



SELF-CONSISTENT CALCULATIONS OF SOLAR CNO NEUTRINO CAPTURE-RATES FOR ^{115}In .

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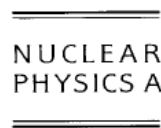
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Gamow–Teller strength functions of superfluid
odd- A nuclei and neutrino capture reactions

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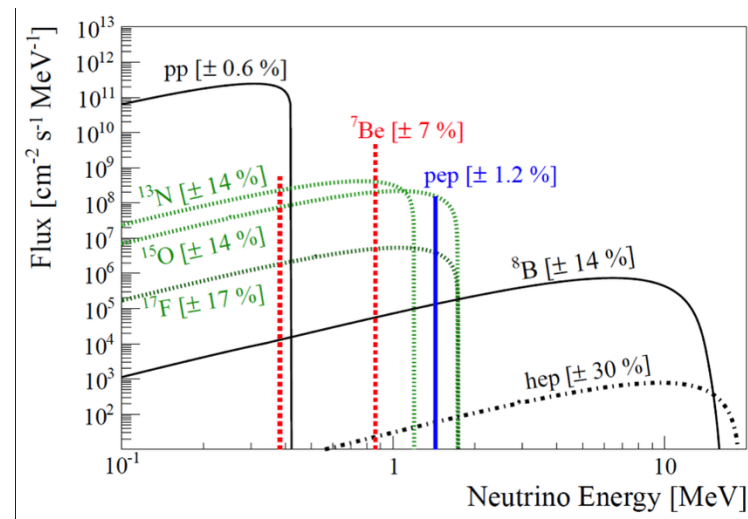
SELF-CONSISTENT APPROACH TO NEUTRINO CAPTURE

The practical aspect of the problem concerns solar-neutrino [15] and reactor-anti-neutrino capture rates in detectors based on the inverse β^\pm -decay processes, i.e. on the $\nu(\bar{\nu})$ -capture reactions, which take place due to the electroweak charge currents,

$$\nu_e + A_Z(I^\pi) \rightarrow e^- + A_{Z+1}(I'\pi'), \quad (1)$$

$$\bar{\nu}_e + A_Z(I^\pi) \rightarrow e^+ + A_{Z-1}(I'\pi'). \quad (2)$$

A number of nuclei were suggested for neutrino detectors. The most interesting nuclei for theoretical analysis are the odd- A nuclei with lowest reaction thresholds: ${}^{71}\text{Ga}$ ($Q_\beta = -236$ keV), ${}^{81}\text{Br}$ (-471 keV), ${}^{115}\text{In}$ (-120 keV) which are sensitive both to the low- and to high-energy portions of the solar-neutrino spectrum.



Motivation

- *The first direct detection of the neutrinos from carbon-nitrogen-oxygen (CNO) fusion cycle in the Sun by the BOREXINO Collaboration [1].*
- *An estimate of “CNO-like” events from geo-antineutrino of ^{40}K decay in the Hydride model of the Earth [2].*
- *Additional experimental prospects given by improved ^{115}In detector system (LENS Project [3]).*

1. *M. Agostini M., K. Altenmuller, S. Appel et al. for BOREXINO Collaboration , arXiv:2006.15115. 2020.*
2. *L.B. Bezrukov et.al., Izvestia RAN **85**, (4), 566 (2021).*
3. *R.S. Raghavan, The LENS Experiment: Spectroscopy of Low Energy Solar Neutrinos, Neutrino 2010, Athens, Greece, 2010*

Theory background

- Ground states of odd-A and odd-odd nuclei are described with the **latest version** of the Fayans density functional.
- The main feature is **self-consistent CQRPA for odd-A and odd-odd nuclei** within full $SO(8)$ framework including the **isoscalar ($T=0$) effective pp-interaction**.
- The latter enables more precise description of the low-energy tail of the GT and FF strength functions in the daughter nuclei. That is important for estimating low energy **pp and CNO** neutrino rates in ^{82}Se , ^{100}Mo as well as ^{115}In .

Specific problems in odd-A and odd-odd nuclei:

- *blocking of pairing*
- *odd quasiparticle in the g.s. with correct J^π*

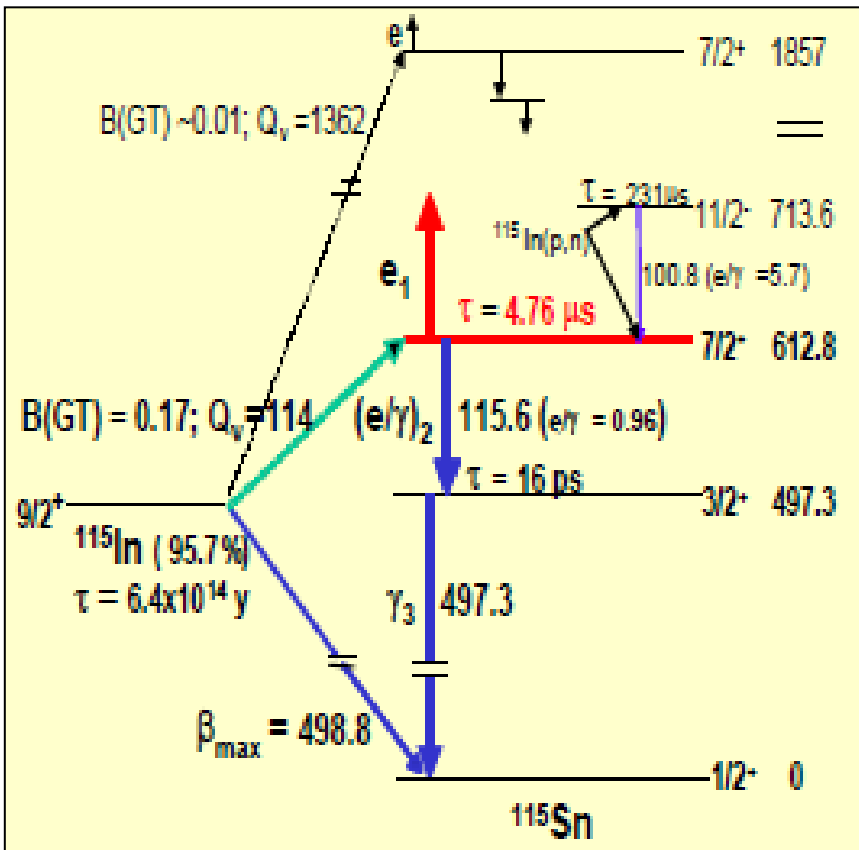
***Unpleasant feature - violation of time-reversal invariance
by odd nucleon !***

The usual way out: Equal Filling Approximation (EFA)

$^{115}\text{In} \rightarrow ^{115}\text{Sn}$ LENS Project

No-threshold, retarded
 ν -induced decay
 $9/2^+ \rightarrow 1/2^+ \quad 6.4 (14)\text{y}$

Low-threshold *inverse GT β -decay*
 $9/2^+ \rightarrow 7/2^+$
 $T_{1/2} = 3.26 \mu\text{sec}$
 $Q_U = 114 \text{ keV}$,,
 $Ex = 0.6128 \text{ MeV}$
 $B(GT) = 0.17$



A unique delayed-coincidence signature

A low-threshold, high-efficiency, direct-counting detector for solar neutrinos from the p-p fusion reaction, pep as well as CNO neutrinos.

R. S. Raghavan
 Phys. Rev. Lett. 37, 259



Three main families of phenomenological EDF

DF3... -a, -b, -f, ...FANDF⁰

S.A. Fayans + collaborators, KI, Moscow

BCPM - Barcelona–Catania–Paris–Madrid
(*originating from an early work by Baldo et al.)*

SeaLL - Seattle–Livermore .

- *directly parametrize the nuclear EoS by series of powers of the density ;*
- *add corrective terms to account for finite-size and many-body effects ;*
- *add terms accounting for the Coulomb potential and pairing corrections.*

*Fayans and SeaLL functionals : the Kohn-Sham type EDF
independent-particle kinetic energy, $m^*=1$*

A detailed explanation of Skyrme and Fayans EDF as used here is given in Ref. [1]. We summarize here the essential features for the present paper. Both functionals are composed from volume, surface, spin-orbit, pairing, Coulomb and center-of-mass terms as

$$\mathcal{E} = \mathcal{E}^v(\rho, \tau) + \mathcal{E}^s(\rho) + \mathcal{E}^{\text{ls}}(\rho, \vec{J}) + \mathcal{E}^{\text{Coul}}(\rho) + \mathcal{E}^{\text{pair}}(\rho) + \mathcal{E}^{\text{c.m.}}(\rho) \quad .$$

These read in detail

	Skyrme	Fayans
volume:	$\mathcal{E}_{\text{Sk}}^v = \sum_{t=0}^1 [(C_{t0}^{\rho\rho} + C_{tD}^{\rho\rho}\rho_0^\alpha)\rho_t^2 + C_t^{\rho\tau}\rho_t\tau_t]$ $C_{t0}^{\rho\rho}, C_{tD}^{\rho\rho}, \alpha, C_t^{\rho\tau} \leftrightarrow E/A_{\text{eq}}, \rho_{\text{eq}}, K, J, L, \frac{m^*}{m}, \kappa_{\text{TRK}}$	$\mathcal{E}_{\text{Fy}}^v = \frac{1}{3}\varepsilon_F\rho_{\text{sat}} \left[a_+^v \frac{1-h_{1+}^v x_0^\sigma}{1+h_{2+}^v x_0^\sigma} x_0^2 + a_-^v \frac{1-h_{1-}^v x_0}{1+h_{2-}^v x_0} x_1^2 \right]$ $a_\pm^v, h_{1\pm}^v, h_{2\pm}^v \leftrightarrow E/A_{\text{eq}}, \rho_{\text{eq}}, K, J, L, h_{2-}^v$
surface :	$\mathcal{E}_{\text{Sk}}^s = \sum_{t=0}^1 C_t^{\rho\Delta\rho} \rho_t \Delta\rho_t$	$\mathcal{E}_{\text{Fy}}^s = \frac{1}{3}\varepsilon_F\rho_{\text{sat}} \frac{a_+^s r_s^2 (\vec{\nabla}x_0)^2}{1+h_+^s x_0^\sigma + h_\nabla^s r_s^2 (\vec{\nabla}x_0)^2}$
spin-orbit:	$\mathcal{E}_{\text{Sk}}^{\text{ls}} = \sum_{t=0}^1 C_t^{\rho\nabla J} \rho_t \nabla \cdot J_t$	$\mathcal{E}_{\text{Fy}}^{\text{ls}} = \sum_{t=0}^1 C_t^{\rho\nabla J} \rho_t \nabla \cdot J_t$
pairing:	$\mathcal{E}_{\text{Sk}}^{\text{pair}} = \frac{1}{4} \sum_{q \in \{p, n\}} V_{\text{pair}, q} \left(1 - \frac{\rho_0}{\rho_{\text{pair}}} \right) \check{\rho}_q^2$	$\mathcal{E}_{\text{Fy}}^{\text{pair}} = \frac{2\varepsilon_F}{3\rho_{\text{sat}}} \check{\rho}_q^2 \left[f_{\text{ex}}^\xi + h_+^\xi x_{\text{pair}}^\gamma + h_\nabla^\xi r_s^2 (\vec{\nabla}x_{\text{pair}})^2 \right]$

where $x_t = \rho_t/\rho_{\text{sat}}$ and $x_{\text{pair}} = \check{\rho}_q/\rho_{\text{sat}}$. The γ , $\rho_{\text{sat}} = 0.16 \text{ fm}^{-3}$ and $\varepsilon_F = \varepsilon_F(\rho_{\text{sat}})$ are given, fixed values. The non-linear surface coefficient is fixed as $h_+^s = h_{2+}^s$. Coulomb term and c.m. correction are irrelevant here. Note that the parameters for the volume terms are handled in term of nuclear matter parameters E/A_{eq} etc as is indicted in the line below the volume terms.

The groups of model parameters for the MCC in Fig. 3 are defined as follows:

	Skyrme	Fayans
pairing:	$V_{\text{pair}, p}, V_{\text{pair}, n}, \rho_{\text{pair}}$	$f_{\text{ex}}^\xi, h_+^\xi, h_\nabla^\xi$
surface:	$C_0^{\rho\Delta\rho}, C_1^{\rho\Delta\rho}$	$a_+^s, h_\nabla^s, h_\nabla^\xi$
spin-orbit:	$C_0^{\rho\nabla J}, C_1^{\rho\nabla J}$	$C_0^{\rho\nabla J}, C_1^{\rho\nabla J}$
symmetry energy:	J, L	J, L

I. Self-consistent approach to nuclear ground state .

Thus the total interaction energy of the superfluid nucleus, $E_{\text{int}}[\rho, \nu] = \int d\mathbf{r} \varepsilon_{\text{int}}(\mathbf{r})$, where $\varepsilon_{\text{int}}(\mathbf{r})$ is defined as above, is a functional of two densities, the normal, $\rho(\mathbf{r})$, and the anomalous, $\nu(\mathbf{r})$. Self-consistent calculation with such a functional looks like the standard variational procedure in which the single-particle hamiltonian takes the form

$$\mathcal{H} = \begin{pmatrix} h - \mu & -\Delta \\ -\Delta & \mu - h \end{pmatrix}, \quad (15)$$

where

$$h = \frac{p^2}{2m} + \frac{\delta E[\rho, \nu]}{\delta \rho}, \quad \Delta = -\frac{\delta E[\rho, \nu]}{\delta \nu}. \quad (16)$$

These equations have been solved iteratively as follows. For given densities $(\rho^{(i)}, \nu^{(i)})$, from the above functional, Eq. (4), the elements of the hamiltonian $\mathcal{H}^{(i)}$ were derived, through its eigenvalues and wave functions $(u^{(i)}, v^{(i)})$ the new densities were calculated, and then, as an input $(\rho^{(i+1)}, \nu^{(i+1)})$ for the next iteration, superpositions of previous densities and these new ones were used with the weights of 0.85 and 0.15, respectively.

IAS.

Fully self-consistent calculation.

$$E_0 = \int \mathcal{E}[\rho(\mathbf{r}), \nu(\mathbf{r})] d^3r,$$

$$h(1, 2) = \frac{\delta E[\rho, \nu]}{\delta \rho(2, 1)}$$

$$1 \equiv (\mathbf{r}_1, s_1, t_1).$$

$$\Delta(\mathbf{r}) = \frac{\delta E[\rho, \nu]}{\delta \nu(\mathbf{r})}$$

$$\mathcal{F}^- = \frac{\delta^2 \mathcal{E}}{(\delta \rho_-)^2},$$

$$\rho_- = \rho_n - \rho_p.$$

GT and FF.

Self-consistent g.s.calculation

GT and FF.

Effective NN-interaction.

U: **Particle-hole channel:**
 δ -interaction
with Landau-Migdal constant
 $g' > 0$ (repulsion)
+ π -meson + ρ -meson exchange

V: **Particle-particle channel:**
 $T=0$, δ -interaction
with one parameter:
 $g'_{pp} < 0$ (attraction)

Neglecting the spin-isospin
dependent components in E_0 causes
 ~ 100 KeV error in masses

II. Continuum CQRPA approach to nuclear excited states

$$\begin{aligned}
 \Delta=0 \quad \rho_{pn} &= L_{pn}(\omega) V_{pn} + M_{pn}(\omega) V_{pn}^h + N_{pn}^1(\omega) d_{pn}^1 + N_{pn}^2(\omega) d_{pn}^2, \\
 \rho_{pn}^h &= M_{pn}(\omega) V_{pn} + L_{pn}(-\omega) V_{pn}^h + N_{pn}^2(-\omega) d_{pn}^1 + N_{pn}^1(-\omega) d_{pn}^2, \\
 \varphi_{pn}^1 &= N_{pn}^1(\omega) V_{pn} + N_{pn}^2(-\omega) V_{pn}^h + K_{pn}(\omega) d_{pn}^1 - M_{pn}(\omega) d_{pn}^2, \\
 \varphi_{pn}^2 &= N_{pn}^2(\omega) V_{pn} + N_{pn}^1(-\omega) V_{pn}^h - M_{pn}(\omega) d_{pn}^1 + K_{pn}(-\omega) d_{pn}^2.
 \end{aligned} \tag{26}$$

Here $\omega = \tilde{\omega} - \delta\mu$, where $\delta\mu = \mu^p - \mu^n$ is the difference between the proton and the neutron chemical potentials, and $\tilde{\omega}$ is the nuclear excitation energy. Propagators obtained by integrating various products of the normal and abnormal Green functions G and F

Even-even core:

$$\begin{aligned}
 L_{pn}(\omega) &= \int G_p(\varepsilon) G_n(\varepsilon + \tilde{\omega}) \frac{d\varepsilon}{2\pi i} \\
 &= - \left[\frac{v_p^2 u_n^2}{E_{pn} - \omega} + \frac{v_n^2 u_p^2}{E_{pn} + \omega} \right],
 \end{aligned}$$

Odd nucleon:

$$\begin{aligned}
 L_{p_0n}(\omega) &= \int G_{p_0}(\varepsilon) G_n(\varepsilon + \tilde{\omega}) \frac{d\varepsilon}{2\pi i} \\
 &= - \left[\frac{v_{p_0}^2 u_n^2}{(E_n + E_{p_0}) - \omega} + \frac{u_{p_0}^2 u_n^2}{(E_n - E_{p_0}) - \omega} \right],
 \end{aligned}$$

Continuum+pairing:

$$\begin{aligned}
 L(\mathbf{r}, \mathbf{r}'; \omega) &= A(\mathbf{r}, \mathbf{r}'; \omega) + \sum [L_{pn}(\omega) - \tilde{A}_{pn}(\omega)] \\
 &\quad \times \varphi_n^*(\mathbf{r}_1) \varphi_p(\mathbf{r}_1) \varphi_n(\mathbf{r}_2) \varphi_p^*(\mathbf{r}_2),
 \end{aligned}$$

*Low-lying tail of GT-excitation spectrum
is defined by competition of*

$g' > 0$ (shifts Ex upward)

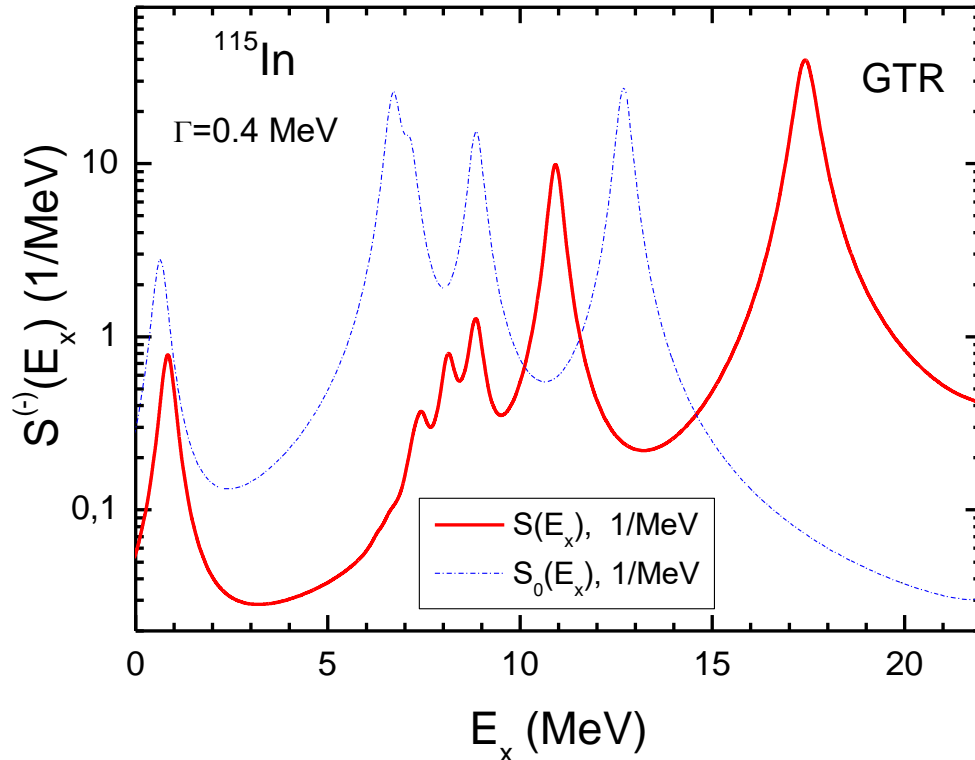


and $g'_{pp} < 0$ (shifts Ex downward)



^{115}In GT strength functions

The structure of the low-lying GT tail is simple



The capture cross section is sensitive to position of the 1st $1+$ state. Its $B(\text{GT})$ is amplified by the Fermi function $F(Z,w)$!

DF3a

The role played by pairing correlations and F^{pp} !

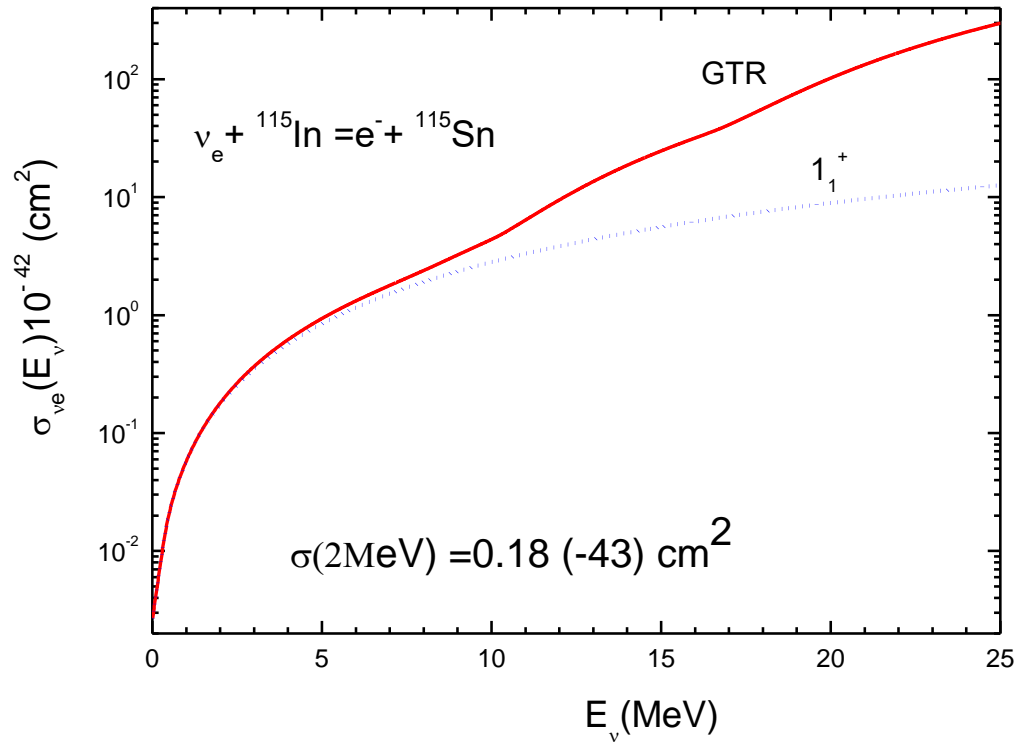
The strength functions in the $d = 0, F^{pp} = 0$ approximation:

For the first GT excited state corresponding to the transition $1g_{9/2} - 1g_{7/2}$ in $^{115}\text{In}-^{115}\text{Sn}$:

- $E_x(\text{GT}_1) = 0.825$ MeV
- $B(\text{GT})$ is higher than the exp. **by 2.5 times.**

$d \neq 0, F^{pp} \neq 0$ $E_x(\text{GT}_1) = 0.628$ MeV
 $B(\text{GT}) = 0.12$

Neutrino capture cross sections, 71Ga , 115In



*Solar-neutrino fluxes and capture rates (in SNU)
for indium detector are not very sensitive to the DF*

Neutrino source	E_{ν}^{\max} [Φ_i]	^{115}In	
		this work	Ref. [15]
1. $^1\text{H}(p, e^+ \nu_e)^2\text{H}$	0.420 [6.00]	460	468
2. $^1\text{H}(p e^-, \nu_e)^2\text{H}$	1.442 (d) [0.014]	7.7	8.1
3. $^3\text{He}(p, e^+ \nu_e)^4\text{He}$	18.773 [7.6×10^{-7}]	0.05	0.05
4. $^7\text{Be}(e^-, \nu_e)^7\text{Li}$	0.862 (d) 0.384 (d) [0.47]	119.8	116
5. $^8\text{B}(e^+ \nu_e)^8\text{Be}^+$	15.00 [5.8×10^{-4}] (2.87×10^{-4})	14.9	14.4
6. $^{13}\text{N}(e^+ \nu_e)^{13}\text{C}$	1.199 [0.061]	(7.5) 13.4	(7.2) 13.6
7. $^{15}\text{O}(e^+ \nu_e)^{15}\text{N}$	1.732 [0.052]	18.0	18.5
		634	639 total

DF3a result
(with no fitting of the
s.p. energies!) is
quite close to DF1

Conclusions

1. *SELF-CONSISTENT calculations based on the new Fayans energy density functional (DF3a) and CQRPA for odd-A nuclei with pairing.*
2. *The total neutrino capture rates are not very sensitive to the EDF chosen. The rates from DF1 are rather close to the ones from DF3a.*

Perspectives and challenges.

1. *The odd-A nuclei . Time reversal symmetry violation.
Framework is needed beyond the Equal filling approximation (EFA).
Schemes enriching the low-lying charge-exchange excitation spectra:

The odd-A nuclei beyond the QRPA (quasiparticle-phonon formalism).
A practical framework –FRSA approach extended for odd-A and odd-odd nuclei:*

1. Complex configuration effects on beta-decay rates *A. P. Severyukhin, V. V. Voronov, I. N. Borzov, N. N. Arsenyev, Nguyen Van Giai*, Journal of Physics: conference series, 580, 012051-1-6, 2015.
2. Proton-neutron quasiparticle RPA with separable Gamow-Teller forces [K. Muto](#), [E. Bender](#), [T. Oda](#) & [H. V. Klapdor-Kleingrothaus](#) *Zeitschrift für Physik A Hadrons and Nuclei* volume 341, pages407–415 (1992)

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V.V. Voronov, BLTP, JINR*

Self-Consistent Ground State. Fayans EDF.

$$\mathcal{E}[\rho(\mathbf{r}), \nu(\mathbf{r})] = \tau + \varepsilon_v + \varepsilon_s + \varepsilon_{\text{Coul}} + \varepsilon_{sl} + \varepsilon_{ss} + \varepsilon_{\text{pair}} .$$

$$E_0^{\text{int}}[\rho] = \int \mathcal{E}(\rho(\mathbf{r})) d^3r = \int \frac{a\rho^2}{2} (1 + \alpha\rho^\sigma) d^3r,$$

Skyrme EDF

vs.

$$\mathcal{E}(\rho) = \frac{a\rho^2}{2} \frac{1 + \alpha\rho^\sigma}{1 + \gamma\rho} .$$

$m=m^*$

new ρ – dependent terms

Fayans EDF

DF3a + Deformed HFBTHO :

*S.V. Tolokonnikov, I.N. Borzov, M. Kortelainen, Yu.S. Lutostansky, E.E.Saperstein
J. Phys.G 35, 291, 2014.*

Neutrino capture cross sections, 71Ga, 115In

I.N. Borzov et al. / Nuclear Physics A 584 (1995) 335–361

351

$\Delta \# 0$

$\sigma^{115}\text{In} (2 \text{ MeV}) \sim 0.4 \cdot 10^{-43} \text{ cm}^2$

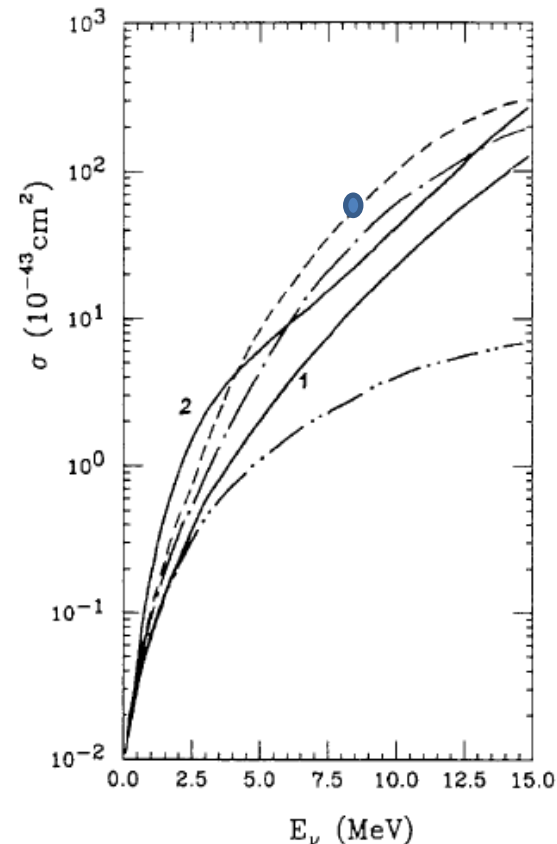
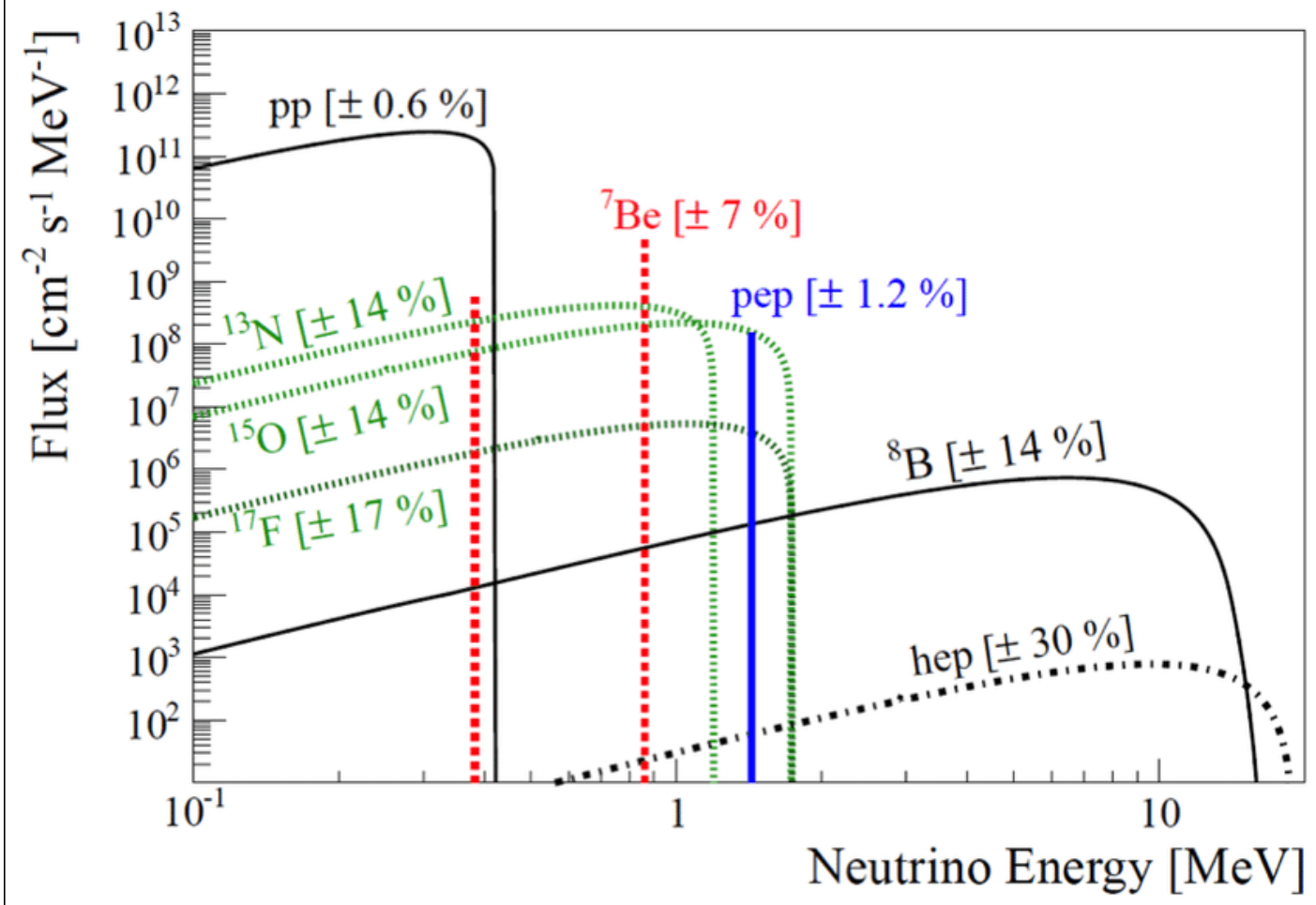


Fig. 2. Neutrino capture cross sections for ^{71}Ga and ^{115}In as functions of E_ν . ^{71}Ga : for the g.s. \rightarrow g.s. (dash-double-dotted line); including all GT transitions up to the neutron separation threshold B_n in ^{71}Ge (curve 1); calculated using the strength function from Ref. [11] (dash-dotted line); taken from Ref. [21] (curve 2). ^{115}In : including all GT transitions up to $E_x = 20 \text{ MeV}$ (dashed line).



Continuum pnQRPA. Full ph-basis, SO(8) symmetry

$$[I - \begin{pmatrix} F^\omega & -F^\xi & F^{\omega\xi} & F^{\xi\omega} \\ -F^\xi & F^\omega & F^{\xi\omega} & F^{\omega\xi} \\ F^{\omega\xi} & F^{\xi\omega} & F^\xi & -F^\omega \\ F^{\xi\omega} & F^{\omega\xi} & -F^\omega & F^\xi \end{pmatrix} \begin{pmatrix} L(\omega) & M(\omega) & N^1(\omega) & N^2(\omega) \\ M(\omega) & L(-\omega) & N^2(-\omega) & N^1(-\omega) \\ N^1(\omega) & N^2(-\omega) & K(\omega) & -M(\omega) \\ N^2(\omega) & N^1(-\omega) & -M(\omega) & K(-\omega) \end{pmatrix}] \begin{pmatrix} V \\ V^h \\ d^{(1)} \\ d^{(2)} \end{pmatrix} = \begin{pmatrix} e_q V_0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

$$L(r, r'; \omega) = A(r, r'; \omega) + \sum_{pn} |L_{pn} - A_{pn}^{\sim}| \varphi_p \varphi_n \varphi_{n^*} \varphi_{p^*}$$



SO(8)

Self-consistent pnQRPA

SU(4)

Standard beta decay Tables

P. Moeller et al., Phys. Rev. (1996, 2003, 2012)