

Background Material for Tuning Discussion

Stephen Mrenna
Fermilab

April 13, 2021

Some CMS Tuning Activities

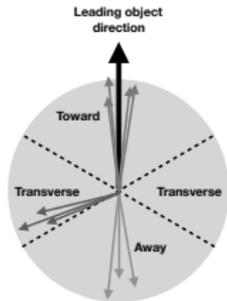
- ★ Both ATLAS and CMS have dedicated tuning groups
- ★ Review process tends to be long and tedious
- ★ Production of new tunes driven by disagreements with data either in UE-observables or some high- p_T matched samples
- ★ Slides shown here from CMS talk at DIS2019

15 Second Overview

- ★ CUETP8M1: Monash + PDF change
- ★ One parameter tweaks:
 - ★ CUETP8M2T4: $\alpha_s(ISR) \neq \alpha_s(FSR)$
 - ★ Jet substructure: $\alpha_s(FSR) < 0.13$
- ★ 6-parameter CPX tunes: match PDF and α_s in ME

Comparisons of predictions for UE observables from previous tunes to 13 TeV data

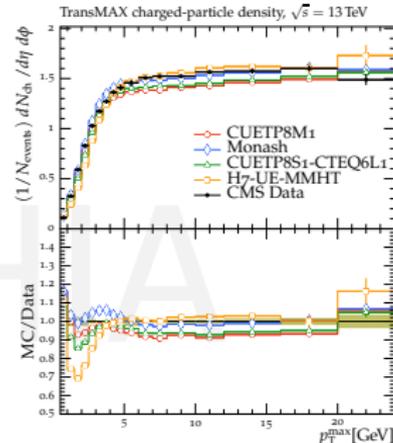
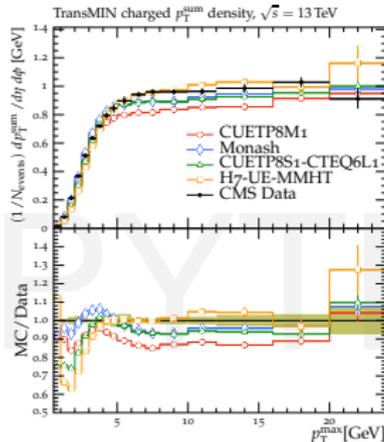
arXiv:1903.12179



$$p_T > 0.5 \text{ GeV}$$

$$|\eta| < 0.8$$

$$p_T^{\text{max}} = \max(p_T^{\text{trk},i})$$



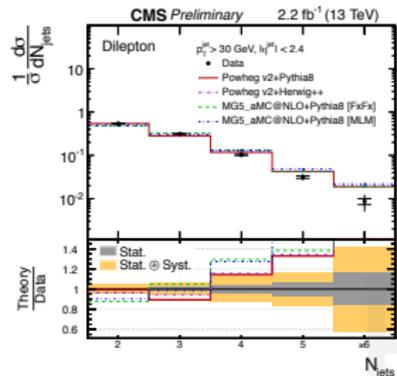
- Main CMS tune (CUETP8M1 based on Monash tune) used until 2017 analyses does not describe well the central values of the data at 13 TeV. [EP J.C. 76 \(2016\) 155](#)
- CUETP8M1: α_s and shower parameters kept as in Monash $\rightarrow \alpha_s^{\text{ISR/FSR}} = 0.1365$ despite the preferred values of 0.130 in LO and 0.118 in NLO matrix elements/ PDF sets.
 - α_s^{FSR} in Monash \rightarrow by fitting Pythia8 predictions to LEP event shapes and α_s^{ISR} is just assumed to be the same as α_s^{FSR} .
 - $\alpha_s^{\text{MPI}} = 0.130$ set to the value preferred in the LO PDF set.

Revisiting Shower Parameters and Tunes

- Starting from parton shower in $t\bar{t}$ events
→ CUETP8M2T4 tune ($\alpha_s^{ISR} \sim 0.11$) [CMS-PAS-TOP-16-021](#)
- UE in $t\bar{t}$ events at 13 TeV
→ $\alpha_s^{FSR} \sim 0.118$ agrees better with data. [EPJC 79 \(2019\) 123](#)
- Jet substructure in $t\bar{t}$ events at 13 TeV
→ $\alpha_s^{FSR} \sim 0.115$. [PRD98 \(2018\) 092014](#)
- New CMS tunes using (N)(N)LO PDF sets in PS
→ CPX tunes (consistent treatment of PDF+ α_s in matrix element and parton shower) [arXiv:1903.12179](#)
- UE in Z+jets events at 13 TeV [JHEP07 \(2018\) 032](#) and [arXiv:1903.12179](#)

CUETP8M2T4 Event Tune

CMS-PAS-TOP-16-021

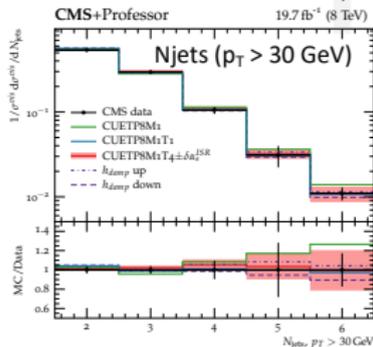


- CUETP8M1 not only bad in describing the UE but its predictions overshoot the data for large jet multiplicities when out of the box parameters are used (in Monash-based tunes: $\alpha_s^{\text{ISR}}=0.1365$)
- Effect also observed with 8 TeV data.

Tune α_s^{ISR} using 8 TeV $t\bar{t}$ N_{jets} (using the parton-shower dominated region) and $t\bar{t}$ jet p_T data →

$$\alpha_s^{\text{ISR}} = 0.1108^{+0.0145}_{-0.0142}$$

$$h_{\text{damp}} = 1.581^{+0.658}_{-0.585} m_t$$



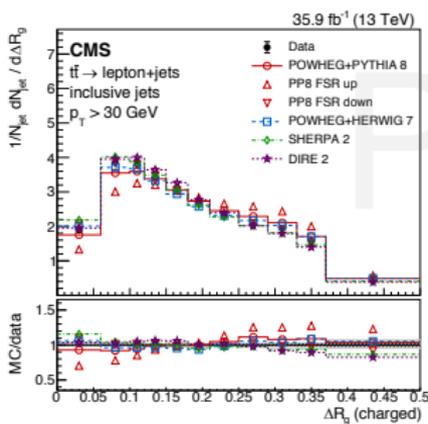
SpaceShower:RapidtyOrdering=on

- Significantly lower α_s^{ISR} cures the overshoot of CUETP8M1 at high jet multiplicities.
- UE event tune starting with fixed lower α_s^{ISR} describes the UE & min-bias (and top quark) significantly data better.

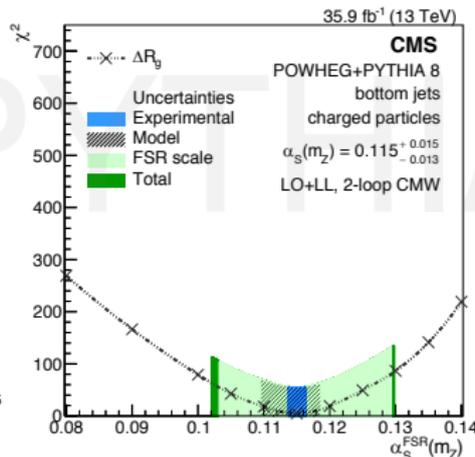
α_s^{FSR} from jet substructure in $t\bar{t}b$ l+jets events

- Measured using charged+neutral and with only charged jet constituents (particle $p_T > 1$ GeV).
- b, light, or gluon jet enriched samples.

PRD98 (2018) 092014



Angle between groomed subjets at particle level (correlated to jet width)



Angle between groomed subjets at particle level.

Pythia8:
 CUETP8M2T4 for $t\bar{t}b$
 CUETP8M1 for the rest.

None of the default tunes yield a good overall description of the data.

Powheg+Pythia8 with LO+LL, 2-loop :

$$\alpha_s^{FSR}(M_Z) = 0.115^{+0.015}_{-0.013}$$

New CMS Tunes using LO PDF

[arXiv:1903.12179](https://arxiv.org/abs/1903.12179)

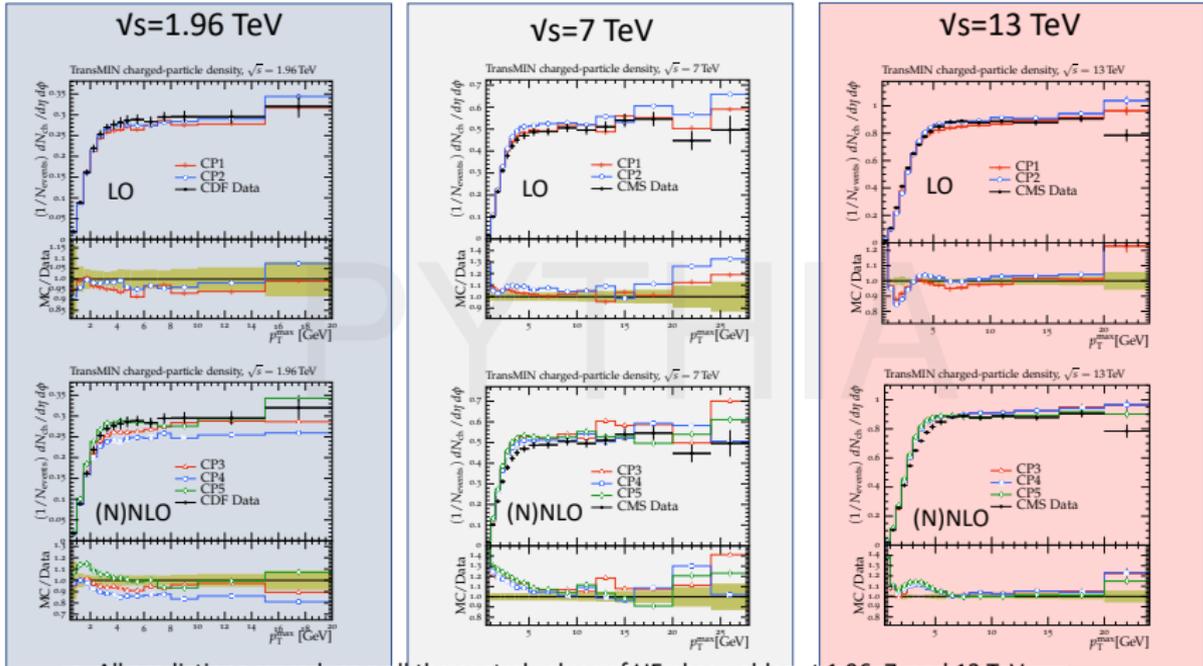
| PYTHIA8 parameter | CP1 | CP2 | |
|--|-------------|-------------|-------------------|
| PDF Set | NNPDF3.1 LO | NNPDF3.1 LO | Fixed inputs |
| $\alpha_S(m_Z)$ | 0.130 | 0.130 | |
| SpaceShower:rapidityOrder | off | off | |
| MultipartonInteractions:EcmRef [GeV] | 7000 | 7000 | |
| $\alpha_S^{\text{ISR}}(m_Z)$ value/order | 0.1365/LO | 0.130/LO | |
| $\alpha_S^{\text{FSR}}(m_Z)$ value/order | 0.1365/LO | 0.130/LO | |
| $\alpha_S^{\text{MPI}}(m_Z)$ value/order | 0.130/LO | 0.130/LO | |
| $\alpha_S^{\text{ME}}(m_Z)$ value/order | 0.130/LO | 0.130/LO | Fitted parameters |
| MultipartonInteractions:pT0Ref [GeV] | 2.4 | 2.3 | |
| MultipartonInteractions:ecmPow | 0.15 | 0.14 | |
| MultipartonInteractions:coreRadius | 0.54 | 0.38 | |
| MultipartonInteractions:coreFraction | 0.68 | 0.33 | |
| ColorReconnection:range | 2.63 | 2.32 | |
| χ^2/dof | 0.89 | 0.54 | |

New CMS Tunes using (N)NLO PDFs

arXiv:1903.12179

| PYTHIA8 parameter | CP3 | CP4 | CP5 | |
|--|--------------|---------------|---------------|-------------------|
| PDF Set | NNPDF3.1 NLO | NNPDF3.1 NNLO | NNPDF3.1 NNLO | Fixed inputs |
| $\alpha_S(m_Z)$ | 0.118 | 0.118 | 0.118 | |
| SpaceShower:rapidityOrder | off | off | on | |
| MultipartonInteractions:EcmRef [GeV] | 7000 | 7000 | 7000 | |
| $\alpha_S^{\text{ISR}}(m_Z)$ value/order | 0.118/NLO | 0.118/NLO | 0.118/NLO | |
| $\alpha_S^{\text{FSR}}(m_Z)$ value/order | 0.118/NLO | 0.118/NLO | 0.118/NLO | |
| $\alpha_S^{\text{MPI}}(m_Z)$ value/order | 0.118/NLO | 0.118/NLO | 0.118/NLO | |
| $\alpha_S^{\text{ME}}(m_Z)$ value/order | 0.118/NLO | 0.118/NLO | 0.118/NLO | |
| MultipartonInteractions:pT0Ref [GeV] | 1.52 | 1.48 | 1.41 | Fitted parameters |
| MultipartonInteractions:ecmPow | 0.02 | 0.02 | 0.03 | |
| MultipartonInteractions:coreRadius | 0.54 | 0.60 | 0.76 | |
| MultipartonInteractions:coreFraction | 0.39 | 0.30 | 0.63 | |
| ColorReconnection:range | 4.73 | 5.61 | 5.18 | |
| χ^2/dof | 0.76 | 0.80 | 1.04 | |

CPX and Energy Dependence



- All predictions reproduce well the central values of UE observables at 1.96, 7, and 13 TeV.
- LO-PDF tunes slightly better in describing the energy dependence.

CPX and Z+Jets

arXiv:1903.12179

Matrix elements:

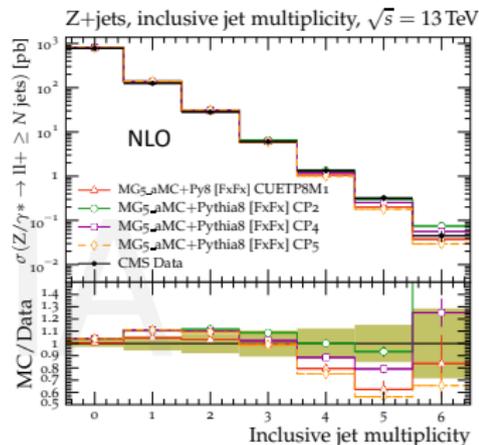
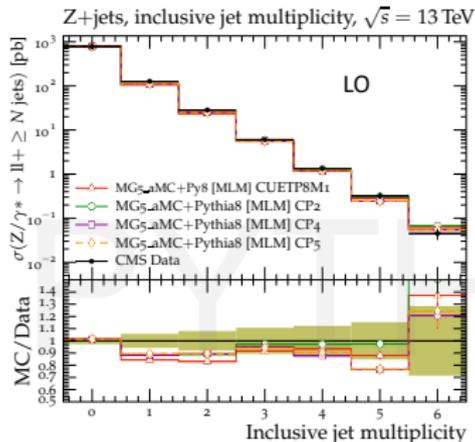
- LO: ≤ 3 partons
 - NLO: ≤ 2 parton at NLO 3rd at LO
 - NNPDF3.1 NNLO
- $\alpha_s(m_Z)=0.118$ for both cases

$$p_T^{\ell\ell} > 20 \text{ GeV}, |y^{\ell\ell}| < 2.4,$$

$$|m^{\ell\ell} - 20| < 91 \text{ GeV}$$

$$p_T^{\text{jet}} > 30 \text{ GeV},$$

$$|y^{\text{jet}}| < 2.4$$



- Little sensitivity to the tune for low multiplicities.
- All tunes describe the central values of Njets reasonably well.
- CP2 has a slightly better description of the central values.
- CUETP8M1 and CP5 undershoot the data at the PS dominated region with at least 4 jets.

Not Mentioned here

Still internal stuff, including

- ★ CR tunes
- ★ *in situ* fit to b-fragmentation in $t\bar{t}$
- ★ ...

ATLAS A14

6.5 years old!

- ★ 10 parameter tune
- ★ Roughly 10 classes of data, 400+ histograms, 7000-ish bins
- ★ Used Professor framework: built surrogate model (polynomial) of Pythia8 predictions based on smallish number of parameter samples
- ★ Aimed at standalone Pythia8 tune good enough for search limits
- ★ Used weighting procedure to adjust importance of various distributions to arrive at certain behavior: very tedious and crying out for automation
- ★ Provided eigentunes
- ★ Fairly well-documented, and even some of the Yoda files were still around

BROOD: Bilevel and Robust Optimization and Outlier Detection for Efficient Tuning of High-Energy Physics Event Generators

<https://arxiv.org/abs/2103.05751>

- ★ Output of a SciDAC project with the thought of doing tuning on HPCs
- ★ However: this work did not rely on that
- ★ Bottom line:
 - ★ One can develop several bi-level algorithms that select weights, minimize, and then iterate
 - ★ These can produce tunes that are comparable or outperform expert tunes
- ★ Evolution of Professor to Apprentice See:
<https://arxiv.org/abs/2103.05748>
 - ★ more choices of optimizers
 - ★ rational approximation surrogate models

Rest is thoughts and comments

WHAT DO WE WANT AN UNCERTAINTY TUNE TO BE?

Can we constructively criticize eigen-tune method and develop an alternative?

What is really χ^2 distributed anyways?

What are the correct degree of freedom? or is it meaningful or necessary to introduce them?

I am pragmatic – don't care about mathematical proofs as long as method is useful and not too conservative

UTOPIAN TUNES

Fit one histogram at a time (yes, even sometimes where $N_{\text{bins}} \sim N_{\text{params}}$)

Surprisingly, this happens. So what?

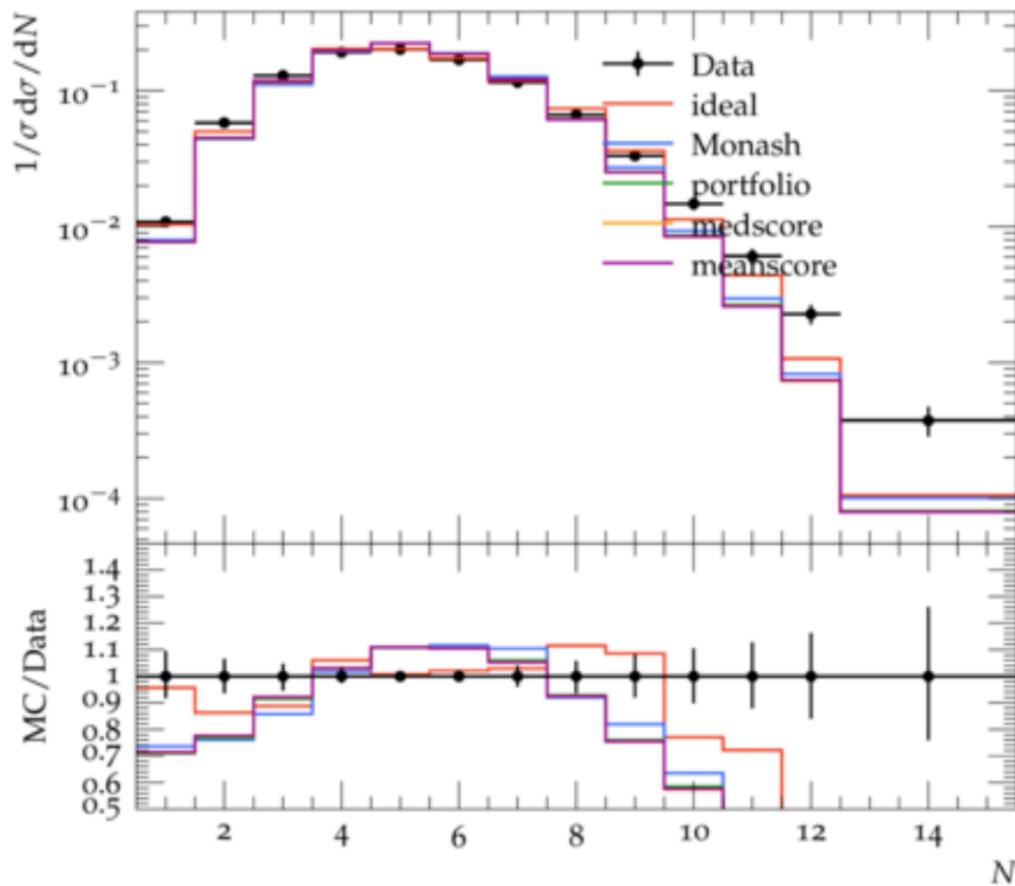
There is a gap between the prediction and data

Build a model of the gap using ML

Can be used to supplement predictions and provide additional uncertainty estimate

Could be expressed symbolically to develop new models

Charged jet multiplicity (anti- k_t , $R = 0.4$, y 0.0-1.9, p_{\perp} 10.0-15.0)



No bounds

With surrogate models and HPCs, we can expand the number of parameters we vary at the LHC

Can/should try to tune hadronization parameters using LHC data

ALICE flavor and hadronization measurements

<https://arxiv.org/abs/1709.08522v1> (no Rivet analysis) Fig 1

<https://arxiv.org/abs/1802.09145v1> (no HepData) Fig 7

<https://arxiv.org/abs/1708.08745v1> (HepData and Rivet available)

<https://arxiv.org/abs/1807.11186> (no HepData)

<https://arxiv.org/abs/1807.11321> (no HepData)

<https://arxiv.org/abs/1811.01535>

arXiv:1011.5531 [pdf, other] Charged particle multiplicities in pp interactions at $\sqrt{s} = 0.9, 2.36, \text{ and } 7 \text{ TeV}$

arXiv:1012.5104 [pdf, ps, other] Charged-particle multiplicities in pp interactions measured with the ATLAS detector at the LHC

arXiv:1307.1094 [pdf, ps, other] Multiplicity dependence of the average transverse momentum in pp, p-Pb, and Pb-Pb collisions at the LHC

Concluding Remarks

- ★ I think we should be in the business of making tunes
- ★ We have expertise on many fronts, including contacts in the relevant experiments
- ★ We can make some use of the tools and computational resources from my SciDAC project
- ★ Providing tunes that fit into the HL-LHC is my primary interest