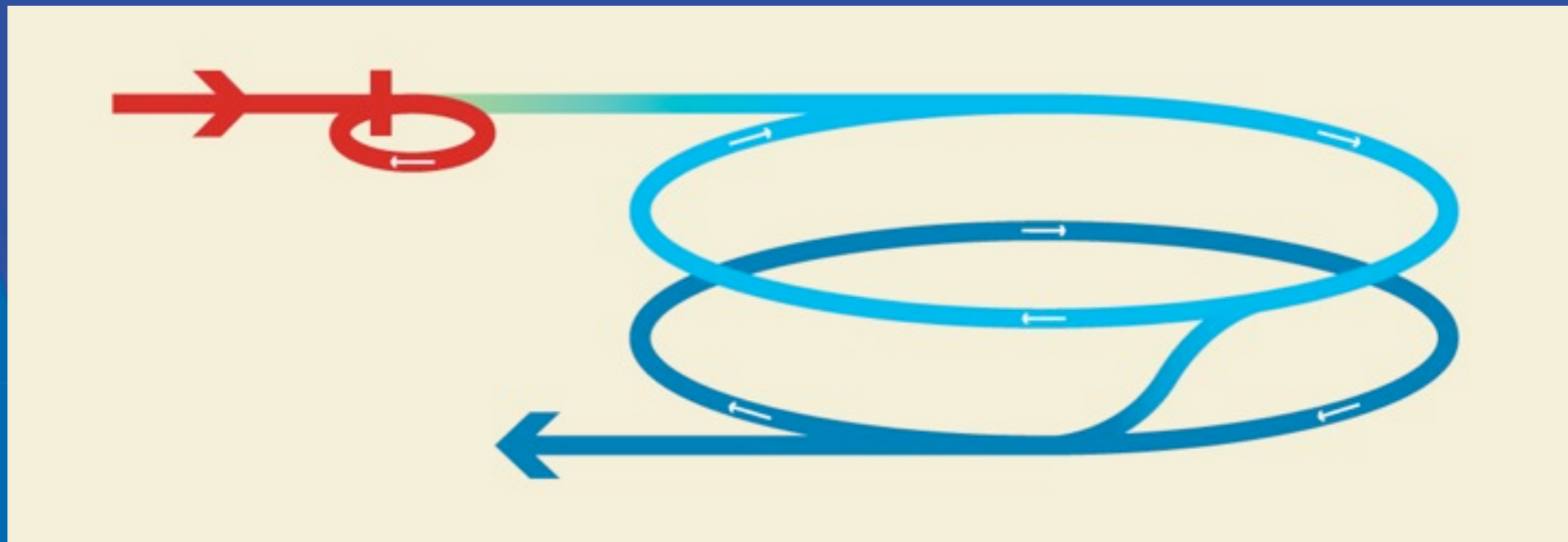
Project X: (Proton Driver)

1 gm protons per year



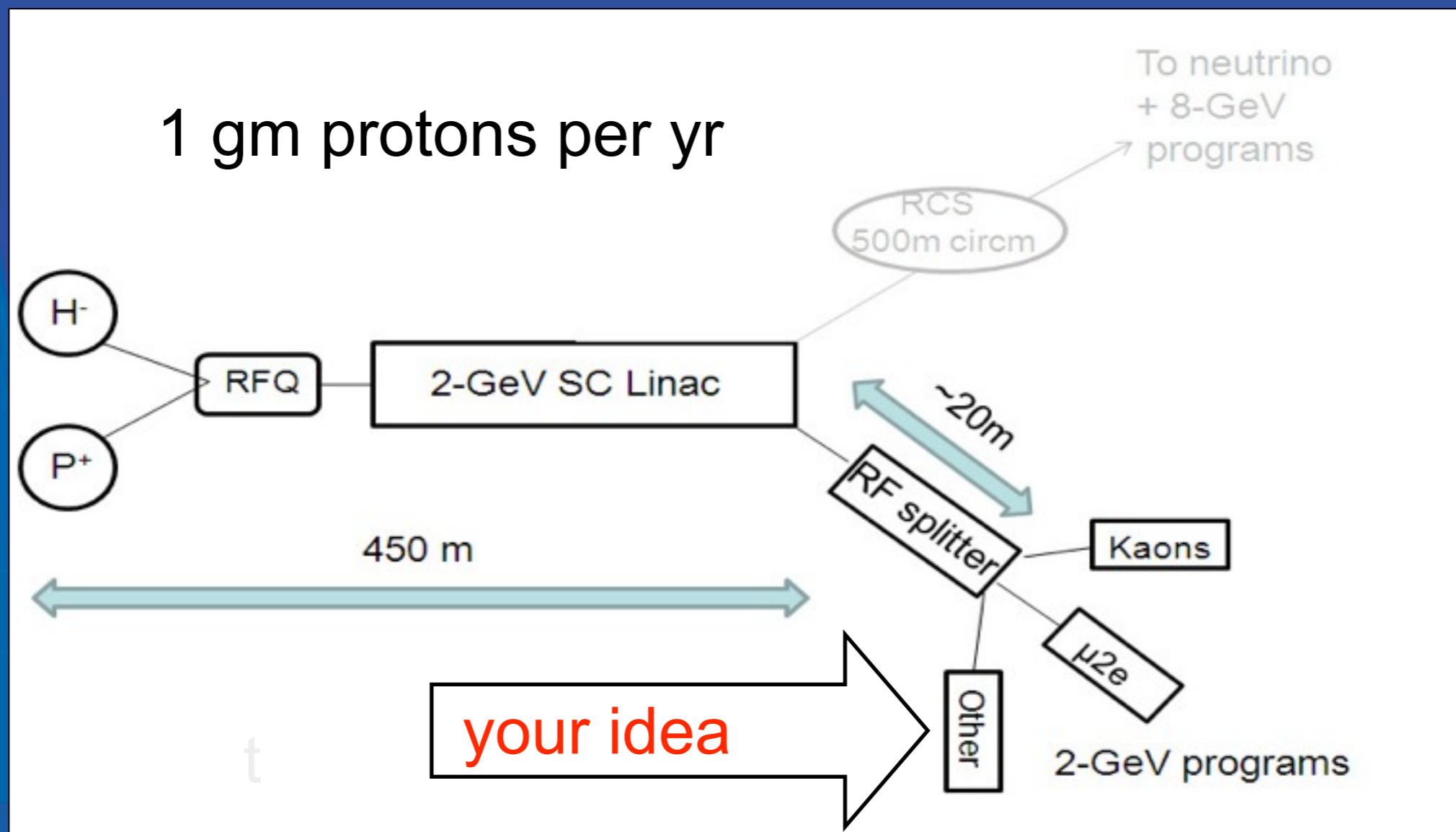
Project X and LBNE to Homestake

- 5% of the time line, the 3 GeV linac feeds a simple Rapid Cycling Synchrotron (RCS), 500m circumference, to strip, accumulate and boost the energy to 8 GeV
- Six pulses of the SAB are transferred to the recycler, filling the existing recycler, and every 1.4 sec transferred to the Main Injector for acceleration to high energies (60 GeV to 120 GeV)



Project X and 2 GeV beams

- The greatest potential for rare processes comes from 2 MW continuous beam. Intensity experiments need continuous beam: pile up is the main limitation in pulsed beams



Neutrinos:

Mixing Matrix:

$$|\nu_e, \nu_\mu, \nu_\tau\rangle_{flavor}^T = U_{\alpha i} |\nu_1, \nu_2, \nu_3\rangle_{mass}^T$$

$$U_{\alpha i} = \begin{pmatrix} 1 & & & \\ & c_{23} & s_{23} & \\ & -s_{23} & c_{23} & \\ & & & 1 \end{pmatrix} \begin{pmatrix} c_{13} & & s_{13}e^{-i\delta} & \\ & 1 & & \\ -s_{13}e^{i\delta} & & c_{13} & \\ & & & 1 \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & \\ -s_{12} & c_{12} & \\ & & 1 \end{pmatrix} \begin{pmatrix} 1 & & \\ & e^{i\alpha} & \\ & & e^{i\beta} \end{pmatrix}$$

Atmos. L/E $\mu \rightarrow \tau$ Atmos. L/E $\mu \leftrightarrow e$ Solar L/E $e \rightarrow \mu, \tau$ $0\nu\beta\beta$ decay

500km/GeV

15km/MeV

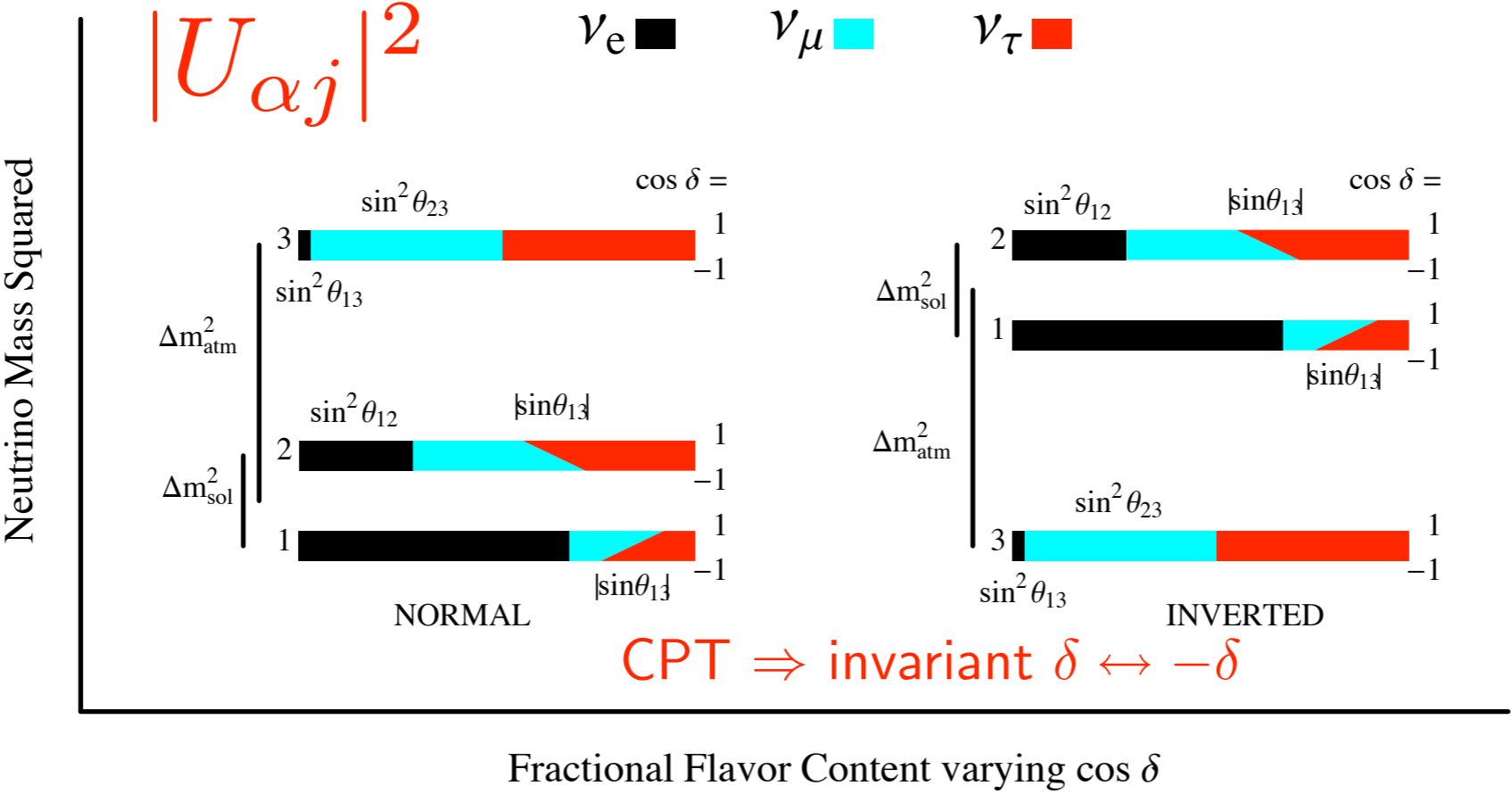
solar/KamLAND

Reactor/LBL

$$= \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{13}s_{23}c_{12}e^{i\delta} & c_{23}c_{12} - s_{13}s_{23}s_{12}e^{i\delta} & c_{13}s_{23} \\ s_{23}s_{12} - s_{13}c_{23}c_{12}e^{i\delta} & -s_{23}c_{12} - s_{13}c_{23}s_{12}e^{i\delta} & c_{13}c_{23} \end{pmatrix}$$

Atmospheric Nus

Neutrinos



$$\delta m_{sol}^2 = +7.6 \times 10^{-5} \text{ eV}^2$$

$$|\delta m_{atm}^2| = 2.4 \times 10^{-3} \text{ eV}^2$$

$$|\delta m_{sol}^2| / |\delta m_{atm}^2| \approx 0.03$$

$$\sin^2 \theta_{12} \sim 1/3$$

$$\sin^2 \theta_{23} \sim 1/2$$

$$\sin^2 \theta_{13} < 3\%$$

$$\sqrt{\delta m_{atm}^2} = 0.05 \text{ eV} < \sum m_{\nu_i} < 0.5 \text{ eV} = 10^{-6} * m_e$$

$$0 \leq \delta < 2\pi$$

At 2σ we have the following limits:

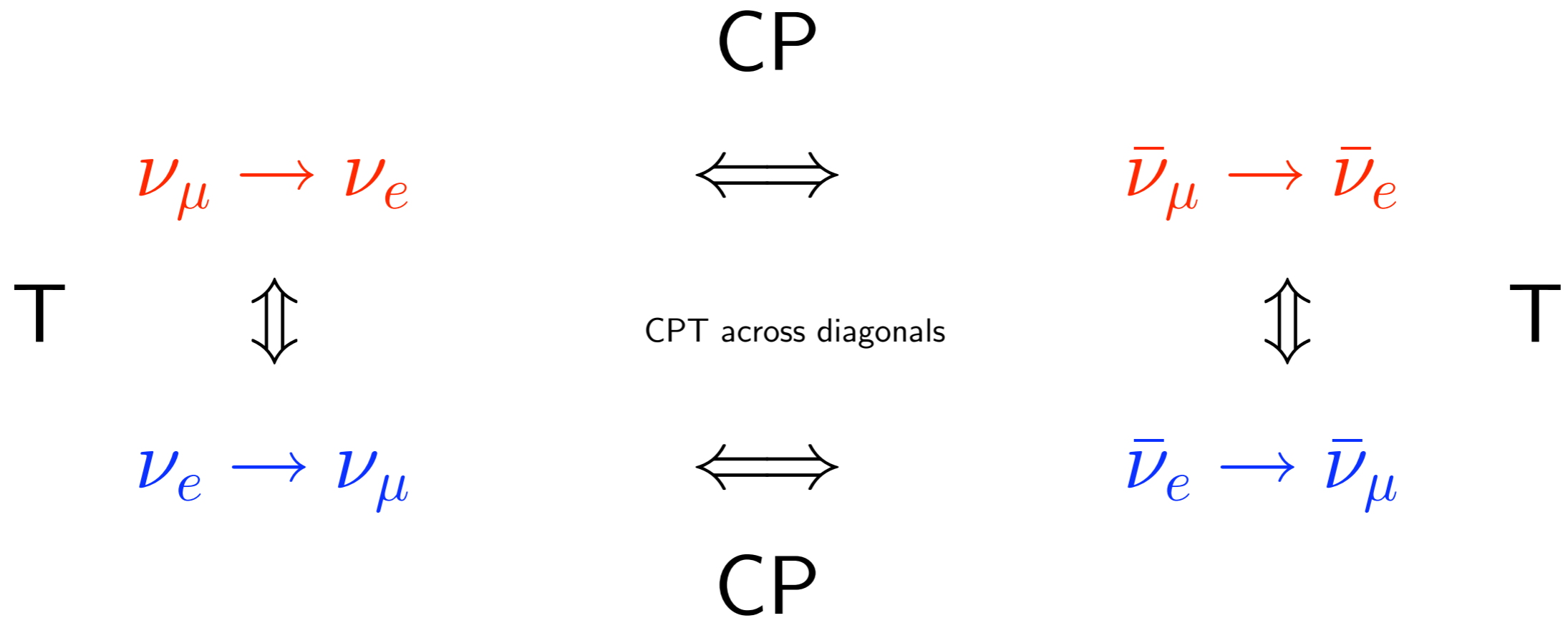
$$\begin{aligned}\sin^2 \theta_{13} &< 0.04 \\ |\sin^2 \theta_{12} - \frac{1}{3}| &< 0.04 \\ |\sin^2 \theta_{23} - \frac{1}{2}| &< 0.12\end{aligned}$$

Close to Tri-Bi-Maximal: accident or symmetry ?

In numerous models:

$$\sin^2 \theta_{13}, |\sin^2 \theta_{12} - \frac{1}{3}|, |\sin^2 \theta_{23} - \frac{1}{2}| \sim \left(\frac{\delta m_{21}^2}{\delta m_{31}^2} \right)^n$$

Experiment has probed down to $n \approx 1/2$ to 1 !!!



- First Row: Superbeams where ν_e contamination $\sim 1\%$
- Second Row: ν -Factory or β -Beams, no beam contamination

Vacuum LBL:

$$\nu_{\mu} \rightarrow \nu_e$$

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

$$\Delta_{ij} = \delta m_{ij}^2 L / 4E$$

CP violation !!!

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \sin \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \sin \Delta_{21}$

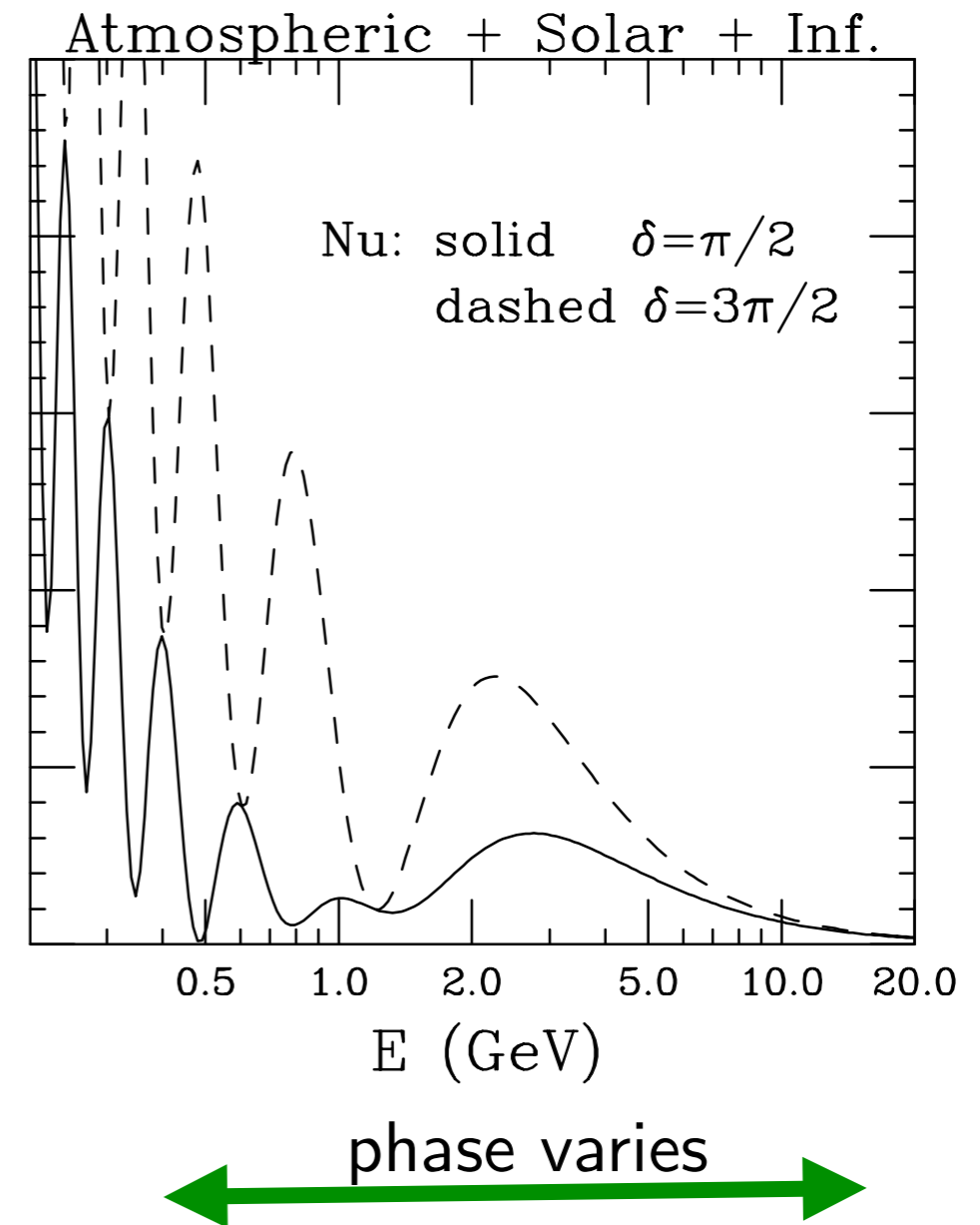
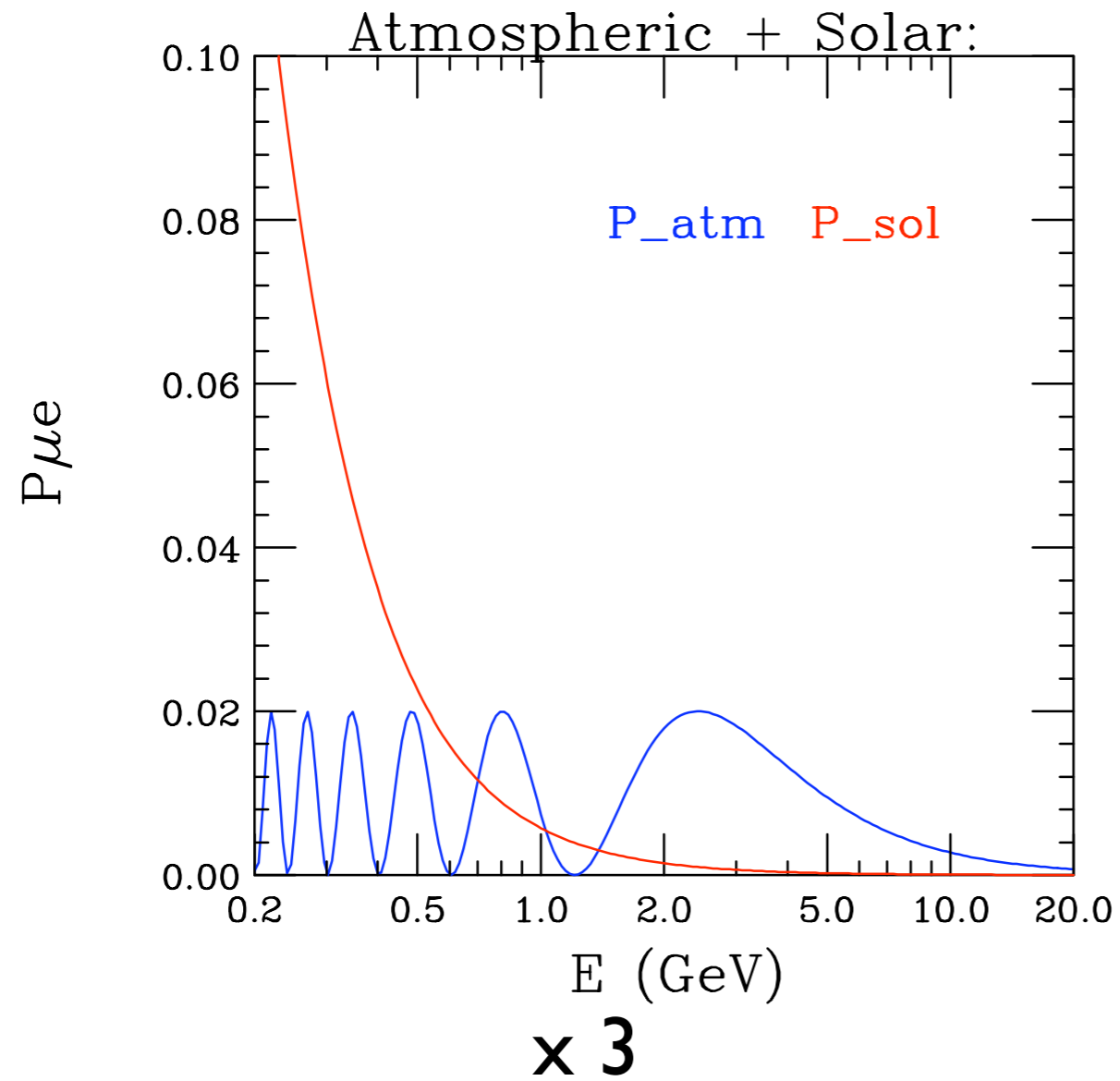
$$P_{\mu \rightarrow e} \approx P_{atm} + 2\sqrt{P_{atm}P_{sol}} \cos(\Delta_{32} \pm \delta) + P_{sol}$$

only CPV

$$\cos(\Delta_{32} \pm \delta) = \cos \Delta_{32} \cos \delta \mp \sin \Delta_{32} \sin \delta$$

$$P(\nu_\mu \rightarrow \nu_e) \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} + \delta)} + \sqrt{P_{sol}} \right|^2$$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$



In Matter:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

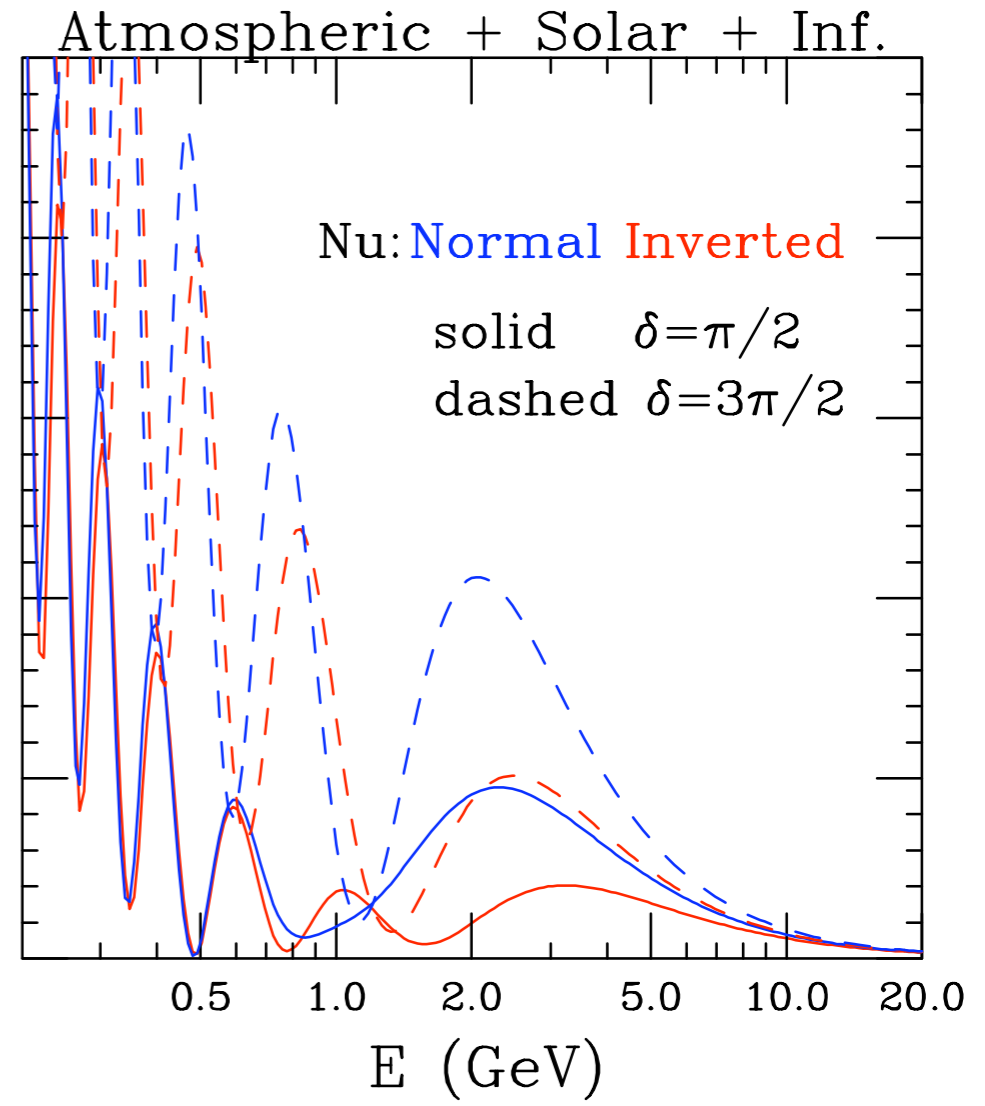
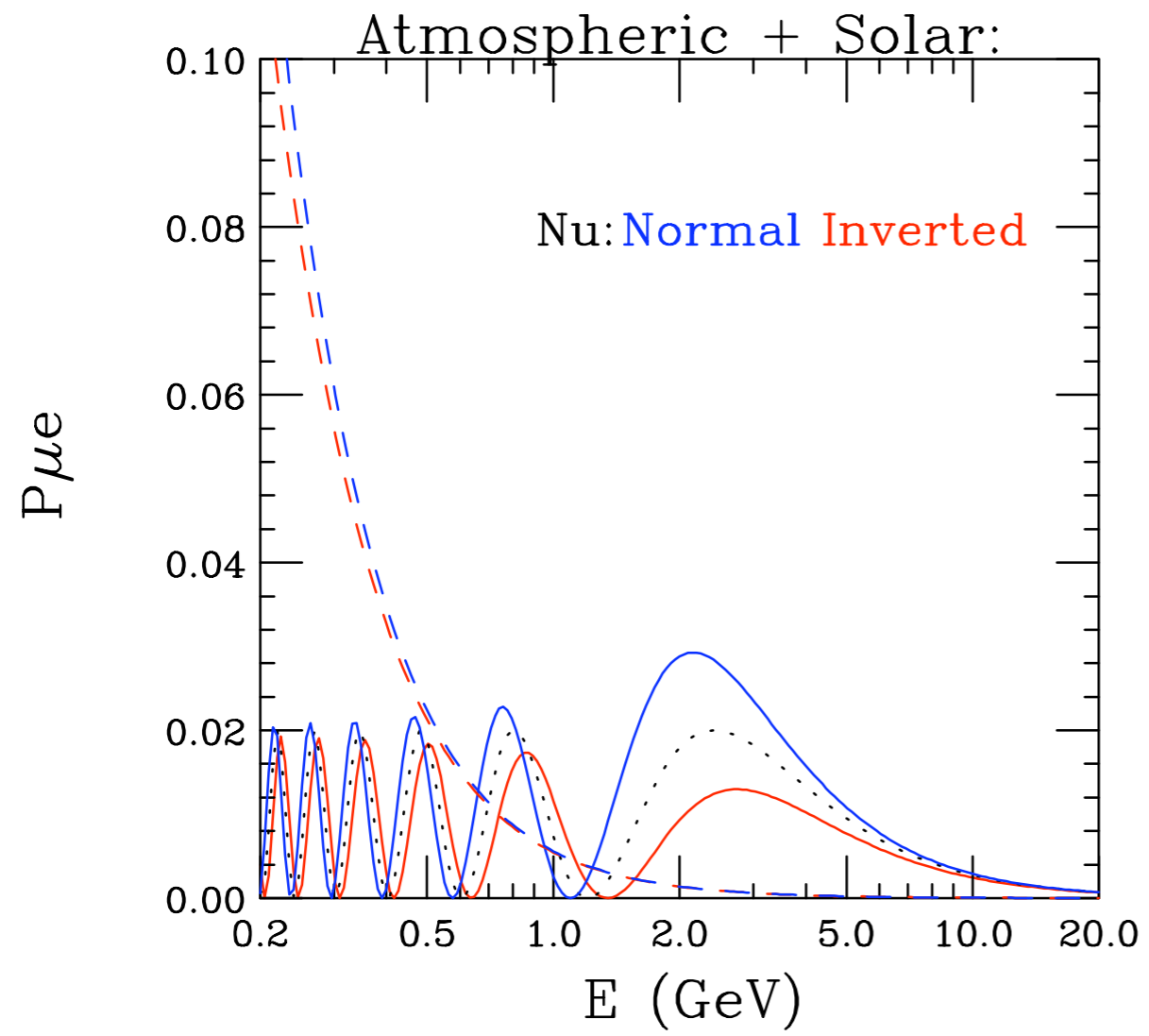
where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

Anti-Nu: Normal Inverted
dashes $\delta = \pi/2$
solid $\delta = 3\pi/2$



$$\nu_\mu \longrightarrow \nu_e$$

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2(\Delta_{31} - aL)}{(\Delta_{31} - aL)^2} \Delta_{31}^2$$

$$+ 2 \sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12} \cos \theta_{13}$$

$$* \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \Delta_{31} \frac{\sin(aL)}{(aL)} \Delta_{21}$$

CPV

$$* (\cos \Delta_{32} \cos \delta - \sin \Delta_{32} \sin \delta)$$

CPC

$$+ \cos^4 \theta_{13} \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(aL)}{(aL)^2} \Delta_{21}^2$$

$$(\Delta_{31} - aL) = \Delta_{31} \left(1 - \frac{aL}{\Delta_{31}}\right) = \Delta_{31} \left(1 - \frac{2\sqrt{2}G_F N_e E}{\delta m_{31}^2}\right)$$

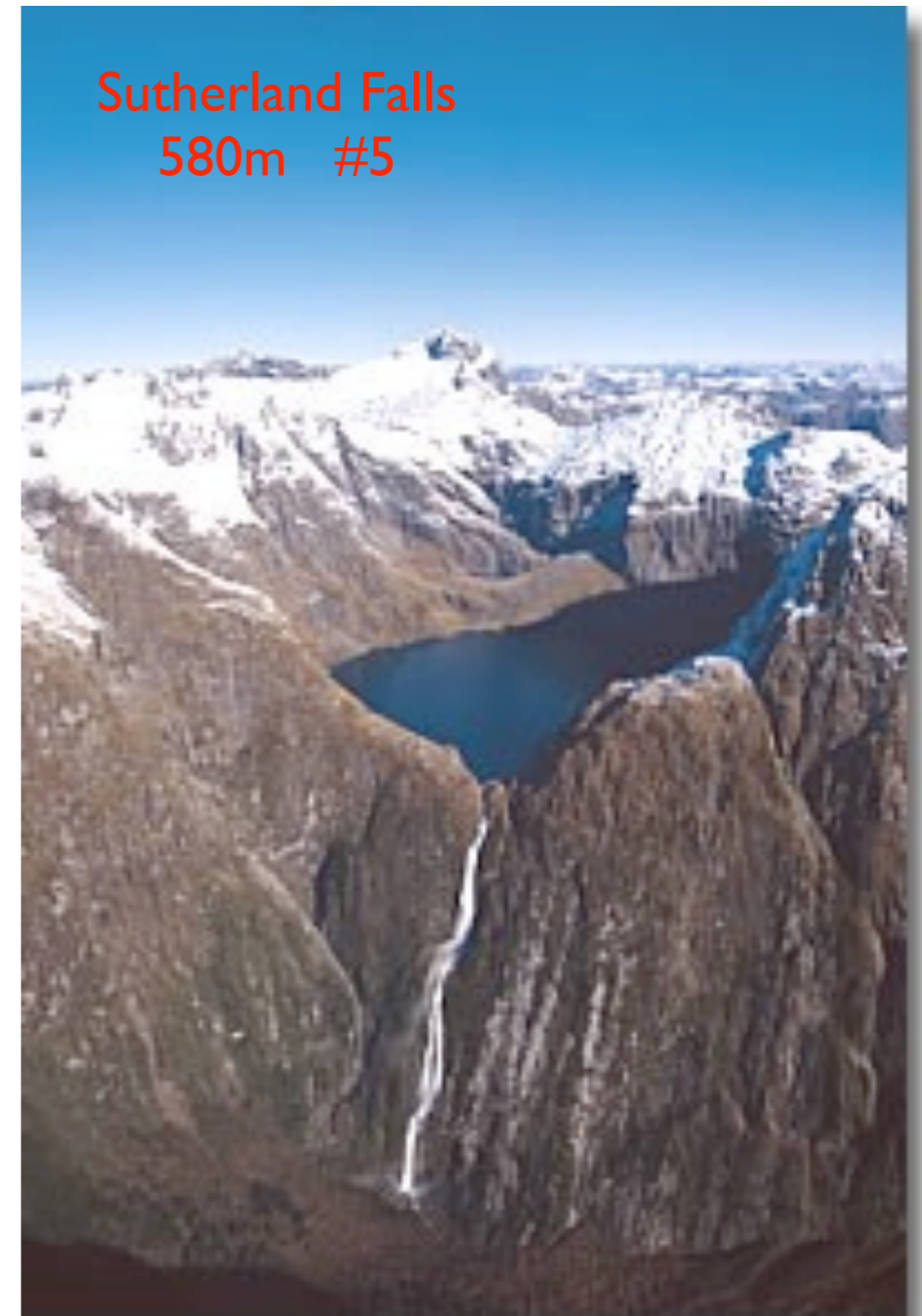
$$\Delta_{32} \approx \Delta_{31}$$

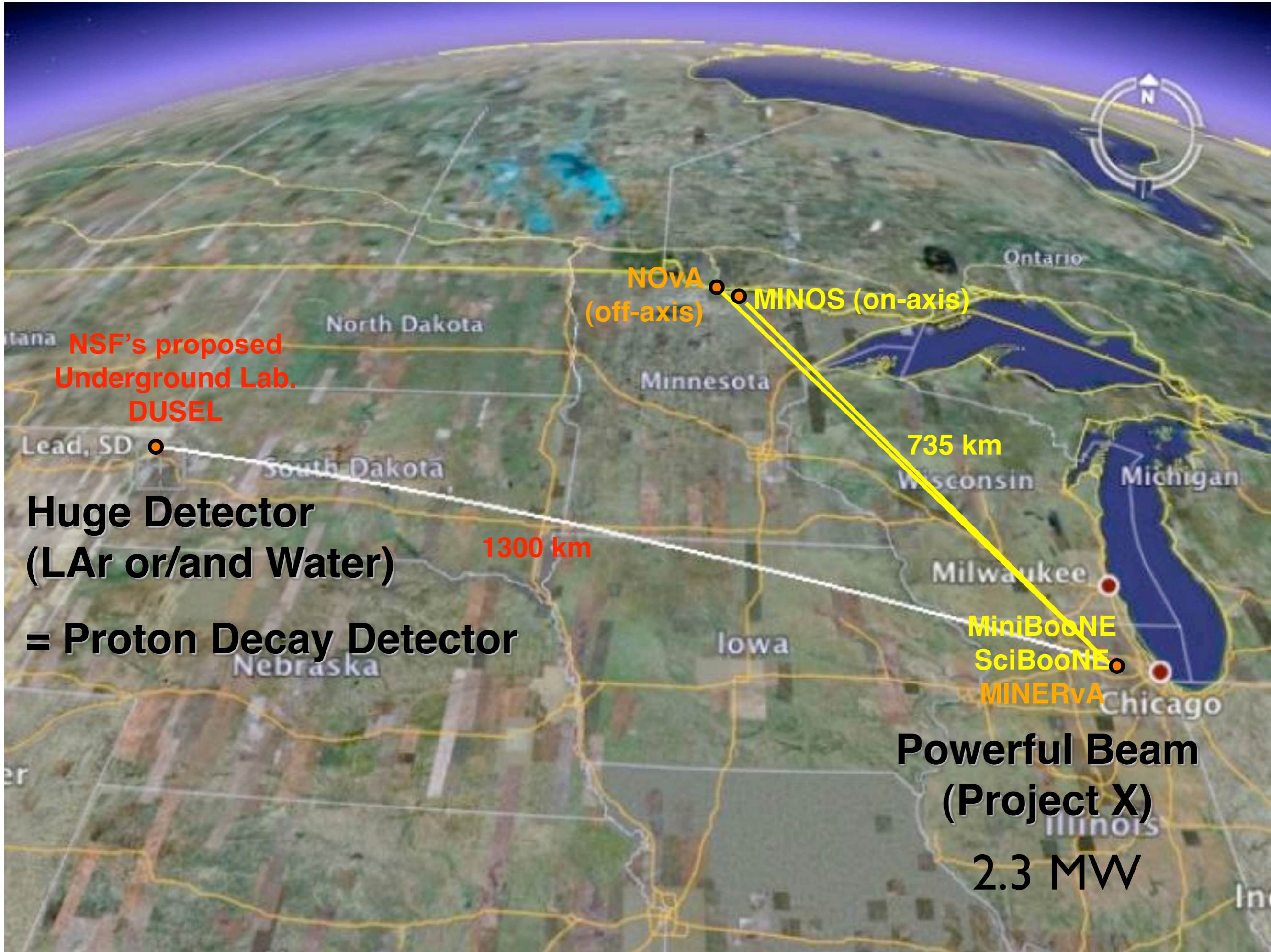
$$J = \sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12} \cos \theta_{13} \sin \delta$$

$$\nu_{\mu} \longrightarrow \nu_e$$

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

Sensitivity to
CPV
and
Mass Hierarchy !!!





NSF's proposed
Underground Lab.
DUSEL

**Huge Detector
(LAr or/and Water)
= Proton Decay Detector**

NOvA
(off-axis) ● ● MINOS (on-axis)

735 km

1300 km

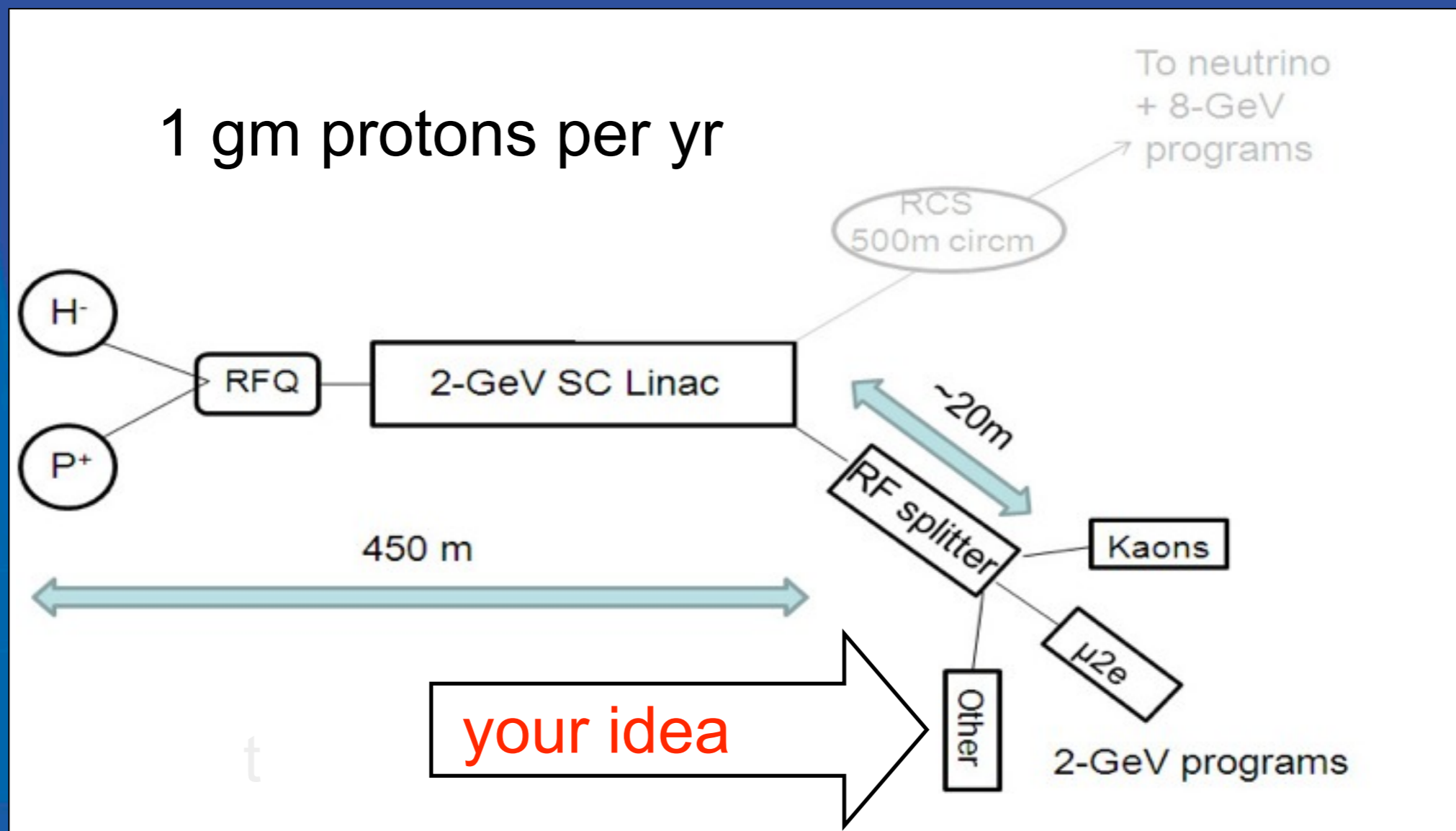
MiniBooNE
SciBooNE
MINERvA

**Powerful Beam
(Project X)**

2.3 MW

Project X and 2 GeV beams

- The greatest potential for rare processes comes from 2 MW continuous beam. Intensity experiments need continuous beam: pile up is the main limitation in pulsed beams

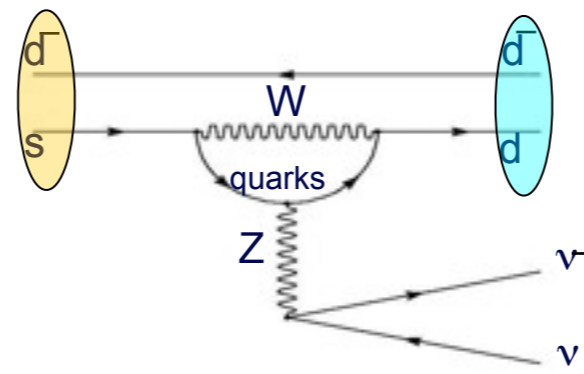
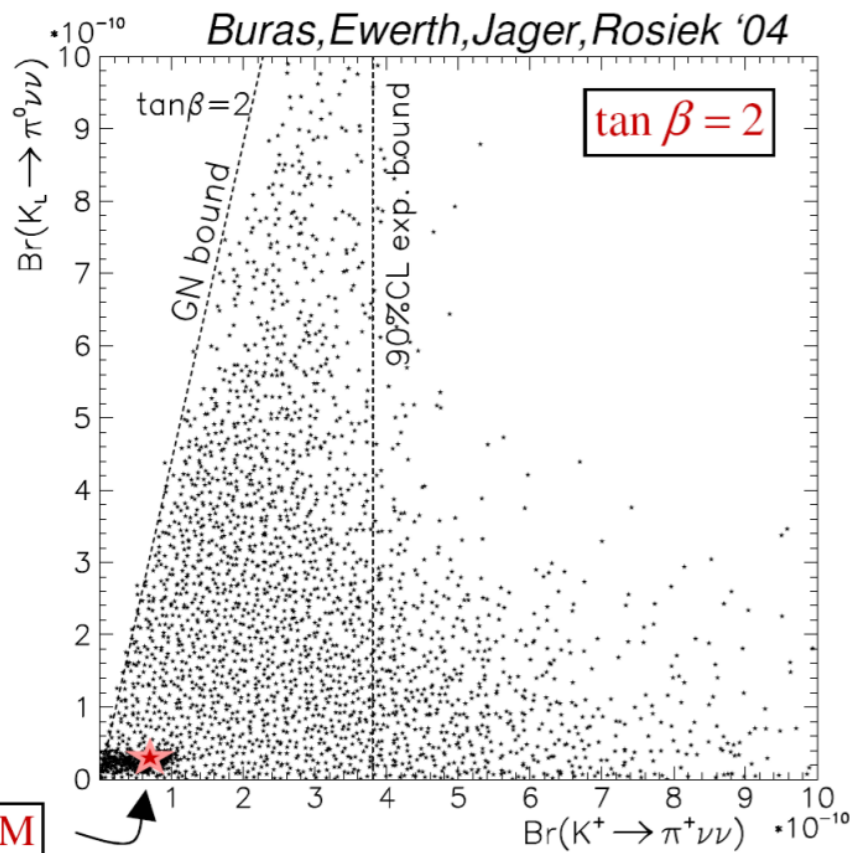


● Kaons

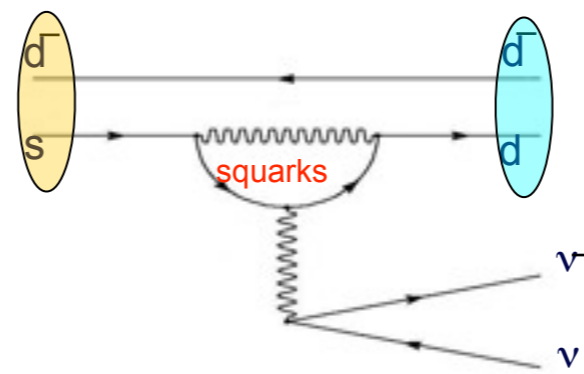
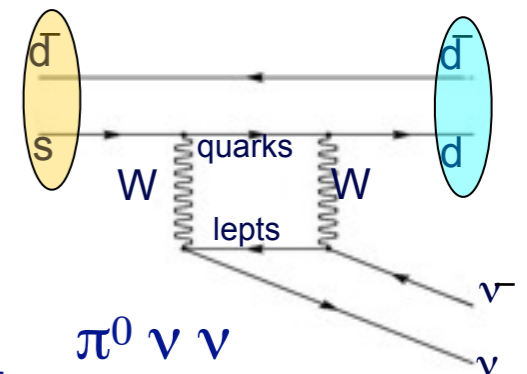
$$K^+ \rightarrow \pi^+ \nu \bar{\nu}$$

$$K_L \rightarrow \pi^0 \nu \bar{\nu}$$

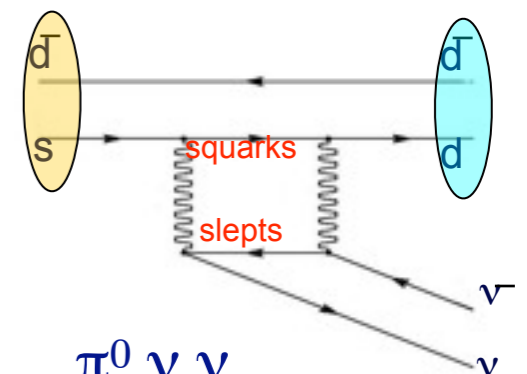
	Theo(SM) $\times 10^{10}$	Exp. $\times 10^{10}$	Experiment
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	0.85 ± 0.07	$1.73^{+1.15}_{-1.05}$	BNL-E787/949
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	0.28 ± 0.04	< 670	KEK-391



SM: $K_L \rightarrow \pi^0 \nu \nu$



BSM: $K_L \rightarrow \pi^0 \nu \nu$



For Statistical Uncertainties \approx Theoretical Uncertainties

~ 1000 events needed in K^+ and K_L !

Project X

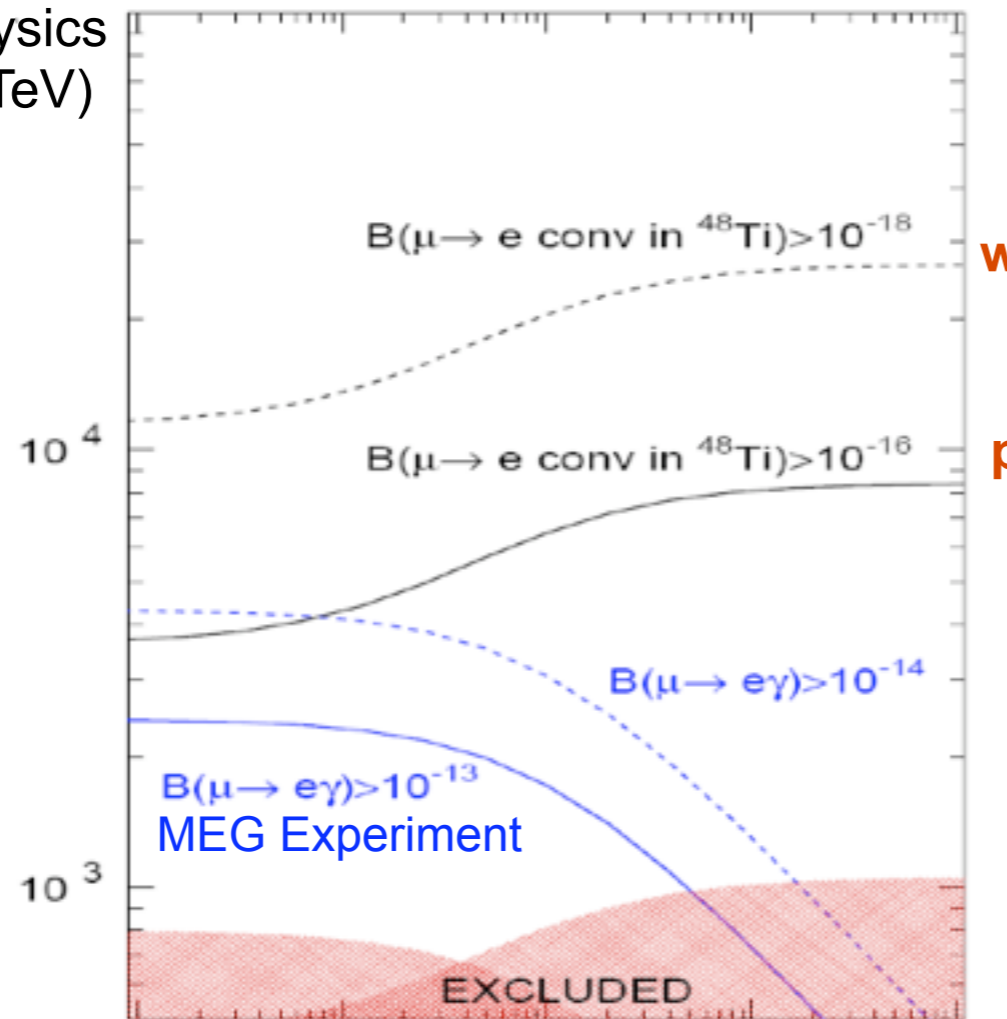
● Muons

$$\mu + N \rightarrow e + N$$

$$(g - 2)_\mu$$

Neutrino Factory

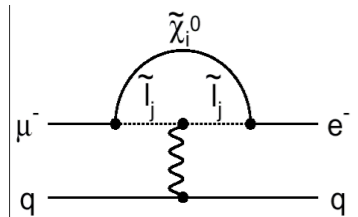
New Physics
Scale (TeV)



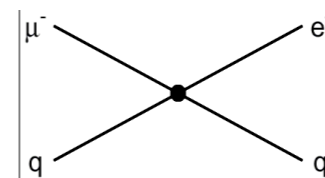
with Project X

pre-Project X

SUSY

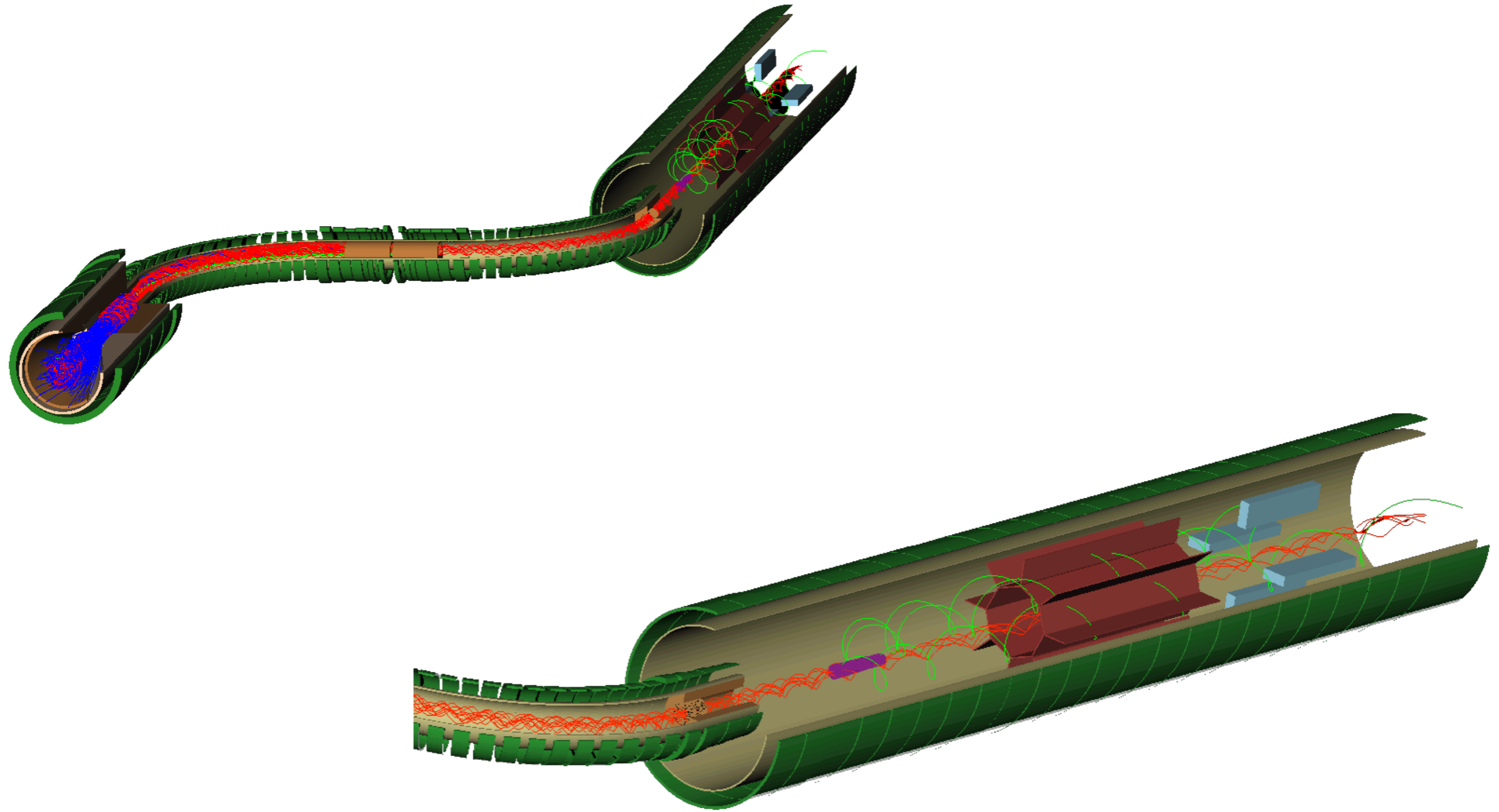


Model Parameter



Mu2e can probe $10^3 - 10^4$ TeV

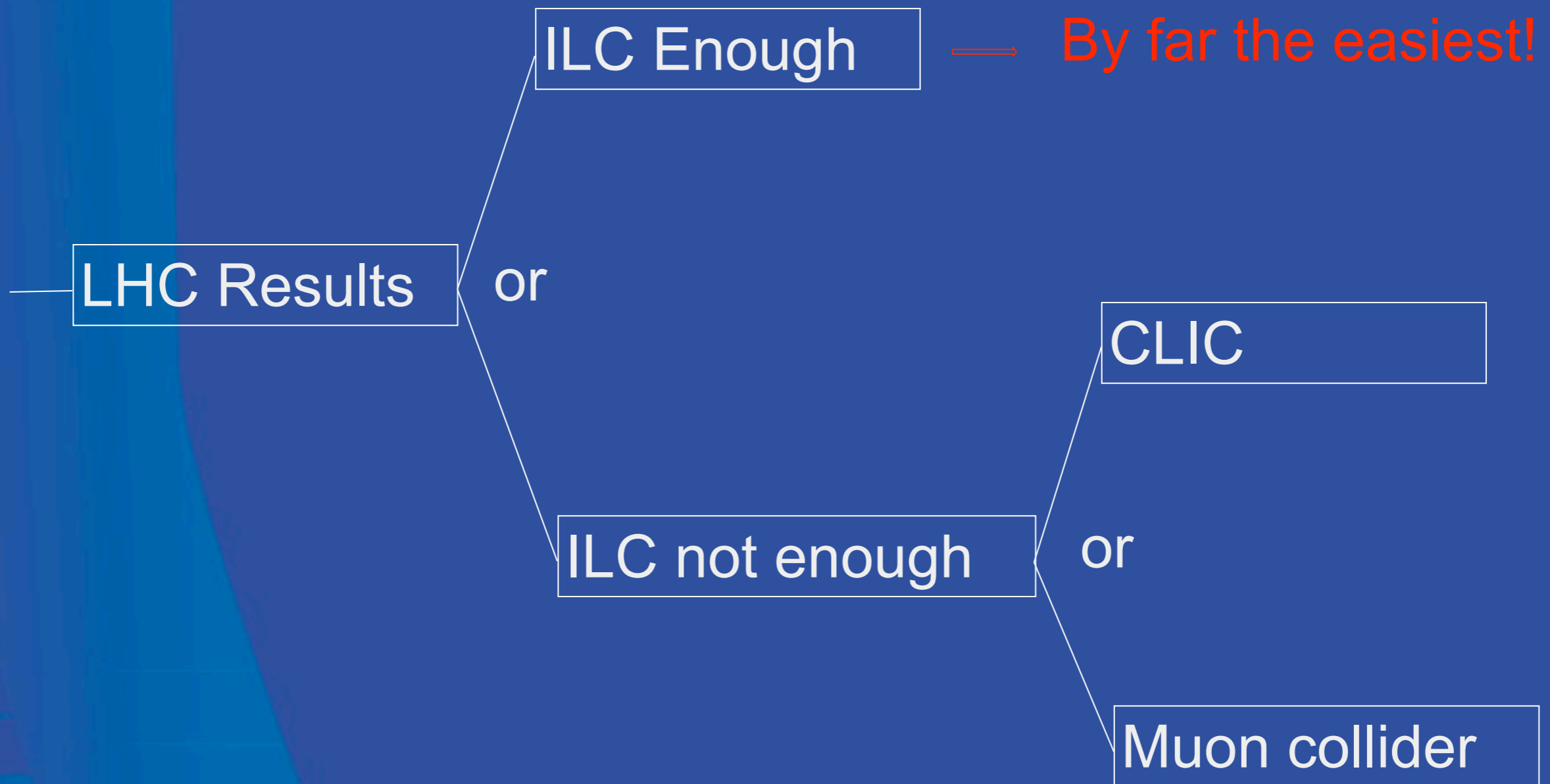
The Mu2e Detector in Detail



Further in the Future:

Muon Collider
and
Neutrino Factory

Lepton colliders beyond LHC

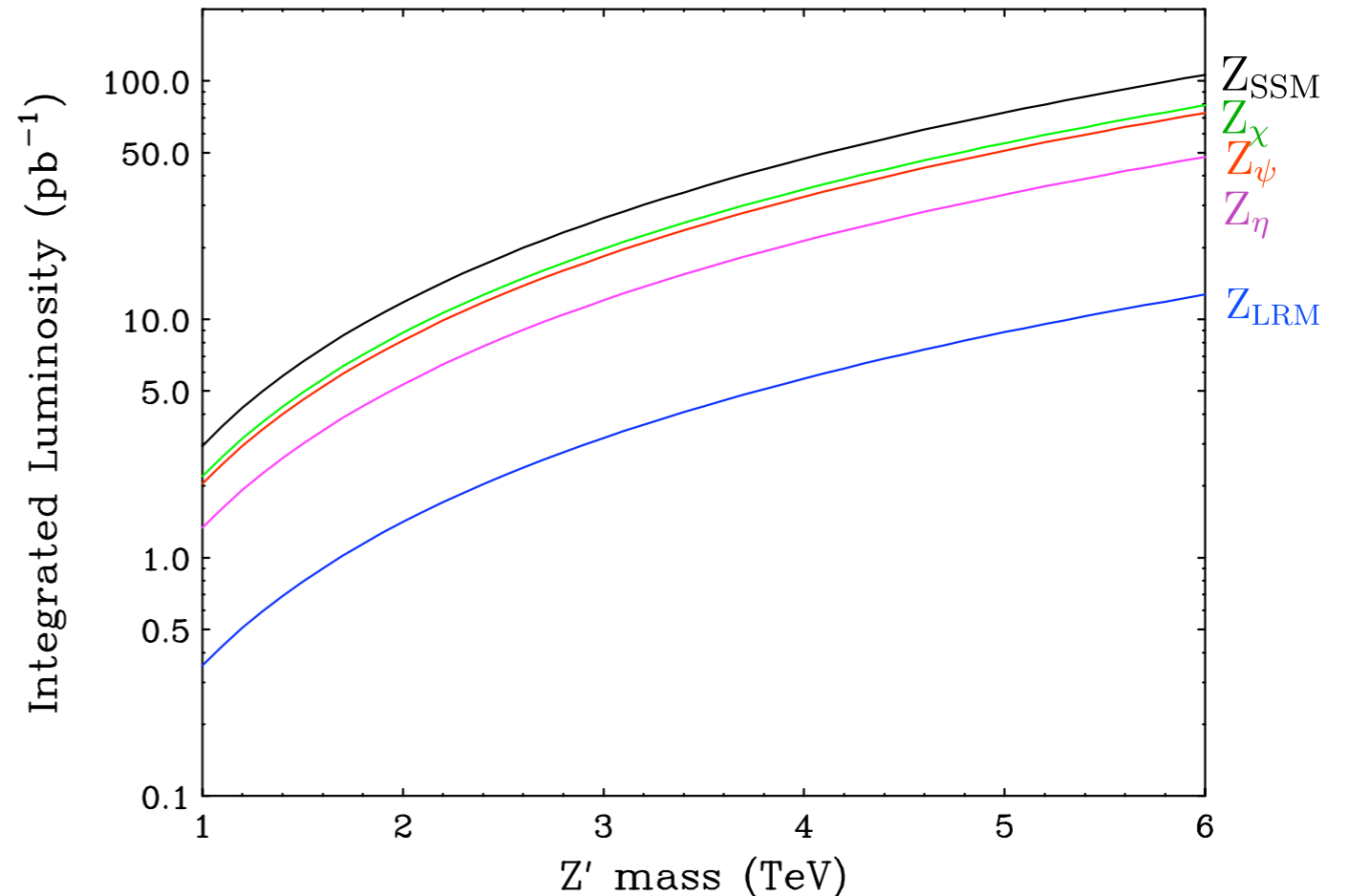


Minimum Luminosity for Physics:

The integrated luminosity required to produce 1000 $\mu^+\mu^- \rightarrow Z'$ events on the peak

- Assuming a new gauge boson: Z'
 - examples: SSM, E6, LRM
 - 5σ discovery limits: 4-5 TeV at LHC (@ 300 fb^{-1})
- For a narrow resonance with $2\Delta E_{\text{beam}} / \Gamma_{\text{resonance}} \ll 1$:

$$\rightarrow R_{\text{peak}} = (2J + 1)3 \frac{B(\mu^+\mu^-)B(\text{visible})}{\alpha_{\text{EM}}^2}$$



Hence minimum luminosity $\rightarrow 0.5\text{--}5.0 \times 10^{30} \text{ cm}^{-2} \text{ sec}^{-1}$
 for $M(Z') \rightarrow 1.5\text{--}5.0 \text{ TeV}$

4 TeV

Muon Collider Conceptual Layout

Project X

Accelerate hydrogen ions to 8 GeV using SRF technology.

Compressor Ring

Reduce size of beam.

Target

Collisions lead to muons with energy of about 200 MeV.

Muon Capture and Cooling

Capture, bunch and cool muons to create a tight beam.

Initial Acceleration

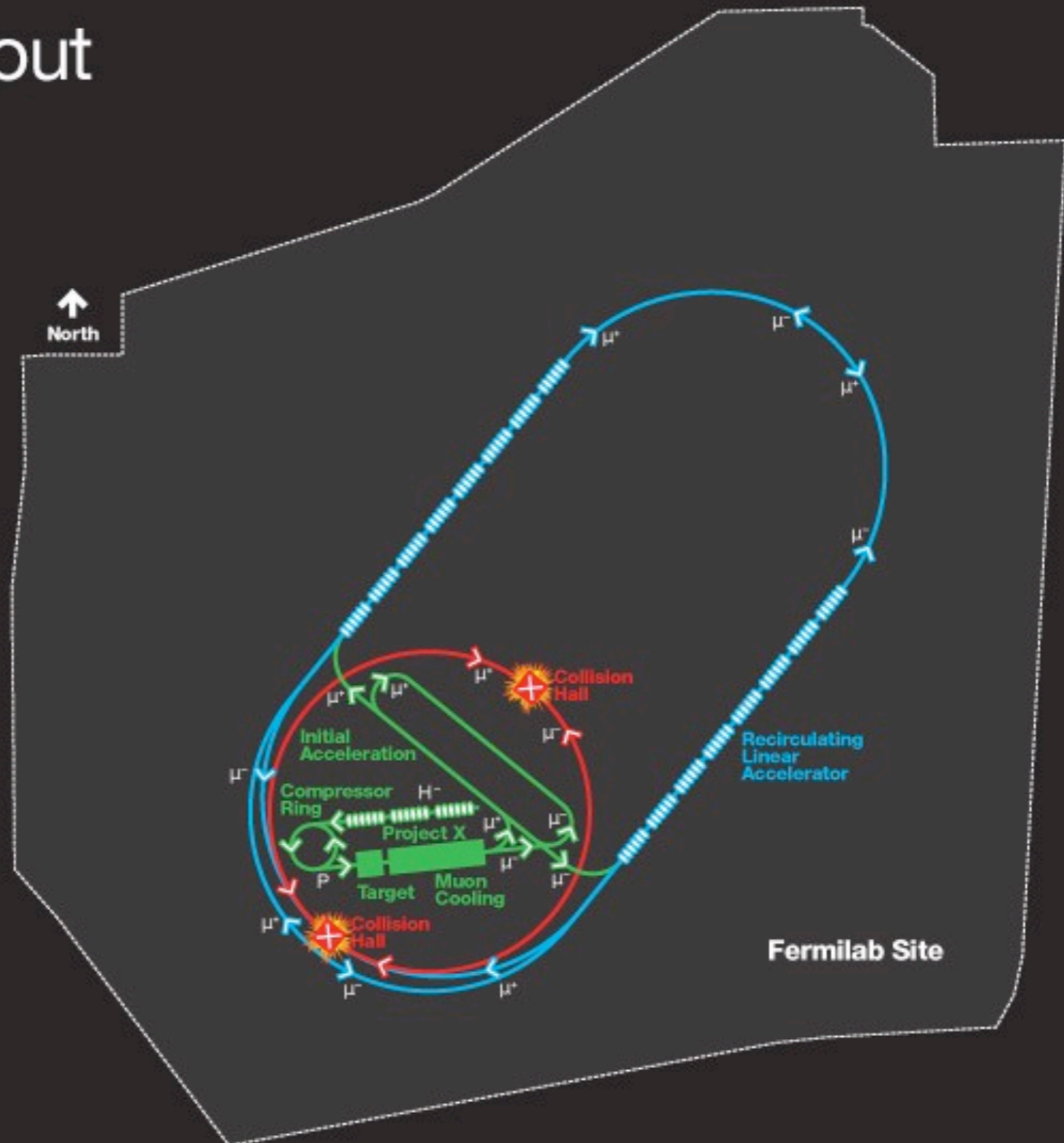
In a dozen turns, accelerate muons to 20 GeV.

Recirculating Linear Accelerator

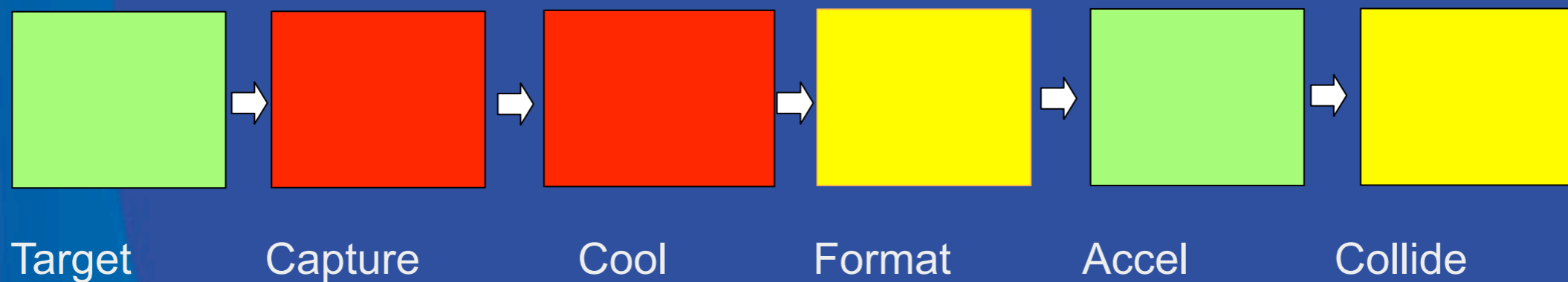
In a number of turns, accelerate muons up to 2 TeV using SRF technology.

Collider Ring

Bring positive and negative muons into collision at two locations 100 meters underground.



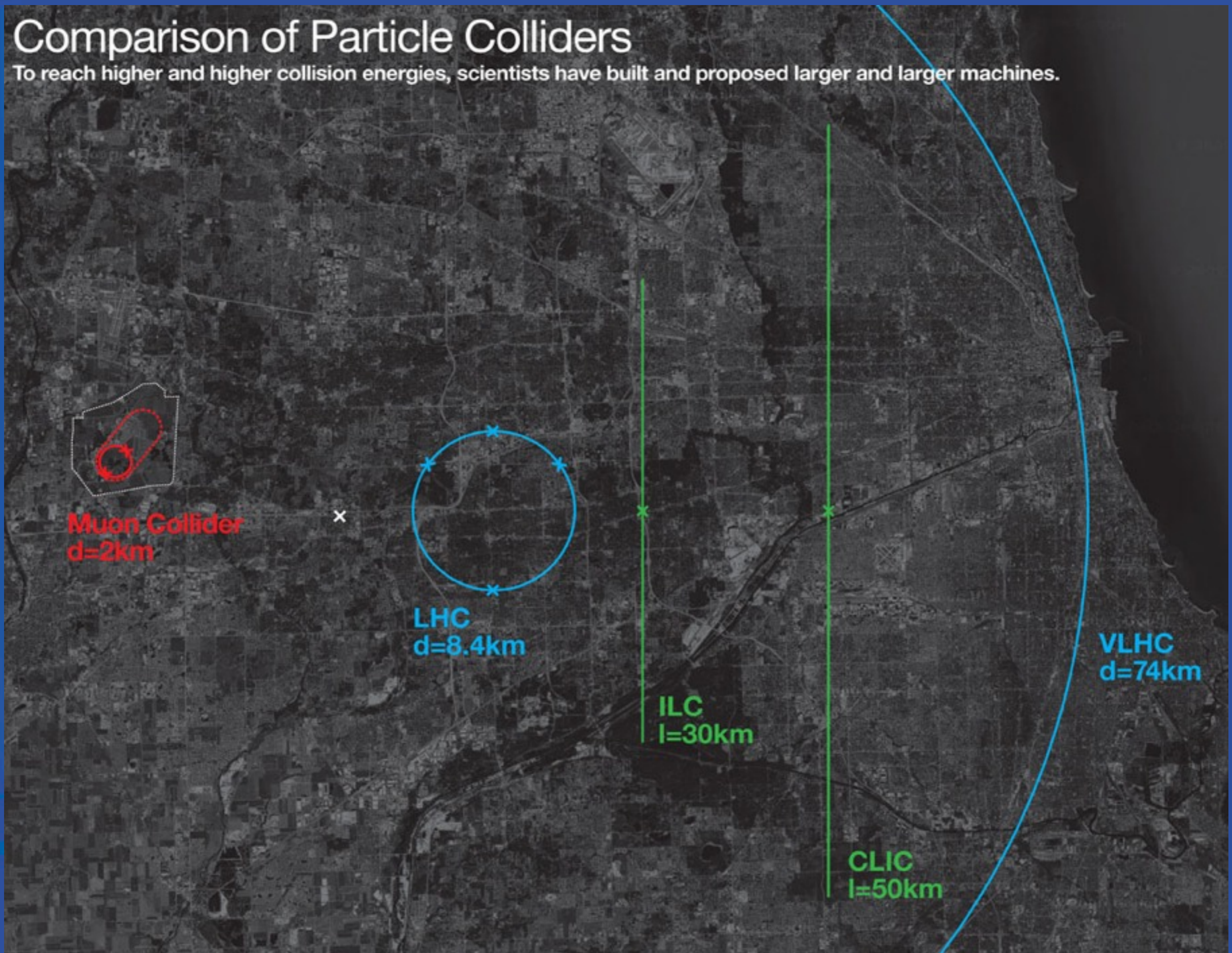
Muon collider functional layout



Color indicates degree of needed R&D (difficulty) and demonstration

Comparison of Particle Colliders

To reach higher and higher collision energies, scientists have built and proposed larger and larger machines.

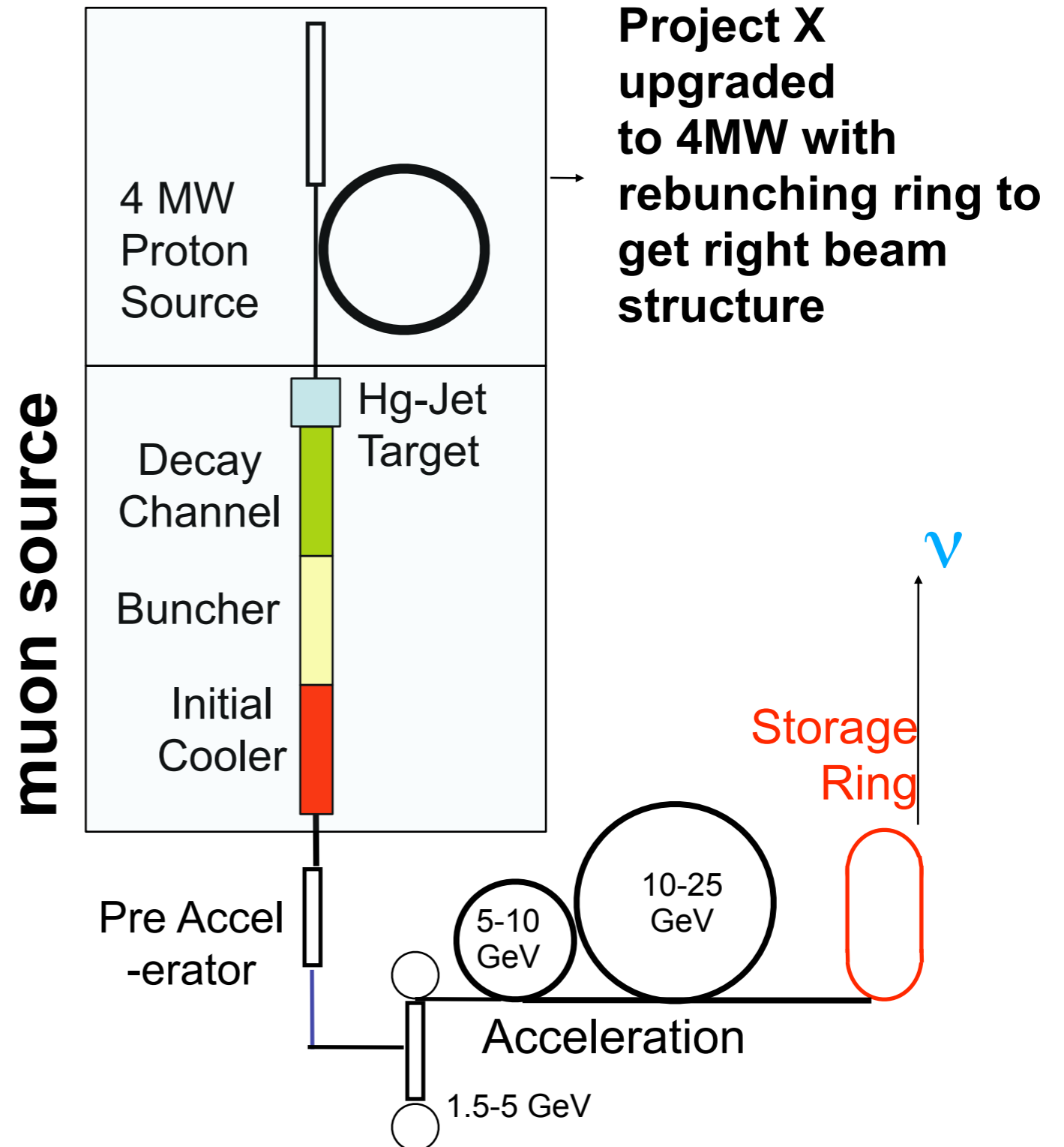




Neutrino Factory Schematic



- Proton Source
 - primary beam on production target
- Target, Capture, and Decay
 - create π ; decay into μ
- Bunching & Phase Rotation
 - reduce ΔE of bunch
- Cooling
 - reduce transverse emittance
- Acceleration
 - $130 \text{ MeV} \rightarrow E_{NF}$
- Storage Ring
 - store for 500 turns; long straight section



Project X in a nutshell:

<http://www.fnal.gov/pub/projectx/>

Highly flexible 3 GeV proton continuous wave linac with a beam power of 2 MW. This can be used for:

- Neutrino Physics: Long Baseline - LBNE (θ_{13} , mass hierarchy, CPV) short baseline - $\sin^2 \theta_W$, anomalous couplings of neutrinos, LSND,
- Kaon Physics: Precision measurements of $K_L \rightarrow \pi^+ \nu \bar{\nu}$ and $K_L \rightarrow \pi^0 \nu \bar{\nu}$. Also precision measurements of K^0 and \bar{K}^0 system, CPT tests.
- Muon Physics: (g-2), electric dipole moments, $\mu \rightarrow e$ conversion, muon cooling leading to Neutrino Factory and maybe a Muon Collider.
- Nuclear Physics: EDM measurements of unstable deformed nuclei.

At Fermilab, our revolution has already started !!!