

# Multicomponent dark matter and the dark photon portal

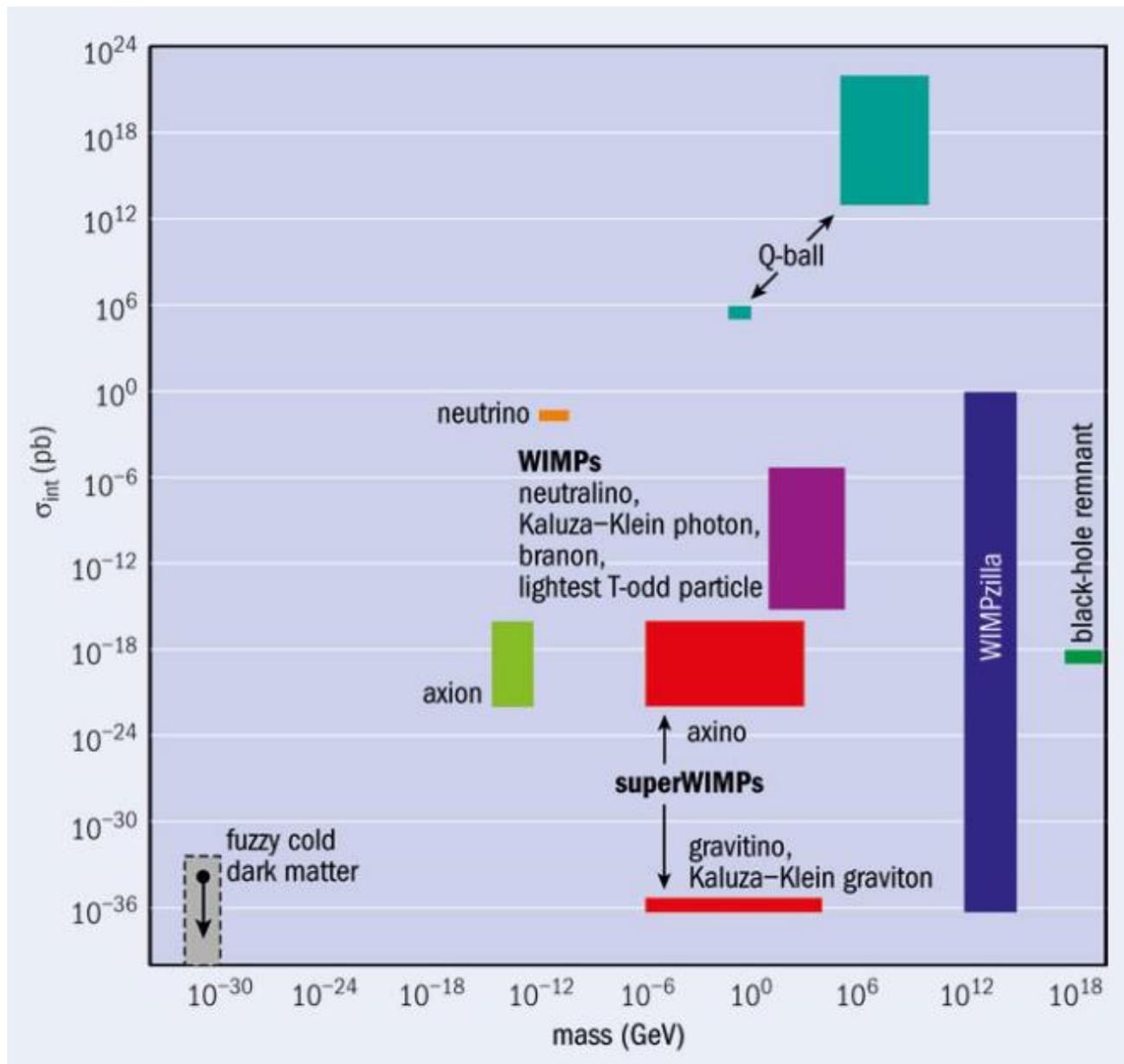
**By: Amalia Betancur Rodríguez**

**In collaboration with Andrés Fernando Castillo, Guillermo Palacio and Juan  
Guillermo Suárez.**

# Outline:

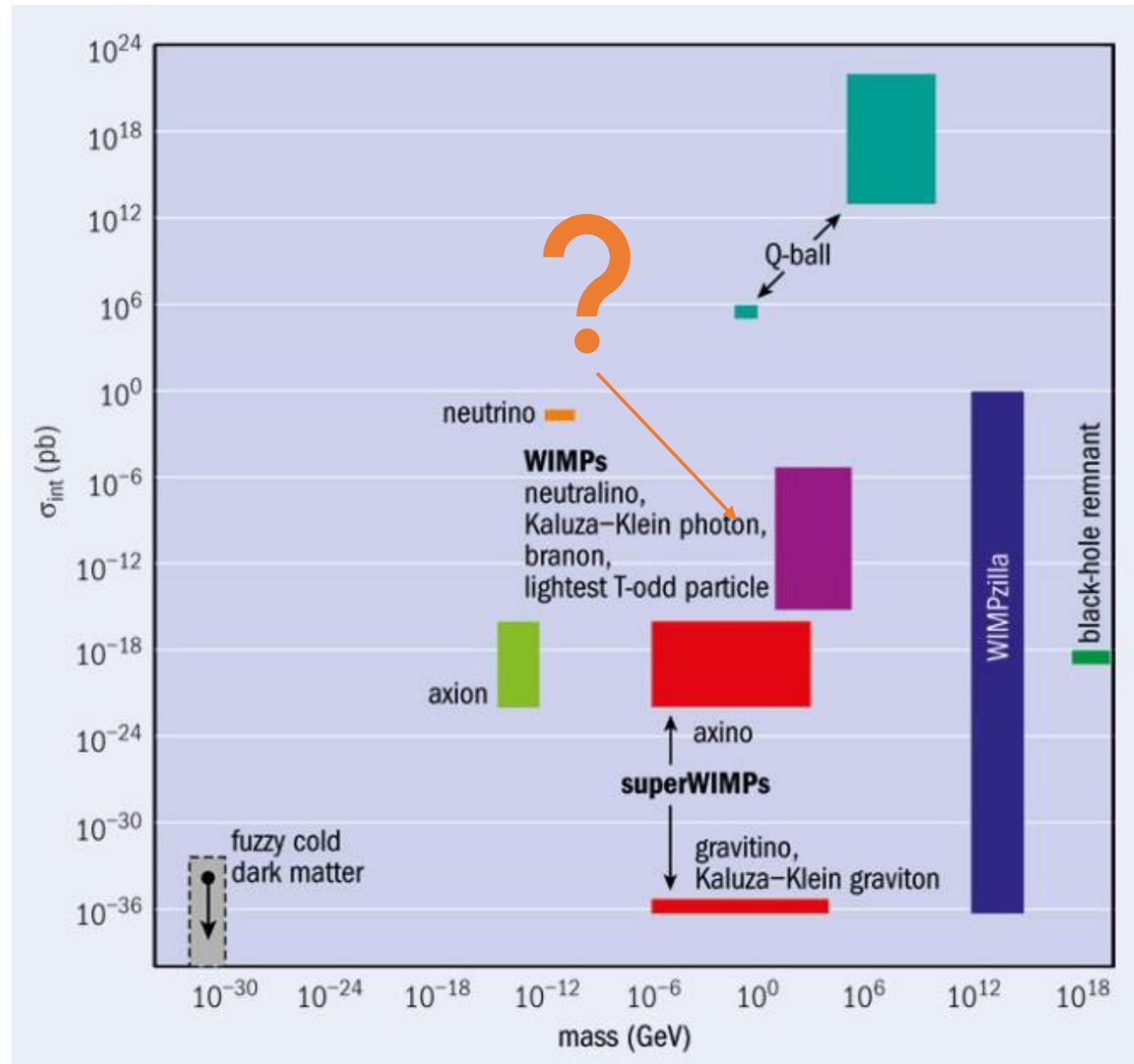
- Motivation
- The model.
- Relic density phenomenology.
- Probing the model.

# Motivation

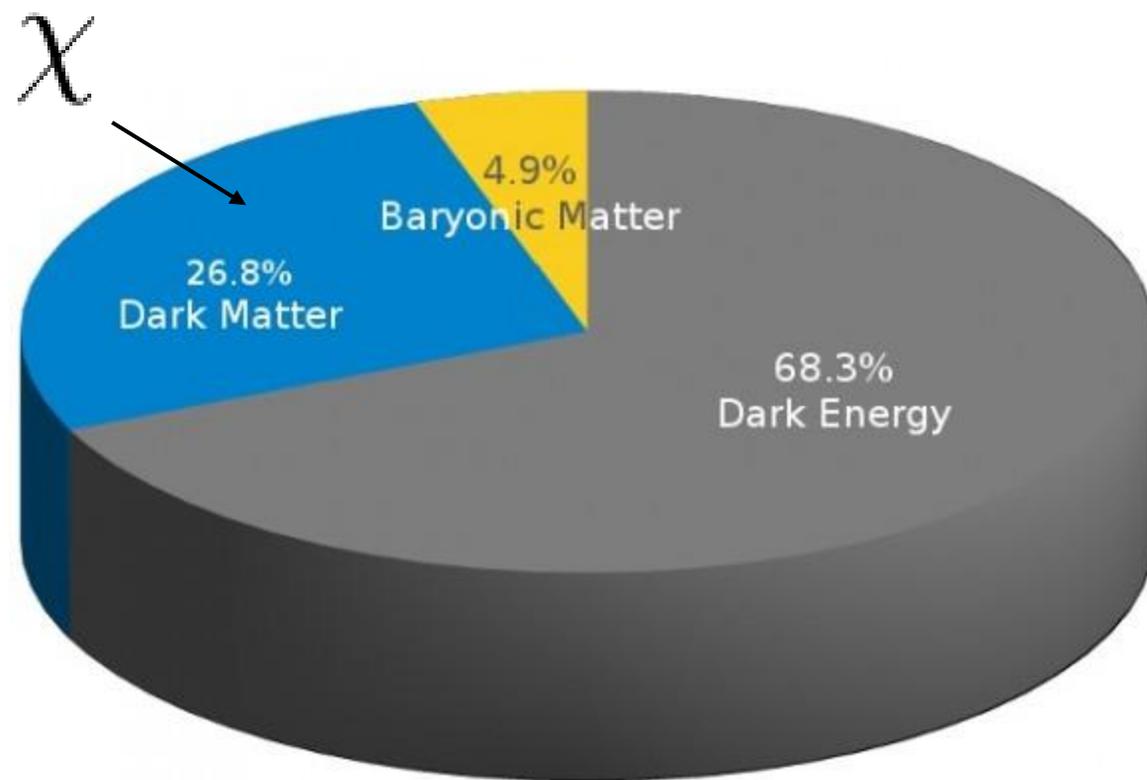


# Motivation

WIMP Dark matter candidates could be below the usual mass scales, but a new annihilation channel is needed.

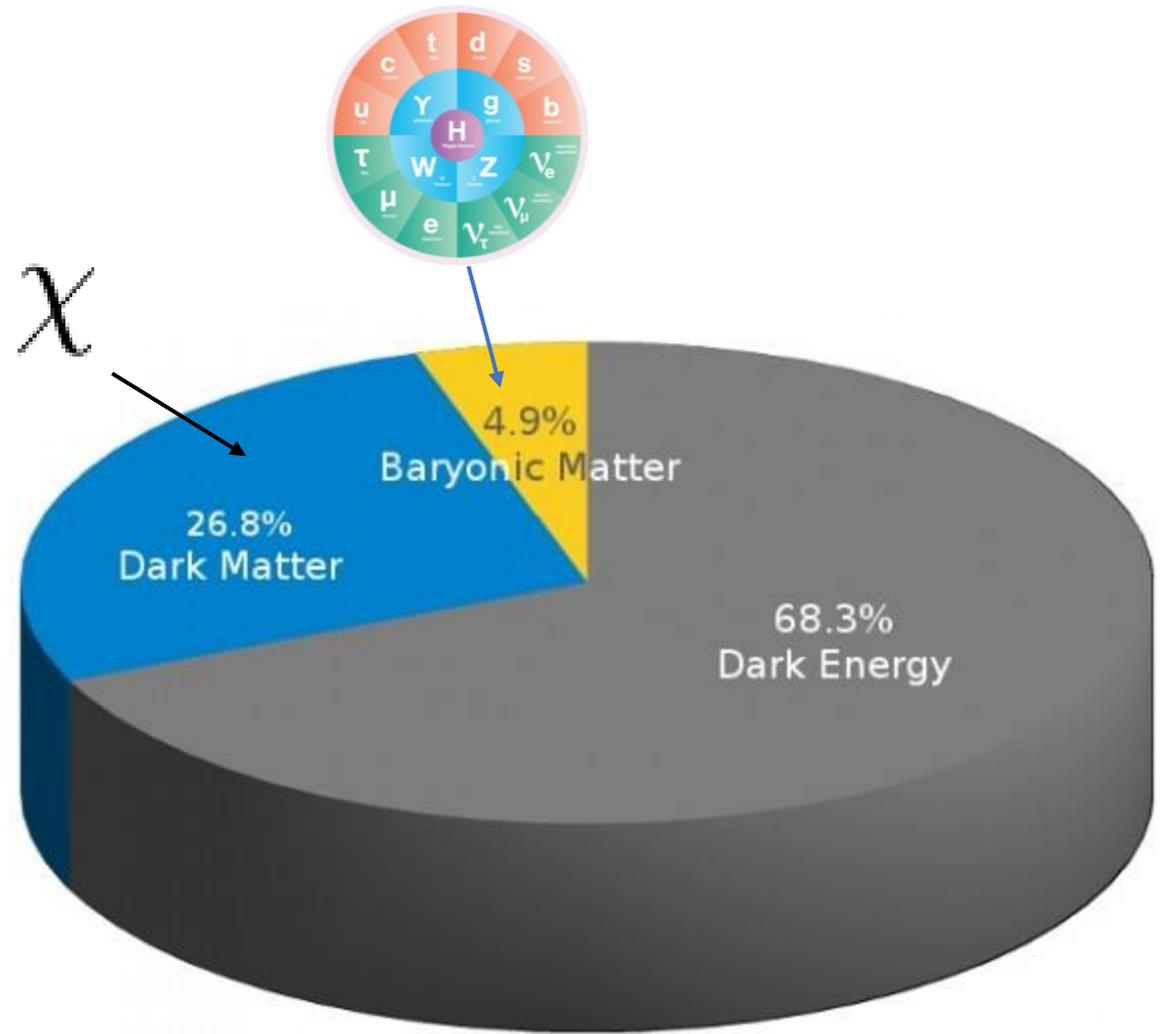


Motivation



# Motivation

Many models focus on one DM candidate, but there could be several, just as the baryonic matter is made of a myriad of particles.



# The Model:

To address the two issues mentioned, we focus on a multicomponent DM model with lighter WIMP particles (SubGeV). The model includes:

Added particle	Symmetries	Spin
$\phi_1$	$U(1)_{D_1} \times Z_2$	0
$\phi_2$	$U(1)_{D_2} \times Z'_2$	0

The gauge Lagrangian associated to the U(1) symmetries:

$$\mathcal{L}_{gauge} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} - \frac{\epsilon}{2}F_{\mu\nu}F'^{\mu\nu}$$

This part of the model based on Pospelov et al 2008, and Pospelov 2009

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The dark U(1) symmetry is broken, thus there is a massive gauge boson, the dark photon.

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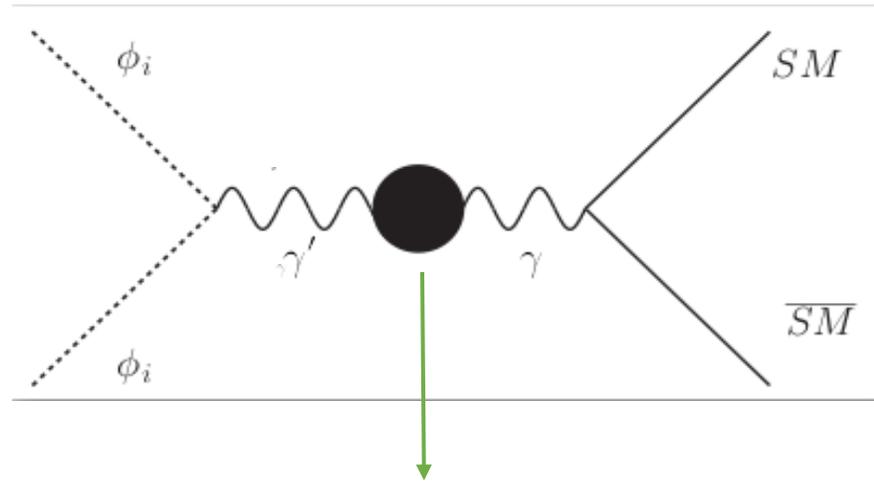
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There is a kinetic mixing between the photon and the dark photon

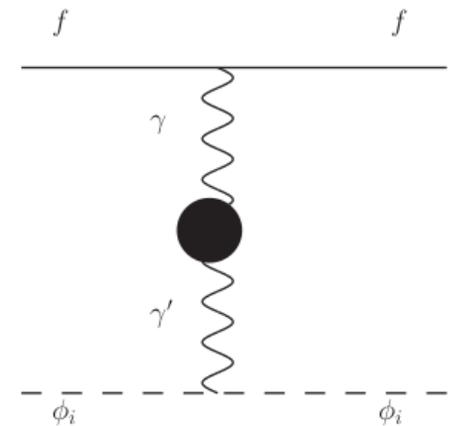
# The Model: Feynman diagrams

DM annihilation into SM

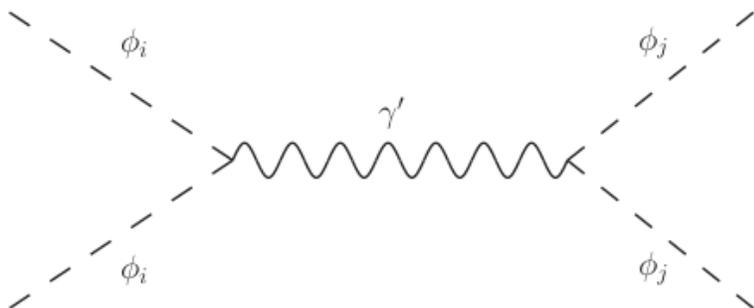


Kinetic Mixing

T-channel interaction with SM



DM conversion



To study the interactions we use SARAH, based on the model of D. Aristizabal et al. 2015

# Relic density phenomenology

Free parameters:

$\epsilon$ ,  $\alpha_D$ ,  $m_{\gamma'}$ ,  $m_{\phi_1}$ , and  $m_{\phi_2}$

**To calculate the relic density, we use micrOmegas.**

Choose:

$\alpha_D=0.1$ ,  $m_{\gamma'}=3m_{\phi_1}$ ,  $m_{\phi_2} > m_{\phi_1}$

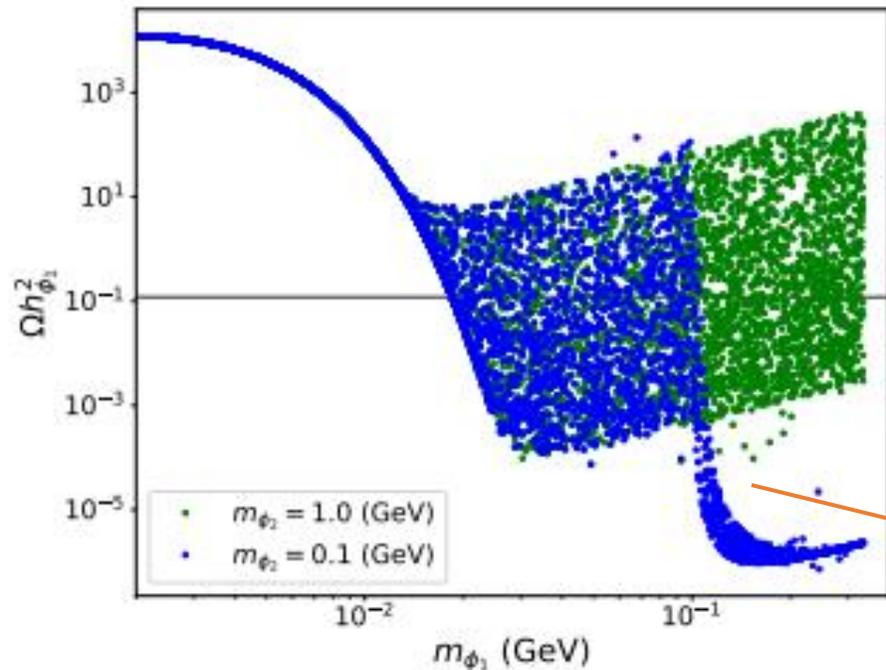
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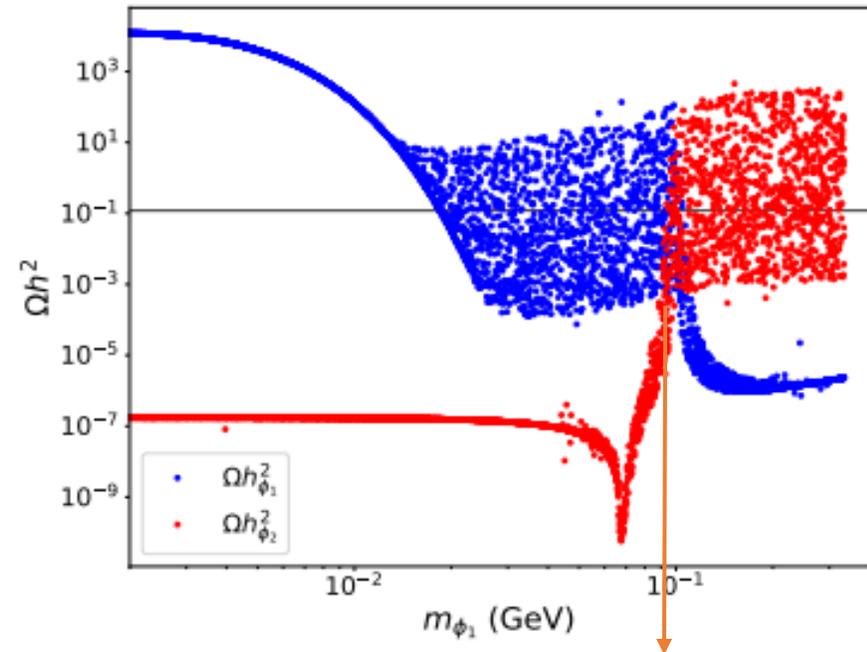
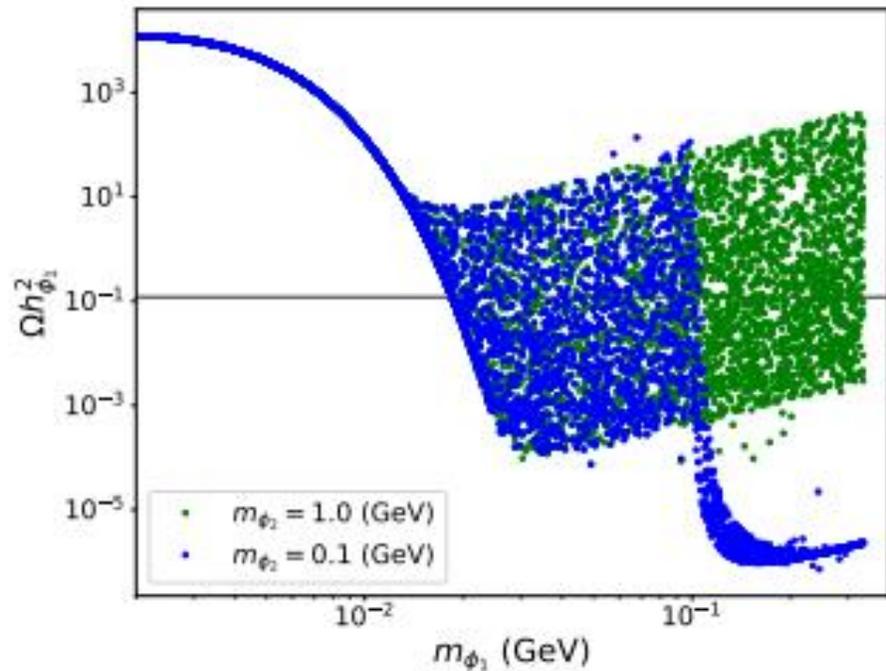
Benchmark similar to  
D.Romeri et al,  
2019 and M.  
Breitbach et al 2021.

Heavier DM gets depleted!

# Relic density phenomenology

Choose:

$$\alpha_D = 0.1, m_{\gamma'} = 3m_{\phi_1}, m_{\phi_2} > m_{\phi_1}$$

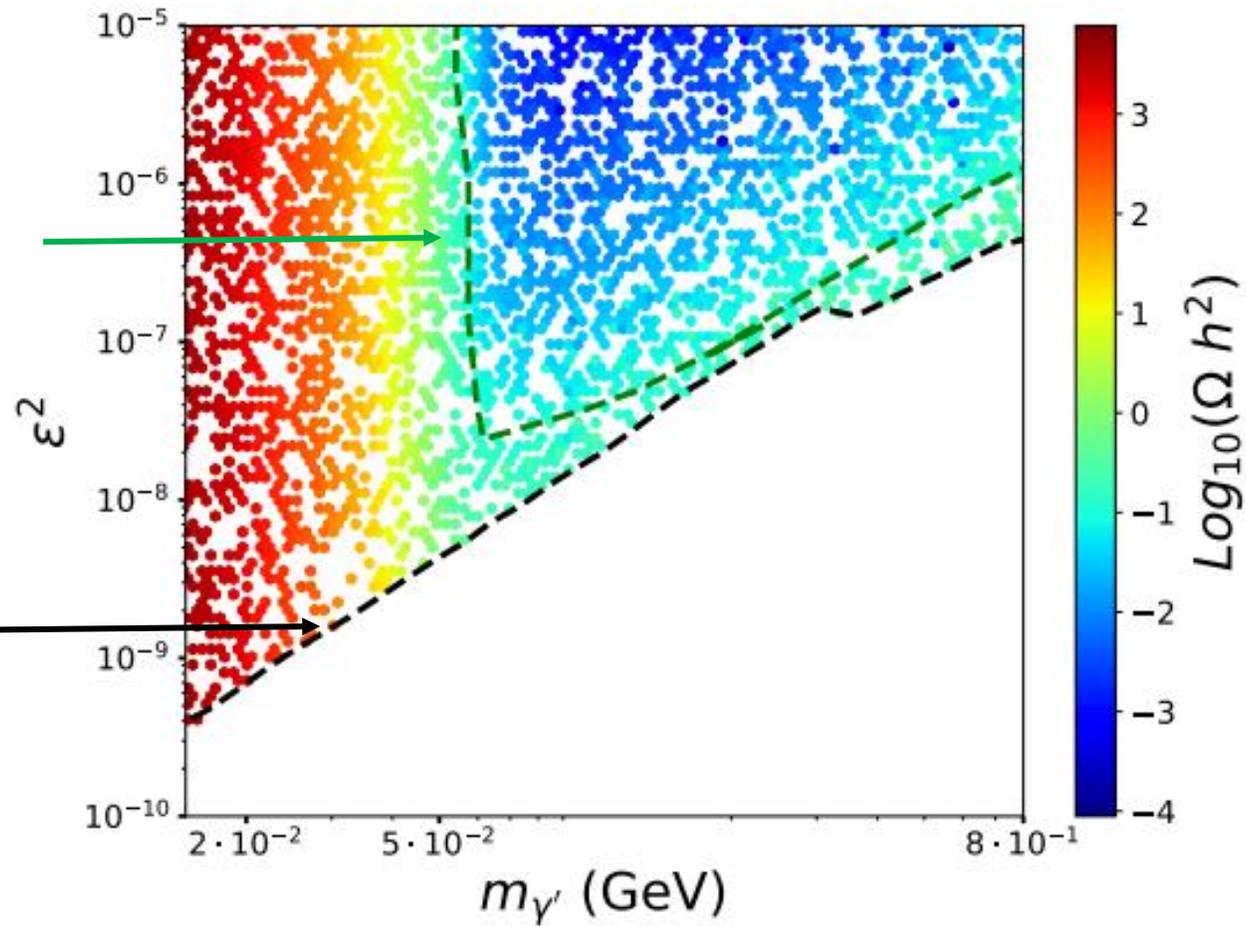


DM particles switch roles,  
only one DM particle remain,  
the heaviest

# Relic density phenomenology

Relic abundance for the multicomponent model.

Correct relic abundance for one DM particle, such as the model studied in D. Romeri et al 2019.

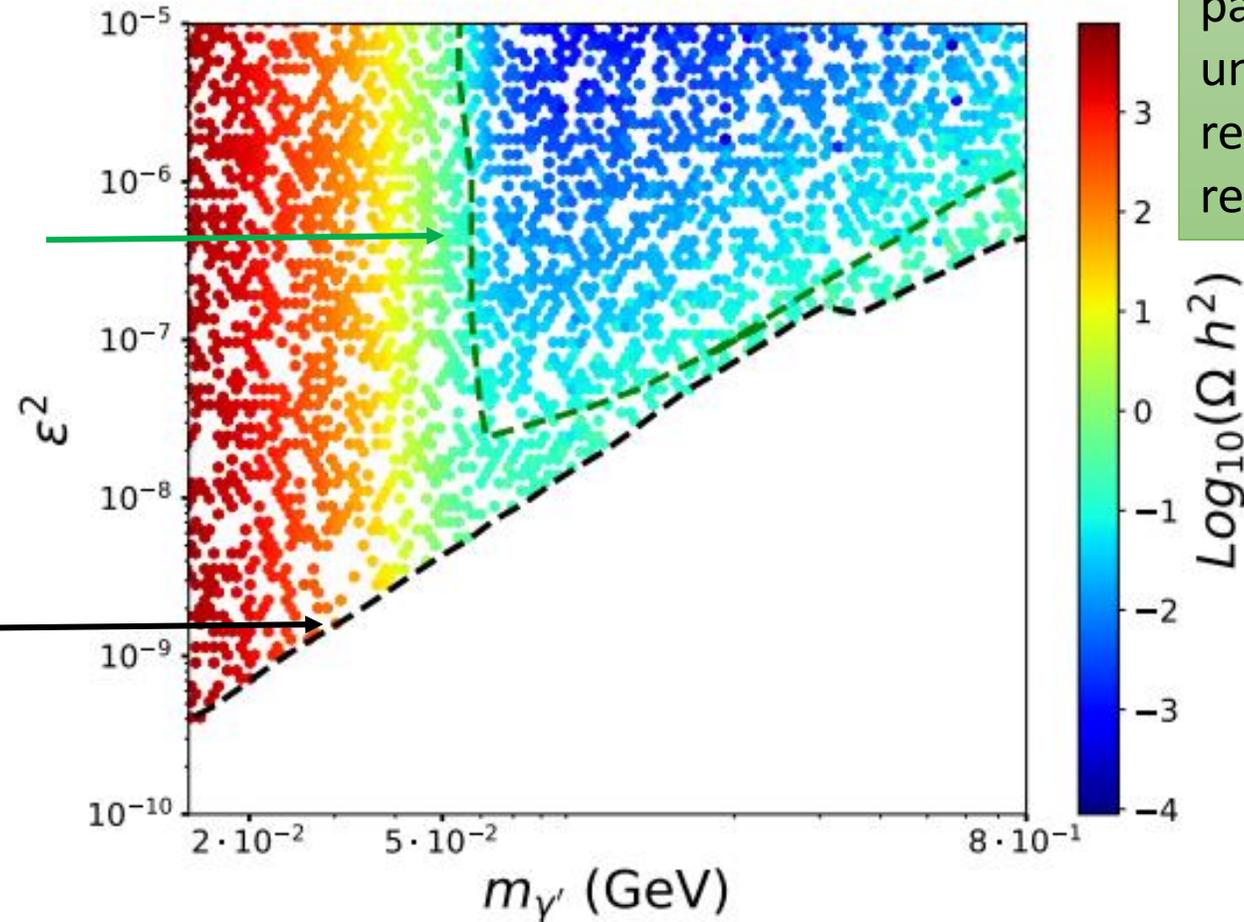


# Relic density phenomenology

Even though only one DM remains, the existence of the other particle in the early universe, affects the relic density of the remaining particle

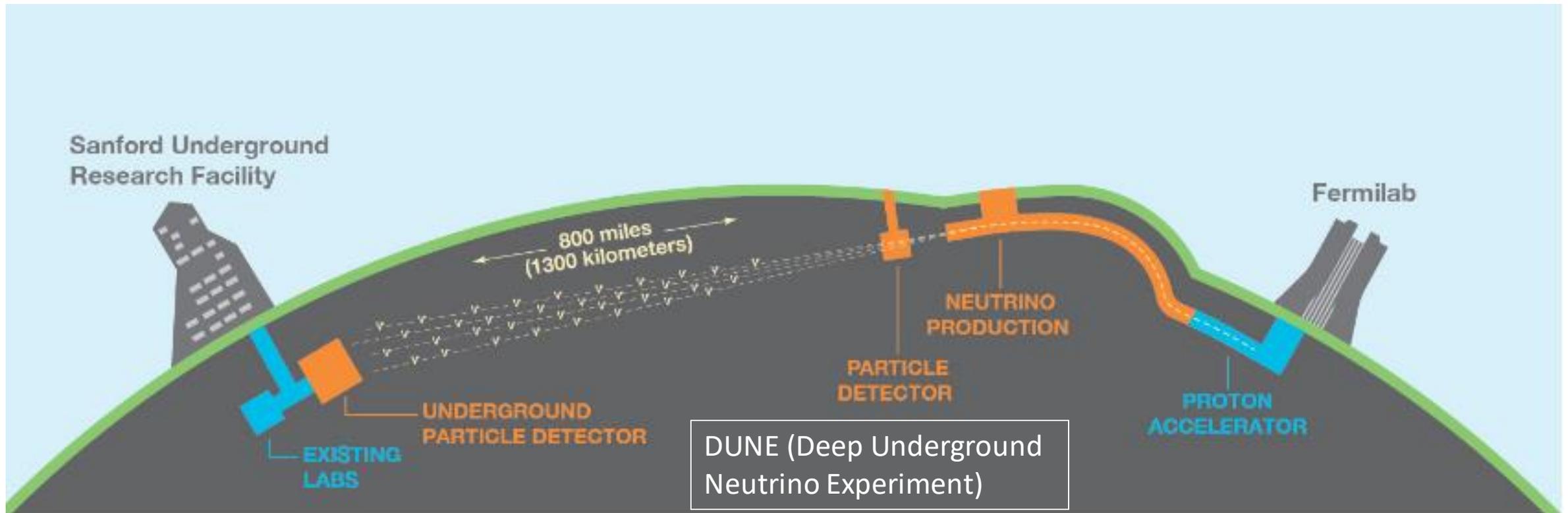
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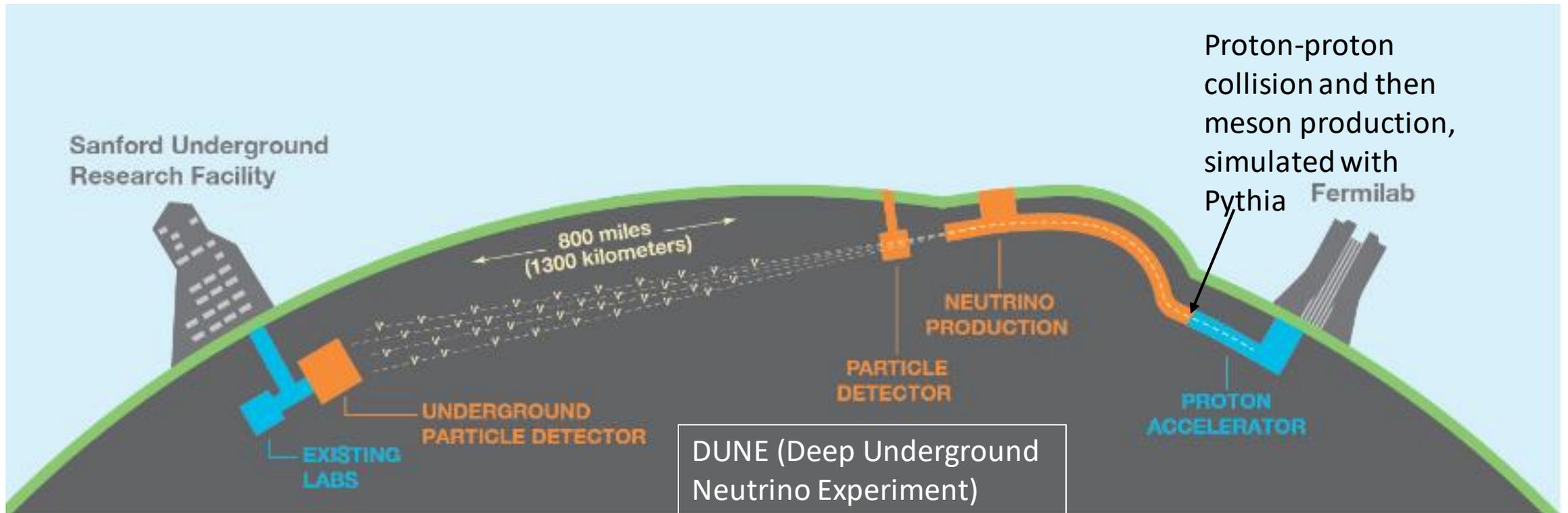
# Probing the model: DUNE

- High intensity proton beams present an opportunity to produce and observe the interactions of these DM particles.

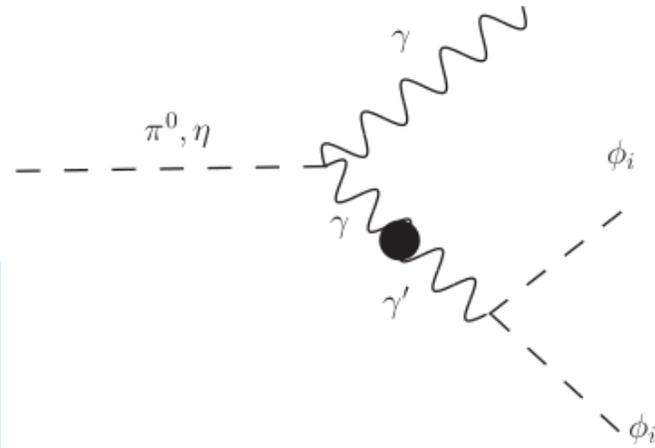


# Probing the model: Meson production

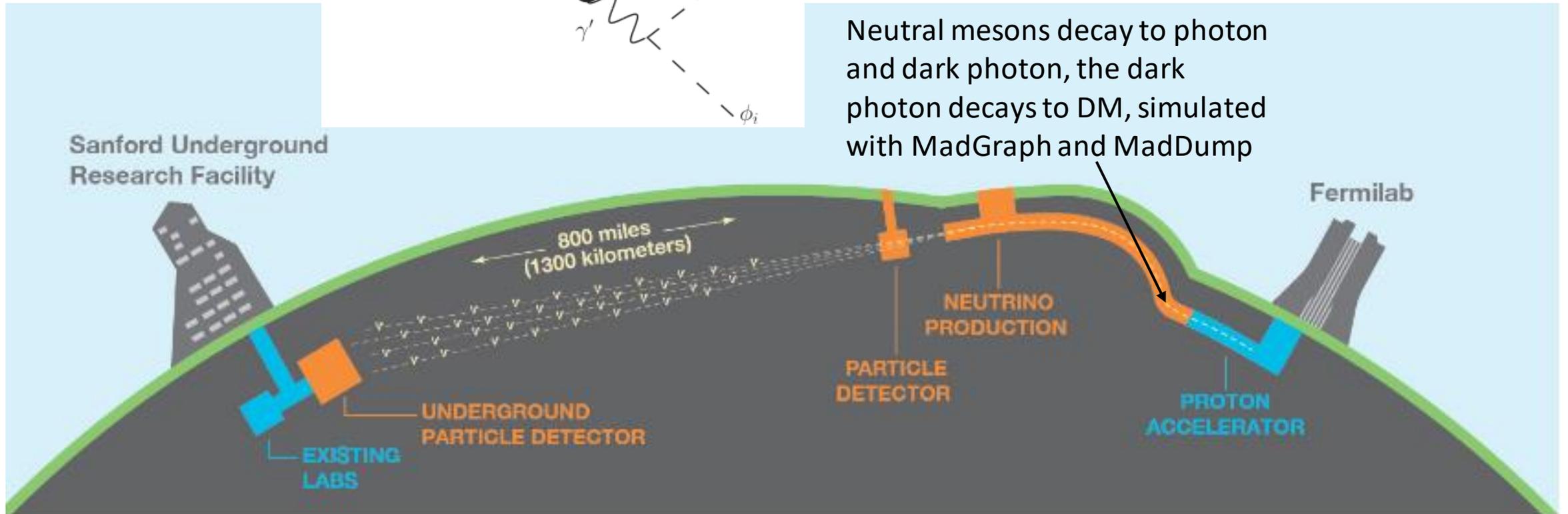
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# Probing the model: Meson decay, DM production

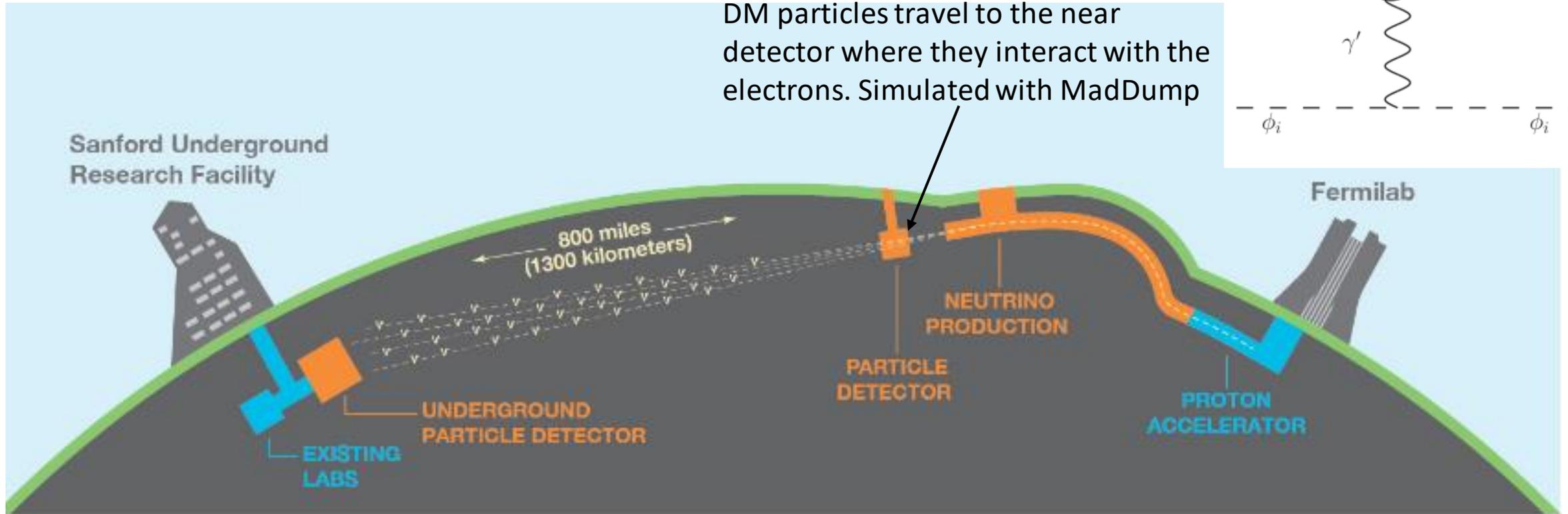
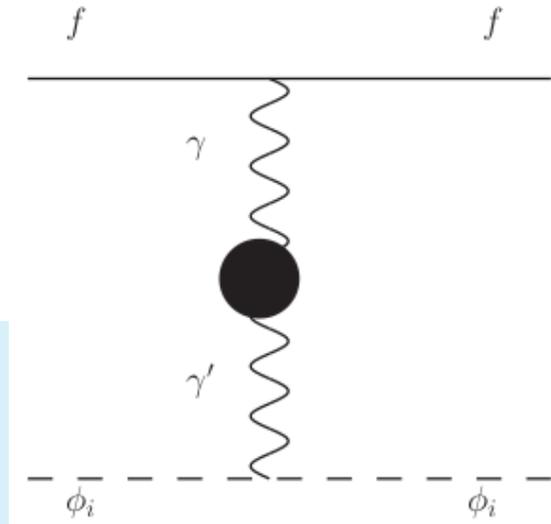


Neutral mesons decay to photon and dark photon, the dark photon decays to DM, simulated with MadGraph and MadDump



# Probing the model: DM interaction

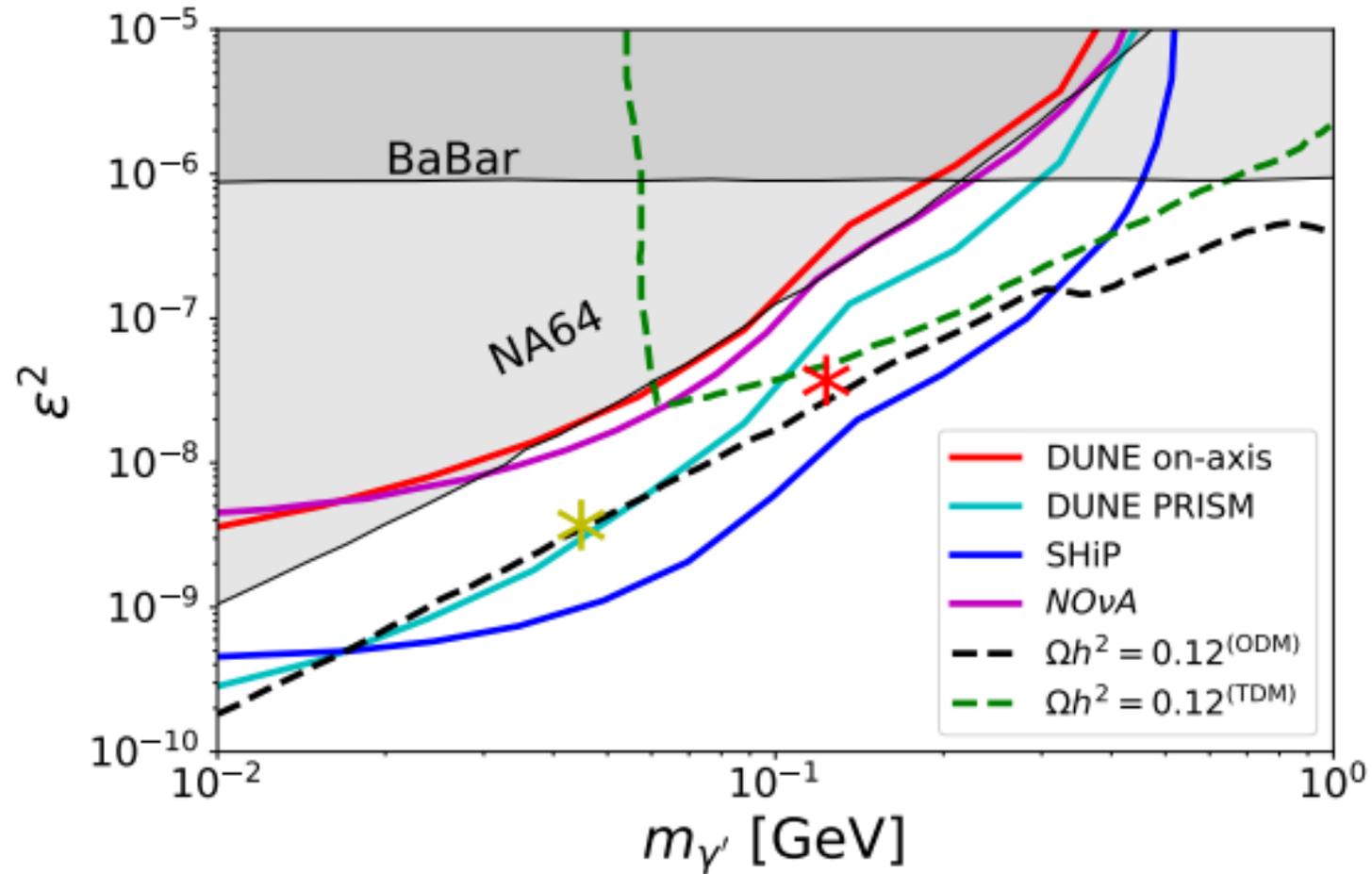
DM particles travel to the near detector where they interact with the electrons. Simulated with MadDump



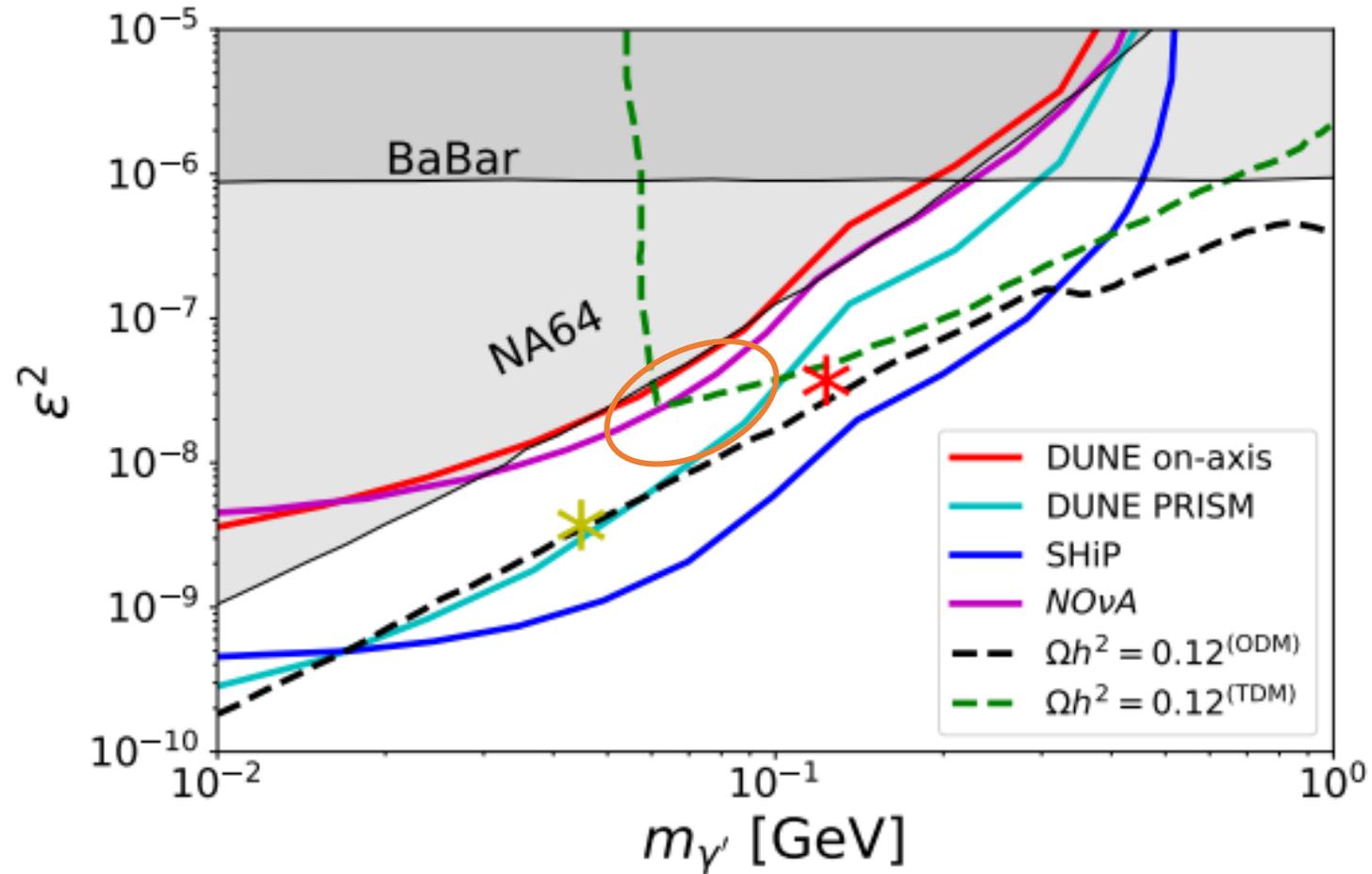
# Probing the model:

- Through this process (proton-proton collision and meson decay) neutrinos are also produced, **huge background!** To simulate it, we used NuWro ([more at NuCo](#)).
- To estimate the parameter space that DUNE may probe, [we analyze the kinematic shape of both: background and DM+background. For this we used a Poisson likelihood, following a procedure similar to D.Romeri et al, 2019 and M. Breitbach et al 2021.](#)

# Probing the model: Results



# Probing the model: Results



# Conclusion:

- In the multicomponent subGeV DM model, the dark photon portal plays a very important role.
- It allows for DM conversion, which is important in setting the relic abundances.
- The model may be probed at DUNE (and also SHIP) with a better exclusion of the parameter space by using the near detector off-axis configuration.

# References:

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