

Study of Earth's Stratospheric γ-Ray Emission in Geographical Coordinates with *Fermi* LAT

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Cosmic-Ray (CR)-Induced y-Ray Emission of Earth



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Earth's γ-Ray Zenith Profile



- Peak of profile moved over time due to LAT orbital decay
- Use data from Peak < Θ_{Zen} < Peak + 2.0° because we can assume that these γ rays were produced at ~50-km altitude (top of stratosphere)
- γ rays with $\Theta_{\rm Zen}$ < peak were produced at unknown altitude

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Vertical Geomagnetic Cutoff Rigidity



Earth's magnetic field blocks CRs below cutoff rigidities for certain locations

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Data Set



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Earth's Stratospheric γ-Ray Intensity Maps



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Earth's γ-Ray Intensity vs Cutoff Rigidity



- Cutoff rigidity = minimum rigidity for CRs to reach the top of atmosphere at certain location (here for the horizontal direction)
- Earth's γ -ray intensity decreases for increasing cutoff rigidity because CR flux falls steeply with energy
- Minimum CR energy to produce γ rays = the lower bound of the energy bin

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Earth's Stratospheric γ-Ray Yield between 0.2 – 20 GeV



• Assume only CR protons and He, and $Y_{\text{He}} = 1.6Y_p$, $Y_p(P) =$

$$P) = -\frac{[dI/dP_c]_P}{J_p(P) + 1.6J_{\text{He}}(P)}$$

• Interpretation: Y_p = Earth's γ -ray intensity per rigidity emitted from the stratosphere (~50 km) as observed by the LAT divided by CR proton intensity per rigidity near Earth

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Earth's γ-Ray Spectrum



Summary

- Using *Fermi* LAT data, we obtain the first geographical maps of the Earth's stratospheric γ-ray emission from CR air showers
- We study associations between γ-ray emission intensity and the geomagnetic cutoffs for CRs
- We calculate the stratospheric γ-ray yield function between 0.2 – 20 GeV due to CR protons
- This work presents unique data on CR interactions with the heliosphere, the geomagnetic field, and the atmosphere



Back Up 1: High-Cutoff Neutron Monitor Count Rate



Back Up 2: Atmospheric Column Density

