

An ultraviolet completion for the Scotogenic model

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Work in collaboration with
Pablo Escribano
[\[arXiv:2107.10265\]](https://arxiv.org/abs/2107.10265)

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VNIVERSITAT
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CSIC
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There are **MANY** Majorana neutrino mass models...

Tree-level

Radiative: 1-loop, 2-loop, 3-loop, ...

High scale

Low scale

Dimension-5: Weinberg operator

Higher dimensions: dim-7, dim-9, ...

There are **MANY** Majorana neutrino mass models...

Review: [Cai, Herrero-García, Schmidt,
AV, Volkas, 2017]

See talk by
Antonio
Cárcamo

Tree-level

Radiative: 1-loop, 2-loop, 3-loop, ...

High scale

Loop suppression

Dark matter candidate

Low scale

Dimension-5: Weinberg operator

Higher dimensions: dim-7, dim-9, ...

Outline

Introduction

Finished already!



The Scotogenic model

A quick review of the well-known Scotogenic model

An UV completion for the Scotogenic model

An UV extension that addresses some of its drawbacks and leads to new pheno

The Scotogenic model

Also known as...

The inert doublet model

The radiative seesaw

Ma's model

The Scotogenic model

[Ma, 2006]

σκότος
skotos = darkness



	gen	$SU(2)_L$	$U(1)_Y$	\mathbb{Z}_2
η	1	2	1/2	—
N	3	1	0	—



Inert (or dark) doublet

**Dark
Matter!**

$$\mathcal{L}_N = \overline{N}_i \not{\partial} N_i - \frac{M_{R_i}}{2} \overline{N}_i^c N_i + y_{i\alpha} \eta \overline{N}_i \ell_\alpha + \text{h.c.}$$

$$\begin{aligned} \mathcal{V} = & m_H^2 H^\dagger H + m_\eta^2 \eta^\dagger \eta + \frac{\lambda_1}{2} (H^\dagger H)^2 + \frac{\lambda_2}{2} (\eta^\dagger \eta)^2 + \lambda_3 (H^\dagger H) (\eta^\dagger \eta) \\ & + \lambda_4 (H^\dagger \eta) (\eta^\dagger H) + \frac{\lambda_5}{2} \left[(H^\dagger \eta)^2 + (\eta^\dagger H)^2 \right] \end{aligned}$$

The Scotogenic model

[Ma, 2006]

$$\mathcal{V} = m_H^2 H^\dagger H + m_\eta^2 \eta^\dagger \eta + \frac{\lambda_1}{2} (H^\dagger H)^2 + \frac{\lambda_2}{2} (\eta^\dagger \eta)^2 + \lambda_3 (H^\dagger H) (\eta^\dagger \eta) \\ + \lambda_4 (H^\dagger \eta) (\eta^\dagger H) + \frac{\lambda_5}{2} \left[(H^\dagger \eta)^2 + (\eta^\dagger H)^2 \right]$$

Inert scalar sector: η^\pm $\eta^0 = (\eta_R + i\eta_I)/\sqrt{2}$

$$\begin{aligned} m_{\eta^+}^2 &= m_\eta^2 + \lambda_3 \langle H^0 \rangle^2 \\ m_R^2 &= m_\eta^2 + (\lambda_3 + \lambda_4 + \lambda_5) \langle H^0 \rangle^2 \\ m_I^2 &= m_\eta^2 + (\lambda_3 + \lambda_4 - \lambda_5) \langle H^0 \rangle^2 \end{aligned} \quad \Rightarrow \quad m_R^2 - m_I^2 = 2 \lambda_5 \langle H^0 \rangle^2$$

Radiative neutrino masses

[Ma, 2006]

Tree-level:

Forbidden by the \mathbb{Z}_2 symmetry

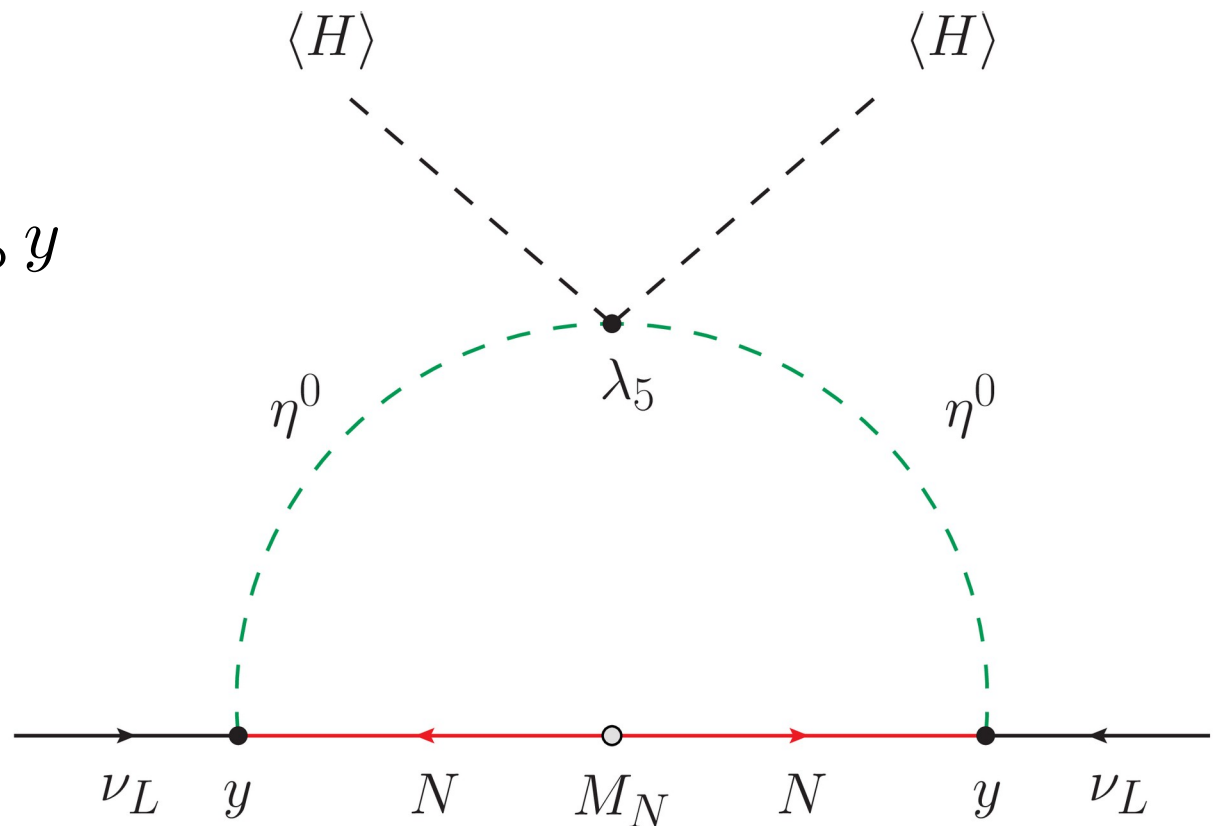
Radiative generation of neutrino masses

$$m_\nu = \frac{\lambda_5 v^2}{32\pi^2} y^T M_R^{-1} f_{\text{loop}} y$$

Dark particles in the loop

[Other variations in Restrepo, Zapata, Yaguna, 2013]

1-loop neutrino masses



Radiative neutrino masses

[Ma, 2006]

Tree-level:

Forbidden by the \mathbb{Z}_2 symmetry

Radiative generation of neutrino masses

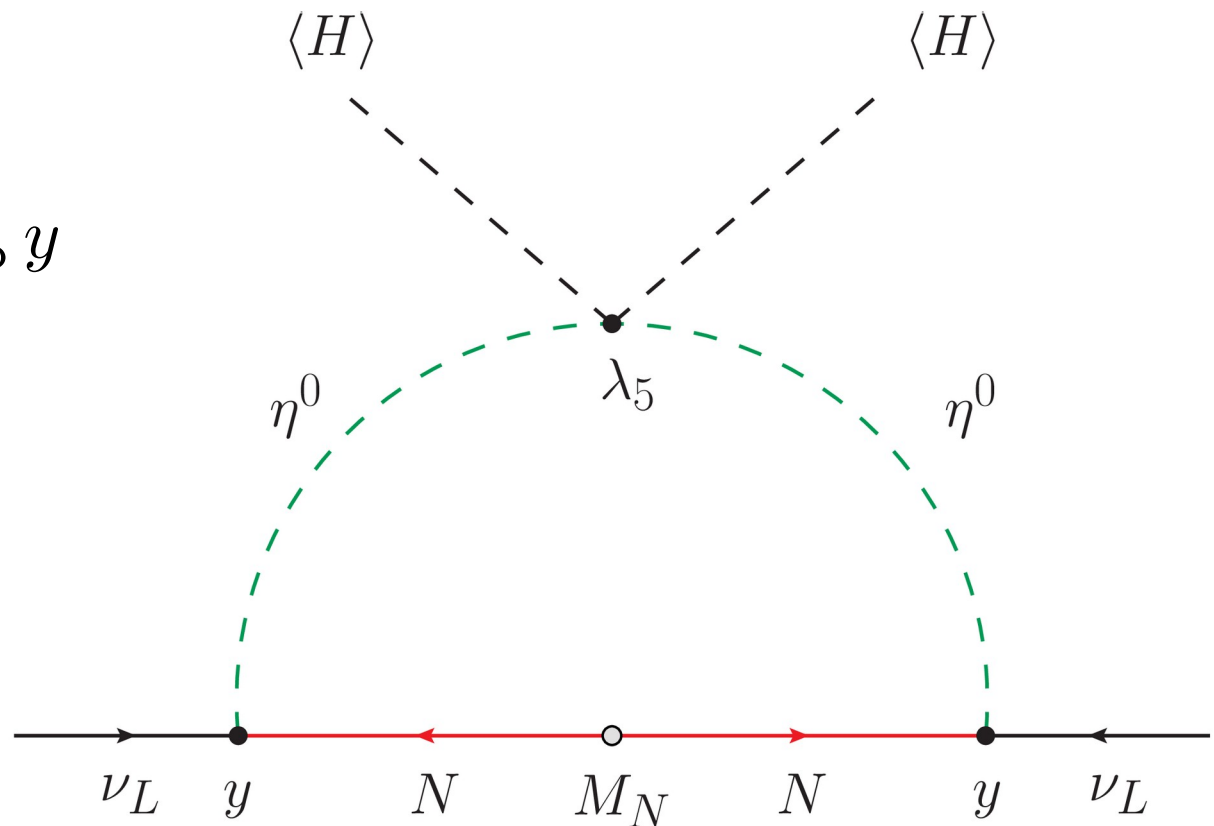
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Dark particles in the loop

[Other variations in Restrepo, Zapata, Yaguna, 2013]



1-loop neutrino masses



Dark matter

The lightest particle charged under \mathbb{Z}_2 is stable: dark matter candidate

Fermion Dark Matter: N_1

- It can only be produced via **Yukawa** interactions
- Potential problems with lepton flavor violation: is it compatible with the current bounds?

Scalar Dark Matter: the lightest neutral η scalar, η_R or η_I

- It also has **gauge** interactions
- Not correlated to lepton flavor violation

Drawbacks



There is no explanation for the smallness of the λ_5 parameter

Although it's natural ['t Hooft, 1980]



The \mathbb{Z}_2 Scotogenic parity is an *ad-hoc* symmetry

These issues can be simultaneously addressed in **ultraviolet completions** of the Scotogenic model

Chuck Norris fact of the day

*Chuck Norris can divide by
zero.*



An UV completion for the Scotogenic model

Work in collaboration with **Pablo Escribano**

[\[arXiv:2107.10265\]](https://arxiv.org/abs/2107.10265)

Ultraviolet theory

NEW! →

NEW! →

Field	Generations	SU(3) _c	SU(2) _L	U(1) _Y	U(1) _L
ℓ_L	3	1	2	-1/2	1
e_R	3	1	1	-1	1
N	3	1	1	0	$\frac{1}{2}$
H	1	1	2	1/2	0
η	1	1	2	1/2	$-\frac{1}{2}$
Δ	1	1	3	1	-1
S	1	1	1	0	1

$$\Delta = \begin{pmatrix} \Delta^+/\sqrt{2} & \Delta^{++} \\ \Delta^0 & -\Delta^+/\sqrt{2} \end{pmatrix}$$

Scalar triplet

S

Scalar singlet

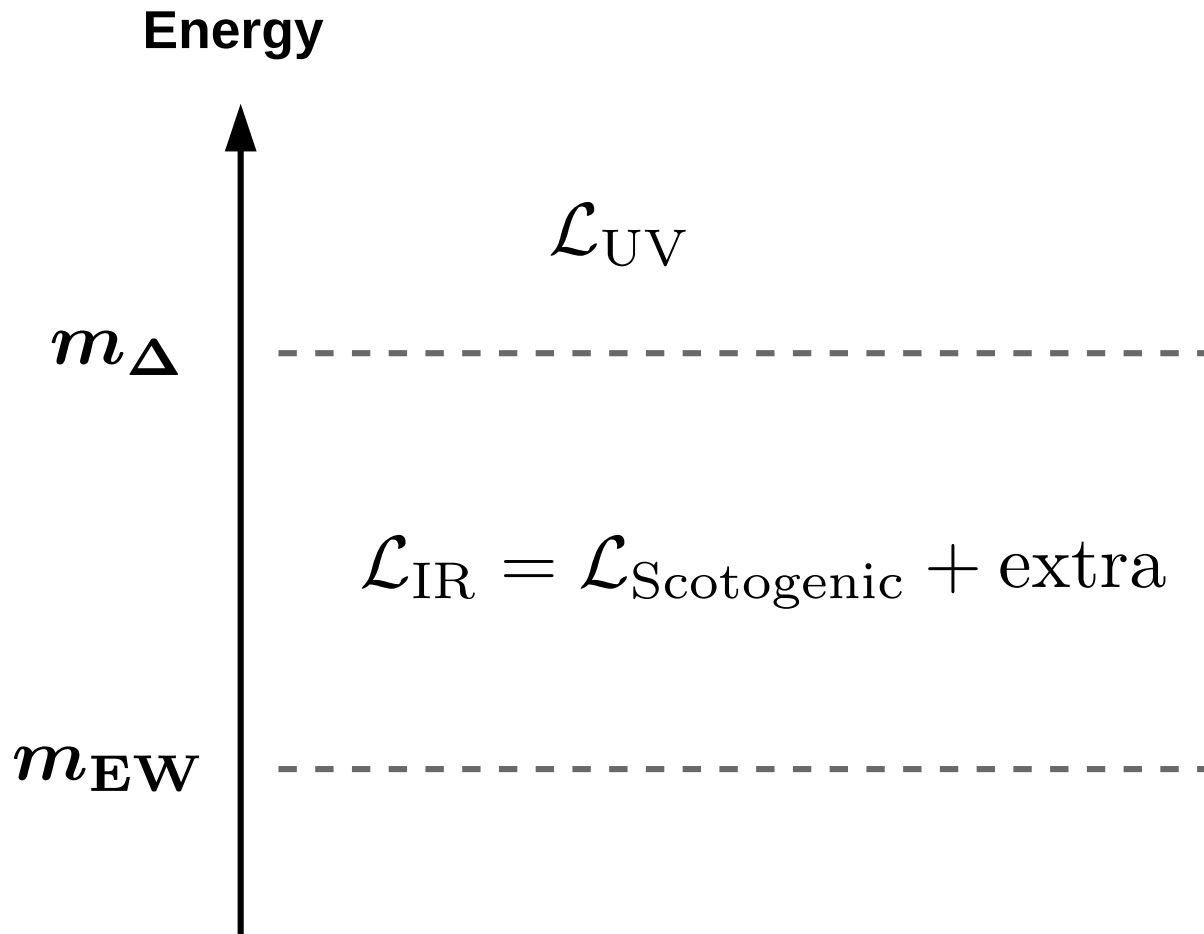
Ultraviolet theory

$$\mathcal{L}_{\text{UV}} = \mathcal{L}_{\text{SM}} + \textcolor{red}{y} \overline{N} \eta i \sigma_2 \ell_L + \textcolor{red}{\kappa} S^* \overline{N^c} N + \text{h.c.} - \mathcal{V}_{\text{UV}}$$

$$\begin{aligned} \mathcal{V}_{\text{UV}} = & m_H^2 H^\dagger H + m_S^2 S^* S + m_\eta^2 \eta^\dagger \eta + m_\Delta^2 (\Delta^\dagger \Delta) \\ & + \frac{1}{2} \lambda_1 (H^\dagger H)^2 + \frac{1}{2} \lambda_S (S^* S)^2 + \frac{1}{2} \lambda_2 (\eta^\dagger \eta)^2 \\ & + \frac{1}{2} \lambda_{\Delta 1} (\Delta^\dagger \Delta)^2 + \frac{1}{2} \lambda_{\Delta 2} (\Delta^\dagger \Delta)^2 + \lambda_3^S (H^\dagger H) (S^* S) + \lambda_3 (H^\dagger H) (\eta^\dagger \eta) \\ & + \lambda_3^\Delta (H^\dagger H) (\Delta^\dagger \Delta) + \lambda_3^{\eta S} (\eta^\dagger \eta) (S^* S) + \lambda_3^{\eta \Delta} (\eta^\dagger \eta) (\Delta^\dagger \Delta) \\ & + \lambda_3^{S \Delta} (S^* S) (\Delta^\dagger \Delta) + \lambda_4 (H^\dagger \eta) (\eta^\dagger H) + \lambda_4^\Delta (H^\dagger \Delta^\dagger \Delta H) \\ & + \lambda_4^{\eta \Delta} (\eta^\dagger \Delta^\dagger \Delta \eta) + [\textcolor{red}{\lambda}_{HS\Delta} S (H^\dagger \Delta i \sigma_2 H^*) + \textcolor{red}{\mu} (\eta^\dagger \Delta i \sigma_2 \eta^*) + \text{h.c.}] . \end{aligned}$$

Other terms are forbidden by **lepton number**

Low-energy theory



Assumption

$$m_\Delta \gg \text{other mass scales}$$

Strategy

Integrate out Δ and keep contributions up to **dimension 6**

$$\mathcal{O}\left(\frac{1}{m_\Delta^2}\right)$$

Low-energy theory

$$\mathcal{L}_{\text{IR}} = \mathcal{L}_{\text{SM}} + \textcolor{red}{y} \bar{N} \eta i \sigma_2 \ell_L + \textcolor{red}{\kappa} S^* \bar{N}^c N + \text{h.c.} - \mathcal{V}_{\text{IR}}$$

$$\begin{aligned} \mathcal{V}_{\text{IR}} = & m_H^2 H^\dagger H + m_S^2 S^* S + m_\eta^2 \eta^\dagger \eta + (H^\dagger H)^2 \left[\frac{\lambda_1}{2} - \frac{|\lambda_{HS\Delta}|^2}{m_\Delta^2} (S^* S) \right] + \frac{\lambda_S}{2} (S^* S)^2 \\ & + (\eta^\dagger \eta)^2 \left(\frac{\lambda_2}{2} - \frac{|\mu|^2}{m_\Delta^2} \right) + \lambda_3^S (H^\dagger H) (S^* S) + \lambda_3 (H^\dagger H) (\eta^\dagger \eta) + \lambda_3^{\eta S} (\eta^\dagger \eta) (S^* S) \\ & + \lambda_4 (H^\dagger \eta) (\eta^\dagger H) - \left[\frac{\lambda_{HS\Delta} \mu^*}{m_\Delta^2} S (H^\dagger \eta)^2 + \right] + \mathcal{O} \left(\frac{1}{m_\Delta^4} \right) \end{aligned}$$

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$$\langle H^0 \rangle = \frac{v_H}{\sqrt{2}} \quad \langle S \rangle = \frac{v_S}{\sqrt{2}}$$

$$\text{U}(1)_L \xrightarrow{v_S} \mathbb{Z}_2$$

Symmetry breaking

[Ma, 2015]

[Centelles Chuliá et al, 2019]


Low-energy theory

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$$\frac{\lambda_5}{2} \equiv - \frac{\lambda_{HS\Delta} \mu^* v_S}{\sqrt{2} m_\Delta^2} \ll 1$$

Automatically small
 λ_5 coupling

Low-energy theory

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Automatically small

λ_5 coupling



\mathbb{Z}_2 -even scalars

$$H^0 = \frac{1}{\sqrt{2}} (v_H + \phi + i A) \qquad S = \frac{1}{\sqrt{2}} (v_S + \rho + i J)$$

\mathbb{Z}_2 -even scalars

$$H^0 = \frac{1}{\sqrt{2}} (v_H + \phi + i A) \quad S = \frac{1}{\sqrt{2}} (v_S + \rho + i J)$$

CP-even scalars

$$\mathcal{M}_R^2 = \begin{pmatrix} v_H^2 \left(\lambda_1 - \frac{v_S^2 |\lambda_{HS\Delta}|^2}{m_\Delta^2} \right) & v_H v_S \left(\lambda_3^S - \frac{v_H^2 |\lambda_{HS\Delta}|^2}{m_\Delta^2} \right) \\ v_H v_S \left(\lambda_3^S - \frac{v_H^2 |\lambda_{HS\Delta}|^2}{m_\Delta^2} \right) & v_S^2 \lambda_S \end{pmatrix}$$

$$V_R^T \mathcal{M}_R^2 V_R = \text{diag} (m_h^2, m_\Phi^2)$$

$$\begin{array}{c} \uparrow \\ \approx 125 \text{ GeV} \end{array}$$

Higgs-like boson @ LHC

Mixing angle

$$\tan(2\alpha) \approx 2 \frac{\lambda_3^S}{\lambda_S} \frac{v_H}{v_S}$$

\mathbb{Z}_2 -even scalars

$$H^0 = \frac{1}{\sqrt{2}} (v_H + \phi + i A)$$

$$S = \frac{1}{\sqrt{2}} (v_S + \rho + i J)$$

CP-odd scalars

A : would-be Goldstone absorbed by the Z

J : massless Goldstone boson, **the majoron**

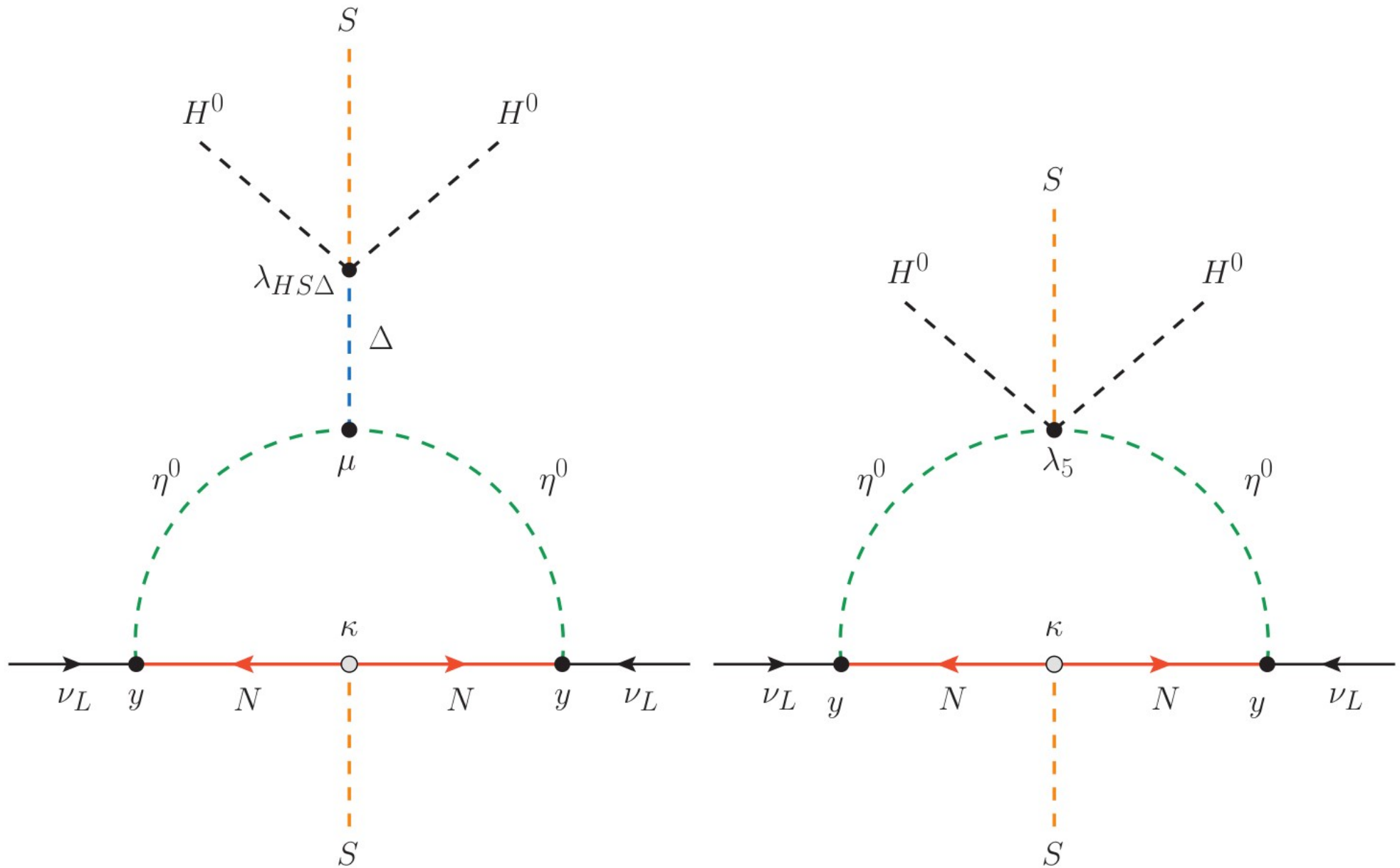


Singlet state

Must have **suppressed**
couplings to matter

The low-energy theory is the
Scotogenic model supplemented with
additional scalar fields

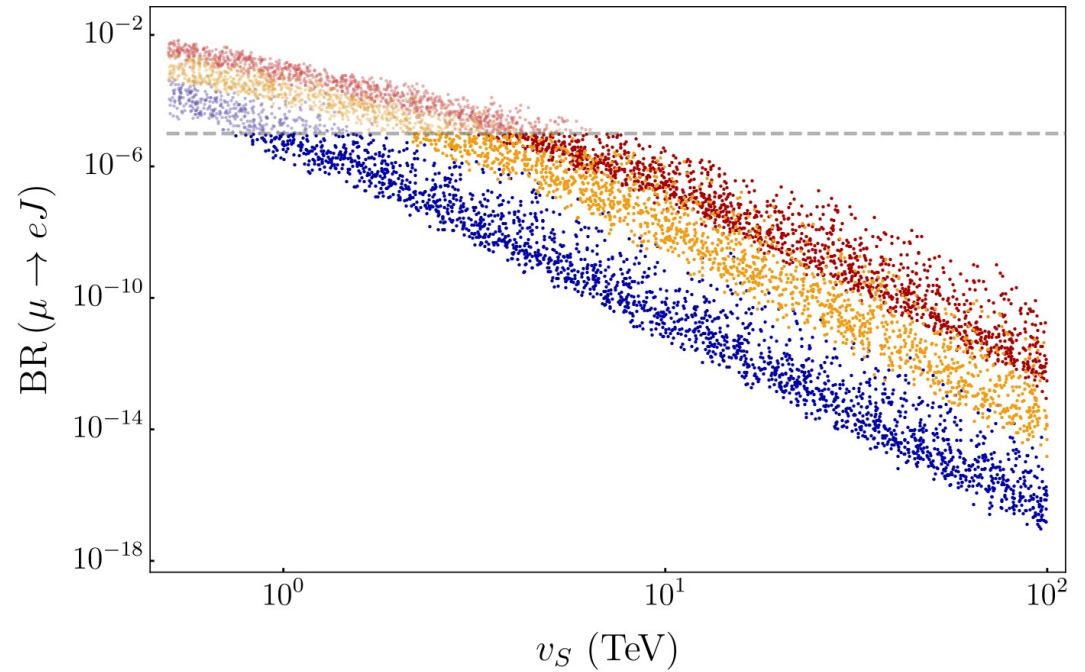
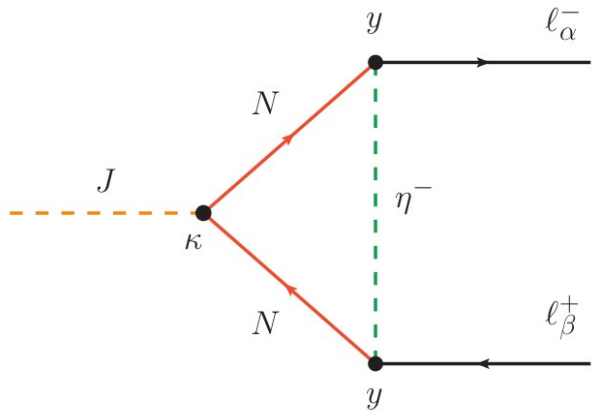
Neutrino masses



Phenomenology

Majoron couplings to charged leptons

1-loop coupling



$$\mathcal{L}_{J\ell\ell} = \frac{i J}{32 \pi^2 v_S} \bar{\ell} \left(M_{\ell} y^{\dagger} \Gamma y P_L - y^{\dagger} \Gamma y M_{\ell} P_R \right) \ell$$

$$\Gamma_{mn} = \frac{M_{N_n}^2}{\left(M_{N_n}^2 - m_{\eta^+}^2 \right)^3} \left(M_{N_n}^4 - m_{\eta^+}^4 + 2 M_{N_n}^2 m_{\eta^+}^2 \log \frac{m_{\eta^+}^2}{M_{N_n}^2} \right) \delta_{mn}$$

Sizable rates for low
lepton number breaking
scales

Phenomenology

Collider signatures

Invisible Higgs decay : $h \rightarrow J J$

$$g_{hJJ} = v_S \lambda_S \sin \alpha + \left(\lambda_3^S - \frac{v_H^2 |\lambda_{HS\Delta}|^2}{m_\Delta^2} \right) v_H \cos \alpha$$
$$\Rightarrow \lambda_3^S \lesssim 10^{-2}$$

Dark matter

New production mechanisms
(particularly relevant for fermion DM)

$$N_1 N_1 \longleftrightarrow \text{SM SM}$$

$$N_1 N_1 \longleftrightarrow J J$$

Mediated by the
new scalars

[Bonilla et al, 2020]

Final discussion

Final discussion

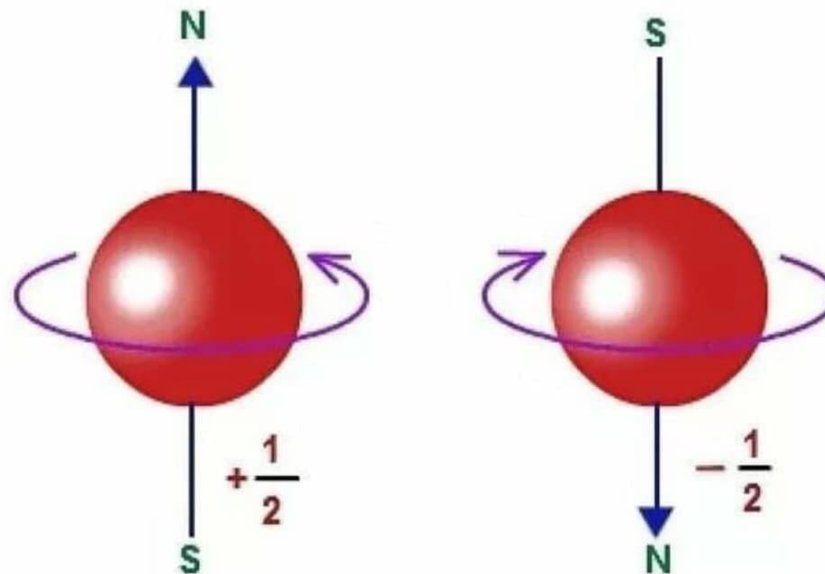
The **Scotogenic model** is a very economical scenario for **neutrino masses** that includes a **dark matter** candidate

An **ultraviolet completion** for the Scotogenic model:

- **Automatically small** λ_5 coupling due to large scale suppression
- \mathbb{Z}_2 parity from **spontaneous lepton number breaking**
- Additional degrees of freedom at low energies: massive scalar + massless **majoron**
- **Novel phenomenological predictions**: lepton flavor violation (with majorons), Higgs physics and dark matter

Thanks for your attention!

Electron spin explained: imagine a ball that's rotating, except it's not a ball and it's not rotating



Backup slides

A philosophical moment

Occam's razor:

The simplest explanation is the correct one

Occam's laser:

The most awesome explanation is the correct one

Occam's hammer:

My explanation is the correct one

All credit goes to
Alberto Aparici