



Di-electron differential cross section in Au+Au collisions at different beam energies at STAR



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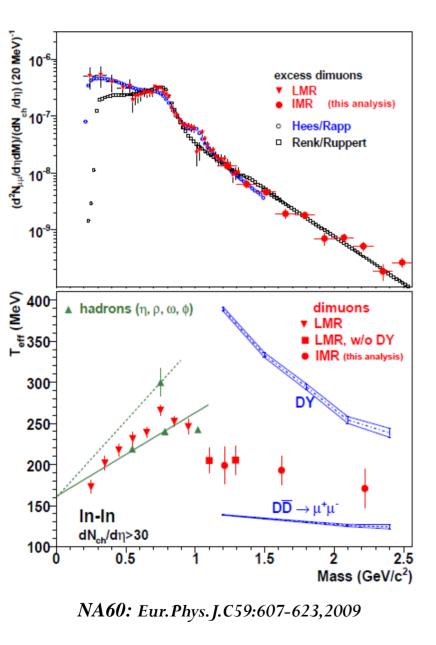


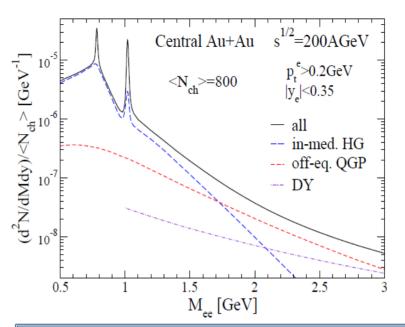
Outline



- Motivation
- Analysis
- Results in Au+Au at 200 GeV
 - Centrality dependence of di-electron production
 - Ellipitic flow v_2 of di-electron
- Di-electron production in Au+Au at 62.4, 39, 19.6 GeV
 Summary

Motivation







R. Rapp Phys.Rev. C63 (2001) 054907.

Low mass range (LMR):

In-medium modifications of vector mesons. Possible link to chiral symmetry restoration. Intermediate mass range (IMR): QGP thermal radiation.

Heavy flavor modification.

➤ Di-electron production from QGP thermal radiation increases from low to high beam energy.

STAR Beam Energy Scan program enables us to systematically study it at different beam energies.

Electron identification

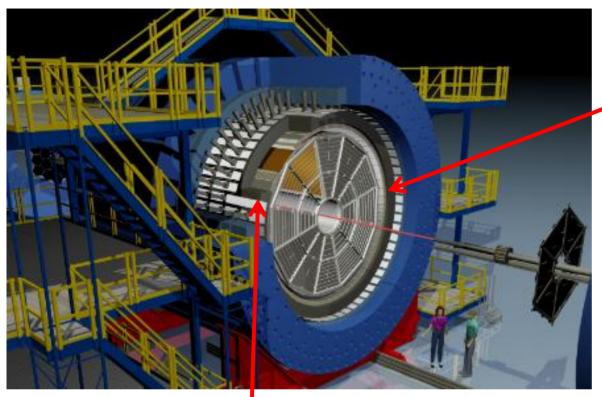
Particle ID: TOF

1. Timing resolution (<100ps)

3. Completed in 2010 (72% in 2009)

2. Coverage: $-0.9 < \eta < 0.9$

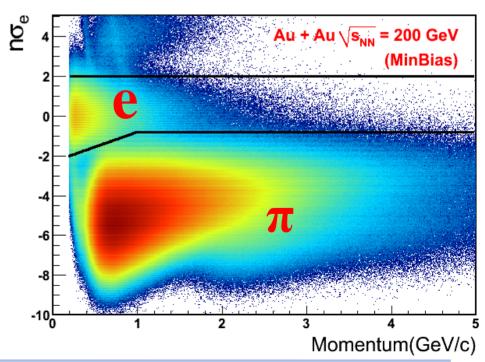
Time Of Flight ----



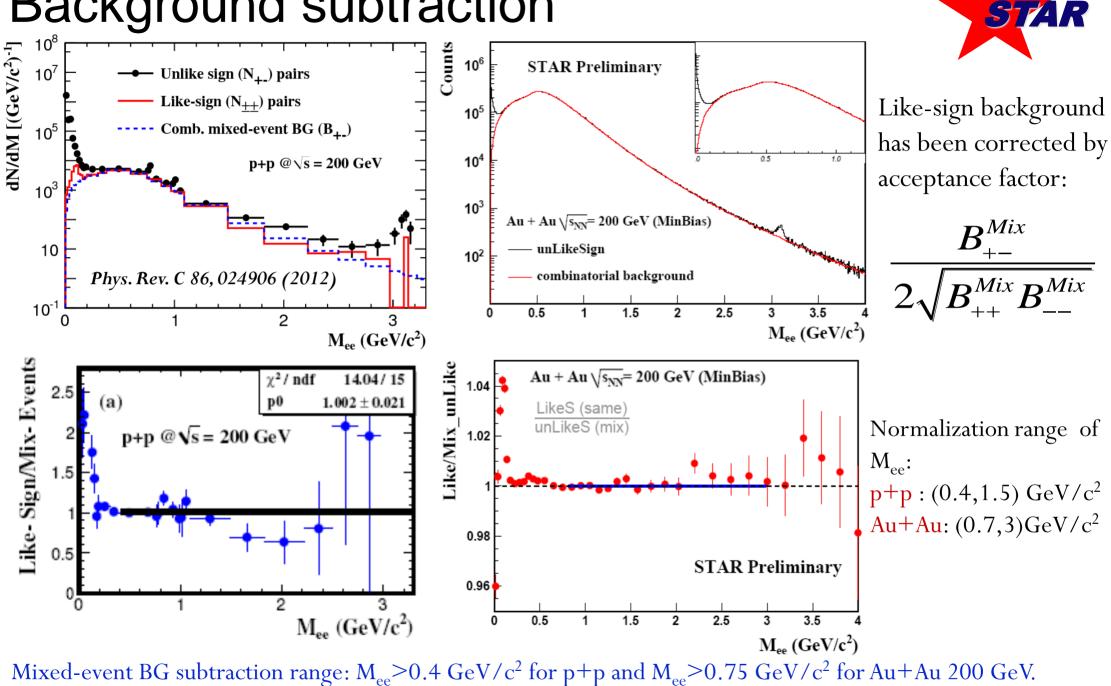


Tracking: TPC

Time Projection Chamber
1. Tracking
2. Ionization energy loss (dE/dx PID)
3. Coverage -1<η<1

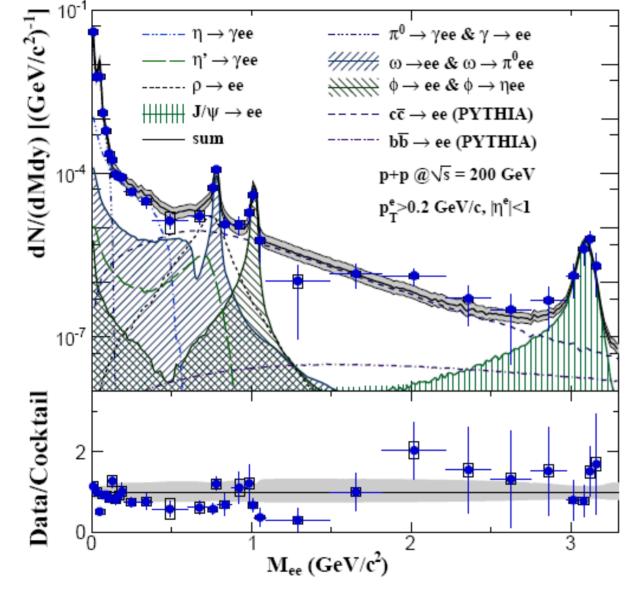


Background subtraction



Di-electron results in p+p 200 GeV



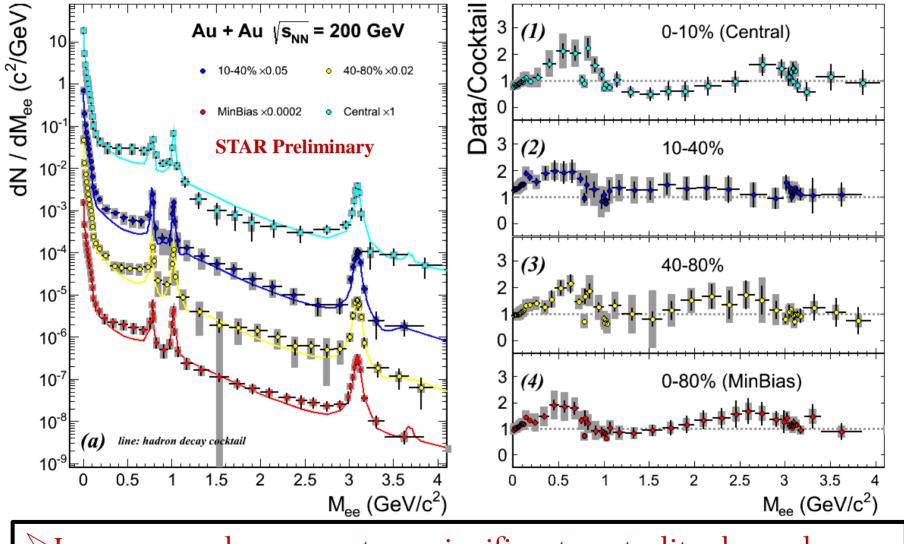


 Cocktail simulation is consistent with di-electron spectrum within quoted uncertainties.
 Intermediate mass region is dominated by charm correlation contribution.

Phys. Rev. C 86, 024906 (2012)

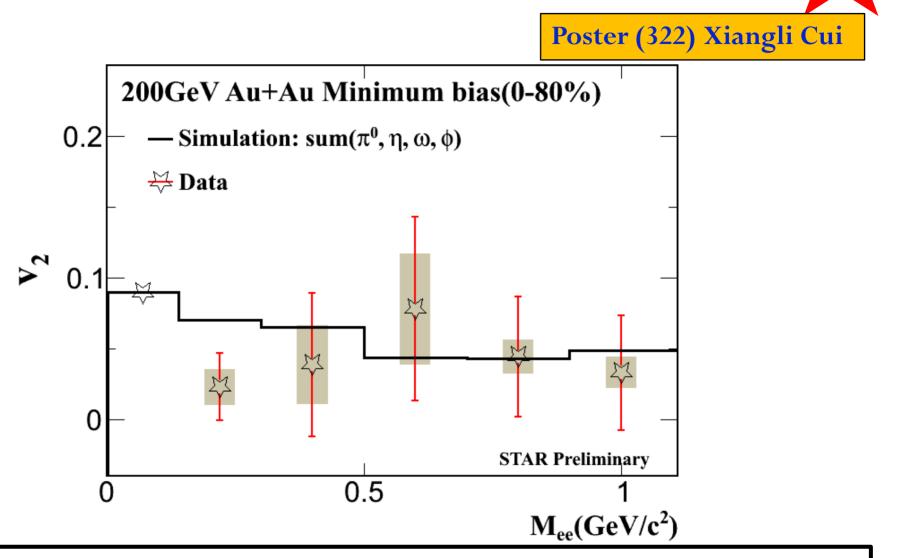
Centrality dependence in Au+Au 200 GeV STAR

Poster (153) Jie Zhao



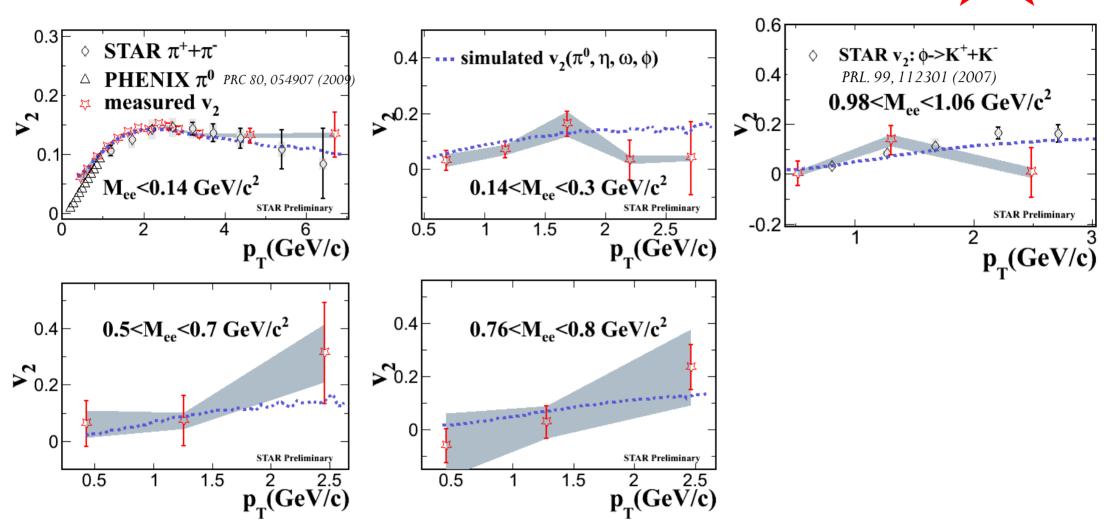
Low mass enhancement: no significant centrality dependence.

Elliptic flow v₂ in Au+Au 200 GeV



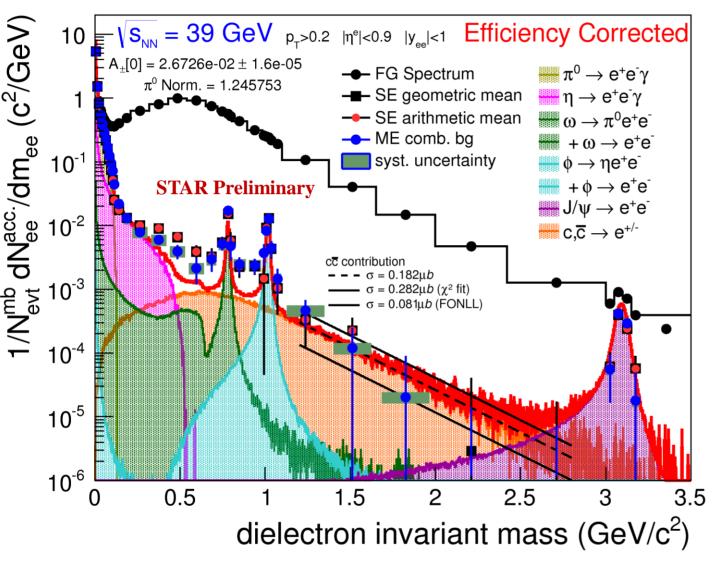
▷ Di-electron v_2 is consistent with simulation using the measured meson v_2 as input in Au+Au at 200 GeV.

p_T dependence of v_2 in Au+Au 200 GeV



Simulated v_2 is consistent with measured di-electron v_2 within uncertainties in 0-80% Au+Au collisions.

Di-electron in Au+Au at 39 GeV





Poster (113) Patrick Huck

Cocktail simulation:

• Unknown p_T distributions are taken from AMPT model calculations. The according dN/dy is extrapolated from measurements at 200 GeV based on the energy dependence given by AMPT.

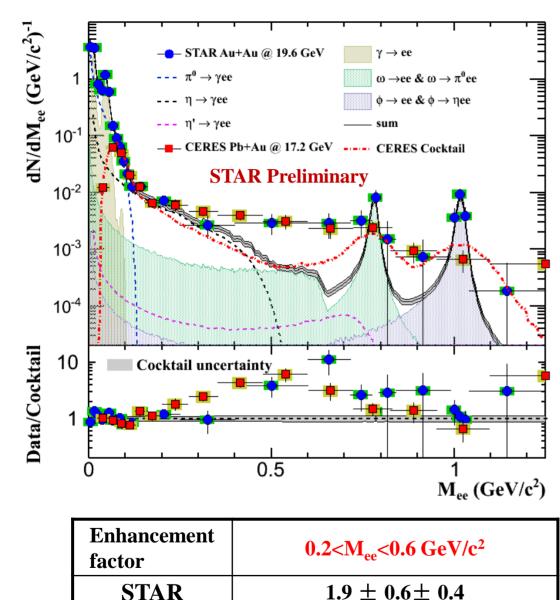
 Correlated charmed hadron contribution are simulated using
 PYTHIA and scaled to Au+Au by the number of binary collisions.

• Charm cross sections are used as center value between $\chi 2$ fits to the IMR data and FONLL prediction.

ME subtraction $M_{ee} > 0.9 \text{ GeV/c}^2$, LS subtraction $M_{ee} < 0.9 \text{ GeV/c}^2$.

Di-electron in Au+Au at 19.6 GeV

 $2.73 \pm 0.25 \pm 0.65 \pm 0.82$ [decays]



CERES

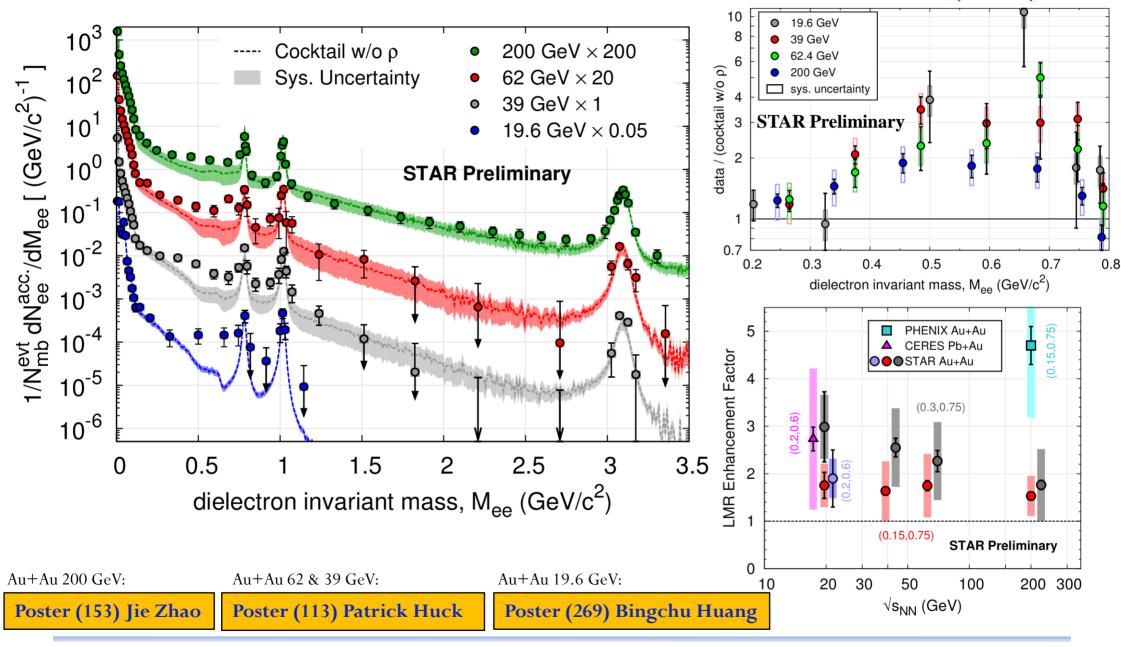
Poster (269) Bingchu Huang

• π yield is from STAR π^{+-} measurement, other meson yields derived from SPS meson/ π^0 ratio.

STAR Au+Au: 0-80% centrality, $p_T > 0.2 \text{ GeV/c}$, $|\eta| < 1$, $|y_{ee}| < 1$. CERES Pb+Au: 0-28% centrality. $p_T > 0.2 \text{ GeV/c}$, $2.1 < \eta < 2.65$, $\theta_{ee} > 35 \text{mrad}$.

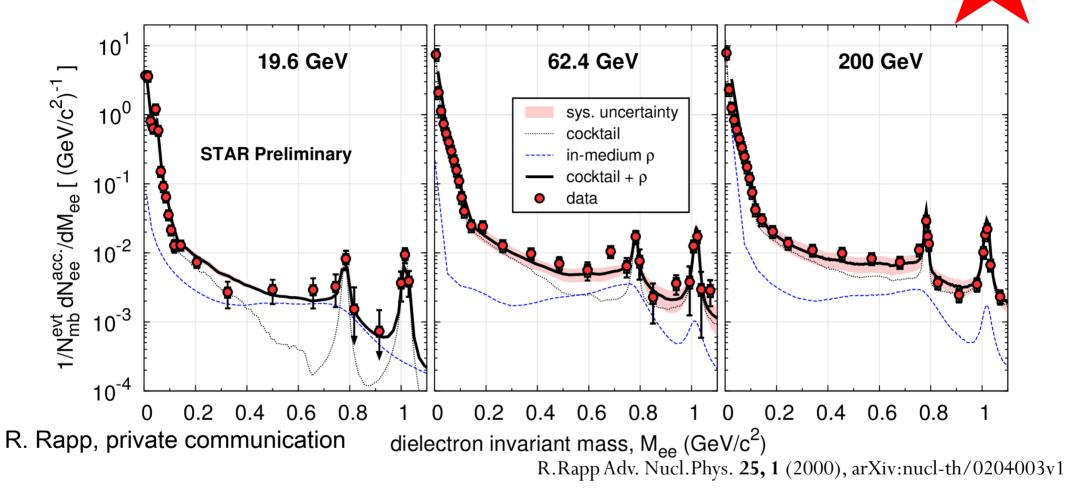
➢ Low mass enhancement comparable to CERES.

Di-electron beam energy dependence



Quark Matter 2012, Washington DC

Comparison with theory calculations



> A broadened ρ spectral function scenario can consistently describe the low mass enhancement observed from 19.6 to 200 GeV.

Summary

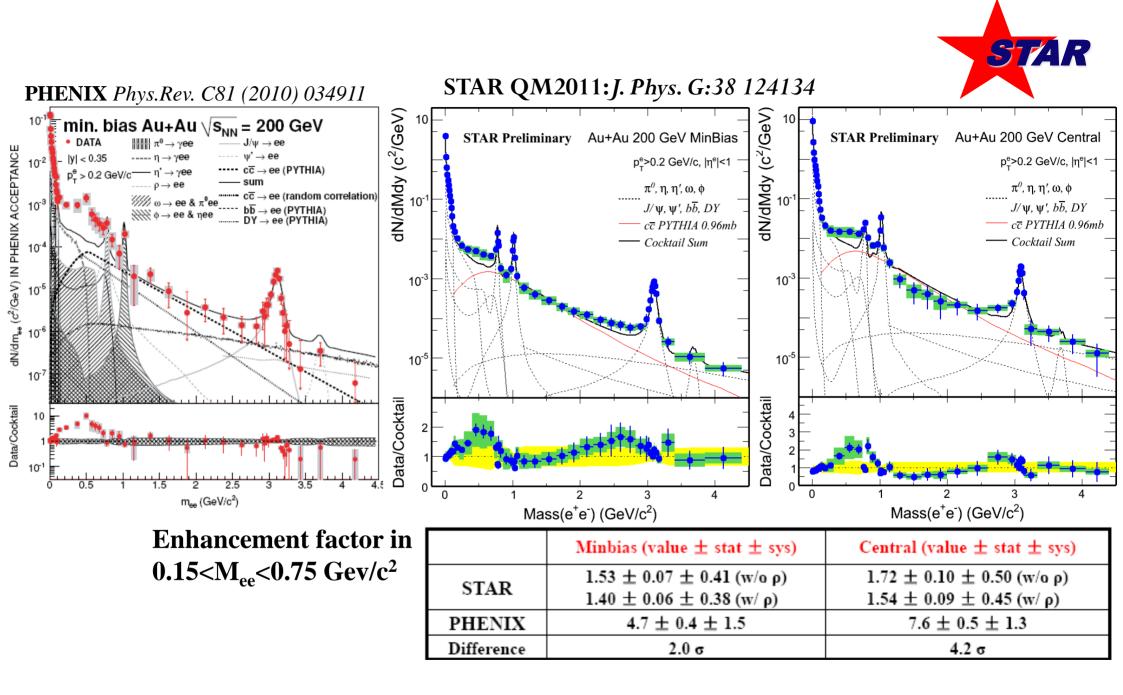


- Centrality-dependent di-electron production has been measured in Au+Au at 200 GeV.
 - No significant centrality dependence of LMR enhancement.
- Di-electron elliptic flow v₂ has been measured in Au+Au 200 GeV.
 - The simulated cocktail v_2 are consistent with the measured di-electron v_2 within uncertainties in 0-80% Au+Au collisions.
- Di-electron productions have been measured in Au+Au at 19.6, 39, and 62 GeV.
 - At 19.6 GeV, we observed comparable low mass enhancement as CERES measurements.
- A broadened ρ spectral function scenario can consistently describe the low mass enhancement observed from 19.6 to 200 GeV.



Backup



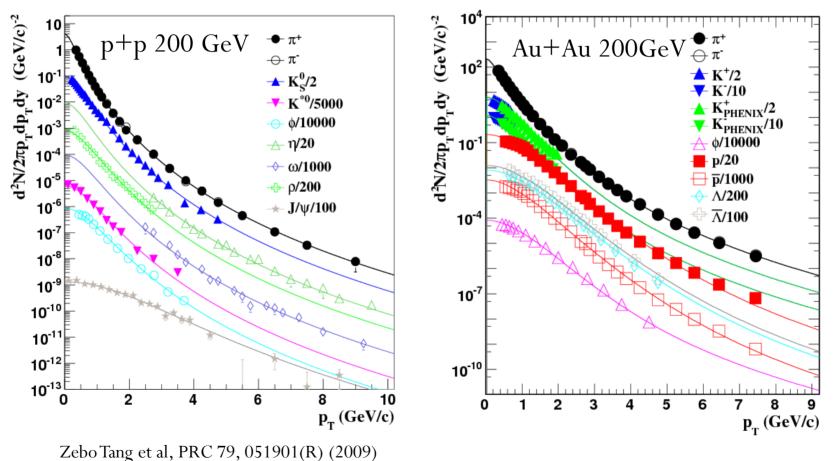


Simulation

Inputs:

Flat rapidity (-1,1), flat Φ (0, 2π), p_T: use Tsallis function fit for all measured particles.
 Hadron sources: using form factors from measurements.

> Heavy flavor sources: using PYTHIA.





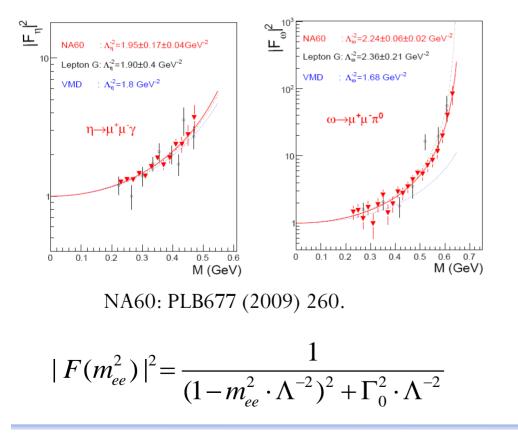
Cocktail simulation

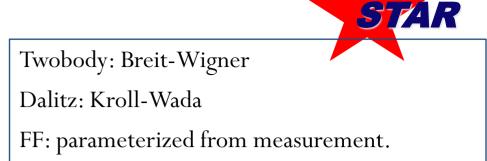
Kroll-Wada Formula:

N.M.

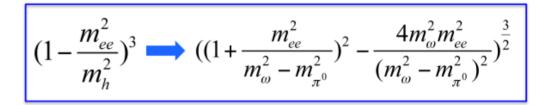
$$\frac{dN}{dm_{ee}} \propto \sqrt{1 - \frac{4m_e^2}{m_{ee}^2}} \cdot \left(1 + \frac{2m_e^2}{m_{ee}^2}\right) \cdot \frac{1}{m_{ee}} \cdot \left(1 - \frac{m_{ee}^2}{M_h^2}\right)^3 |F(m_{ee}^2)|^2$$

$$\boxed{\text{QED}}$$
Phase Form Factor
Kroll, et al., Phys Rev, 98 (1955) 5.





Phase Space term for ω , ϕ :



$$\frac{dN}{dm_{ee}dp_T} \propto \frac{m_{ee}M_{\rho}\Gamma_{ee}}{(M_{\rho}^2 - m_{ee}^2)^2 + M_{\rho}^2(\Gamma_{\pi\pi} + \Gamma_{ee}\Gamma_2)^2} \times PS,$$

 ρ line shape:

P-wave of $\pi\pi$ channel:

$$\Gamma_{\pi\pi} = \Gamma_0 \frac{M_{\rho}}{m_{ee}} (\frac{m_{ee}^2 - 4M_{\pi}^2}{M_{\rho}^2 - 4M_{\pi}^2})^{3/2},$$

S-wave of **ee** channel:

PRC 78, 044906 (2008)

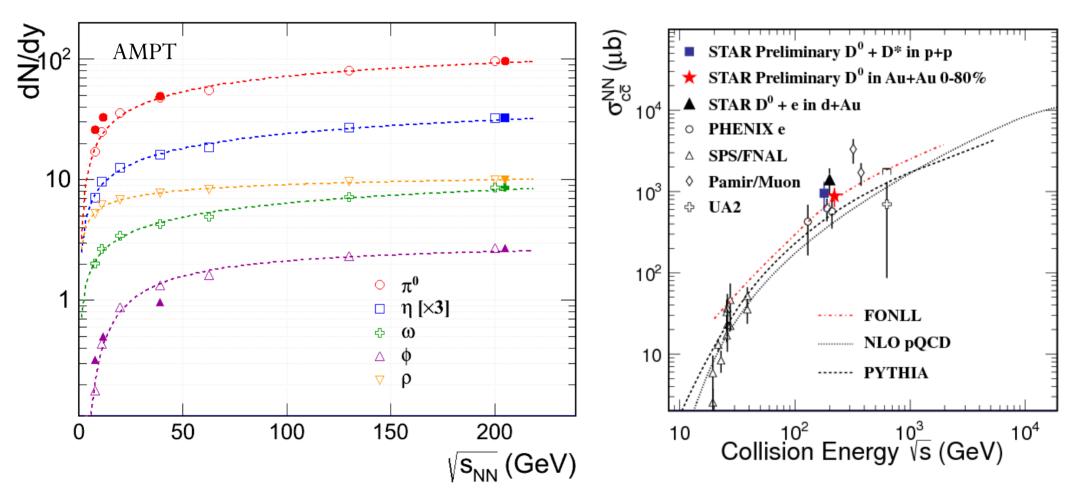
$$\Gamma_{ee} = \Gamma_0 \frac{M_\rho}{m_{ee}} (\frac{m_{ee}^2 - 4m_e^2}{M_\rho^2 - 4m_e^2})^{1/2},$$

 $PS = \frac{m_{ee}}{\sqrt{m_{ee}^2 + p_T^2}} e^{-\frac{\sqrt{m_{ee}^2 + p_T^2}}{T}}$

Extrapolation to lower energies



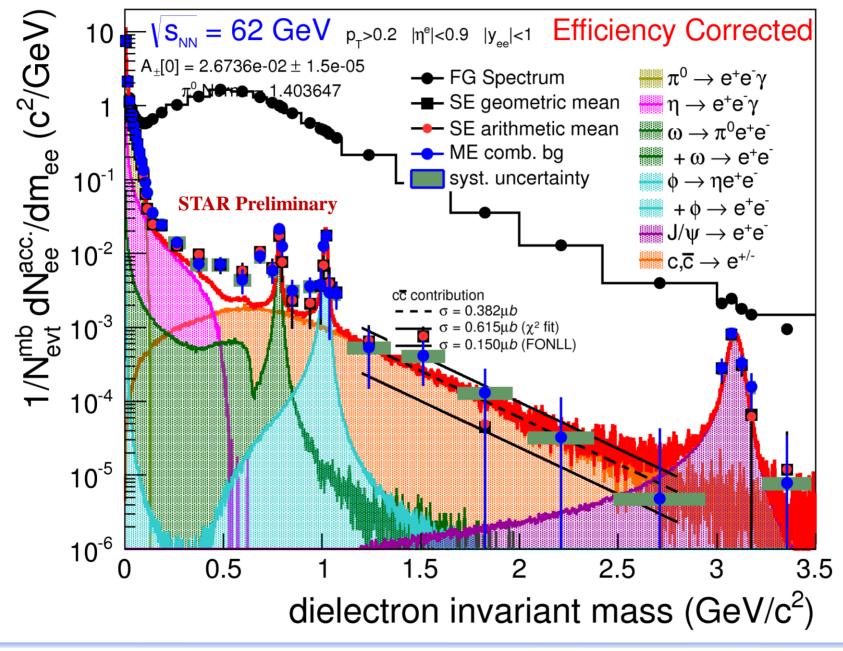
dN/dy and p_T distributions from AMPT calculations. Charm cross section: NLO calculation for low energies.



Scale to measurements at 200 GeV (solid symbols), some difference at lower energy.

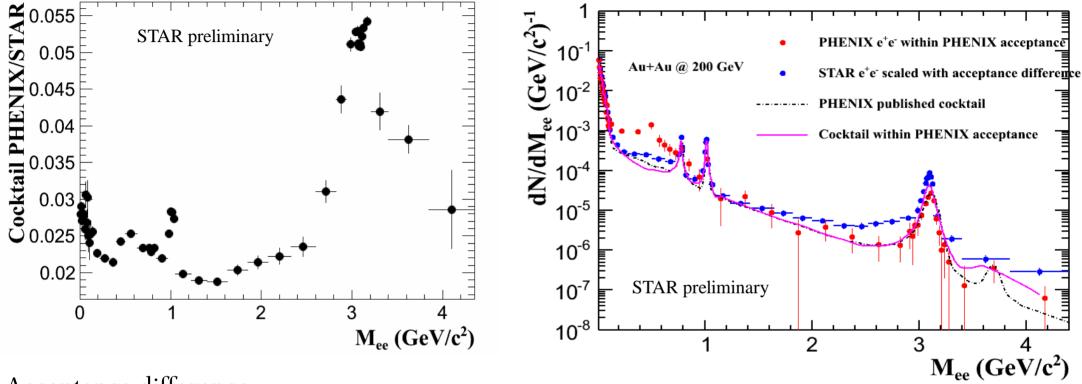
62 GeV





Check with acceptance difference





Acceptance difference:

Cocktail in PHENIX acceptance

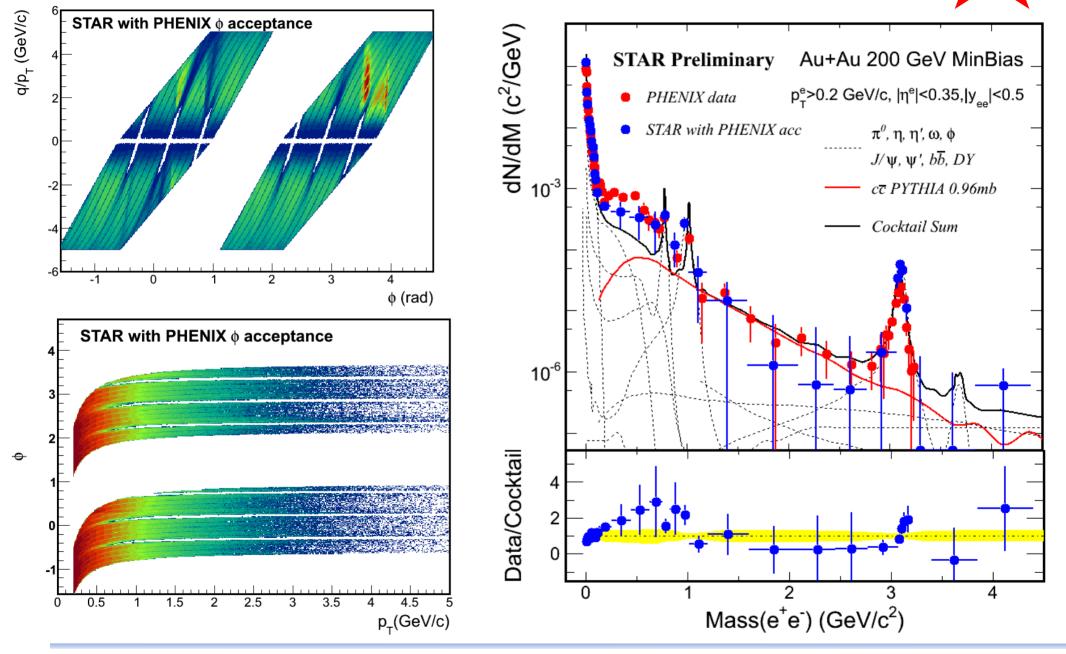
Cocktail in STAR acceptance

Scaled by same meson and charm yields.

Difference at low mass is not from the simulation but from the measurements.

Scaled by the acceptance difference

STAR with PHENIX Φ acceptance



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