

# Energy Scaling of MB Tunes



P. Skands (CERN)  
with H. Schulz

# Two Issues

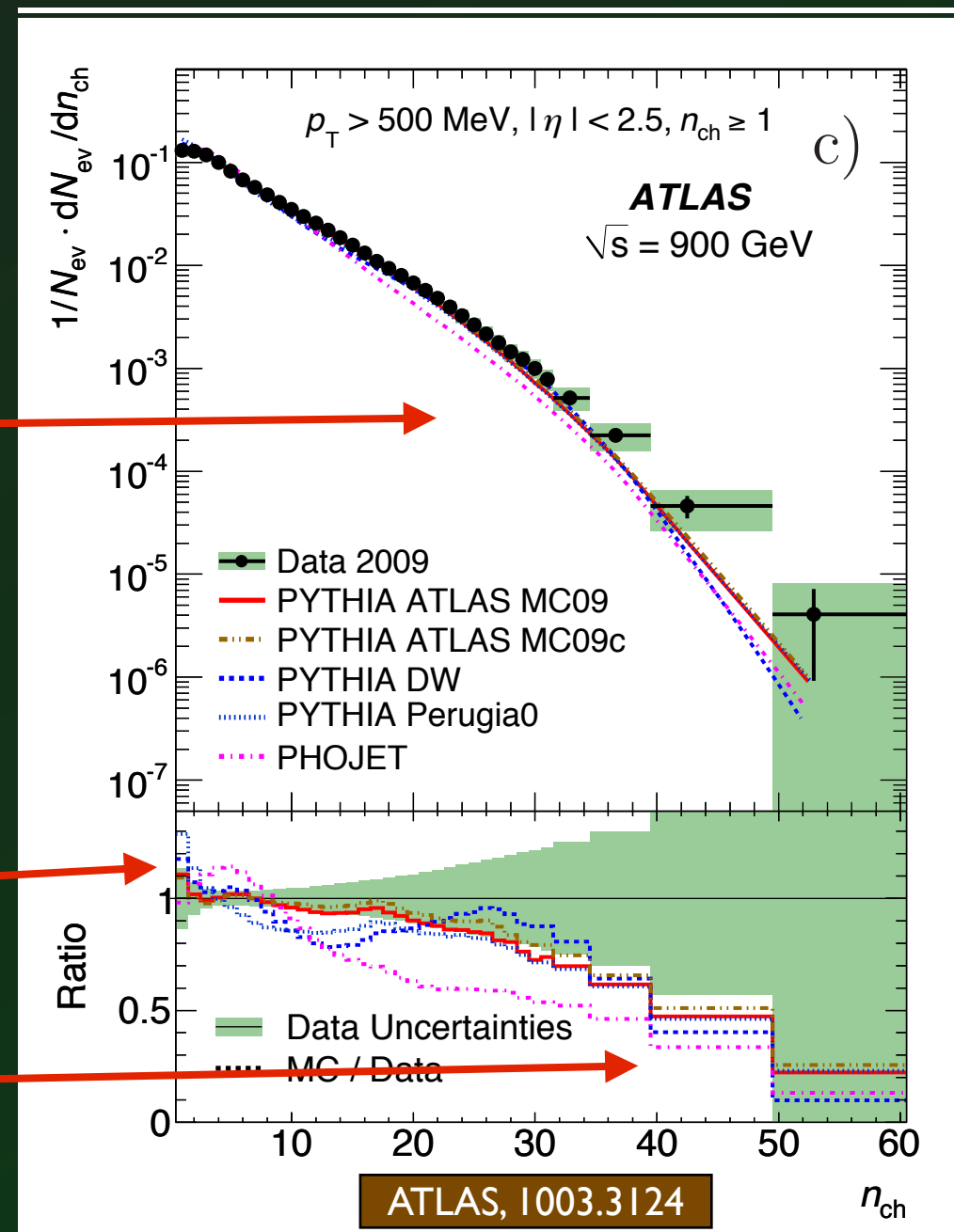
## Multiplicities are TOO LOW

Many models/tunes are  
slightly low even at 900 GeV

(though  $\approx 20\%$  on IR  
sensitive quantity not bad)

Diffraction?

Slightly wrong asymptotic slope?



Note: can't see very much from  $dN/d\eta$  alone

# Two Issues

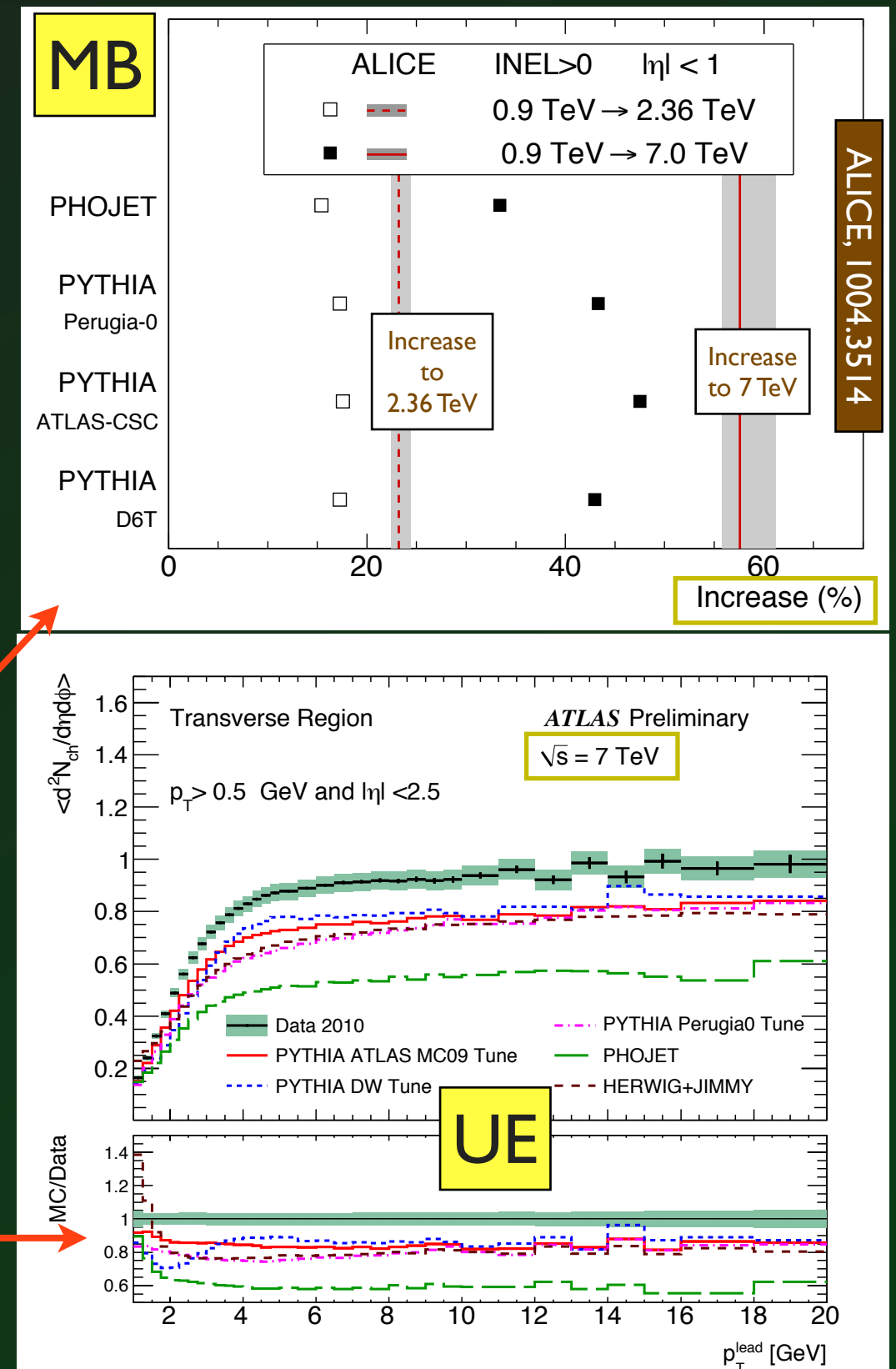
## Multiplicities are TOO LOW

Many models/tunes are slightly low even at 900 GeV  
(though  $\approx 20\%$  on IR sensitive quantity not bad)

## + SCALE TOO SLOWLY

→ Even lower at 7 TeV

→ too low UE



# Beyond Multiplicities

**ESSENTIAL to consider several distributions simultaneously:**

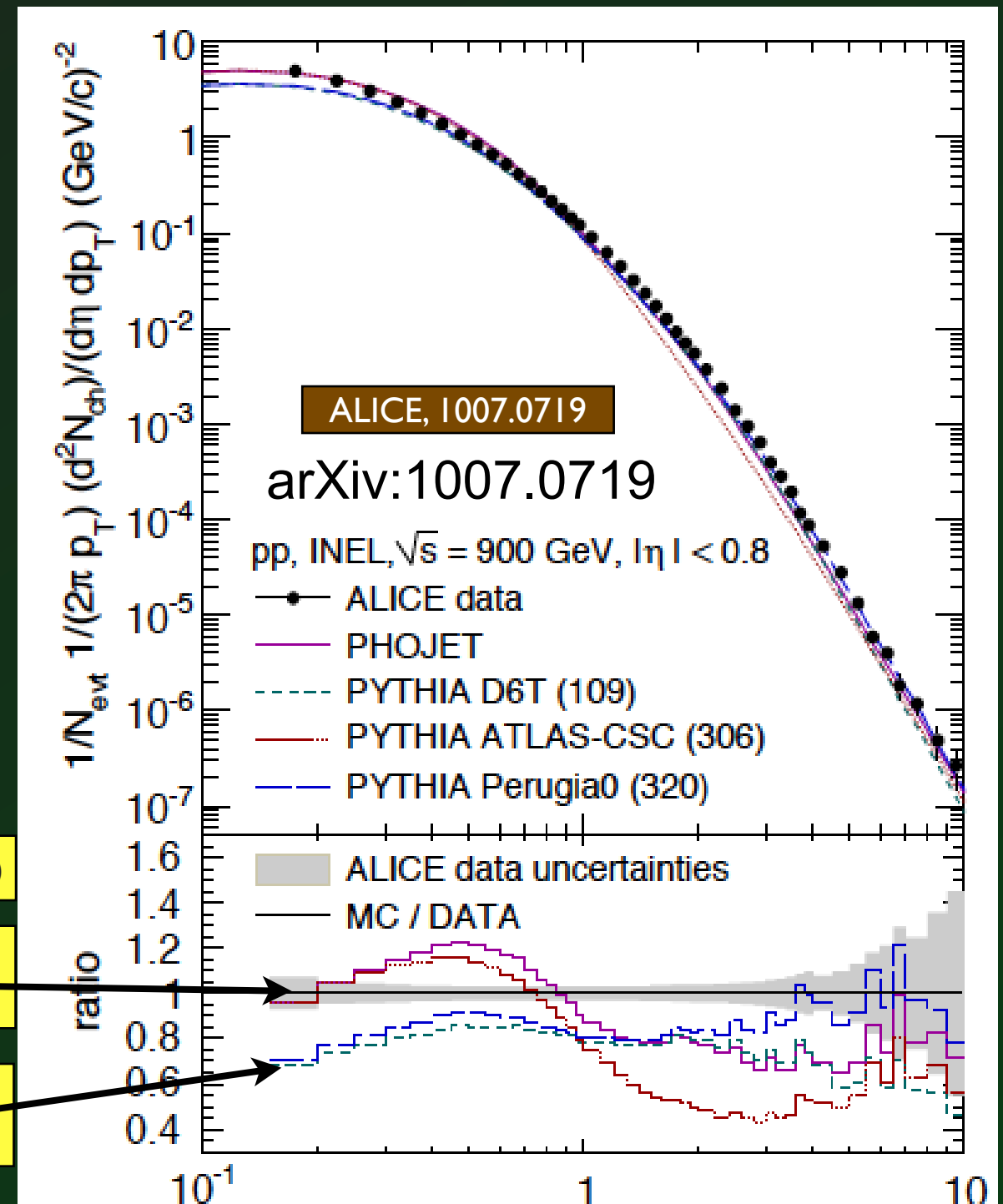
*“Those that reproduce the multiplicity don’t reproduce the  $p_T$  distributions and vice versa”*

J. Fiete Grosse-Oetringhaus

+ C. Zampolli (different tune in different PS region)

Normalization fine, shape wrong

Normalization wrong, shape fine



# An Organized View

## 1. Where is the energy going?

Note: only linearized Sphericity is IR safe

*Sum( $p_T$ ) densities, event shapes, mini-jet rates, energy flow correlations...*

## 2. How many tracks is it divided onto?

*$N_{\text{tracks}}$ ,  $dN_{\text{tracks}}/dp_T$ , Associated track densities, track correlations...*

## 3. What kind of tracks?

*Strangeness per track, baryons per track, ...*

*Further: strange baryons per strange, strange-antistrange correlations, ... ..*

IR Safe

IR Sensitive

More IR Sensitive

# Action Items

## 1. Need better models for diffraction

Tuning is fast - but modeling takes time

*Physical observables, in diffractively enriched samples*

*+ data preservation (HEPDATA/Rivet) → can test any future model*

cf., e.g.,  
ATLAS (L. Tompkins)  
CMS (H. Jung, M. Velasco)

## 2. Get Organized

**Global View:** Consider each model on several observables in several phase-space regions simultaneously → better conclusions

**Factorized:** Order observables from IR safe to IR sensitive

## 3. Need better understanding of E-scaling

E-scaling allows to consolidate measurements from different colliders  
→ powerful cross check on physics model

*While waiting for better model of diffraction, isolate and continue testing non-diffractive tail of MB + Systematically compare to LEP (jet fragmentation) & UE*

# Energy Scaling

**Can we be more general than this-tune-does-this, that-tune-does-that?**

Yes.

*The new automated tuning tools allow us to get an Unbiased optimization at each collider separately*

- counter-check the model assumptions on energy scaling
- + counter-check the consistency of the interpolations
- + differences give a new kind of uncertainty estimate

**Critical for this task:**

*“Comparable” data set at each different collider*

# Scaling according to Holger

(Schulz)

**MCnet/LPCC Summer Student**

(+co-author of Professor)

**Used CDF, UA5, and ATLAS data**

$P(N_{ch}), dN_{ch}/dp_T, \langle p_T \rangle(N_{ch})$

*+ can even focus on  $N_{ch} \geq 6$  sample separately!*

From 630 GeV to 7 TeV (we would have liked to add STAR at 200 GeV, but we did not have a complete obs set from them)

**Reduce model to 3 main parameters:**

*Starting point = Perugia 0*

1. Infrared Regularization Scale

PARP(82)

2. Proton Transverse Mass Distributions

PARP(83)

3. Strength of Color Reconnections

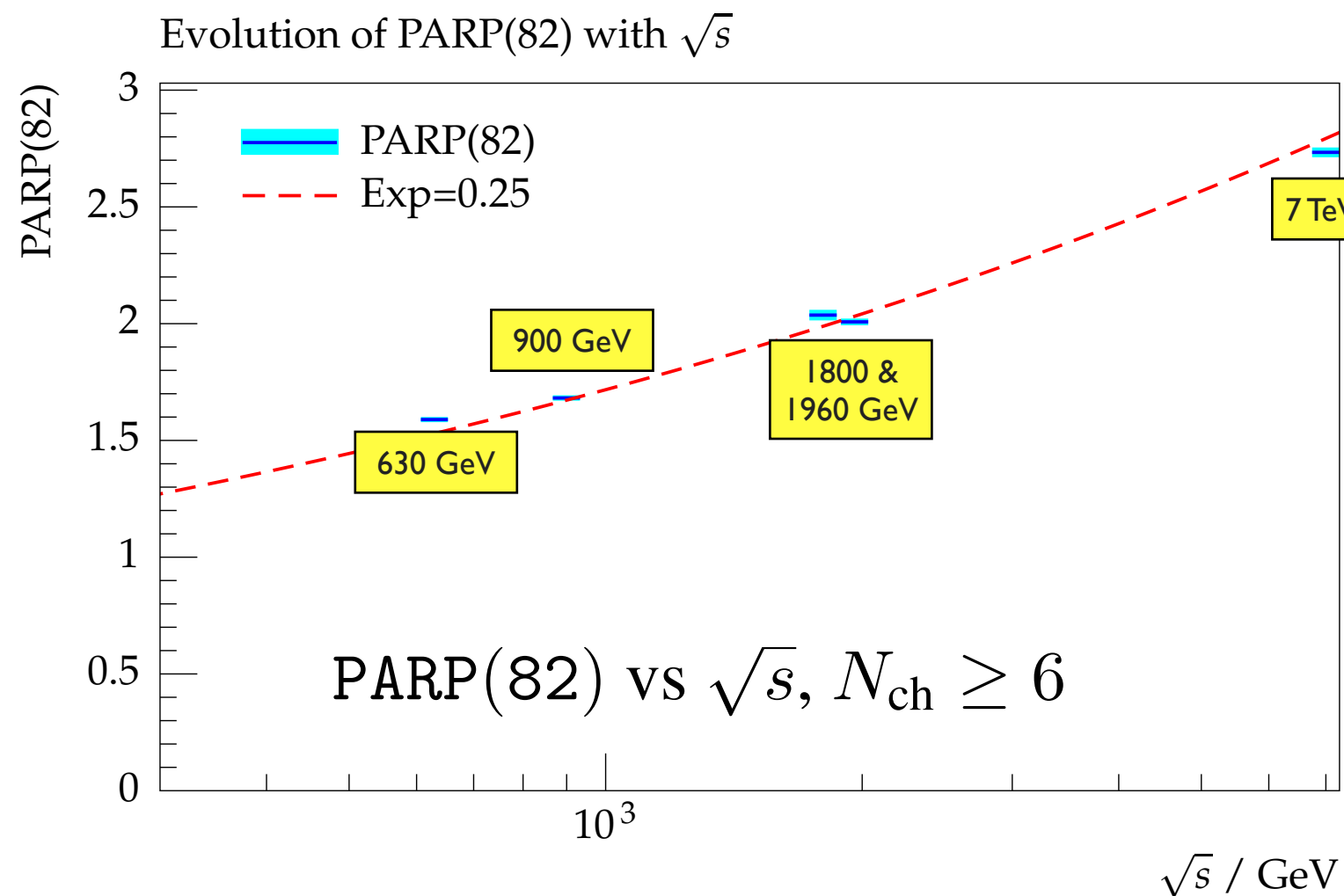
PARP(78)



# Infrared Regularization

## Independent tunings compared to Perugia 0

Rather striking agreement with the assumed functional form (Perugia-0 uses  $\text{PARP}(90) = 0.25$ )

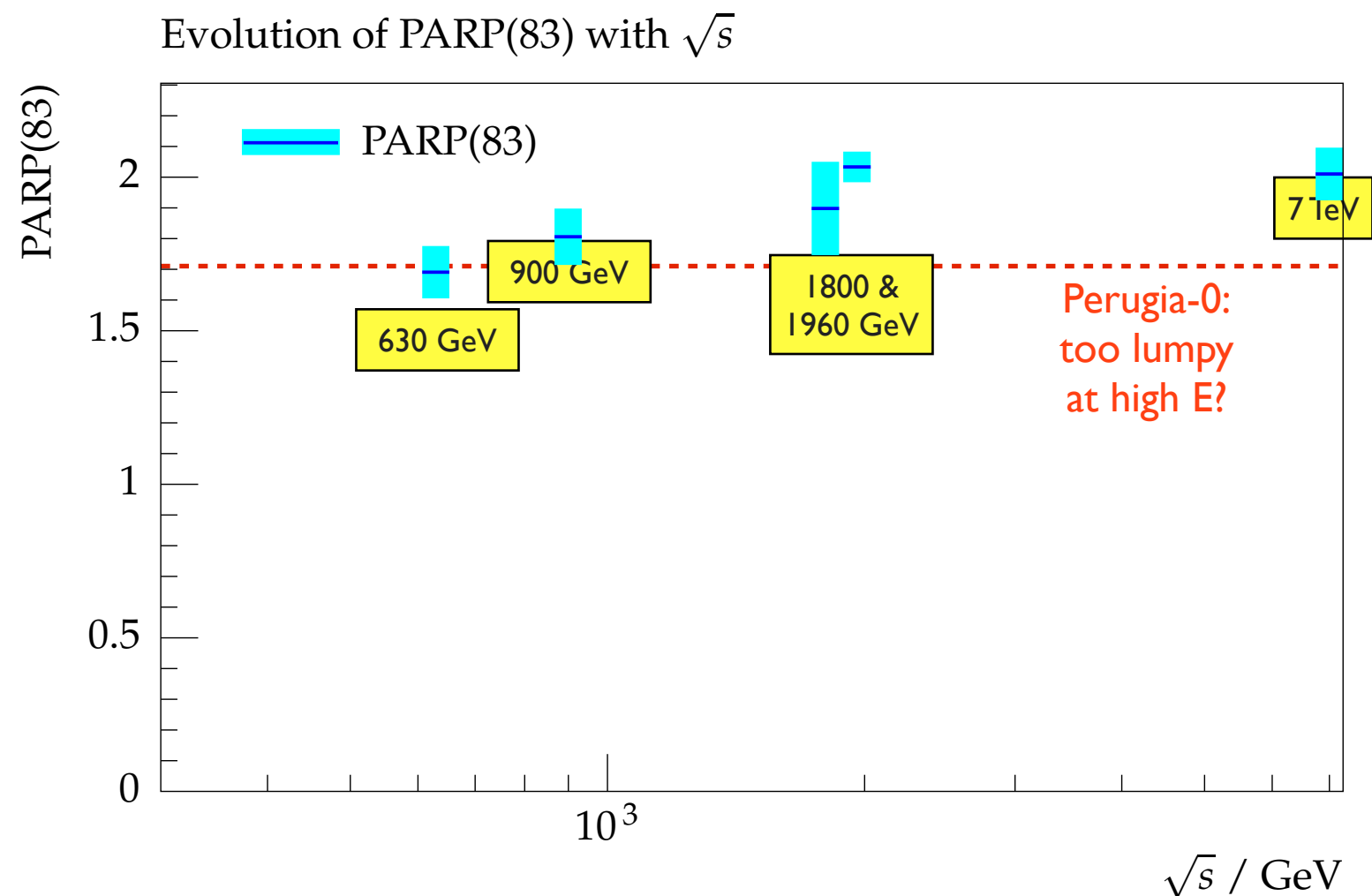


“Energy Scaling of MB Tunes”, H. Schulz + PS, in preparation

# Mass Distribution

## Independent tunings compared to Perugia 0

Hint of departure from Gaussian (PARP(83)=2.0) at lower energies? Consistent with higher  $x \rightarrow$  more lumpy?



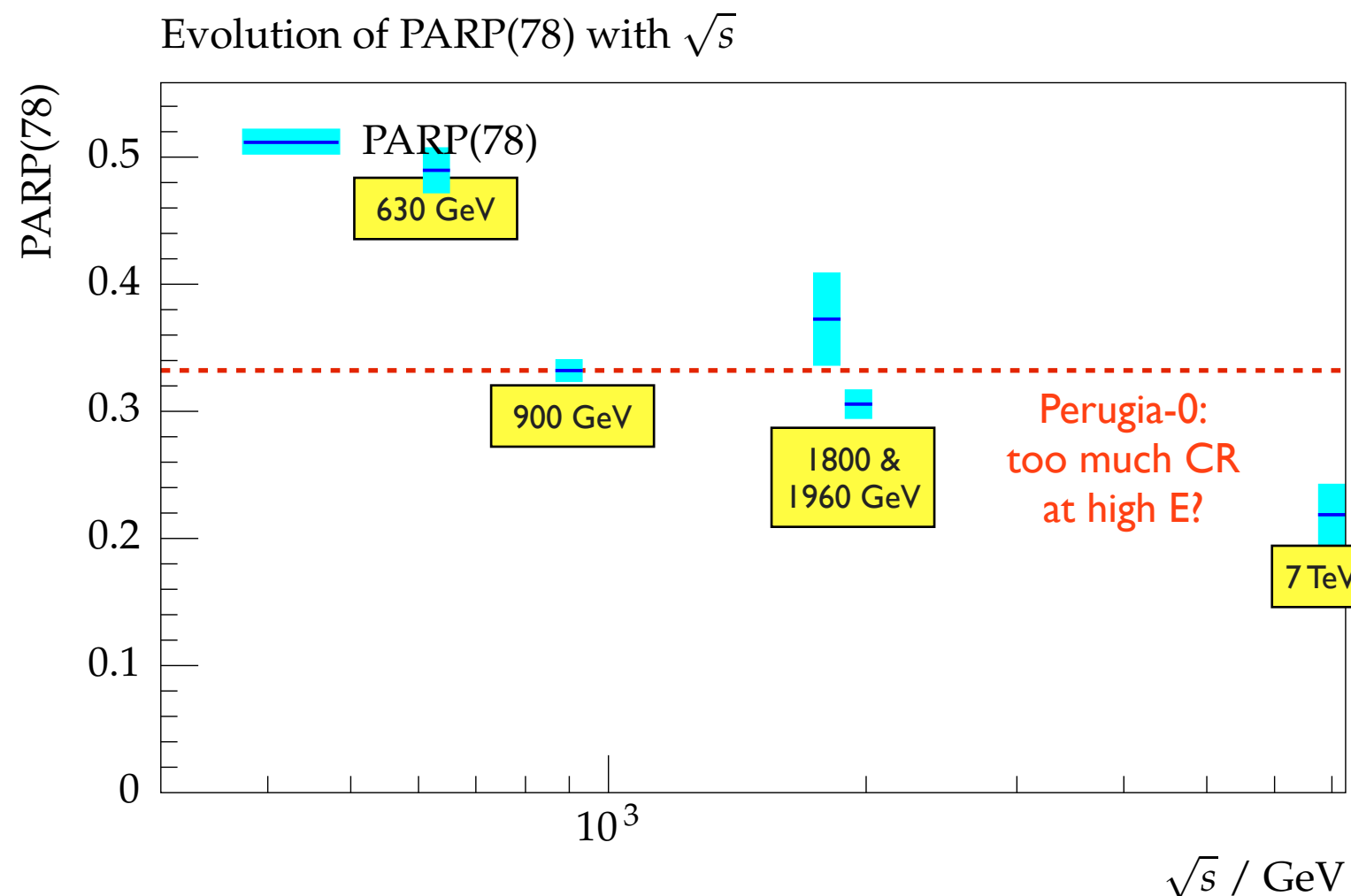
“Energy Scaling of MB Tunes”, H. Schulz + PS, in preparation

# Color Reconnections

## Independent tunings compared to Perugia 0

CR are the most poorly understood part of these models

Assumption of constant strength not supported by data!



“Energy Scaling of MB Tunes”, H. Schulz + PS, in preparation

# PYTHIA Updates



with input from R. Corke, T. Sjöstrand

# PYTHIA 6

## The Perugia Tunes

PS, arXiv:1005.3457v2

Intended to provide reasonable starting points for tuning efforts of the  $p_T$ -ordered framework

Mark the last development effort from the authors

## Diffraction

Obsolete Model: no diffractive jet production

→ PYTHIA 8: S. Navin, arXiv:1005.3894

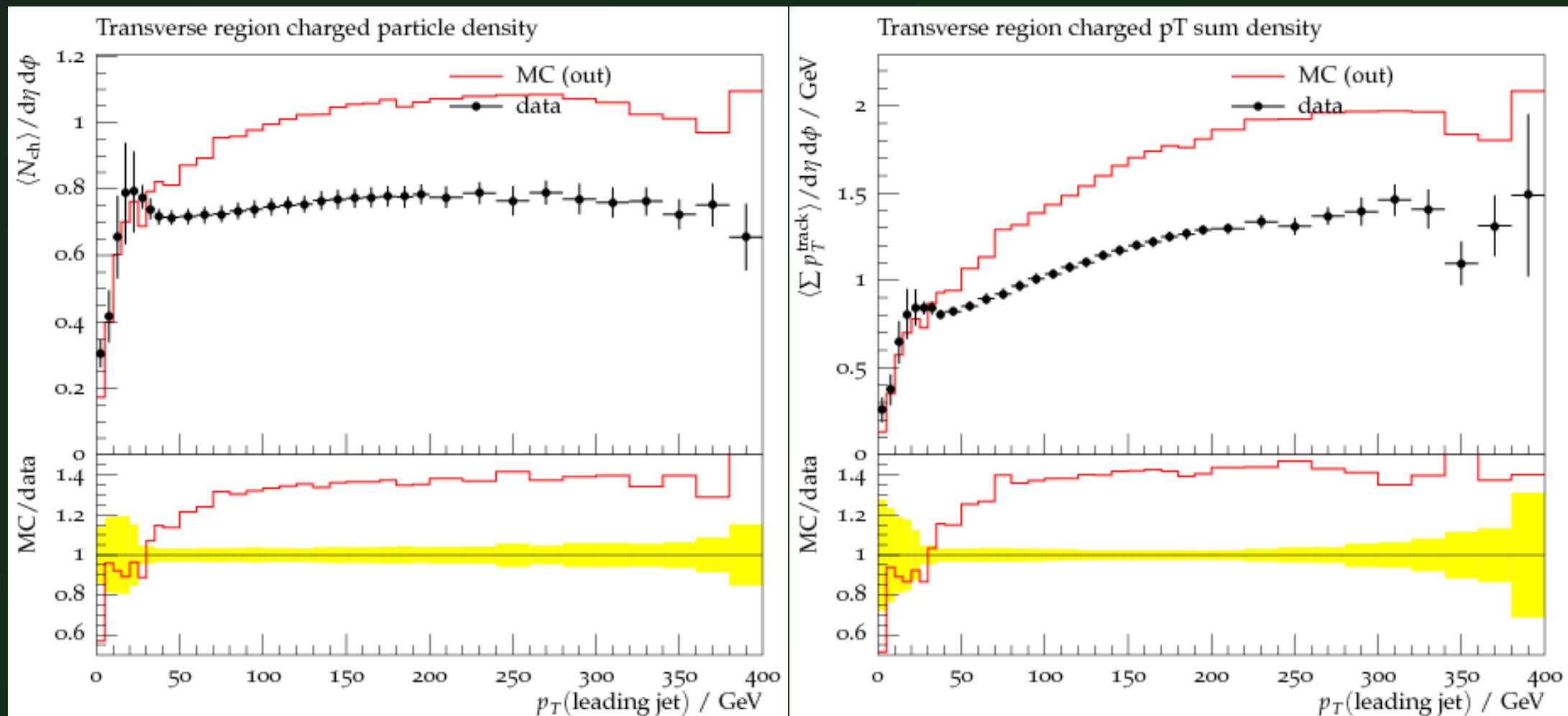
## Status

No longer actively developed

# PYTHIA 8

**Already significant improvements  
but there was one snag...**

cf., e.g., yesterday's  
ATLAS talk (L. Tompkins)

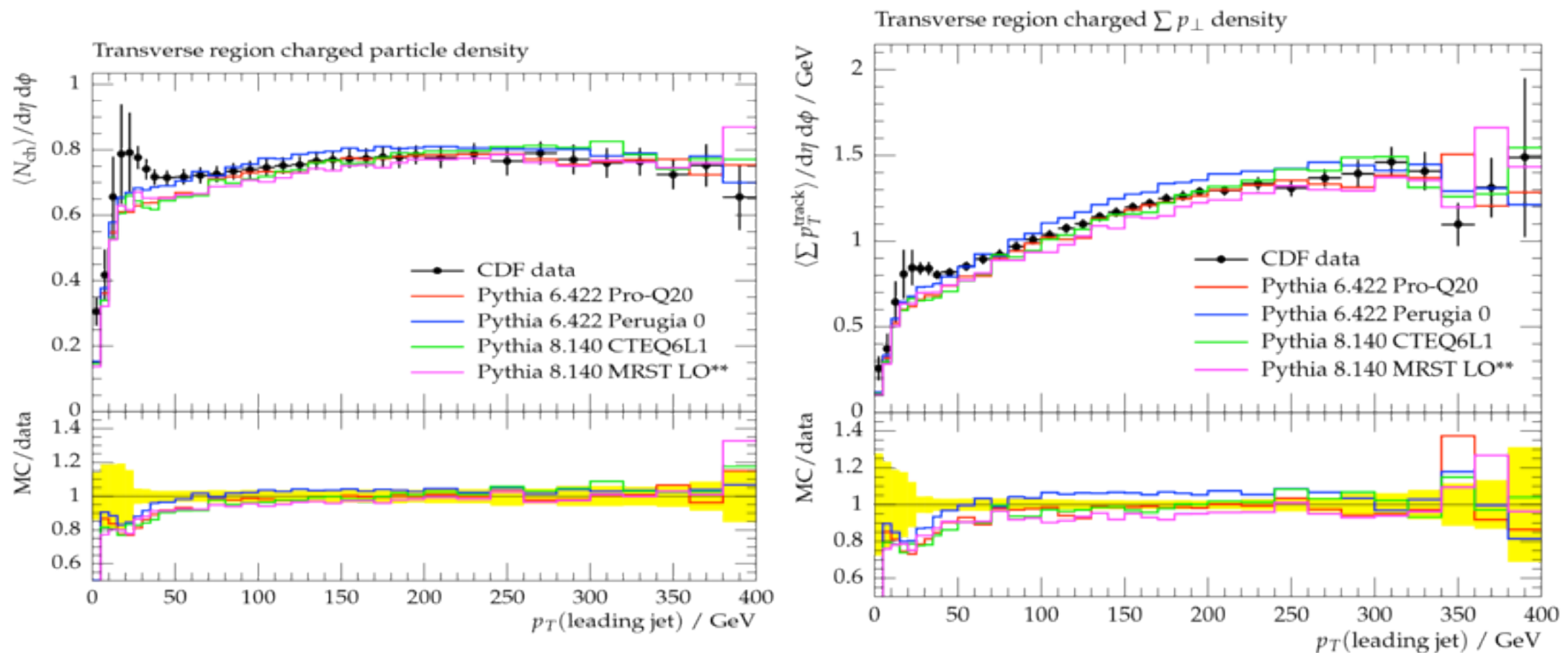


**Where did we go wrong?**



# PYTHIA 8

**A problem with Initial-Final Dipoles  
(missing coherence), now addressed →**



**PYTHIA 8 now ready to replace PYTHIA 6 also for UE**

# Summary

## A new way of using tuning tools

→ Check of consistency and universality of the model

*Not just the best tune*

**Power + Flexibility** of automated tools allow  
*independent optimizations in complementary phase space regions*

We used different beam energies as our complementary regions  
(→ tests of energy scaling assumptions)  
Other complementary sets could be used to test other aspects

Crucial: Need complete and comparable data sets in each region!

+ *get a data-driven idea of any non-universalities as a bonus → better uncertainties*

**+ Time to move to PYTHIA 8**





# Backup Slides

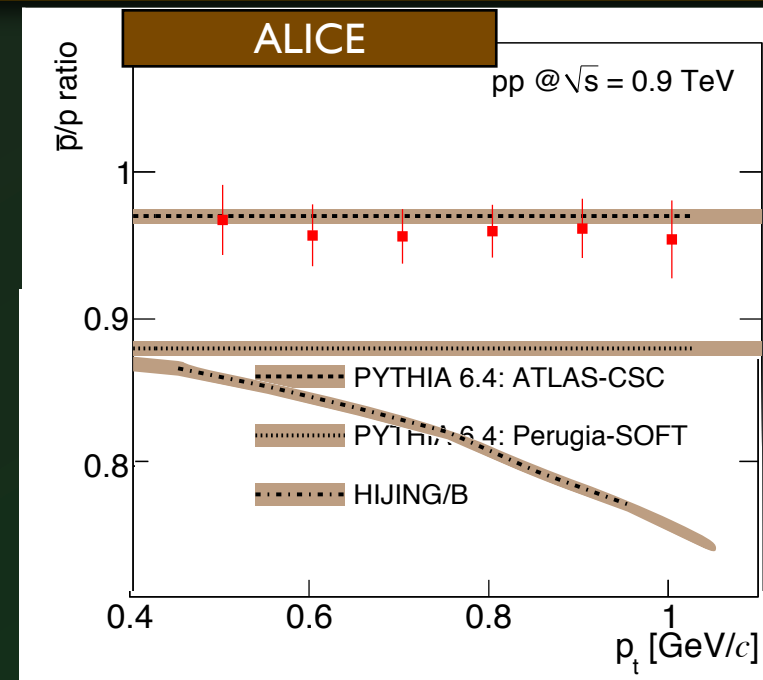
# Baryon Transport

**LESS than  
Perugia-SOFT**

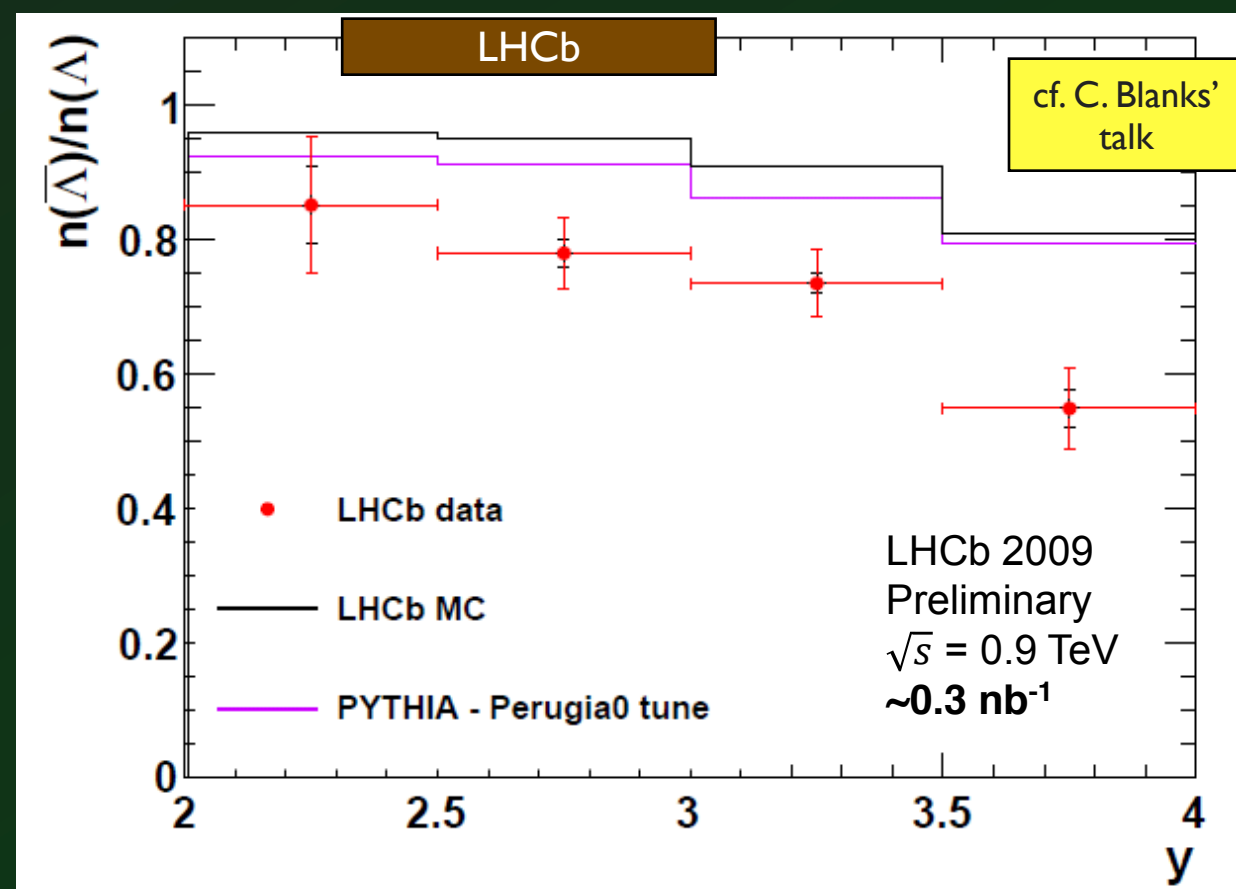
(at least for  
protons, in central  
region)

**But MORE  
than Perugia-0**

(at least for  
Lambdas, in  
forward region)



cf. J. Fiete's  
talk



cf. C. Blanks'  
talk