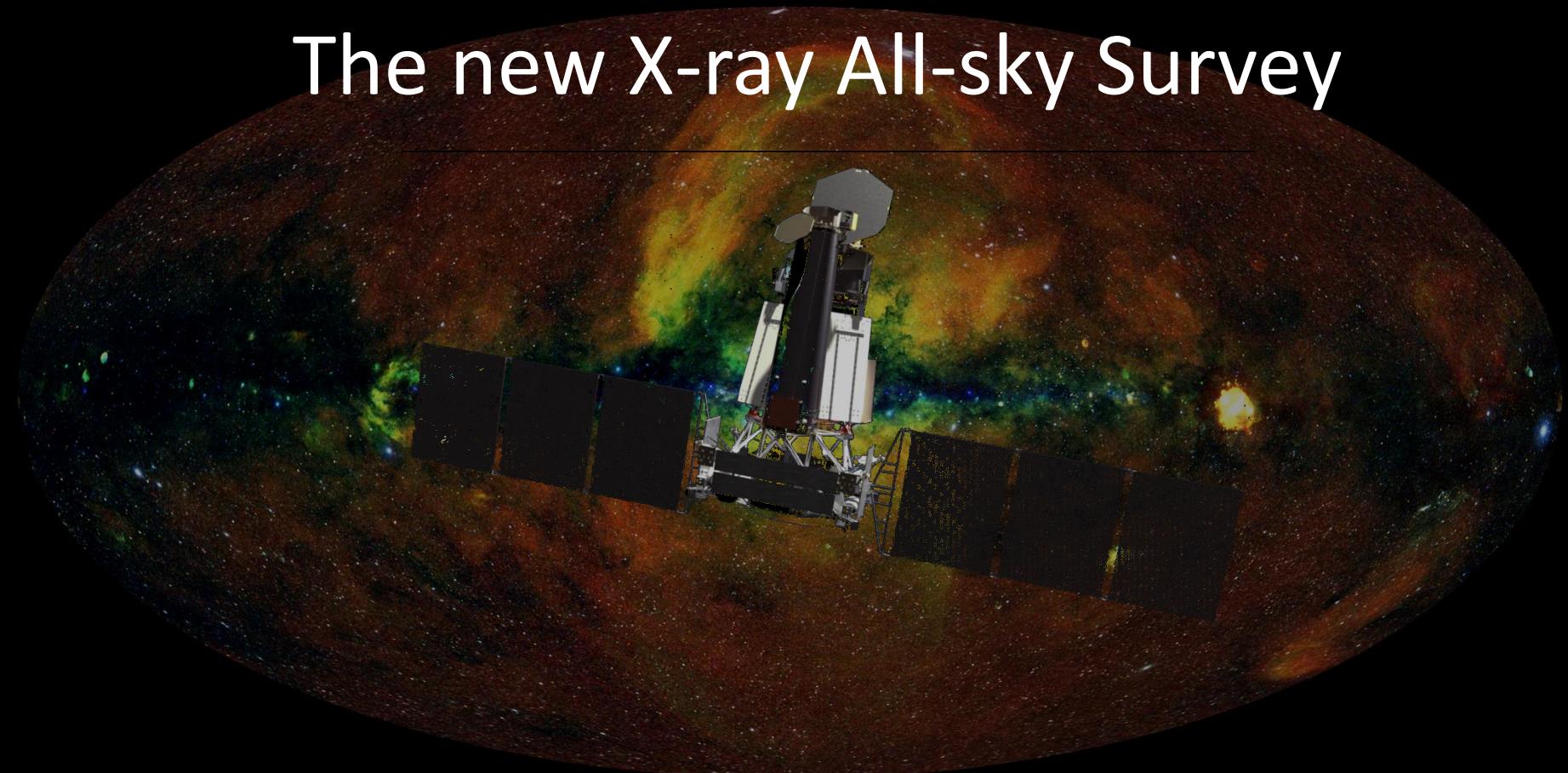
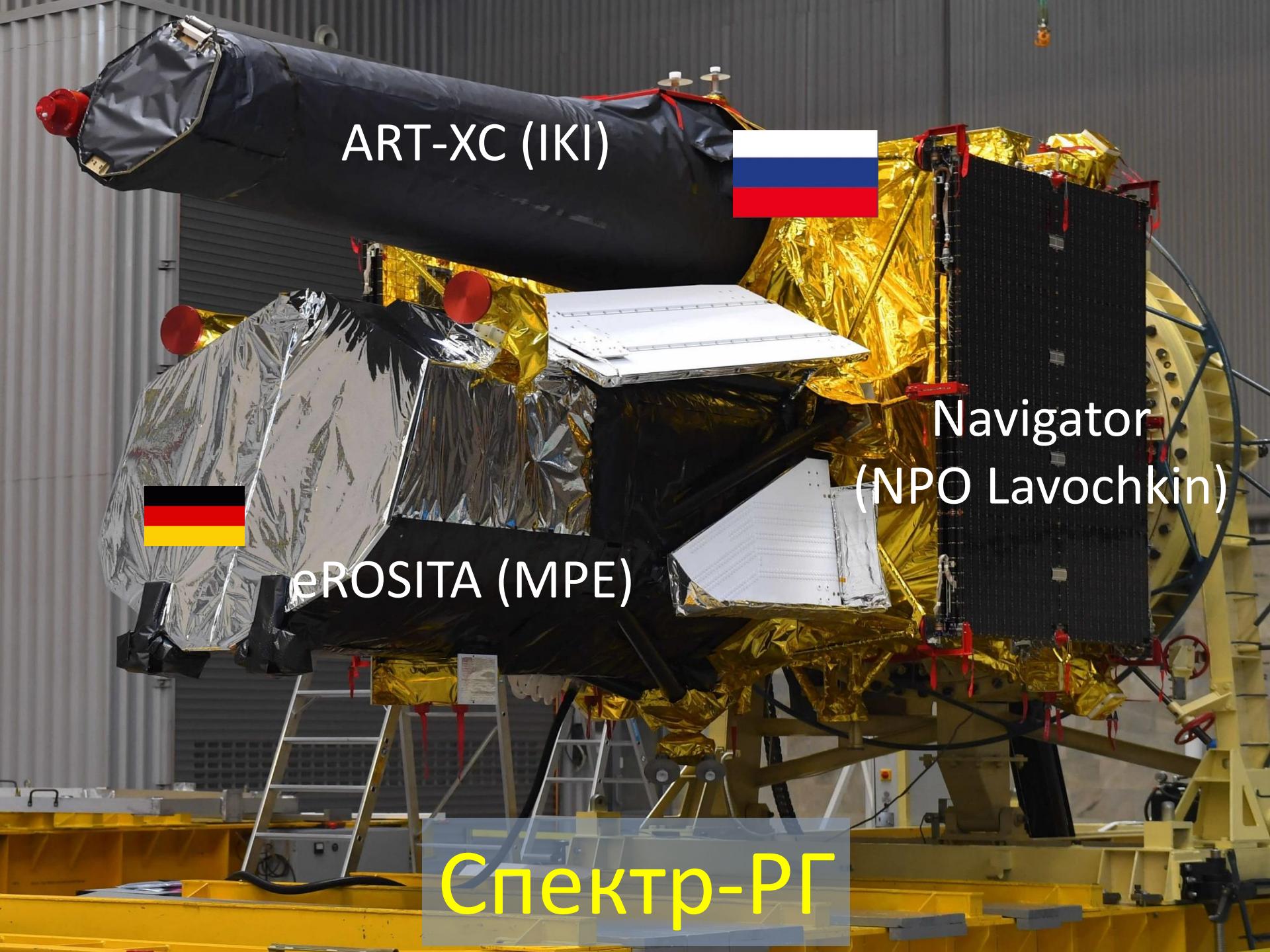


eROSITA on SRG

The new X-ray All-sky Survey



Peter Predehl (MPE)
Fermi Symposium, April 2021



ART-XC (IKI)

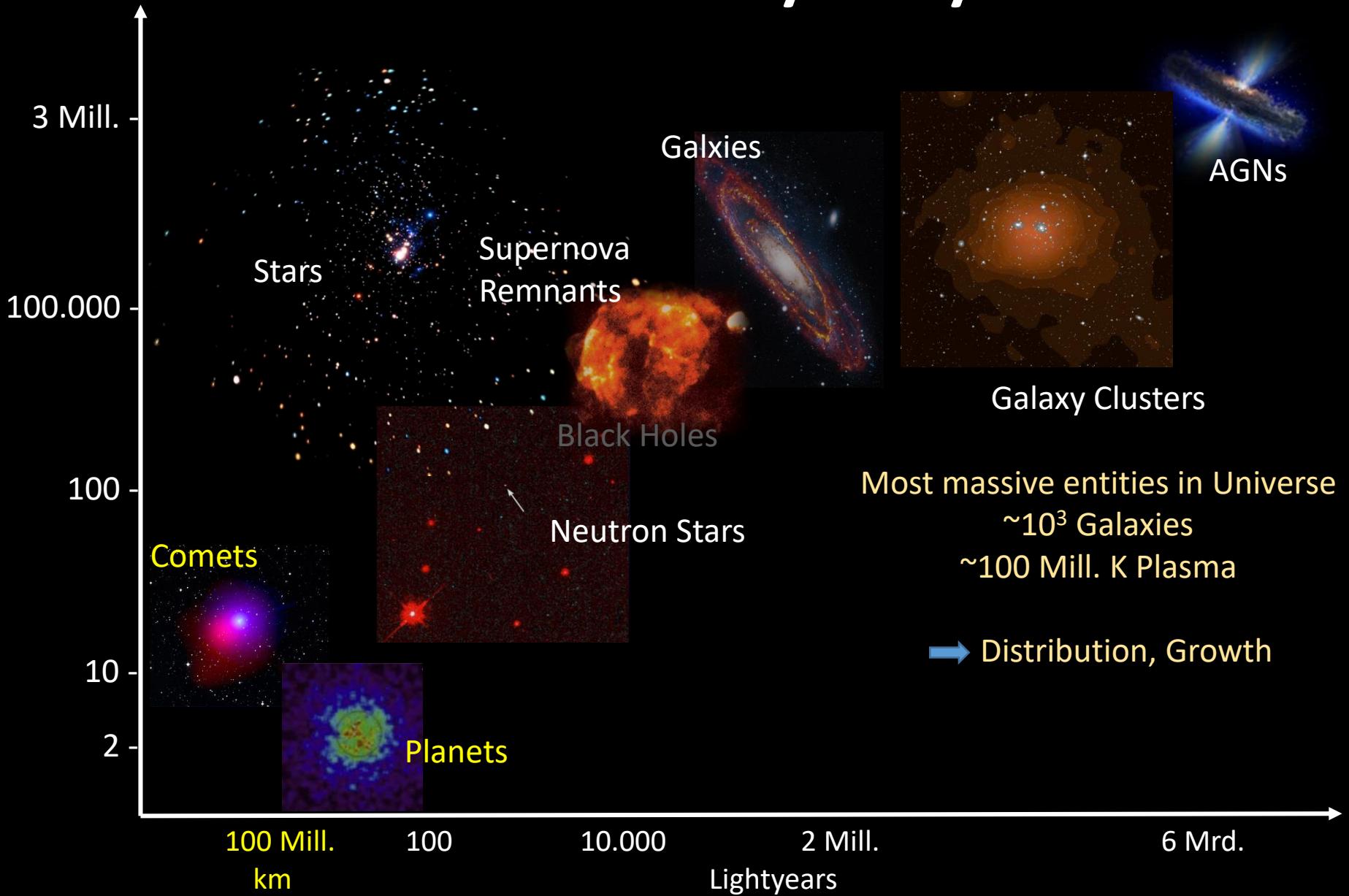


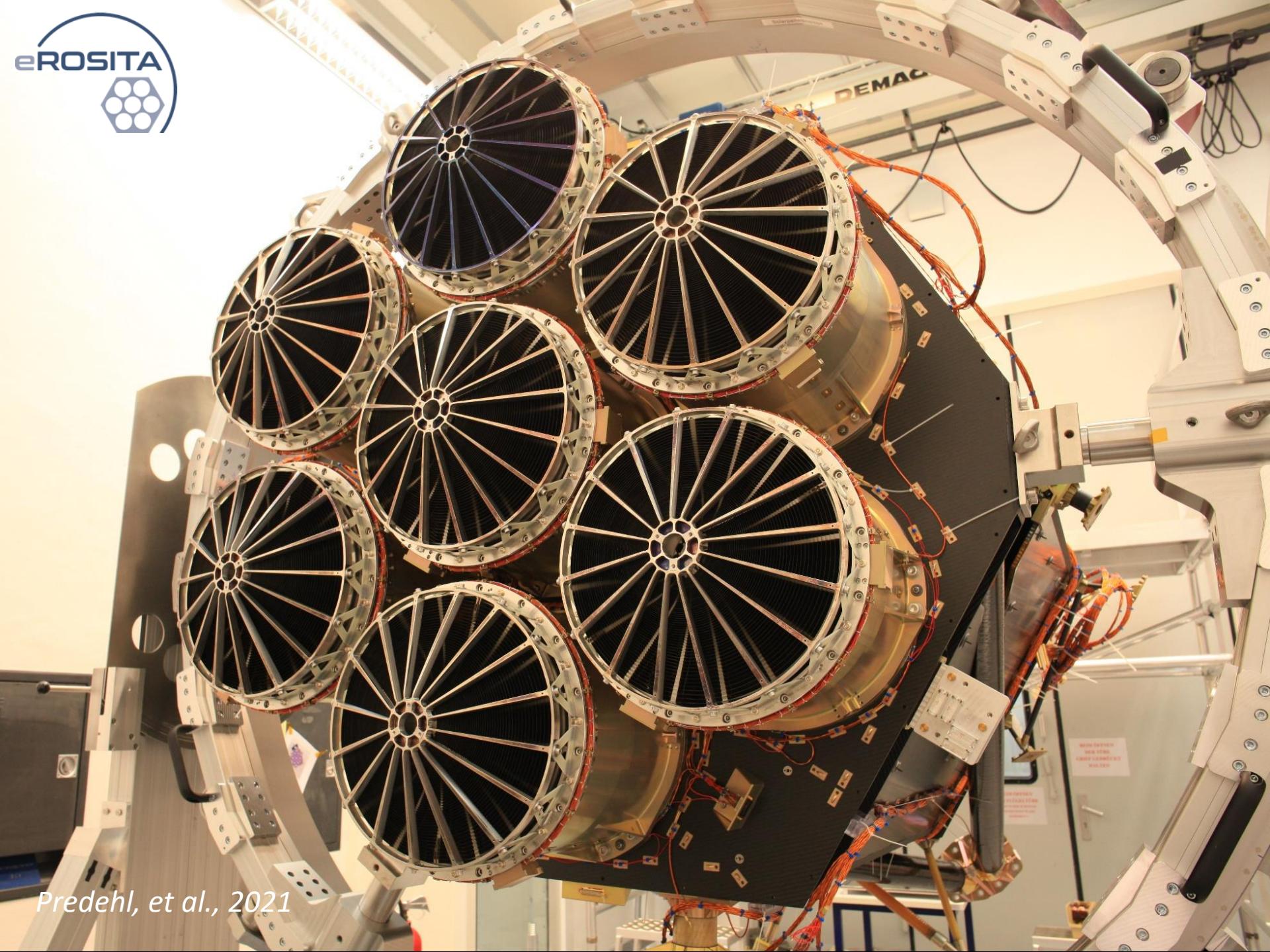
eROSITA (MPE)

Navigator
(NPO Lavochkin)

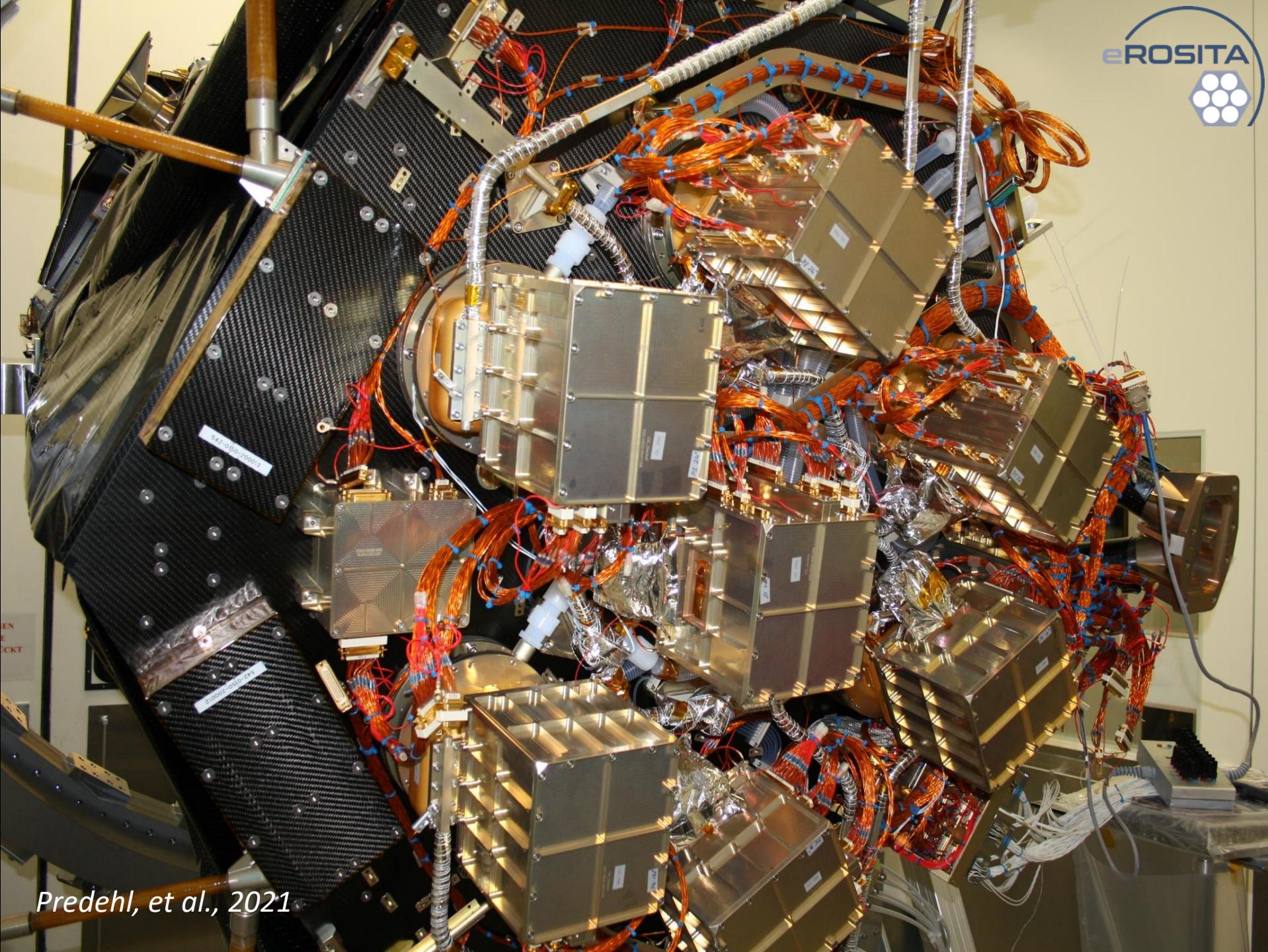
Спектр-РГ

The X-ray Sky





Predehl, et al., 2021

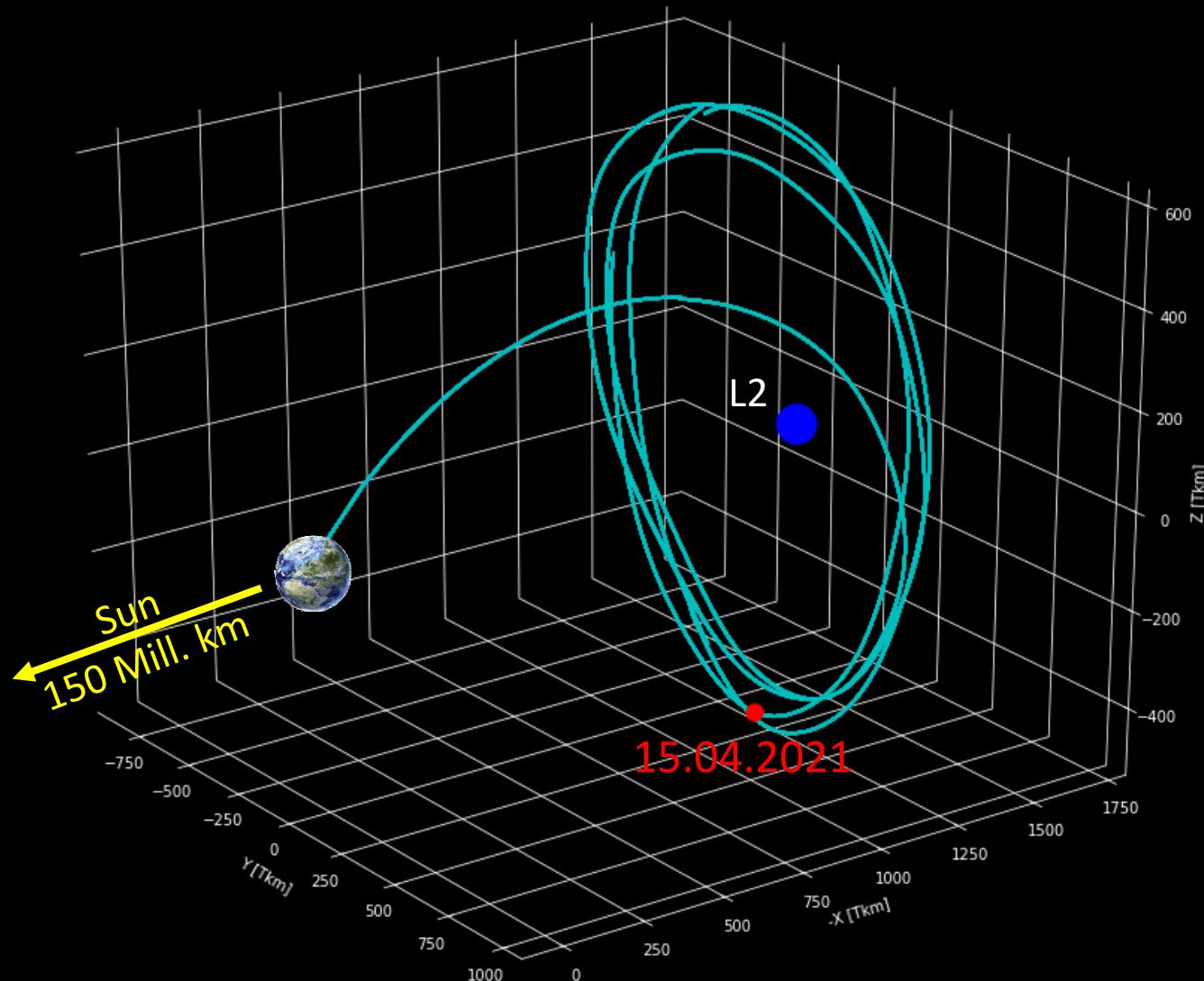


13.07.2019, 17:31

Cosmodrom Baikonur/Kazakhstan
Proton-M / BLOK-DM03



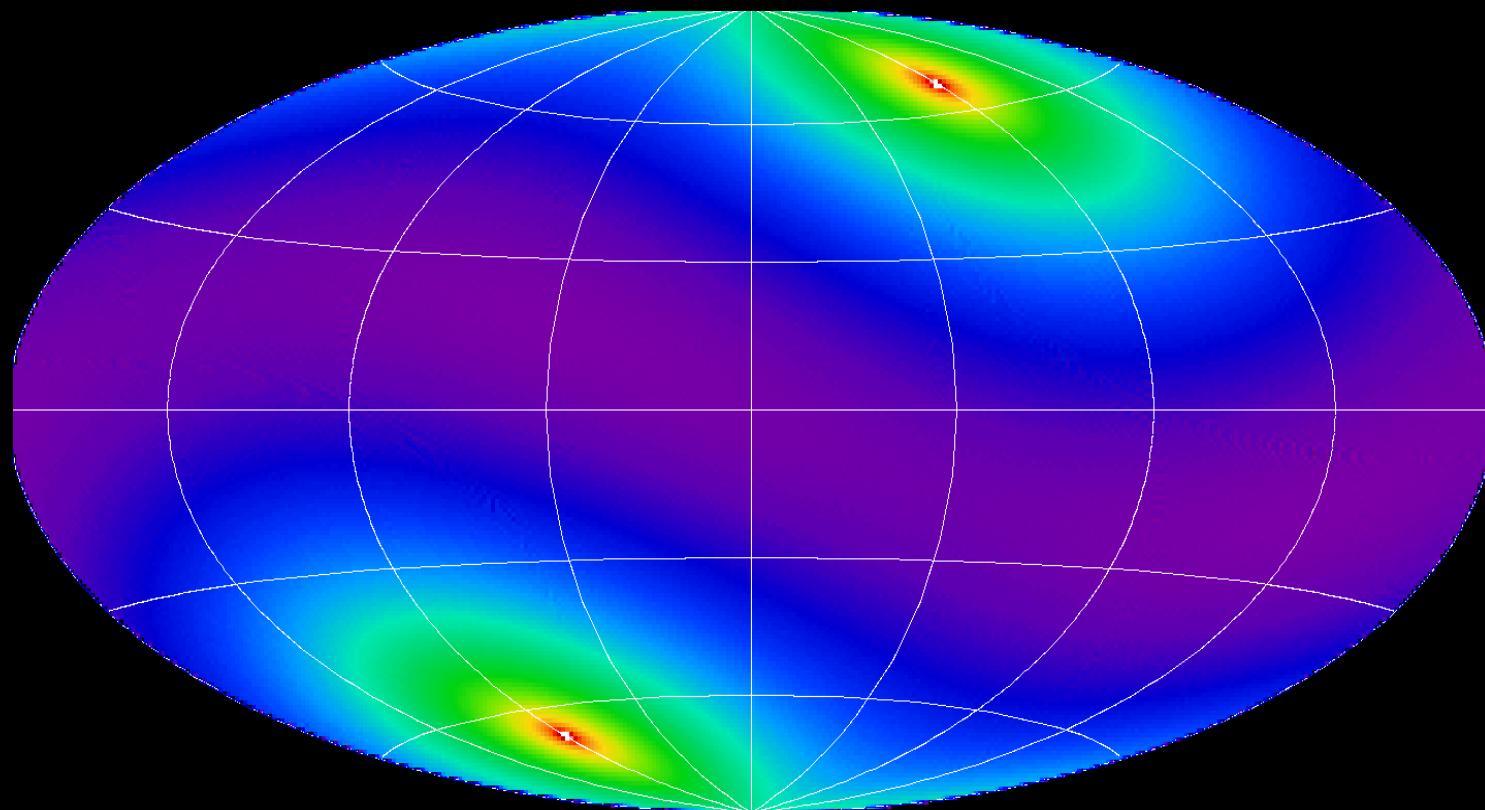
Credit: Roscosmos



- 4 years:
- 2.5 years:

8 all sky surveys (6 rotations/day)
pointed observations

eROSITA Cadence Map



→ # of daily eROSITA visits over 4yrs

Exp.: 10^3 sec

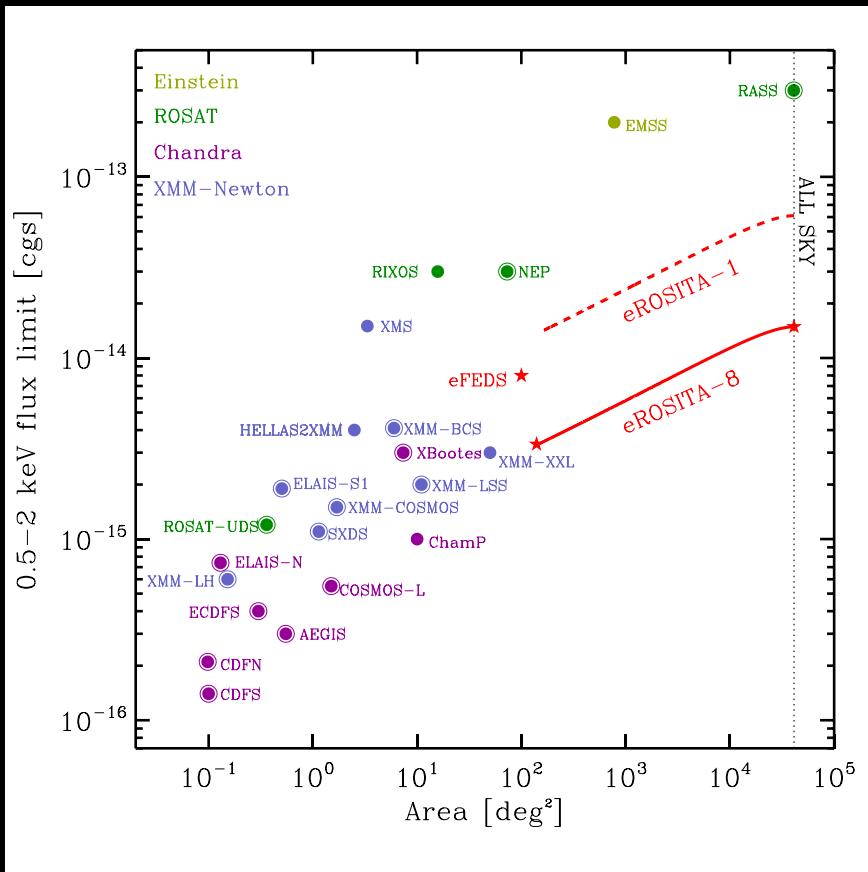
10^4 sec

src. conf.

eROSITA surveys in context

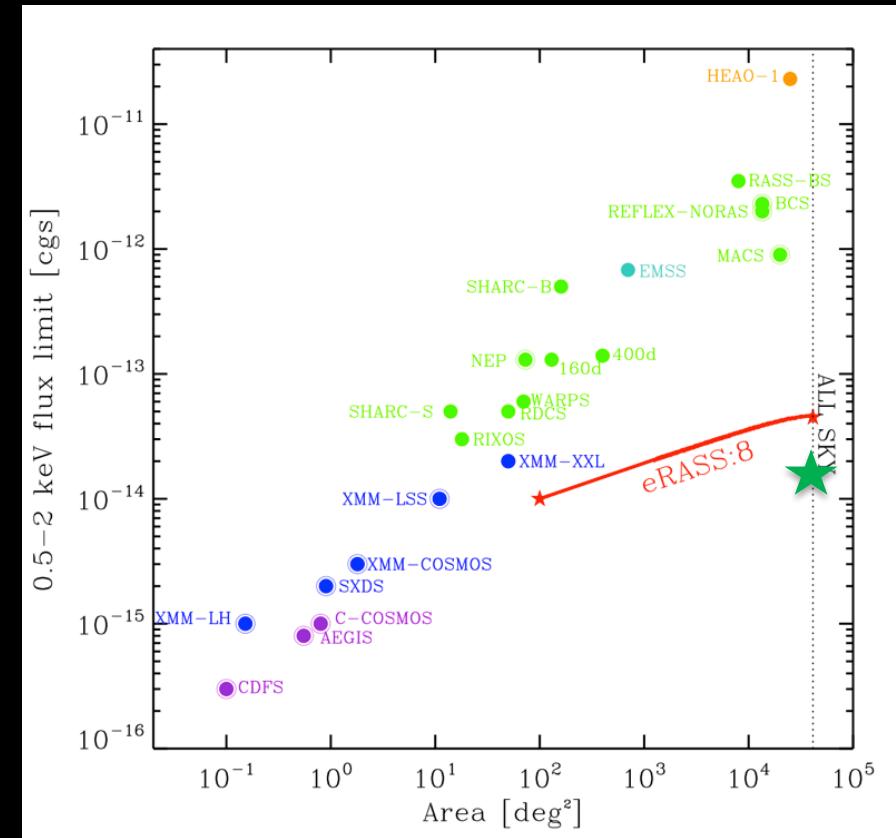


Point sources sensitivity



All sky: 10^{-14} [$\text{erg}/\text{cm}^2/\text{s}$] (0.5-2 keV)

Extended sources sensitivity



All sky: 3.4×10^{-14} [$\text{erg}/\text{cm}^2/\text{s}$] (0.5-2 keV)
 <1.5 *Liu et al. 2021*

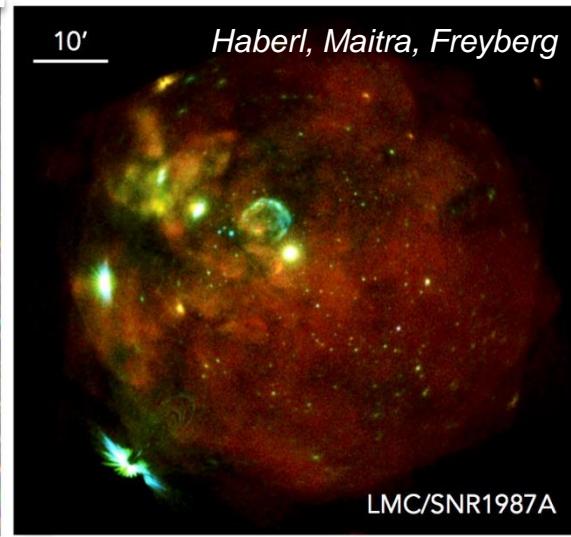
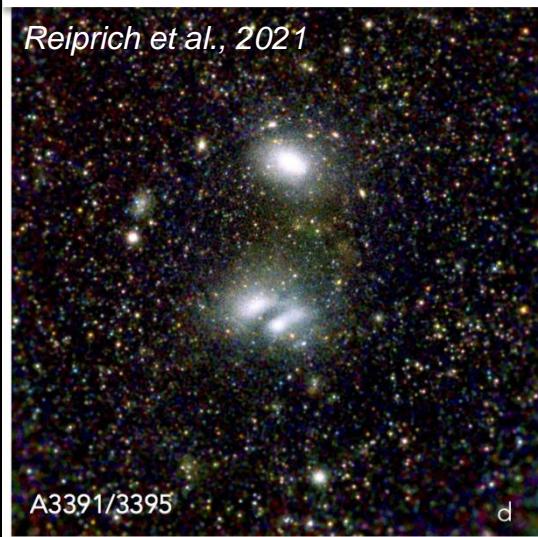
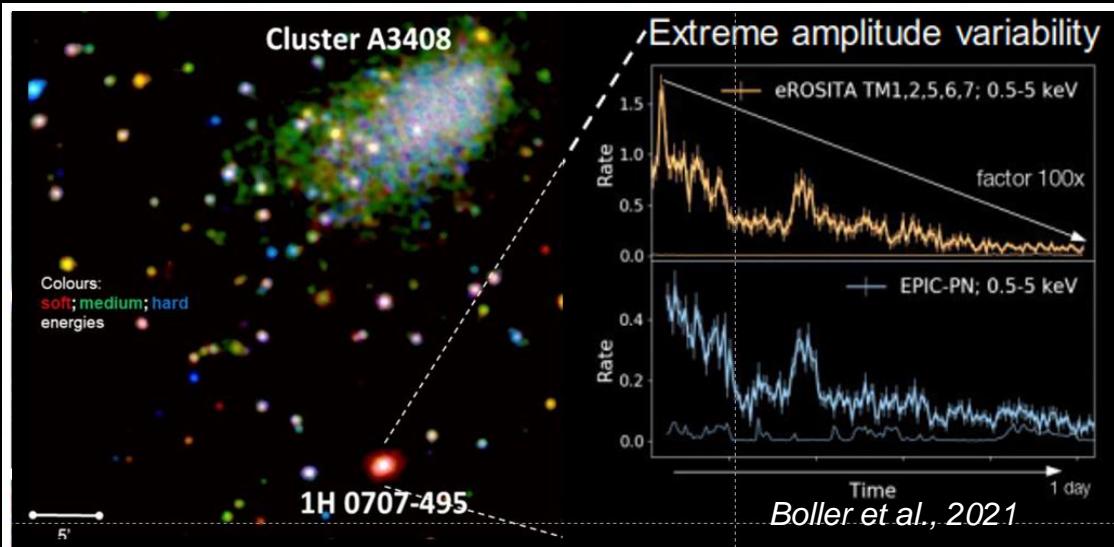
Merloni et al. 2012

Time Domain Astrophysics with eROSITA

- 50msec time resolution
- 40sec scanspeed + 1 deg FoV
- 4 hours rotation period of SRG
- 1 day overlapping scans at ecliptic equator
every day at the ecliptic poles
- half year one complete survey
- 4 years 8 surveys

Pulsars, GRBs, Flares from stars and AGN, TDEs

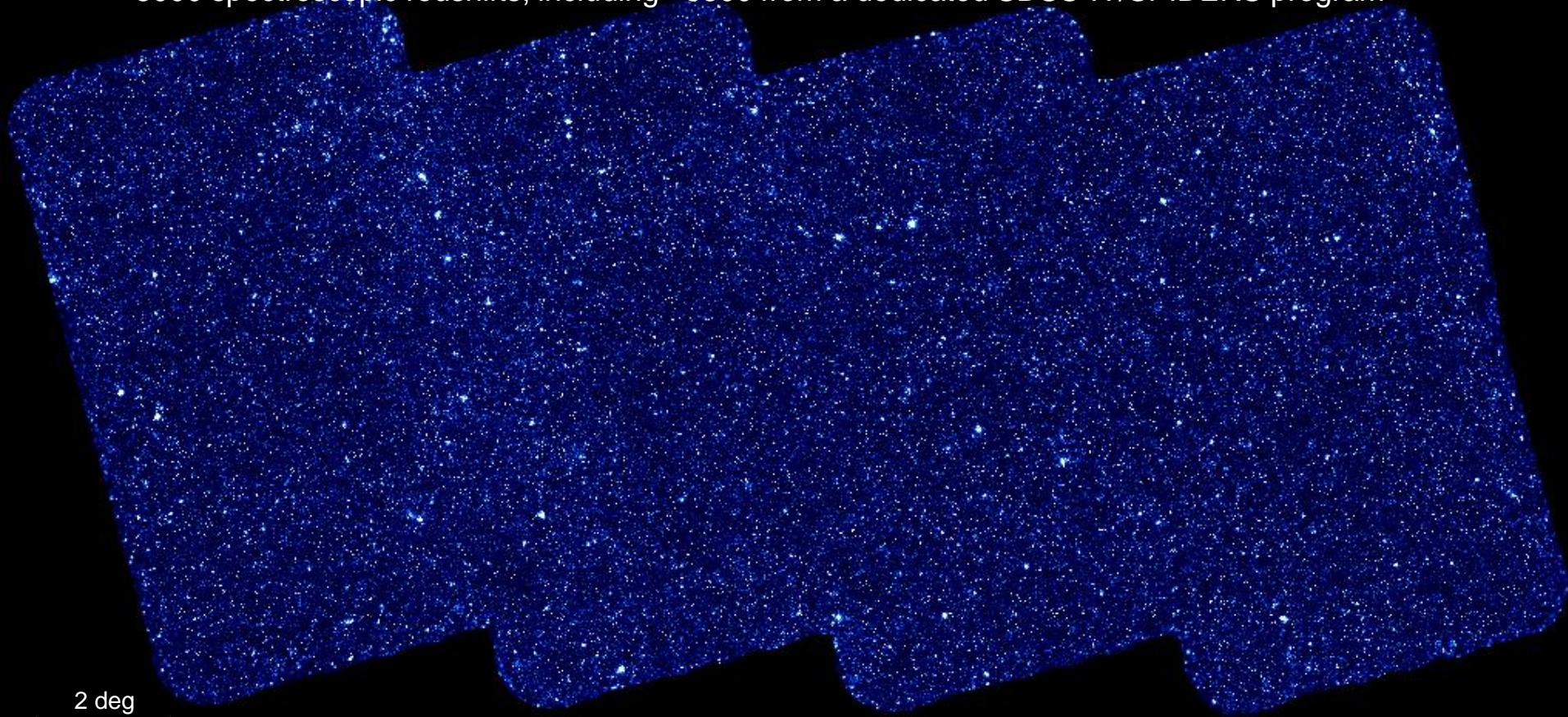
Highlights from CalPV phase



eFEDS AGN

More than 25k point-sources detected

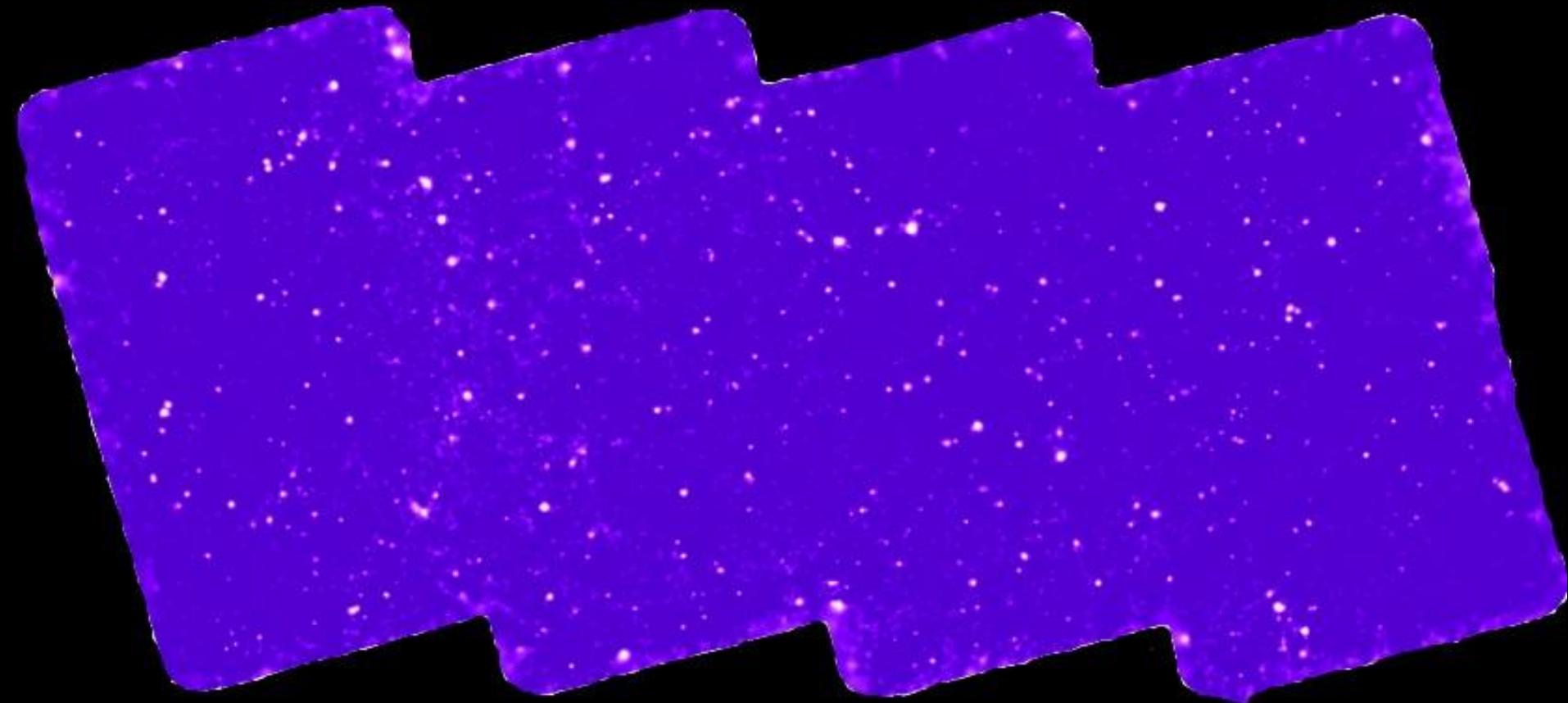
- ~8000 spectroscopic redshifts, including ~3800 from a dedicated SDSS-IV/SPIDERS program



0.2-2.3keV, exposure and vignetting corrected, 1.2ksec

eFEDS Clusters

542 galaxy clusters detected by eROSITA
~ 440 already optically confirmed $0.1 < z < 1.3$

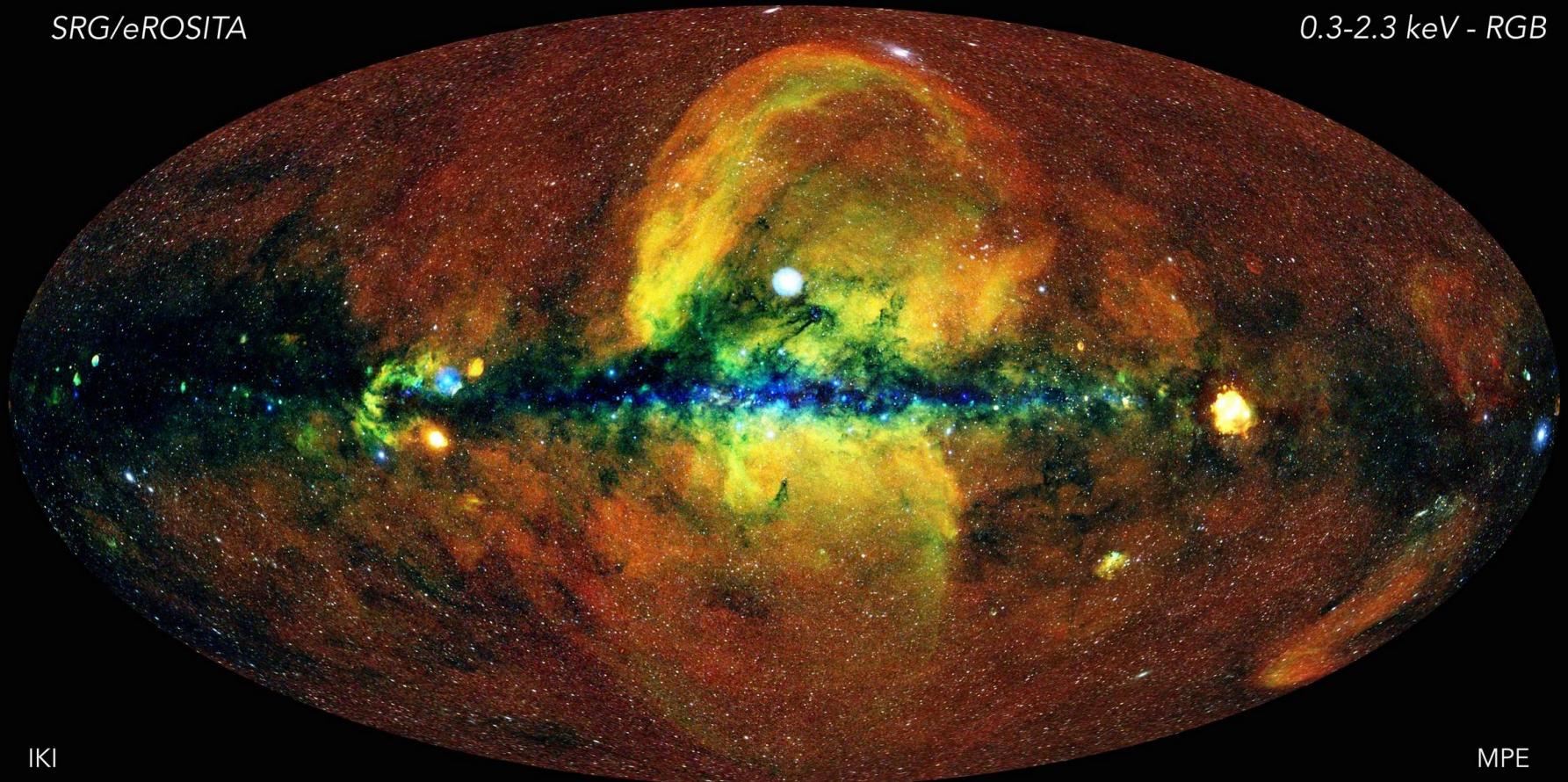


0.2-2.3keV, exposure and vignetting corrected, 1.2ksec

eRASS1

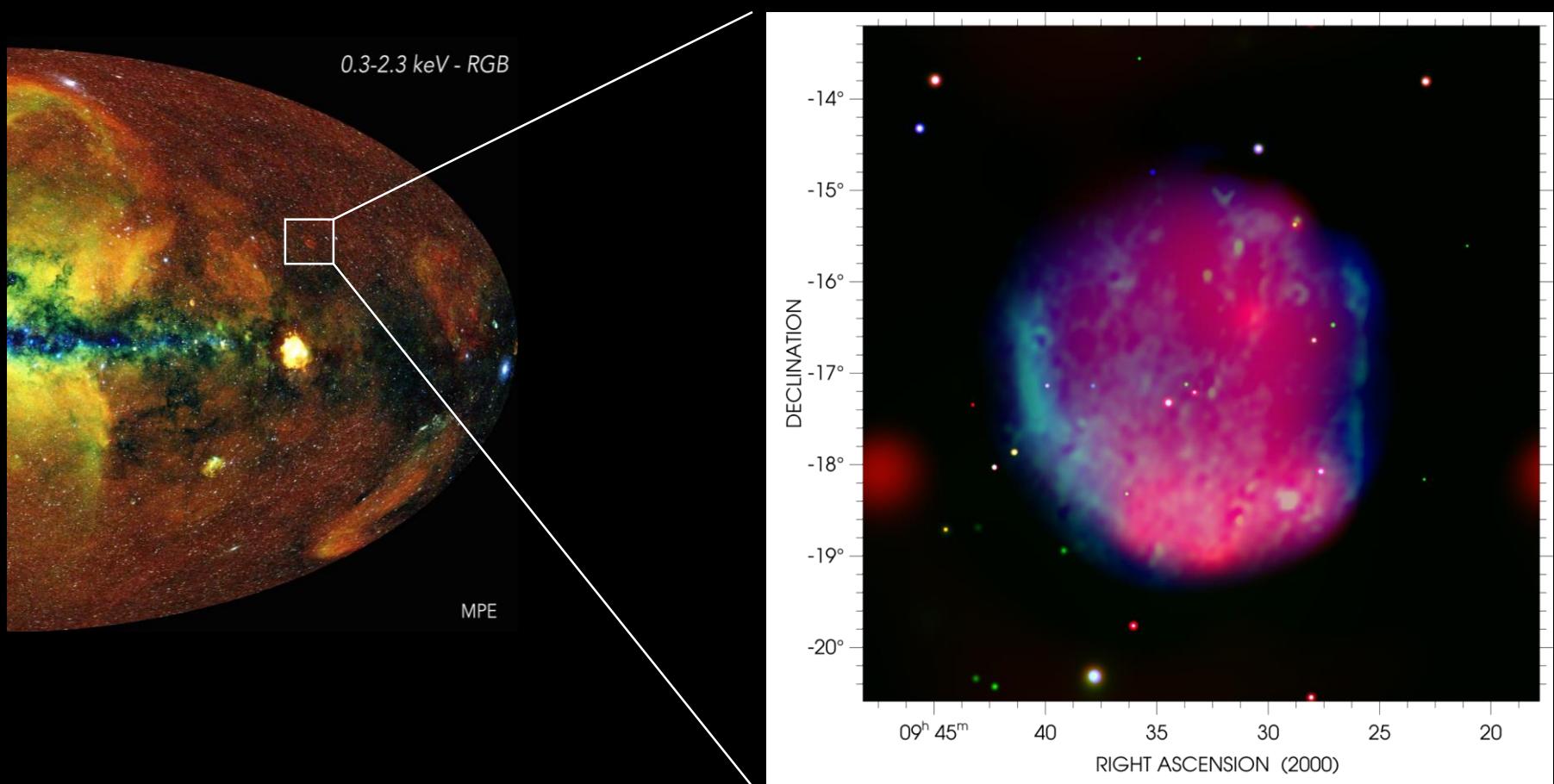
SRG/eROSITA

0.3-2.3 keV - RGB



> 1.000.000 X-ray sources found

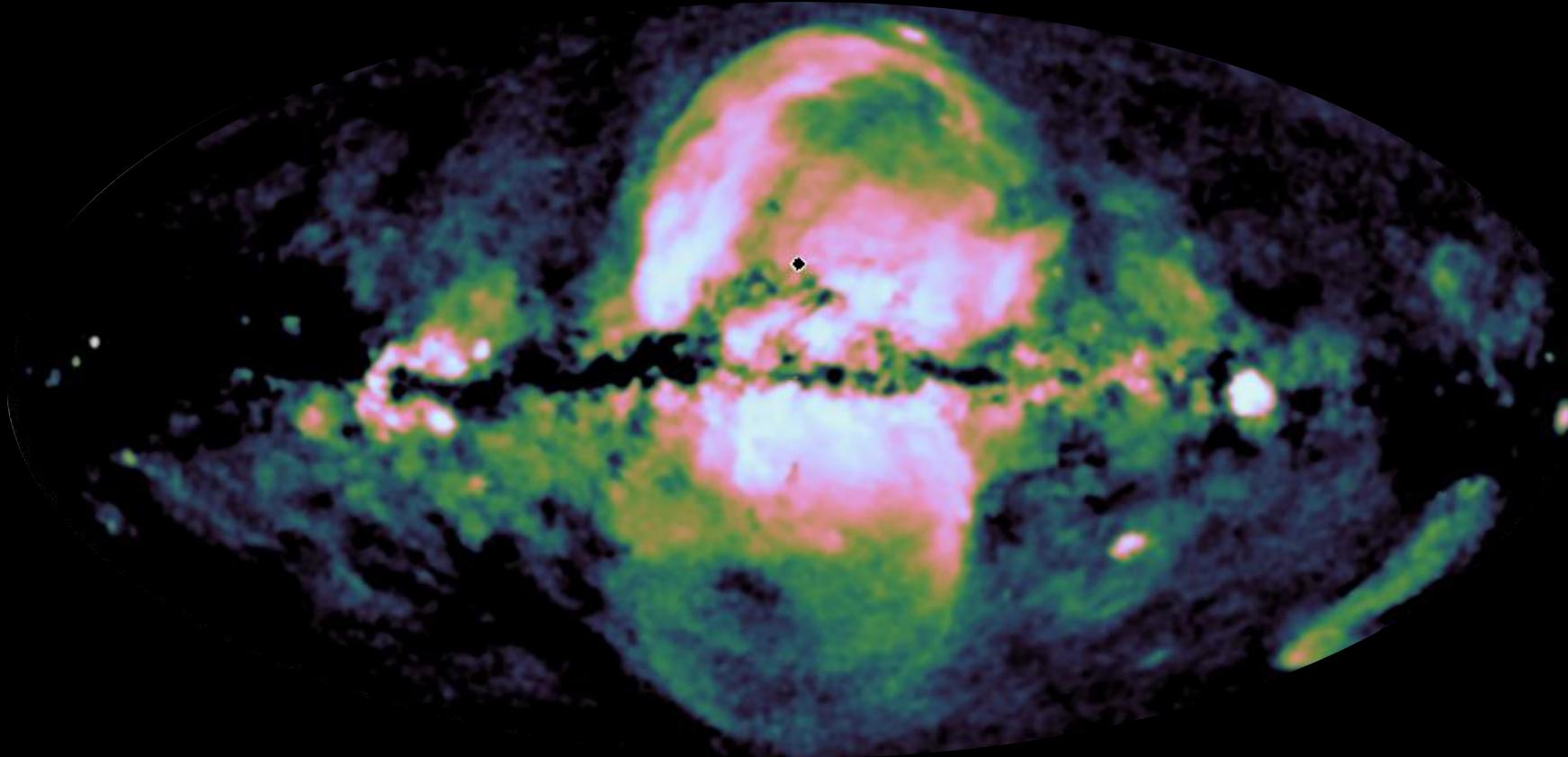
Hoinga – the largest supernova remnant ever discovered with X-rays



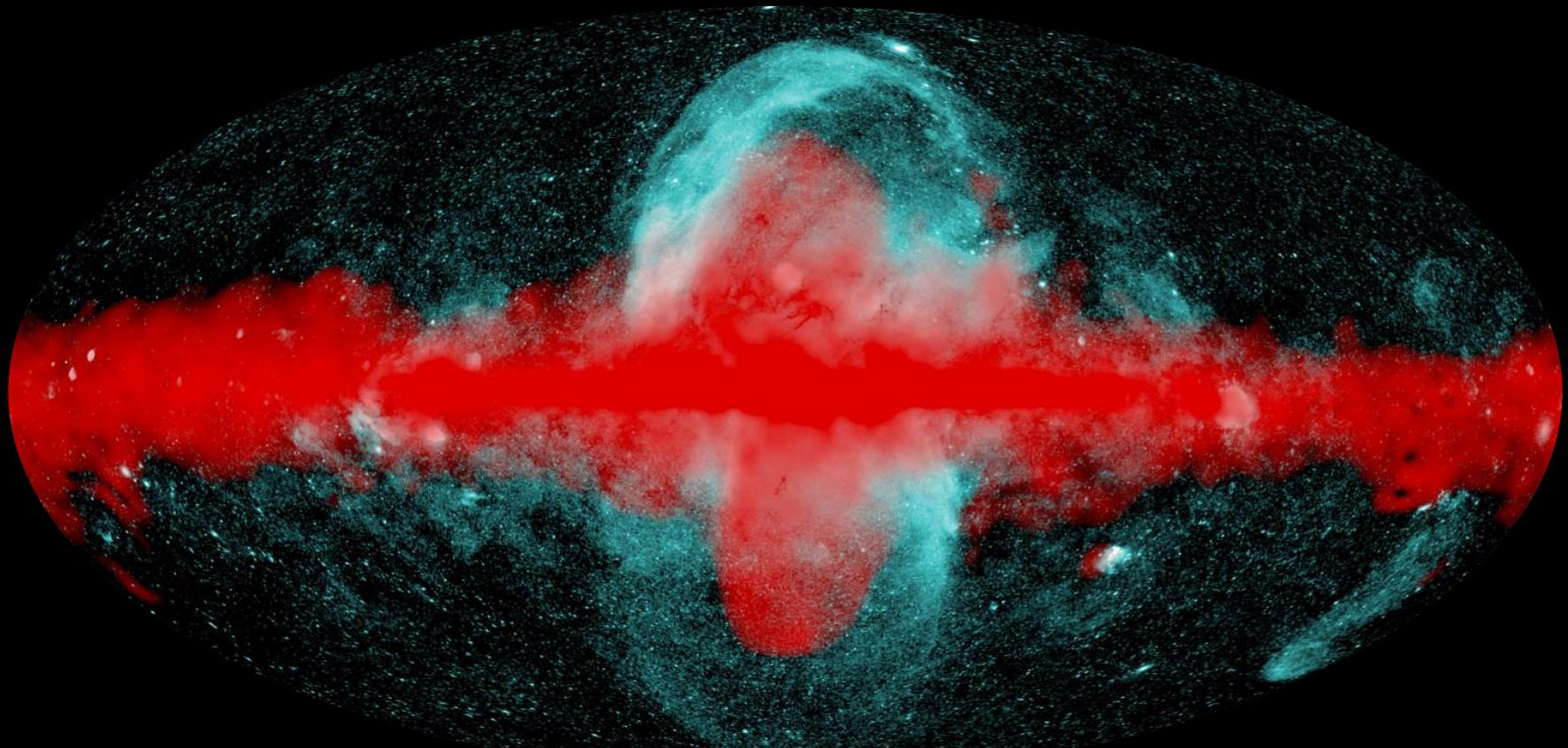
SNR: $4.4^\circ \varnothing$, d~500pc, 17 - 30.000 yrs

Becker et al., 2021

„eROSITA Bubbles“

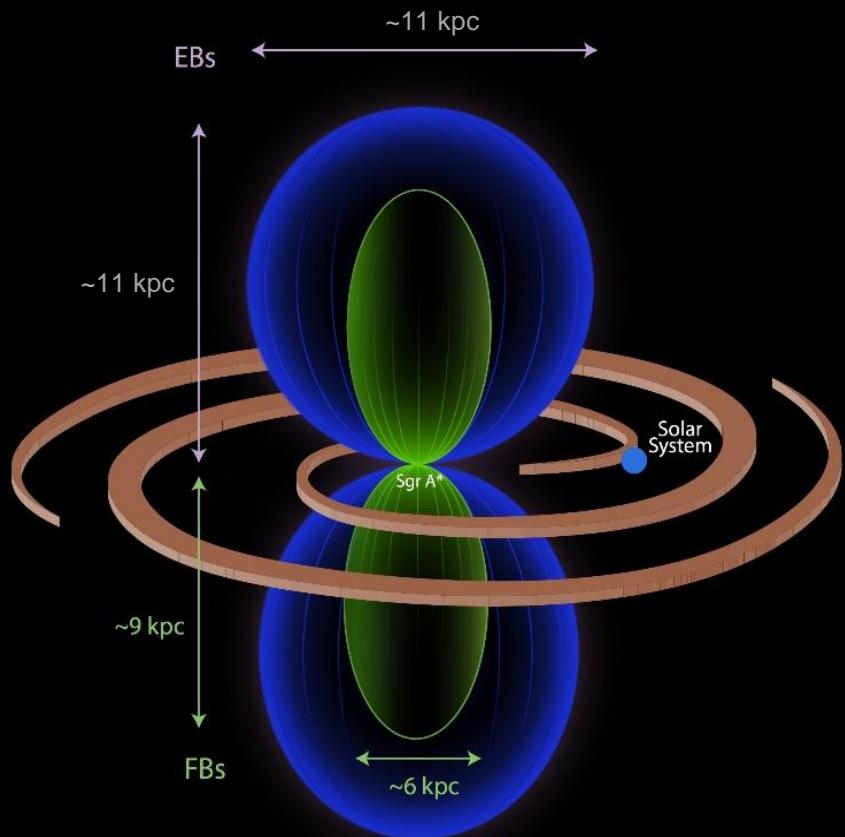


eROSITA & Fermi Bubbles



- Similar morphology
 - Boundary between eRO & Fermi is not clear
 - Fermi = symmetric, eRO = not
 - Starburst or AGN activity
- common GC origin?
 - 2 consecutive events?
 - NPS/Loop I somewhat peculiar
 - MV quiescent only now?

eROSITA Bubble



Some numbers:

distance	~ 10 kpc
brightness	6×10^{38} ergs $^{-1}$
kT	0.3 keV
t _{cool}	1.9×10^8 yr
v _{exp}	~ 300 kms $^{-1}$
age	~ 20 Mio yr
E _{thermal}	$\sim 10^{56}$ erg
abundance	0.2 \odot (halo)

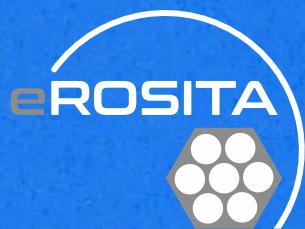
since April 2020



Operations @Home



Coutinho et al., 2021

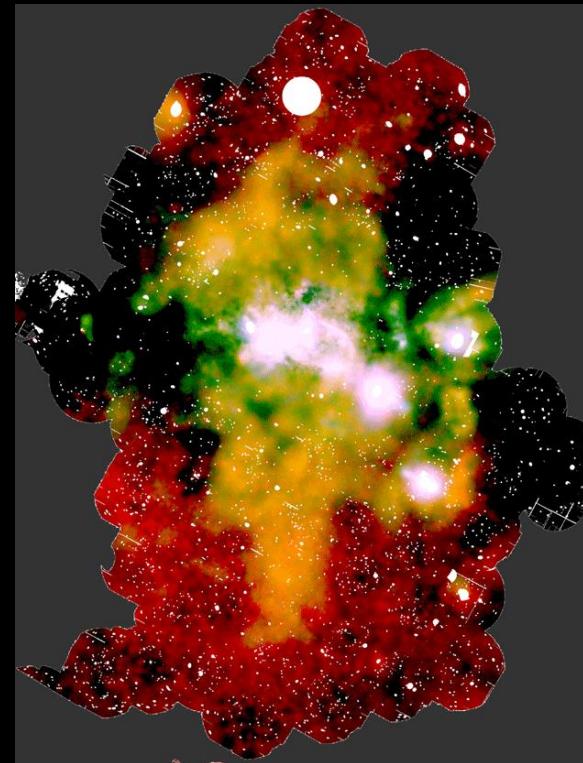
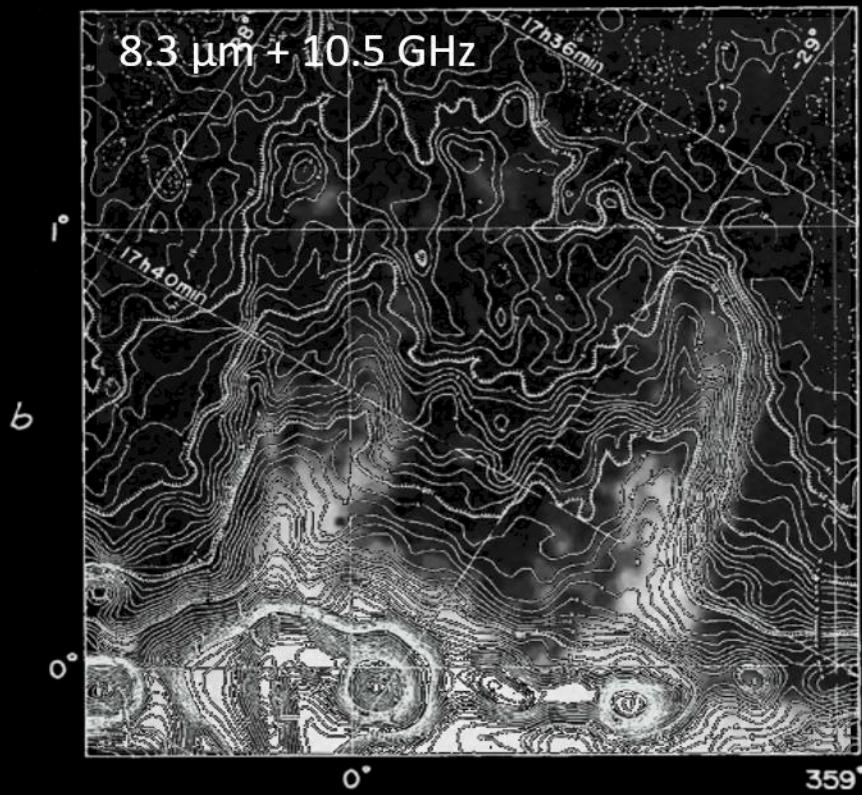


Thank you very much
for your attention



Photo: V. Burwitz (MPE)

Channeled outflow from GC



Bland-Hawthorn *et al.*, 2003

Ponti *et al.*, 2019

eROSITA Collaboration

Core Institutes (DLR funding):

MPE, Garching
University Erlangen-Nuremberg
IAAT (University Tuebingen)
SB (University Hamburg)
Leibniz-Institute for Astrophysics Potsdam

Associated Institutes:

USM (LMU Munich)
AIFA (University Bonn)

Russian Partner Institute:

IKI, Moscow

Industry:

Media Lario/I Mirrors, Mandrels
Tecnotron/D PCBs
Kayser-Threde/D Mirror Structures
Carl Zeiss/D ABRIXAS-Mandrels
Invent/D Telescope Structure
pnSensor/D CCDs
IberEspacio/E Heatpipes
RUAG/A Mechanism
HPS/D,P MLI
+ many small companies

NPOL – Lavochkin Association

MPE: Scientific Lead Institute, Project Management

Instrument Design, Manufacturing, Integration & Test

Operation, Data Handling & Processing, Archive etc.

