

DARK MATTER IN EARLY LHC DATA

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West Coast ATLAS Forum
24 March 2010

OUTLINE

- Long-Lived Charged Particles
 - SuperWIMP DM
 - Feng, Rajaraman, Takayama (2003)
- Delayed Photons/Diphotons
 - Light Gravitino DM
 - Lee, Feng, Kamionkowski (in preparation)
- Exotic 4th Generation Quarks
 - WIMPless DM
 - Alwall, Feng, Kumar, Su (1002.3366 hep-ph)

DARK MATTER

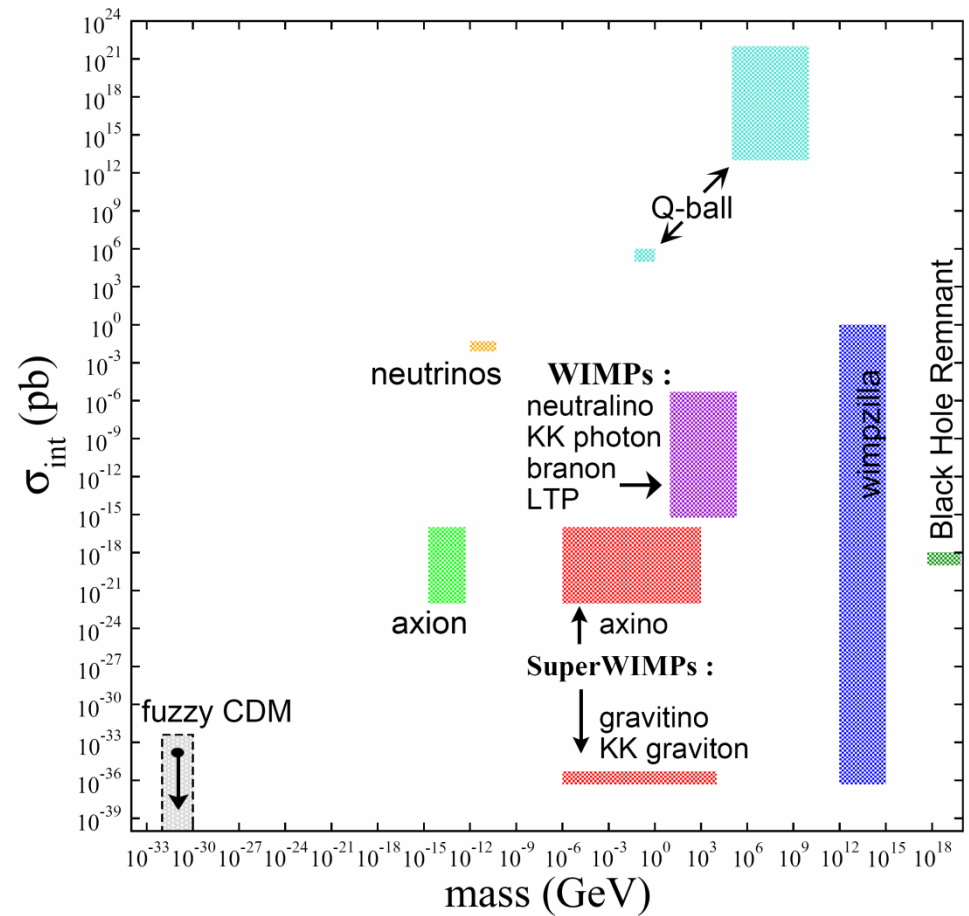
As motivation for new collider physics:

- Not only an aesthetic problem
 - Unlike cosmological constant, gauge hierarchy problems
- Indications it is linked to weak-scale physics
 - Unlike neutrino mass, baryogenesis
- Interesting
 - Highly interdisciplinary, inner space-outer space connections

DARK MATTER CANDIDATES

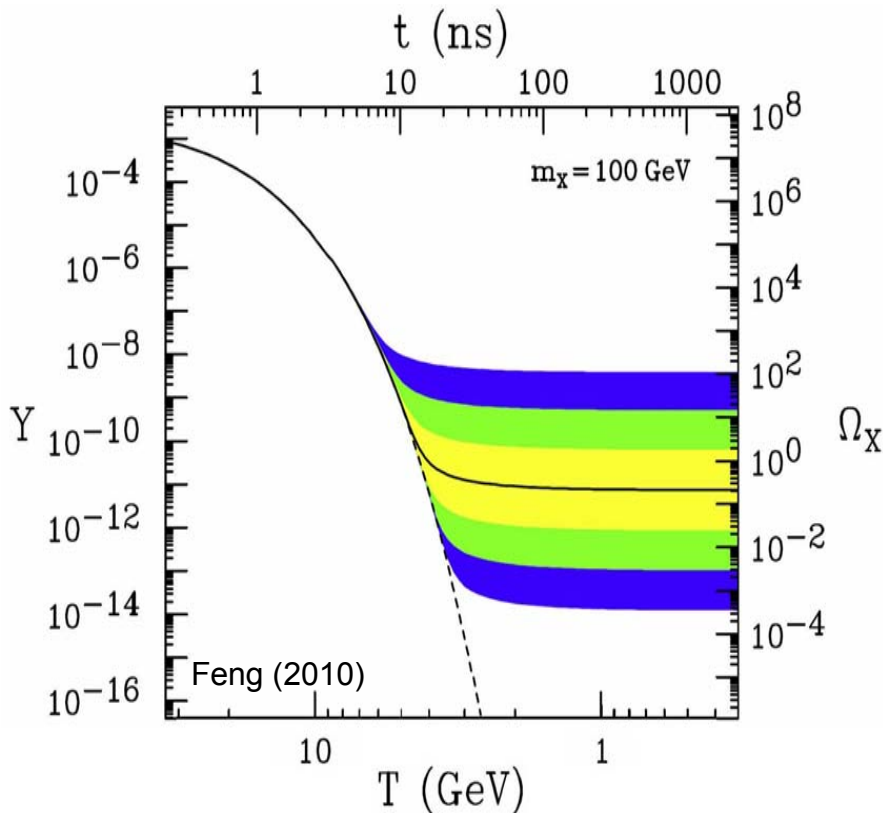
- The observational constraints are no match for the creativity of theorists
- It could be axions, WIMPzillas, etc., in which case the LHC will have very little to say

“Dark Matter Particle Candidates and Methods of Detection,” 1003.0904, to appear in ARAA



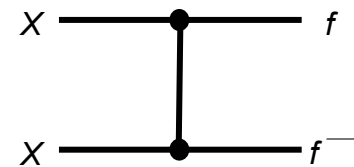
HEPAP/AAAC DMSAG Subpanel (2007)

DM @ LHC: RELIC DENSITY



- The resulting relic density is

$$\Omega_X \propto \frac{1}{\langle \sigma v \rangle} \sim \frac{m_X^2}{g_X^4}$$

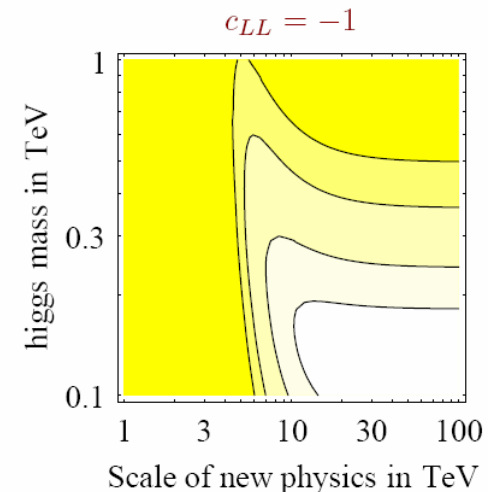
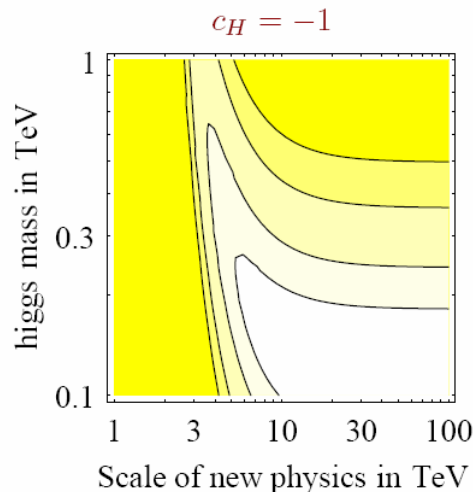
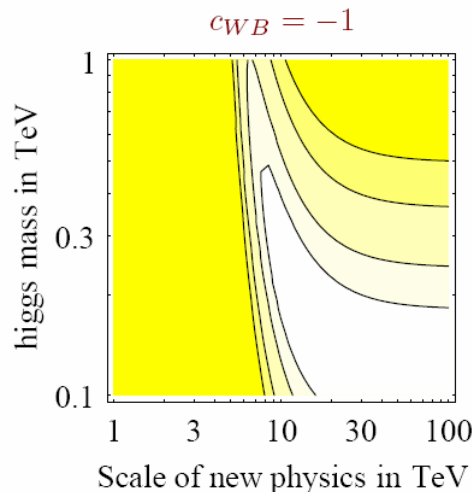


- For a WIMP, $m_X \sim 100$ GeV and $g_X \sim 0.6 \rightarrow \Omega_X \sim 0.1$

- WIMP miracle: weak-scale particles naturally have the right density to be dark matter**

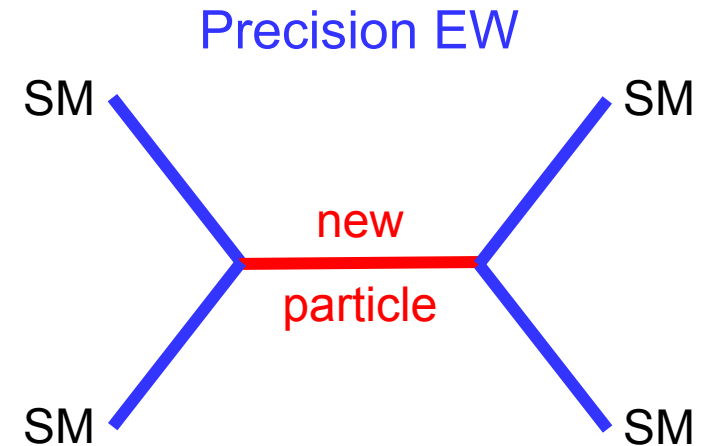
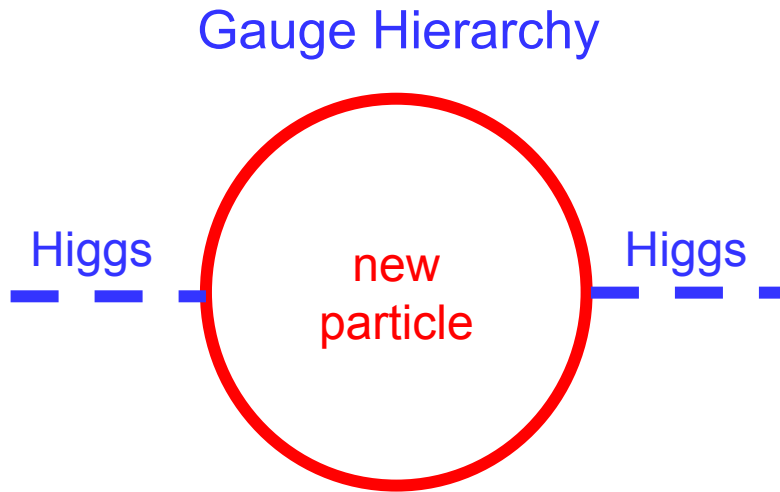
DM @ LHC: STABILITY

- The WIMP miracle assumes a new particle is stable. Why should this be?
- LEP and SLC confirmed the standard model, stringently constrained effects of new particles
- Problem: Gauge hierarchy \rightarrow new particles ~ 100 GeV
LEP/SLC \rightarrow new particles > 3 TeV
(even considering only flavor-, CP-, B-, and L-conserving effects)



Barbieri, Strumia (2000)

LEP'S COSMOLOGICAL LEGACY



- Simple solution: impose a discrete parity, so all interactions require pairs of new particles. This also makes the lightest new particle stable.

Cheng, Low (2003); Wudka (2003)

- This is a general argument that the LHC may make DM.
- In specific contexts, this may be augmented by additional arguments. E.g., in SUSY, proton decay \rightarrow R-parity \rightarrow stable LSP.

EXAMPLES

Supersymmetry

- R-parity
- Neutralino DM

Fayet, Farrar (1974)

Goldberg (1983)
Ellis et al. (1984)

Universal Extra Dimensions

- KK-parity
- Kaluza-Klein DM

Appelquist, Cheng, Dobrescu (2000)

Servant, Tait (2002)

Cheng, Feng, Matchev (2002)

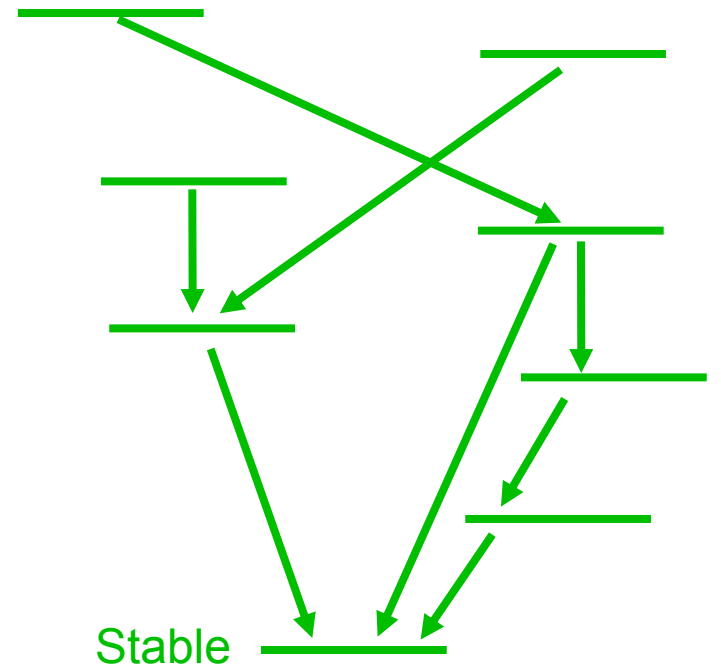
Branes

- Brane-parity
- Branons DM

Cembranos, Dobado, Maroto (2003)

...

New Particle States



Standard Model Particles

THE STANDARD LORE

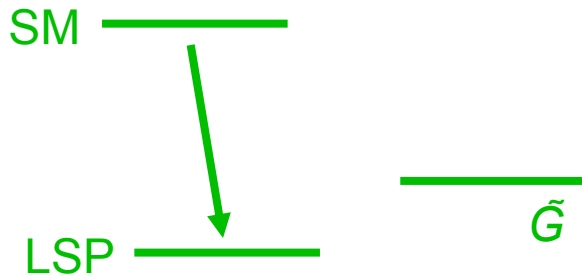
- These observations have led to the commonly heard statement that DM implies MET at the LHC.
- This is unfortunate: MET is not so great for early discoveries, even worse for precision studies.
- Fortunately, it's also wrong. There are great many other possibilities; focus here on possibilities for early discoveries, spectacular signals.

SUPERWIMP DM

Feng, Rajaraman, Takayama (2003)

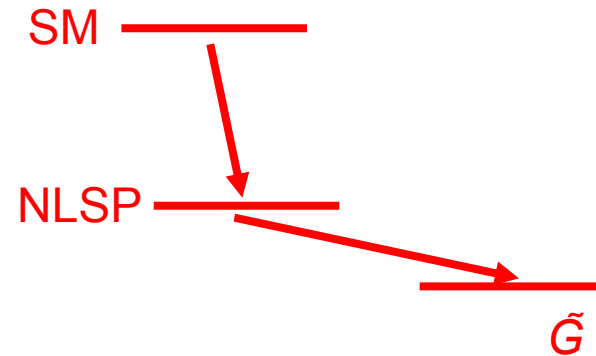
Consider (high-scale) supersymmetry. There is a gravitino, mass ~ 100 GeV, couplings $\sim M_W/M_{Pl} \sim 10^{-16}$

- \tilde{G} not LSP



- Assumption of most of literature

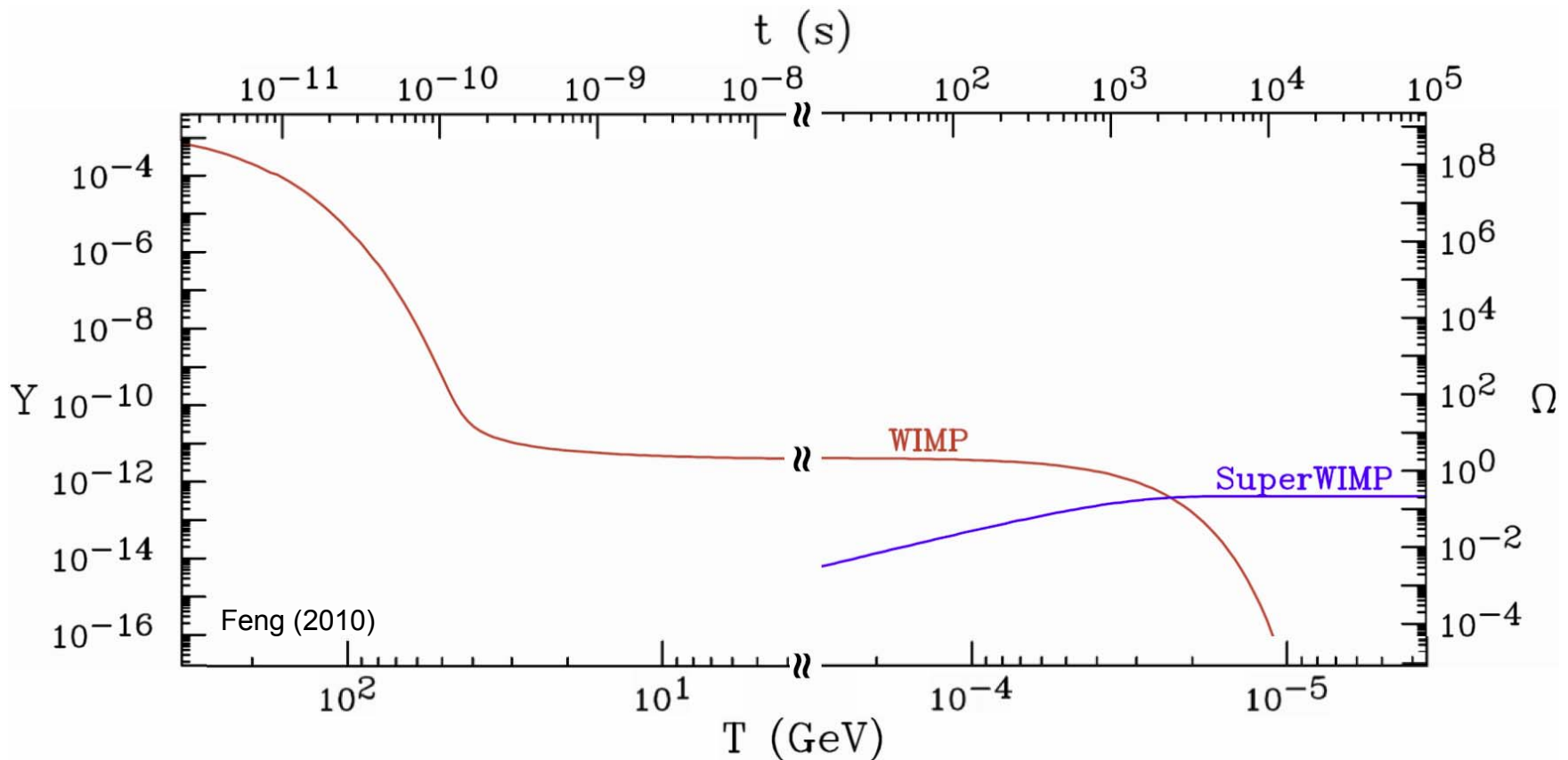
- \tilde{G} LSP



- Completely different cosmology and particle physics

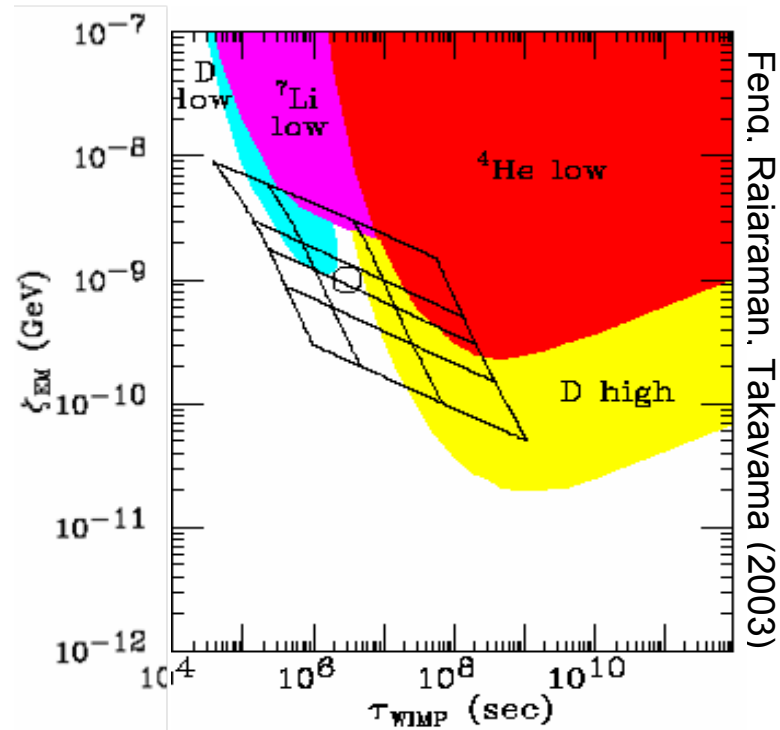
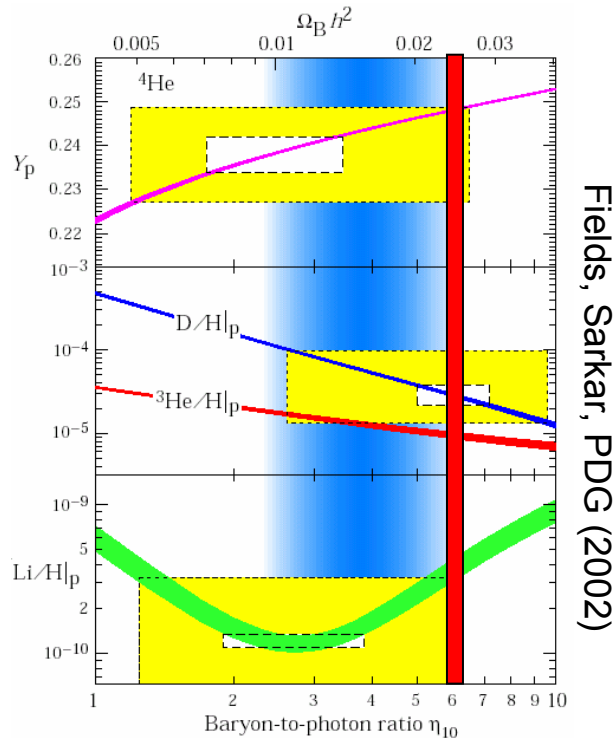
SUPERWIMP RELICS

- Consider \tilde{G} LSPs: WIMPs freeze out as usual, but then decay to \tilde{G} after $M_{\text{Pl}}^2/M_W^3 \sim$ seconds to months



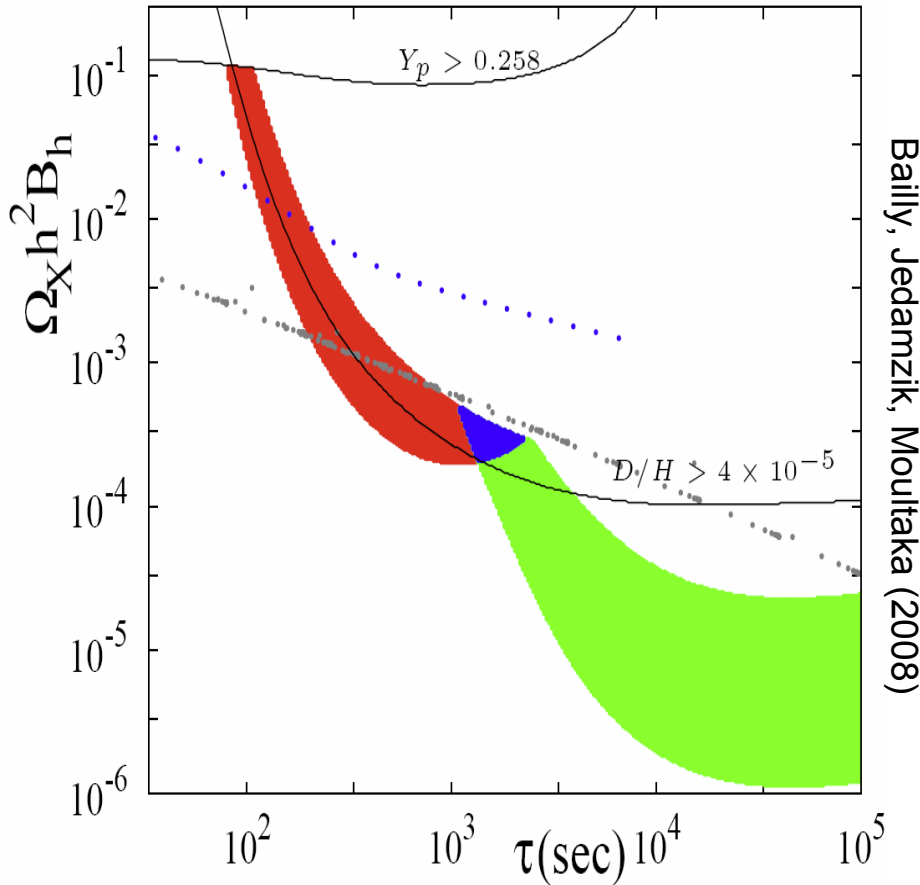
COSMOLOGY OF LATE DECAYS

Late decays impact light element abundances



- Lots of complicated nucleoparticlecosmochemistry
- BBN typically excludes very large lifetimes
- BBN excludes $\chi \rightarrow Z \tilde{G}$, but $\tilde{I} \rightarrow I \tilde{G}$ ok

LATE DECAYS AND ${}^7\text{Li}/{}^6\text{Li}$



- ${}^7\text{Li}$ does not agree with standard BBN prediction
 - Too low by factor of 3, $\sim 5\sigma$ at face value
 - May be solved by convection in stars, but then why so uniform?
- ${}^6\text{Li}$ may also not agree
 - Too high
- Late decays can fix either one or both
- For mSUGRA, fixing both, and requiring $\Omega_{\tilde{G}} = 0.1 \rightarrow$ heavy sleptons $> \text{TeV}$

BOTTOM LINE

- This scenario
 - Solves the gauge hierarchy problem
 - Has DM with naturally the right relic density
 - Is consistent with EW precision constraints
 - Is consistent with all other constraints
 - Collider signal is CHAMPs, not MET

MODEL FRAMEWORKS

- mSUGRA's famous 4+1 parameters:

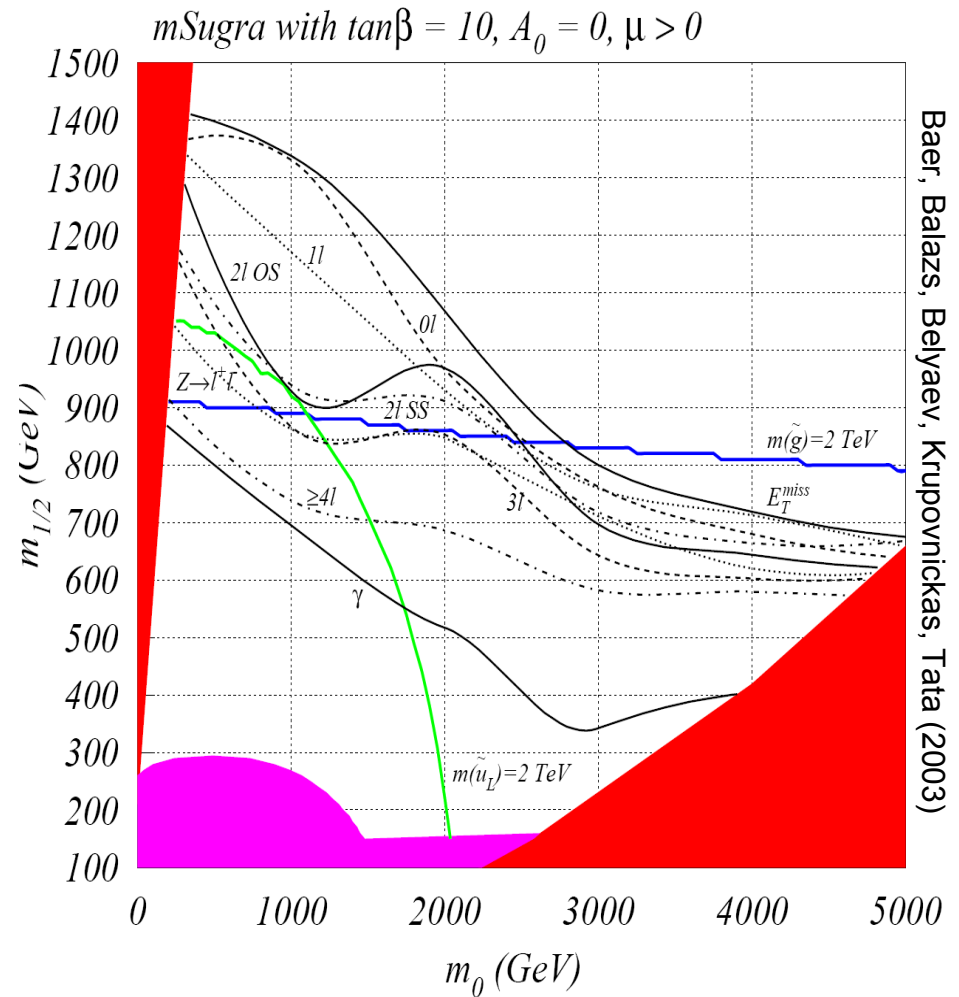
$$m_0^2, M_{1/2}, A_0, \tan \beta, \text{sign}(\mu)$$

- Excluded regions

- LEP limits
- Stau LSP

- But this is incomplete

- Missing $m_{\tilde{G}}$
- Assumes $m_0^2 > 0$



THE COMPLETE MSUGRA

- Extend the mSUGRA parameters to

$$m_0^2, M_{1/2}, A_0, \tan\beta, \text{sign}(\mu), \text{ and } m_{3/2}$$

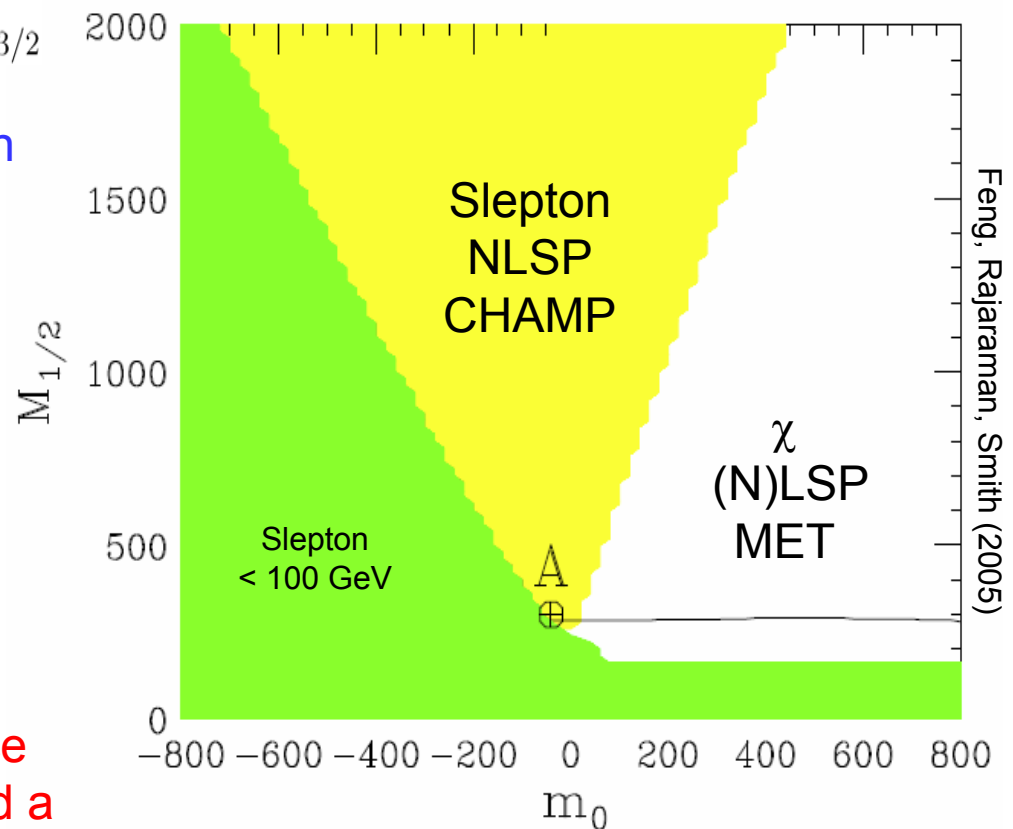
- If LSP = gravitino, then no reason to exclude stau NLSP region

- Also include small or negative

$$m_0 \equiv \text{sign}(m_0^2) \sqrt{|m_0^2|}$$

- This includes no-scale/gaugino-mediated models with $m_0 = 0$

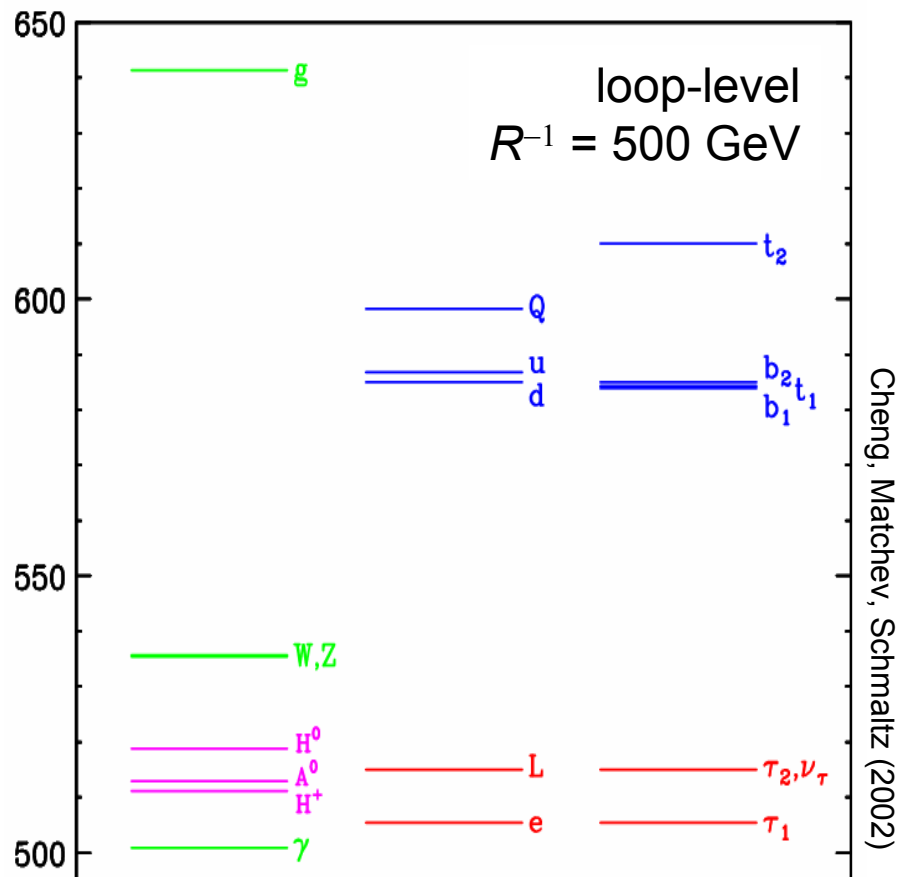
- Much of the new parameter space is viable with a slepton NLSP and a gravitino LSP



UNIVERSAL EXTRA DIMENSIONS

Appelquist, Cheng, Dobrescu (2000)

- Assume 1 extra dimension, where the 5th dimension is a circle with radius R
- All Kaluza-Klein level 1 states have mass R^{-1}
- This is broken by many effects, but the lightest KK states are still highly degenerate



UED COMMON LORE

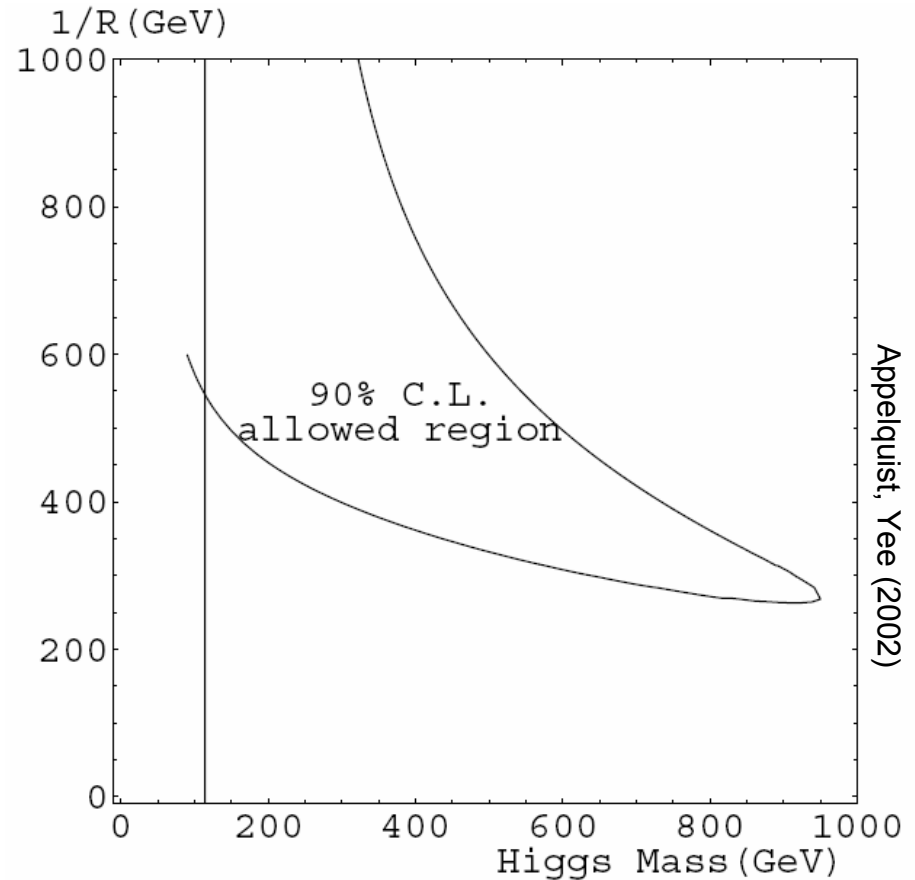
- UED looks like SUSY
 - $n=2$ and higher levels typically out of reach
 - $n=1$ Higgses $\rightarrow A, H^0, H^\pm$
 - Colored particles are heavier than uncolored ones
 - LKP is stable $B^1 \rightarrow \text{MET}$ at LHC
- Spectrum is more degenerate, but basically similar to SUSY

“Bosonic supersymmetry”

Cheng, Matchev, Schmaltz (2002)

BUT THERE'S MORE

- R is the only new parameter, but it is not the only free parameter: the Higgs boson mass is unknown
- Original collider studies set $m_h=120$ GeV, but it can be larger (KK towers modify EW precision constraints)
- H^0, A, H^\pm masses depend on m_h
- Also, there's another state in the theory: the KK graviton G^1

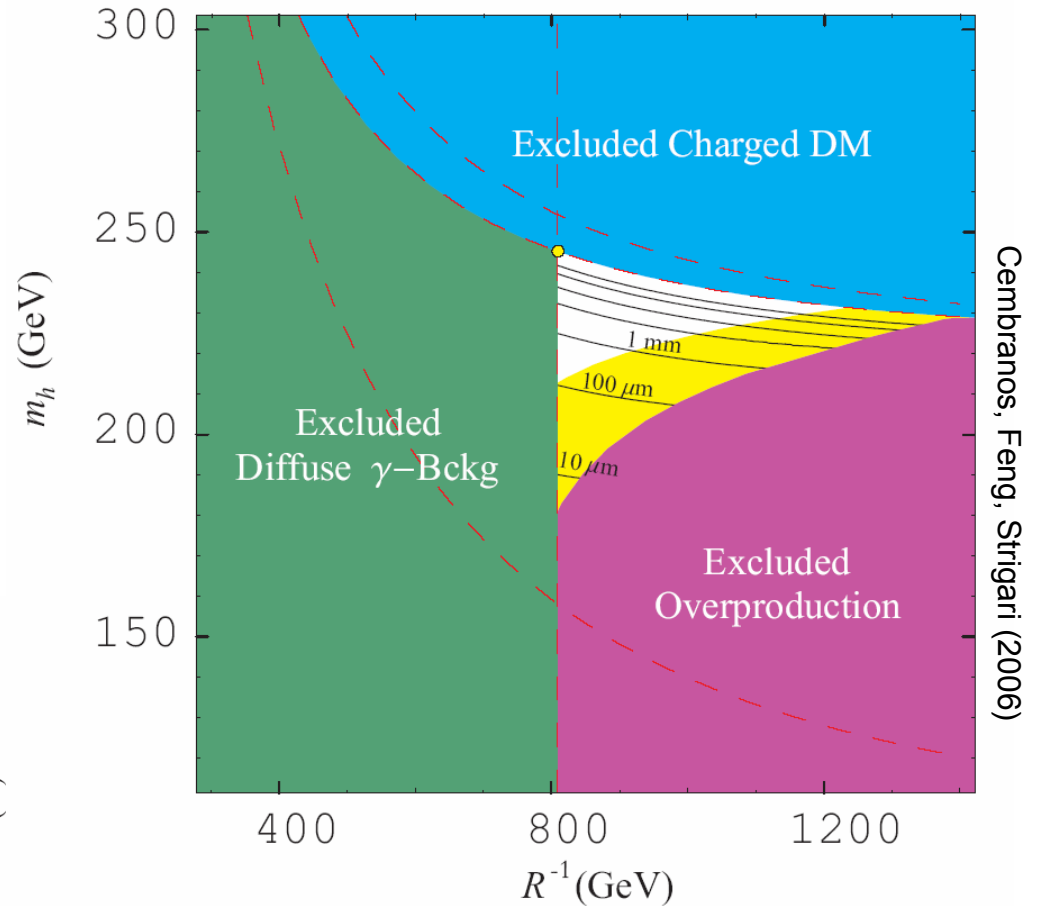
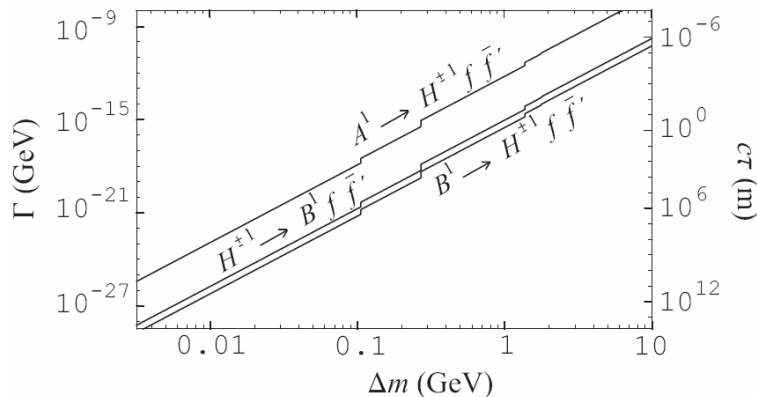


THE COMPLETE (MINIMAL) UED

- In minimal UED, after all particle and astrophysical constraints, NLKP \rightarrow LKP is

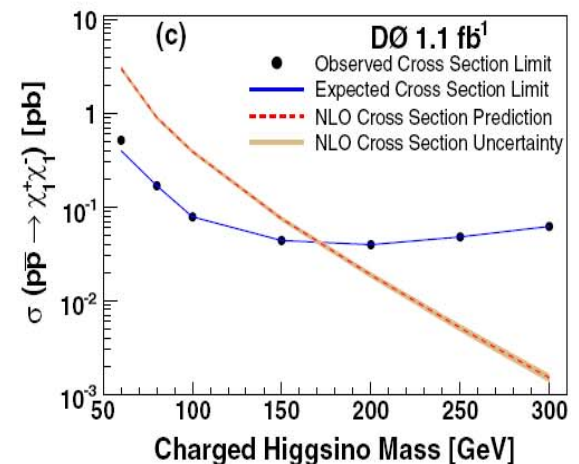
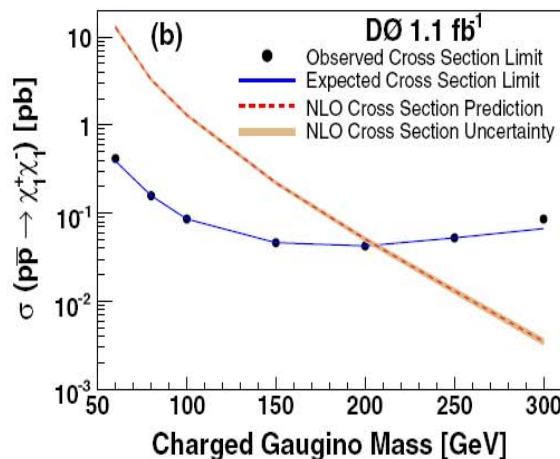
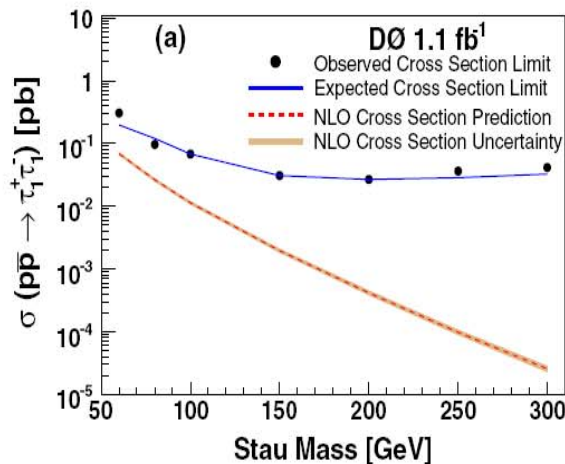
$$H^{\pm 1} \rightarrow B^1 f f'$$

- Mass splitting $\Delta m < 7$ GeV
- Decay length $c\tau > 10 \mu\text{m}$



CURRENT BOUNDS

- Current Bounds
 - LEP: slepton mass > 97.5 GeV, chargino > 102.5 GeV
 - CDF Run I: slepton cross section < 1 pb
 - D0 Run II: slepton cross section < 0.1 pb
 - assumes only Drell-Yan pair production (no cascades)
 - require 2 slow, isolated “muons”
 - about a factor of 5 from unexplored mass territory



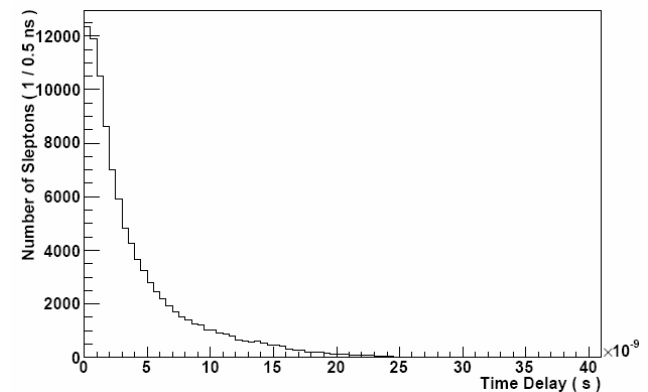
LHC DISCOVERY POTENTIAL

Rajaraman, Smith (2006)

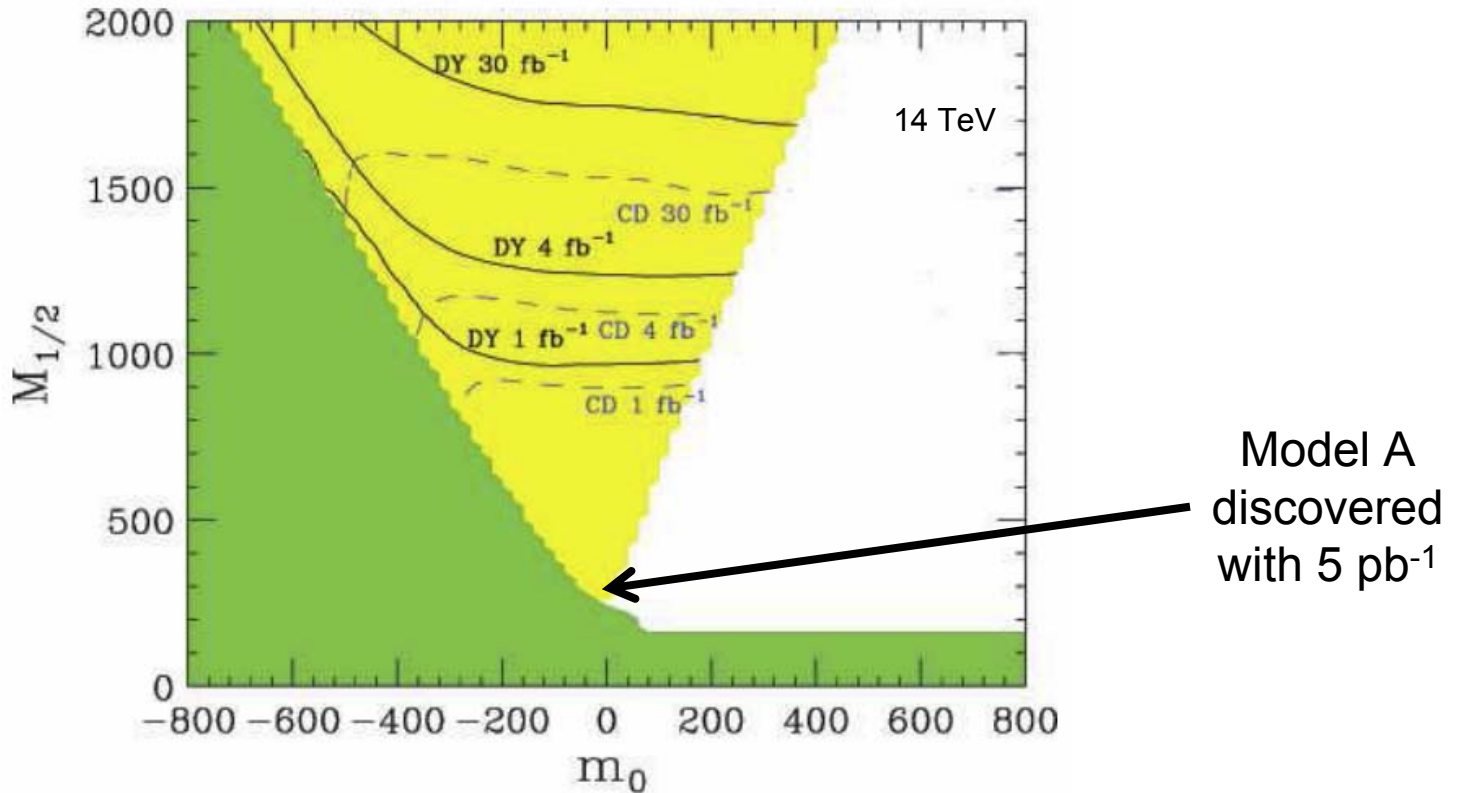
- Look for Drell-Yan slepton pair production
- Require events with 2 central, isolated “muons” with
 - $p > 100 \text{ GeV}$
 - $p_T > 20 \text{ GeV}$
- Finally assume TOF detector resolution of 1 ns, require both muons to have TOF delays $> 3 \text{ ns}$

	Total cross-section	After Drell-Yan cuts
Model A	18pb	9pb
Model B	43fb	28fb
QCD	10^2mb	$< 1 \text{pb}$
$\gamma^*/Z \rightarrow \mu\mu$	100nb	3pb
W+jet	360nb	$< 40 \text{fb}$
Z+jet	150nb	7pb
$t\bar{t}$	800pb	430fb
WW,WZ,ZZ	2.5nb	150fb

Time delay of	0ns	1 ns	2ns	3ns	4ns	5ns
Drell-Yan; background	10pb	1.35pb	3.3fb	0.2ab	$< 0.1 \text{ab}$	$< 0.1 \text{ab}$
Drell-Yan; Model A	9pb	5.2pb	2.9pb	1.8pb	1.1 pb	750fb



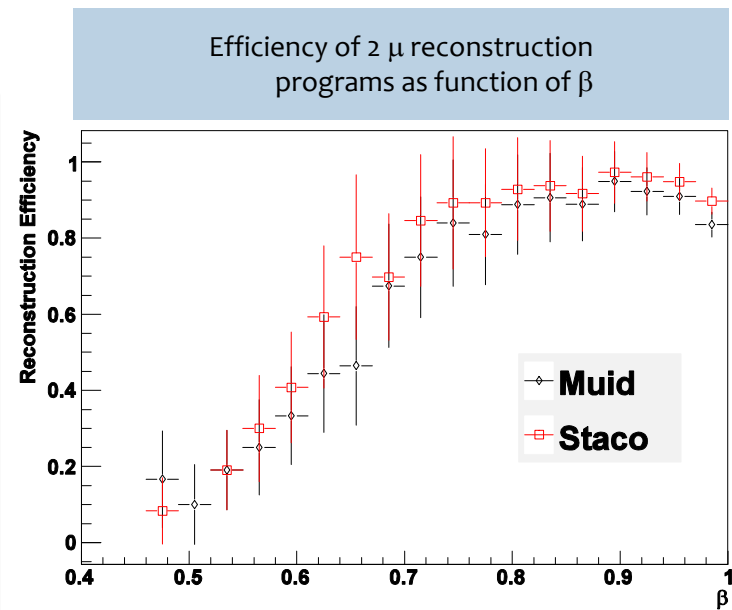
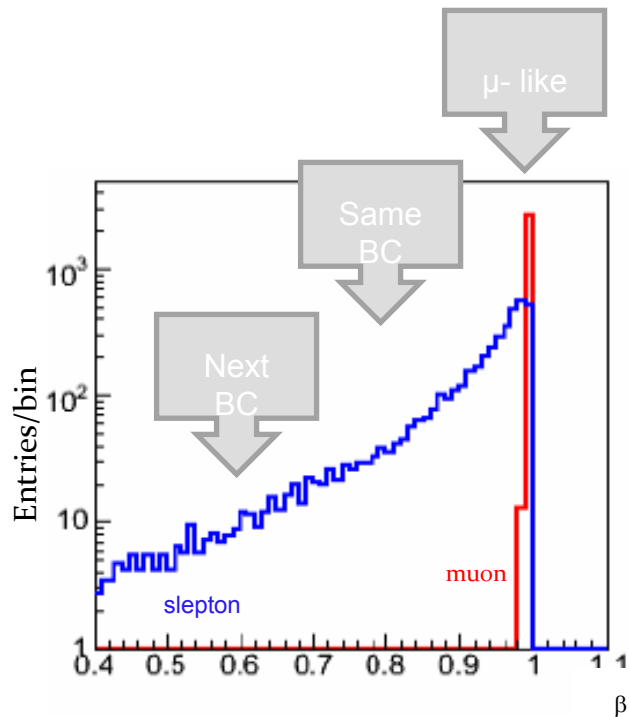
- Require 5σ signal with $S > 10$ events for discovery



- Model A is “best case scenario”
- Lesson: Very early on, the LHC will probe new territory

CMS/ATLAS ANALYSES

- Ongoing work on CHAMP search and reconstruction
 - ATLAS (Tarem et al.): added ToF calculation to level 2 trigger to improve reconstruction efficiency
 - CMS (Rizzi): studied both dE/dx and ToF (Analysis Note (2006))



Tarem, Bressler, Nomoto, De Mattia (2008)

LIGHT GRAVITINO DM

- The original SUSY DM scenario
 - Universe cools from high temperature
 - Gravitinos decouple while relativistic, $\Omega_{\tilde{G}} h^2 \approx m_{\tilde{G}} / 800 \text{ eV}$
 - Favored scenario: keV gravitinos

Pagels, Primack (1982)

- This minimal scenario is now excluded
 - $\Omega_{\tilde{G}} h^2 < 0.1 \rightarrow m_{\tilde{G}} < 80 \text{ eV}$
 - Gravitinos not too hot $\rightarrow m_{\tilde{G}} > \text{few keV}$
 - Disfavored scenario: keV gravitinos

Viel, Lesgourgues, Haehnelt, Matarrese, Riotto (2005)
Seljak, Makarov, McDonald, Trac (2006)

- Two ways out
 - Λ WDM: $m_{\tilde{G}} > \text{few keV}$. Gravitinos are all the DM, but thermal density is diluted by low reheating temperature,...
 - Λ WCDM: $m_{\tilde{G}} < 16 \text{ eV}$. Gravitinos are only part of the DM, mixed warm-cold scenario

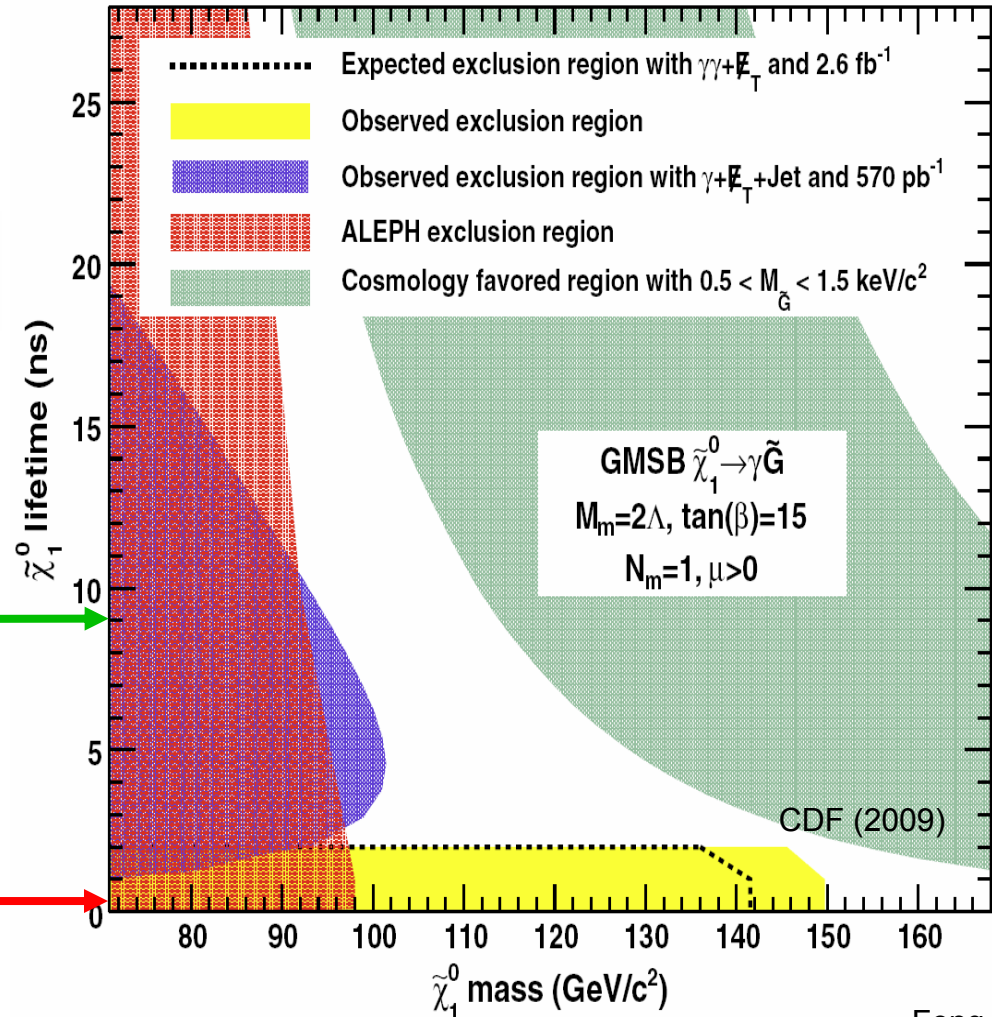
CURRENT BOUNDS

- Remarkably, this lifetime difference is observable at colliders!

$$c\tau_{\text{NLSP}} \approx 50 \text{ cm} \left(\frac{200 \text{ GeV}}{m_{\text{NLSP}}} \right)^5 \left(\frac{m_{\tilde{G}}}{\text{keV}} \right)^2$$

- $m_{\tilde{G}} > \text{few keV}$:
Delayed photon signatures

- $m_{\tilde{G}} < 16 \text{ eV}$:
Prompt photon signatures

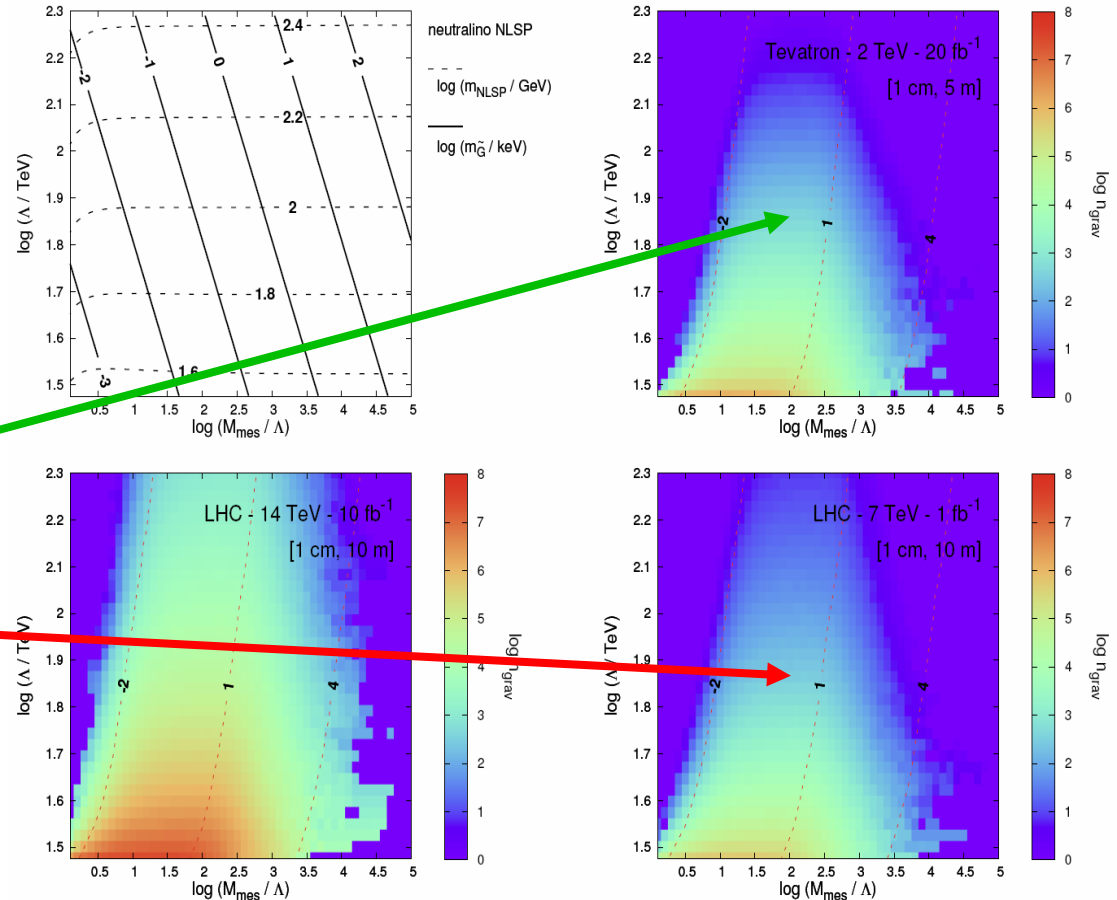


LIGHT GRAVITINOS AT THE LHC

- Hundreds to thousands of delayed photon events are possible at

Tevatron (20 fb⁻¹)

LHC (7 TeV, 1 fb⁻¹)



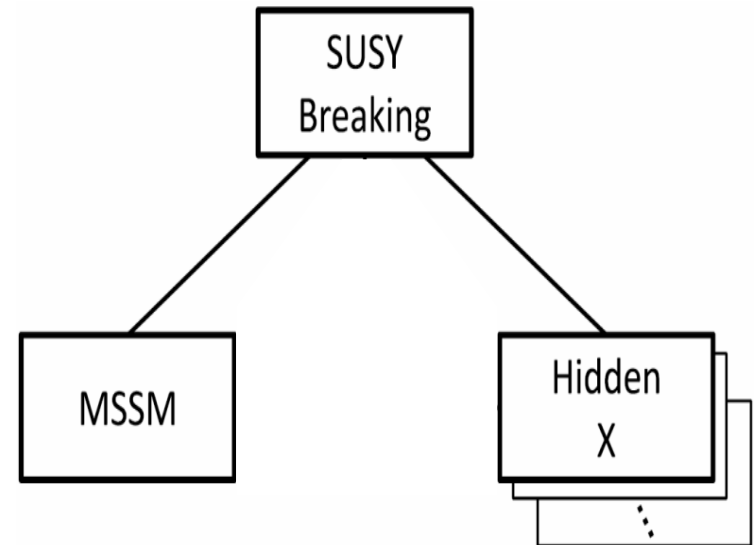
Lee, Feng, Kamionkowski (in preparation)

WIMPLESS DM

Feng, Kumar (2008); Feng, Tu, Yu (2008)

- Consider hidden DM, that is, DM with no SM gauge interactions. What can we say?
- Generically, nothing
- But in SUSY models that solve the new physics flavor problem (gauge-mediated models, anomaly-mediated models) the superpartner masses are determined by gauge couplings

$$m_X \sim g_X^2$$

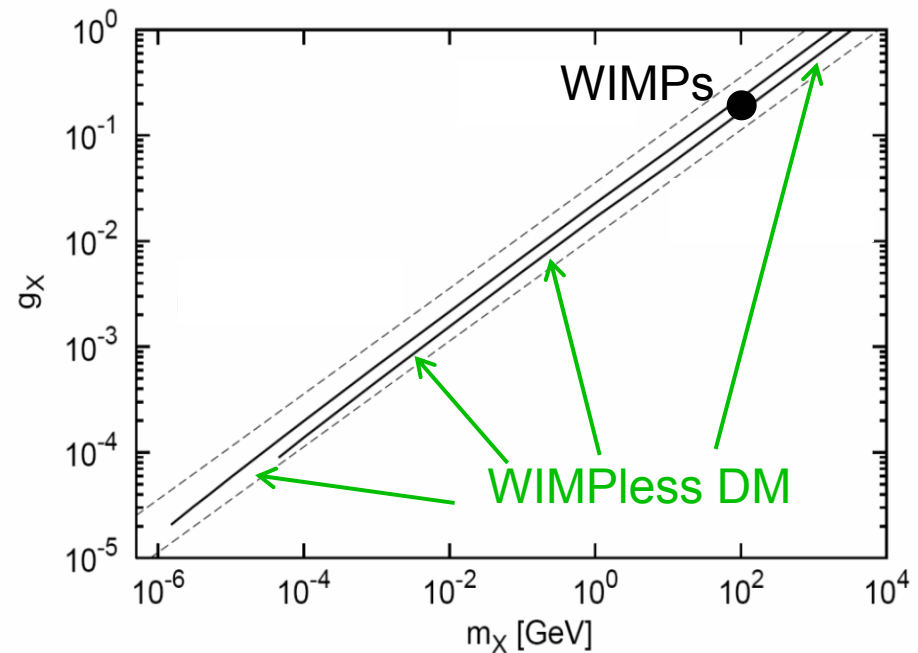


THE WIMPLESS MIRACLE

- This leaves the thermal relic density invariant:

$$\Omega_X \propto \frac{1}{\langle \sigma v \rangle} \sim \frac{m_X^2}{g_X^4}$$

- The thermal relic density constrains only one combination of g_X and m_X . These models map out the remaining degree of freedom; candidates have a range of masses and couplings, but always the right relic density.



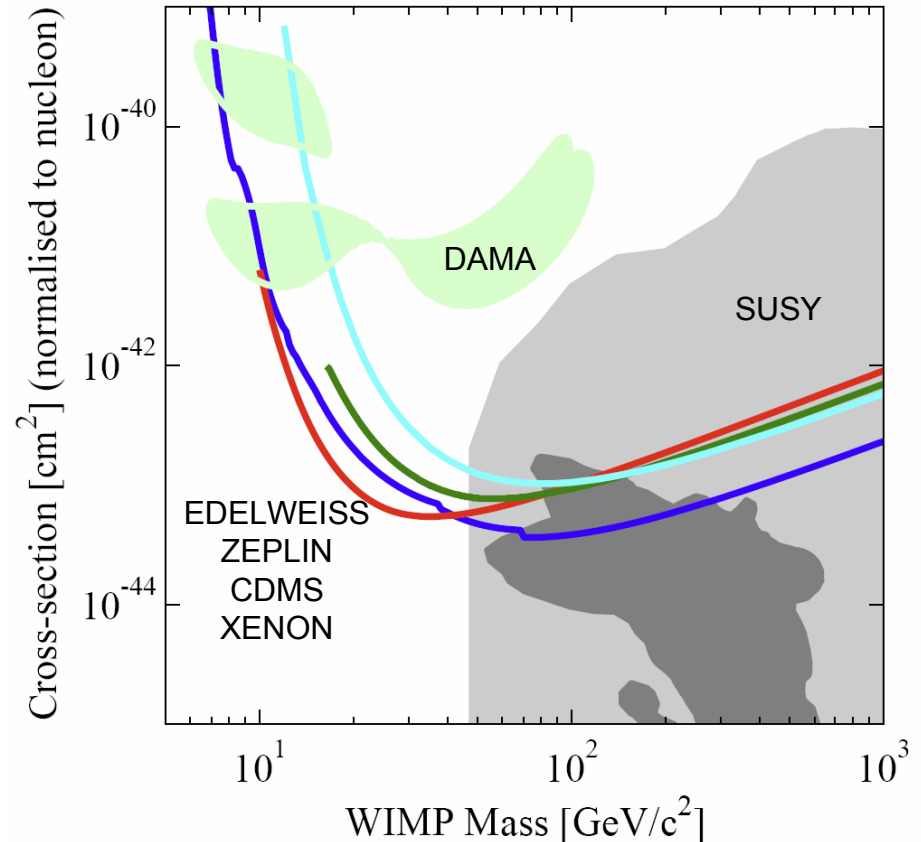
- This decouples the WIMP miracle from WIMPs (is this what the flavor problem is really trying to tell us?)

DAMA

- The DAMA signal is now at 8.9σ
- The $m_\chi \sim 100$ GeV favored region is excluded by many other experiments
- One marginally viable solution is light dark matter with very large cross section

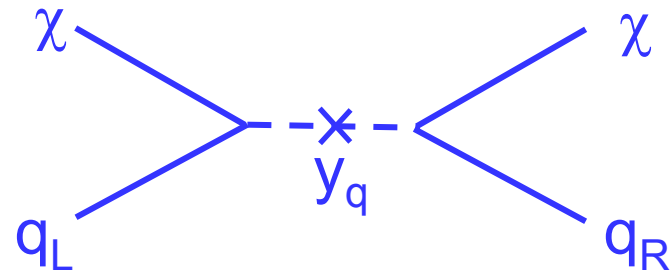
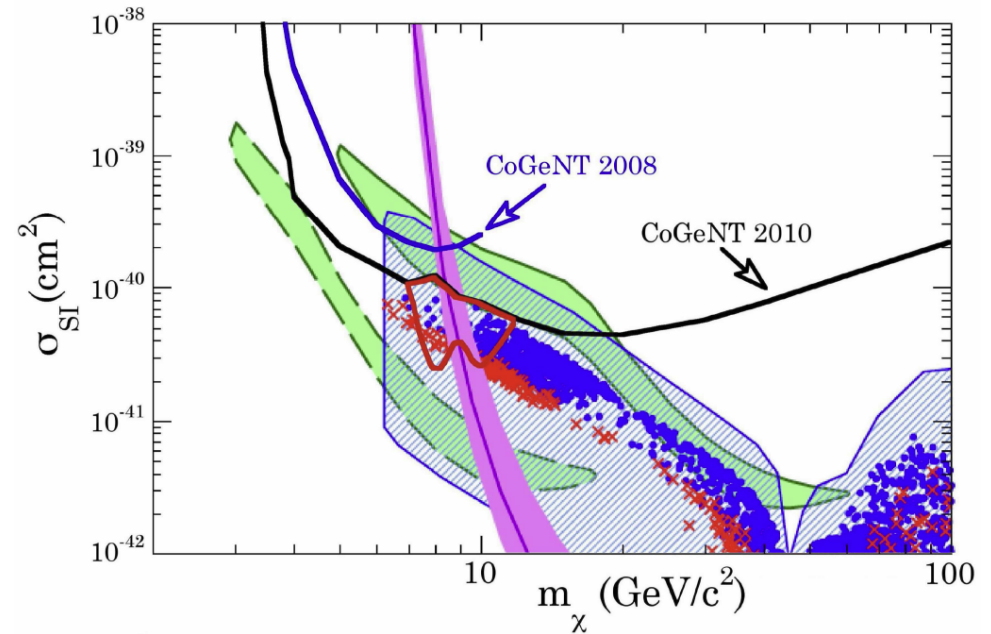
$$m_\chi \sim 5-10 \text{ GeV}$$

$$\sigma_{SI} \sim 10^{-40} \text{ cm}^2$$



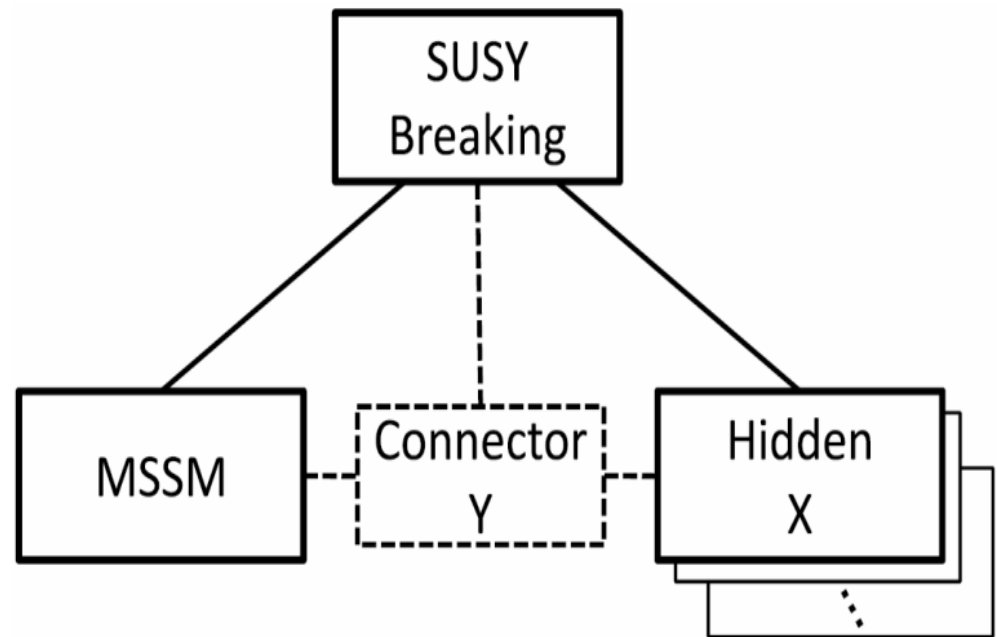
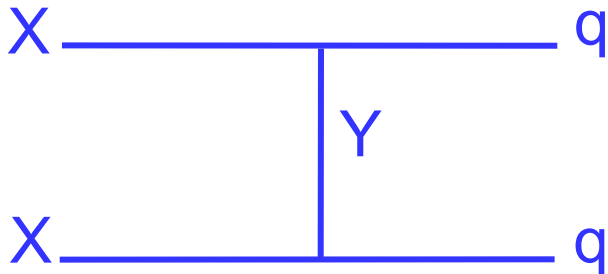
COGENT

- The light DM scenario is, however, now tentatively supported by CoGeNT
- This region of parameter space is very hard to obtain with conventional WIMPs: for example, for neutralinos, chirality flip implies large suppression



WIMPLESS SIGNALS

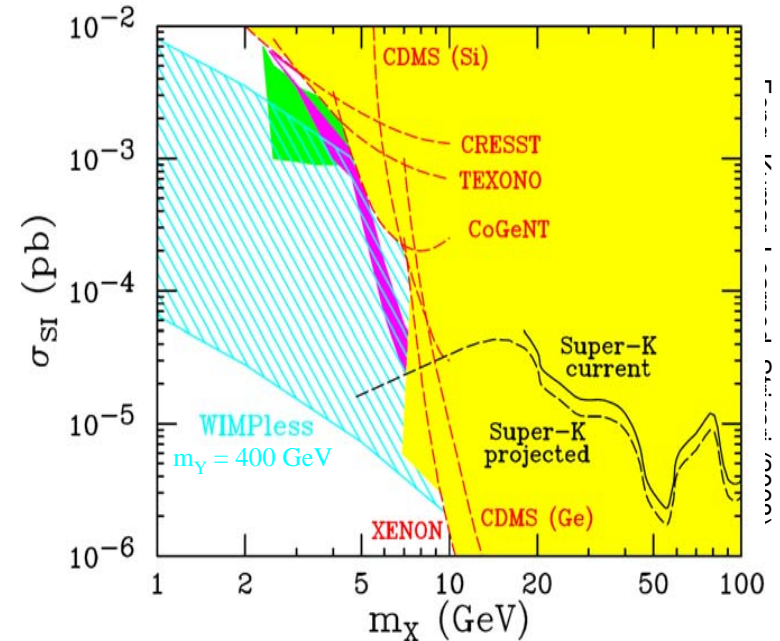
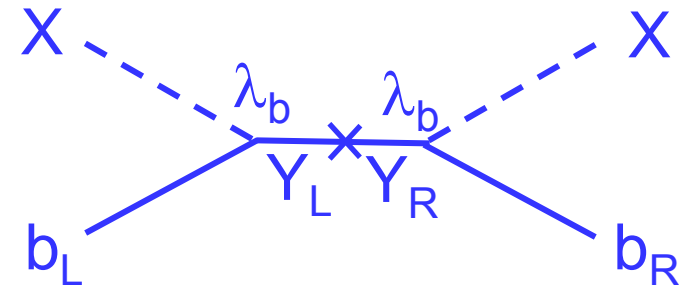
- Hidden DM may interact with normal matter through non-gauge interactions



WIMPLESS DIRECT DETECTION

- The DAMA/CoGeNT region is easy to reach with WIMPless DM
- E.g., assume WIMPless DM X is a scalar, Y is a fermion, interact with b quarks through

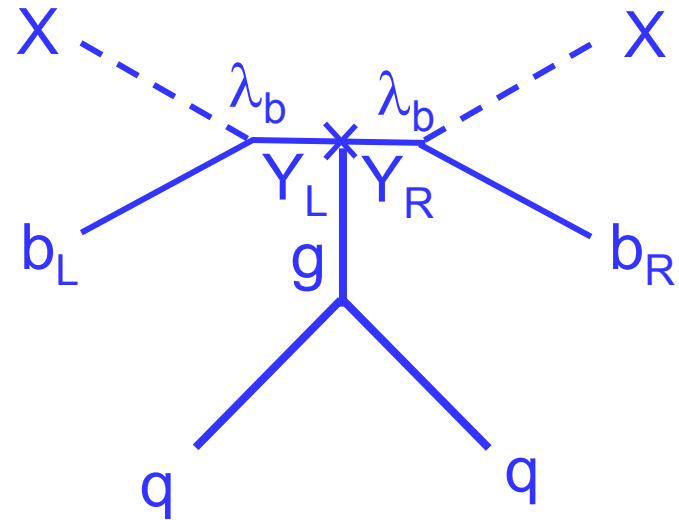
$$\lambda_b (XY_L b_L + XY_R b_R) + m_Y Y_L Y_R$$
- Naturally correct mass, cross section
 - $m_X \sim 5\text{-}10$ GeV (WIMPless miracle)
 - large σ_{SI} for $\lambda_b \sim 0.3 - 1$ (flip chirality on heavy Y propagator)



EXOTIC T' , B' AT LHC

Alwall, Feng, Kumar, Su (2010)

- Gauge invariance \rightarrow Y 's are colored, similar to 4th generation quarks
- EW precision studies, direct searches, perturbativity \rightarrow
 $300 \text{ GeV} < m_Y < 600 \text{ GeV}$



- **Signals**
 - $T' T' \rightarrow tXtX$, $B' B' \rightarrow bX bX$
 - different from standard 4th generation quarks ($T' \rightarrow b W$)
 - best channel is all hadronic

MadGraph, MadEvent, Pythia 6.4.20, PGS4)

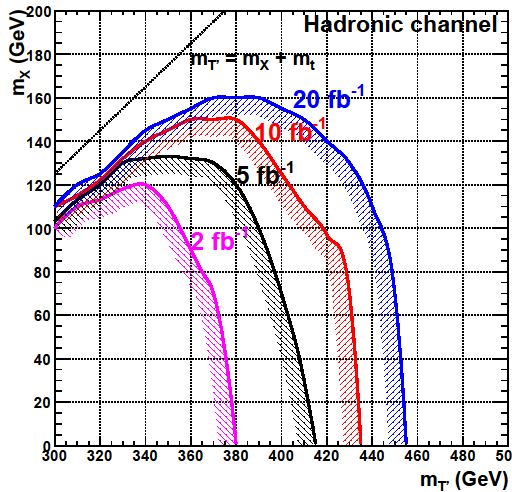
Cut	T' (300)	T' (400)	T' (500)	$t\bar{t}$ (1 τ)	$t\bar{t}$ (1 e/μ)	$t\bar{t}$ (had)	W +jets	Z +jets
No cut	14.89	3.16	0.922	43.96	66.67	104.59	(42.28)	(18.86)
0 isolated leptons	6.75	1.5	0.45	16.88	13.11	72.29	(16.8)	(15.71)
$E_T > 100 \text{ GeV}$	4.15	1.21	0.394	3.91	2.67	0.097	(11.25)	(11.48)
≥ 5 jets	1.34	0.406	0.135	0.664	0.47	0.031	0.305	0.212
$\Delta\phi$ cuts	1.19	0.374	0.125	0.56	0.41	0.01	0.265	0.187

EXOTIC 4TH QUARKS AT LHC

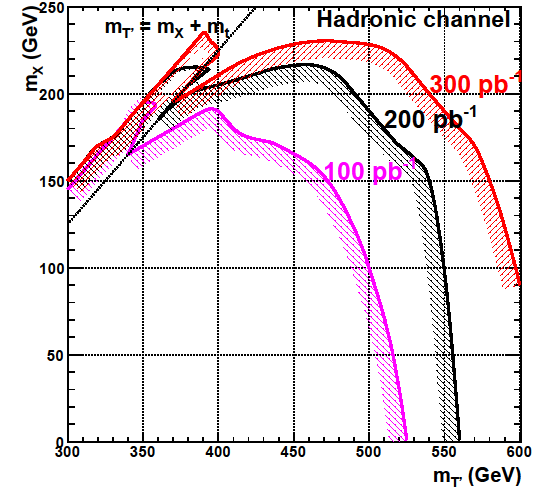
- Entire $m_X \sim 10$ GeV region can be excluded by 10 TeV LHC with 300 pb^{-1} (~ 7 TeV LHC with 1 fb^{-1})

- Significant discovery prospects with early LHC data

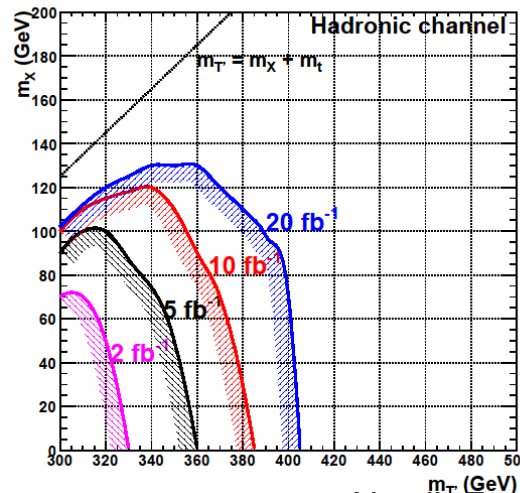
Exclusion for $T' \bar{T}' \rightarrow t X \bar{t} X$ at the Tevatron



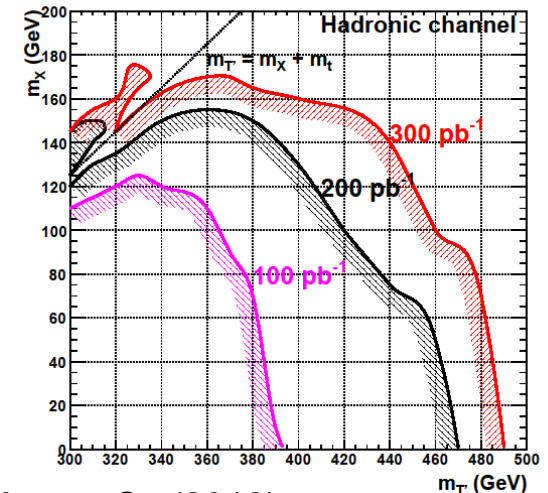
Exclusion for $T' \bar{T}' \rightarrow t X \bar{t} X$ at 10 TeV LHC



Discovery of $T' \bar{T}' \rightarrow t X \bar{t} X$ at the Tevatron



Discovery for $T' \bar{T}' \rightarrow t X \bar{t} X$ at 10 TeV LHC



Alwall, Feng, Kumar, Su (2010)

CONCLUSIONS

- There are many signals that are just as well-motivated by dark matter, the gauge hierarchy problem, and all other existing constraints as MET
 - CHAMPS
 - Delayed photons, diphotons
 - Exotic 4th generation quarks
- Real possibilities for early LHC data (and in some cases, for the Tevatron)