

Higher Order Corrections to the Drell-Yan Cross Section in the Mellin Space

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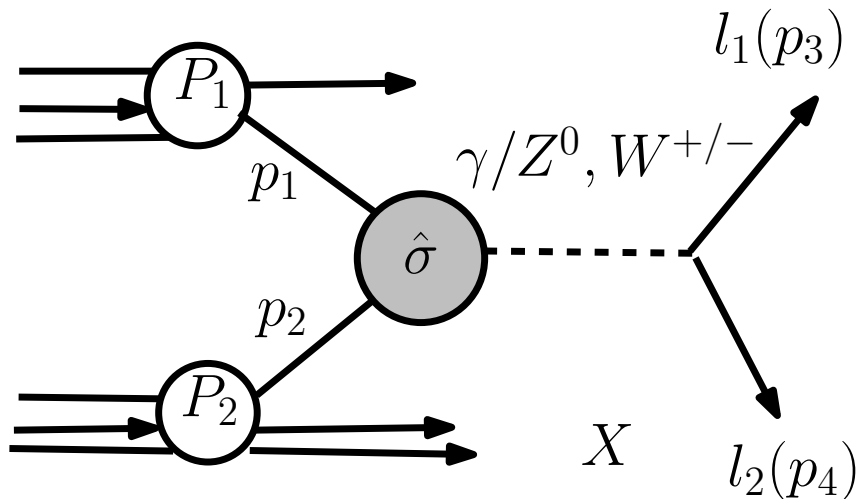
Outline

- The Drell-Yan mechanism
- The Mellin transform
- Calculation of the Drell-Yan structure functions in Mellin space up to NNLO
 - Coefficient functions
 - Parton distribution functions
- Partial results and checks
- Conclusions and outlook

The Drell-Yan Process

Massive lepton pair production in hadron-hadron collision, $M_{l_1 l_2}^2 \gg 1 \text{ GeV}^2$

[Drell, Yan 1970]



- $M_{l_1 l_2} = Q^2 = (p_2 + p_3)^2$

- CM energy of hadrons
 $s = (P_1 + P_2)^2$

- $p_i = x_i P_i \quad x_i \in (0, 1)$

- CM energy of partons
 $\hat{s} = (p_1 + p_2)^2 = s x_1 x_2$

- Neutral Current

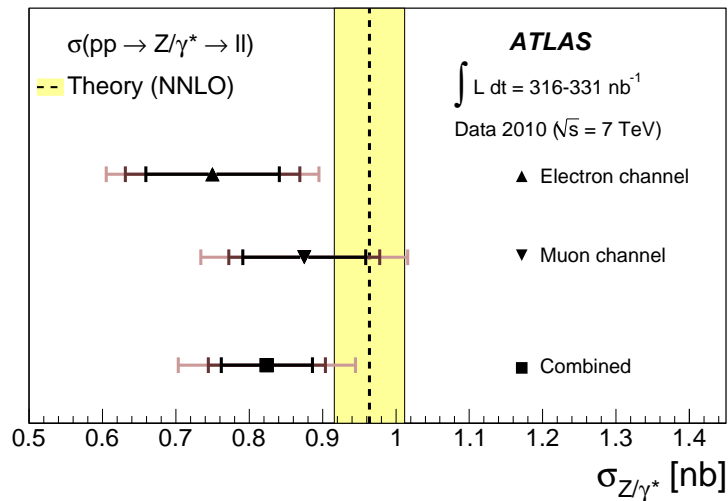
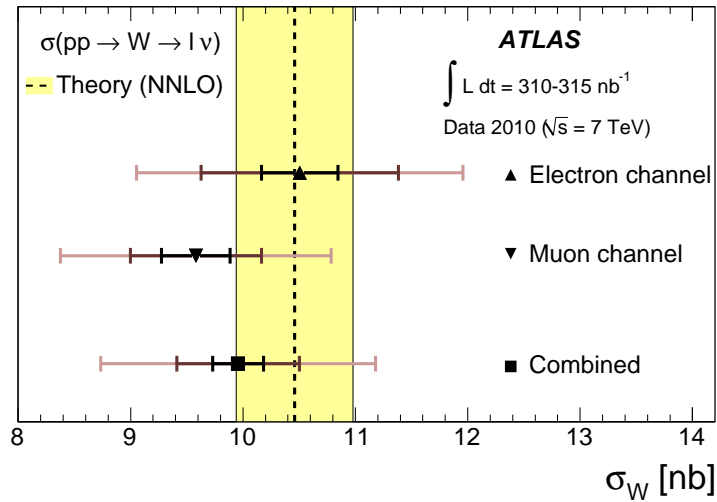
- $pp \rightarrow \gamma^* \rightarrow l\bar{l}X \quad M_{l\bar{l}} \ll M_Z \quad M_Z \sim 91.2 \text{ GeV}$

- $pp \rightarrow Z^0 \rightarrow l\bar{l}X \quad M_{l\bar{l}} \sim M_Z$

- Charged Current

- $pp \rightarrow W^\pm \rightarrow l\nu X \quad M_{l\bar{l}} \sim M_W \quad M_W \sim 80.4 \text{ GeV}$

The Drell-Yan process at the LHC



- Large total cross sections for W and Z production
- Clear signal
- Background for new physics measurements
- Test of the Standard Model
- Constraints on parton distribution functions

The Drell-Yan Cross Section

$$\frac{d\sigma_{DY}^V(Q^2)}{dQ^2} = \sum_{a,b} \int_0^1 dx_1 \int_0^1 dx_2 \underbrace{f_a(x_1, Q^2) f_b(x_2, Q^2)}_{\text{PDFs}} \underbrace{\frac{d\hat{\sigma}_{ab}^V(Q^2)}{dQ^2}}_{\text{Hard Cross section}}$$

$$V = \gamma^*, Z^0, W^\pm \quad a, b = q, \bar{q}, g \quad x_{1,2} \in (0, 1)$$

- Higher order NLO QCD corrections [Altarelli, Ellis, Martinelli 1979]
- NNLO [Hamberg, Matsuura, van Neerven 1990; Harlander, Kilgore 2002]
- NNLO fully differential cross section [Anastasiou, Dixon, Melnikov, Petriello 2004; Catani, Ferrera, Grazzini 2010]
⇒ new constraints on PDFs [Alekhin, Melnikov, Petriello 2006]
- Electroweak corrections up to NLO [Hollik, Wackerroth 1996; Vicini et.al. 2009]

The Drell-Yan Cross Section

$$\begin{aligned} \frac{d\sigma_{DY}^V(Q^2)}{dQ^2} &= \sum_{a,b} \int_0^1 dx_1 \int_0^1 dx_2 f_a(x_1, Q^2) f_b(x_2, Q^2) \frac{d\hat{\sigma}_{ab}^V(z, \alpha_s)}{dQ^2} \\ &= x N^V \underbrace{\sum_{a,b} C_{ab}^V(f_a \otimes f_b \otimes \Delta_{ab})(x)}_{\text{Structure function } W_{DY}(x, Q^2)} ; \quad x = \frac{Q^2}{s} \quad z = \frac{x}{x_1 x_2} \end{aligned}$$

$$\Delta_{ab} = \sum_k \left(\frac{\alpha_s}{4\pi} \right)^k \Delta_{ab}^{(k)} \quad \text{Coefficient functions}$$

● Convolution

$$\begin{aligned} &(f_1 \otimes f_2 \otimes \dots \otimes f_k)(x) \\ &= \int_0^1 dx_1 \int_0^1 dx_2 \dots \int_0^1 dx_k \delta(x - x_1 x_2 \dots x_k) f_1(x_1) f_2(x_2) \dots f(x_k) \end{aligned}$$

The Convolution and The Mellin Transform

- The Mellin transform - integral transform

$f(x)$ (“ x -space”) $\rightarrow \tilde{f}(N)$ (“ N -space”)

$$\mathbf{M}[f(x)] = \int_0^1 dx x^{N-1} f(x) = \tilde{f}(N)$$

- Mellin transform of the convolution is a product of functions in N -space

$$\mathbf{M}[(f_1 \otimes f_2 \otimes \cdots \otimes f_k)(x)](N) = f_1(N) f_2(N) \cdots f_k(N)$$

The Convolution and The Mellin Transform

- The structure function

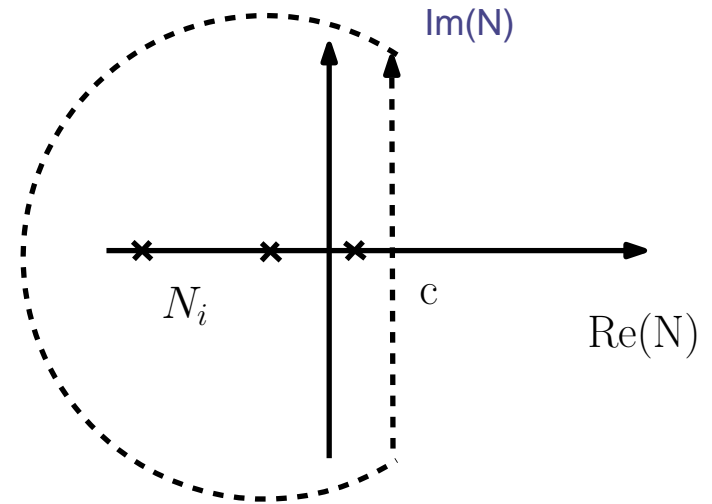
$$W(x, Q^2) = (f_a \otimes f_b \otimes \Delta_{ab}^V)(x) \quad \rightarrow \quad f_a(N) f_b(N) \Delta_{ab}^V(N) = \tilde{W}(N, Q^2)$$

- Two integrations \rightarrow simple product
- Need analytic expressions for all input functions in x -space for which the Mellin transforms exist [Vermaseren 1998]
[Moch, Vermaseren, Vogt 2004; Blümlein, Kurth 1998 (2000); Blümlein, Ravindran 2005, ...]
- Original x -space recovered performing the inverse Mellin transform

The Inverse Mellin Transform

- The inverse Mellin Transform - integral over a complex plane

$$W(\tau) = \frac{1}{2\pi i} \int_{c-i\infty}^{c+i\infty} dN x^{-N} \tilde{W}(N)$$



- All poles of $\tilde{W}(N)$ lie to the left from the contour c
- In general inverse MT are performed numerically

The Inverse MT - numerical evaluation

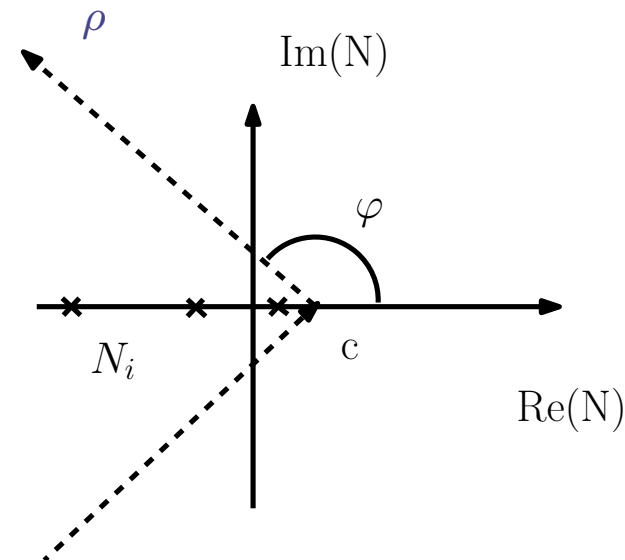
Rewriting the integral

$$\frac{1}{2\pi i} \int_{c-i\infty}^{c+i\infty} dN x^{-N} \tilde{f}(N) = \frac{1}{\pi} \int_0^{\infty} d\rho \operatorname{Im}[e^{i\varphi} x^{-c-\rho e^{i\varphi}} f(N)]$$

→ integral over a real variable ρ

$N = c + \rho e^{i\varphi}$ becomes complex variable

- Integration only up to ρ_{\max} due to the factor $\exp[\rho \ln(1/x) \cos \varphi]$
- φ , c and number of N 's can be tuned to get better accuracy
- Need analytic continuations of all N -space functions to the complex plane



Coefficient functions in N -space

- Starting from x -space expressions [Ravindran, unpublished]
Mellin transforms obtained using `harmopol` package [Remiddi, Vermaseren 2000]
- N -space results (NNLO) have been published [Blümlein, Ravindran 2005]
- \rightarrow N -space results mostly in terms of harmonic sums $S_k(N)$
with “most right” pole at $N = 1$.
- Analytic continuations of harmonic sums in terms of polygamma functions
- Finite set of more complicated harmonic sums needs special treatment in
order to get the analytic continuations, e.g. `ancont` [Blümlein, 2000]

Evolution of Parton Distribution Functions

- Scale dependence of $f_a(x, Q^2)$ described by evolution equations (DGLAP)

$$\frac{\partial f_a(x, Q^2)}{\partial \ln Q^2} = \sum_b \underbrace{P_{ab}(x, Q^2)}_{\text{Splitting functions}} \otimes f_b(x, Q^2)$$

- The solution

$$f_a(x, Q^2) = E \left[P_{ab}(x), \alpha_s(Q^2), \underbrace{f_b(x, Q_0^2)}_{Q_0^2 \sim \text{few GeV}} \right]$$

Parton Distribution Functions in N -space

- Evolution in N -space

- Choose initial distribution $f(x, Q_0^2)$, typical parametrization

$$xf(x, \mu_0^2) = ax^b(1-x)^c(1+dx^f+gx), \quad a, \dots, g \in \mathbb{R}$$

$$f(N, Q_0^2) = a \left[\beta(N+b-1, c+1) + \beta(N+b+f-1, c+1) + \beta(N+b, c+1) \right]$$

- Splitting functions in N -space

[Moch, Vermaseren, Vogt 2004]

- Evolve PDFs in N -space

[QCD-PEGASUS, Vogt 2004]

$$f_a(N, Q^2) = E \left[P_{ab}(N), \alpha_s(Q^2), f_b(N, Q_0^2) \right]$$

- Evolution in x -space (use LHAPDF) \rightarrow find an interpolation function for $f(x, Q^2)$ and transform it to N -space.

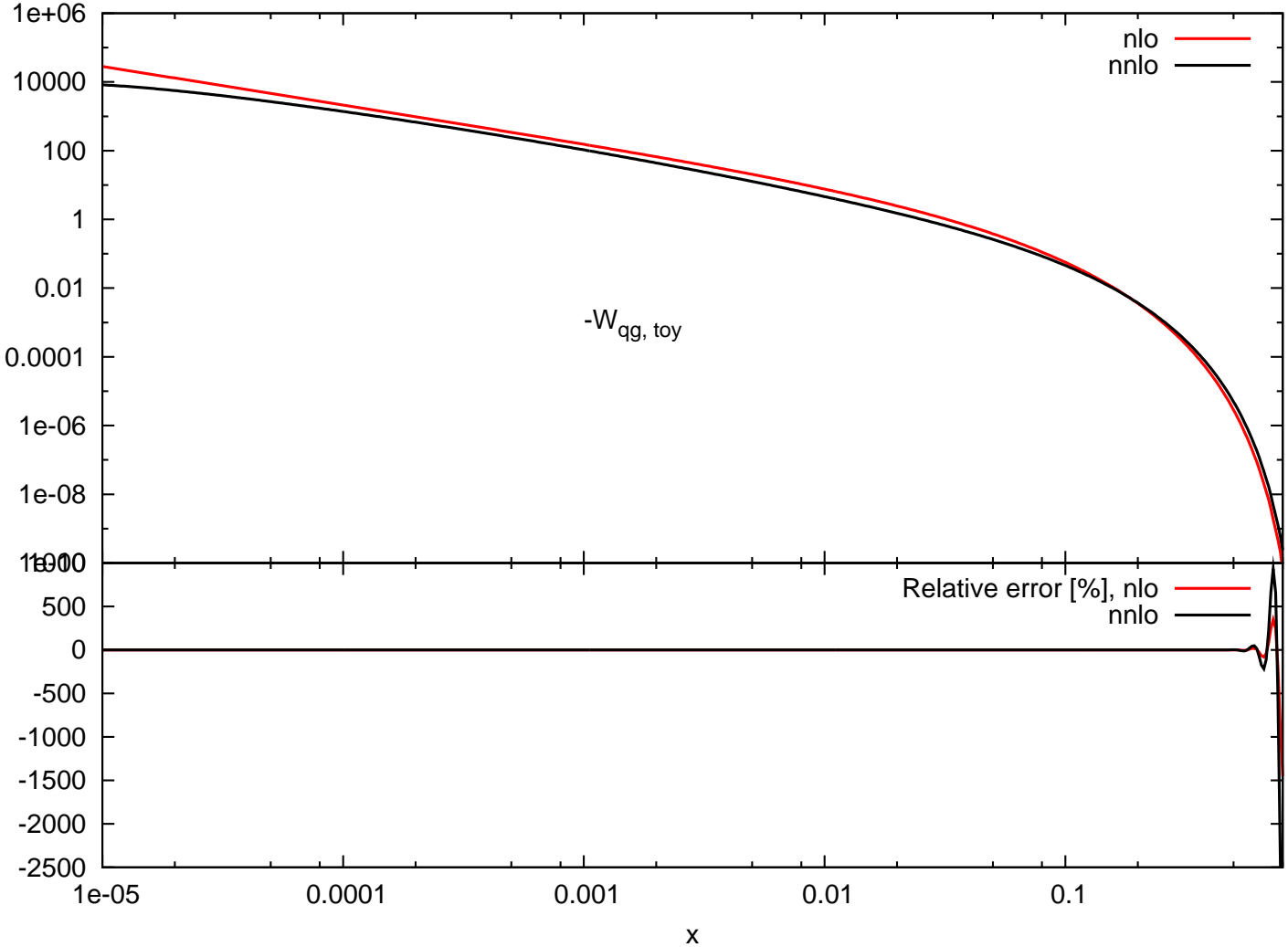
The Calculation - Indirect Check

- Choose a toy initial parton distribution function, perform evolution in N -space with QCD-PEGASUS $\rightarrow \tilde{\sigma}_{\text{toy}}(N, Q^2)$
- Choose the same initial distribution in x -space and do a standard convolution $\rightarrow \sigma_{\text{toy}}(x, Q^2)$ [CUBA, Hahn 2005]
- Compare $\mathbf{M}[\tilde{\sigma}_{\text{toy}}(N, Q^2)] \Leftrightarrow \sigma_{\text{toy}}(x, Q^2)$.
- Replace toy pdfs in x -space with real ones (LHAPDF)
 $\sigma_{\text{toy}}(x, Q^2) \rightarrow \sigma_{\text{LHAPDF}}(x, Q^2)$ and compare to publicly available codes
 - MCFM (NLO) [Campbell, Ellis]
 - ZWPROD [Hamberg, Matsuura, van Neerven, 2002]
 - DYNNLO [Catani, Cieri, Ferrera, de Florian, Grazzini 2009; Catani, Grazzini 2007]
 - FEWZ [Melnikov, Petriello]

The Calculation - Direct Check

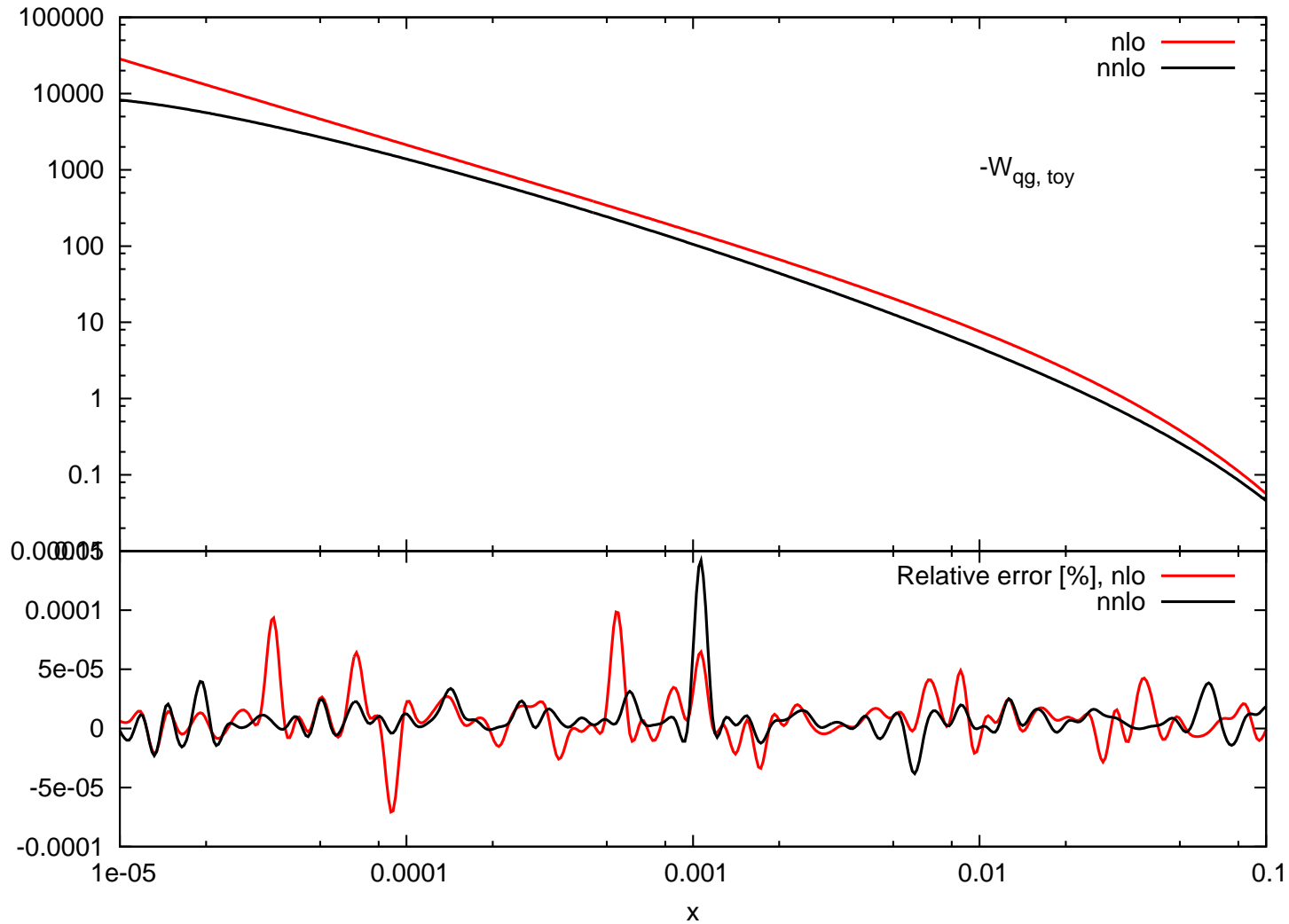
- Choose a real initial parton distribution function \rightarrow Mellin transform, perform evolution in N -space with QCD-PEGASUS and compare $\mathbf{IM}[\sigma(N, Q^2)]$ to cross section calculated by publicly available codes
- Needs to extend slightly QCD-PEGASUS
- Not possible for any distributions in LHAPDF - some initial parametrizations are not in form that Mellin transform on them can be performed
- Alternatively, access directly evolved PDFs from LHAPDF grids \rightarrow interpolation \rightarrow Mellin transform

QG contribution to the DY structure function



$$W^{n^k lo}(N)_{qg, toy} = \left(\frac{\alpha_s}{4\pi}\right)^k C_{qg}(N) f_{q, toy}(N) f_{g, toy}(N), \quad W_{qg, toy} = \text{IM}[W_{qg, toy}(N)]$$

QG contribution to the DY structure function



$$W^{n^k lo}(N)_{qg, toy} = \left(\frac{\alpha_s}{4\pi}\right)^k C_{qg}(N) f_{q, toy}(N) f_{g, toy}(N), \quad W_{qg, toy} = \text{IM}[W_{qg, toy}(N)]$$

Summary and Outlook

- Higher order corrections for Drell-Yan process are necessary for precise predictions
- Calculation in Mellin space is fast and accurate

- NNLO cross section
- Interface for PDFs accessing directly LHAPDF grids
- Implementation of DIS
- Release the code

- Analytic continuation of a Harmonic Sum

$$S_k(N) = \frac{(-1)^{k+1}}{(k-1)!} \psi^{k-1}(N+1) + \zeta(k)$$

- The Euler Beta function

$$\beta(a, b) = \frac{\Gamma(p)\Gamma(q)}{\Gamma(p+q)}$$