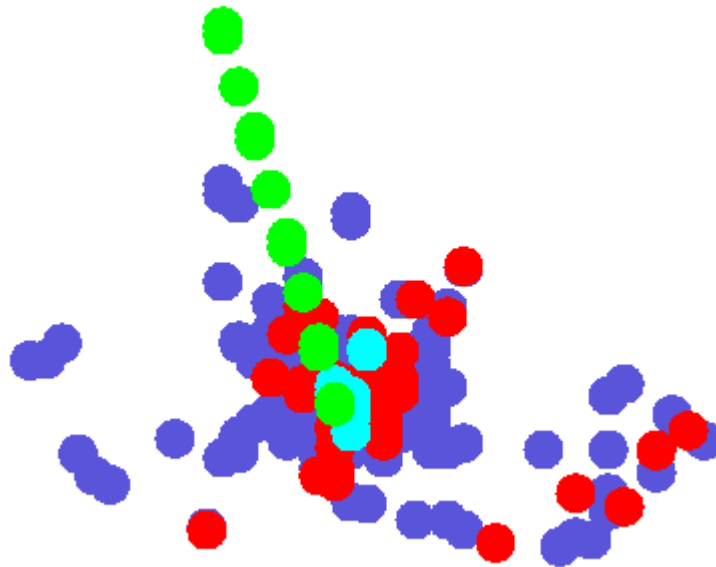


# Searching for Atmospheric Neutrino Oscillations at MINOS



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***Cambridge University***

***April 2004***

# The MINOS Experiment

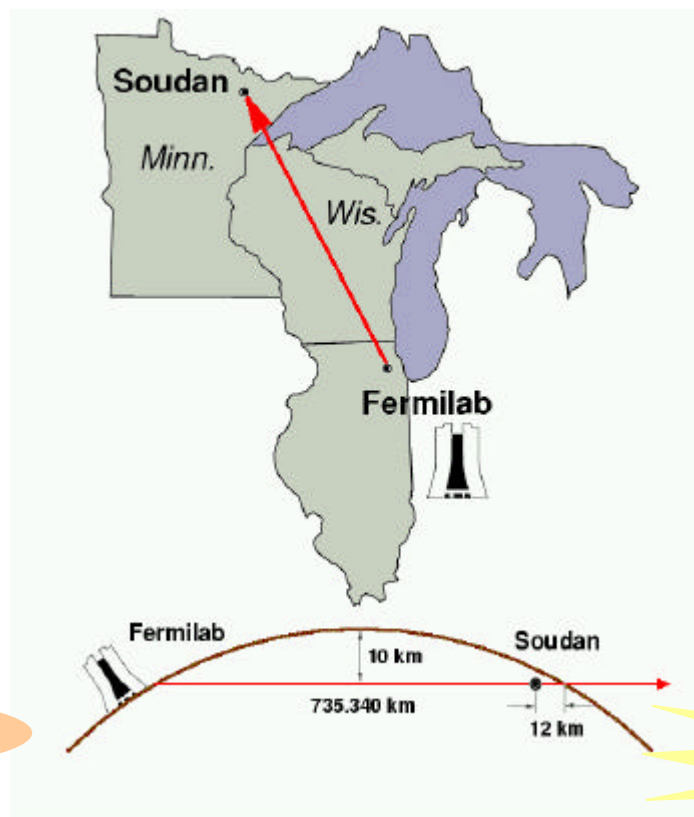
*Main Injector Neutrino Oscillation Search*

Fermi Laboratory



Near Detector (2004)

Neutrino Beam (2005)



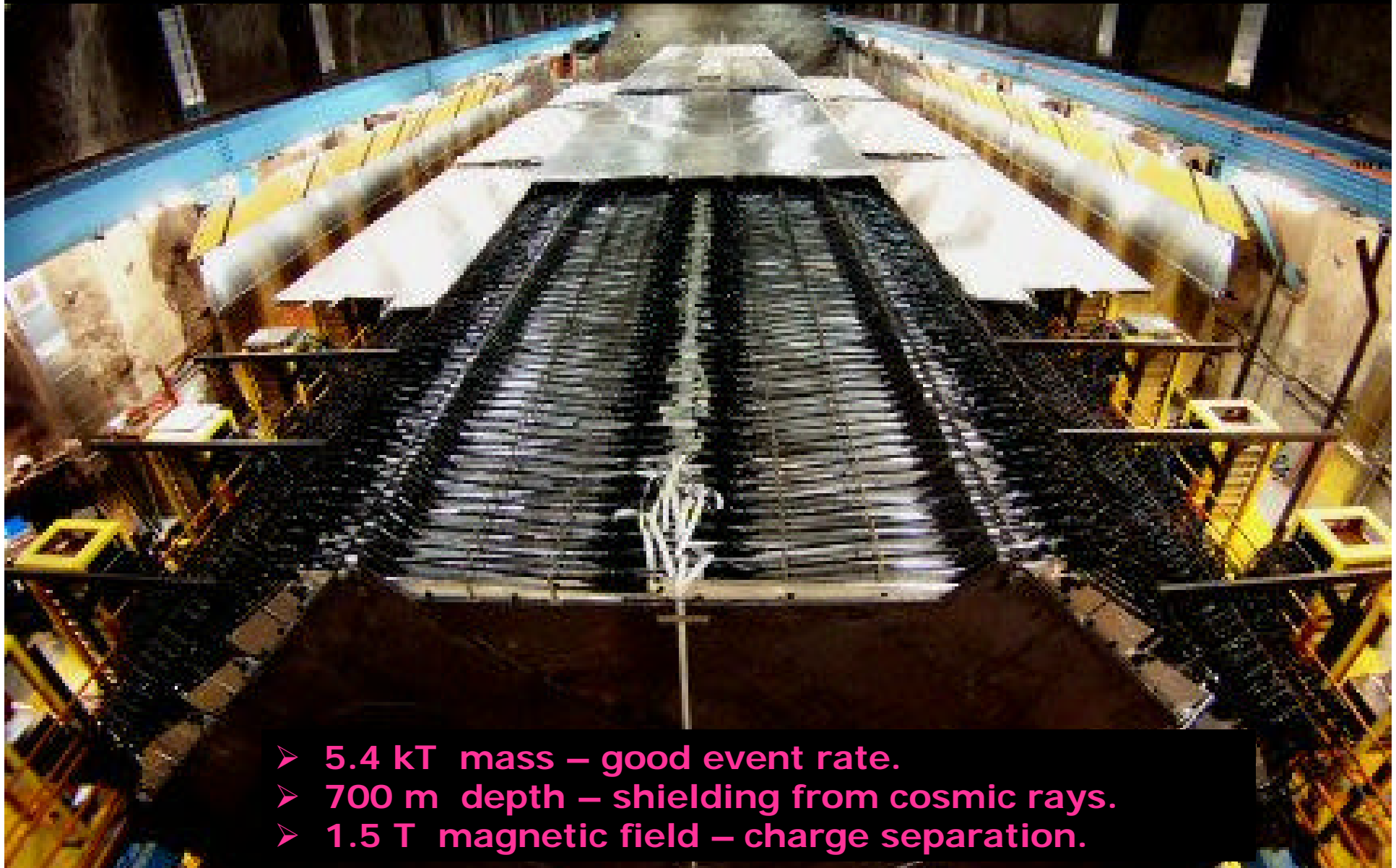
Soudan Mine



Far Detector

COLLECTING DATA!

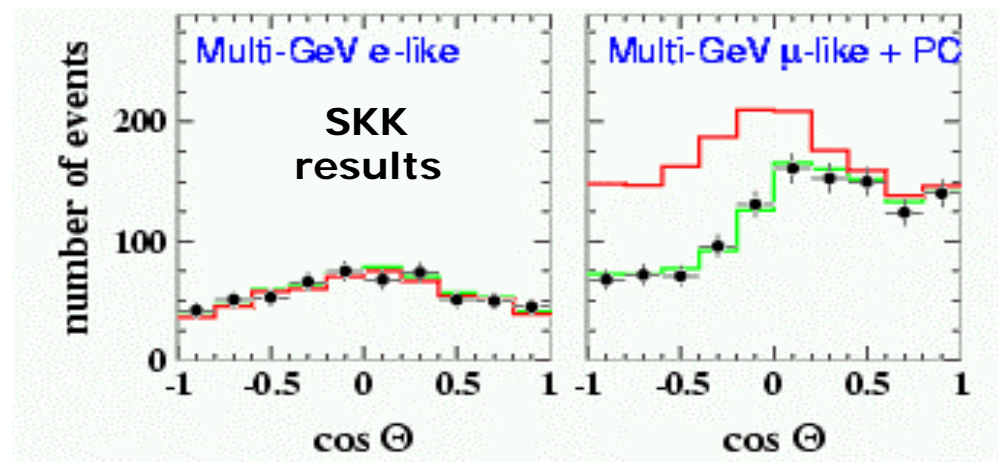
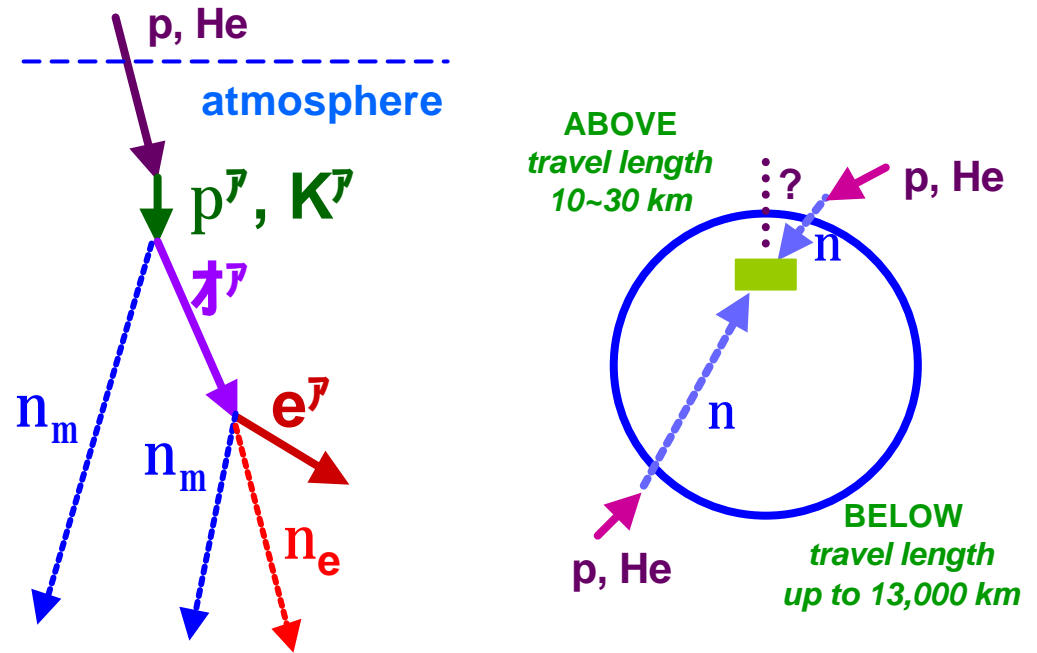
## MINOS Far Detector will measure atmospheric neutrino oscillations



- 5.4 kT mass – good event rate.
- 700 m depth – shielding from cosmic rays.
- 1.5 T magnetic field – charge separation.

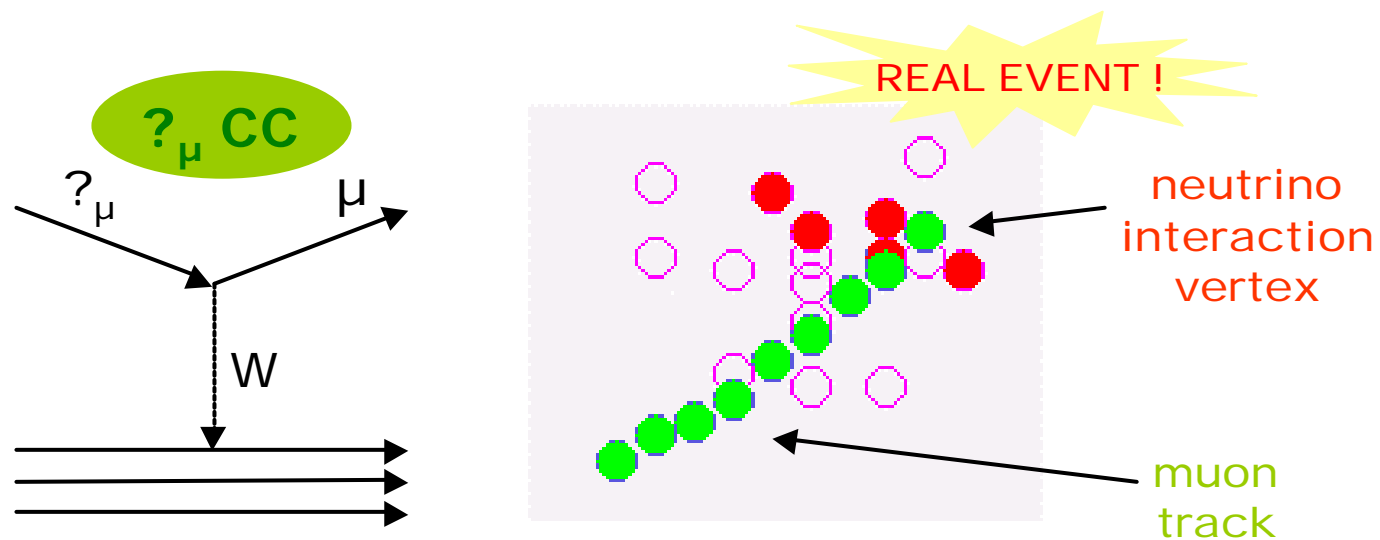
# Atmospheric Neutrino Oscillations

- Compelling evidence for  $\nu_m \leftrightarrow \nu_t$  oscillations – lots of experiments have measured up/down asymmetry in atmospheric  $\nu_m$  flux.
- MINOS magnetic field can distinguish  $\mu^-$  from  $\mu^+$ , and thus  $\nu_m$  from  $\bar{\nu}_m$ .
- MINOS will make separate measurements of  $\nu_m \leftrightarrow \nu_t$  and  $\bar{\nu}_m \leftrightarrow \bar{\nu}_t$  oscillations in atmospheric neutrinos.



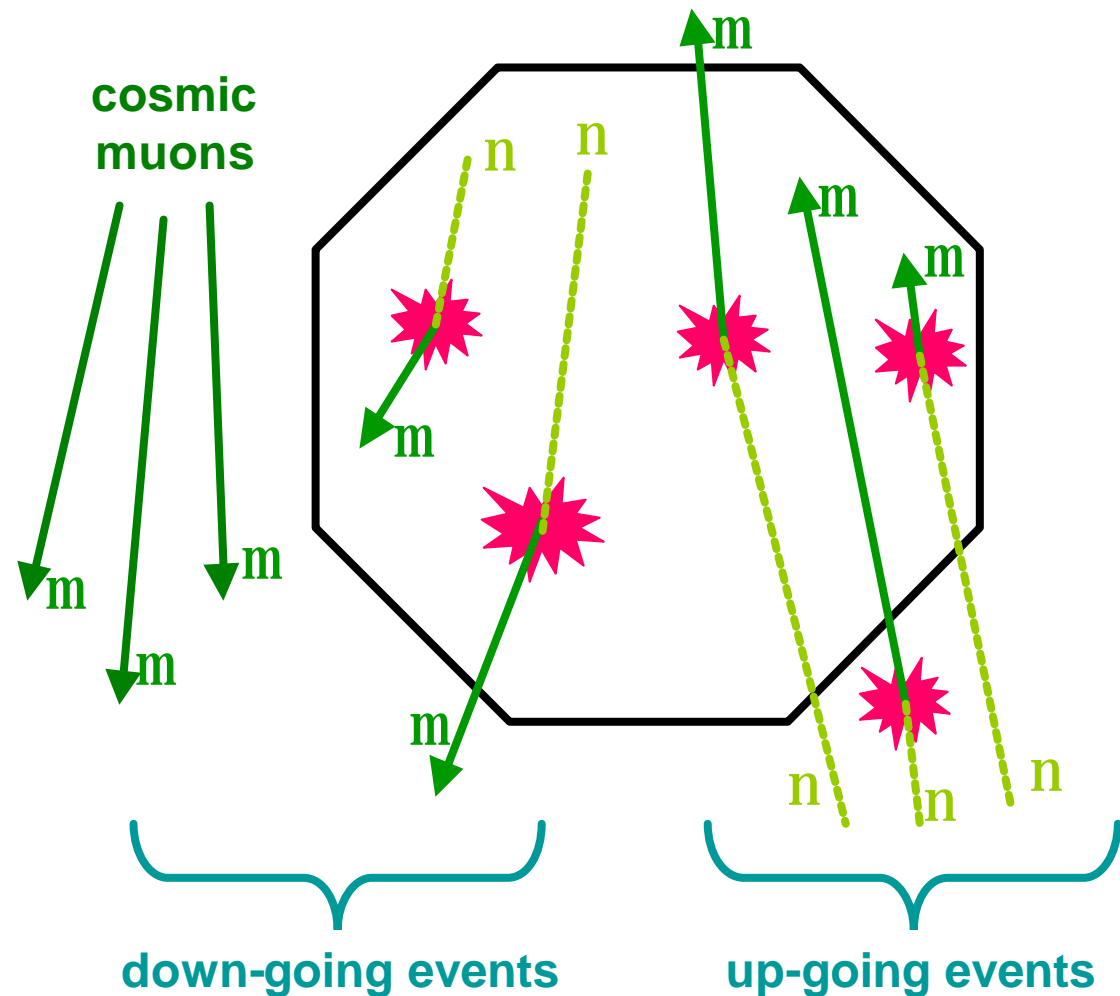
# Detecting Atmospheric Neutrinos

- **The MINOS detectors are sampling calorimeters.**
  - Far Detector comprises ~500 interleaved planes of inch-thick magnetized steel and plastic scintillator.
- **Need clean sample of  $n_m$  CC events to measure oscillations.**
  - select events with muon track and contained interaction vertex.
- **Cambridge is developing event reconstruction software.**
  - reconstruct particle tracks + showers.
  - measure particle direction + charge + energy.



# Detecting Atmospheric Neutrinos

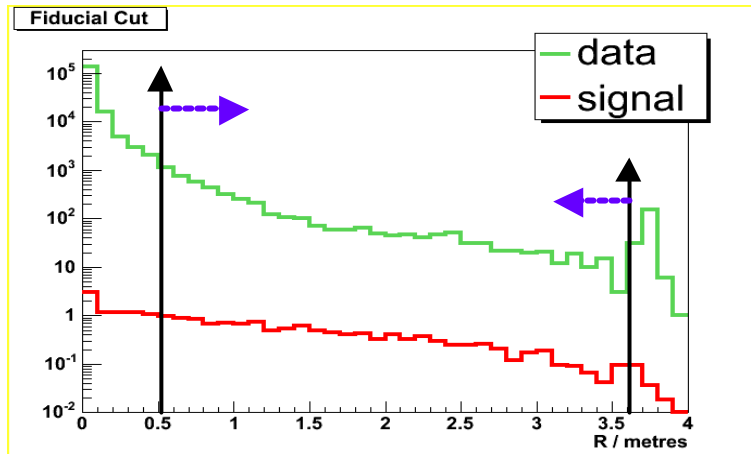
- **Neutrino event signatures:**
  - contained interaction vertex.
  - upward-going muons.
- **Dominant background from cosmic muons**
  - sneak between detector planes, appear contained ( ? containment cuts )
  - mis-reconstructed as upward-going muons ( ? direction cuts )
- **Need to achieve  $\sim 10^5$  cosmic muon rejection to separate neutrino signal.**



# Contained Events

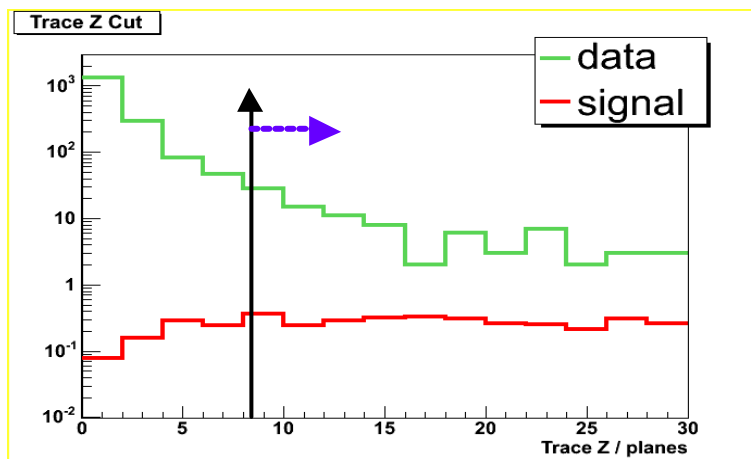
## (1) Fiducial Cuts

- consider top track vertex
- must be  $>50$  cm from detector edge



## (2) Trace Cuts

- trace track back to detector edge
- must cross  $>50$  cm steel



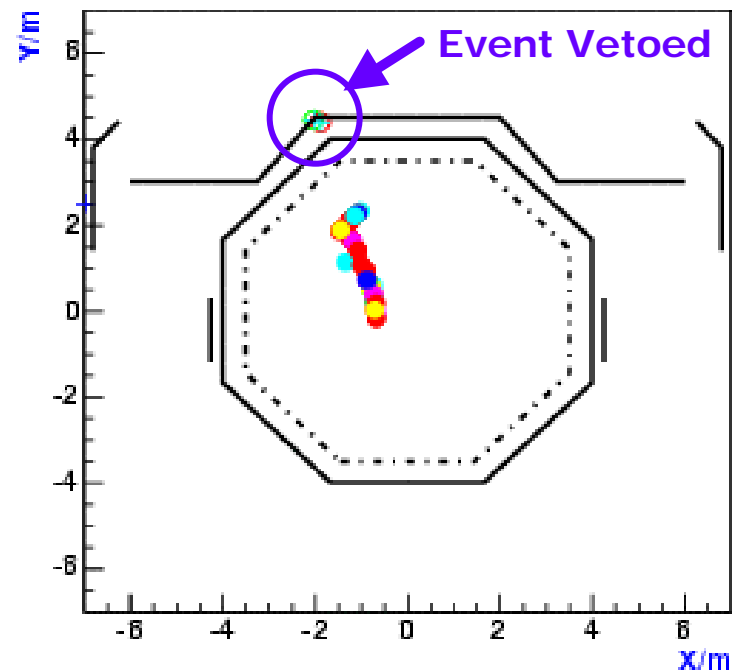
## (3) Topology Cuts

- Cut on steep muon topologies
- 1:1 signal-to-background achieved.

## (4) Veto Shield Cuts

- planes of scintillator positioned above detector to tag cosmic muons
- $>95\%$  rejection rate.

**Event X/Y View**



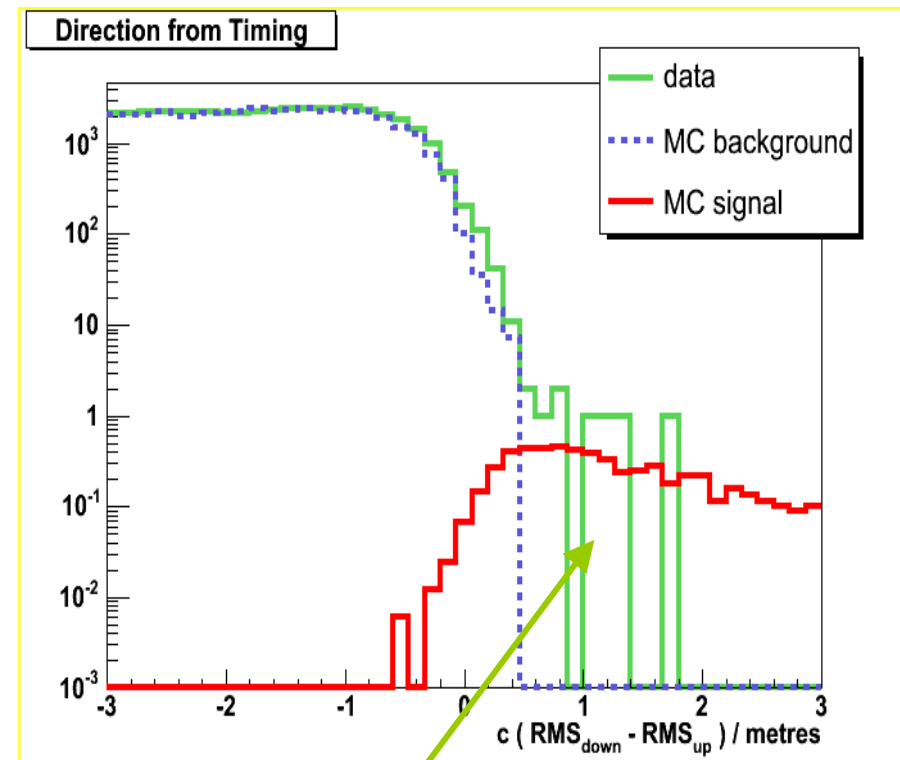
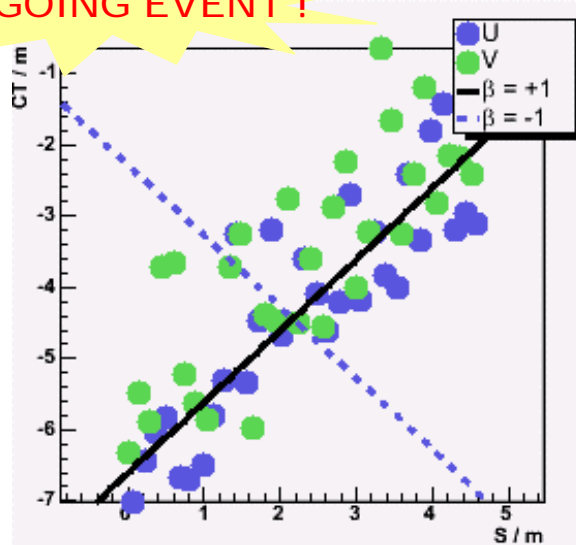
# Up-Going Events

MINOS Far Detector Timing Resolution  $\sim 2.5$  ns

## • Direction-Finding Algorithm:

- consider distance vs time for track
- force fits with  $\beta = \pm 1$
- calculate RMS about each fit
- $\text{RMS}_{\text{down}} - \text{RMS}_{\text{up}} > 0$  for up-going tracks.

UP-GOING EVENT !



up-going  
neutrinos



# Charge Reconstruction

MINOS Far Detector Magnetic Field ~ 1.5 T

- Charge-Finding Algorithm:

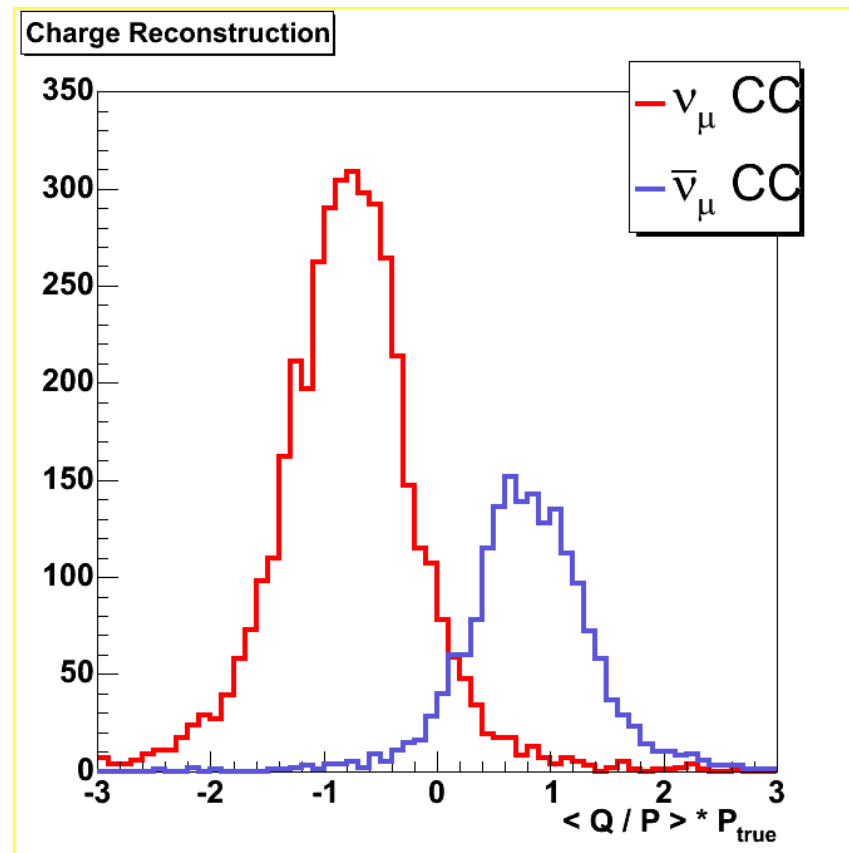
changing momentum given by:

$$\frac{d\mathbf{p}}{ds} = 0.3Q\mathbf{p} \times \mathbf{B} + \hat{\mathbf{p}} \frac{dp_{\text{loss}}}{ds}$$

$$\frac{Q}{p} = \frac{\frac{d\hat{\mathbf{p}}}{ds} \times \hat{\mathbf{p}} \times \mathbf{B}}{0.3 |\hat{\mathbf{p}} \times \mathbf{B}|^2}$$

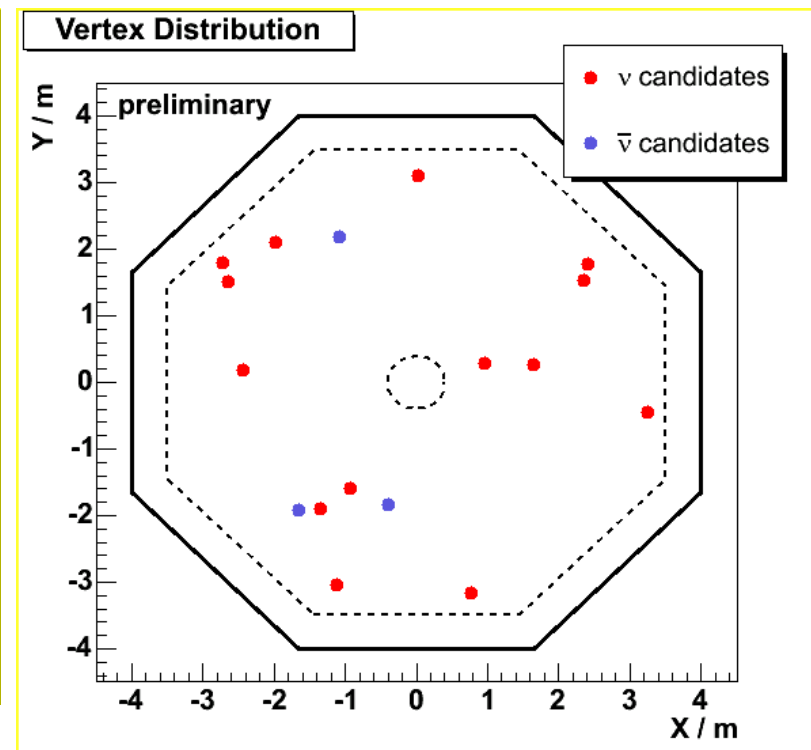
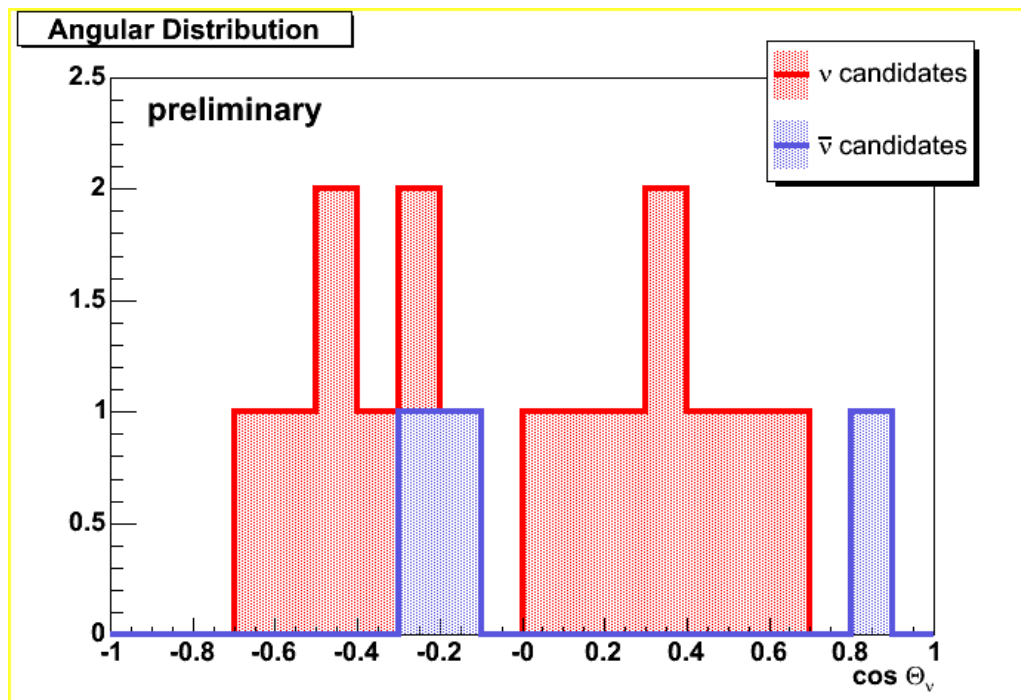
- need to measure curvature of muon track in magnetic field.
- parametrize track segments using polynomial fits.
- calculate  $\langle Q/p \rangle$  along track.

> 90% charge separation



# Current Status

- **Currently analysing subset of data – 17 candidate events so far!**
  - selecting events with contained interaction vertices.
  - 14 neutrinos + 3 anti-neutrinos (consistent with expected ratio of ~3:1).
- **Distributions of zenith angle + interaction vertex look sensible.**
- **Expect >100 contained events per year.**



# Conclusion

- **MINOS Far Detector is able to identify atmospheric neutrino events.**
- **MINOS will carry out charge-separated atmospheric neutrino oscillation analysis.**
- **Data is now accumulating – first physics results expected soon!**