

Search for Chiral Magnetic Wave (CMW) with ALICE at the LHC



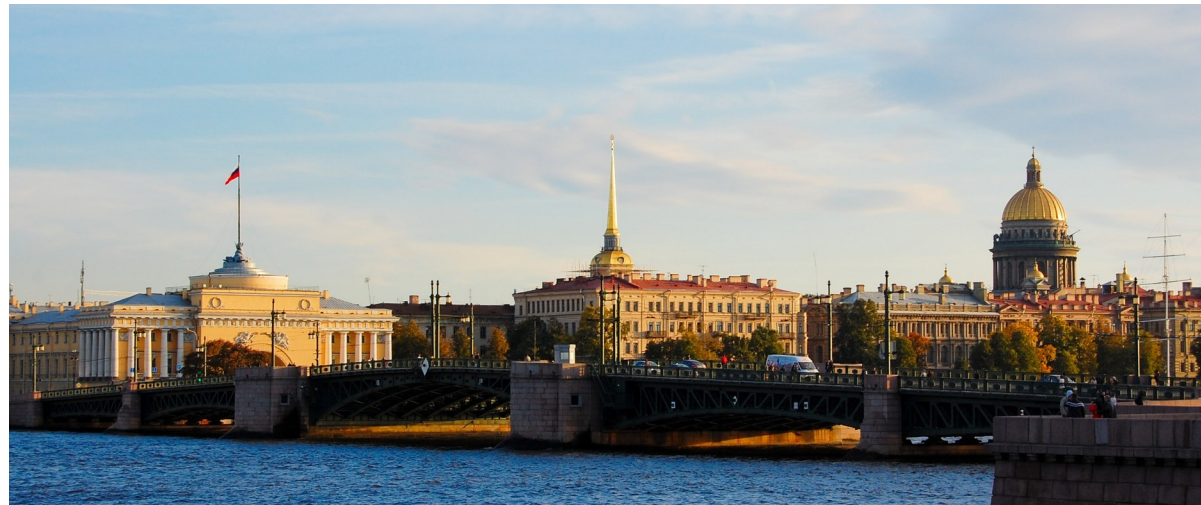
Prattay Das (for the ALICE Collaboration)
National Institute of Science Education and Research
HBNI



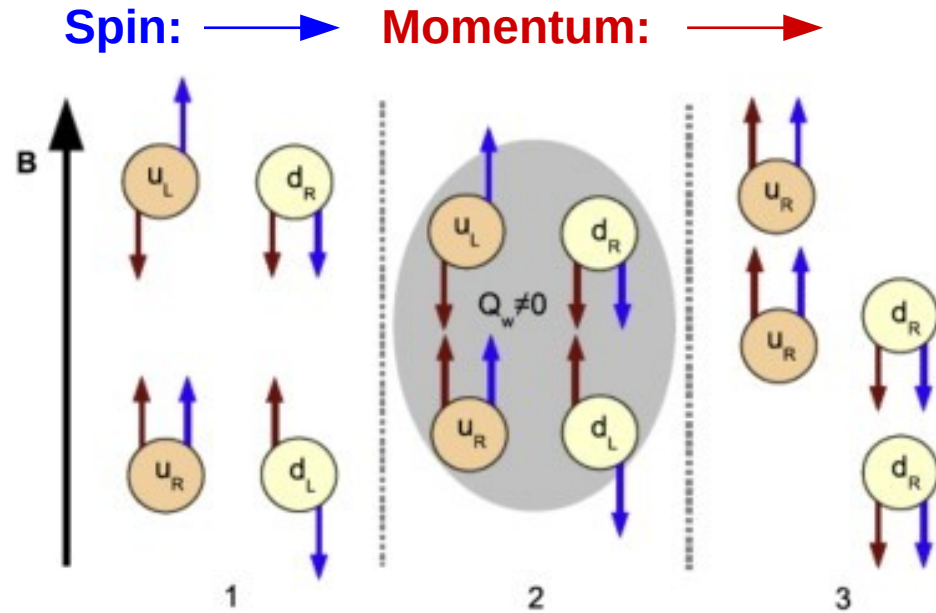
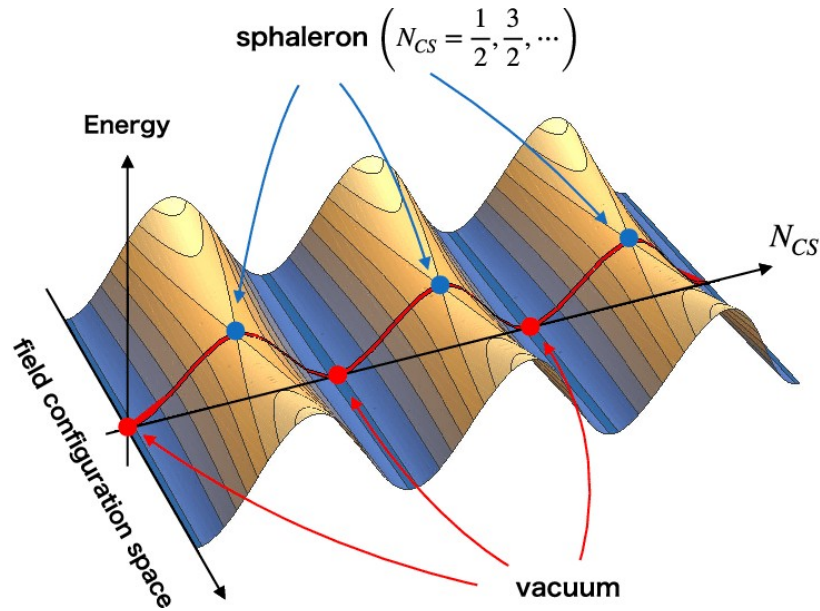
Outline:

- ❖ Motivation
- ❖ Experimental observable
- ❖ ALICE detectors
- ❖ Analysed data, event and track cuts
- ❖ Results
- ❖ Summary and outlook

LXXI International Conference NUCLEUS – 2021



Motivation



☞ QCD vacuum: degenerate

☞ Generates chirality imbalance:

$$N_L^f - N_R^f = 2Q_W$$

☞ Axial and vector currents

☞ Induces **parity odd domains**

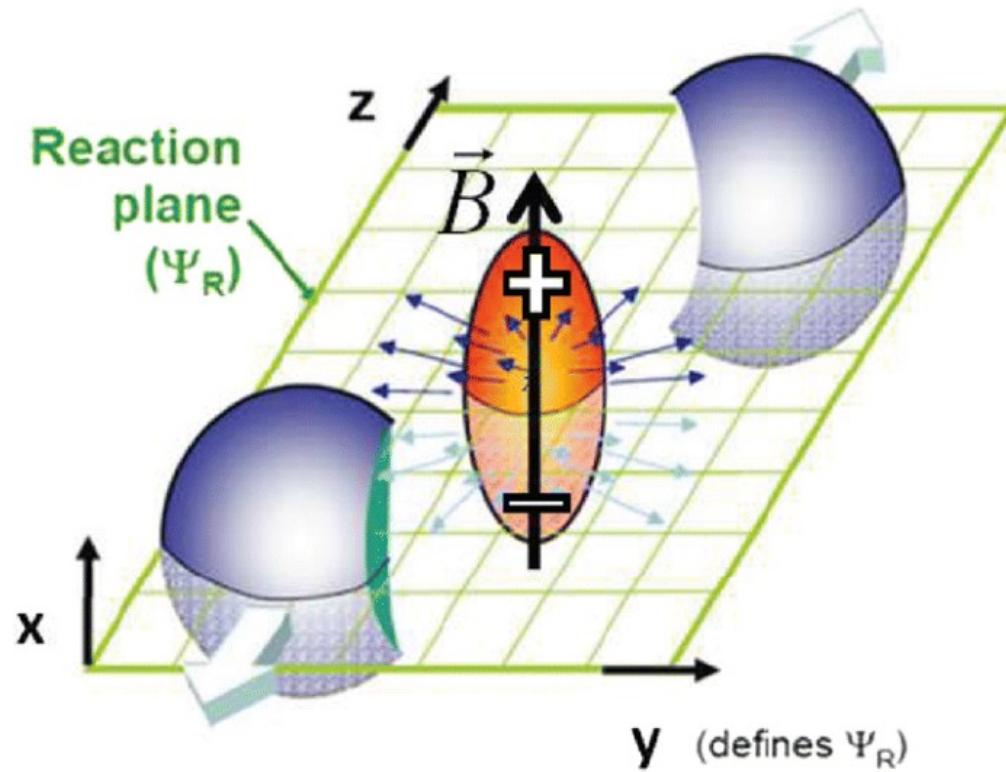
Chiral Magnetic Effect (CME): $j_v = \frac{N_c e}{2\pi^2} \mu_A B$

Chiral Separation Effect (CSE): $j_A = \frac{N_c e}{2\pi^2} \mu_v B$

Chiral Magnetic Wave: CME + CSE

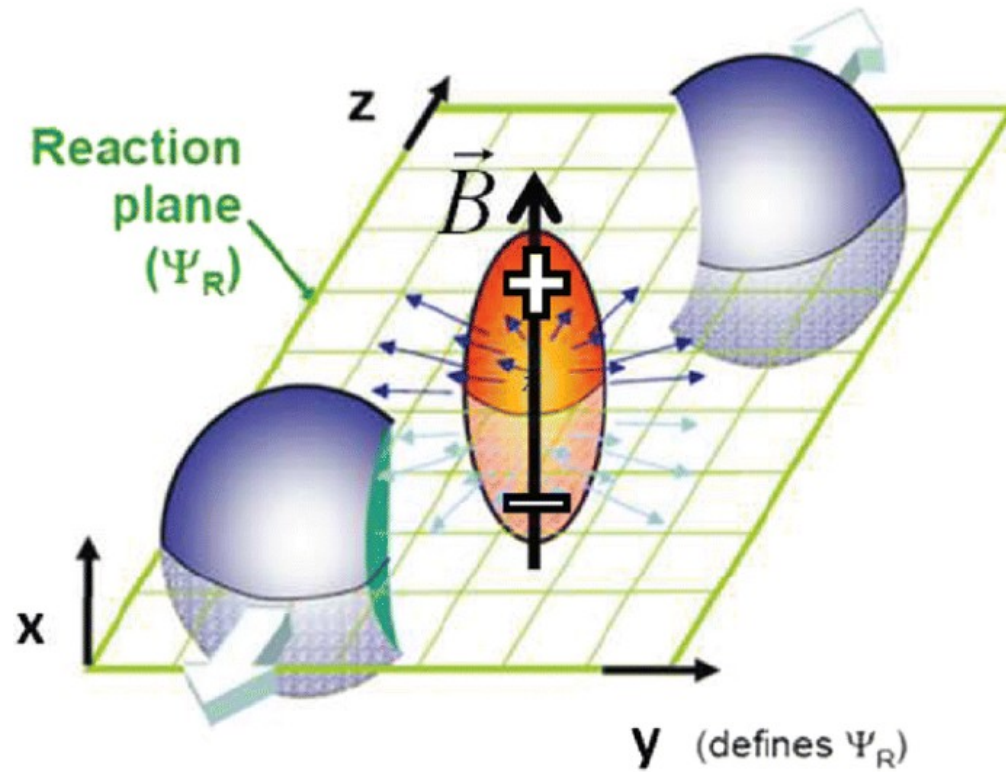
[1] Phys.Rev.Lett. 81 (1998) 512-515
[2] Phys.Rev.D. 101 (2020) 096014

☞ Chiral symmetry restoration

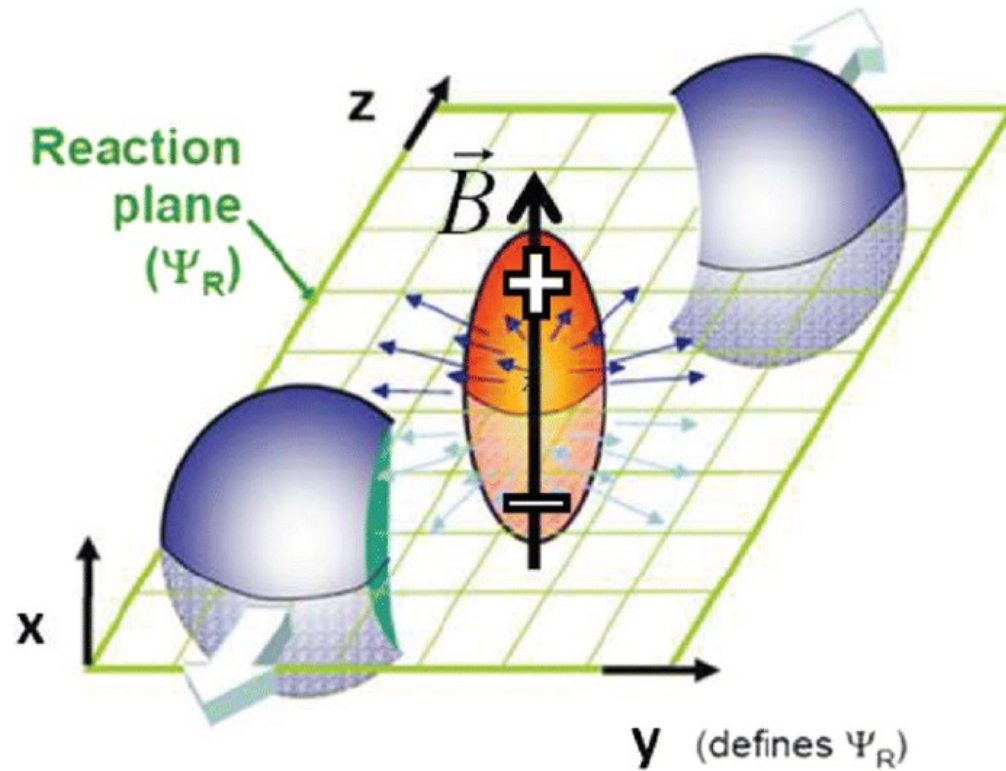


Heavy-ion collisions

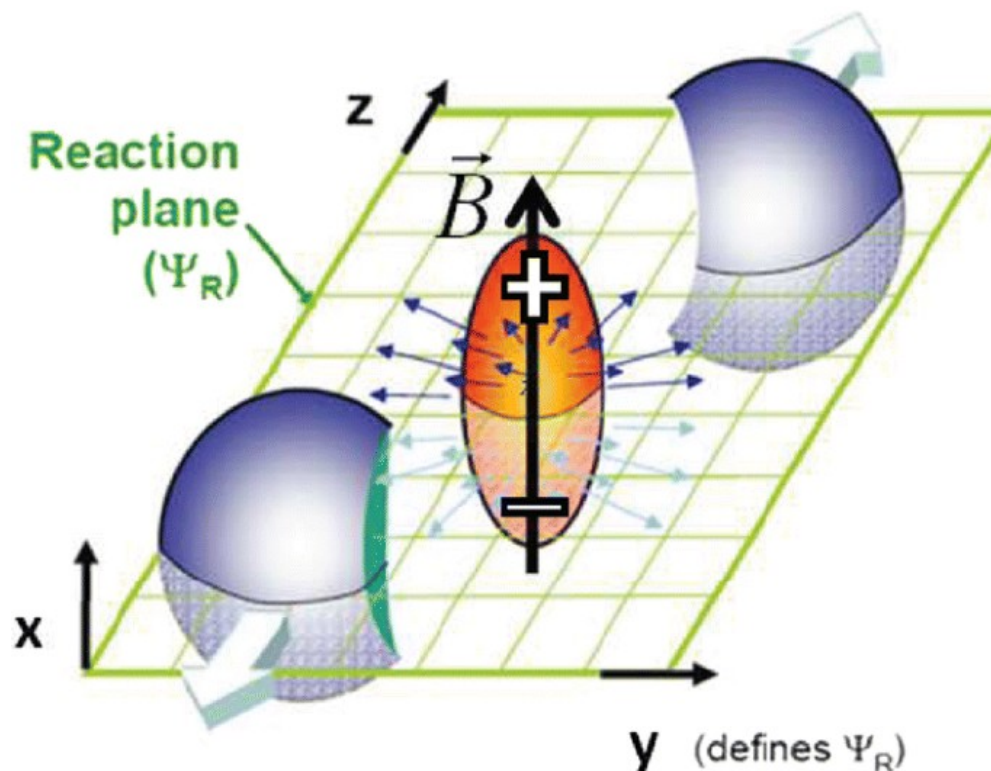
- ☞ Chiral symmetry restoration
- ☞ Deconfinement



- ☞ Chiral symmetry restoration
- ☞ Deconfinement
- ☞ QCD vacuum transitions

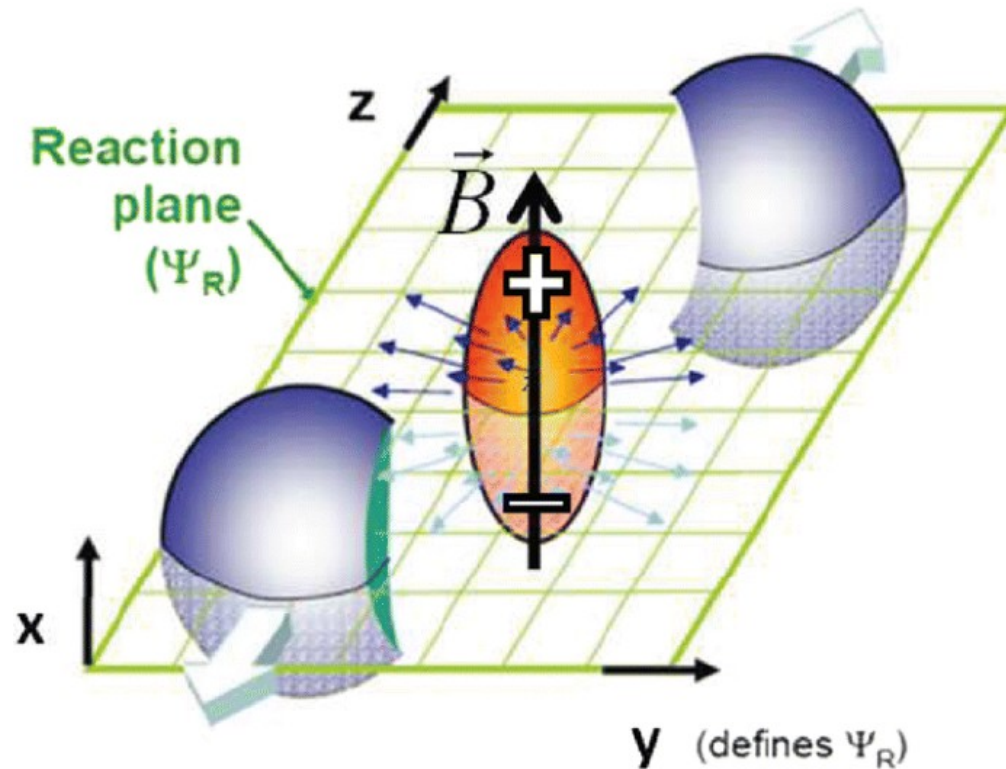


- ☞ Chiral symmetry restoration
- ☞ Deconfinement
- ☞ QCD vacuum transitions
- ☞ Extremely strong magnetic field ($\sim 10^{15}$ T)



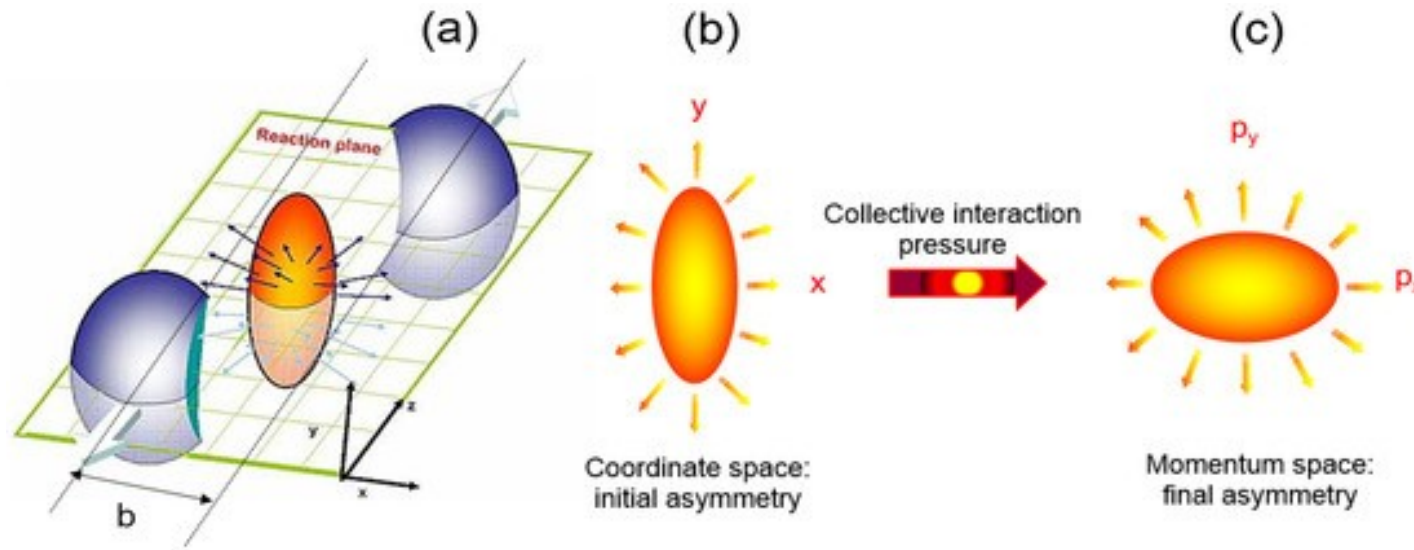
Heavy-ion collisions

- ☞ Chiral symmetry restoration
- ☞ Deconfinement
- ☞ QCD vacuum transitions
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All the necessary conditions are possible to be achieved in heavy-ion collisions

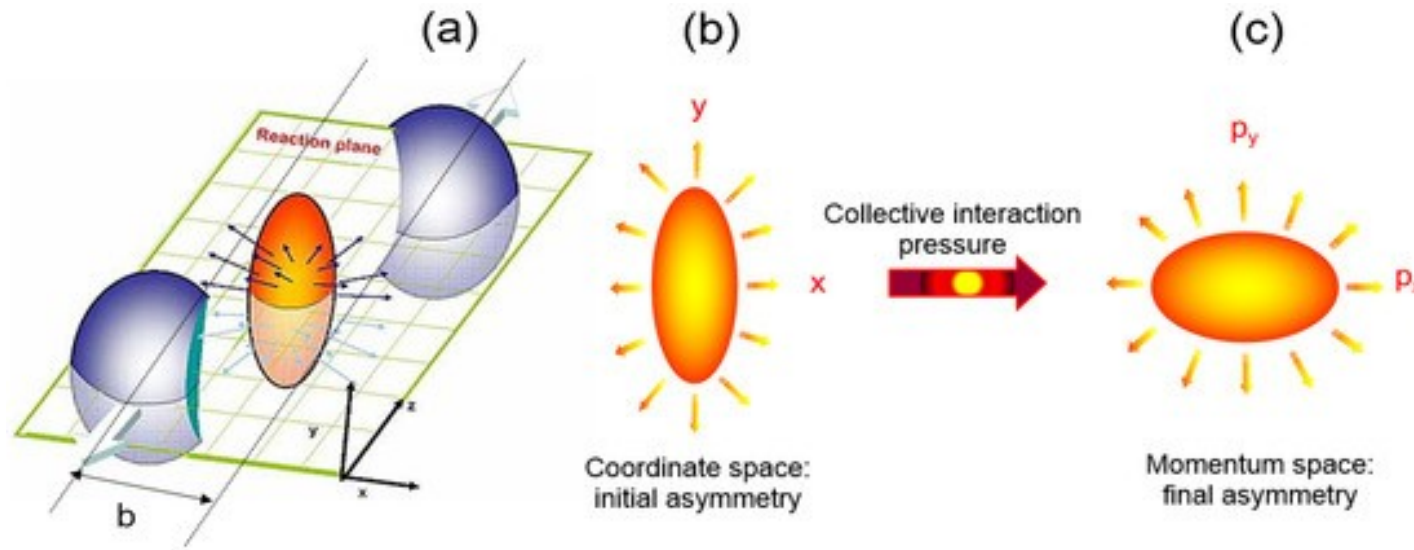
Anisotropic flow



👉 Spatial anisotropy → momentum anisotropy

Phys.Rev.D 48 (1993) 1132-1139

Anisotropic flow



👉 Spatial anisotropy \longrightarrow momentum anisotropy

👉 Characterised by:

$$E \frac{d^3 N}{d^3 p} = \frac{d^2 N}{2 \pi p_T dp_T dy} (1 + \sum 2 v_n \cos[n(\varphi - \Psi_{n,R})])$$

Fourier coefficients

Phys.Rev.D 48 (1993) 1132-1139

For illustration purpose

☞ Charge dependent elliptic flow

$$v_2^{h^\pm} = v_2 \mp r \frac{A_{ch}}{2} \quad A_{ch} = \frac{N^+ - N^-}{N^+ + N^-}$$

☞ **CMW observable:** $r_{\Delta v_2}^{Norm} = \frac{d\left(\frac{\Delta v_2}{\langle v_2 \rangle}\right)}{d A_{ch}}$

Normalised slope ,

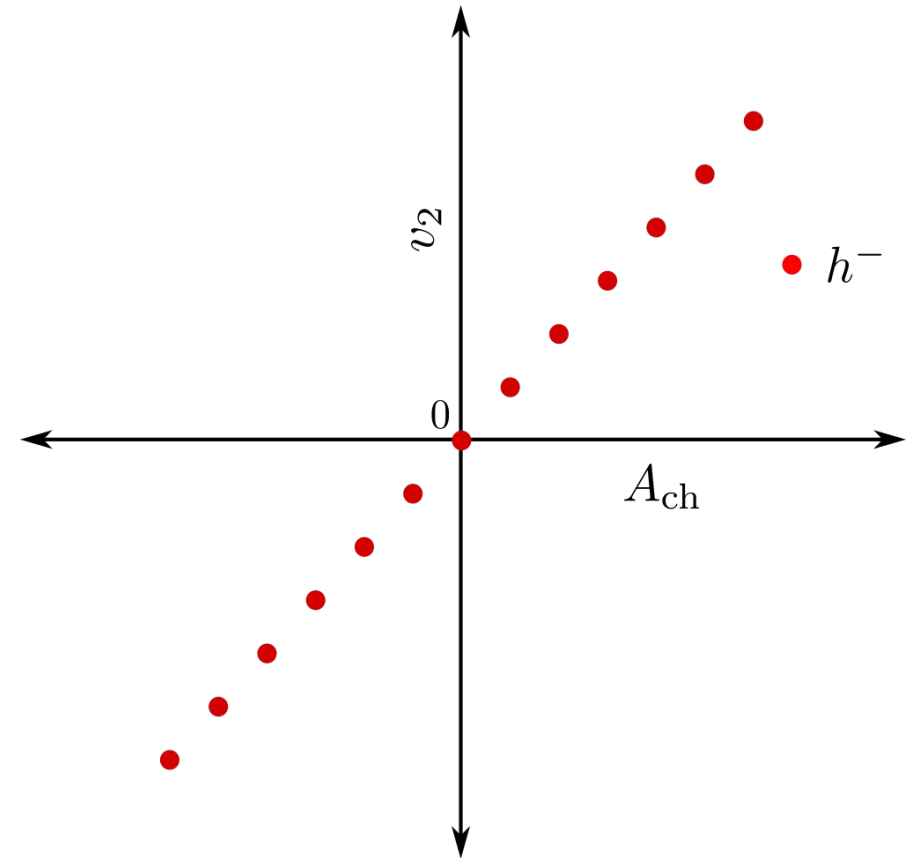
$$\Delta v_2 = v_2^{h^-} - v_2^{h^+} \quad \langle v_2 \rangle = \frac{v_2^{h^-} + v_2^{h^+}}{2}$$

☞ **Possible background:**
Local charge conservation (LCC)

- Probe the background:
Similar measurement with v_3

☞ **Signal:**

- $r_{\Delta v_2}^{Norm} > 0$
- $r_{\Delta v_2}^{Norm} > r_{\Delta v_3}^{Norm}$



[1] Phys.Rev.Lett. 107 (2011) 052303
 [2] Phys.Rev.C 100 (2019) 6, 064908
 [3] Phys. Rev. C 103 (2021) 034906

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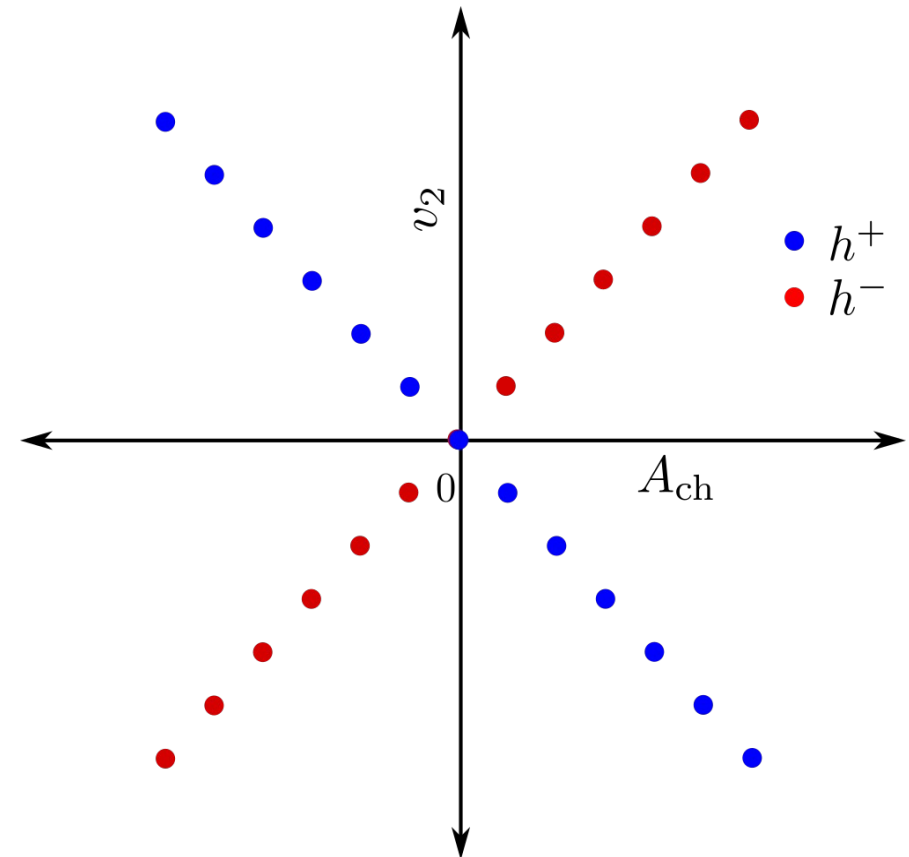
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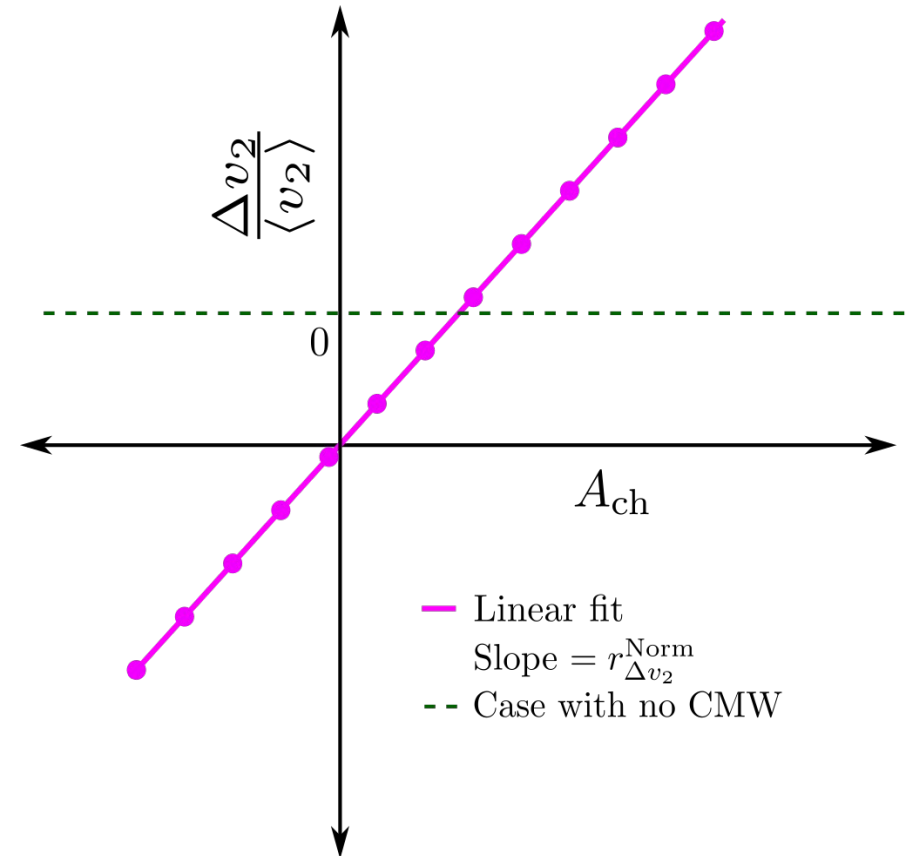
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Creator:cairo 1.14.6 (<http://cairographi>)
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CreationDate:Mon Mar 6 15:35:55 2017
LanguageLevel:3

V0

TPC

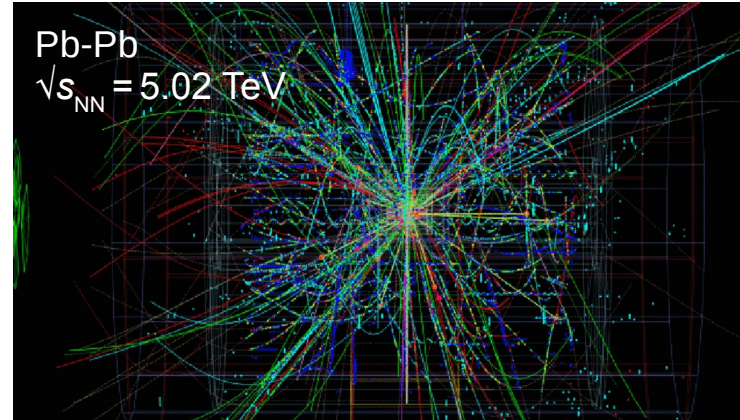
Time Projection Chamber (TPC): ($|\eta| < 0.9$)

- ☞ Primary vertex and tracking
- ☞ Momentum measurement
- ☞ PID through dE/dx

V0: V0A ($2.8 < \eta < 5.1$) &
V0C ($-3.7 < \eta < -1.7$)

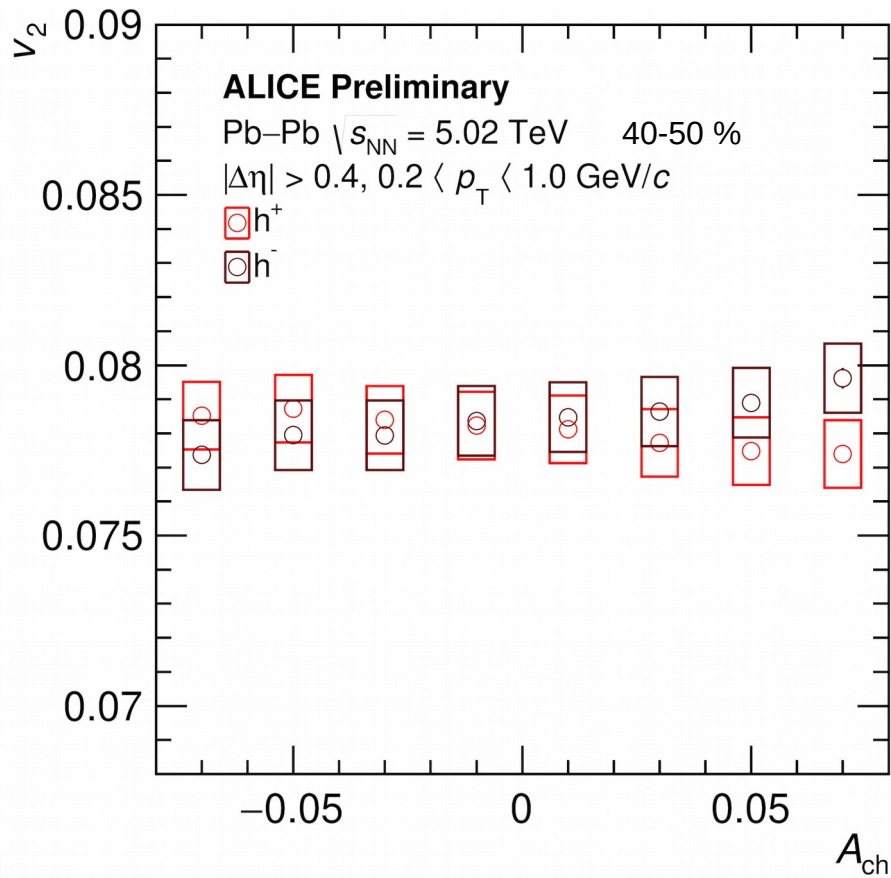
- ☞ Trigger, centrality

Analysis details

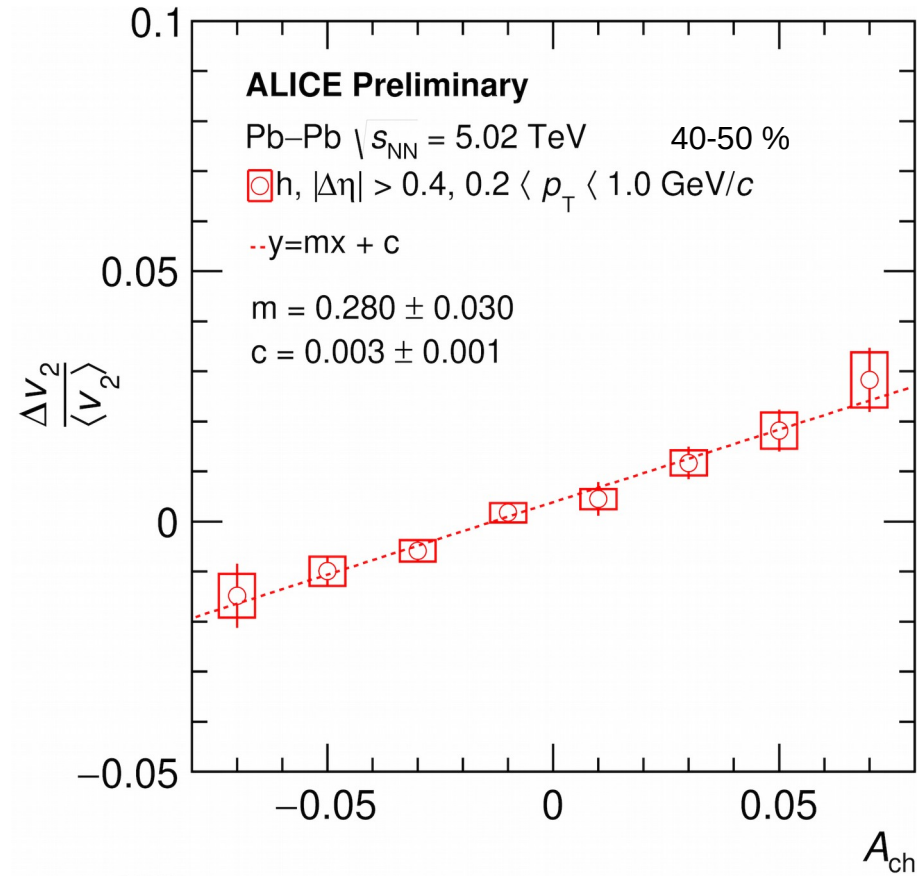


No. of events	$\sim 45 \times 10^6$
Kinematic range	$ \eta < 0.8$ $0.2 < p_T < 0.5$ GeV/c (pions) $0.2 < p_T < 1.0$ GeV/c (hadrons)
Non flow suppression	$ \Delta\eta > 0.4$ between subevents
Charge asymmetry (A_{ch})	$0.2 < p_T < 10$ GeV/c, $ \eta < 0.8$, 10 uniform bins (-0.1 to 0.1)
Centrality (%)	0 - 80

Elliptic flow (v_2) vs charge asymmetry (A_{ch})



ALI-PREL-365984



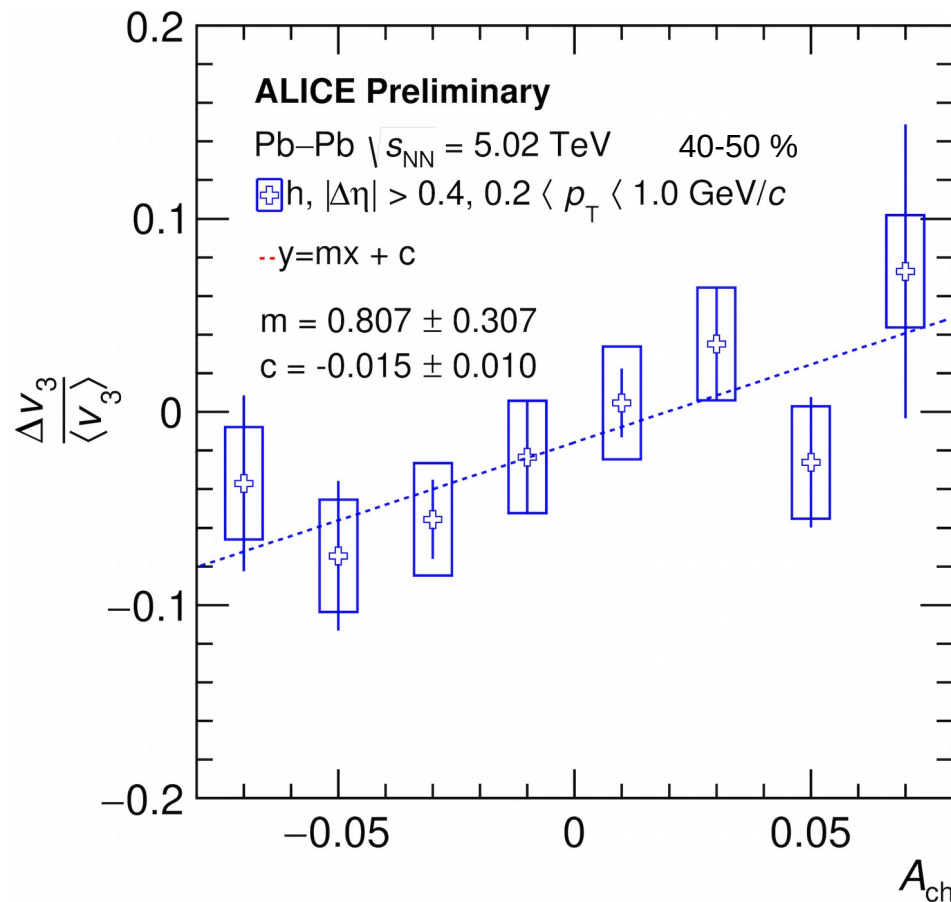
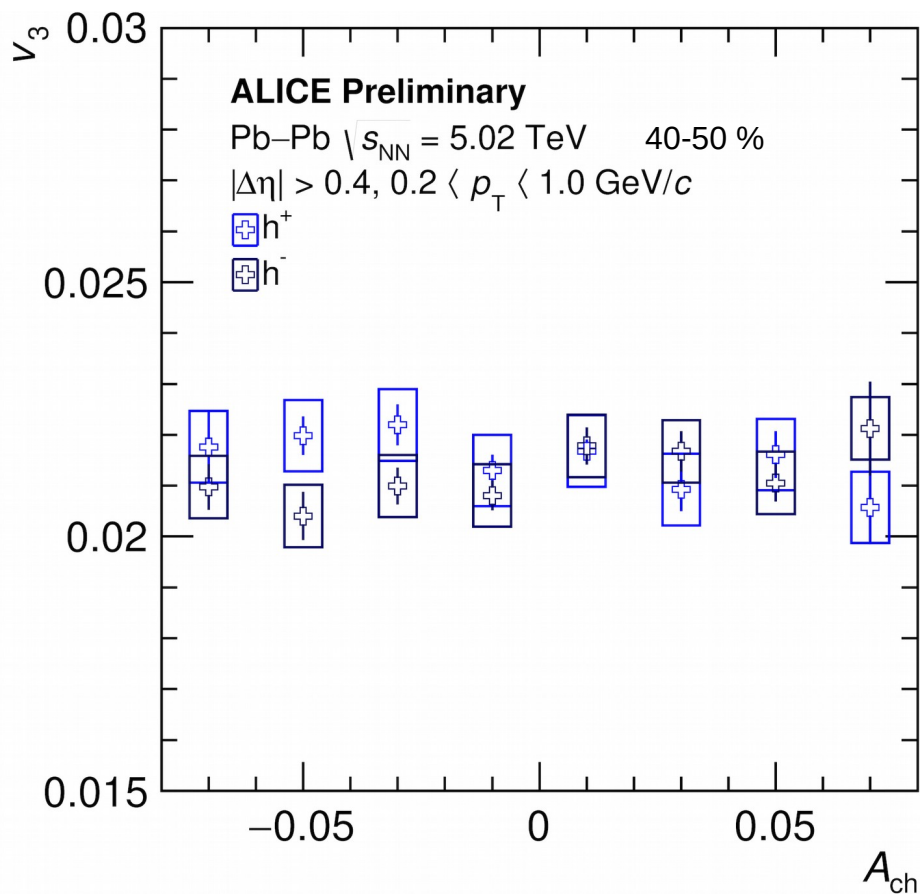
ALI-PREL-366004

Finite $r_{\Delta v_2}^{Norm}$ is observed

$$r_{\Delta v_n}^{Norm} = \frac{d\left(\frac{\Delta v_n}{\langle v_n \rangle}\right)}{d A_{ch}}$$

$$A_{ch} = \frac{N^+ - N^-}{N^+ + N^-}$$

Triangular flow (v_3) vs charge asymmetry (A_{ch})



ALI-PREL-365992

ALI-PREL-366008

Finite $r_{\Delta v_3}^{Norm}$ is observed

$$r_{\Delta v_n}^{Norm} = \frac{d\left(\frac{\Delta v_n}{\langle v_n \rangle}\right)}{d A_{ch}}$$

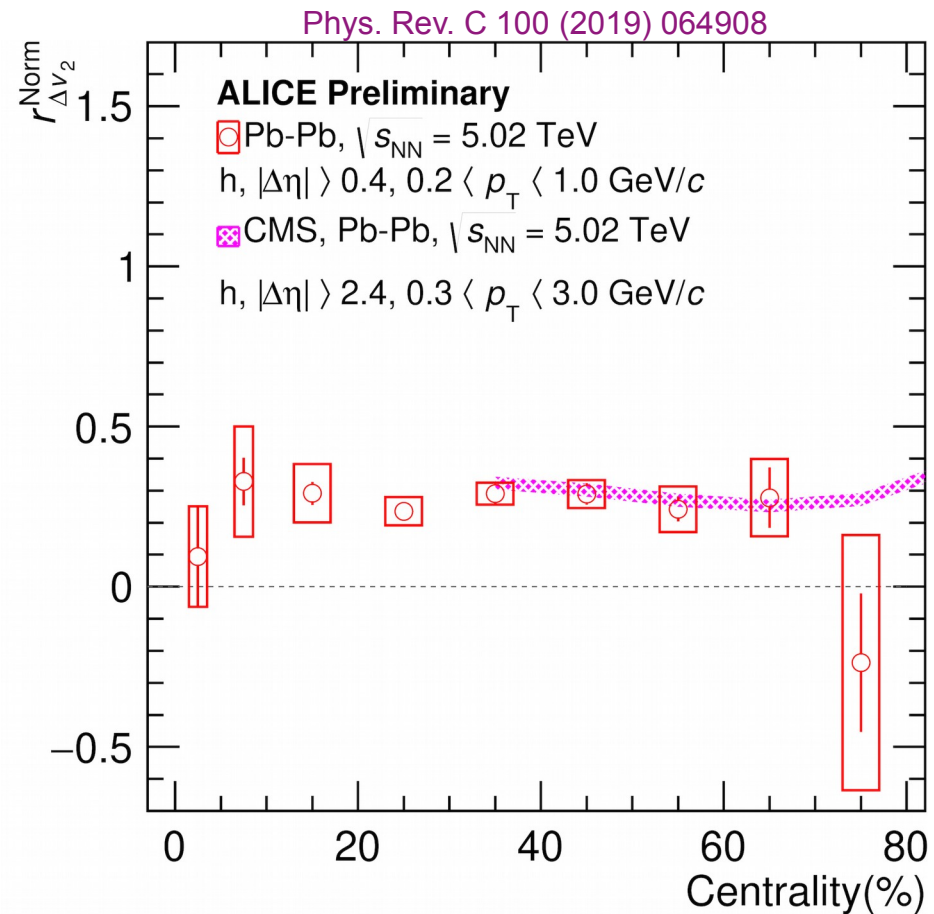
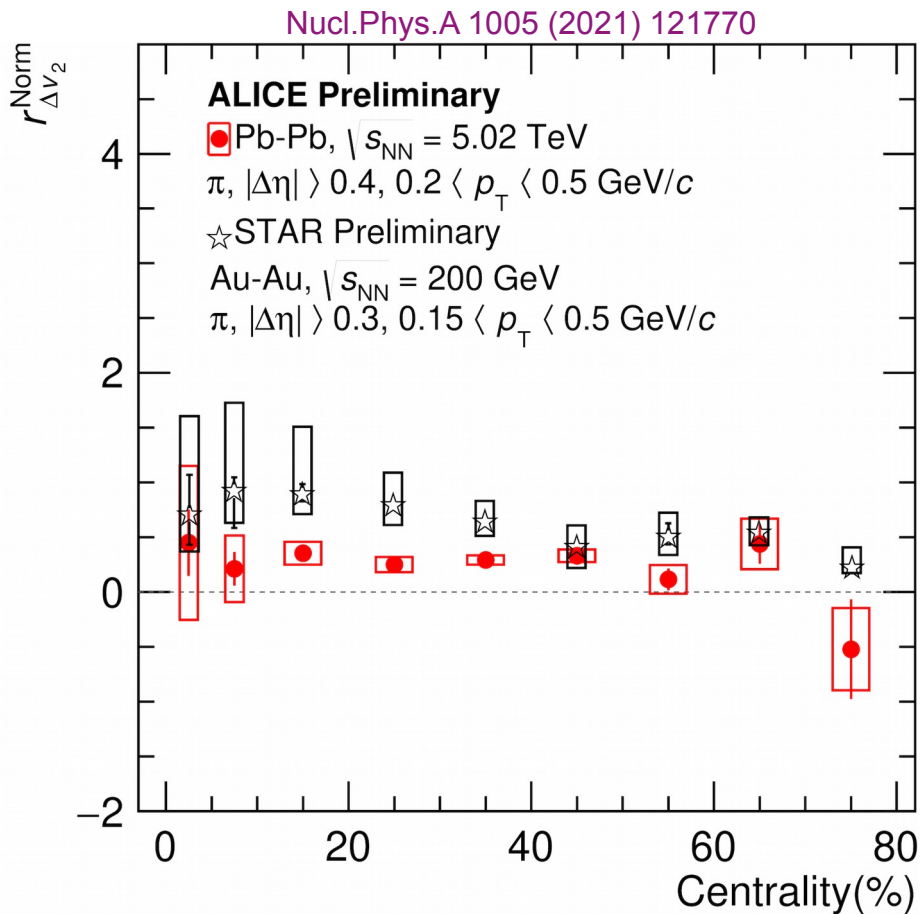
$$A_{ch} = \frac{N^+ - N^-}{N^+ + N^-}$$

Comparison of $r_{\Delta v_2}^{\text{Norm}}$ in ALICE to STAR and CMS



STAR

CMS



ALI-PREL-365968

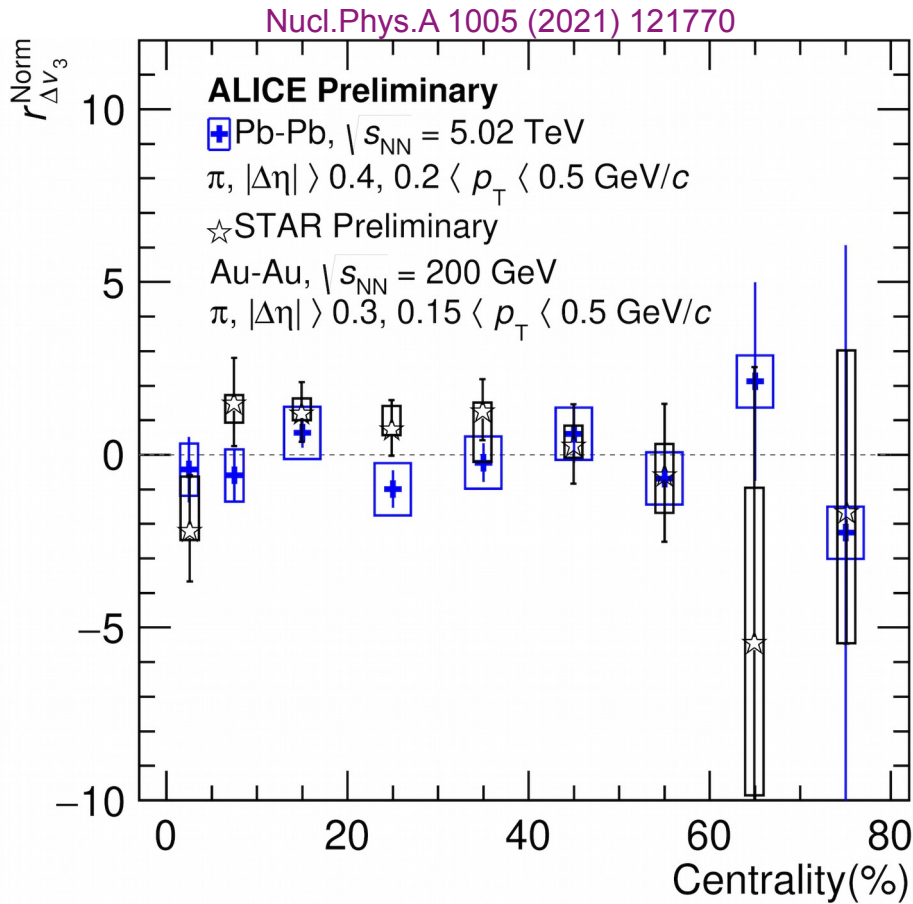
ALI-PREL-365976

$r_{\Delta v_2}^{\text{Norm}} h(\text{ALICE}) \approx r_{\Delta v_2}^{\text{Norm}} h(\text{CMS})$, but note different p_T ranges
 $r_{\Delta v_2}^{\text{Norm}} \pi(\text{ALICE}) < r_{\Delta v_2}^{\text{Norm}} \pi(\text{STAR})$

Comparison of $r_{\Delta v_3}^{\text{Norm}}$ in ALICE to STAR and CMS

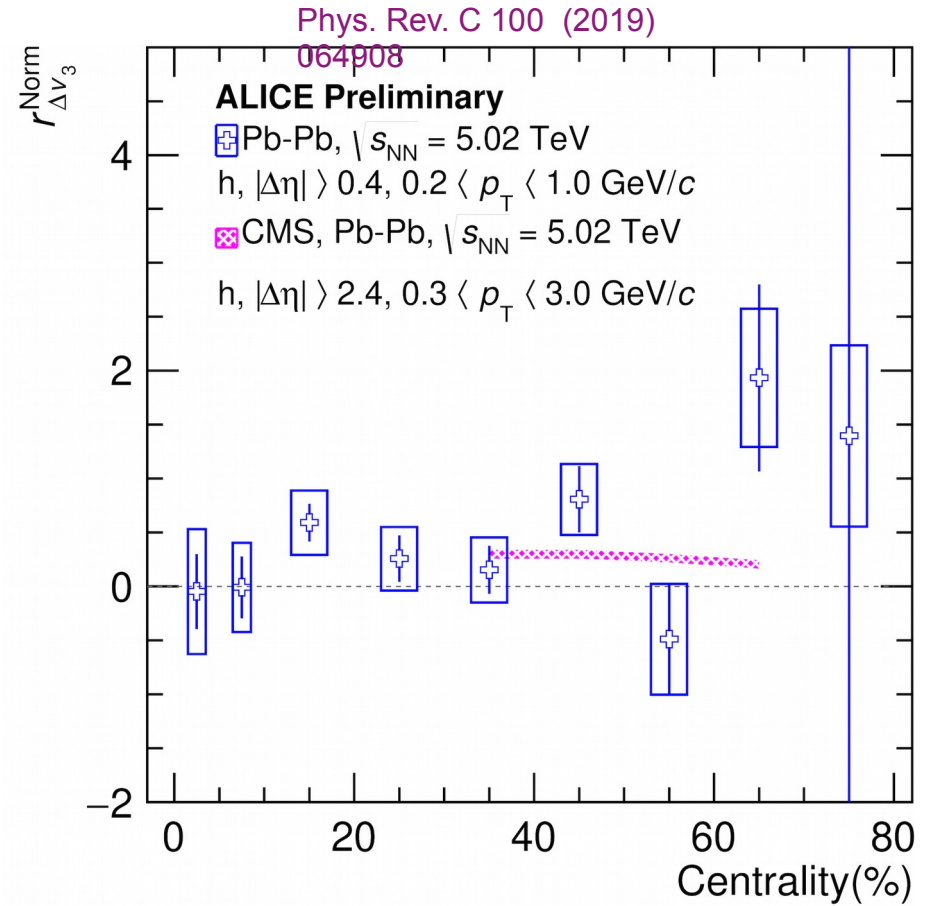


STAR



ALI-PREL-365972

CMS

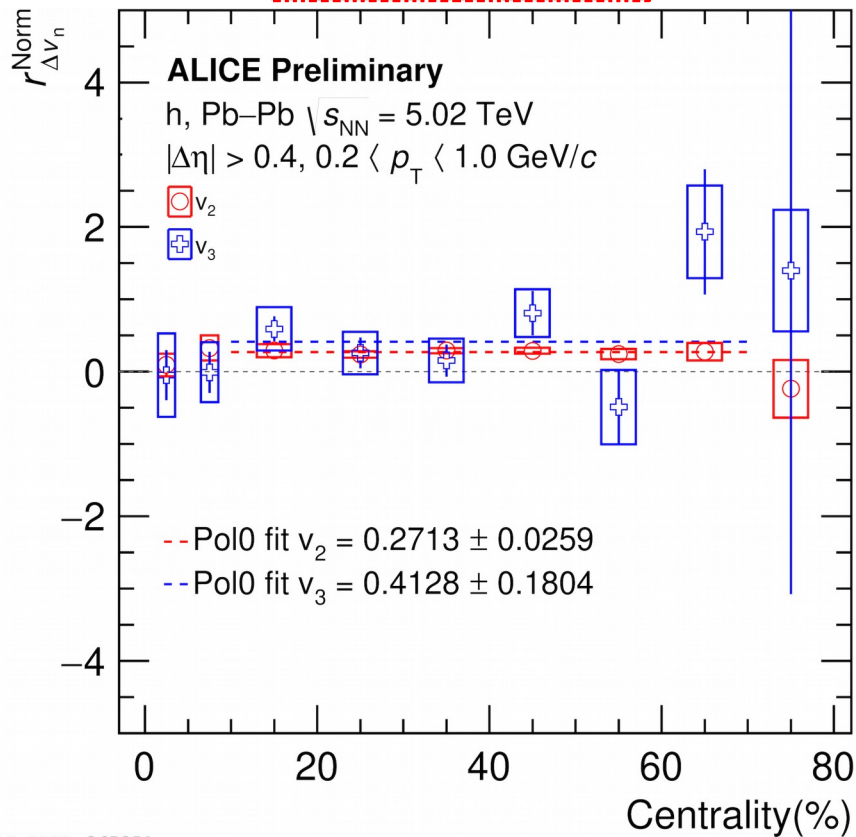


ALI-PREL-365980

👉 No observed discrepancies in $r_{\Delta v_3}^{\text{Norm}}$ between ALICE, STAR and CMS, but uncertainties are large

Centrality dependence of $r_{\Delta v_n}^{Norm}$

Hadrons



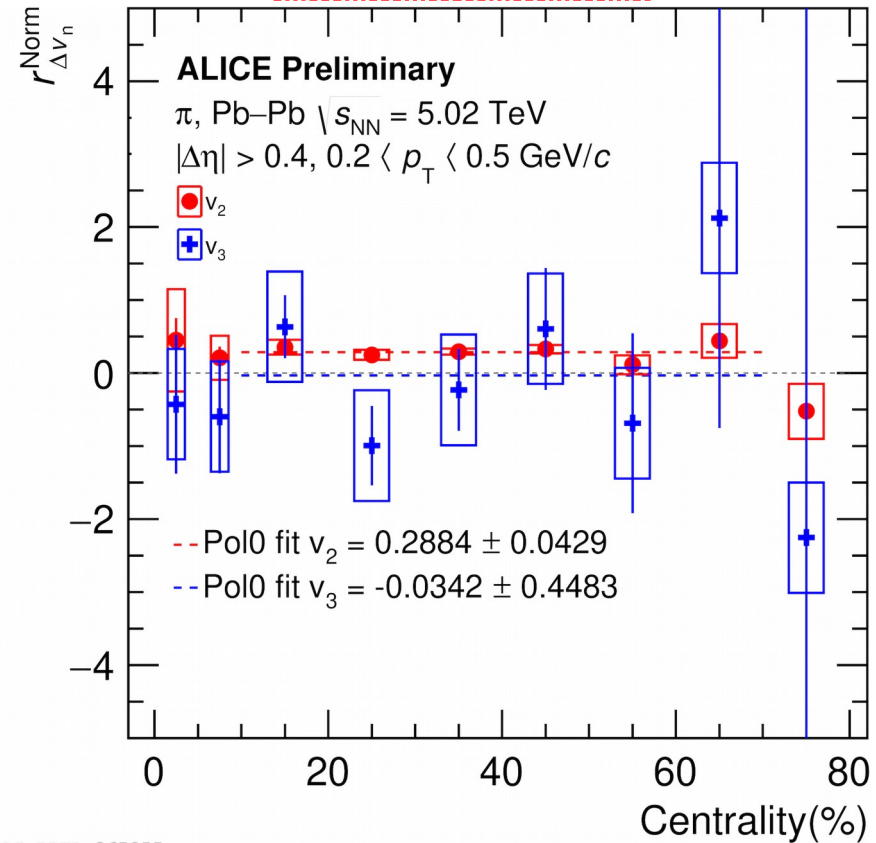
ALI-PREL-365951

☞ $r_{\Delta v_3}^{Norm}$ has large uncertainties

☞ $r_{\Delta v_2}^{Norm}$ is compatible with $r_{\Delta v_3}^{Norm}$ for both hadrons and pions

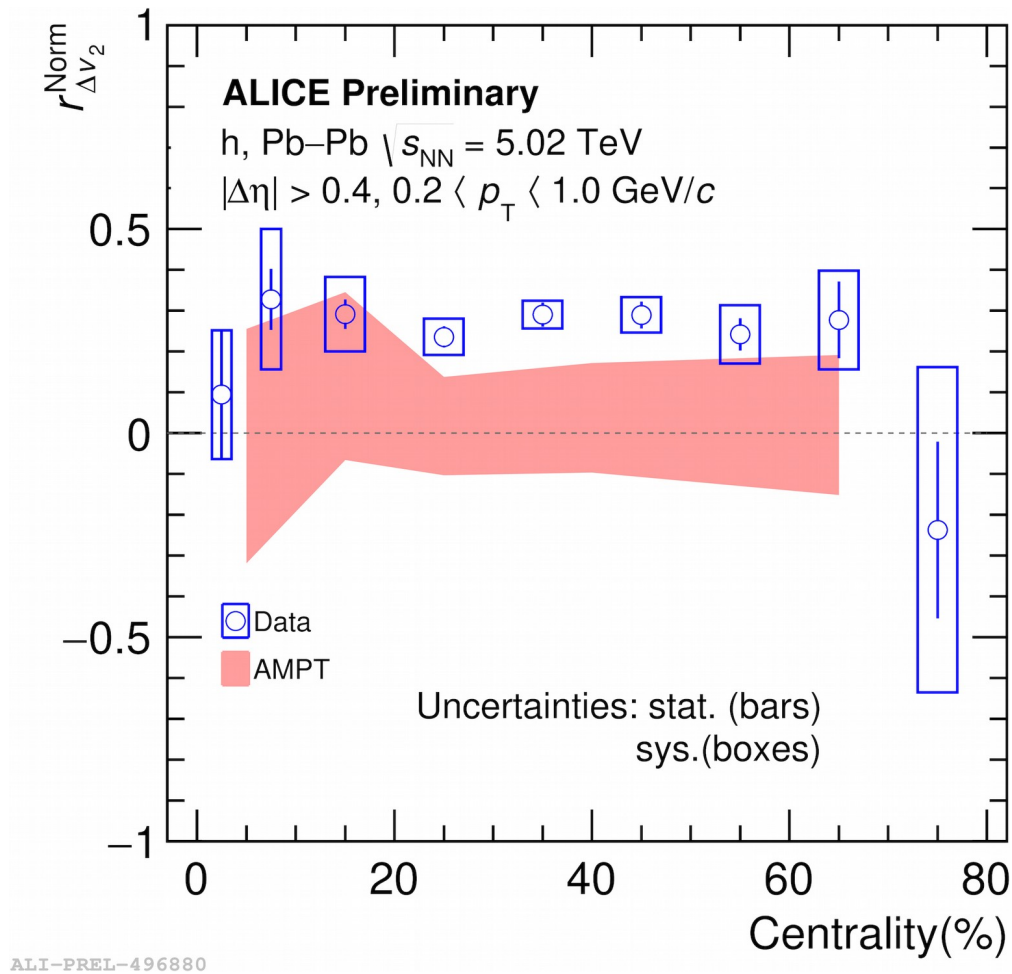
☞ Background dominates the signal

Pions



ALI-PREL-365955

Comparison of $r_{\Delta V_2}^{\text{Norm}}$ with AMPT model



☞ AMPT: No CMW and violation of charge conservation -> Slope consistent with zero

Summary



First measurement of normalised Δv_2 and Δv_3 slope of charged hadrons and pions in Pb-Pb collisions in ALICE.

$r_{\Delta v_2}^{Norm} > 0$ but compatible with $r_{\Delta v_3}^{Norm}$ indicates that background dominates the

signal

$r_{\Delta v_3}^{Norm}$ has large uncertainties

$r_{\Delta v_2}^{Norm}$ (ALICE) \approx $r_{\Delta v_2}^{Norm}$ (CMS)

$r_{\Delta v_3}^{Norm}$ (ALICE) $<$ $r_{\Delta v_3}^{Norm}$ (STAR)

$r_{\Delta v_2}^{Norm}$ (ALICE) \approx (CMS) \approx (STAR)

consistent with zero from AMPT model

Outlook

Analysis ongoing in high statistics data taken in 2018

THANK YOU