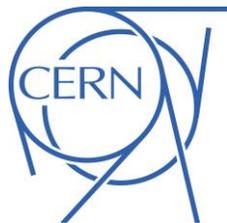
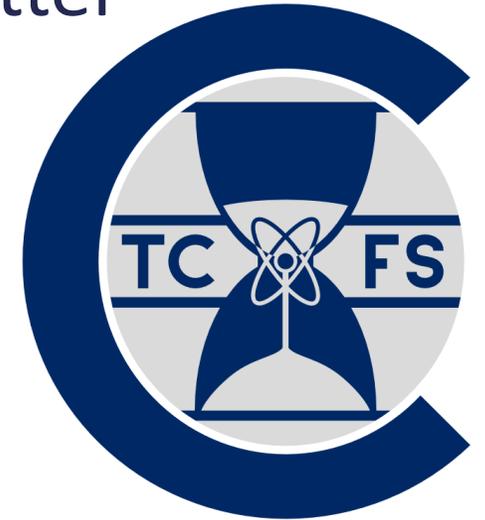


Constraining the coupling between axion-like dark matter and photons using an antiproton superconducting detection circuit in a cryogenic Penning trap

Jack Devlin (CERN, RIKEN)



東京大学
THE UNIVERSITY OF TOKYO

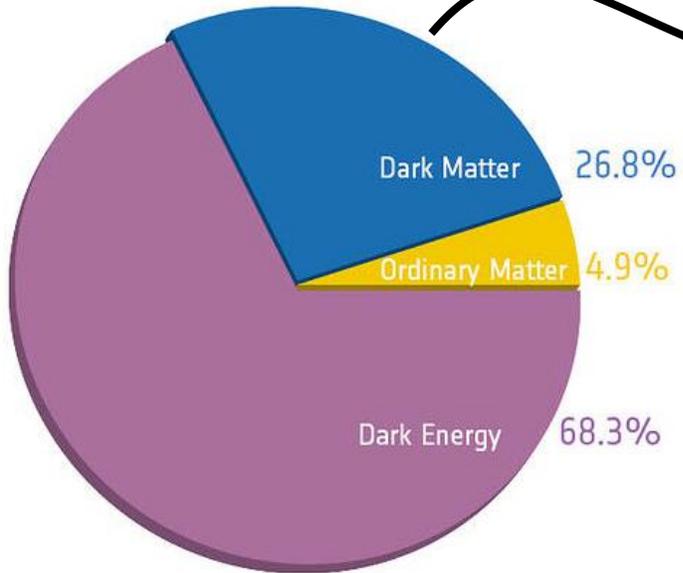


JOHANNES GUTENBERG
UNIVERSITÄT MAINZ

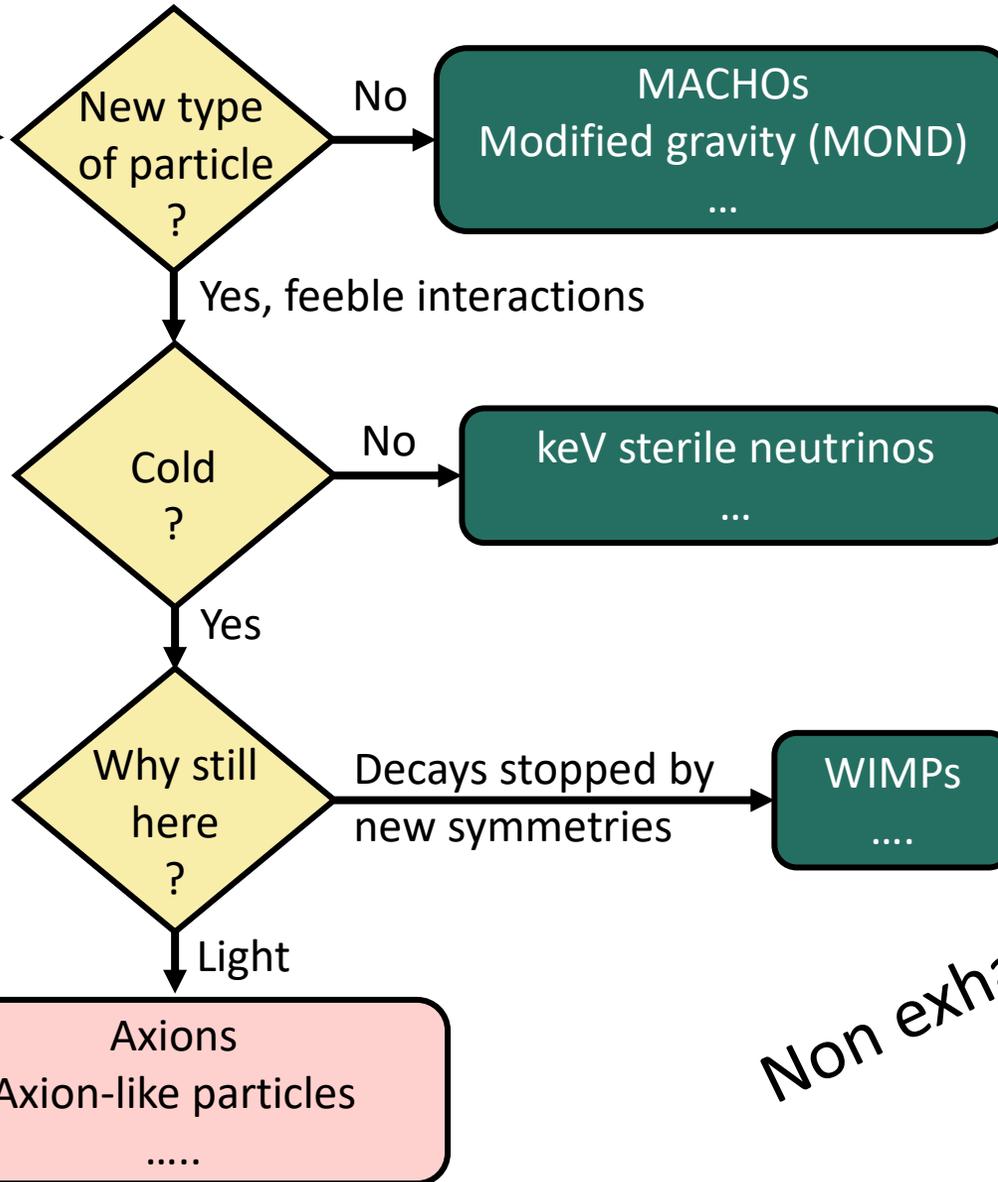


1. Axion-like dark matter
2. The BASE experiment
3. Detecting ALPs at BASE
4. Future perspectives

Why axion-like dark matter?



https://www.nasa.gov/mission_pages/planck/multimedia/pia16878.html



Motivated by symmetry →

Non exhaustive !

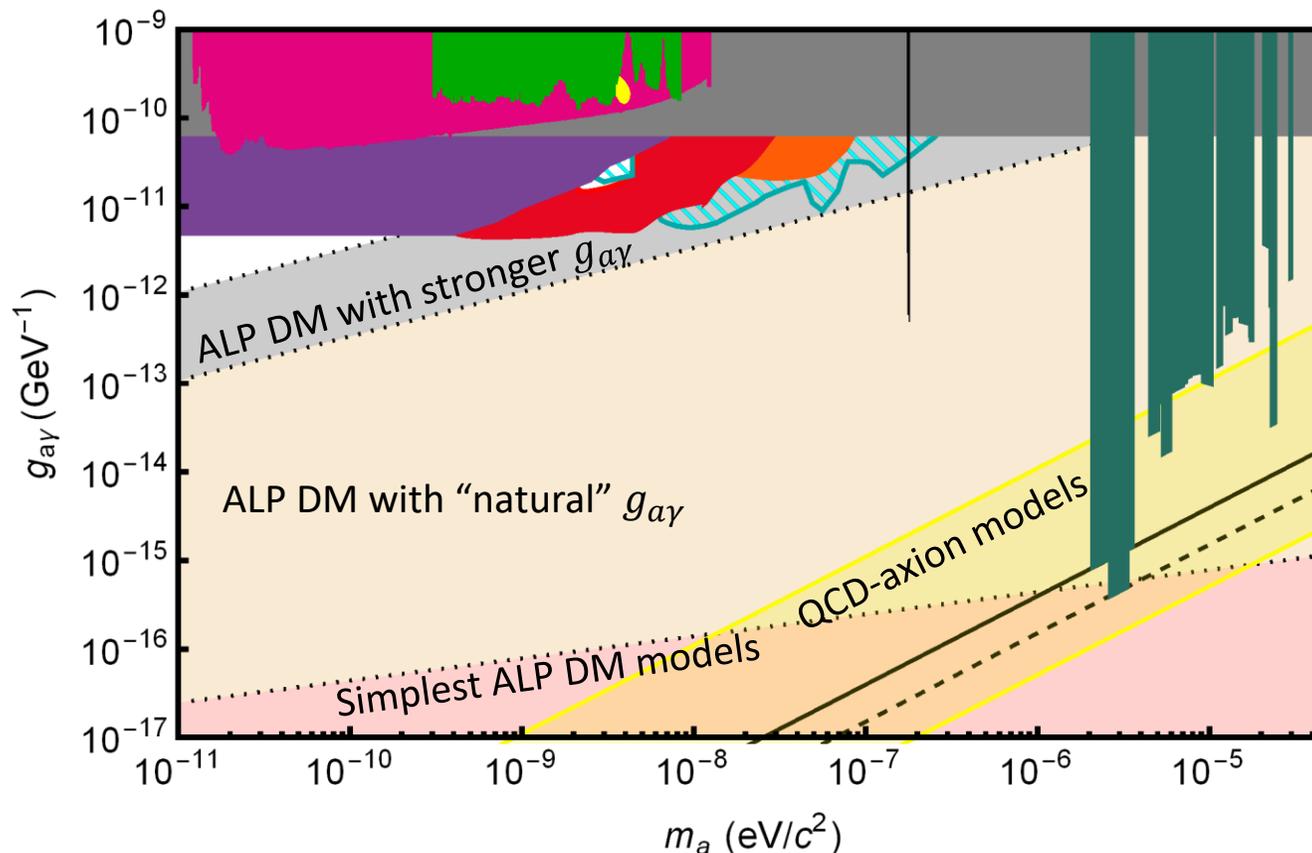
ALP-photon coupling

Axion-like particles are (light) pseudoscalar bosons which, if generated via the misalignment mechanism, form a BEC and behave like a classical scalar field

ALP-photon coupling

$$\mathcal{L}_{\text{int}} = -\frac{g_{a\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

“Easy” to make an ALP candidate consistent with observed dark matter density and early universe evolution



Limits					Hints	
SN-1987A	Cavities	CAST	ADMX-SLIC	FERMI-LAT	Excess γ -rays	Pulsars
H.E.S.S.	SHAFT	BASE	ABRACADABRA			

How to detect very low mass ALPs

Axions can couple to photons via the interaction term $\mathcal{L}_{\text{int}} = -\frac{g_{a\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$

This modifies Maxwell's equations

$$\nabla \cdot \vec{E} = \rho - g_{a\gamma} \vec{B} \cdot \nabla a$$

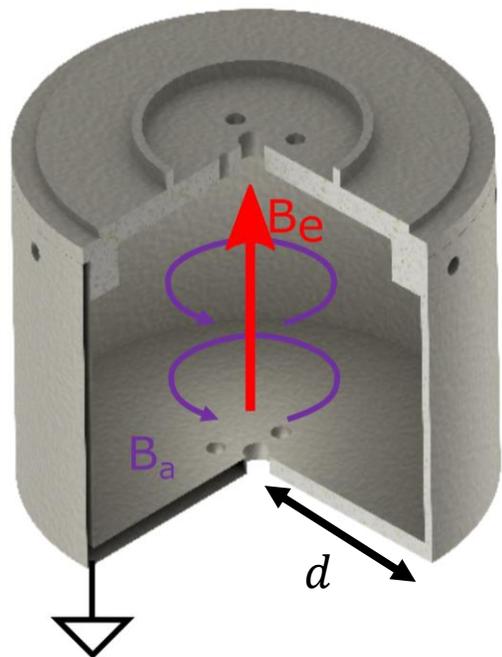
$$\nabla \times \vec{B} - \partial_t \vec{E} = \vec{J} + g_{a\gamma} (\vec{B} \partial_t a - \vec{E} \times \nabla a)$$

$$\nabla \cdot \vec{B} = 0$$

$$\nabla \times \vec{E} + \partial_t \vec{B} = 0$$

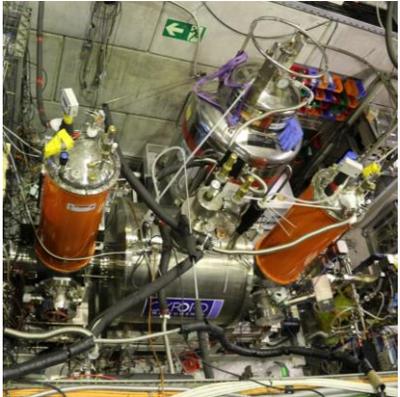
Inside a conducting housing with $d \ll \lambda_a$, and where there is a strong field B_e , the axions source a magnetic field

$$|\vec{B}_a| = \frac{1}{2} r g_{a\gamma} |\vec{B}_0| \sqrt{\rho_a \hbar c}$$

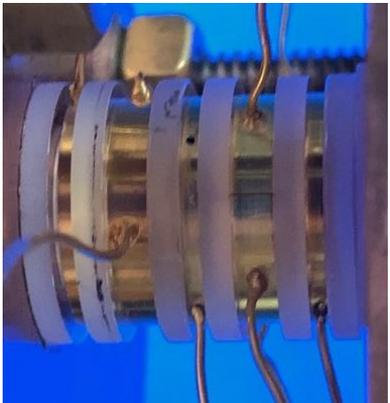


The BASE experiment

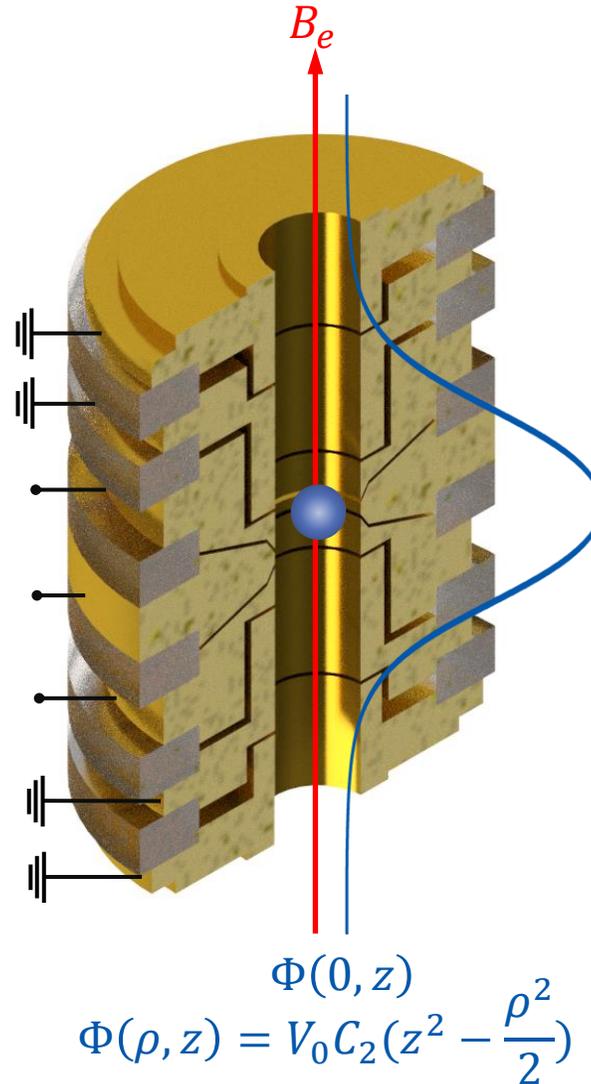
1.95 T B field from solenoid



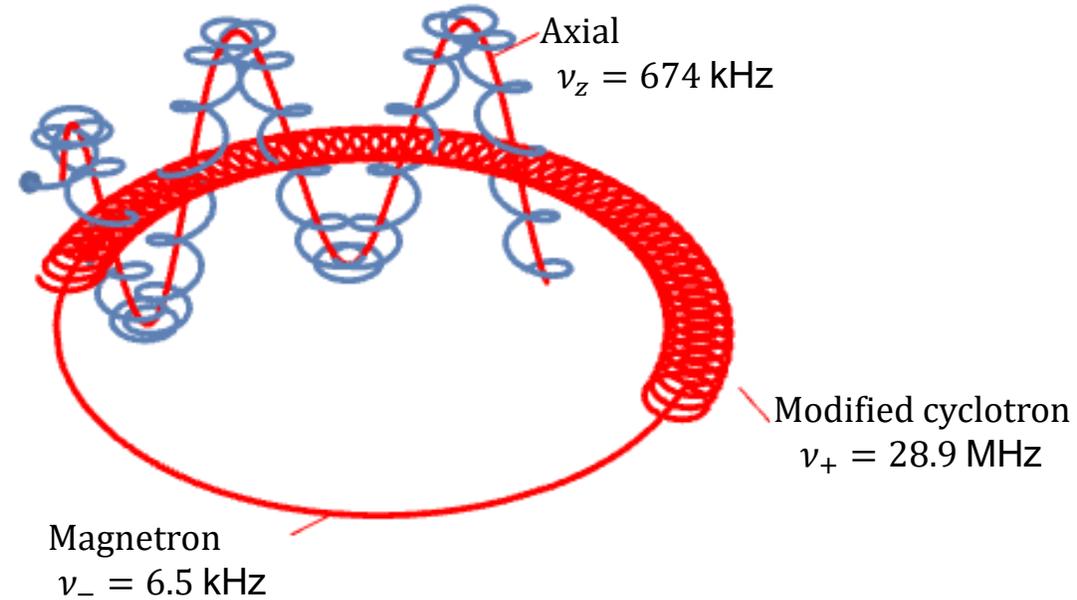
Voltages applied to ring-shaped electrodes



A Penning trap



Orbit is sum of three normal modes



$$\sqrt{\nu_z^2 + \nu_+^2 + \nu_-^2} = \nu_c = \frac{q}{2\pi m} B_e$$

Measure frequencies and get access to charge-to-mass ratio and magnetic field

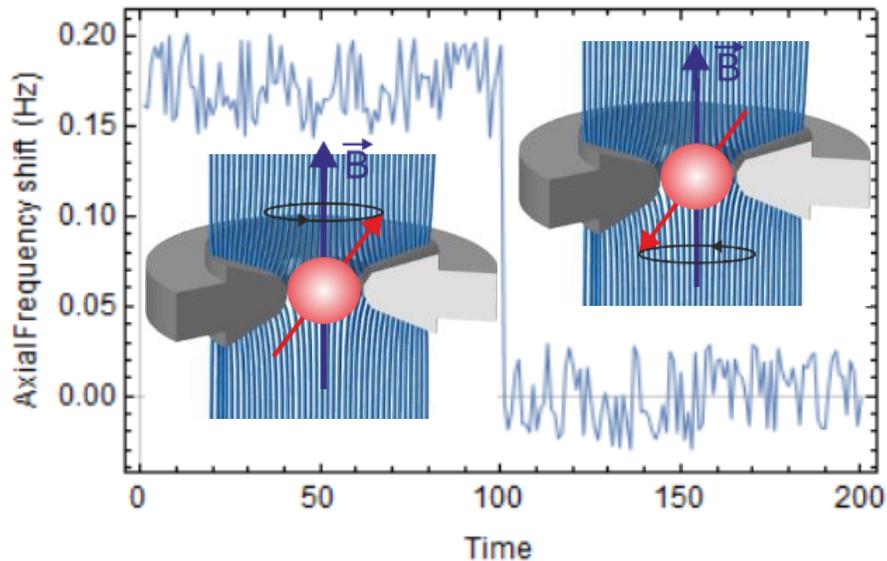
Measuring magnetic moments

$$\frac{g}{2} = \frac{2\pi \nu_L}{B_e \mu_N} = \frac{\nu_L}{\nu_c}$$

ν_L is the Larmor frequency, the frequency at which rf can most easily flip the spin

To measure ν_L , just need a way to identify the antiproton spin state

Detecting changes in the direction or size of μ



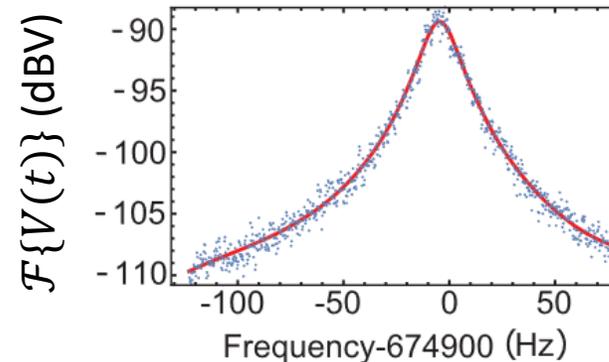
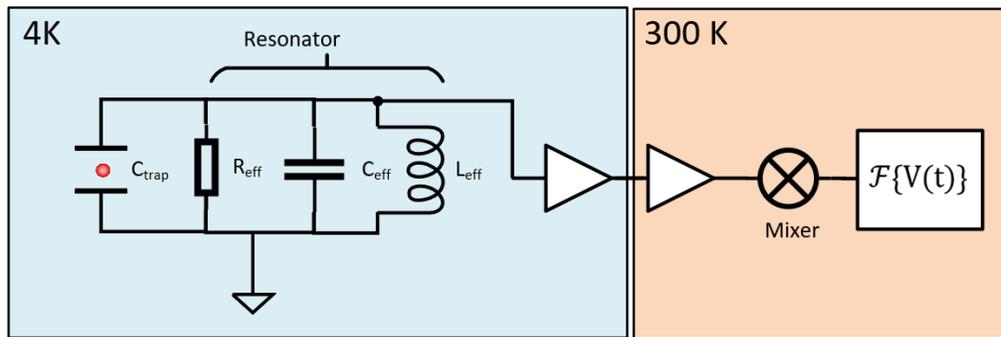
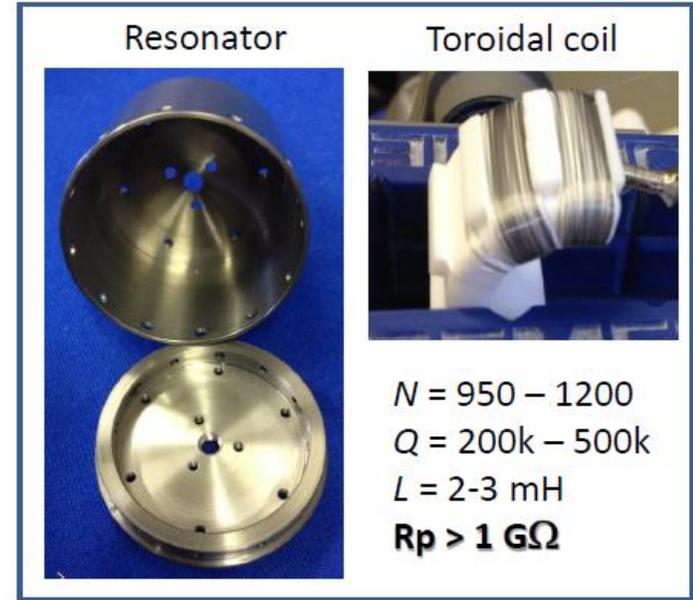
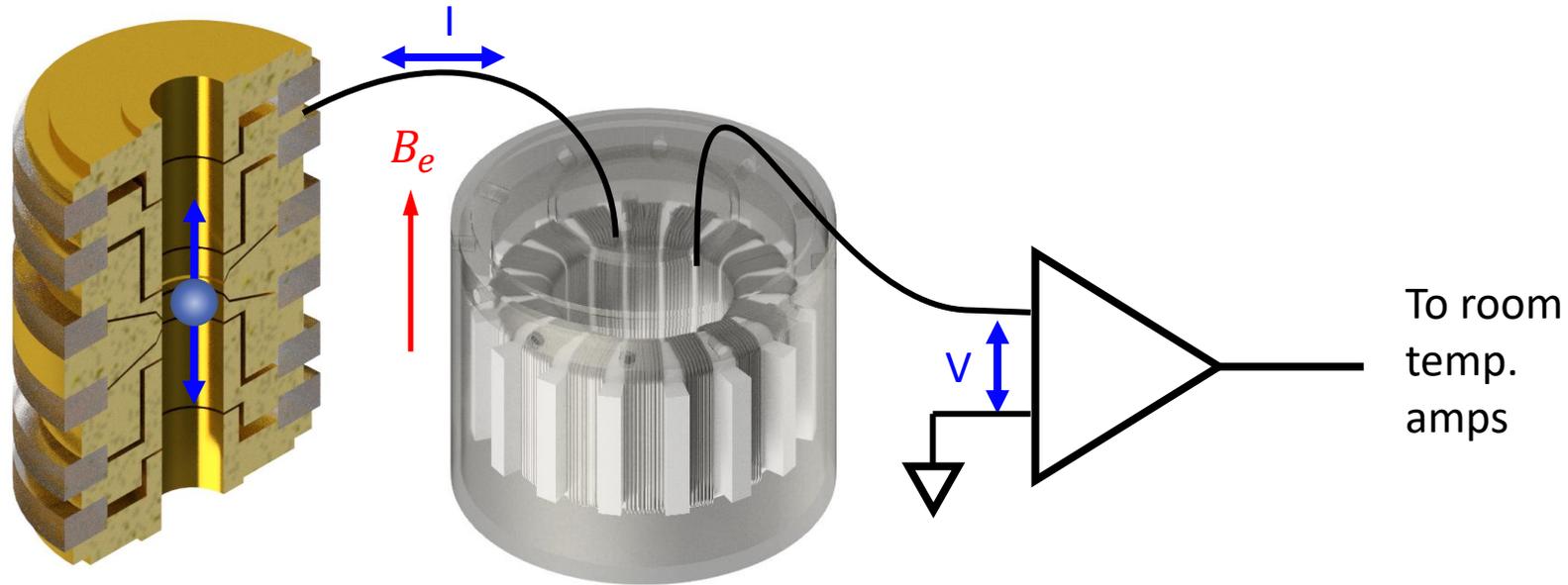
The particle has a total magnetic moment μ (spin + orbital)

If we add an inhomogeneous field $B = B_0 + B_2(z^2 - \rho^2/2)$ then there is an additional force $\nabla(\mu \cdot B)$ which contributes to the axial trapping force

This shifts the axial frequency depending on μ , so the spin state can be identified

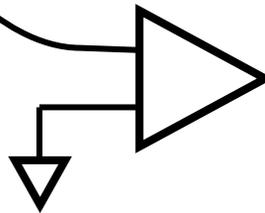
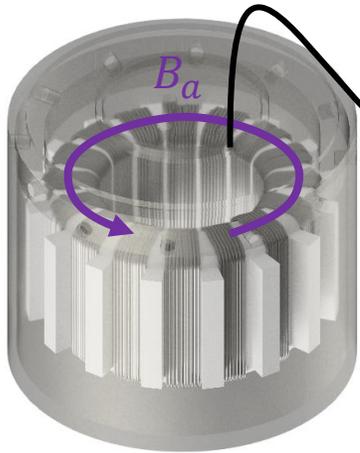
The BASE frequency detection system

Measurement of ~ 1 fA image currents induced in trap electrodes

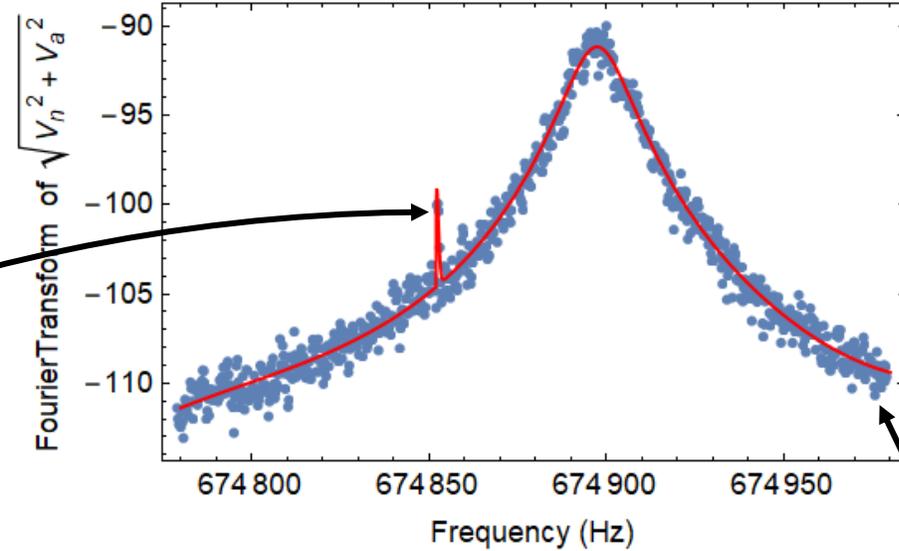


Resolving single antiproton spin flips requires the highest Q and lowest temperature LC resonant detectors ever built- **BASE-CERN is the state of the art**

Expected ALP signal



$$V_{\text{rms}} = \sqrt{V_n^2 + V_a^2}$$



The axion signal

$$V_a = \frac{\pi}{2} Q \sqrt{f(\nu, Q, \mathbf{q})} \kappa \nu_a l N_T (r_2^2 - r_1^2) g_{a\gamma} \| \mathbf{B}_e \| \sqrt{\rho_a \hbar c}.$$

The resonator background

$$V_n = \sqrt{e_n^2 \Delta\nu + \kappa^2 4k_B T_z \Delta\nu R_p f(\nu, Q, \mathbf{q})}$$

$f(\nu, Q, \mathbf{q})$ is a lorentzian line-shape function proportional to $\text{Re}\{Z\}$

e_n is the equivalent input noise of the amplifier

κ is the coupling constant

Q is the resonator Q-factor

N_T is the number of turns

l is the length of the toroid along the magnet B field

r_1 is the inner radius of the toroid

r_2 is the outer radius

$g_{a\gamma}$ is the coupling constant

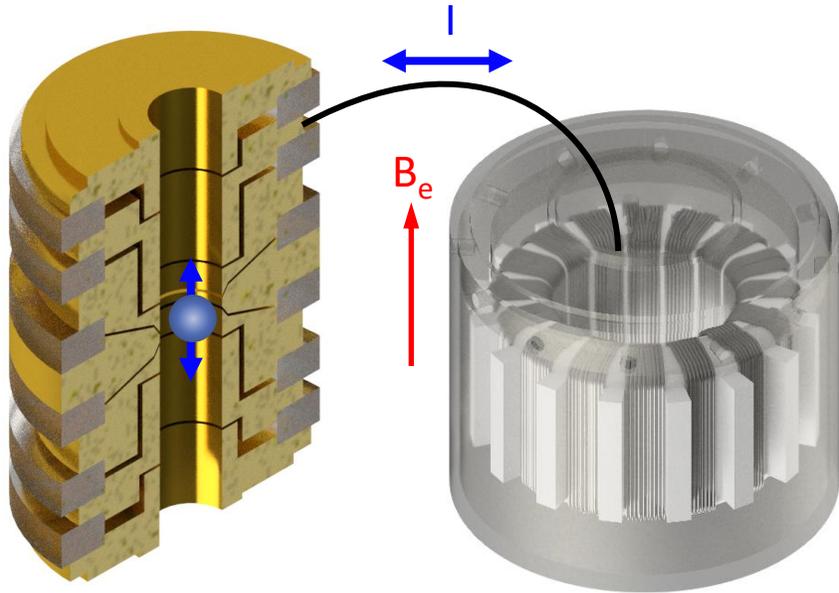
B is the static magnetic field

ρ_a is the dark matter density

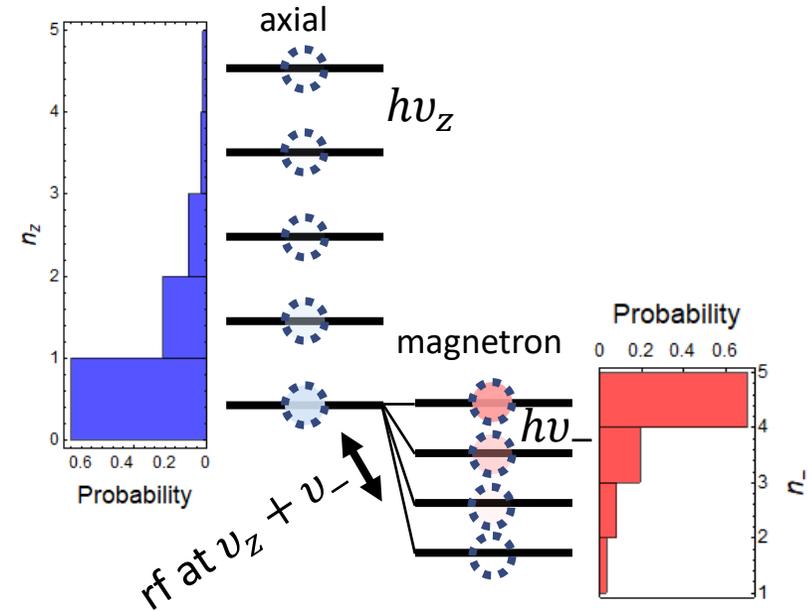
How to measure T_z ?

A "Boltzmann" thermometer

1.



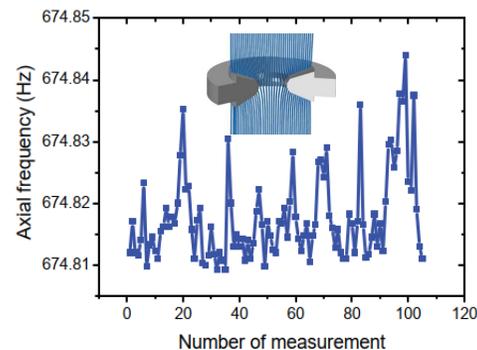
2.



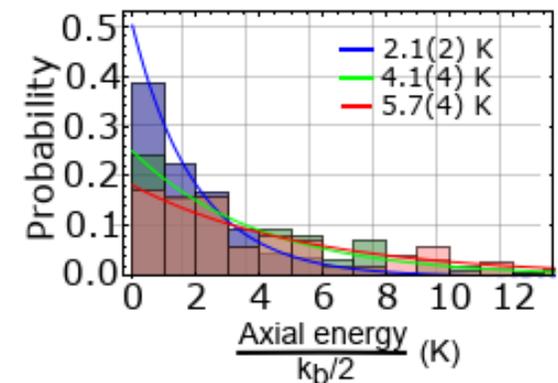
Trapped antiproton's axial motion reaches thermal equilibrium with the detector- can use it as a "quantum" sensor

3.

Magnetic inhomogeneity gives axial frequency shift proportional to $\mu \propto n_-$

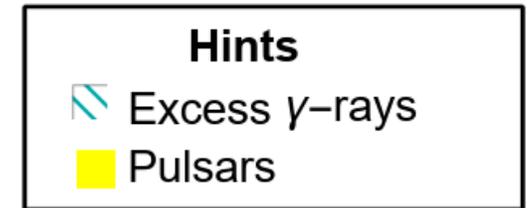
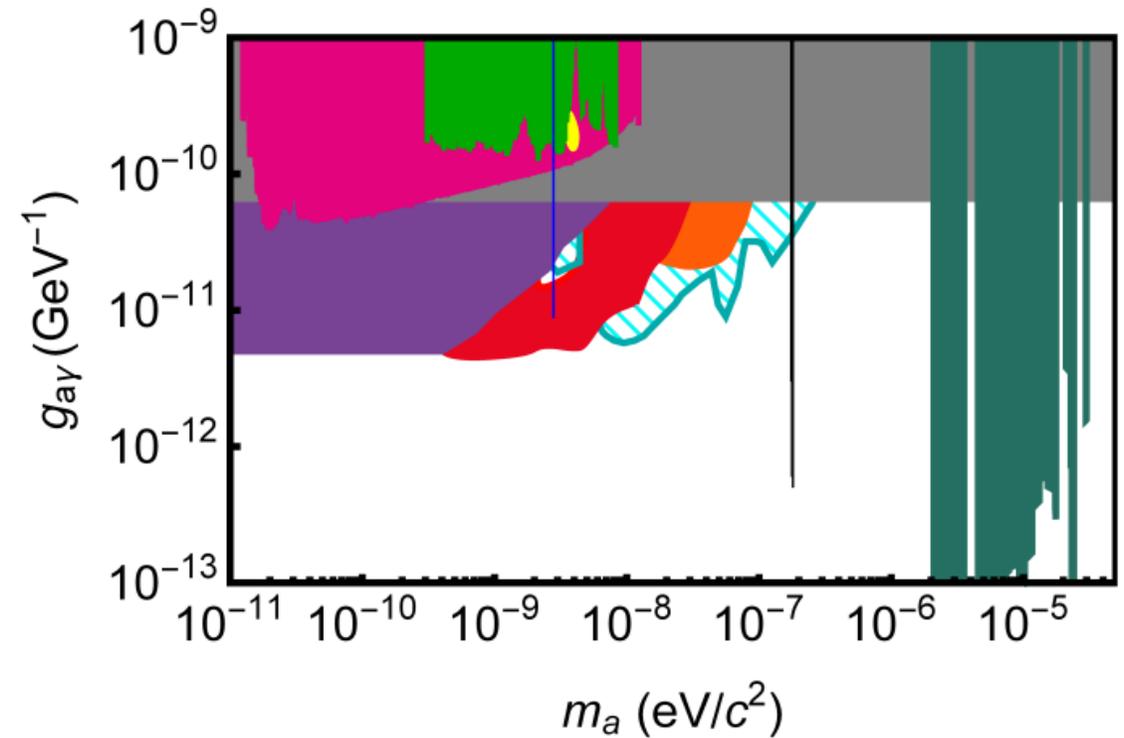
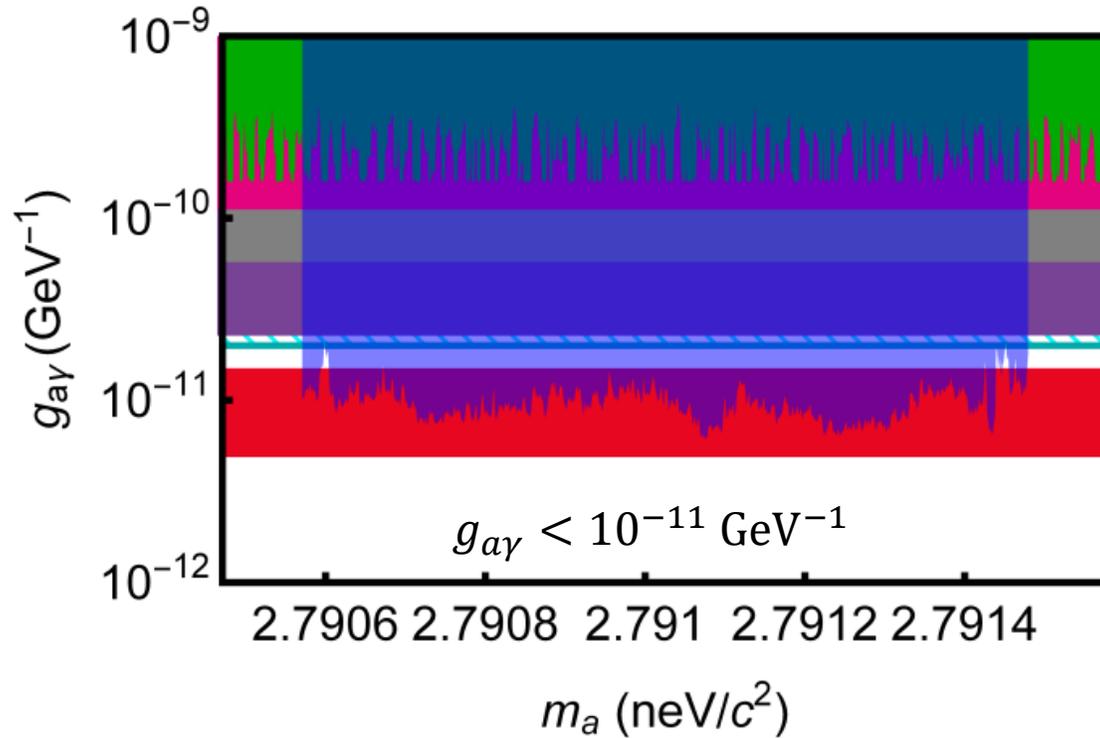


Knowing strength of inhomogeneity, can determine T_z



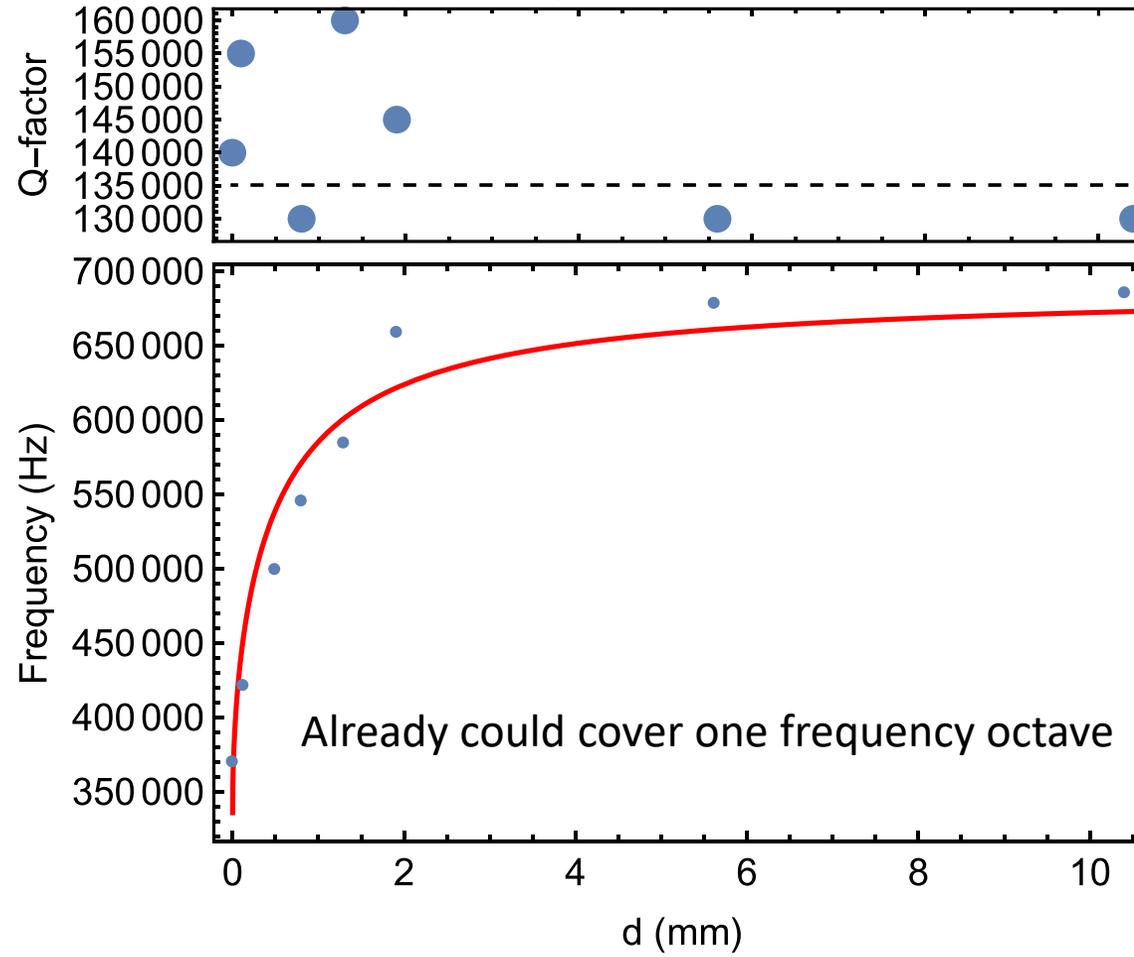
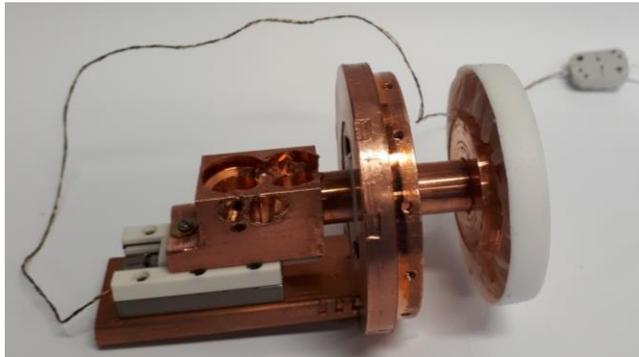
Limits

We use maximum likelihood estimation to estimate $g_{a\gamma}$, and then decide if we have a discovery, or set limits



Next step: frequency tuning

Cryogenic adjustable capacitance with no loss of Q already developed



← No capacitor

Already could cover one frequency octave

Frederik Volksen



Small detector(s): 5 cm long, 5 cm diameter

 **Immediately realizable with BASE technology today**, 6-9 months assembly time
 $(\frac{e_n}{\kappa} = 2.4 \text{ nV}/\sqrt{\text{Hz}}, 7 \text{ T}, 1 \text{ year acquisition})$

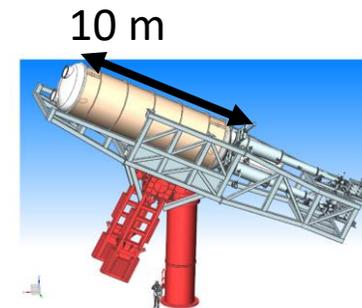
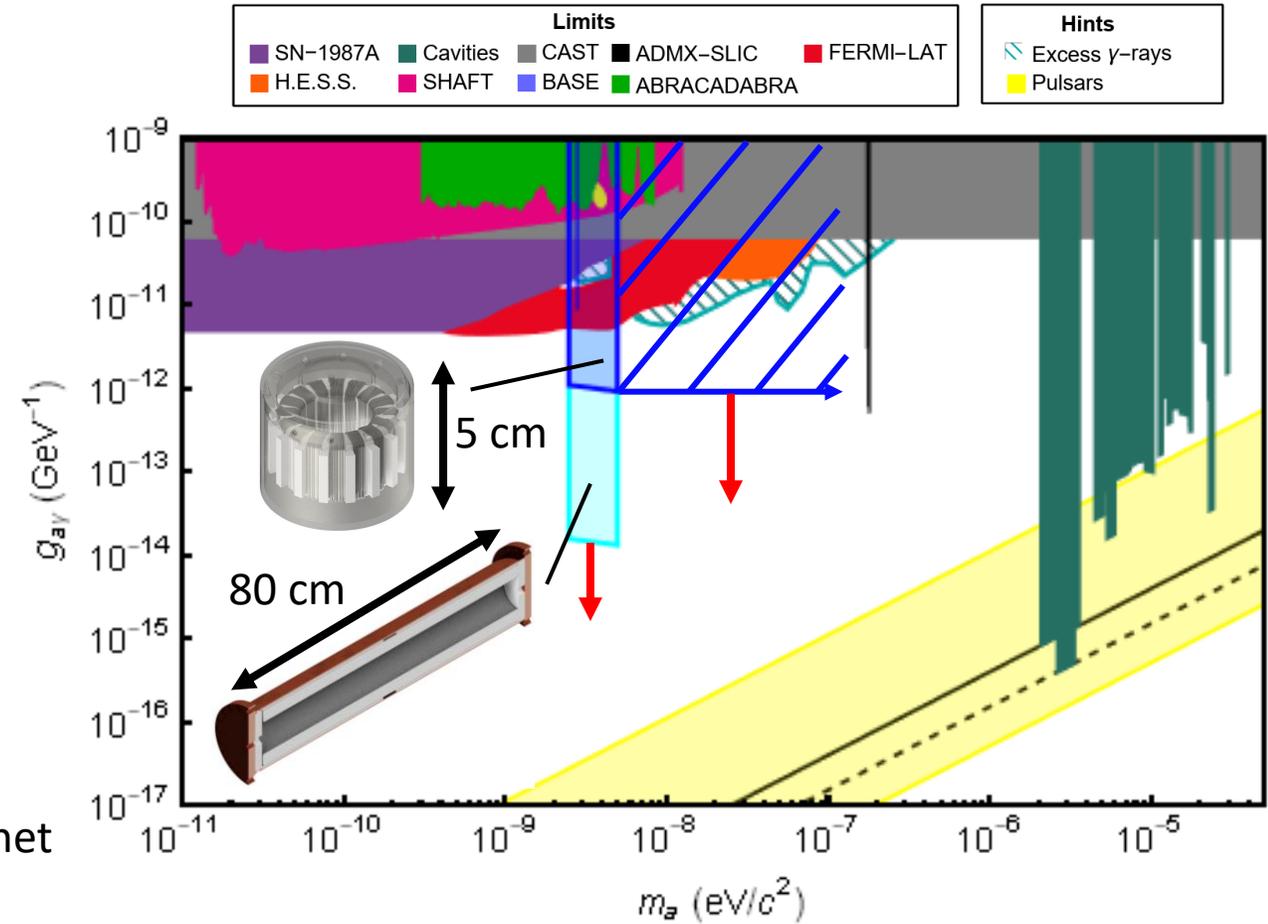
 Possible in the short term with detector RnD work using BASE know-how
 $(\frac{e_n}{\kappa} = 2.4 \text{ nV}/\sqrt{\text{Hz}}, 7 \text{ T}, 1 \text{ year acquisition, use multiple higher frequency coils without excessive Q-loss})$

Large detector: 80 cm long, 14 cm diameter

 Optimistic projection, “normal” Penning trap magnet
 $(\frac{e_n}{\kappa} = 2.4 \text{ nV}/\sqrt{\text{Hz}}, 7 \text{ T}, 1 \text{ year acquisition, } 10 \text{ mK}, Q=200,000)$

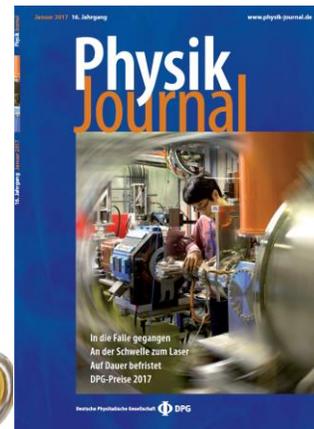
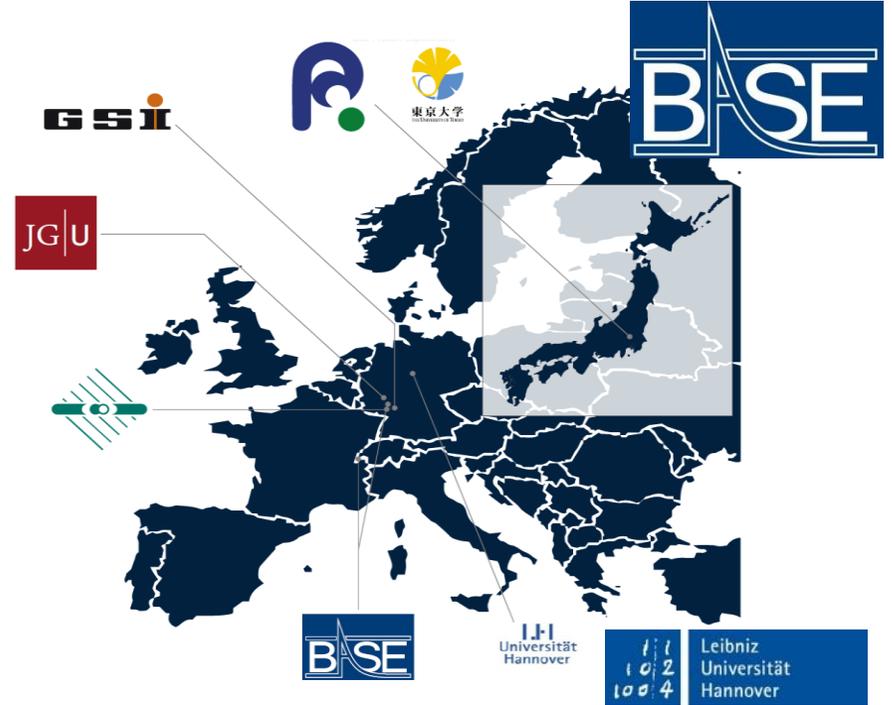
Pushing the sensitivity further

-  Much large detector volumes- **in discussion with RADES/babyIAXO**
- Colder detectors- **laser cooled resonators?**
- Lower noise amplifiers – particle assisted readout?



Thank you!

- **Mainz:** Measurement of the magnetic moment of the proton, implementation of new technologies
- **CERN-AD:** Measurement of the magnetic moment of the antiproton and proton/antiproton q/m ratio
- **Hannover/PTB:** QLEDS-laser cooling project, new technologies



Institutes: RIKEN, MPIK, CERN, University of Mainz, Tokyo University, GSI Darmstadt, University of Hannover, PTB Braunschweig



Team at CERN