

Prospects for (part of) Astroparticle Physics

Pierre Auger Observatory air shower radio detection the situation in Germany

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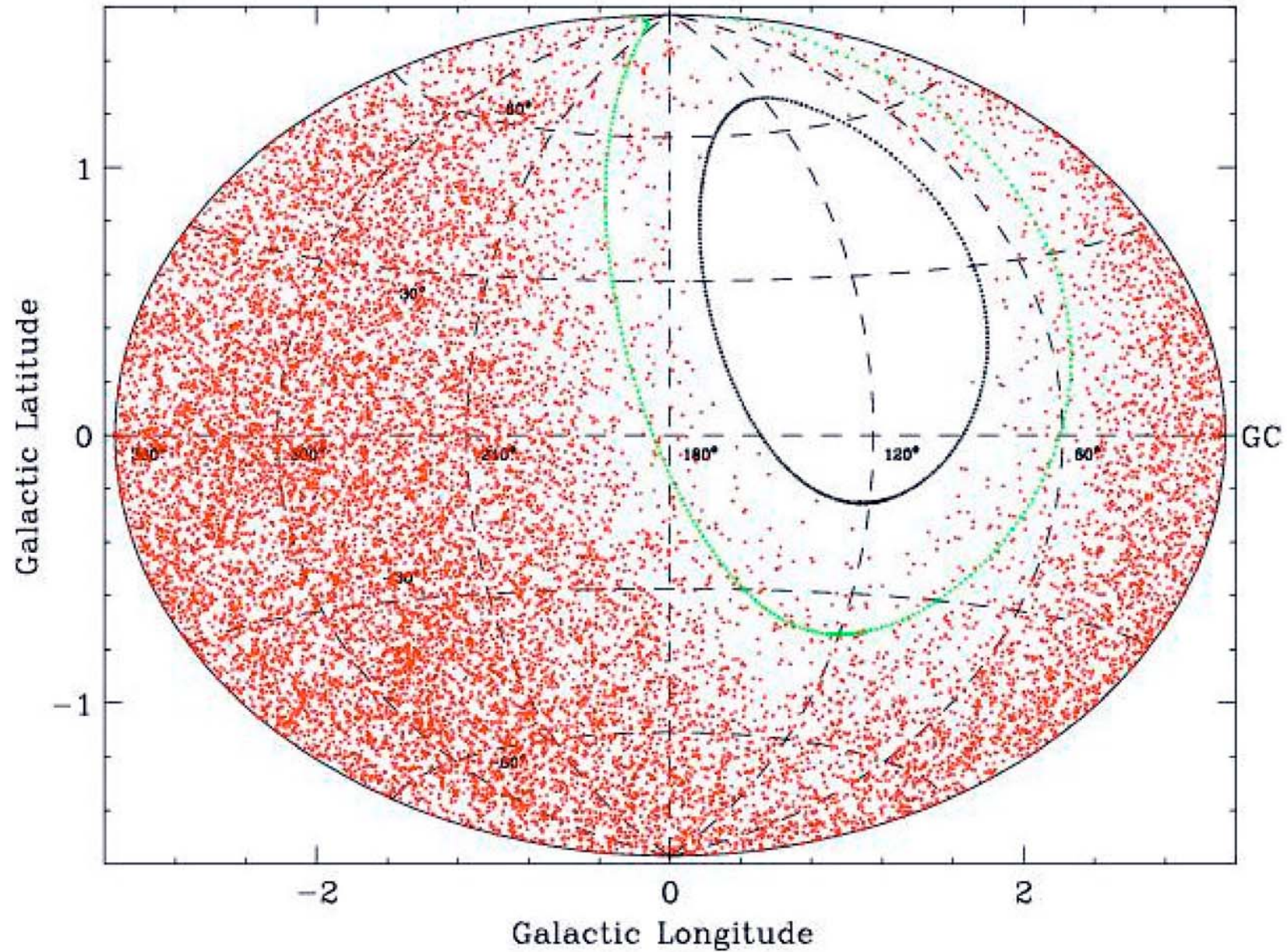
Towards full-sky coverage

- top priority: completion of the Southern Observatory
- Auger North scientific case
 - matter distributions N/S are very different \Rightarrow effect on possible point sources & energy spectra
 - anisotropy searches much better \Rightarrow new physics?
- site: Utah or Colorado
 - decision June 2005
 - one (1) Auger Collaboration running two observatory sites
- observatories be functionally equivalent
 - for seamless data integration
 - options:
 - larger area
 - novel technologies
 - optimized designs



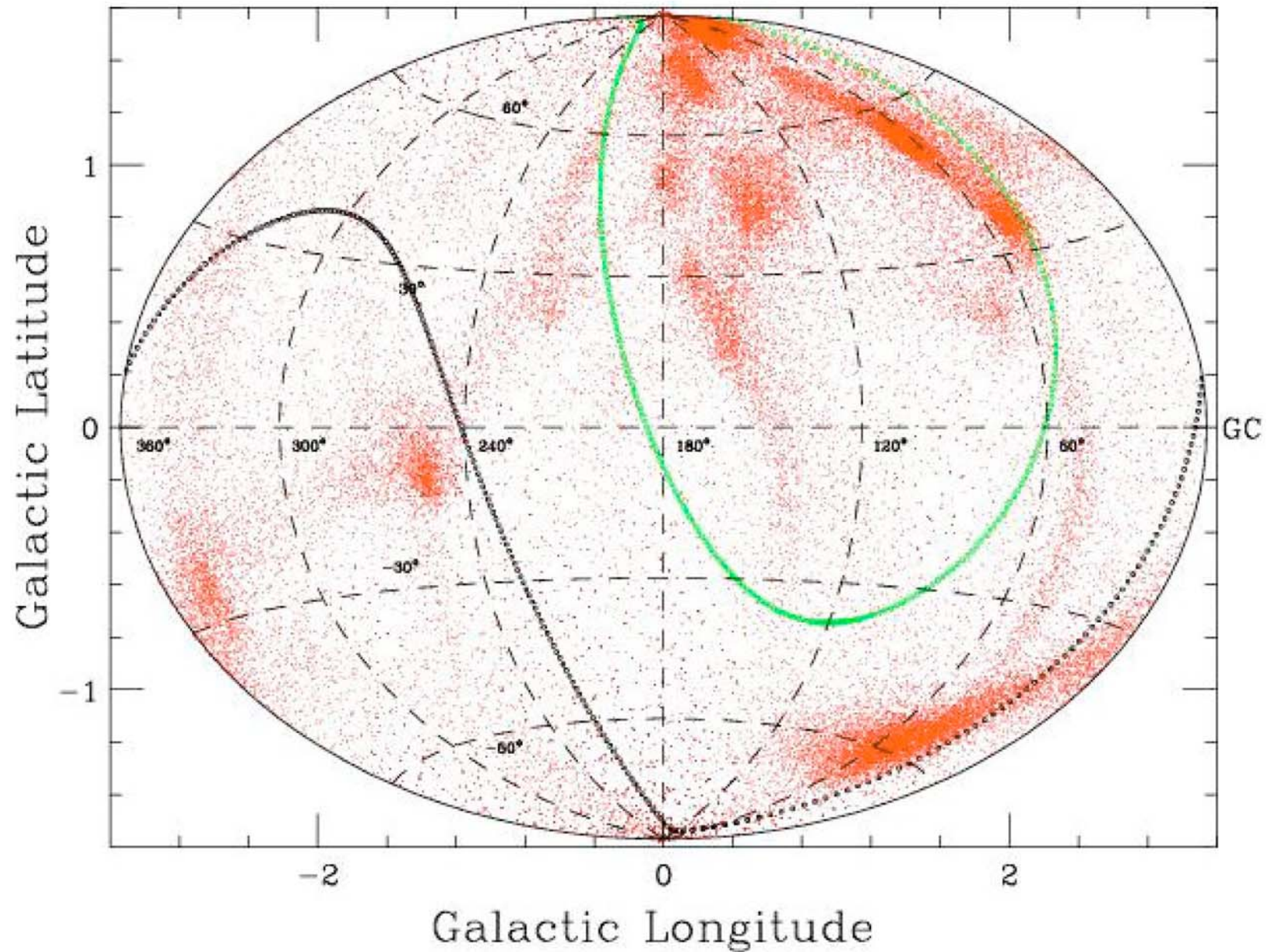
Auger sky map

Distribution of Auger events: 60 deg bound (green), 85 deg bound (black)



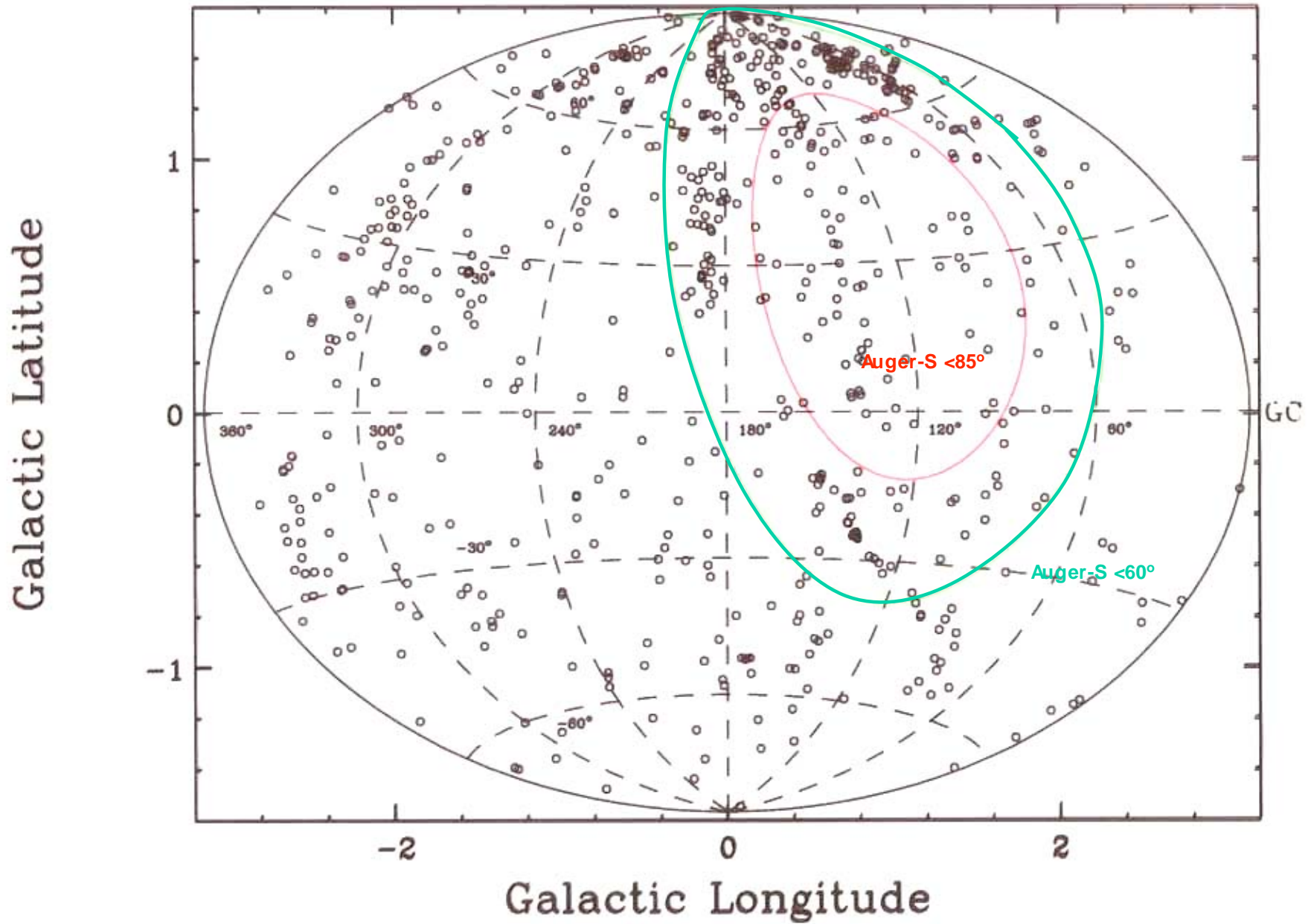
Northern and Southern matter distributions...

Matter distribution 7–21 Mpc. Exclusion zones; north array (black), south array (green)



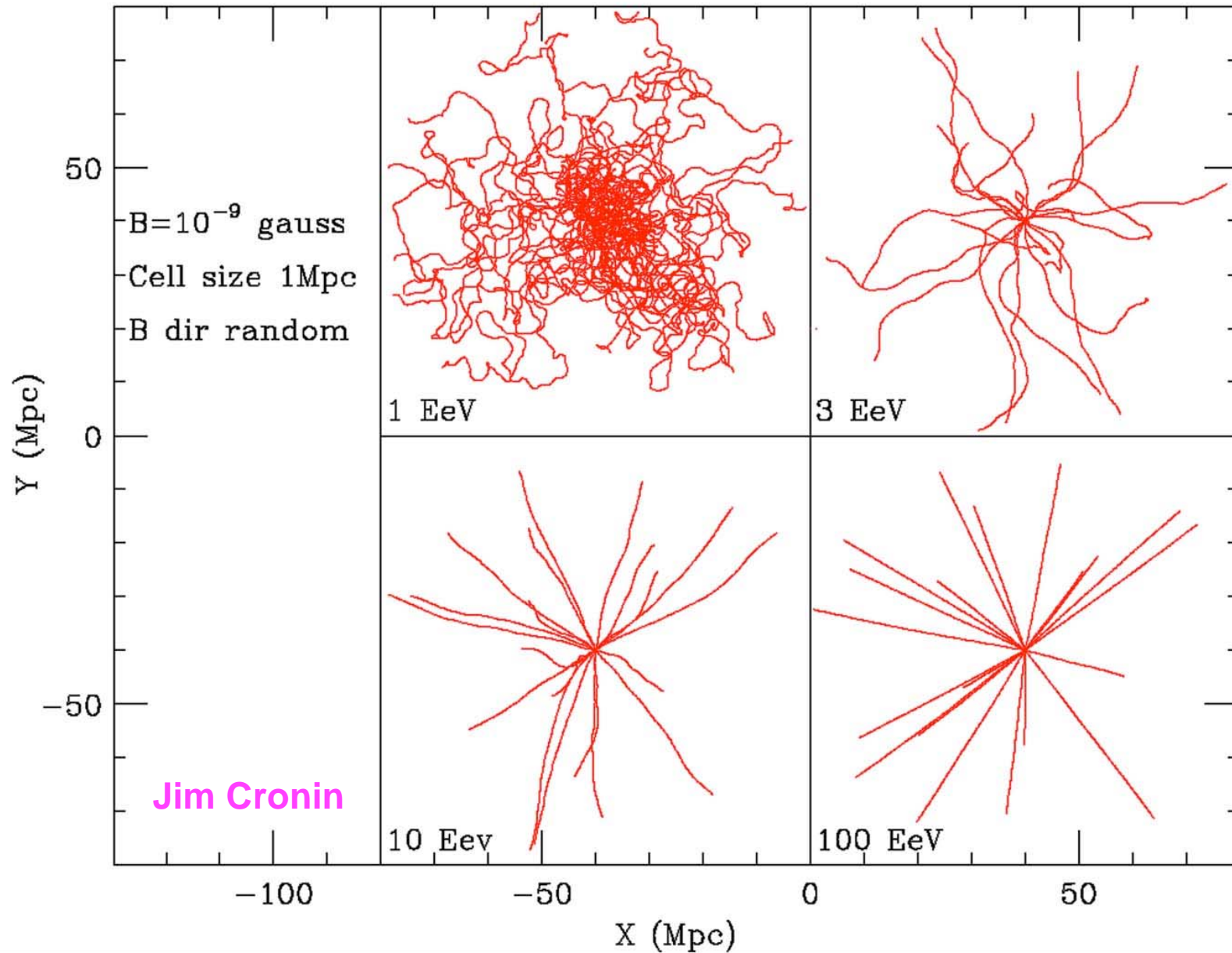
Northern and Southern matter distributions...

Distribution of Radio Galaxies from IPAC Catalog ($0 < Z < 0.02$)

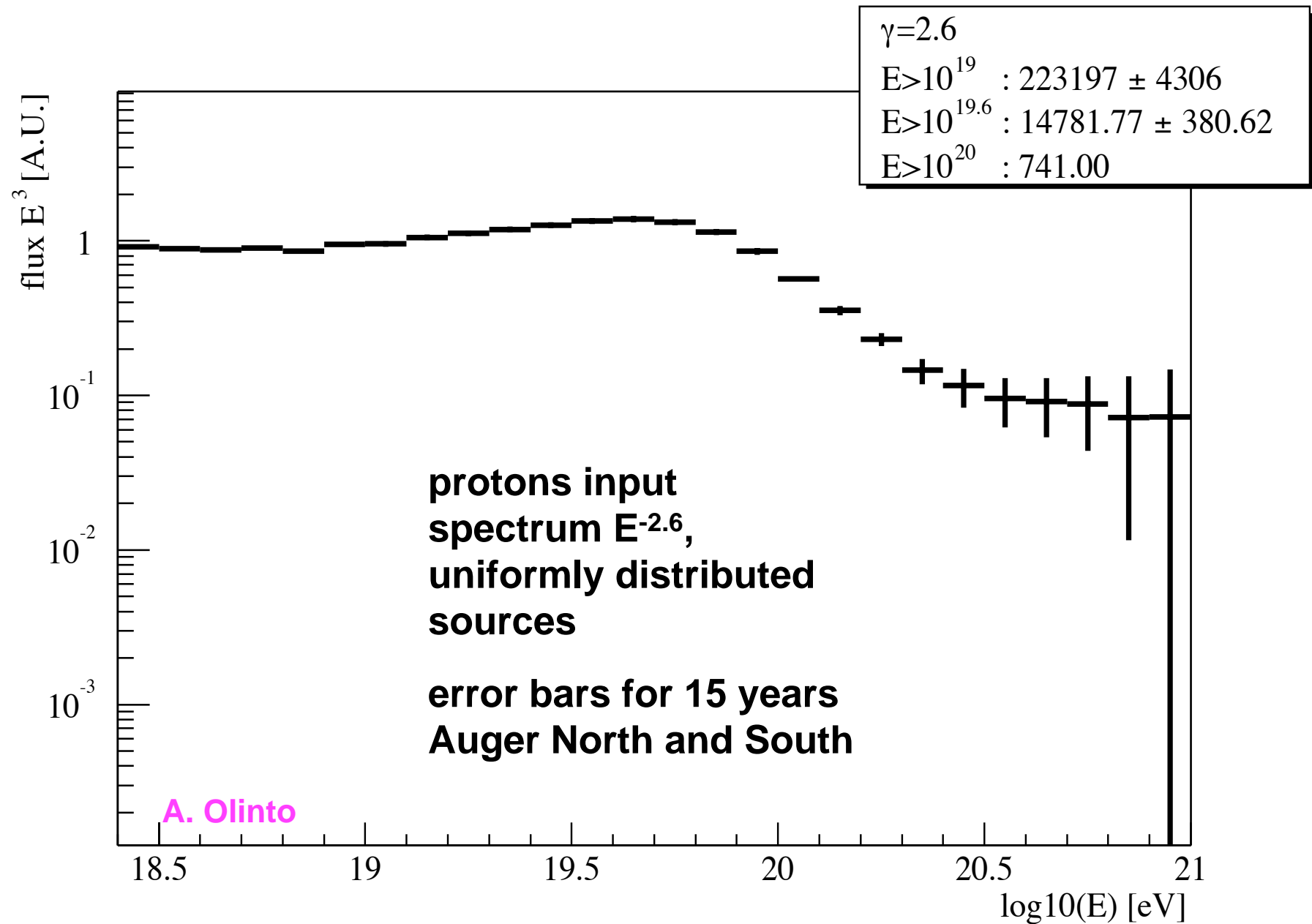


proton astronomy

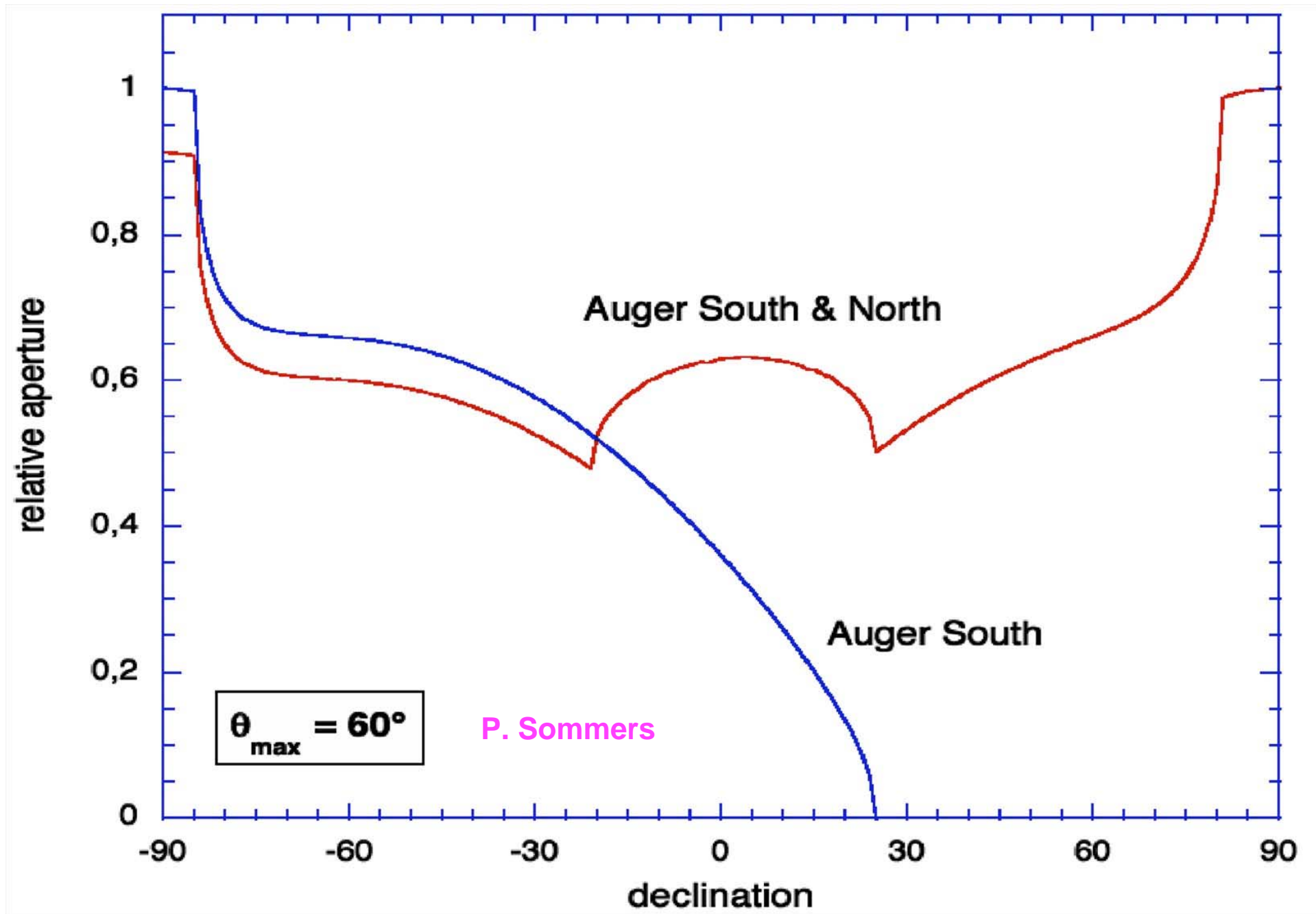
3D trajectories projected on X-Y plane



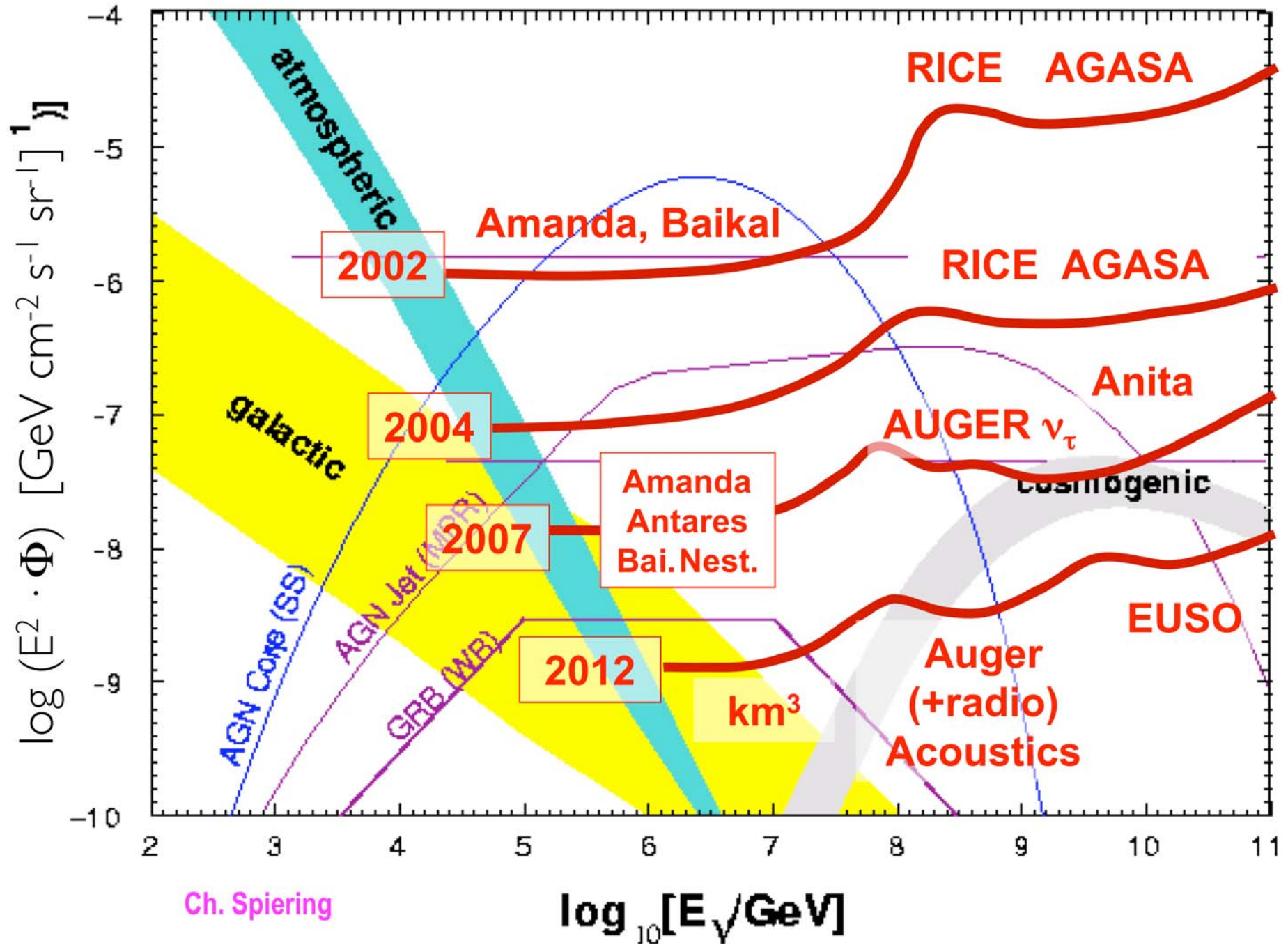
Auger N+S statistical power



relative apertures S+N for anisotropy study



neutrino sensitivity roadmap



Ch. Spiering



Site Comparison

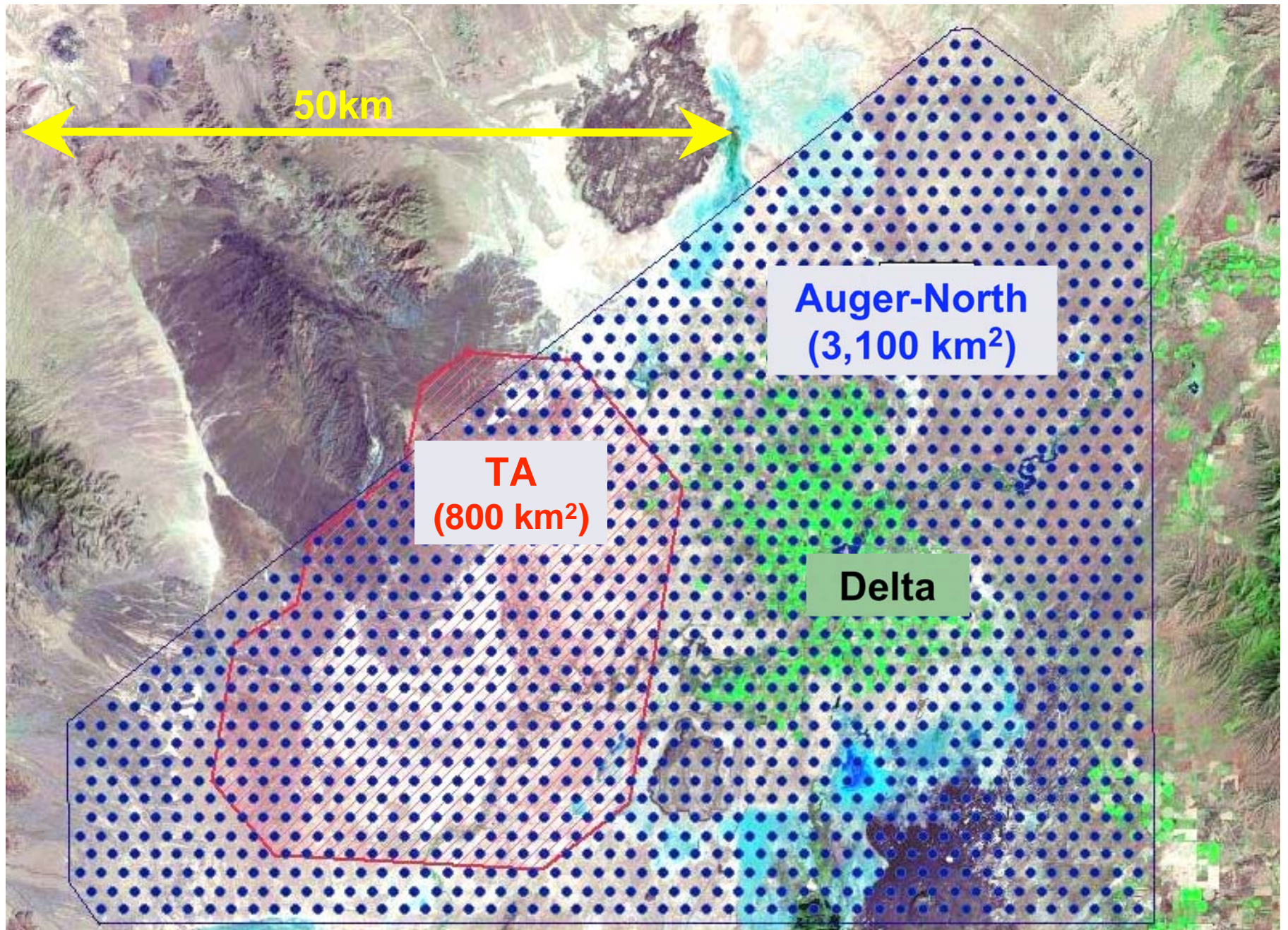
	Southern Site	Northern Sites under Consideration	
Location/Geography			
Country	Argentina	USA	USA
State (Province)	Mendoza	Utah	Colorado
Latitude	35 -35.3 South	39 North	37 45' North
Longitude	68.9 – 69.4 West	112 West	102 45' West
Altitude	1300 m (E) 1500 m (W)	1400-1450m	1330 m (SW) 1220 m (NE)
Maximum Area (confirmed)	5,800 km ²	7,200 km ²	15,000 km ²
Geography	Flat with a gentle slope	Great Basin	High plains, gentle slope
Land Usage	Ranching (goat, cattle), small farming	Ranching, farming	Ranching, farming
Infrastructure			
Nearby Local Town	Malargue	Delta / Fillmore	Lamar
Population	18,000	3,000 / 2,000	9,000
Major City/ Airport	Mendoza, San Rafael	Salt Lake City	Denver
Distance to Site	420 km, 200 km	210 km	320 km
Weather			
Temperature (Summer)	January 20°C	July mean 24°C	July mean 25°C
(Winter)	July 4°C	January mean -3°C	January mean -2°C
"Cloudiness"		≈ 10%	≈ 7.5%
Irradiance	≈ 380 W/m ²	≈ 253 W/m ²	≈ 250 W/m ²
Land Ownership			
Private	≈100 owners (95%)	≈200 owners (20%)	≈280 owners (95%)
Federal-Owned	0 %	BLM (75%)	negligible
State-Owned	≈ 5%	≈ 5%	≈ 5%

Site Comparison

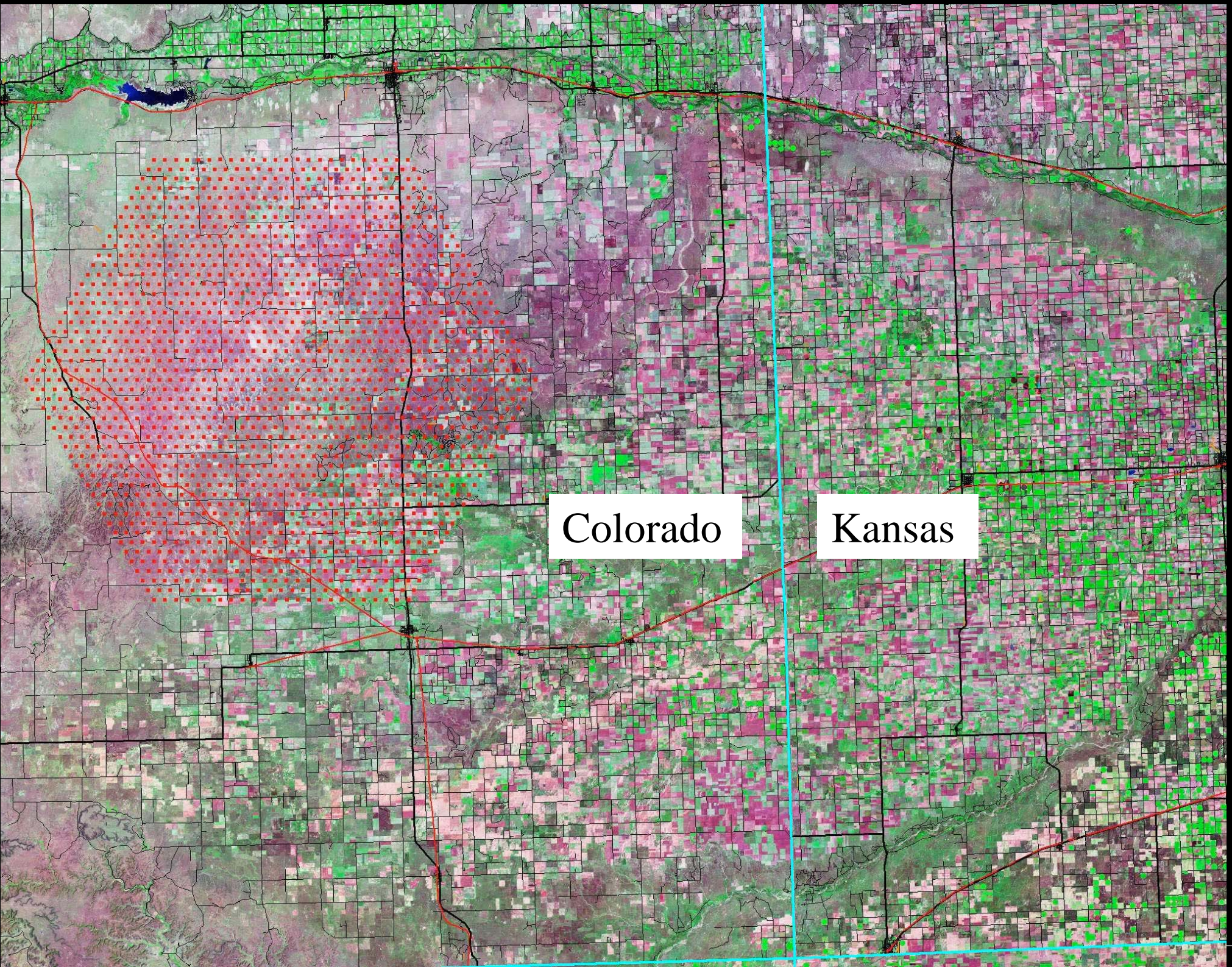
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Utah Site



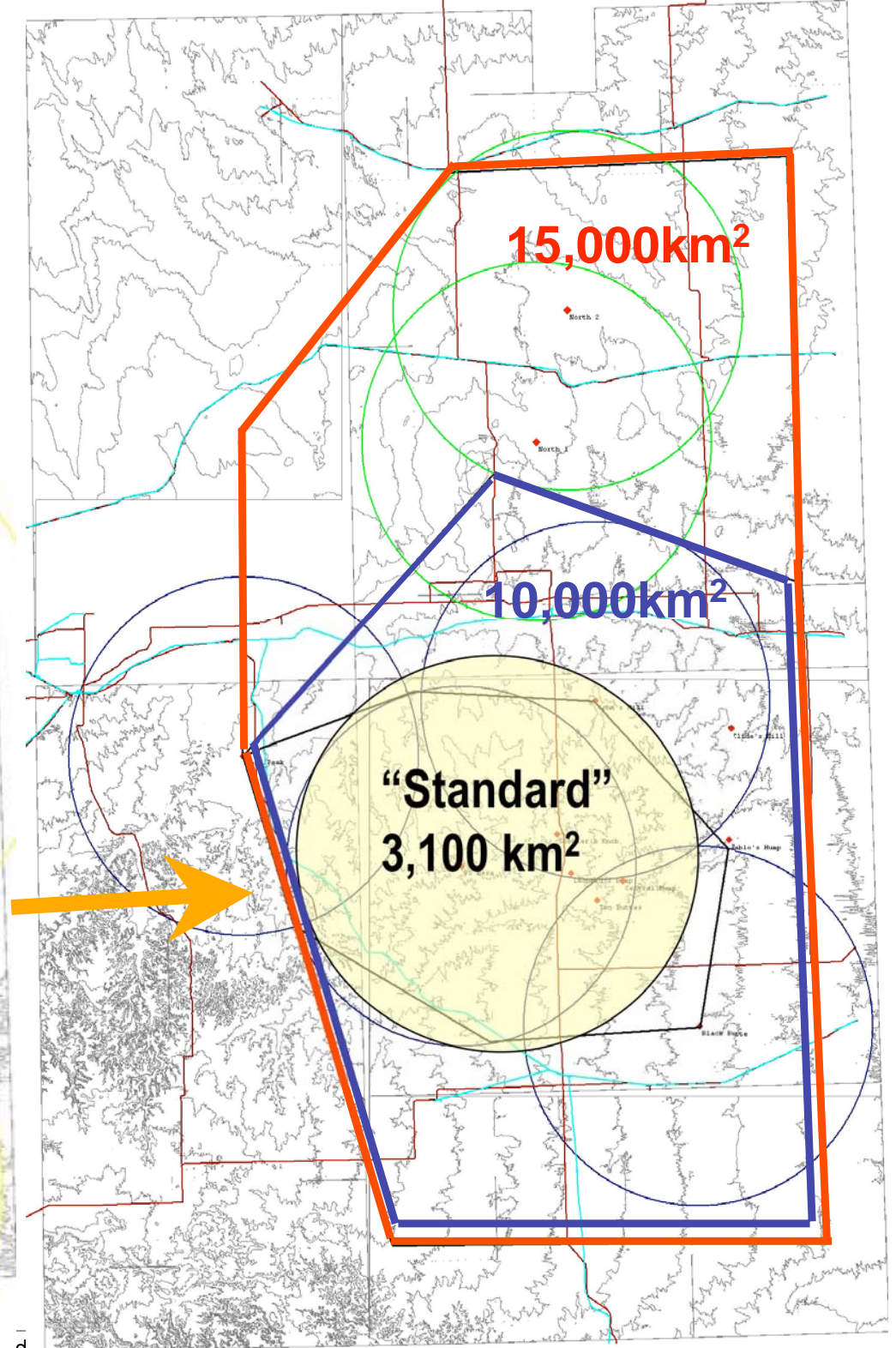
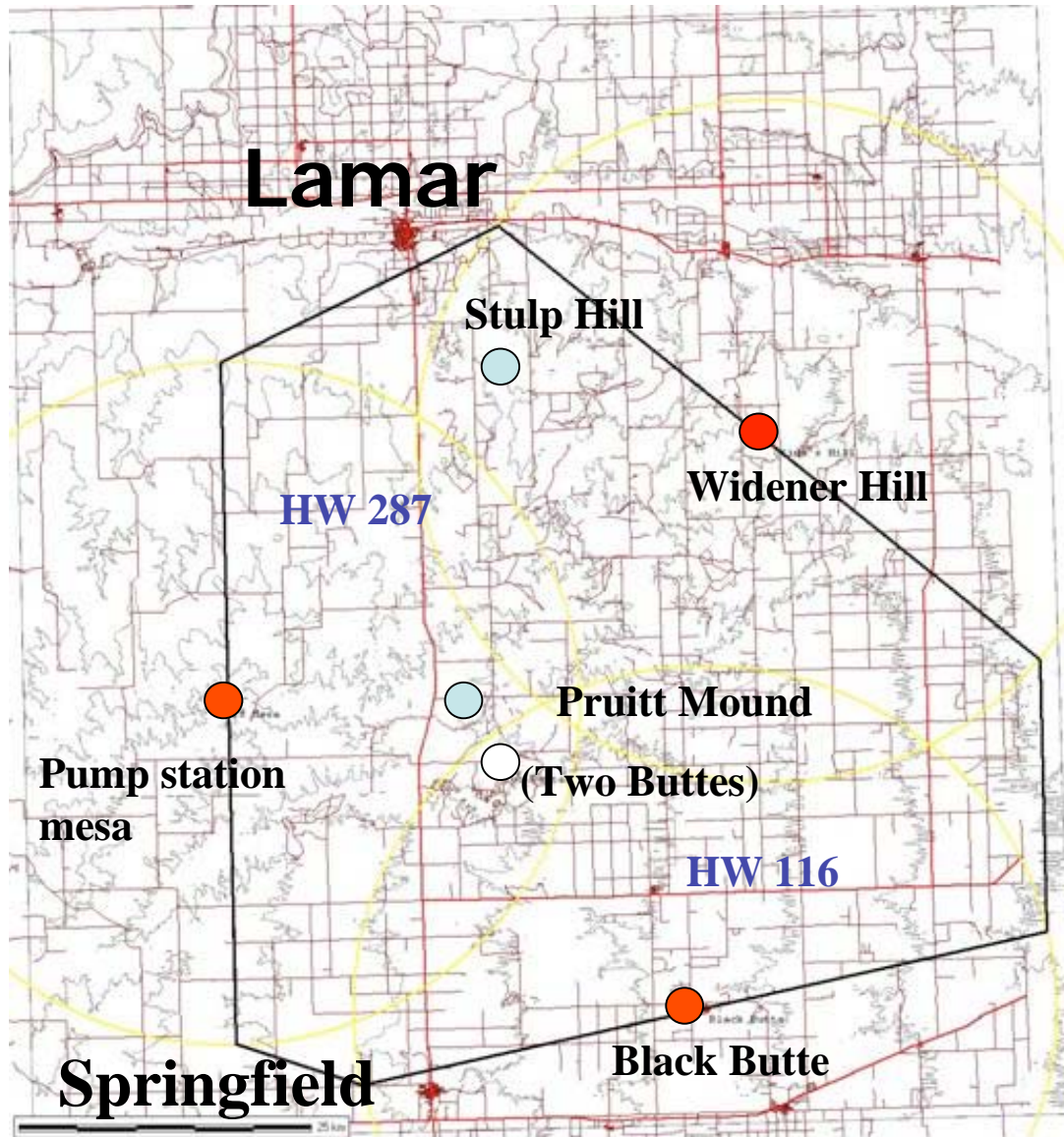




Colorado

Kansas

Colorado Site





air shower radio detection

history

LOFAR

LOPES & KASCADE

it works !

air shower radio detection: historic plots

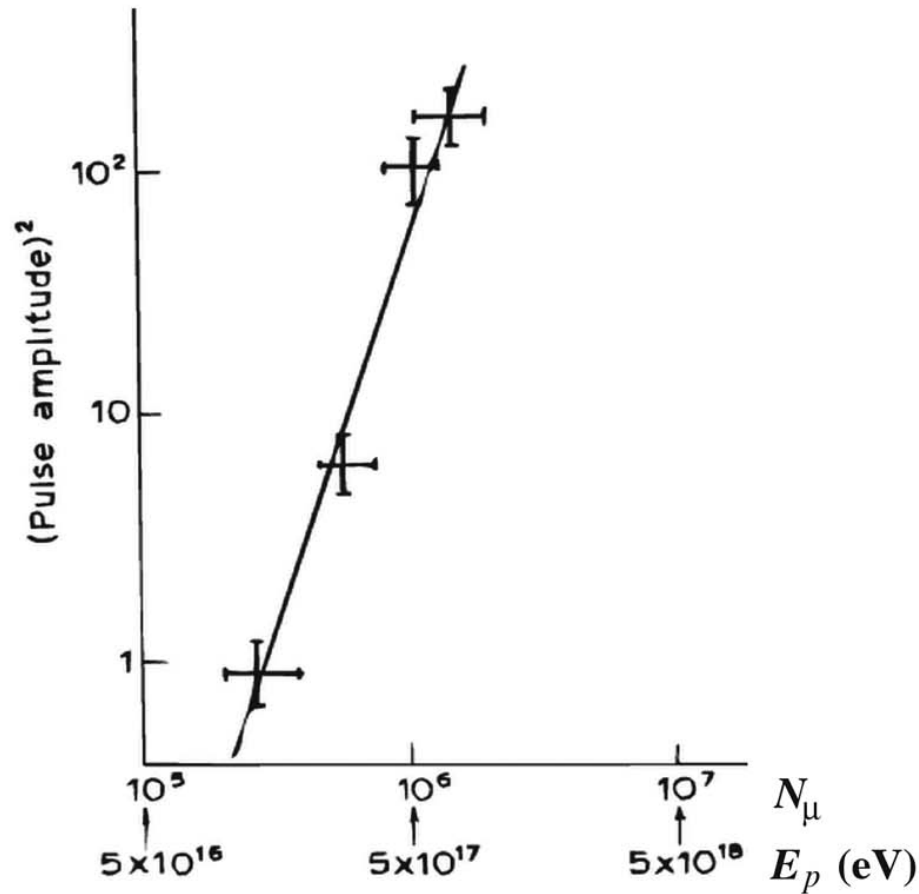
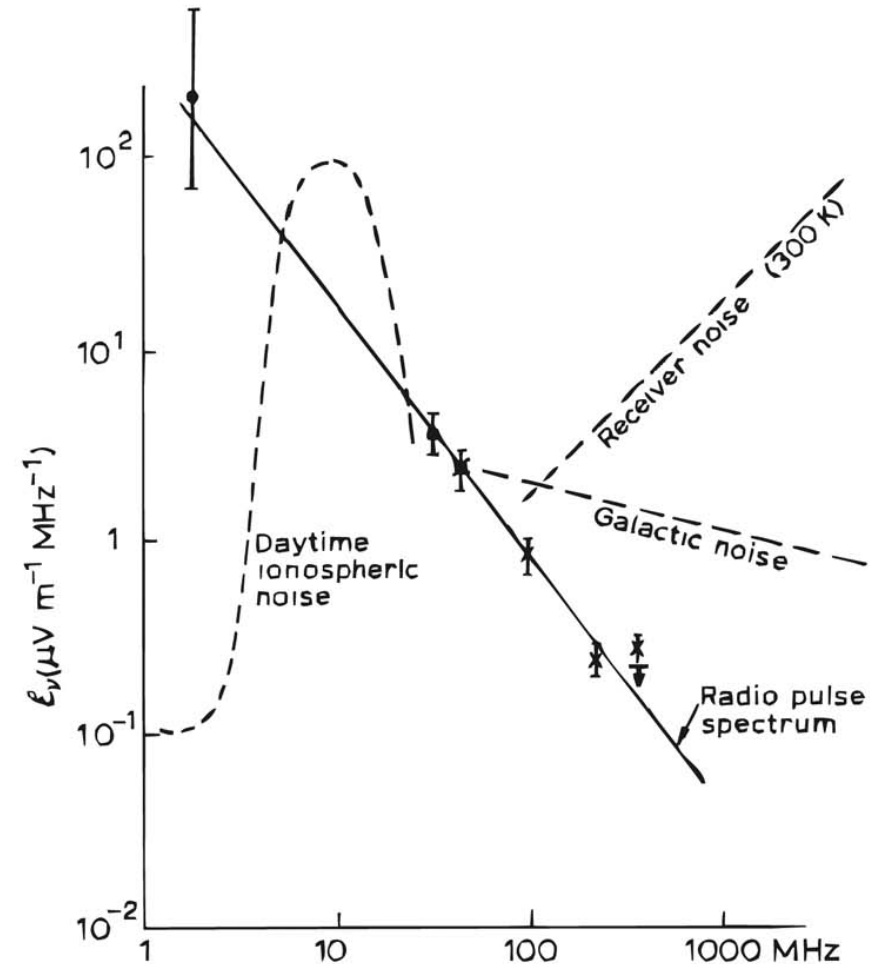


Figure 6. The dependence of EAS radio flux on the primary particle energy as measured by Vernov et al. (1968) following roughly a E_p^2 power-law. Some earlier papers found somewhat flatter dependencies.

taken from Falcke et al astro-ph/0409229

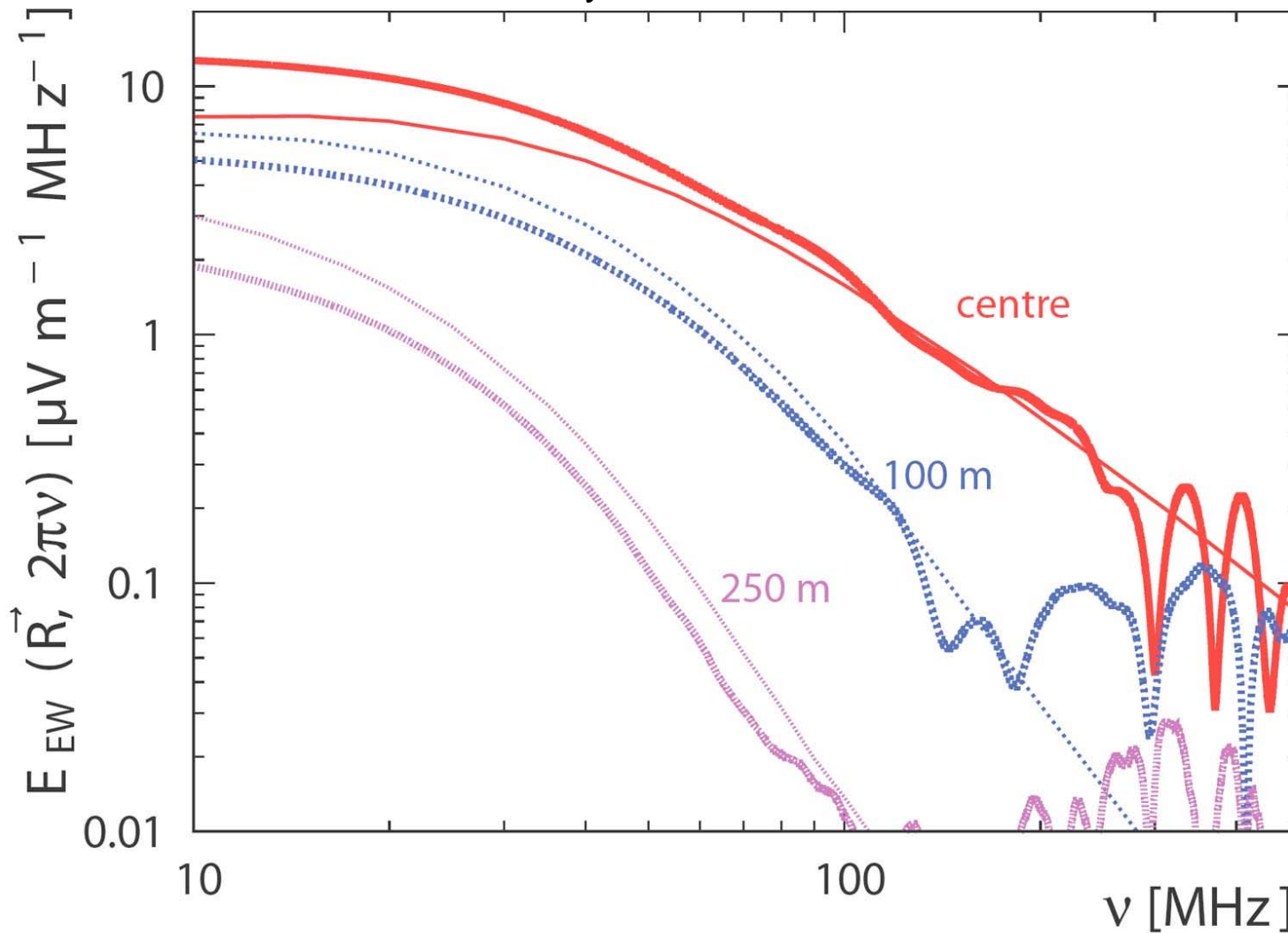


- ⊕ Haverah Park points
- ⊗ Jodrell Bank points
- ⊥ Upper limit

Figure 7. A tentative radio pulse spectrum for 2 MHz to 520 MHz. The data are not simultaneous. From Allan (1971) and Spencer (1969).

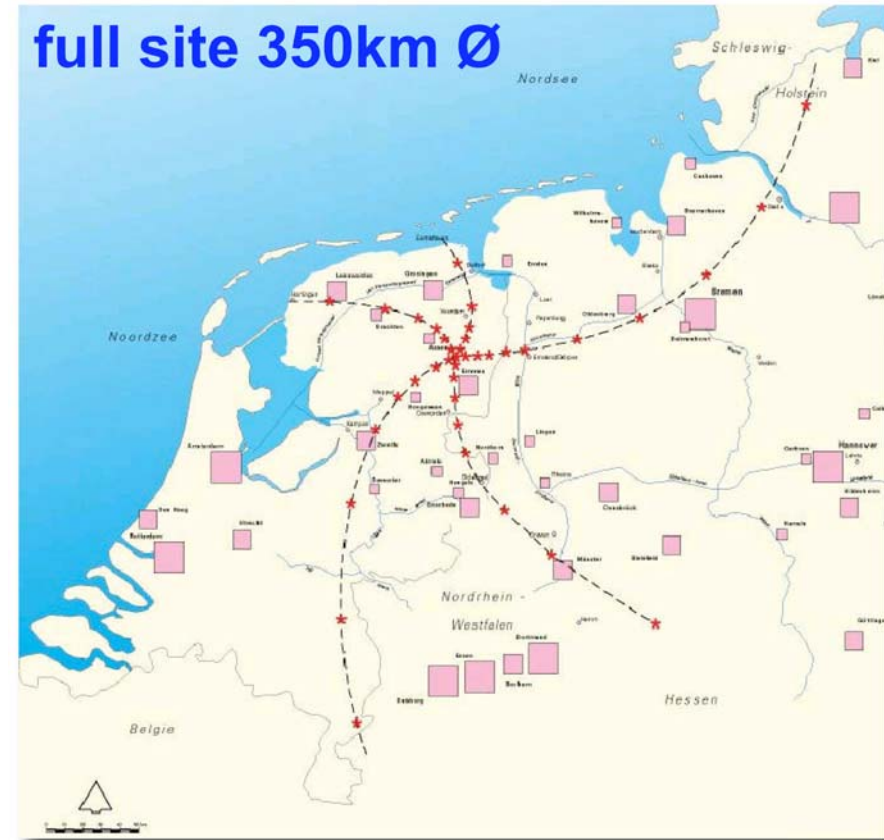
Radio detection of air showers

spectra from 10^{17} eV vertical shower for 3 distances to North
thin: analytic shifted x0.5; thick: MC



the Geosynchrotron paper:
Huege T., Falcke H. 2003, Astronomy & Astrophysics, Vol. 412, p. 19-34

LOFAR: Low Frequency Array



The five key areas of science to be addressed by LOFAR are:

1. **The High Redshift Universe:** the study of the most distant radio galaxies and quasars
2. **The Epoch of Reionization:** detection of the global signature, and mapping of structures
3. **Mapping Galactic Cosmic Rays:** to map the 3D distribution of the Galactic cosmic ray electron gas
4. **The Bursting and Transient Universe:** to detect short lived transient events – bursts from Jupiter-like planets, merging and interacting compact objects.
5. **Solar-Terrestrial Relationships:** to detect Coronal Mass Ejections possibly in combination with a Solar Radar, and to study the Earth's Ionosphere.

Radio detection of air showers



LOPES collaboration:

- KASCADE-Grande
- U Nijmegen, NL
- MPIfR Bonn
- Astron, NL
- FZK-IPE

2003:

10 antennas inside KASCADE

frequency band 40-80 MHz

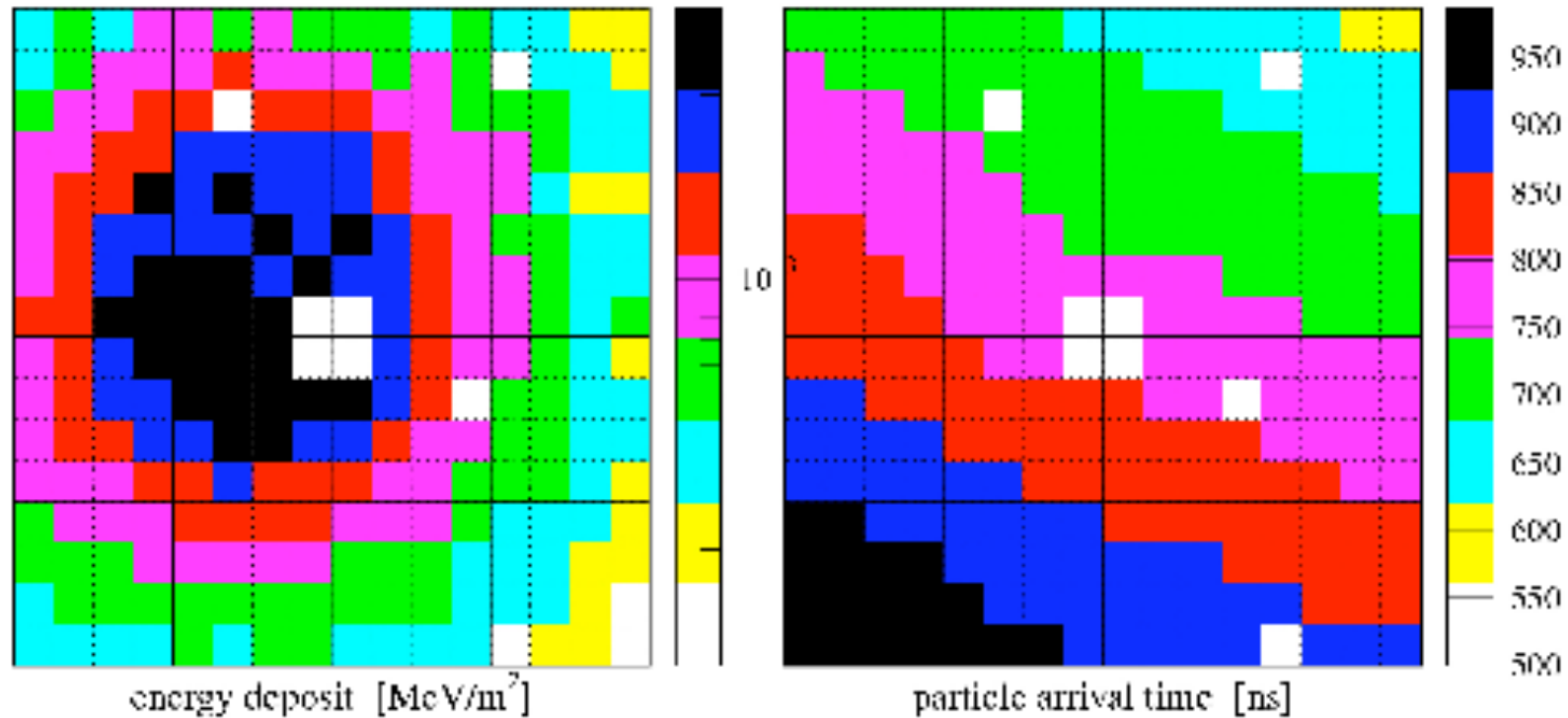
trigger: >10/16 cluster of KASCADE, $E_0 > 10^{16}$ eV

the LOPES project



first promising event: KASCADE view

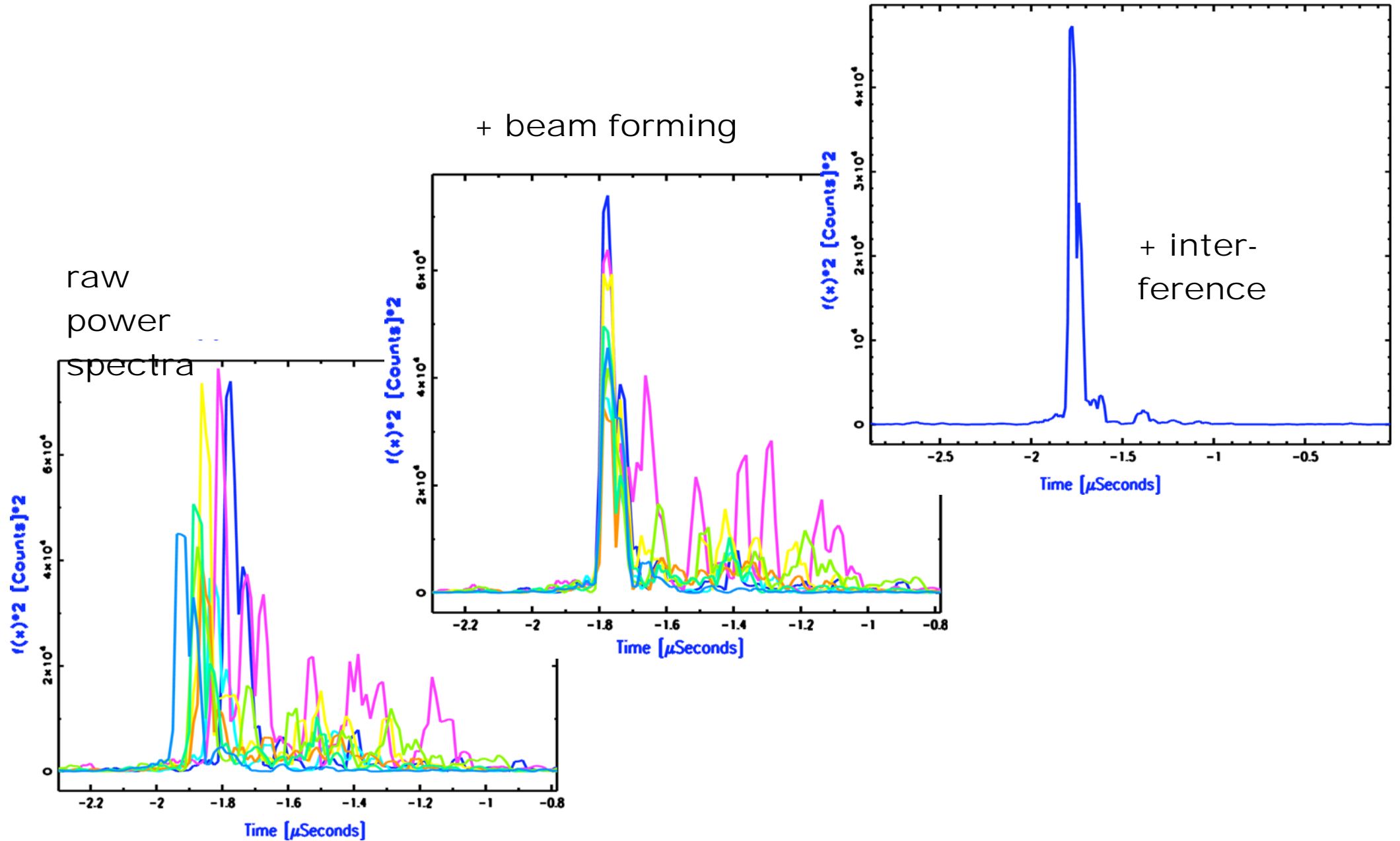
e/ γ -detector, run 004702 event 0294563



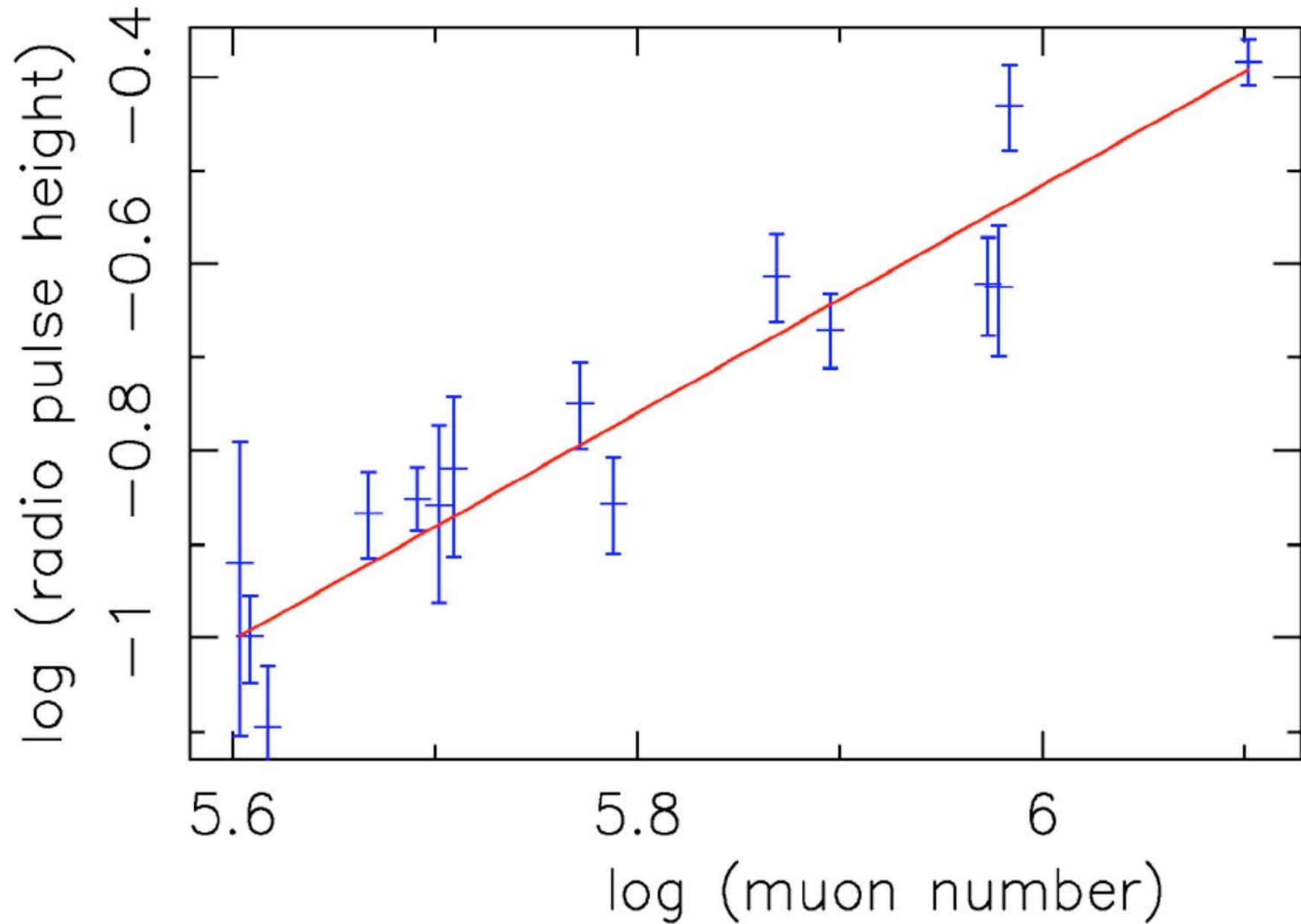
event in KASCADE

- energy $\approx 10^{17}$ eV
- EAS core inside antennas
- $\Theta = 25.5^\circ$, $\Phi = 42.5^\circ$
- 8 antennas were working and show signals
- signal is coherent

first promising event: LOPES view



correlation with KASCADE measurements





Pierre Auger Observatory

- the apparatus works extremely well, beyond expectations
- 636 tanks and 6+6 telescopes in data acquisition + more in the field
- Feb 2005: 85% of AGASA exposure (1700 km² sr yr)
- data look very interesting indeed
- data set for first analysis:
 - present run until March 2005
 - statistical power about "1 AGASA"
 - spectral shape by SD, energy scale by FD (both are work in progress)
 - release for the ICRC in August 2005
- continuing deployment
 - tanks until mid-2006
 - telescopes: 18 within few months, 4th building done by Mendoza Province
- further strengthening of the collaboration towards full-sky coverage
 - German Universities, Netherlands, Korea, ...

related work to be done

- **Preparation for Auger North**
 - surface detectors
 - climate impact
 - Liners
 - PMT enclosures
 - communications
 - wireless broadband
 - GPS receivers
 - frontend electronics redesign
 - addition of Cherenkov and radio detection

- **upgrade Southern site**
 - low energy extension
 - upward-looking fluorescence telescopes
 - denser partial array
 - muon counters

both are part of the same endeavour
strengthen the collaboration
additional funds

Astroparticle Physics in Germany

community

structures

Helmholtz Virtual Institutes

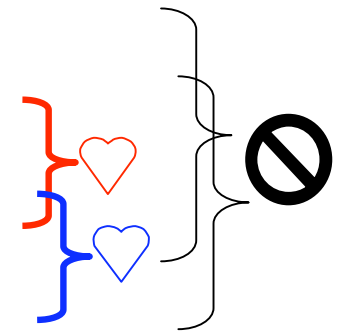
KAT

definition of an astroparticle physicist (m/f)

I think to be an astroparticle physicist, **so I am**
an astroparticle physicist

Komitee für Astroteilchenphysik

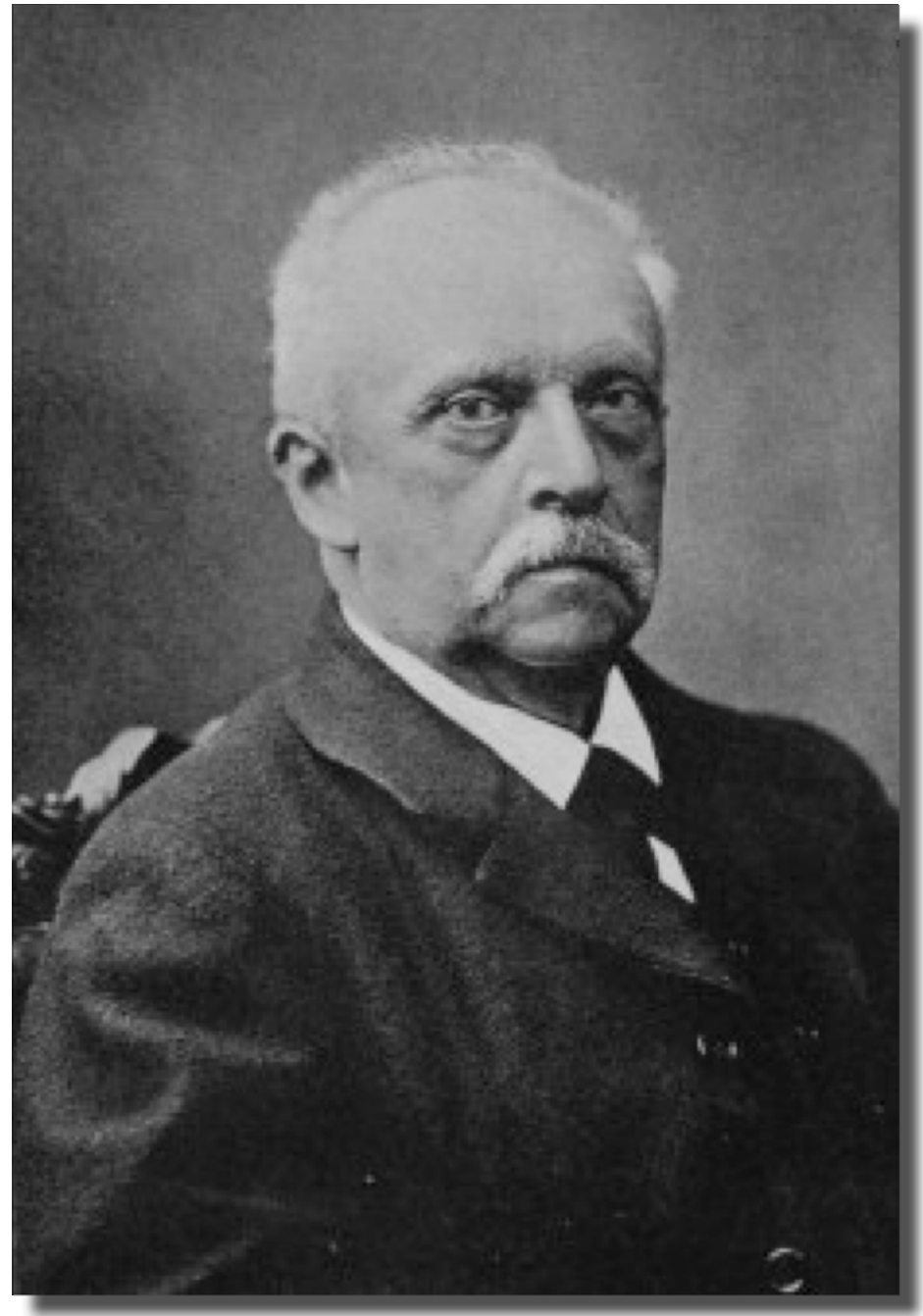
- one member for each research field, elected by the community
- ex-officio members representing other bodies:
 - Komitee für Hadronen und Kerne, KhuK
 - Komitee für Elementarteilchenphysik, KET
 - Council of German Astronomers
 - APPEC
 - Deutsche Forschungsgemeinschaft DFG €
 - Max Planck-Gesellschaft MPG €€€€
 - Verbundforschung Astroteilchenphysik €€€€€€€ → Universities!
 - Helmholtz Association HGF €€€€€€€€€€



KAT election January 2004

Anton	Gisela	HE neutrinos	Erlangen
Blümer	Johannes	cosmic rays	Karlsruhe
Danzmann	Karsten	gravitational waves	Hannover
Drexlin	Guido	neutrino properties	Karlsruhe
Hofmann	Werner	gammy astronomy	Heidelberg
Jochum	Josef	dark matter	Tübingen
Raffelt	Georg	theory	Munich
Rolfs	Claus	nuclear astrophysics	Bochum
von Feilitzsch	Franz	LE neutrinos	Munich

Anton	Gisela	Komitee für Hadronen- und Kernphysik (KHuK)
Bartelmann	Matthias	Rat Deutscher Sternwarten RDS
Berghöfer	Thomas	Verbundforschung Astroteilchenphysik
Koepcke	Rainer	BMBF, Referat 411
Koepke	Lutz	Komitee für Elementarteilchenphysik KET
Kleinschmidt	Michael	Deutsche Forschungsgemeinschaft DFG
Krückeberg	Stefan	
Völk	Heinz	APPEC
Schmidt	Sebastian	Helmholtz-Gemeinschaft HGF
Echinger	Stefan	Max Planck-Gesellschaft MPG





Ernest Rutherford 1871 - 1937

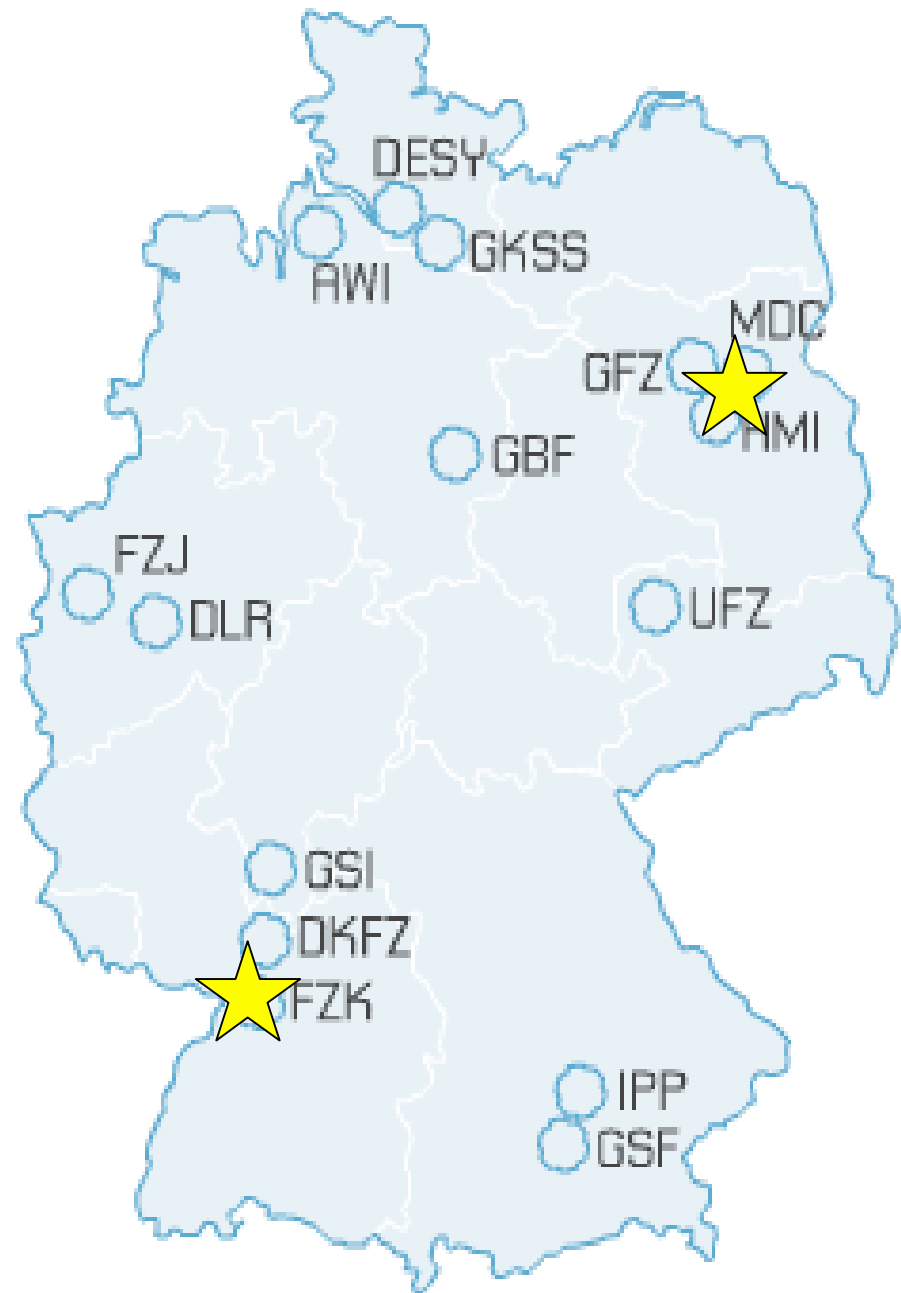


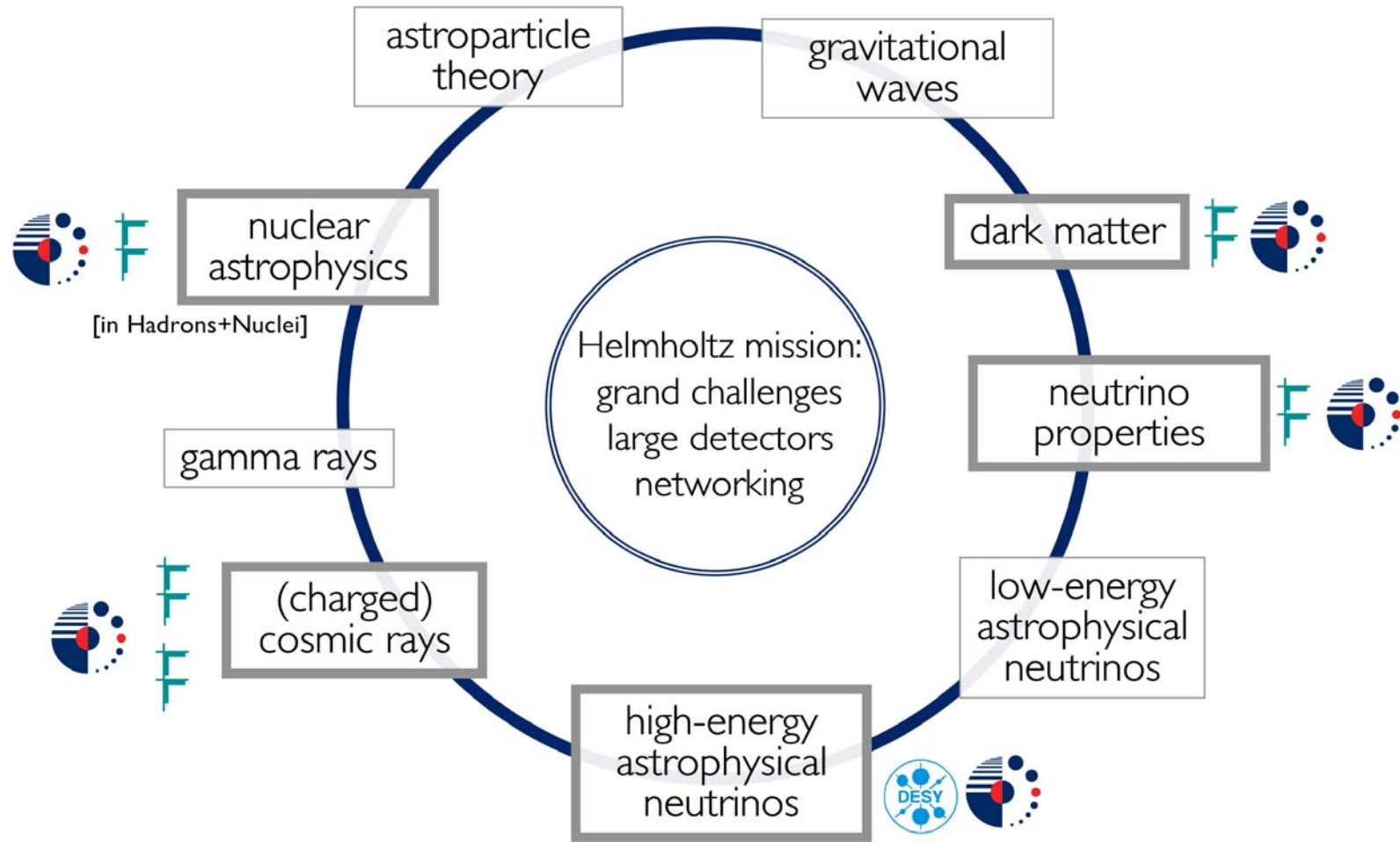
Hermann von Helmholtz 1821-94

Helmholtz Research Centres (HGF)

Six Research fields

- Energy,
- Earth and Environment,
- Health,
- Key Technologies,
- Transport and Space
- **Structure of Matter**
 - Elementary Particles
 - Hadrons and Nuclei
 - Photons, Ions and Neutrons
 - Condensed Matter
 - **Astroparticles**
 - DESY Zeuthen
 - BAIKAL
 - AMANDA, IceCube
 - FZK
 - KASCADE-Grande
 - Pierre Auger
 - CORSIKA/CONEX
 - KATRIN
 - EDELWEISS





Helmholtz timelines

	1998	1999	2000	2001	Evaluation periods			BMBF			2009	BMBF			2014	2015
					2002	2003	2004	HGF				HGF				
KASCADE-Grande																
Pierre Auger																
BAIKAL																
AMANDA																
IceCube																
KATRIN																
Edelweiss																

Helmholtz Virtual Institutes

funded by the Helmholtz President's Networking and Innovation
fund with 240k€ each



VIDMAN

Eine gemeinsame Einrichtung des Helmholtz-Zentrums Forschungszentrum Karlsruhe mit den Universitäten Bonn, Darmstadt, Karlsruhe, Mainz, TUM München, Tübingen sowie dem Max Planck-Institut MPK Heidelberg

Virtuelles Institut für Dunkle Materie und Neutrinophysik

1 Helmholtz Centre, 7 Universities, 2 Max Planck, 100 people



VIHKOS

Eine gemeinsame Einrichtung der Helmholtz-Zentren Karlsruhe und DESY mit den Universitäten Aachen, HU Berlin, Bochum, Erlangen-Nürnberg, Dortmund, Frankfurt, Hamburg, LSW Heidelberg, Karlsruhe, Mainz, Siegen, Wuppertal, Würzburg sowie den Max Planck-Instituten MPIfR Bonn, MPK Heidelberg und MPI München

Virtuelles Institut für Hochenergiestrahlungen aus dem Kosmos

2 Helmholtz Centres, 12 Universities, 3 Max Planck, 180 people

concluding remarks

- astroparticle physics is growing fast, still continuing
- money follows people with long delays
- some self-organization is desirable
 - bottom-up basis like KAT works ok in Germany
 - European: APPEC
 - EU FP6, FP7, ILIAS, ...?
 - World: PANAGIC
- UK should (continue to) play a highly visible role in the Pierre Auger project
- inter-disciplinarity is extreme in astroparticle physics (by definition)
- watch at the whole field, opportunities for 1.5-2 decades

