## High-resistivity CMOS pixel sensors and their application to the STAR PIXEL detector

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#### Contents

 Requirements for the Heavy Flavour Tracker of the STAR experiment (see also L.Greiner's talk)

- Monolithic Active Pixel Sensors, EUDET Beam telescope
- EUDET BT sensor as a prototype for STAR implemented in a high resistivity epitaxial layer

 Prototyping the pixel optimisation for the STAR sensor (called ULTIMATE)

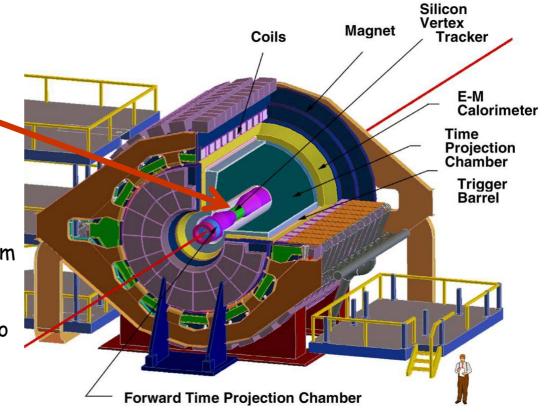
Conclusions and perspectives

#### STAR experiment and Heavy Flavour Tracker

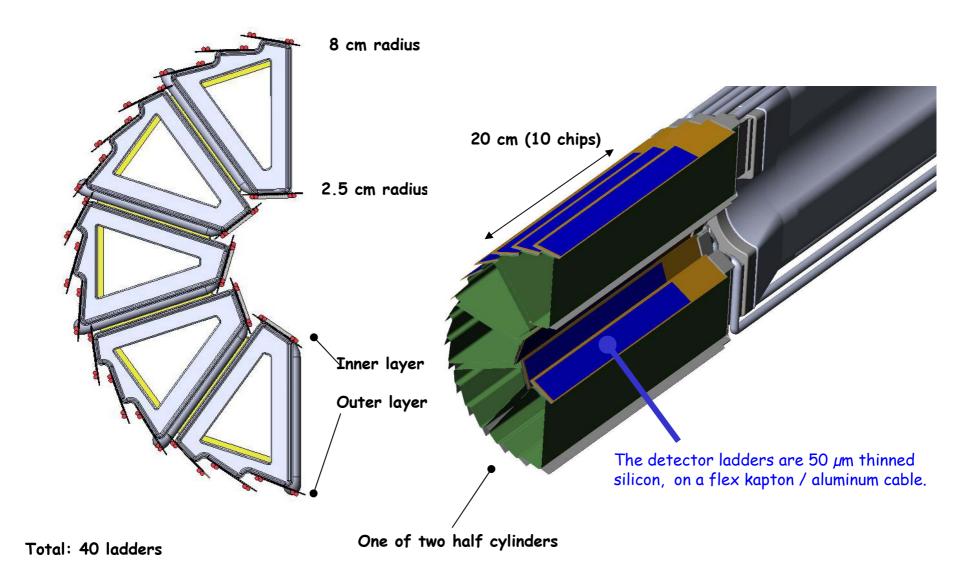
The STAR (Solenoidal Tracker At RHIC) experiment at the Relativistic Heavy Ion Collide (RHIC) studies the new state of matter produced in relativistic heavy ion collisions.

The Heavy Flavour Tracker (HFT) will use 2 cylindrical monolithic active pixel sensors (MAPS) layers for the vertex detector upgrade

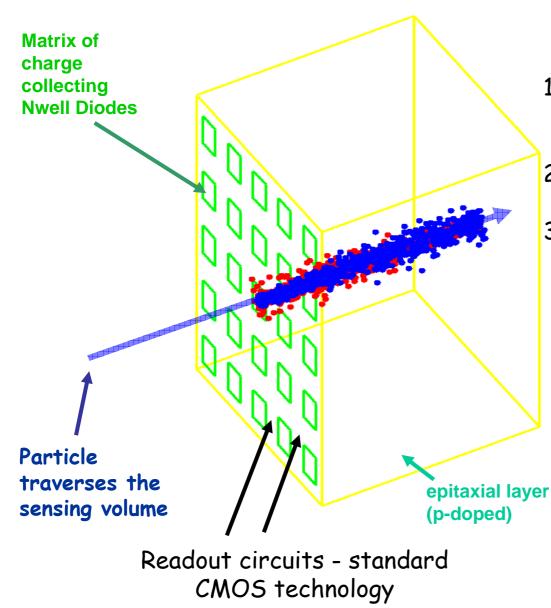
- small radiation length: thinned to 50 um on a flex kapton/aluminium cable to minimize multiple scattering
- sensors positioned close (2.5 8 cm) to the interaction region (~ 150 kRad/yr and 3e12 Neq/cm<sup>2</sup>)
- resolution sufficient to resolve the secondary decay vertices from the primary vertex (< 10 um)</li>
- work at ambient (~ +35 C) temperature



#### PIXEL detector geometry



#### Monolithic Active Pixel Sensors

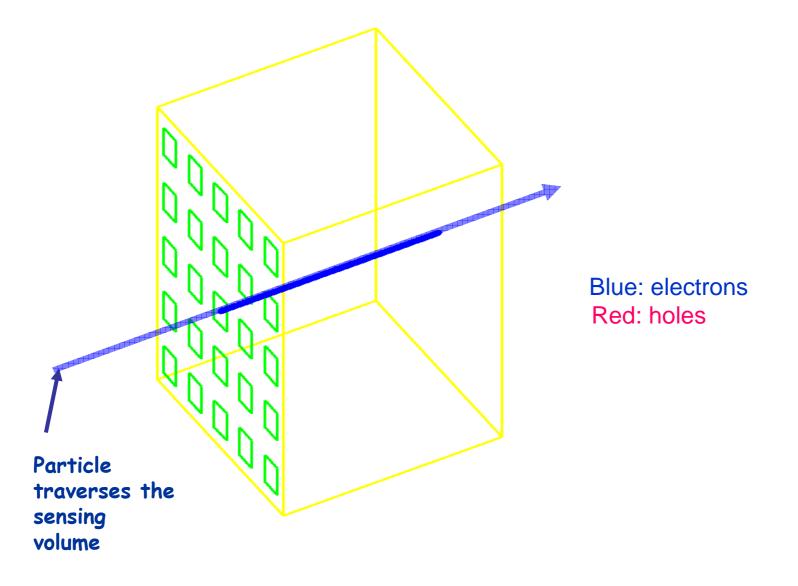


Under development since 1999 at IPHC (Strasbourg):

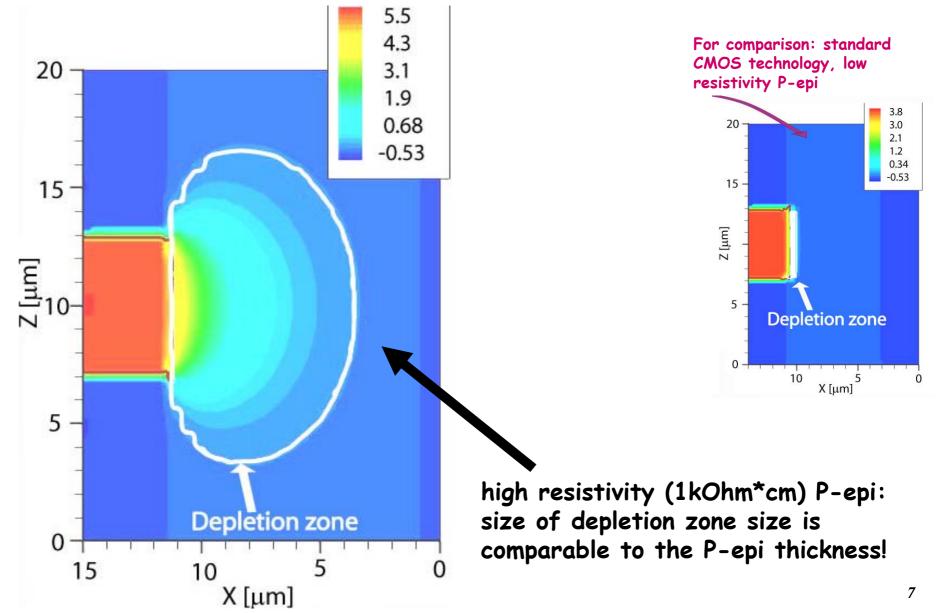
- MAPS implemented in standard CMOS substrate with readout electronics
- use p-epitaxial layer as sensing volume (~10-20 um )
- 3. Nwell Diodes collect e-h pair liberated by the particle

- low cost
- high granularity
- low material budget
- in-pixel signal amplification
- low noise
- integrated readout and signal processing

#### Monolithic Active Pixel Sensors: charge collection



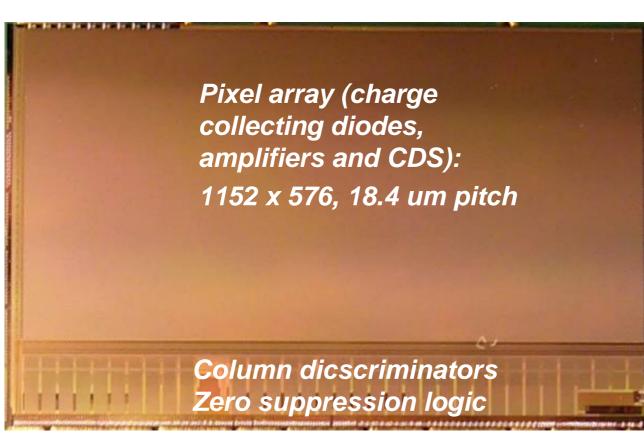
#### MAPS in a high resistivity epitaxial layer



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#### Sensor (Mimosa26) for EUDET beam telescope

#### 10.6 x 21.2 mm<sup>2</sup>



CMOS AMS 0.35 um OPTO process, epilayer 400 Ohm\*cm

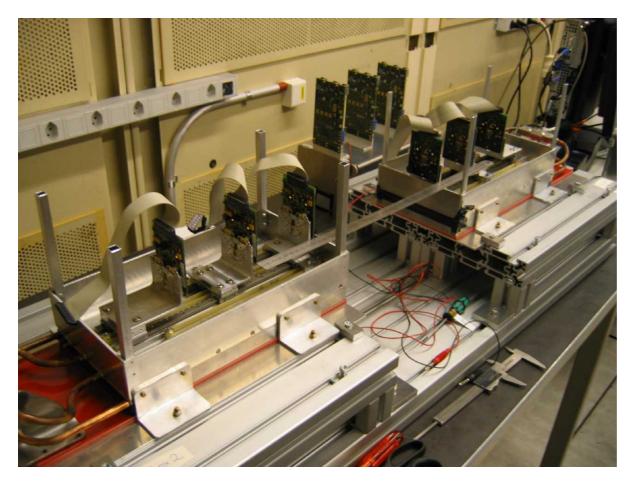
Read-out in rollingshutter mode: 80 MHz clock (can work up to 115MHz), 115.2 us (down to ~85 us) integration time

binary signal readout

zero suppression in 18 groups of 64 columns allowing  $\leq$  9 "pixel strings" / row

#### EUDET beam telescope

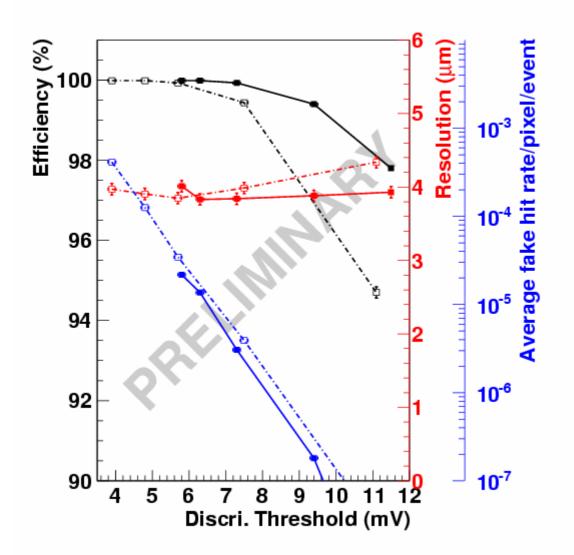
- $\cdot$  2 arms of 3 planes
- $\cdot$  Mimosa26 thinned to 50  $\mu m$
- extrapolated 1-2 µm EVEN with e- (3 GeV, DESY)
- frame read-out frequency O(10^4) Hz
- $\cdot$  running since '07 (demonstrator: analogues outputs) at CERN-SPS & DESY



A. Dorokhov, IPHC, Strasbourg, France

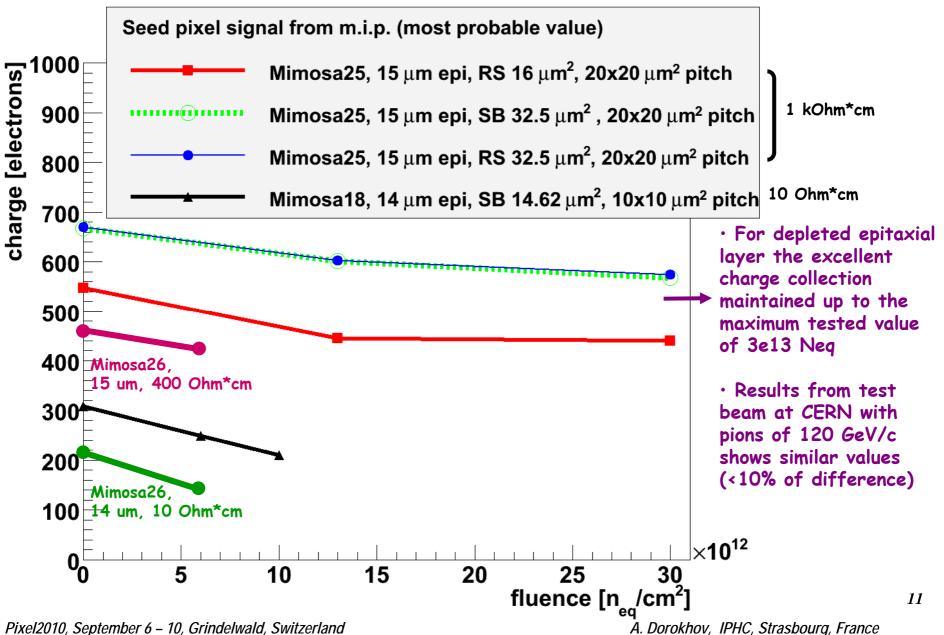
#### EUDET beam telescope sensor performance

Mi26 HR-15 and HR-10 Efficiency, Fake rate and Resolution



- non-irradiated
- $\cdot$  thinned down to 50 um
- 20 C temperature
- 400 Ohm\*cm epitaxial layer
- $\cdot$  15 um and 10 um thickness of epitaxial layer

#### Comparison of charge collection from Ru106 source for irradiated sensors



#### Mimosa26 sensor as a prototype for the STAR HFT

Mimosa26 sensor

ULTIMATE sensor (chip submission October 2010), design in progress

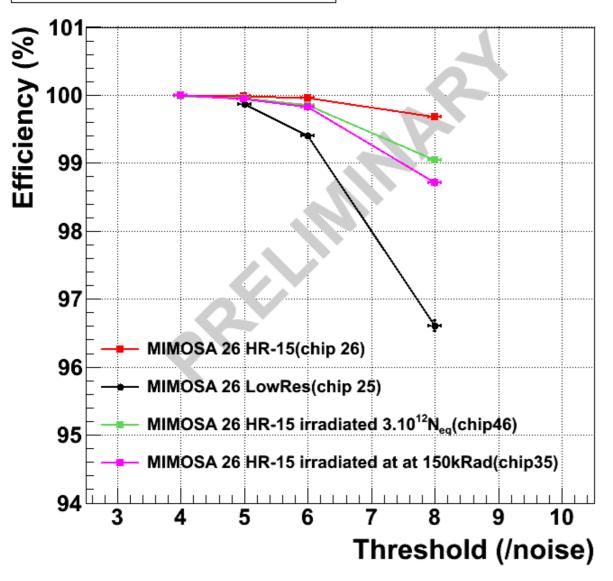
- 80 MHz
- 115.2 us integration time
- temperature ~ 20C
- 1152 x 576 pixels matrix
- 18.4 um pitch
- $\cdot$  no constrains on radiation tolerance

- larger matrix, integration time will be larger (<200 us)</li>
- temperature +30 +35C
- 960 x 928 pixels matrix
- $\cdot$  20.7 um pitch, to fit power consumption constraints (< 135 mW/cm^2)
- $\cdot$  150 kRad ionizing radiation
- 3e12 Neq/cm^2 non-ionizing radiation

- 1. need to improve charge collection, radiation tolerance-> high resistivity epitaxial layer
- 2. improvement of radiation tolerance of pixel design

#### Mimosa26 sensor in a high-resistivity epitaxial layer

#### Efficiency vs Threshold



 $\cdot$  50 MHz to emulate longer integration time

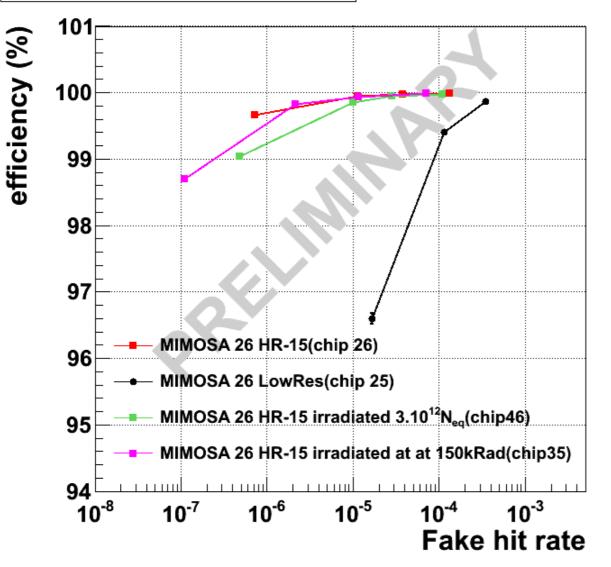
+35 C temperature!

resolution < 5um

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#### Mimosa26 sensor in a high-resistivity epitaxial layer





• 50 MHz to emulate longer integration time

• +35 C temperature !

resolution < 5um

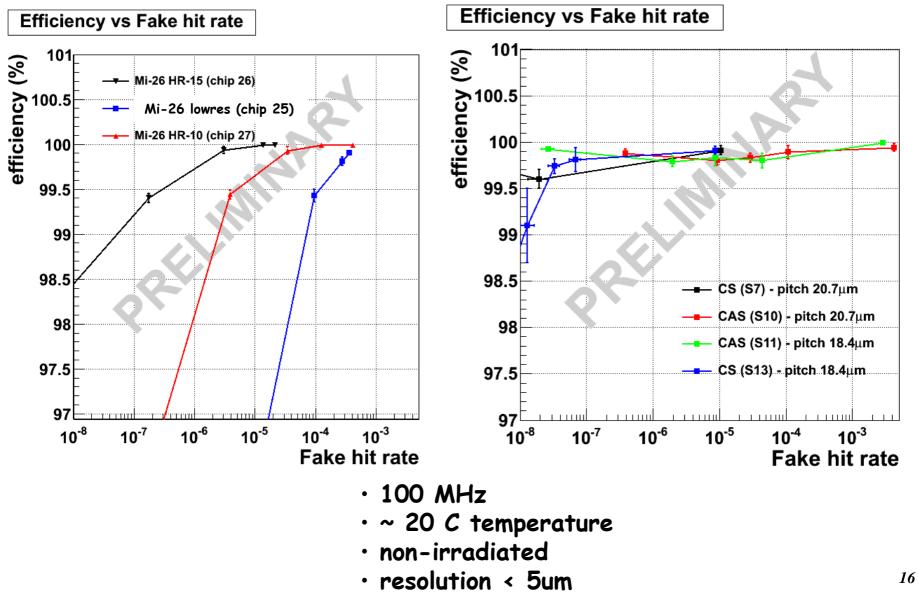
# Prototyping for optimization: Mimosa22ahr – tests of several pixel variants

- pitch 20.7 um and 18.4 um
- few diode sizes , from 11 um<sup>2</sup> to 18 um<sup>2</sup>
- radiation tolerant designs (ELT)
- diode bias voltage increase
- two stage amplifiers, AC coupling

Lab test with Fe55 at +35 C and integration time imposed by the STAR requirements shows:

- $\cdot$  the conversion gain can be improved by factor of two
- $\cdot$  noise before irradiation 10 to 13 e
- $\cdot$  noise after irradiation with 150 kRad is from 13 to 23 e
- $\cdot$  noise after irradiation with 3e12 Neq is from 16 to 23 e
- $\cdot$  SNR up to 30 after irradiation
- for a low resistivity epitaxial layer ~30 before irradiation

#### Mimosa22ahr - beam tests (CERN-SPS)



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### Conclusions

• The high resistivity epitaxial layer improves signal-to-noise ratio of MAPS, confirmed in a beam test with Mimosa26 -> efficiency >99.8% (at 10e-5 fake hit rate)

-> In particular the radiation tolerance is highly improved (NIx100)

•The radiation hardness can still be improved by pixel design, first lab results of Mimosa22ahr + beam test measurements are in progress..

• PIXEL detector for the STAR experiment based on a modified version of Mimosa26 Sensors built in a high resistivity epitaxial layer will fulfill the STAR HFT requirement

-> ULTIMATE sensor is going to be submitted for fabrication autumn 2010