

Neutrino physics

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Two useful and updated references

- Carlo Giunti "Neutrino Unbound"

<http://www.nu.to.infn.it/>

- Alessandro Strumia and Francesco Vissani review

<http://astrumia.home.cern.ch/astrumia/review.pdf>

- 1998-: impressive acceleration of neutrino physics (after 70 yrs)
- Data interpretation (at last) pretty clean
- New physics!
- Cosmology (baryogenesis, CMB, LSS, BBN,...)
- Astrophysics (probe of SUN, SNe, HE sources,...)
- Particle physics (access $\Lambda \approx M_{\text{GUT}}$, unification, flavour, LFV)

Outline

I. Neutrino observables

- Neutrino parameters
- What do we know
- How do we know

II. Theoretical impact

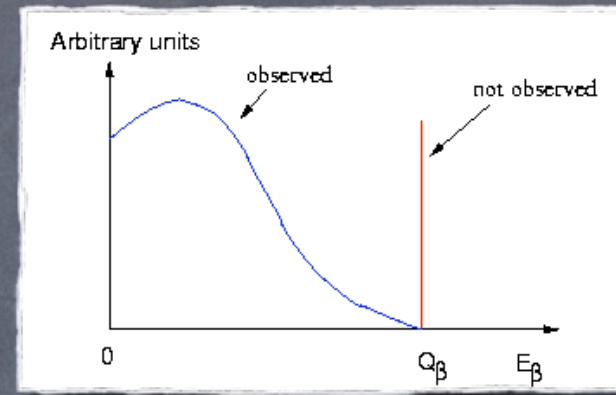
- Origin of neutrino masses
- Origin of the pattern of neutrino masses and mixing

The birth of neutrino physics (ν_e)

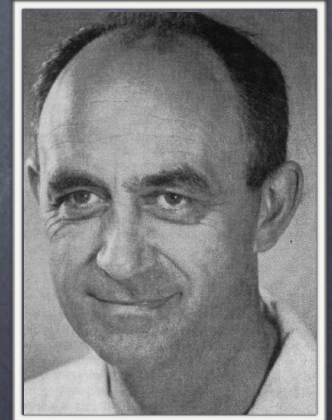
- Pauli 1930 postulated a new particle as a “desperate remedy” to save energy conservation

$$(A, Z) \rightarrow (A, Z+1) + e^- + \bar{\nu}$$

$$\nu \equiv \nu_e$$



- Fermi 1934 “speculation too remote from reality”: theory of neutrino (weak) interactions; ν = “neutrino”



- Reines and Cowan 1956 detected (reactor) neutrinos through inverse beta decay $\bar{\nu}_e + p \rightarrow n + e^+$

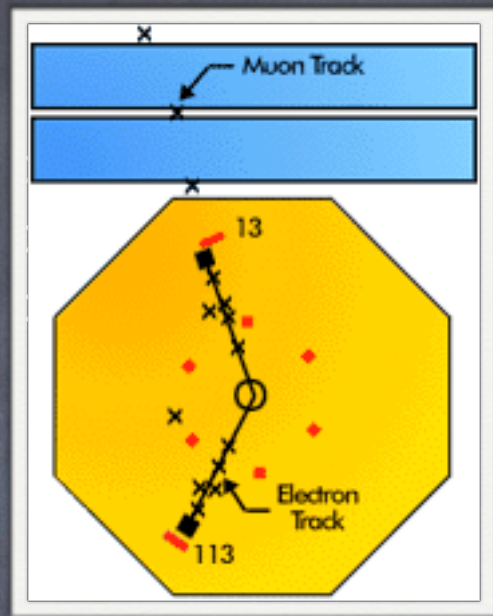
$$\nu_{\mu}$$

- “Missing momentum” in cosmic $\pi^{\pm}$ tracks: $\pi^+ \rightarrow \mu^+ \nu$



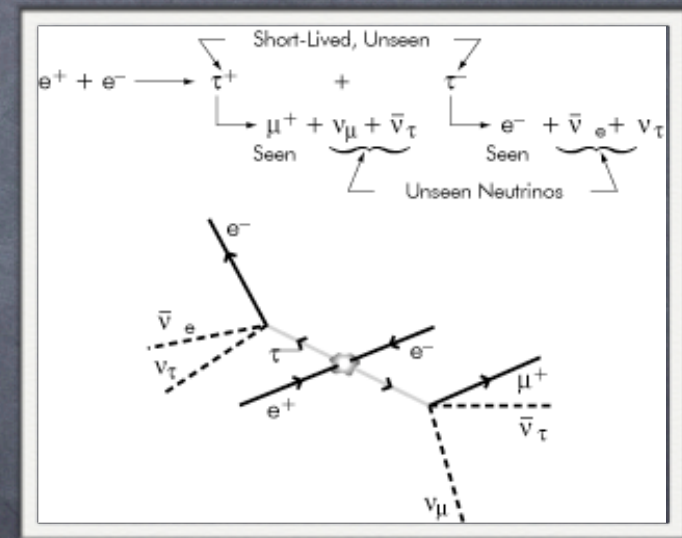
- ν from π decay detected by Lederman Schwartz Steinberger 1964: first (and prototypical) neutrino beam!
 - Protons on target $\rightarrow$ pions $\rightarrow$ neutrinos
 - Big detectors with strong shielding
- Neutrinos produced μ , not e : $\nu \equiv \nu_{\mu}$. **Lepton** (electron, muon) **number conservation** extended to neutrinos
- $(\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_{\mu})$

ν_τ



τ discovery at SPEAR, 1975:
anomalous events $e^+e^- \rightarrow e \mu X$
(beginning of third family physics)

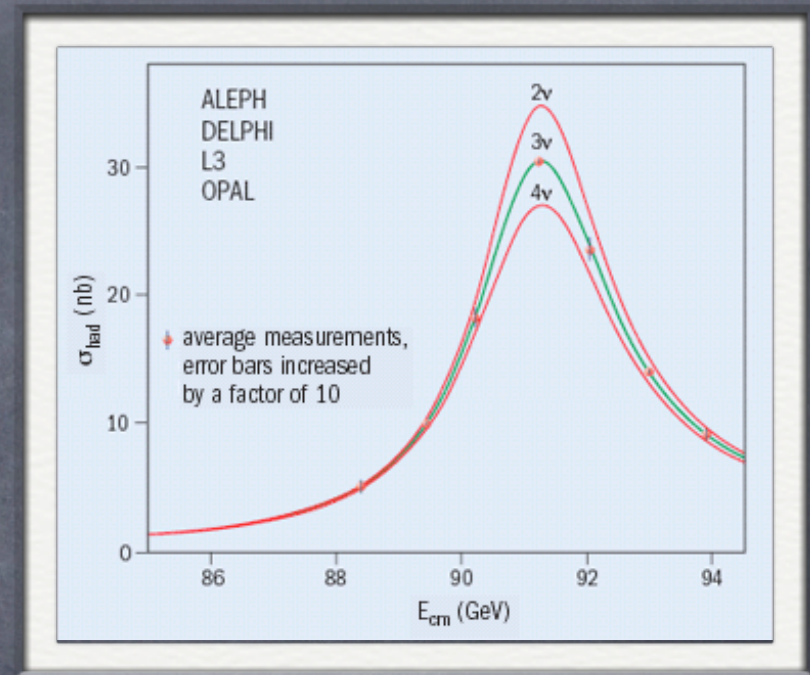
- L_e, L_μ violated? No:
(τ decay too fast to be observed)



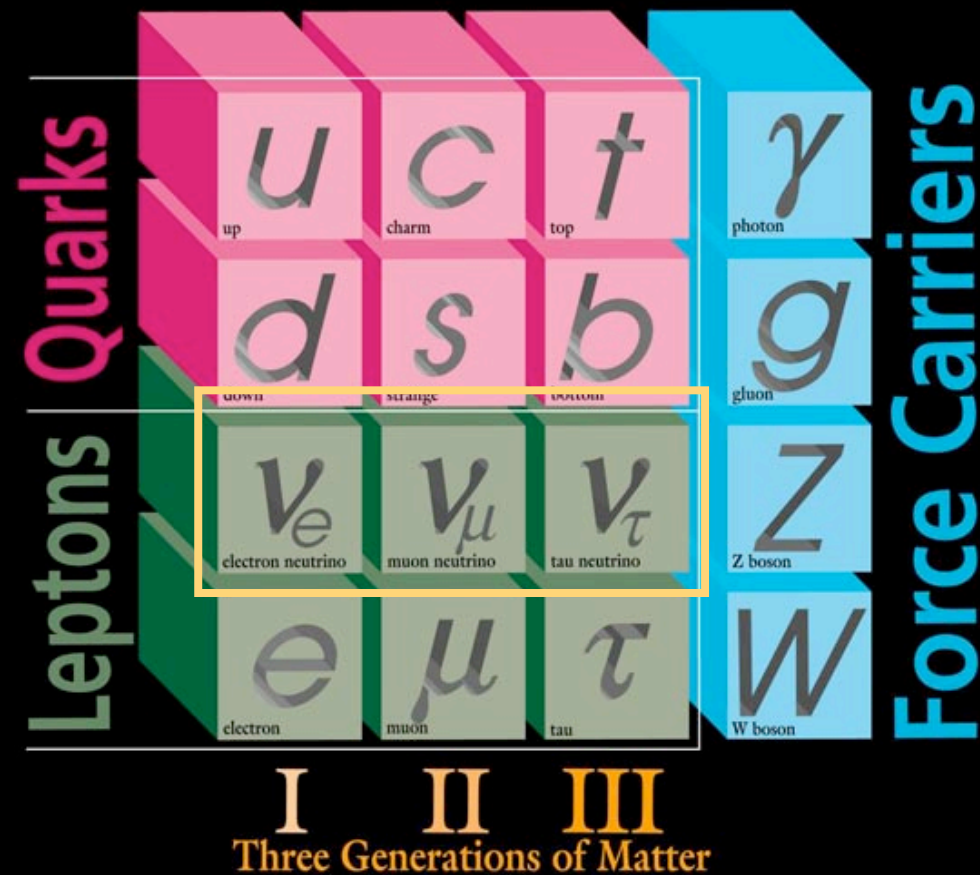
- ν_τ observed at FNAL in 2000
(beam dump; $D_s \rightarrow \tau \nu_\tau$; $\nu_\tau + N \rightarrow \tau + X$)

More ν 's?

- Z with at LEP I
- Depends on the number of (kinematically accessible) Z decay channels
- Measurement corresponds to Z -decay into the known charged particles with $m < 45$ GeV and 3 (2.98 ± 0.01) neutrinos with
 - $m < 45$ GeV
 - standard weak interactions ("sterile neutrinos" not constrained)
- 3 families (with light neutrinos)

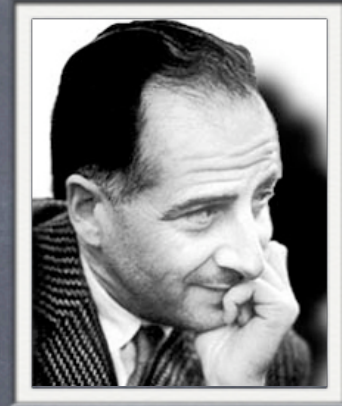


ELEMENTARY PARTICLES



Neutrino oscillations and anomalies

- Pontecorvo 1957: ν - $\bar{\nu}$ oscillations
(in analogy with $\bar{K}^0$ - K^0 oscillations)
- Maki-Nakagawa-Sakata (MNS) 1962:
 ν_e - ν_μ mixing and $\mu \rightarrow e \gamma$
(in a model of leptons bound into nuclei)
- Davis 1968: solar neutrino anomaly
(exp technique suggested by Pontecorvo)
solar neutrino flux $\approx 1/2$ of Bahcall 1963 prediction
- “Modern” era 1998-: getting rid of flux uncertainties



Neutrino parameters

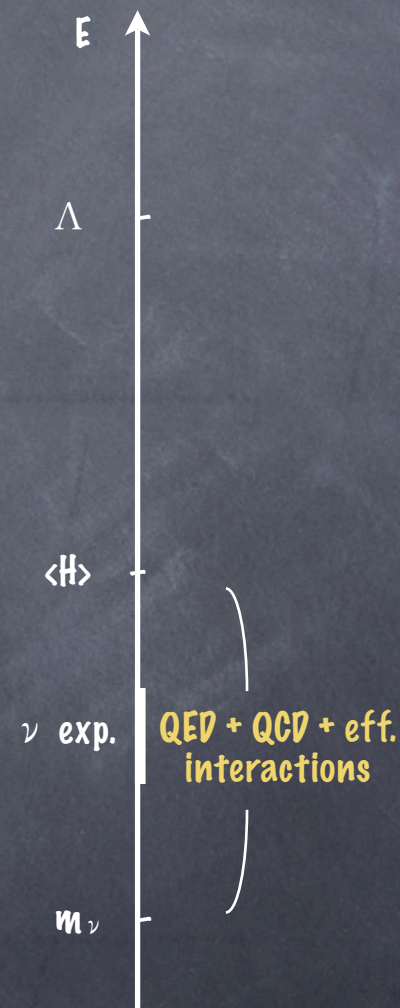
The framework

- $E \ll \langle H \rangle \approx 174 \text{ GeV}$
- Relevant symmetries: QED ($U(1)_{em}$), QCD ($SU(3)_c$)
- Neutrino interactions described by 4 fermion Fermi Gamow-Teller Sudarshan-Marshak effective hamiltonian

$$\mathcal{L}_{E \ll M_Z}^{\text{eff}} = \mathcal{L}_{\text{QED+QCD}}^{\text{ren}} + 4 \frac{G_F}{\sqrt{2}} j_c^\mu j_{c\mu}^\dagger + \text{N.C.} + \dots$$

$$j_c^\mu = \overline{\nu_{e_i}} \gamma^\mu P_L e_i + \overline{u_i} \gamma^\mu P_L d_i$$

- Neutrinos are allowed to have masses $m_1 m_2 m_3$
- Mass eigenstates $\nu_1 \nu_2 \nu_3$ diagonalize the mass matrix
Flavour eigenstates $\nu_e \nu_\mu \nu_\tau$ paired to charged leptons
in CC (also denoted ν_{ei} , $i = 1, 2, 3$)



$$V_{ei} = U_{ij} V_j$$

- U = PMNS (unitary) mixing matrix
- “Dirac” neutrinos: $4 = 3+1$ physical parameters (as CKM)
Standard parameterization:

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

- “Majorana” neutrinos: $6 = 3+3$ physical parameters
(2 extra CP-violating “Majorana” phases)

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}$$

Dirac and Majorana neutrinos

- Dirac neutrinos:

- $\nu \neq \bar{\nu}$

- lepton number conserved

- $\nu = \nu_L + \nu_R$

- Majorana neutrinos:

- $\nu = \bar{\nu}$

- lepton number violated

- $\nu = \nu_L$

- The distinction is irrelevant

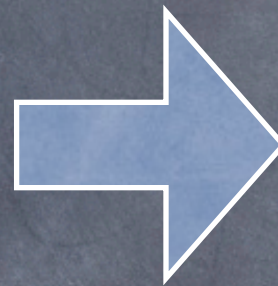
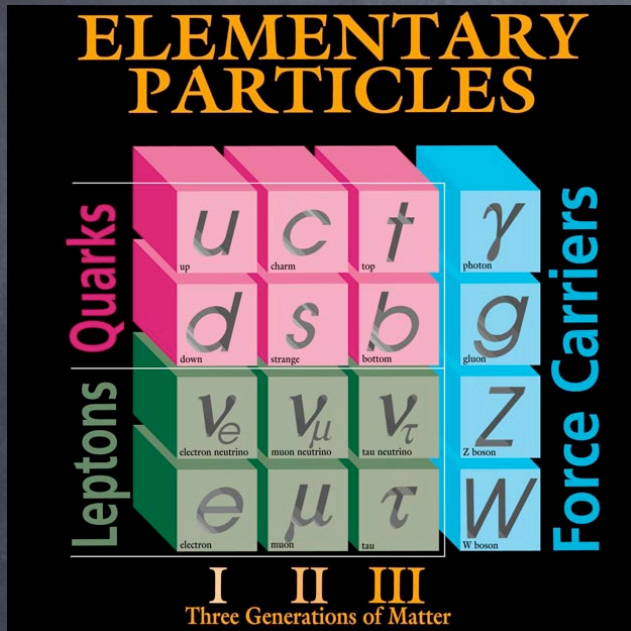
- in oscillation experiments ($E \gg m_\nu$)

- most experiments except when lepton number violation plays a role ($0\nu 2\beta$)

(Interlude: the neutrino
mass term)

Left- and Right-handed

The Dirac spinor e is not elementary: $e = e_L + e_R$



$u_{iL} \ d_{iL} \ \nu_{iL} \ e_{iL}$ **L**

$u_{iR} \ d_{iR} \quad \quad e_{iR}$ **R**

$\bar{u}_{iL} \ \bar{d}_{iL} \ \bar{\nu}_{iL} \ \bar{e}_{iL}$ **R**

$\bar{u}_{iR} \ \bar{d}_{iR} \quad \quad \bar{e}_{iR}$ **L**

Most general gauge transformation can mix all **L**

Most general **mass term** combines **LL** (in gauge invariant way)

(will not specify the Lorentz contraction)

Fermion mass terms

ψ_i “L” fermions

ψ

Most general mass term: $\frac{m_{ij}}{2} \psi_i \psi_j$

$$\frac{m}{2} \psi \psi$$

m_{ij} symmetric

“Majorana”
breaks any charge of ψ

example: $\psi = \nu_L$
 $m \neq 0$ allowed by QED+QCD
breaks lepton number

Fermion mass terms

ψ_i “L” fermions

ψ, ψ^c

Most general mass term: $\frac{m_{ij}}{2} \psi_i \psi_j$

$$\underbrace{\frac{m_1}{2} \psi \psi + \frac{m_2}{2} \psi^c \psi^c}_{\text{“Majorana”}} + \underbrace{m \psi^c \psi}_{\text{“Dirac”}}$$

m_{ij} symmetric

“Majorana”

“Dirac”

$Q(\psi) + Q(\psi^c) = 0$
(all charged SM fermions)

(e.g. Dirac neutrino mass term: $m \overline{\nu_R} \nu_L$)

$$\psi^c = \overline{\nu_R} \quad \psi = \nu_L$$

needs ν_R

does not break lepton number)

(e.g. electron mass term: $m \overline{e_R} e_L$)

$$\psi^c = \overline{e_R} \quad \psi = e_L$$

Most general mass terms (no ν_R)

• L fields:

	u_L	d_L	ν_L	e_L	$\overline{u_R}$	$\overline{d_R}$	$\overline{e_R}$
Q	$\frac{2}{3}$	$-\frac{1}{3}$	0	-1	$-\frac{2}{3}$	$\frac{1}{3}$	1
$SU(3)_c$	3	3	1	1	$\overline{3}$	$\overline{3}$	1

• Gauge invariant LL terms:

$$m_u \overline{u_R} u_L + m_d \overline{d_R} d_L + m_e \overline{e_R} e_L + \frac{m_\nu}{2} \nu_L \nu_L$$

With ν_R

- ν_R : charge and color singlet (as ν_L)
but does not enter weak interactions

$$\frac{m_L}{2} \nu_L \nu_L \rightarrow \frac{m_L}{2} \nu_L \nu_L + \frac{m_R}{2} \overline{\nu_R \nu_R} + m_N \overline{\nu_R} \nu_L$$

- Dirac limit (lepton number conserved): $m_L = 0$, $m_R = 0$

$$\frac{m_L}{2} \nu_L \nu_L \rightarrow m_N \overline{\nu_R} \nu_L$$

- For definiteness (and theoretical prejudice, see later) we will focus on Majorana type mass terms

Masses and mixings: quarks

• Mass eigenstates

$$m^D = U_{d_R}^\dagger m_{\text{diag}}^D U_{d_L} \quad m^U = U_{u_R}^\dagger m_{\text{diag}}^U U_{u_L}$$

$$\begin{cases} d'_{iR} = U_{ij}^{d_R} d_{jR} \\ d'_{iL} = U_{ij}^{d_L} d_{jL} \end{cases}, \begin{cases} u'_{iR} = U_{ij}^{u_R} u_{jR} \\ u'_{iL} = U_{ij}^{u_L} u_{jL} \end{cases}$$

$$m_{ij}^D \overline{d_{iR}} d_{jL} + m_{ij}^U \overline{u_{iR}} u_{jL} = m_{d_i} \overline{d'_{iR}} d'_{iL} + m_{u_i} \overline{u'_{iR}} u'_{iL}$$

• In terms of mass eigenstates:

$$j_{c,\text{had}}^\mu = \overline{u_{iL}} \gamma^\mu d_{iL} = V_{ij} \overline{u'_{iL}} \gamma^\mu d'_{jL}$$

$$j_{n,\text{had}}^\mu = (j_{n,\text{had}}^\mu)'$$

$$j_{\text{em},\text{had}}^\mu = (j_{\text{em},\text{had}}^\mu)'$$

$$V = U_{u_L} U_{d_L}^\dagger$$

Cabibbo Kobayashi Maskawa (CKM) matrix

Physical parameters in V

$$m_{d_i} \overline{d_{iR}} d_{iL} + m_{u_i} \overline{u_{iR}} u_{iL} \quad j_{c,\text{had}}^\mu = V_{ij} \overline{u_{iL}} \gamma^\mu d_{jL}$$

$$V = \underbrace{\begin{pmatrix} e^{i\tau_1} & & \\ & e^{i\tau_2} & \\ & & e^{i\tau_3} \end{pmatrix}}_{\text{unphysical}} \underbrace{\begin{pmatrix} \text{standard par.} \end{pmatrix}}_{\text{standard par.}} \underbrace{\begin{pmatrix} 1 & & \\ & e^{i\sigma} & \\ & & e^{i\rho} \end{pmatrix}}_{\text{unphysical}}$$

$$9 = 3 + 3 + 1 + 2$$

Masses and mixings: leptons

• Mass eigenstates

$$m^\nu = U_{\nu L}^T m_{\text{diag}}^D U_{\nu L} \quad m^E = U_{e R}^\dagger m_{\text{diag}}^E U_{e L}$$

$$\nu'_{iL} = U_{ij}^{\nu L} \nu_{jL}, \quad \begin{cases} e'_{iR} = U_{ij}^{eR} e_{jR} \\ e'_{iL} = U_{ij}^{eL} e_{jL} \end{cases}$$

$$\frac{m_{ij}^\nu}{2} \nu_{iL} \nu_{jL} + m_{ij}^E \overline{e_{iR}} e_{jL} = \frac{m_{\nu_i}}{2} \nu'_{iL} \nu'_{iL} + m_{e_i} \overline{e'_{iR}} e'_{iL}$$

• In terms of mass eigenstates:

$$j_{\text{c,lep}}^\mu = \overline{\nu_{iL}} \gamma^\mu e_{iL} = U_{ij}^\dagger \overline{\nu'_{iL}} \gamma^\mu e'_{jL}$$

$$j_{\text{n,lep}}^\mu = (j_{\text{n,lep}}^\mu)'$$

$$j_{\text{em,lep}}^\mu = (j_{\text{em,lep}}^\mu)'$$

$$U = U_{eL} U_{\nu L}^\dagger$$

Pontecorvo - Maki Nakagawa Sakata (P-MNS) matrix

Physical parameters in U

$$\frac{m_{\nu_i}}{2} \nu_{iL} \nu_{iL} + m_{e_i} \overline{e_{iR}} e_{iL} \quad j_{c,lep}^{\mu\dagger} = U_{ij} \overline{e_{iL}} \gamma^\mu \nu_{jL}$$

$$U = \underbrace{\begin{pmatrix} e^{i\gamma_1} & & \\ & e^{i\gamma_2} & \\ & & e^{i\gamma_3} \end{pmatrix}}_{\text{unphysical}} \underbrace{\begin{pmatrix} \text{standard par.} \end{pmatrix}}_{\text{3 + 1}} \underbrace{\begin{pmatrix} 1 & & \\ & e^{i\alpha} & \\ & & e^{i\beta} \end{pmatrix}}_{\text{physical (Majorana) 2}}$$

9 = 3 + 3 + 1 + 2

(End of interlude)

Physical mass and mixing parameters in the lepton sector

$$\frac{m_{\nu_i}}{2} \nu_{iL} \nu_{iL} + m_{e_i} \overline{e_{iR}} e_{iL} \quad j_{\text{c,lep}}^\mu = U_{ij}^\dagger \overline{\nu}_{iL} \gamma^\mu e_{jL}$$

$$m_e, m_\mu, m_\tau, m_{\nu_1}, m_{\nu_2}, m_{\nu_3}, \theta_{23}, \theta_{12}, \theta_{13}, \delta, \alpha, \beta$$

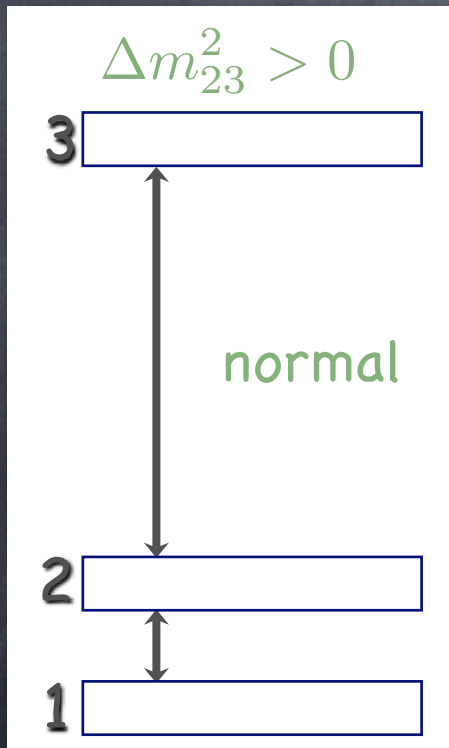
$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}s^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}s^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}$$

$$0 \leq \theta_{23}, \theta_{12}, \theta_{13} \leq \frac{\pi}{2}, \quad 0 \leq \delta < 2\pi, \quad 0 \leq \alpha, \beta < 2\pi$$

Standard labeling of eigenstates

$0 < \Delta m_{12}^2 < |\Delta m_{23}^2|$ uniquely defines the labeling
 $\Delta m_{12}^2 > 0$ by definition; Δm_{23}^2 can have both signs

$$\begin{pmatrix} \Delta m_{\text{SUN}}^2 \equiv \Delta m_{12}^2 \\ \Delta m_{\text{ATM}}^2 \equiv \Delta m_{23}^2 \end{pmatrix}$$



e.g.:

$$m_1 < m_2 \ll m_3$$

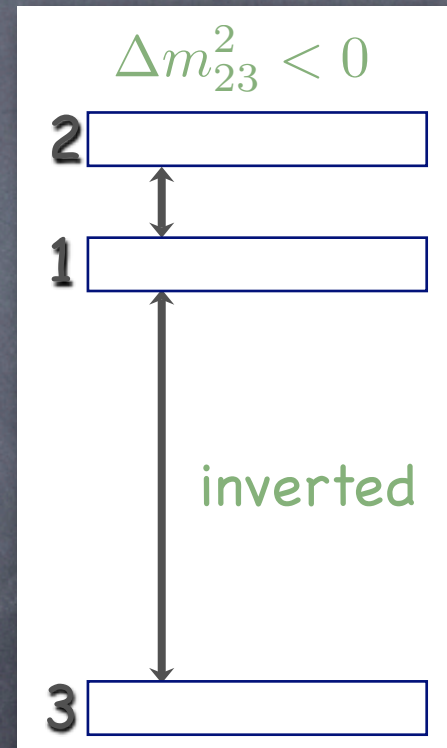
(hierarchical)

$$m_1 \approx m_2 \approx m_3$$

(degenerate)

$$m_1 \approx m_2 < m_3$$

(neither)



e.g.:

$$m_1 \approx m_2 > m_3$$

(inverse
hierarchical)

$$m_1 \approx m_2 \approx m_3$$

(degenerate)

$$\Delta m_{ij}^2 \equiv m_j^2 - m_i^2$$

Neutrino parameters:
what do we know?

Accessible
to oscillations

Not accessible
to oscillations

Charged
sector

$$m_{e,\mu,\tau}$$

$$\Delta m_{12}^2$$

$$|\Delta m_{23}^2|$$

$$\text{sign}(\Delta m_{23}^2)$$

$$\theta_{12}, \theta_{23}, \theta_{13}, \delta$$

$$m_{\text{lightest}}$$

$$\alpha$$

$$\beta$$

$$(\Delta m_{ij}^2 \equiv m_{\nu_j}^2 - m_{\nu_i}^2)$$

Accessible
to oscillations

Not accessible
to oscillations

Charged
sector

$$m_{e,\mu,\tau}$$

Well known

$$\Delta m_{12}^2$$

$$|\Delta m_{23}^2|$$

$$\text{sign}(\Delta m_{23}^2)$$

$$\theta_{12}, \theta_{23}, \theta_{13}, \delta$$

$$m_{\text{lightest}}$$

$$\alpha$$

$$\beta$$

Known

Bounds

$$\Delta m_{\text{ATM}}^2 \sim 2.4 \times 10^{-3} \text{ eV}^2 \quad \theta_{23} \sim 45^\circ$$

(ATM, K2K, Minos)

$$\Delta m_{\text{SUN}}^2 \sim 0.76 \times 10^{-4} \text{ eV}^2 \quad \theta_{12} \sim 35^\circ$$

(SUN, KamLAND)

$$\theta_{13} < 7^\circ \quad (2\sigma)$$

(CHOOZ, Palo Verde + ATM)

$$|m_{ee}| = |U_{ei}^2 m_{\nu_i}| < \mathcal{O}(1) \times 0.4 \text{ eV}$$

(Heidelberg-Moscow)

$$(m^\dagger m)_{ee} = |U_{ei}|^2 m_{\nu_i}^2 < (2.2 \text{ eV})^2$$

(Mainz, Troitsk)

$$\sum_i m_{\nu_i} < \mathcal{O}(1) \text{ eV (priors)}$$

(Cosmology)

$$m_{\nu_i} \ll 174 \text{ GeV}$$

$$\theta_{23} \sim 45^\circ (= 45^\circ?)$$

Guidelines for theory:

$$\theta_{12} \sim 30^\circ - 35^\circ \neq 45^\circ (> 5\sigma)$$

$$\theta_{13} < 10^\circ$$

$$|\Delta m_{12}^2 / \Delta m_{23}^2| \approx 0.035 \ll 1$$