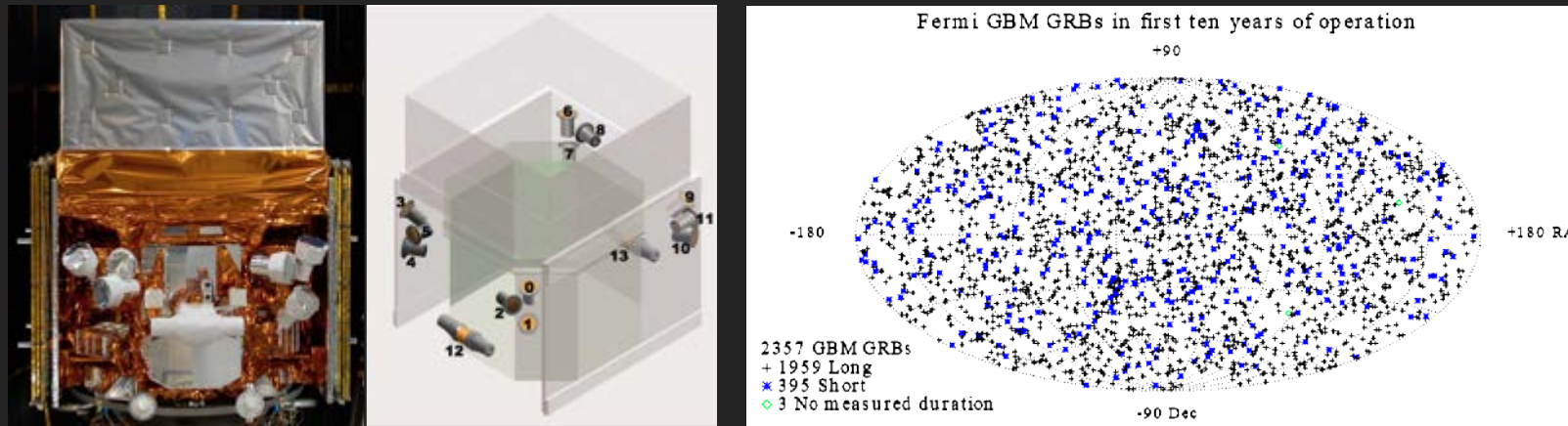


# A Decade of Gamma-ray Bursts observed by Fermi-GBM

4<sup>th</sup> GRB Catalog: von Kienlin, A., et al., 2020, ApJ, 893, 46

3<sup>rd</sup> GRB Spectral Catalog: Poolakkil, S., et al., 2021 ApJ accepted, arXiv:2103.13528v1



Andreas von Kienlin - on behalf of the Fermi-GBM team

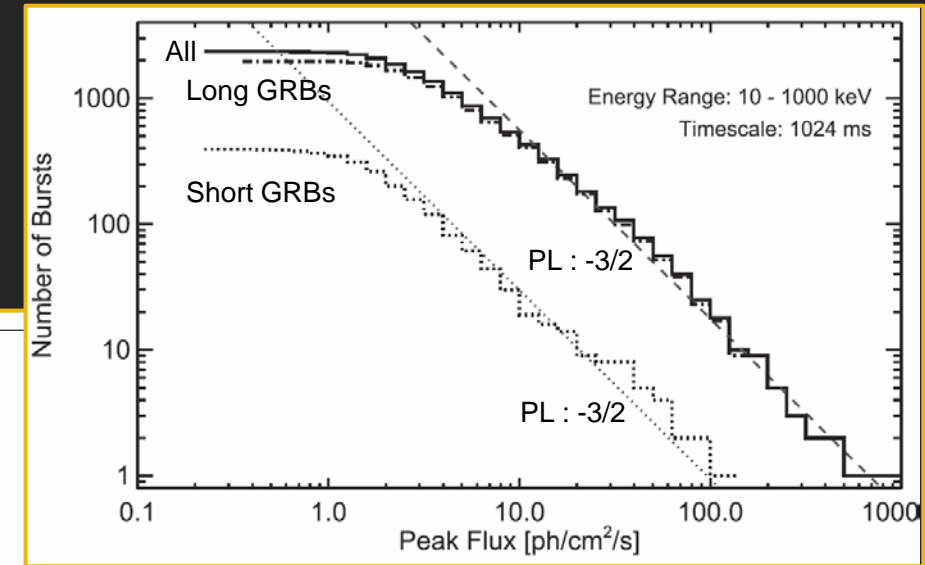
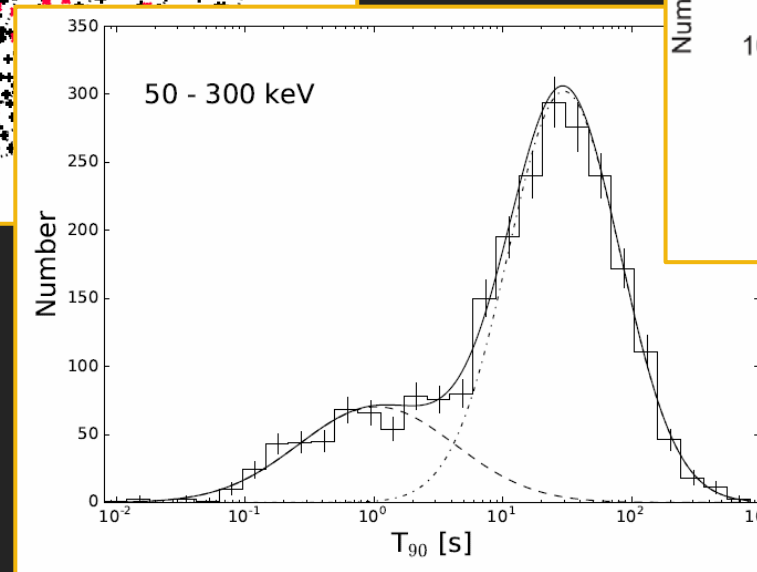
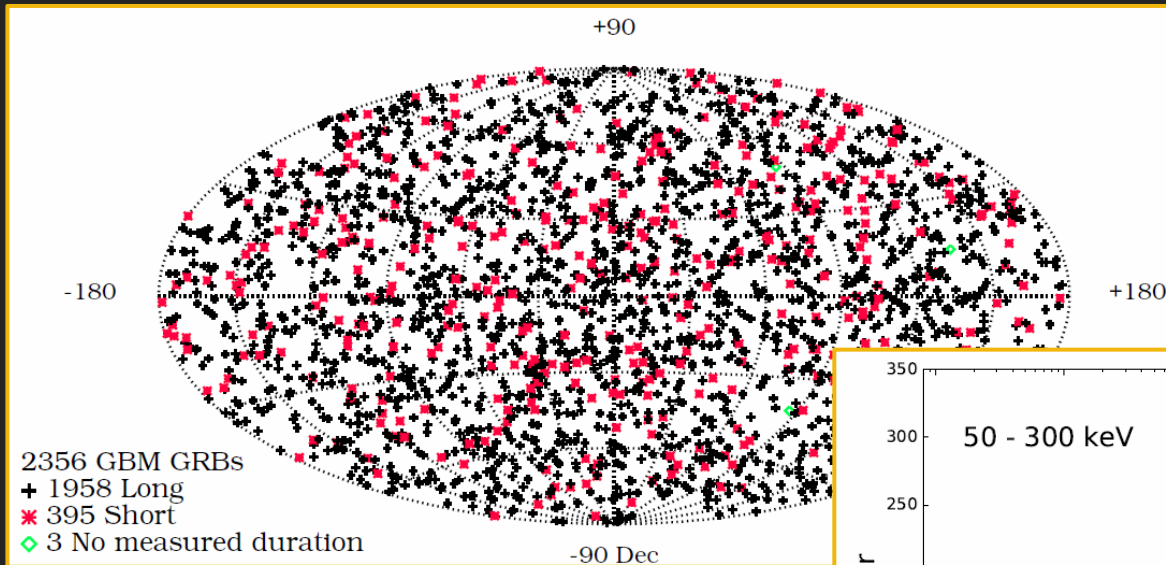
Max-Planck-Institut für extraterrestrische Physik (MPE), Garching



# 4<sup>th</sup> GBM GRB Catalog - 10 years of data

von Kienlin et al., 2020

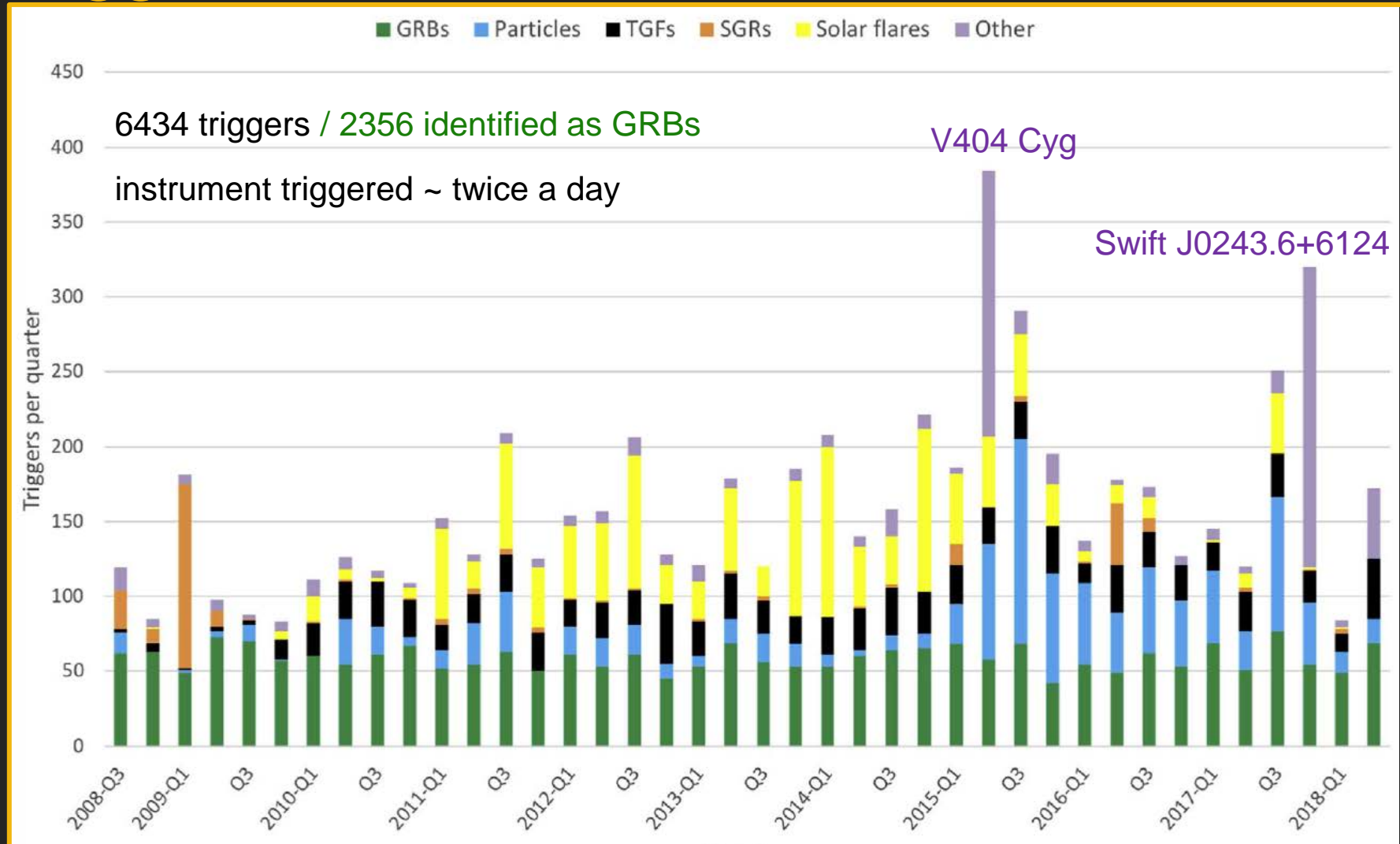
- 1<sup>st</sup>, 2<sup>nd</sup>, 3<sup>rd</sup> catalog: 2, 4, 6 years  $\Rightarrow$  Paciesas et al. 2012, von Kienlin et al. 2014, Bhat et al. 2016
- for each GRB: location, duration, peak flux & fluence (50 – 300 keV, 10 – 1000keV)  $\Rightarrow$  standard tables
- + information on triggering criteria, exceptional operational conditions, GCN products



2008 July 12 – 2018 July 11

complete results available  
@ HEASARC

# 10 Year Trigger Statistics



# GBM Triggers:

GRBs	SGRs	TGFs	SFs	Galactic	CPs	Other	Sum	ARRs	LAT GRBs
2360*	258	880	1176	407	1023	331	6434	220	186

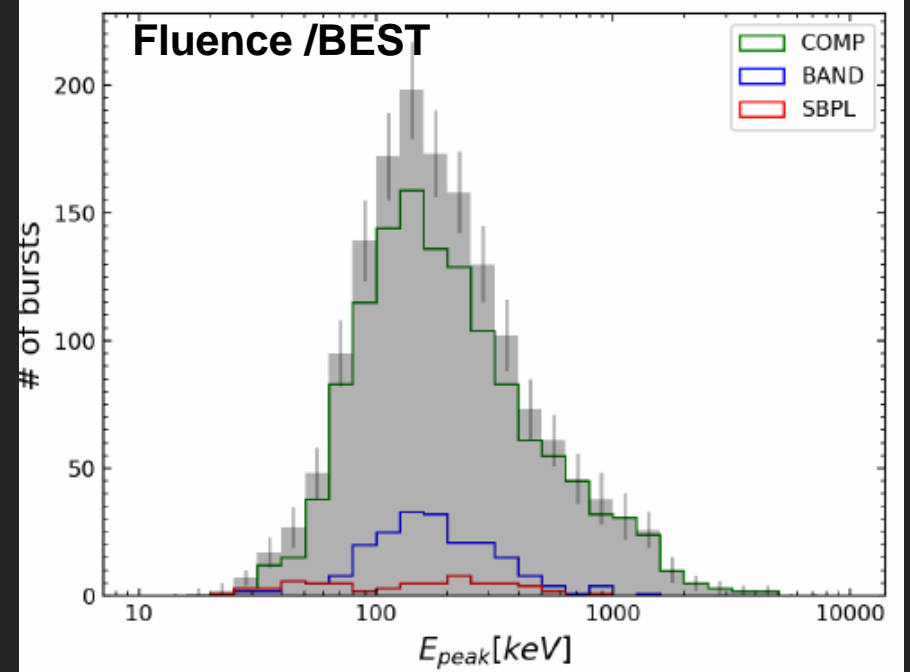
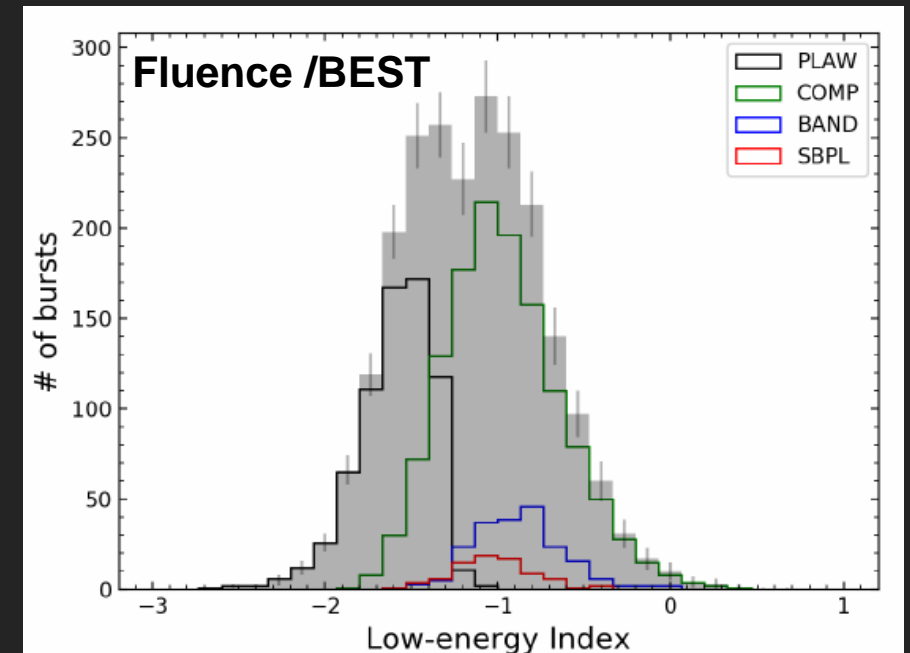
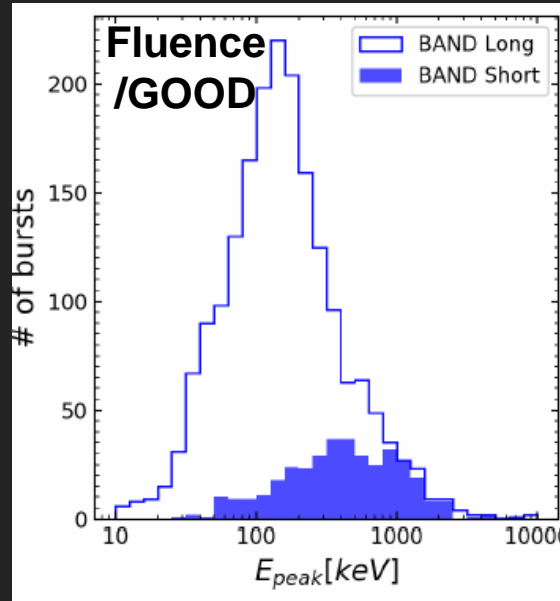
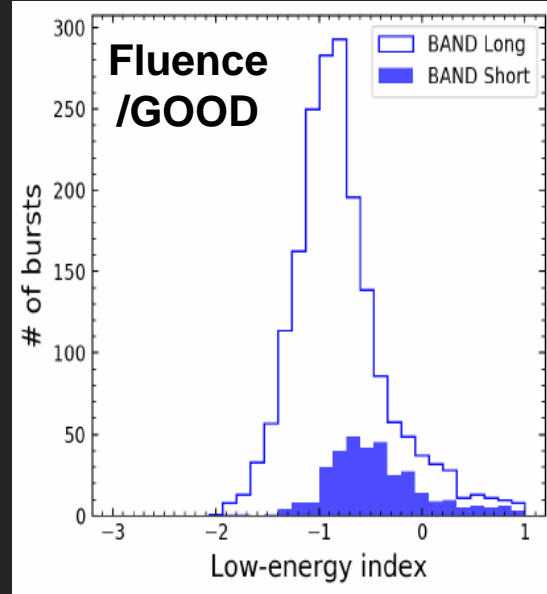
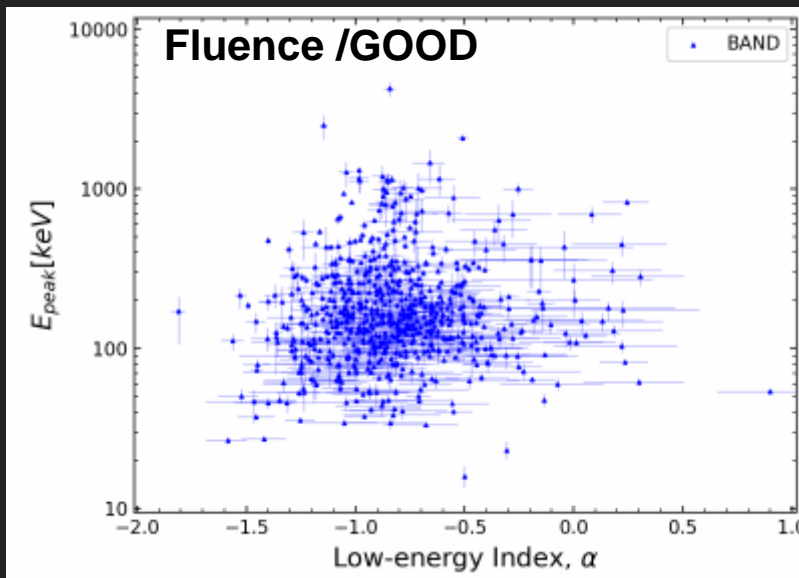
# GRB Spectral Catalog - 10 years of data

Poolakkil, S., et al., 2021

- ◆ Systematic spectral analysis of **2297 GRBs**, two types of spectra:
  - time-integrated spectral fits  $\Rightarrow$  'fluence' spectrum
  - spectral fits at the brightest time bin (1.024s / 64 ms)  $\Rightarrow$  'peak flux' spectrum
  - $\Rightarrow$  resulting in a compendium of over **18000 spectra!**
  - Update of 2-yr (Goldstein et al. 2012) and 4-yr (Gruber et al. 2014) catalogs (data at HEASARC)
- ◆ 4 different empirical spectral models
  - **PLAW** ( $A, \lambda$ ), **COMP** ( $A, \alpha, E_{\text{peak}}$ ), **BAND** ( $A, \alpha, \beta, E_{\text{peak}}$ ), **SBPL** ( $A, \lambda_1, \lambda_2, E_{\text{break}}, \Delta$ )
  - **New**: two-sided uncertainties!
  - Fit ratings:
    - ▶ **GOOD**: parameter error of all model parameters are within certain limits
    - ▶ **BEST**: best representation model, based upon goodness of fit criteria

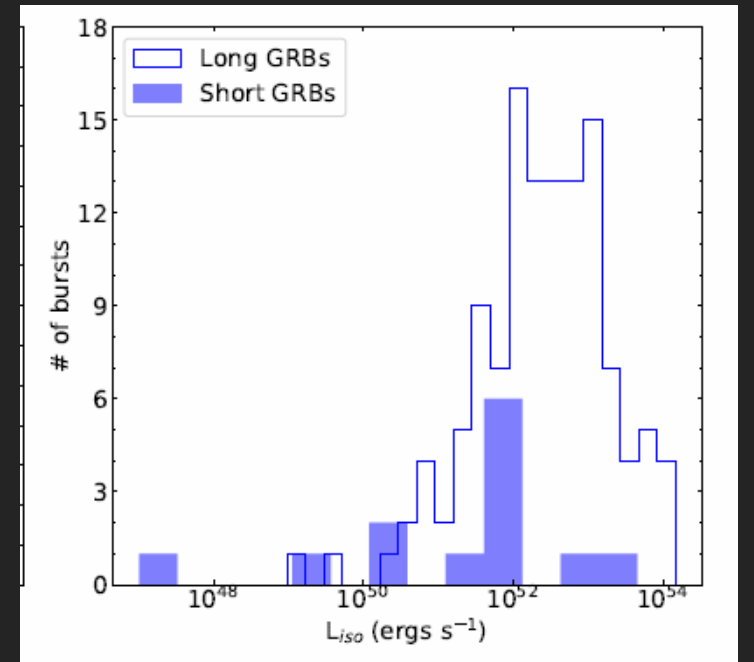
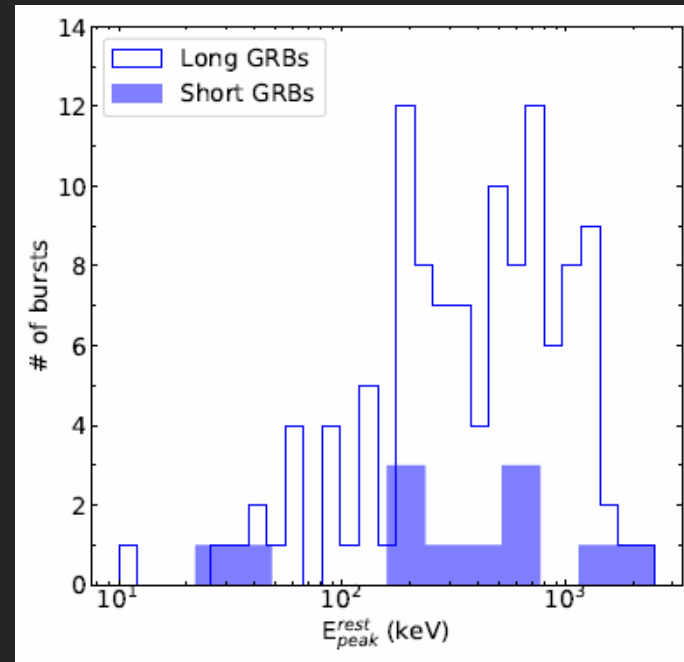
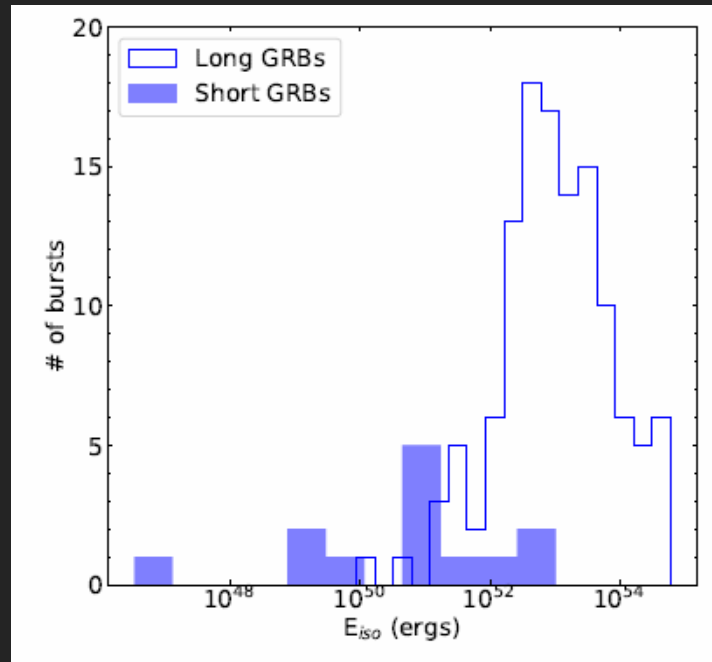
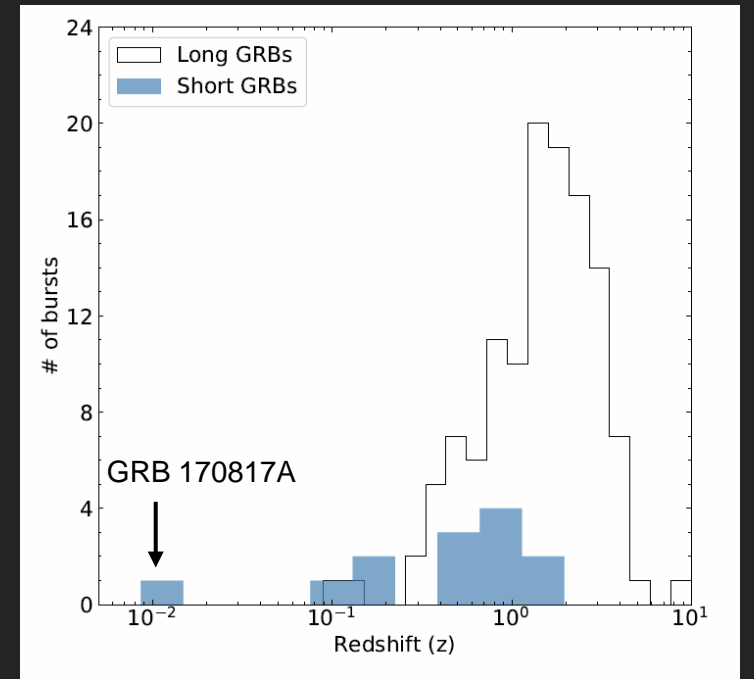
# Parameter Distributions (extract)

- ◆ All parameters: fluence- / peak flux spectra
- ◆ GOOD / BEST categories
- ◆ Observer frame
- ◆ Histogrammed probability density plots

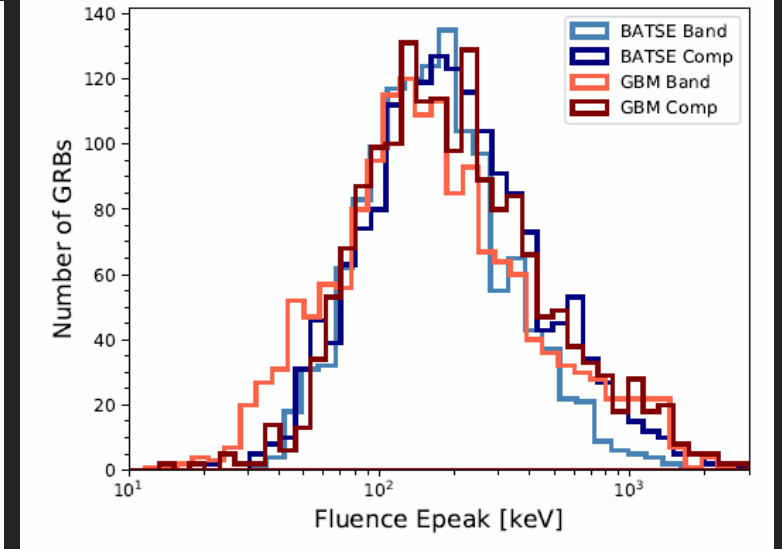
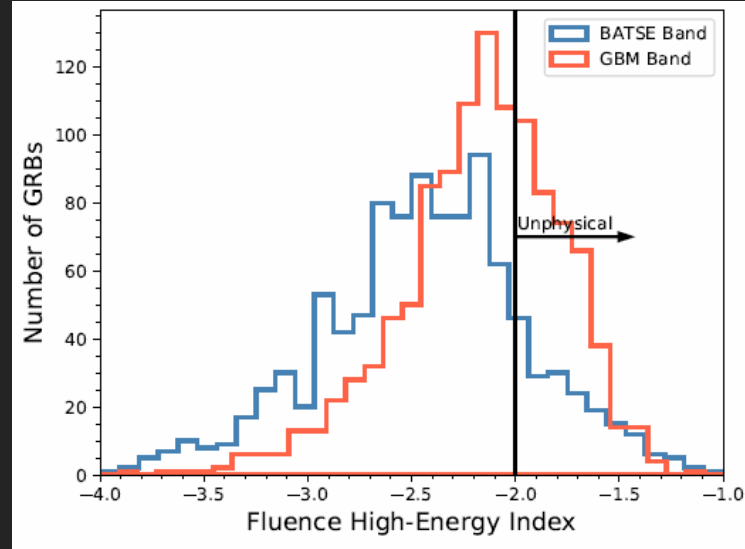
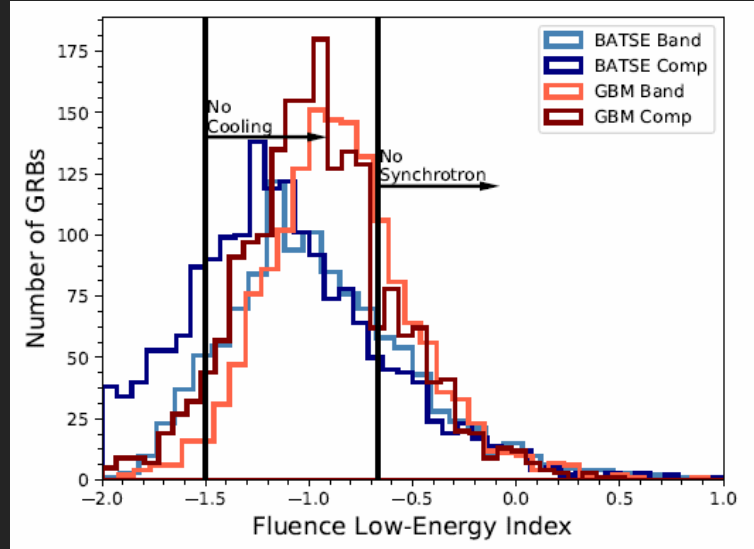


# Spectral Catalog: Rest-Frame Properties

- ◆ Redshift sample: 13 short / 122 long GRBs
- ◆ Rest frame energetics:  $E_{\text{iso}}$ ,  $E_{\text{peak}}^{\text{rest}}$ ,  $L_{\text{iso}}$ 
  - long GRBs:  $E_{\text{iso}}$  centered around  $10^{53}$  erg
  - Neither short nor long GRBs exhibit a preference in  $E_{\text{peak}}^{\text{rest}}$ ,  $L_{\text{iso}}$



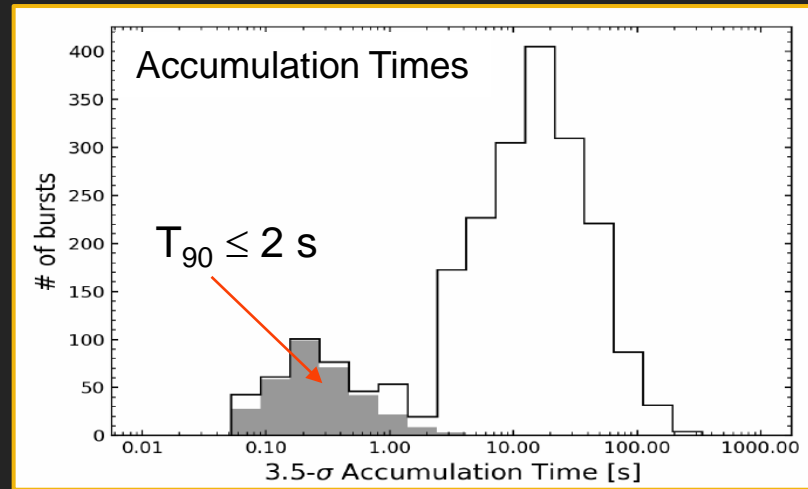
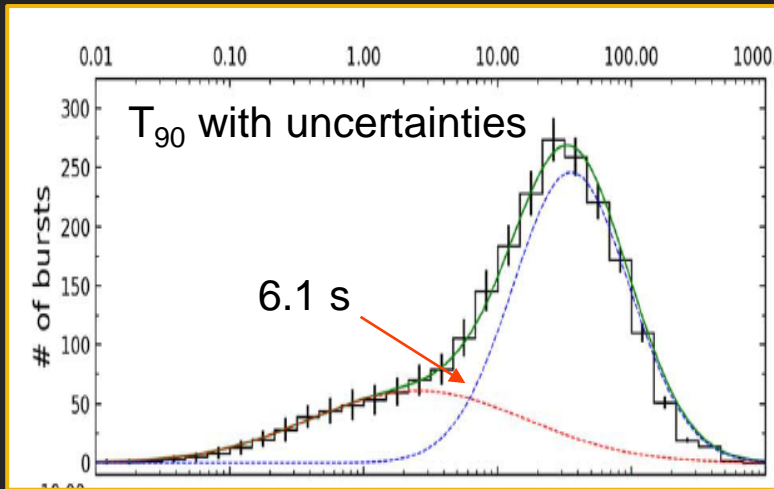
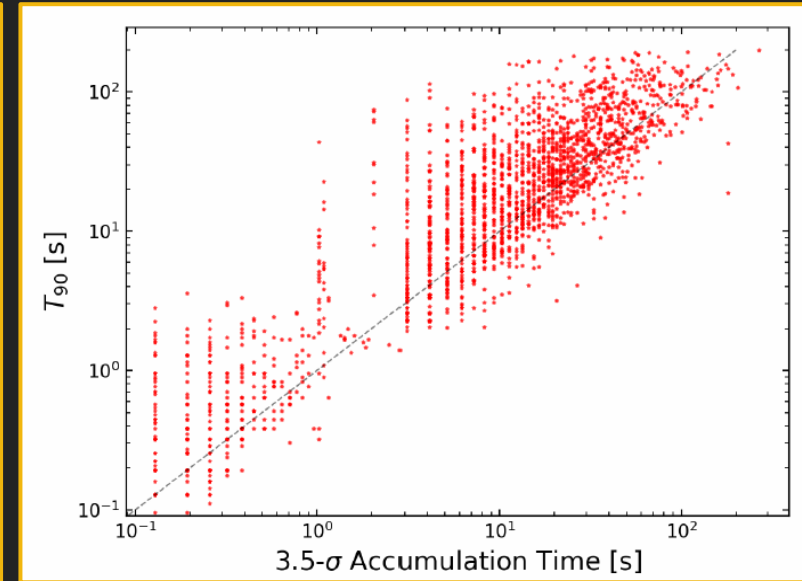
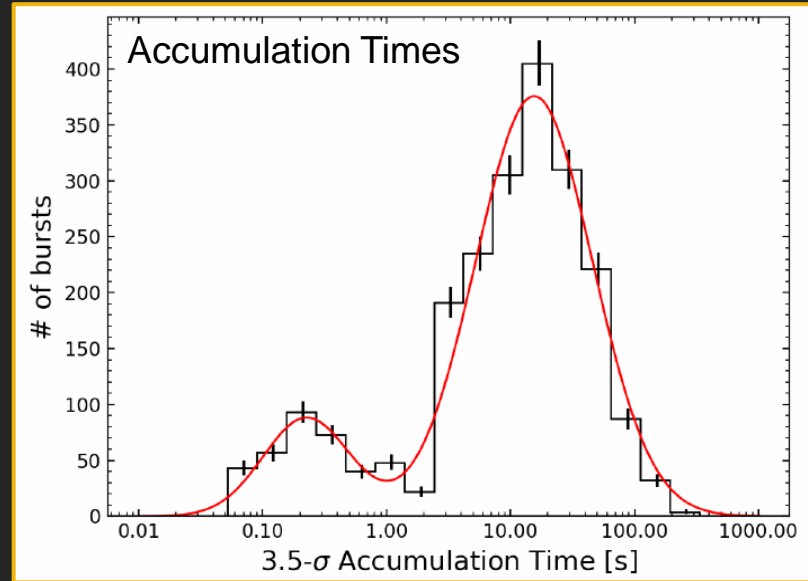
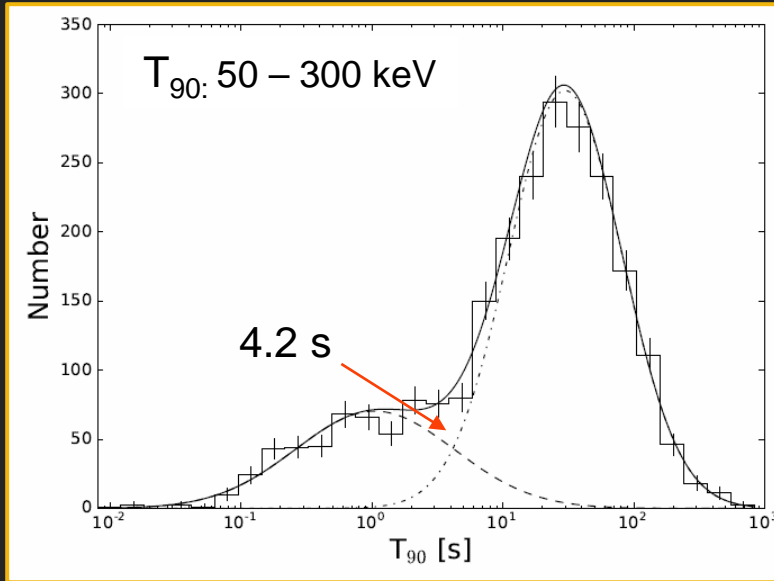
# Comparison to previous Analysis – BATSE 5B



- ◆ Low-energy photon index  $\alpha$ : for synchrotron radiation:
  - $\alpha < -2/3$  in the case of non-adiabatic cooling
  - $\alpha < -3/2$  in the case of adiabatic cooling
  - Significant fraction violate these conditions  $\Rightarrow$  synchrotron 'line-of-death' problem
  - GBM measures a slightly harder index than BATSE
- ◆ High-energy index  $\beta$ : GBM measurements are generally harder
  - $\beta \geq -2$  - unphysical - leads to an infinite flux if extrapolated in energy
- ◆  $E_{\text{peak}}$ : both distributions broadly agree

# GRB Duration Analysis

- ◆ GRB catalog  $\Rightarrow$   $T_{90}$  Durations
- ◆ Spectral catalog  $\Rightarrow$  Accumulation times based on the  $3.5\sigma$  S/N selections

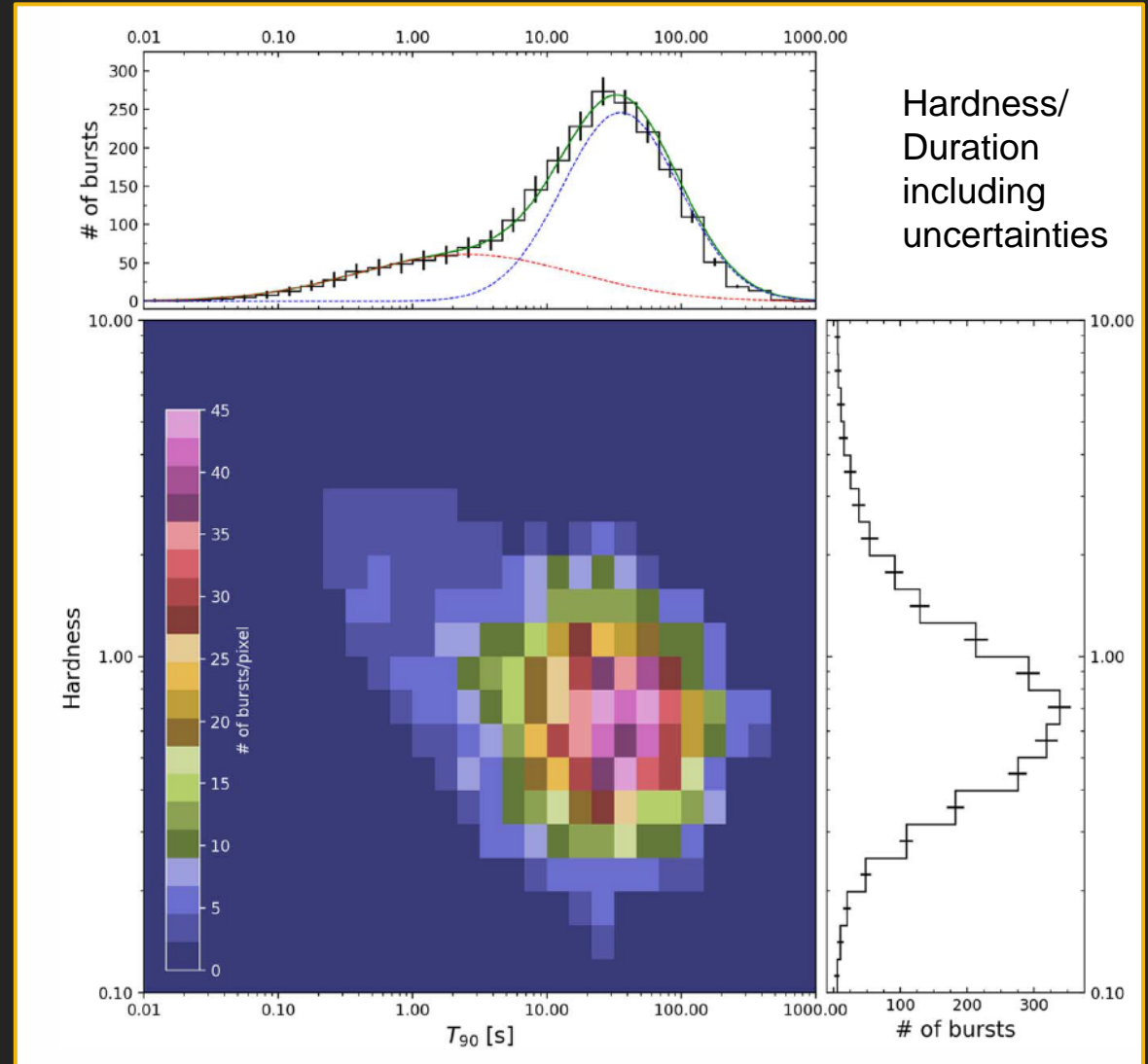
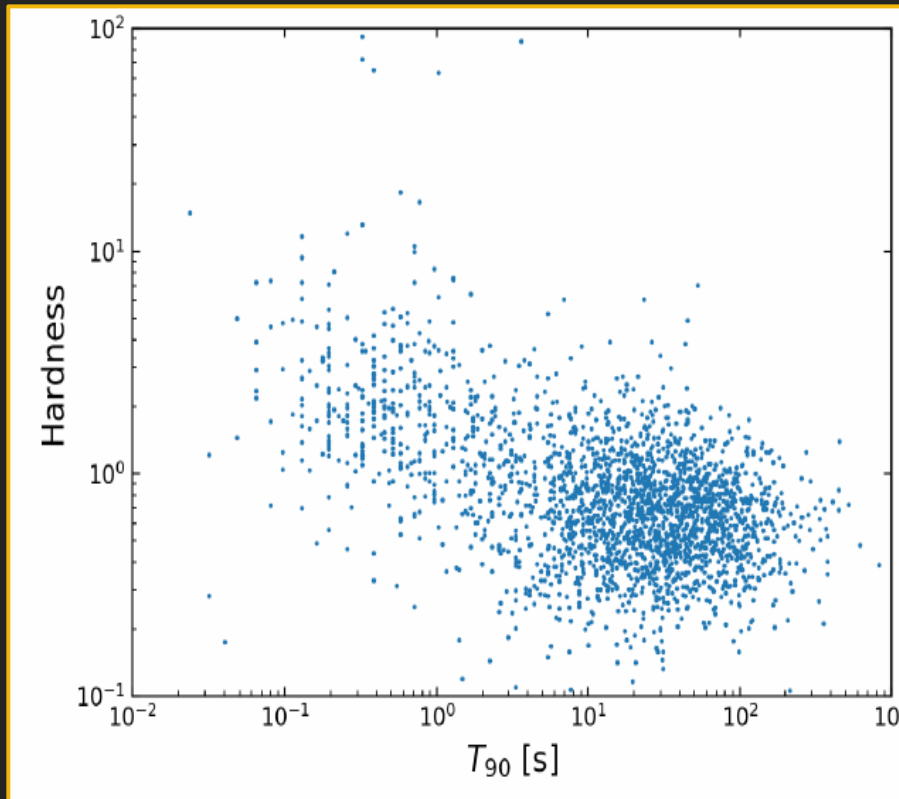




# Spectral Hardness vs. Duration

Hardness, ratio of flux density:

50–300 keV / 10–50 keV



# Conclusion:

4<sup>th</sup> GRB catalog: 2365 GRBs  $\Rightarrow$  Localization, T<sub>90</sub>, peak flux & fluence

3<sup>rd</sup> spectral catalog: 2297 GRBs  $\Rightarrow$  18000 peak flux & fluence spectra (4 models)

! Huge database and reference for catalog-based follow-up analysis

$\Rightarrow$  “Fermi-GBM GRBs with Characteristics Similar to GRB 170817A”

