

# The XENONnT Experiment

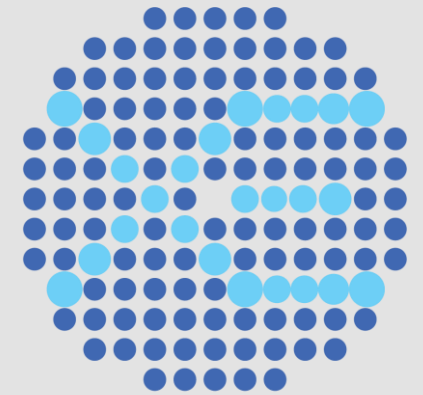
## Detector and science program

SPS/ÖPG annual meeting

Ricardo Peres, University of Zürich

[rperes@physik.uzh.ch](mailto:rperes@physik.uzh.ch)

01/09/2021



**XENON**



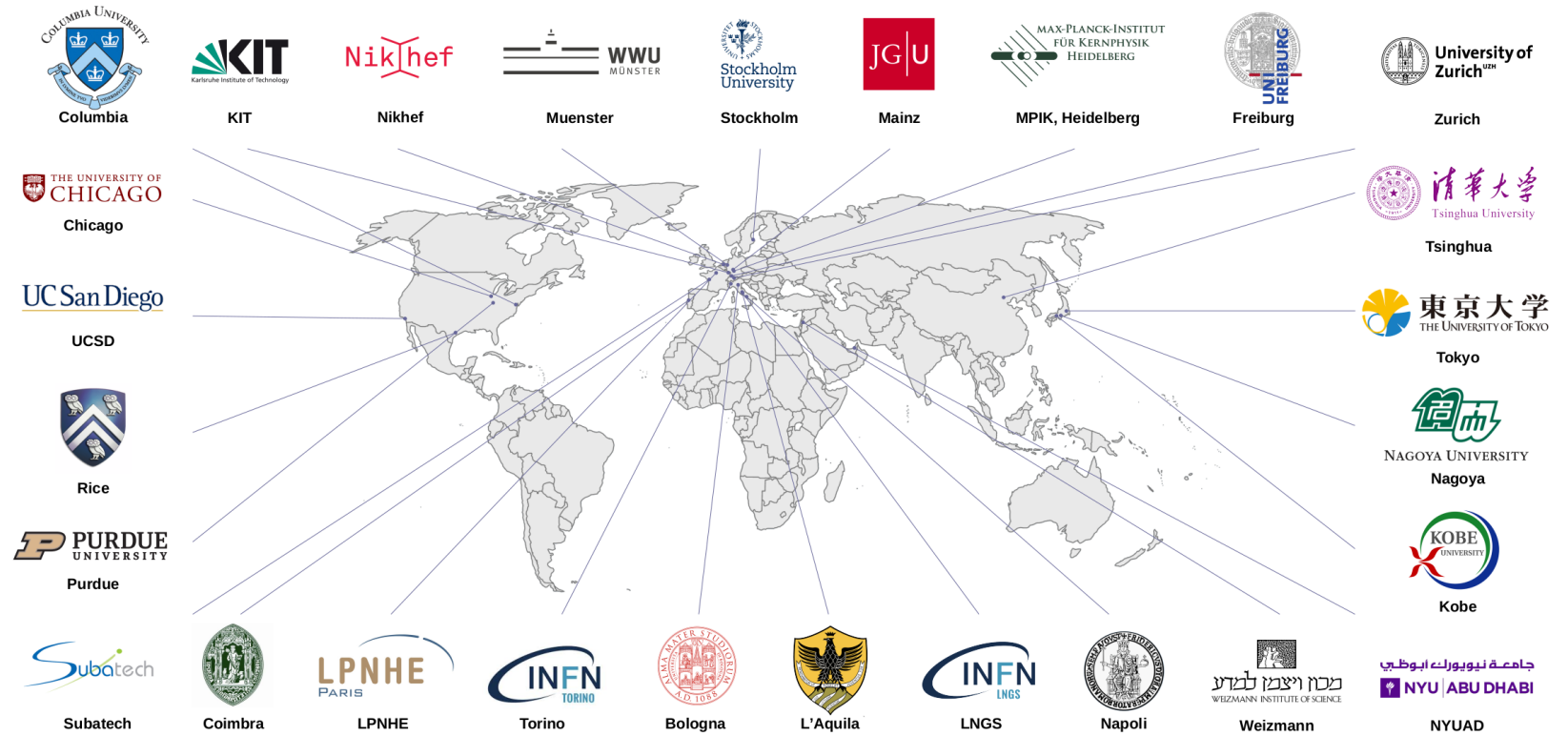
# XENON



University of Zurich <sup>UZH</sup>

# The XENON Collaboration

- 26 institutions worldwide
- ~170 scientists
- Main goal: look for dark matter particles with a xenon TPC





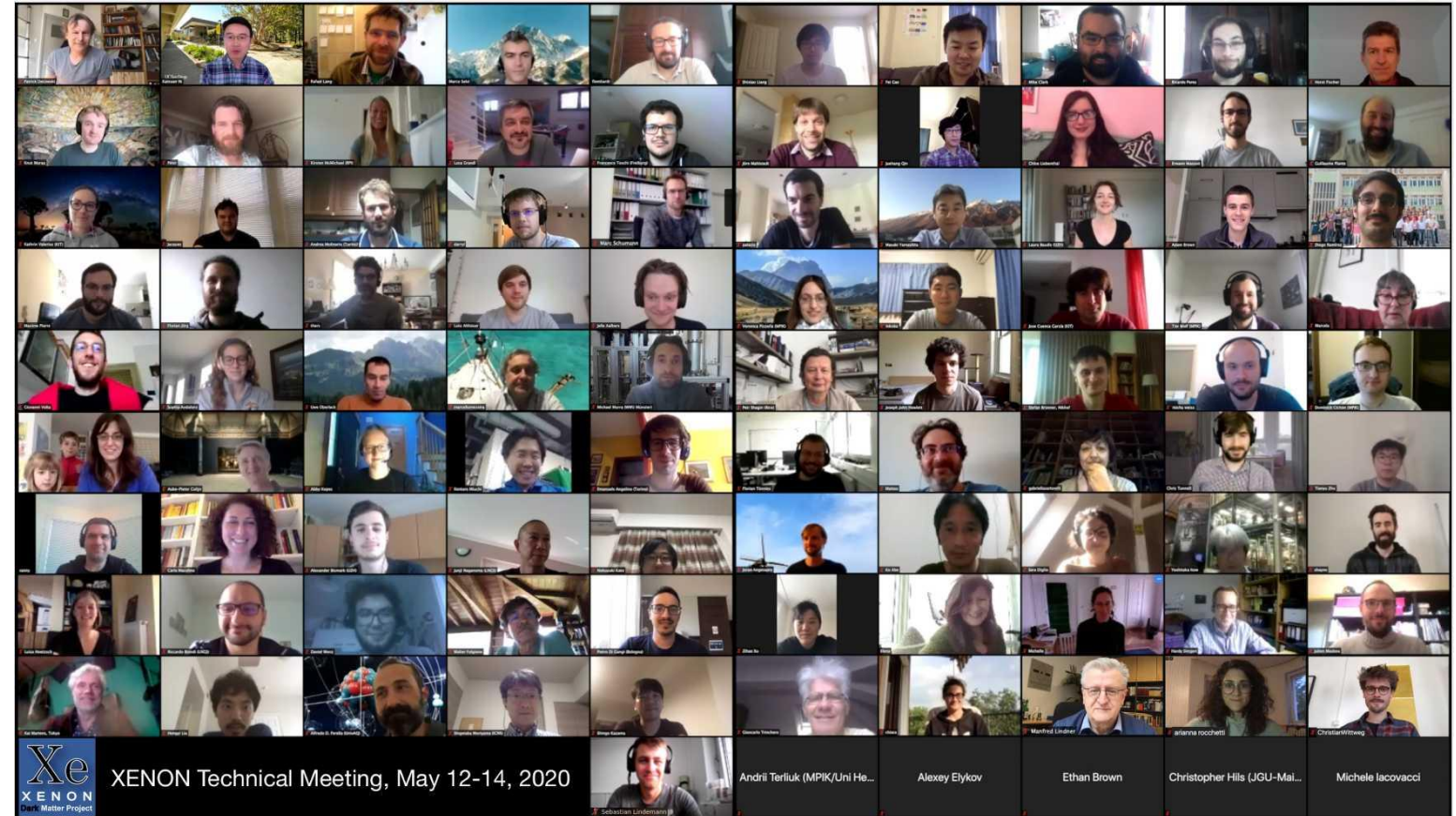
# XENON



University of  
Zurich <sup>UZH</sup>

## The XENON Collaboration

- 26 institutions worldwide
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# The TPC detection principle

- Dual-phase (liquid+gas)
- Energy reconstruction
- 3D event reconstruction
- Event discrimination (electronic recoil vs nuclear recoil)

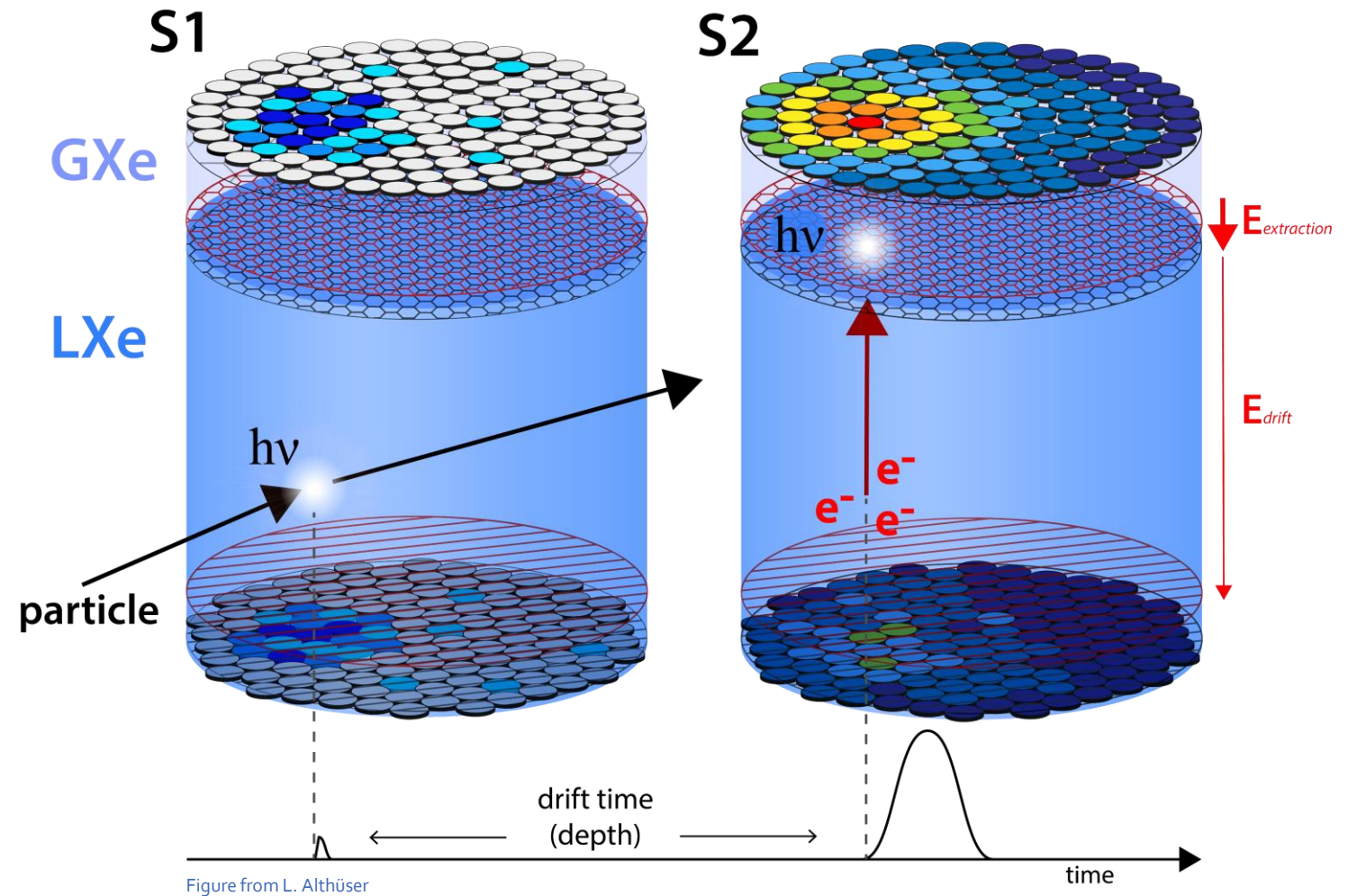
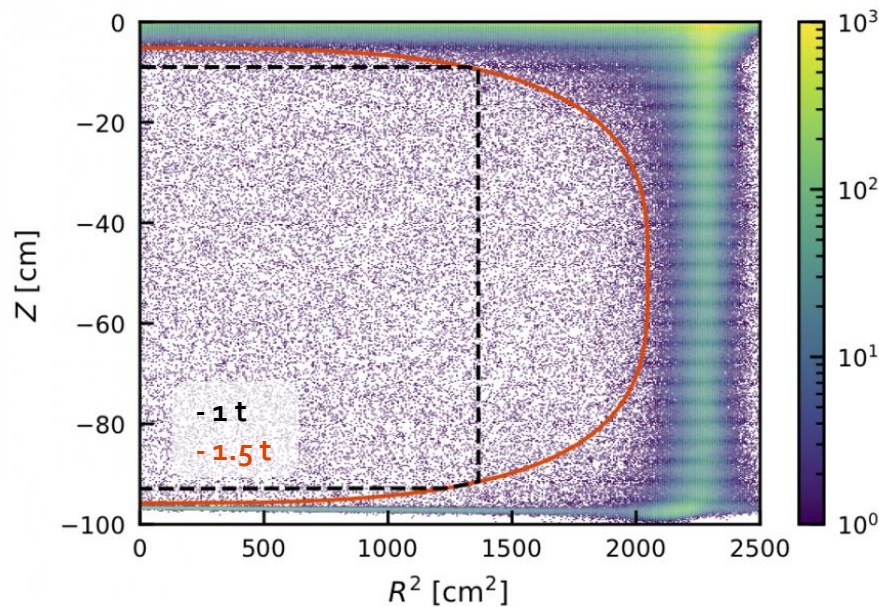
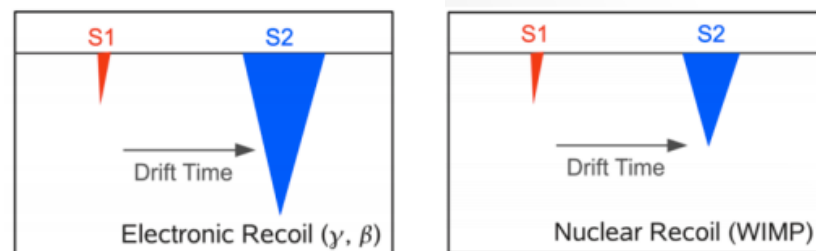


Figure from L. Althüser

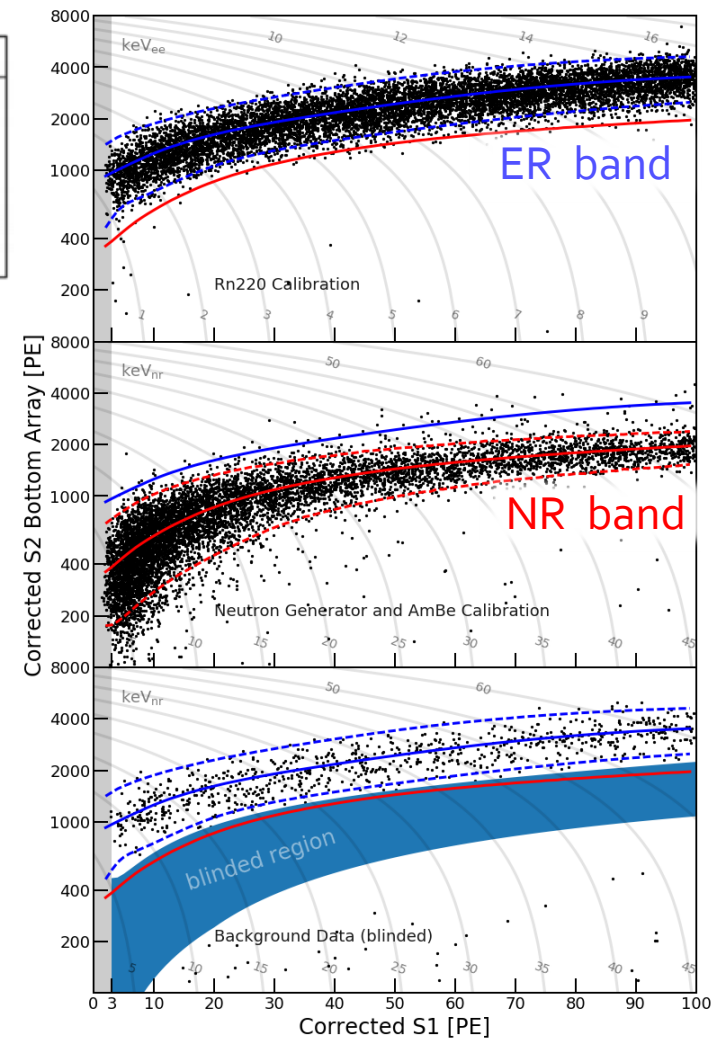


## The TPC detection principle

- Energy reconstruction
- ER/NR events discrimination
- Particle ID ( $\gamma$ ,  $\alpha$ ,  $\beta$ ,  $n$ , WIMP)
- Fiducialization



Phys. Rev. Lett. **119**, 181302 (2017)





## The XENON program

- Since 2005
- Dual-phase xenon TPCs
- XENONnT currently taking science data

XENON<sub>10</sub>

Target mass: 14kg

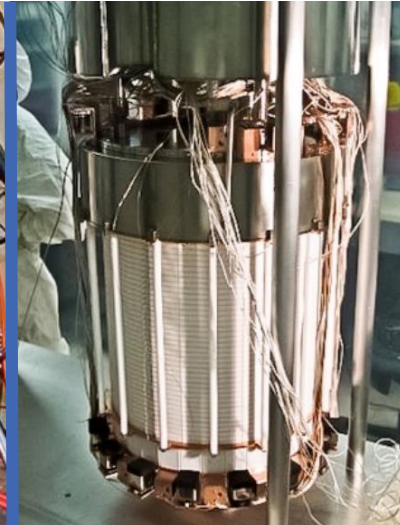
2005-2007



XENON<sub>100</sub>

Target mass: 62kg

2008-2016



XENON<sub>1T</sub>

Target mass: 2t

2013-2018



XENONnT

Target mass: 5.9t

2020 - present



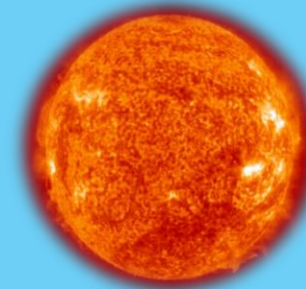


## Physics channels

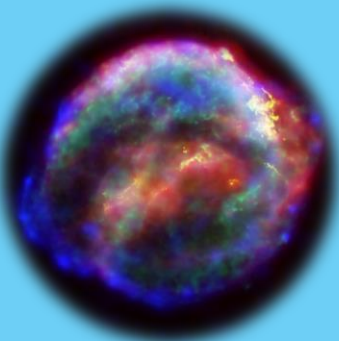
- Dark matter
- Solar neutrinos
- Supernova events
- Neutrino properties
- Atmospheric neutrinos



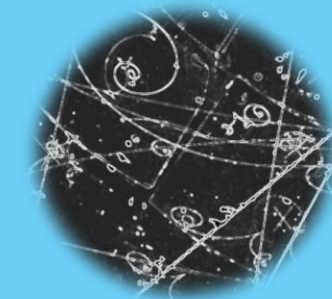
- WIMP-search
  - Spin-independent
  - Spin-dependent
- Sub-GeV
- Dark photons
- Axion-like particles



- Solar neutrinos
  - Boron-8
  - pp neutrinos
- Solar axions



- Supernova neutrinos
- Actively communicate with SNEWS
- Multi-messenger in DM experiments

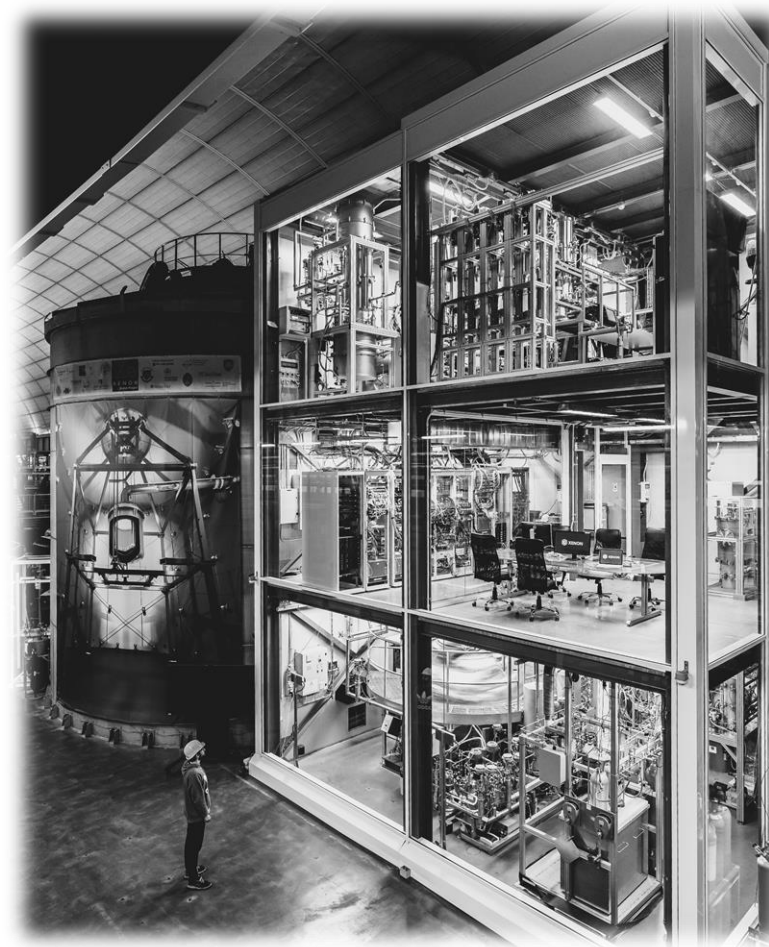
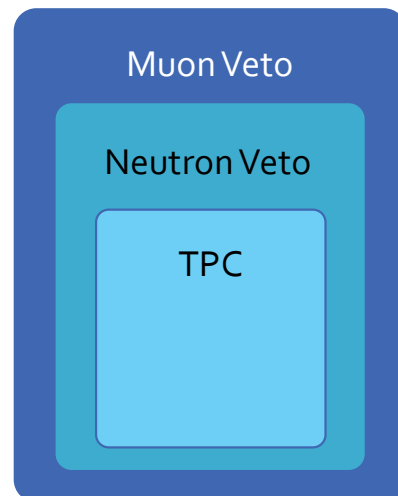


- Neutrino properties
  - Double beta decay of  $^{136}\text{Xe}$
  - Double-electron capture in  $^{124}\text{Xe}$
  - Neutrino magnetic moment



## XENONnT experiment

- 3 complementary detectors:
  - TPC
  - Muon Veto
  - Neutron Veto
- Several auxiliary systems
- Fast upgrade on 1T infrastructure



Cryogenics and  
purification

DAQ and SC

Calibration lines

Kr and Rn distillation

Xenon storage



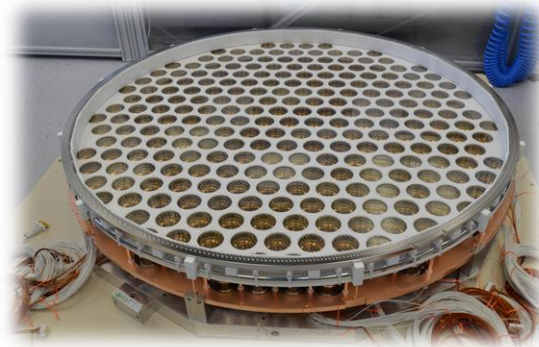


## TPC

- Larger active mass
- 1.5 m height
- 1.3 m diameter
- Finished installation summer 2020

5.9 t active liquid xenon  
(3x XENON1t)

494 3" PMTs



Tunable field shaping  
rings chains

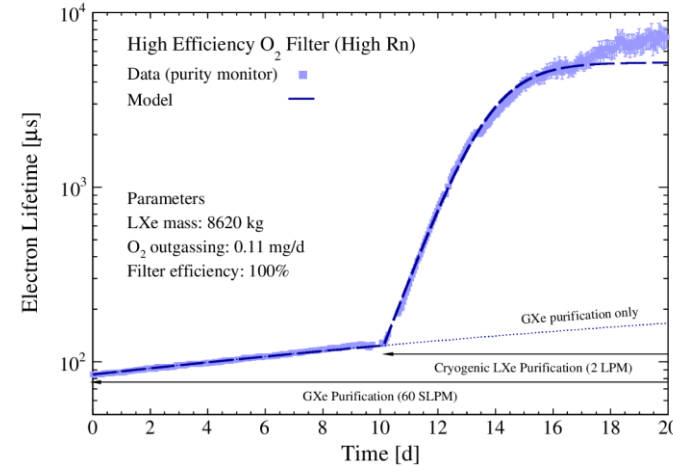
Thoroughly selected and  
screened materials



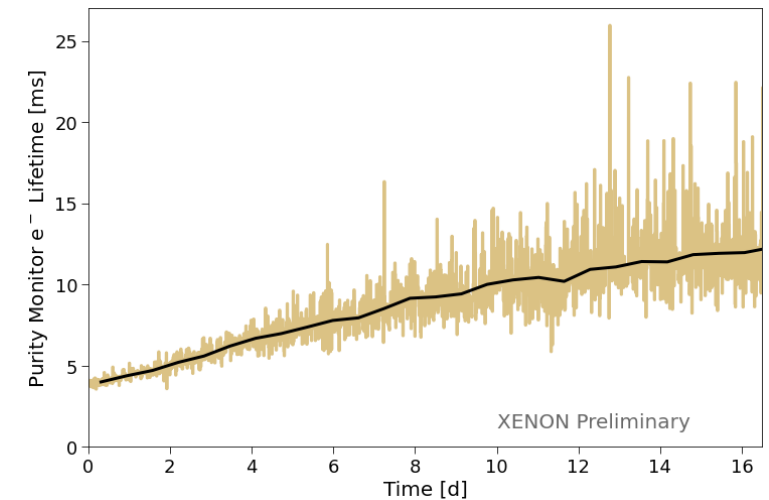


# Liquid Xenon Purification

- Purification directly in liquid
- ~1000 slpm purification speed
- Electron lifetime of >10 ms achieved
- Very low Rn emanation



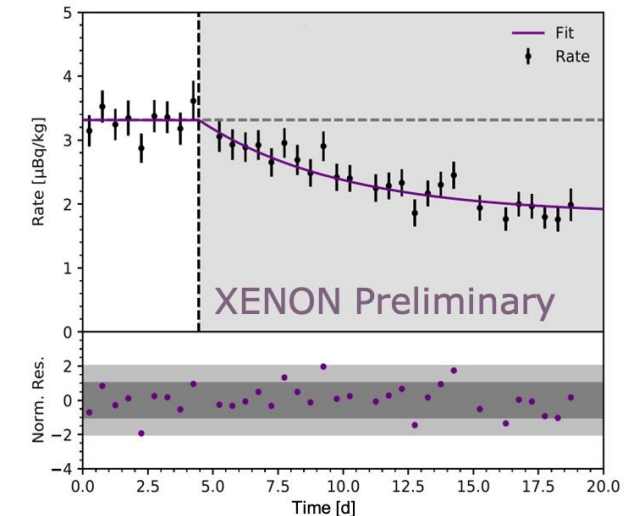
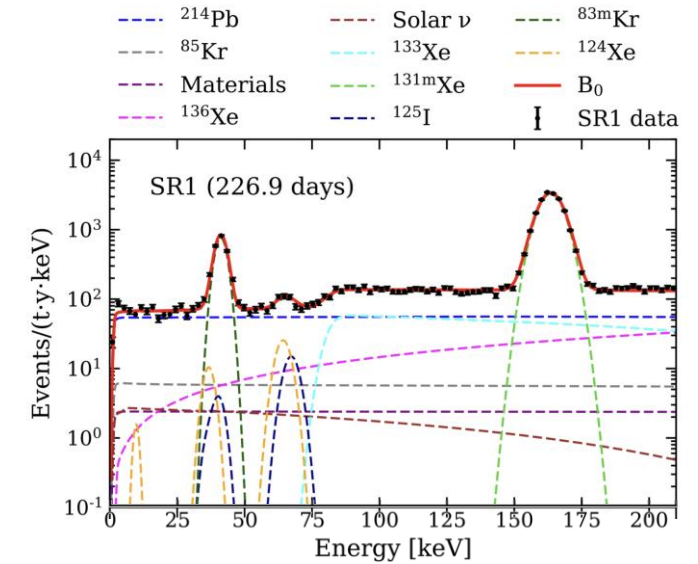
> 90% electron survival probability, improving S<sub>2</sub>-only and neutrino searches





# Radon Distillation Column

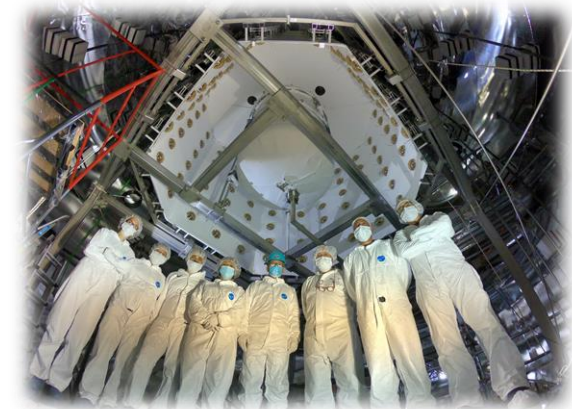
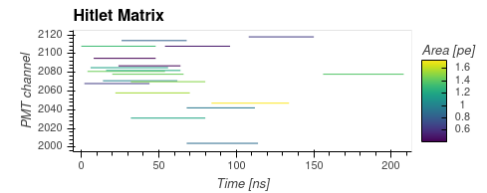
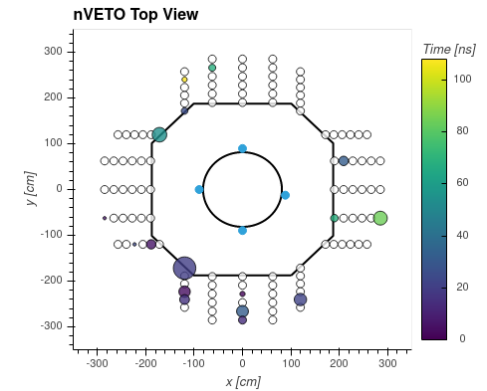
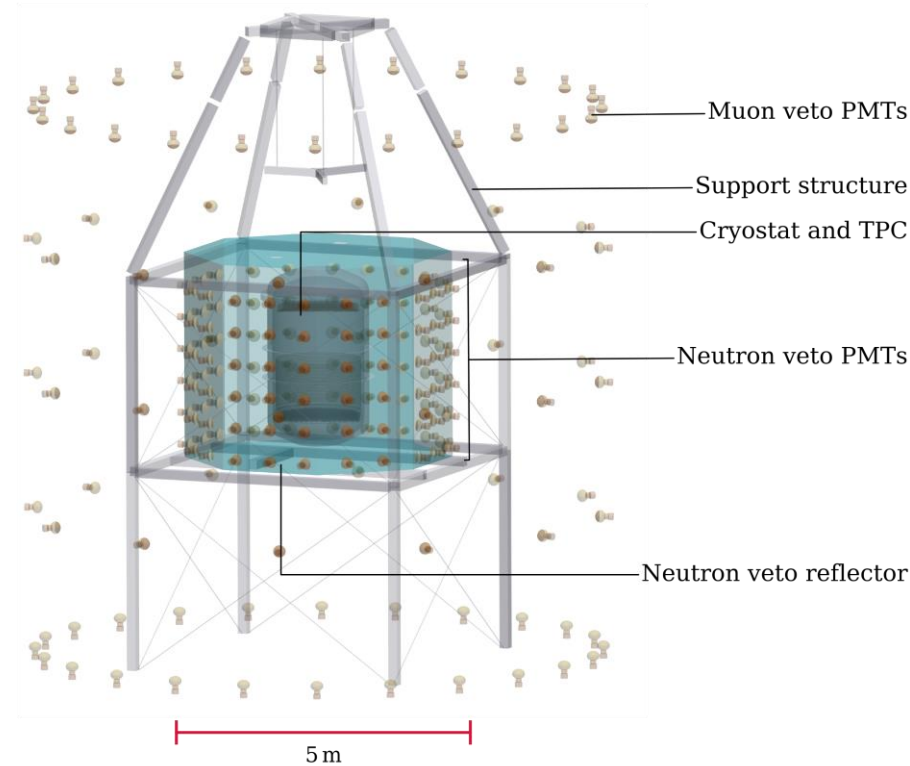
- Main internal ER background in XENON1T
- Developed and installed a Rn distillation column
- Goal of 1  $\mu\text{Bq/kg}$  (13  $\mu\text{Bq/kg}$  in XENON1T)
- 2  $\mu\text{Bq/kg}$  achieved and extra factor of 2 reduction possible





## Neutron and Muon Veto

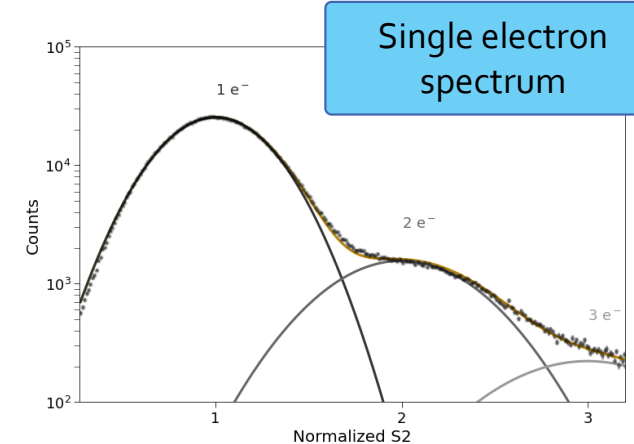
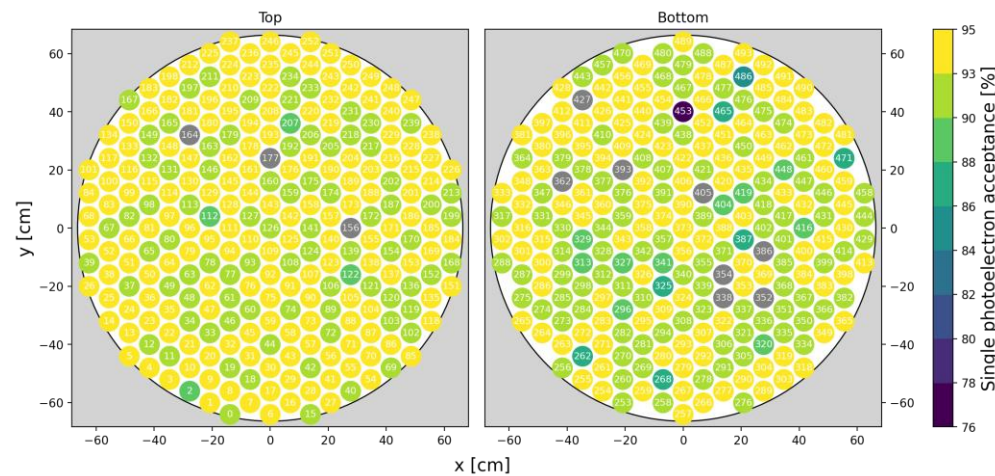
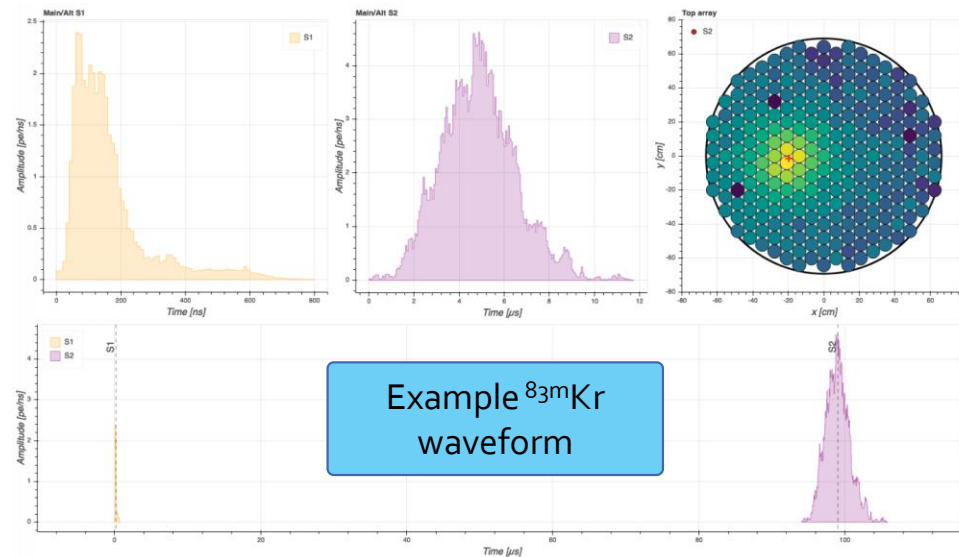
- Neutron Veto:
  - Gd loaded water Cherenkov detector (soon)
  - Optically separated from the MV
  - 120 dedicated PMTs
  - 87% neutron tagging efficiency by design
- Muon Veto:
  - Passive water shield
  - Cherenkov detector





## Data in science run configuration

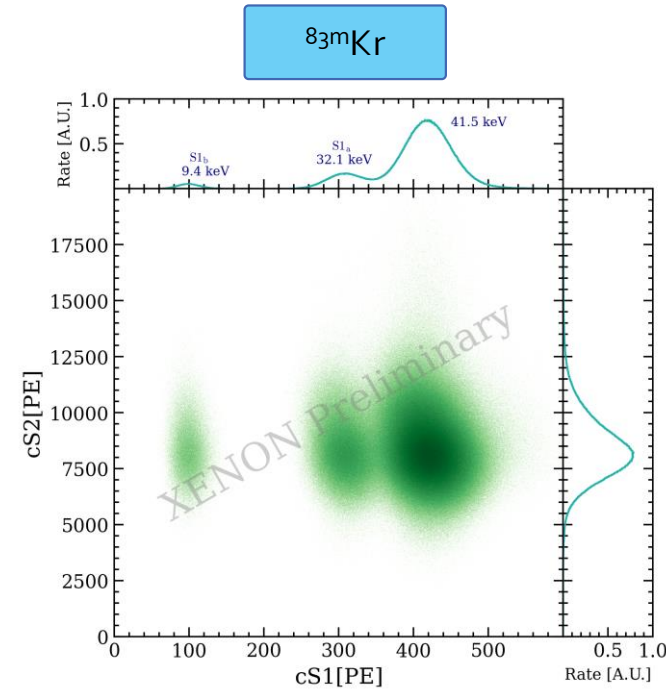
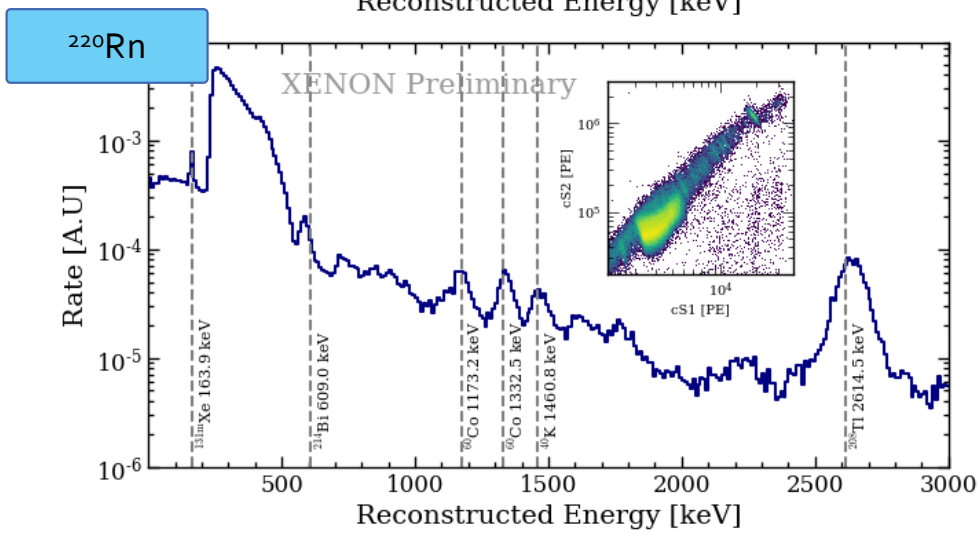
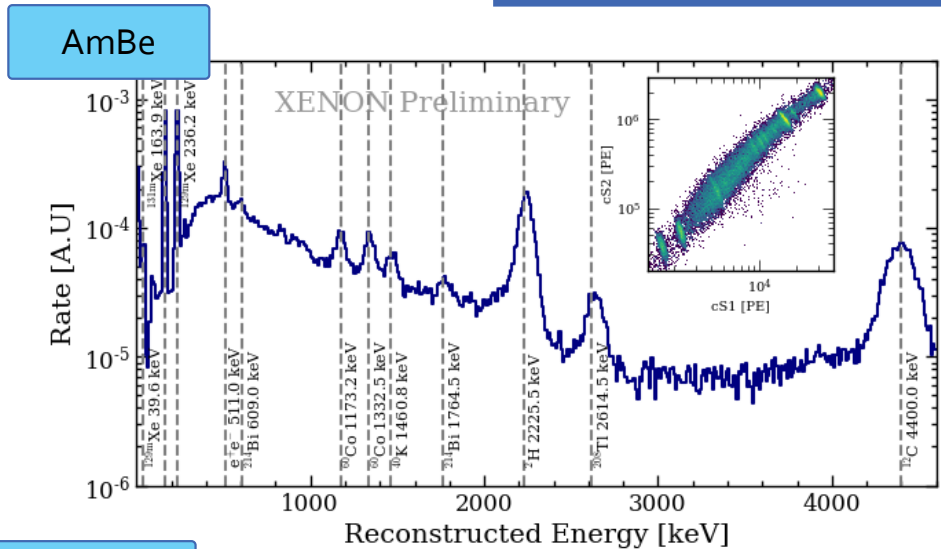
- PMT arrays performing well
- Single electron resolution with triggerless DAQ
- PDE:  $\sim 0.17$  PE/photon





# Data in science run configuration

- Successful calibration runs:
  - NR – AmBe
  - ER –  $^{220}\text{Rn}$
  - Detector effects and energy resolution –  $^{83\text{m}}\text{Kr}$
- 7.6% energy resolution at 41.5 keV

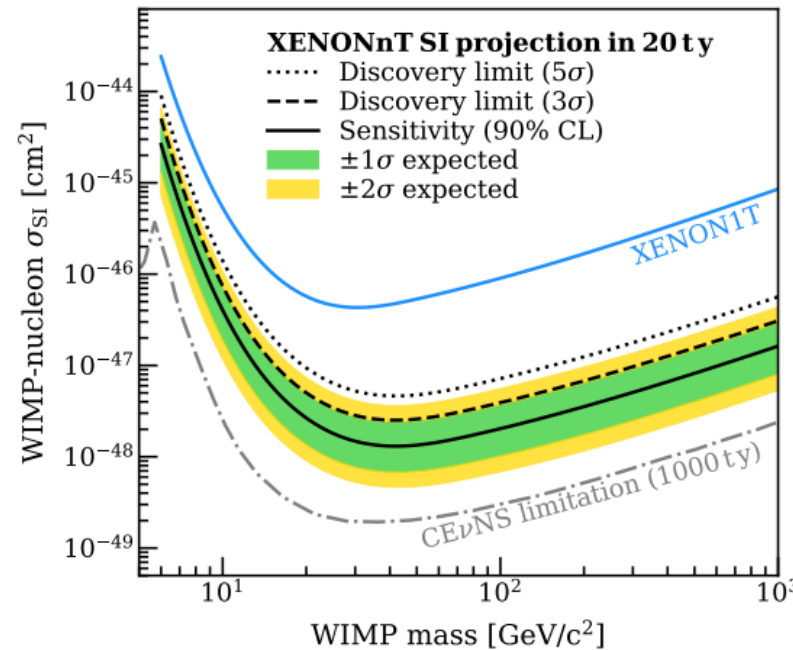




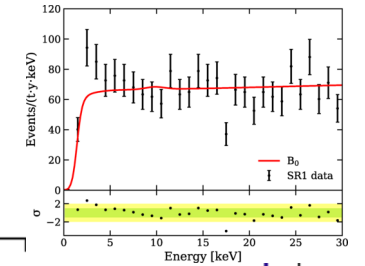
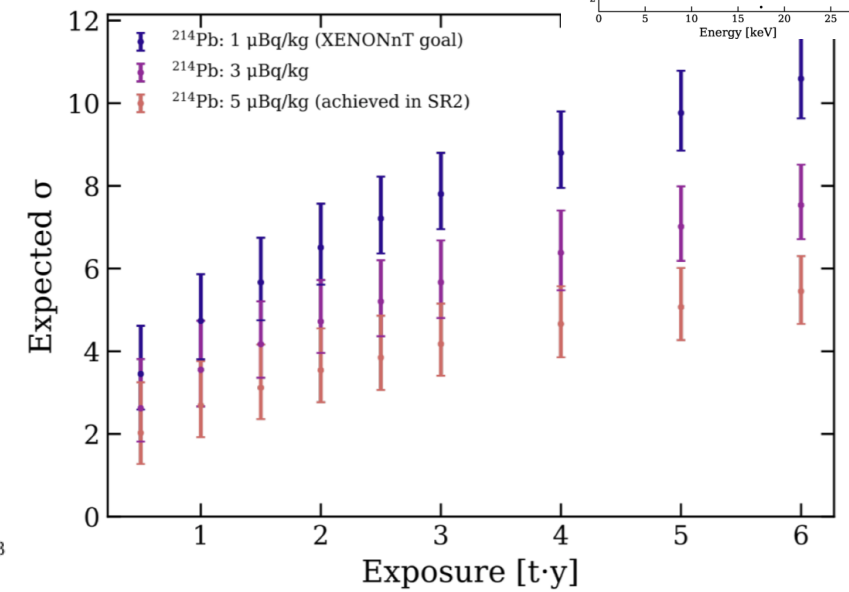
# First science results on the horizon

- Two main physics analysis:
  - WIMP search ( $1.4 \times 10^{-48}$  cm<sup>2</sup>@50 GeV/c<sup>2</sup> for 20 t.y)
  - Low-energy ER excess seen in XENON1T
- Most conditions on goal value (e-lifetime, <sup>222</sup>Rn level)
- First science run ongoing!

JCAP11(2020)031



Evan Shockley PhD Thesis (2020)





## Summary and Outlook



xe-pr@lngs.infn.it



@XENONexperiment



fb.com/XENONexperiment



@xenon\_experiment

- LXe TPCs exceptional to probe several rare-event processes.
- XENON1T made considerable improvements to the field throughout the years, both in dark matter, neutrino physics and technical design for low-background experiments
- XENONnT is taking science data with major improvements over XENON1T configuration:
  - target mass: 5.9 t
  - e-lifetime: > 10 ms (liquid purification)
  - Rn level: < 2  $\mu\text{Bq/kg}$  (Rn distillation column)
  - Neutron veto in place
- First results from XENONnT are on the horizon, shedding light on the XENON1T Low-energy ER excess and leading limits on the SI WIMP-nucleon interaction cross section.



# Backup Slides



**XENON**



## XENON<sub>1</sub>T/nT detectors

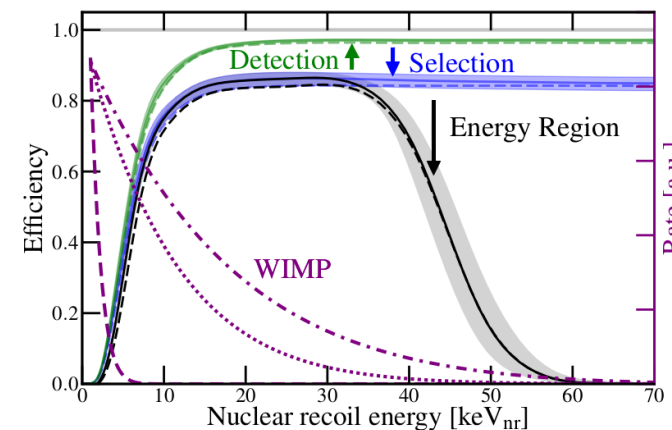
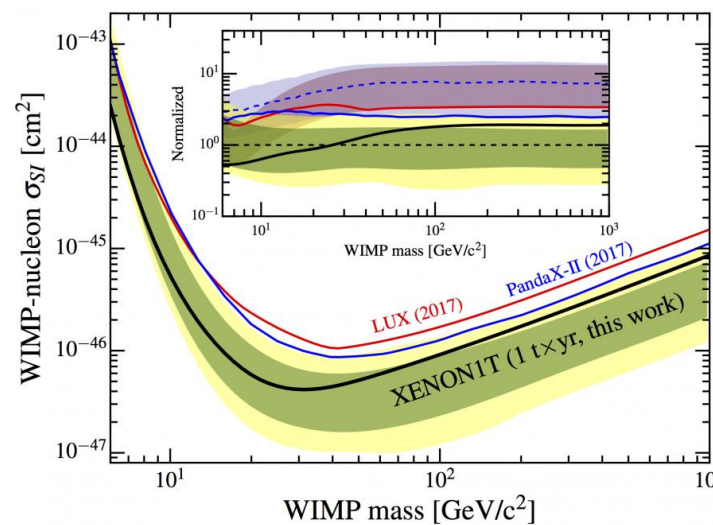
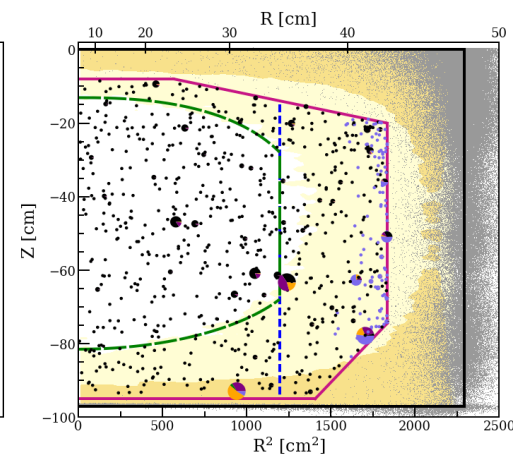
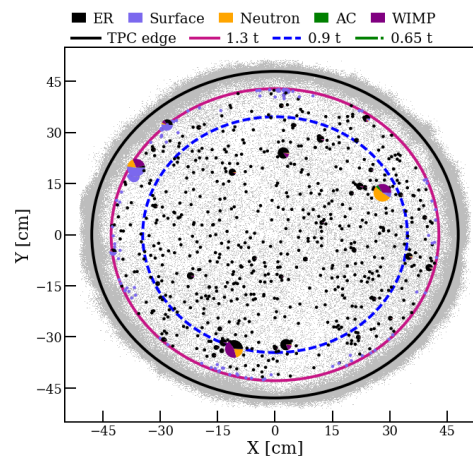
- Dual-phase XENON TPC in a water passive and active muon veto
- Service building:
  - Xenon storage and recuperation
  - DAQ
  - Kr distillation column
  - Cryogenics and calibration system





# Dark matter search

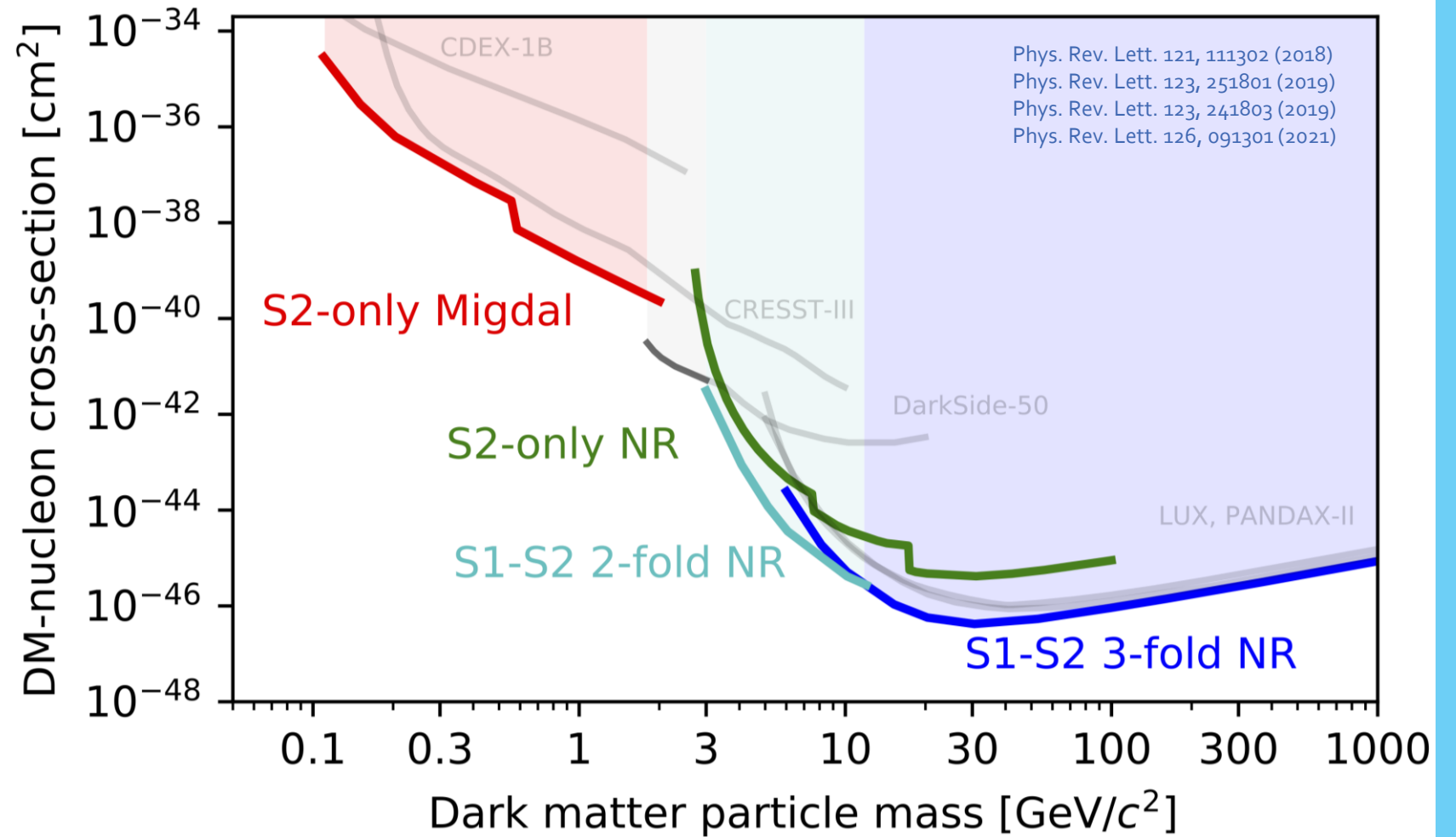
- Phys. Rev. Lett. 121, 111302 (2018)
- No significant excess over background
- Most stringent WIMP-nucleon cross section:  $4.1 \times 10^{-47} \text{ cm}^2$  @ 30 GeV/c<sup>2</sup>, 90% CL





# XENON<sub>1</sub>T SI WIMP results

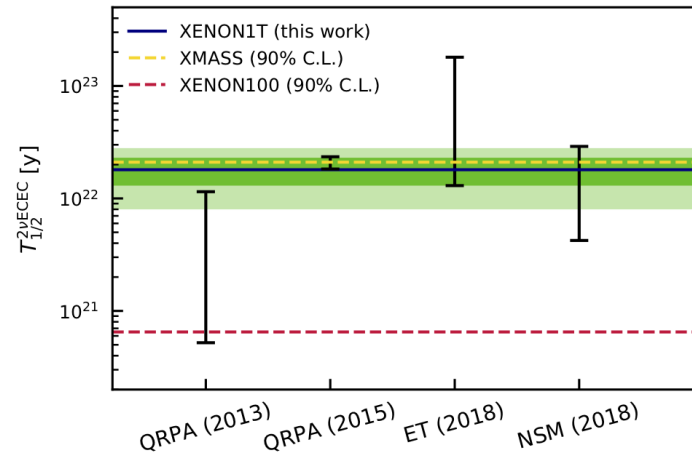
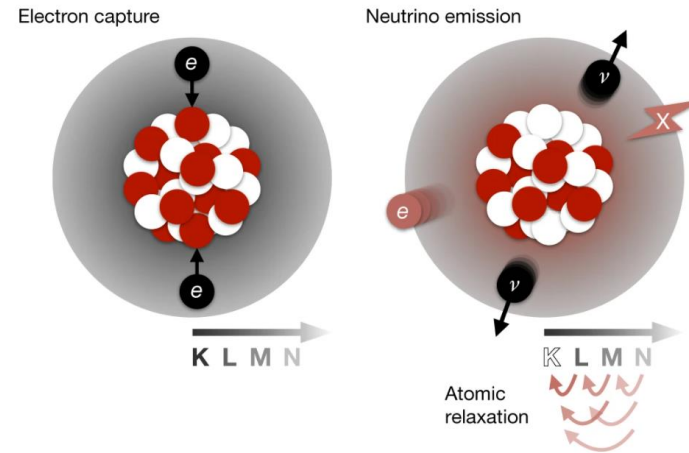
- Extended DM search with ionization-only channel
- No complete background model (only limit setting)
- O(100 eV) energy threshold



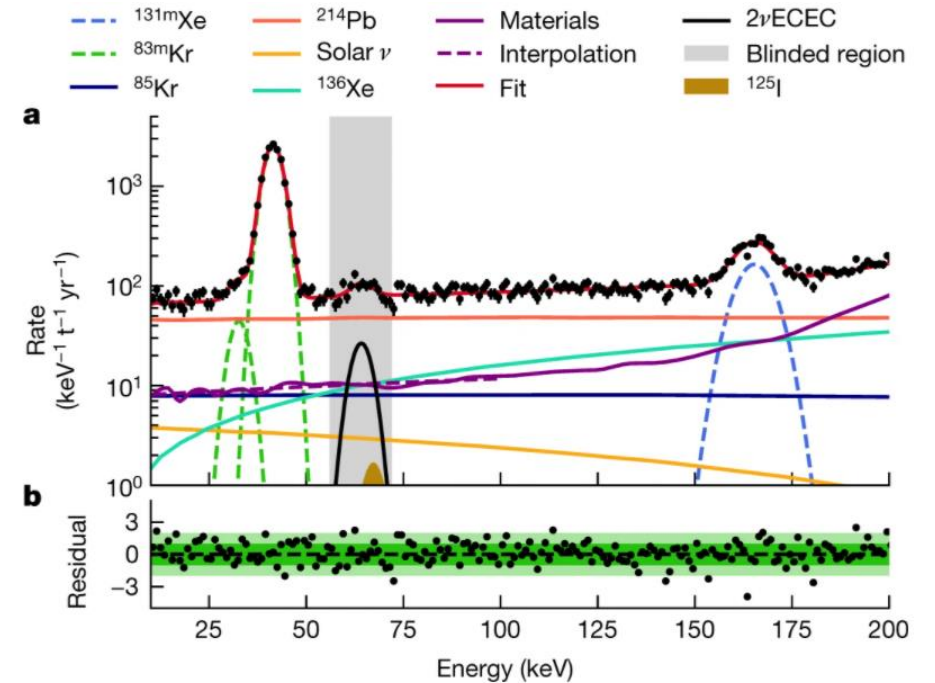


# Double Electron Capture in $^{124}\text{Xe}$

- Observation of X-rays and Auger electrons,  $Q_{\text{value}} = 64,3 \text{ keV}$
- Longest half-life ever observed directly:  $1.8 \times 10^{22}$  years at  $4.4\sigma$  significance
- First step for neutrinoless DEC search



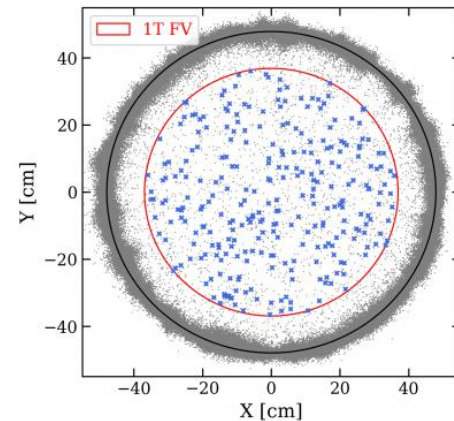
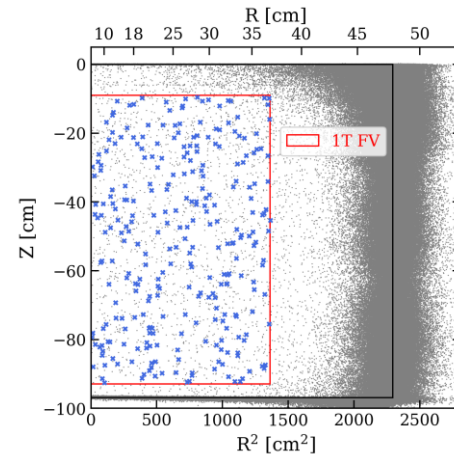
*Nature volume 568, 532–535(2019)*



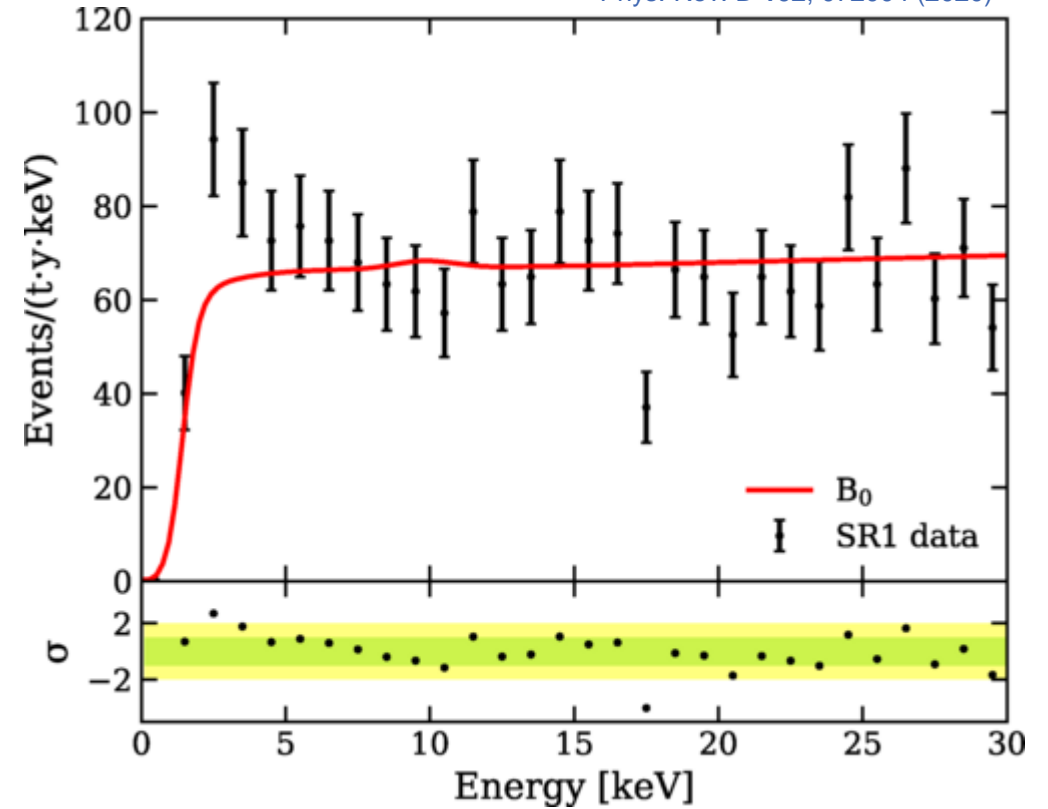


## Low-ER excess

- ER search in  $<30$  keV range
- $3.3\sigma$  fluctuation of background in 1-7 keV
- Several hypothesis:
  - Solar axions ( $3.4\sigma$  over bkg)
  - Neutrino magnetic moment ( $3.2\sigma$  over bkg)
  - Bosonic DM: ALPs and dark photons ( $3.0\sigma$  over bkg)
  - Tritium ( $3.2\sigma$  over bkg)



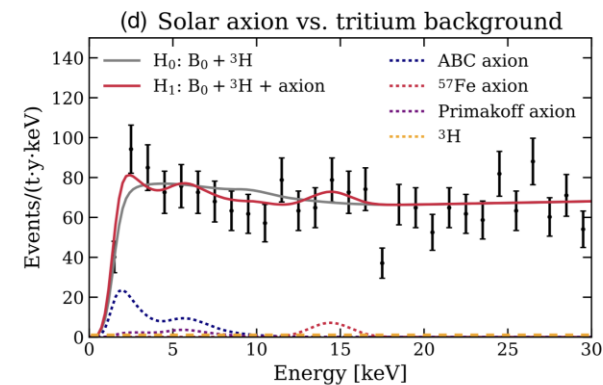
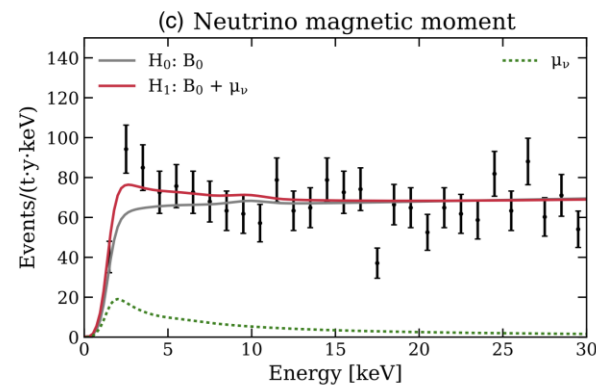
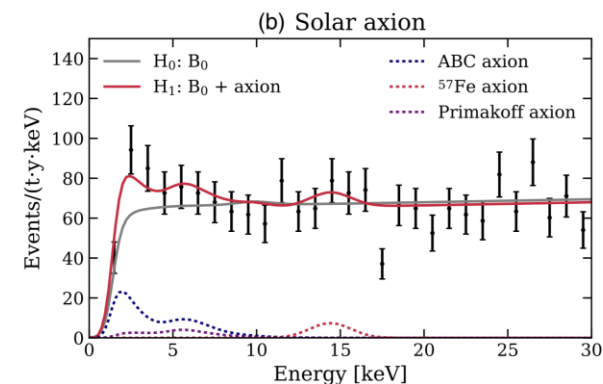
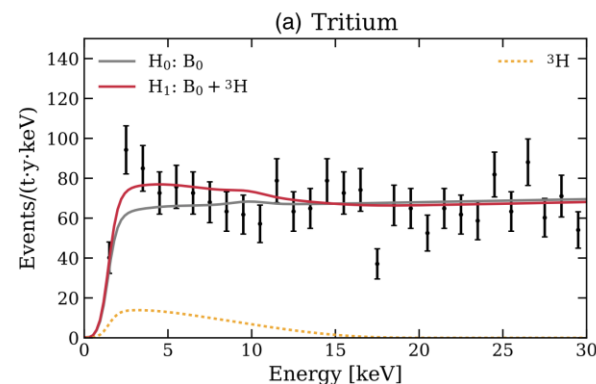
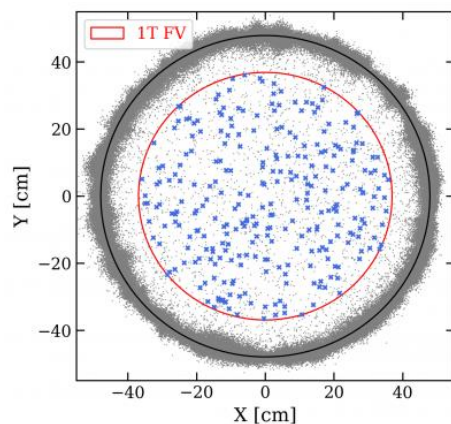
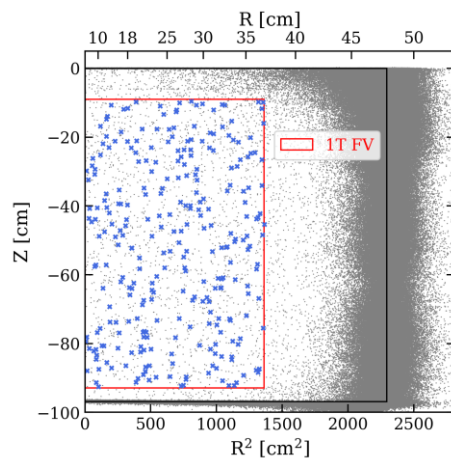
Phys. Rev. D **102**, 072004 (2020)





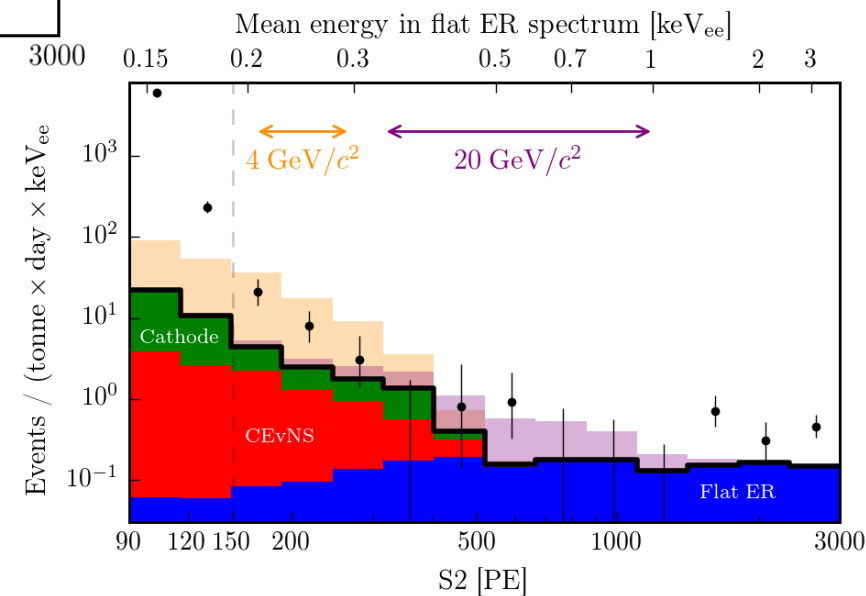
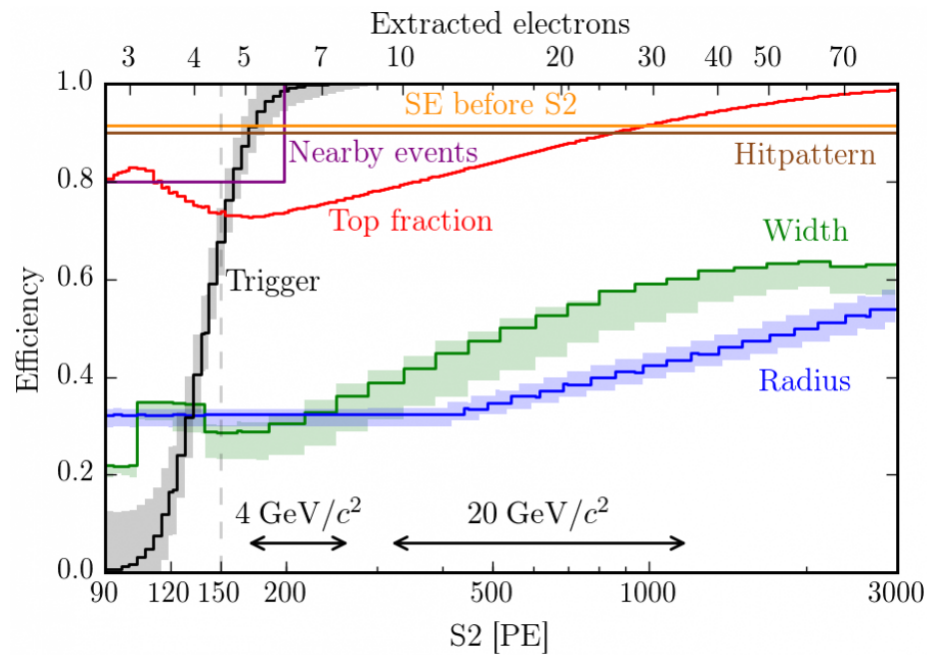
# Low-ER excess

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- Neutrino magnetic moment ( $3.2\sigma$  over bkg)
- Bosonic DM: ALPs and dark photons ( $3.0\sigma$  over bkg)
- Tritium ( $3.2\sigma$  over bkg)





### S2-only

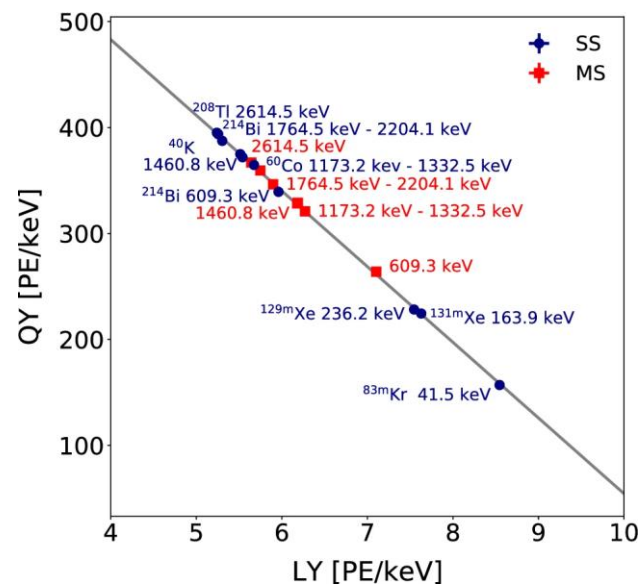






# ER Energy Scale

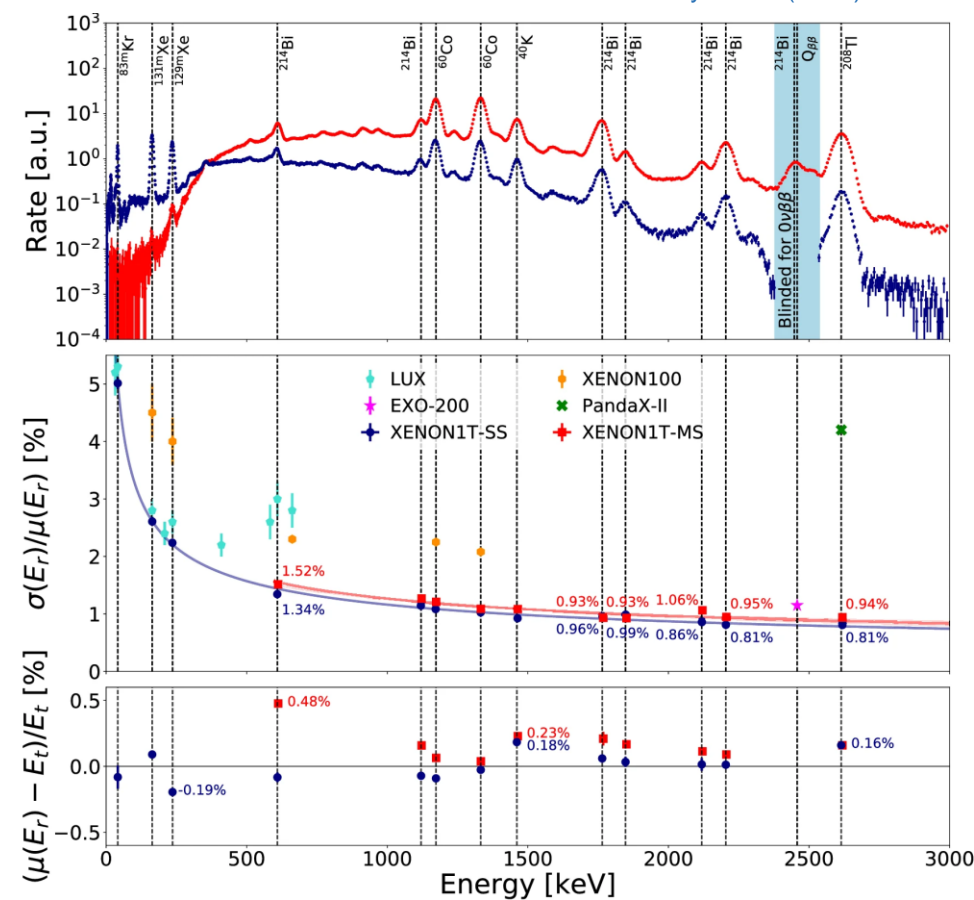
- ER background MC-data matching validates framework
- Very low energy resolution at  $^{136}\text{Xe}$   $Q_{\beta\beta}=2.46\text{MeV}$ :  $(0.80\pm 0.02)\%$
- Promising for near future neutrinoless double-beta decay results!



$$E = (n_{ph} + n_e) \cdot W$$

$$= \left( \frac{S1}{g1} + \frac{S2}{g2} \right) \cdot W$$

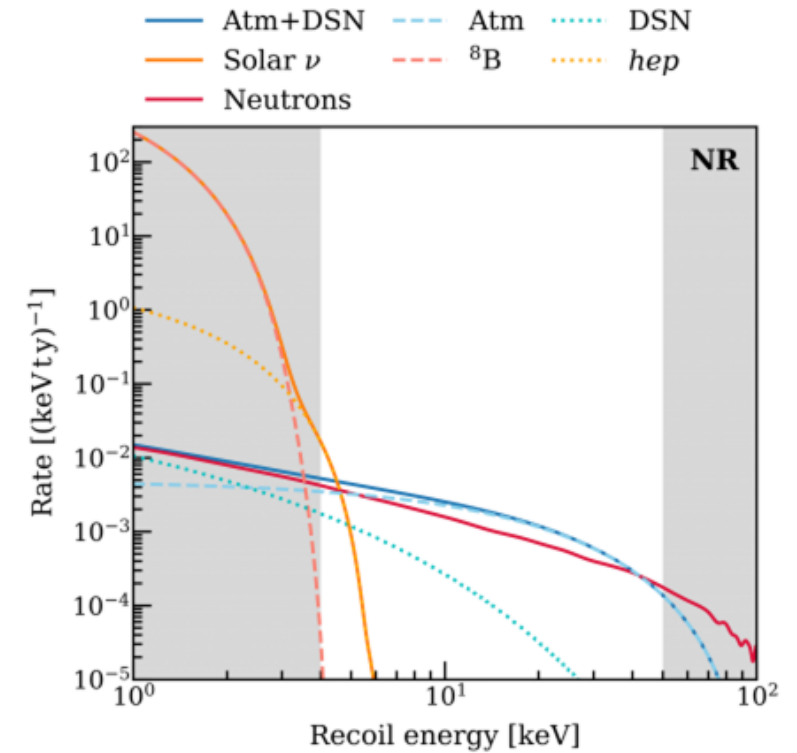
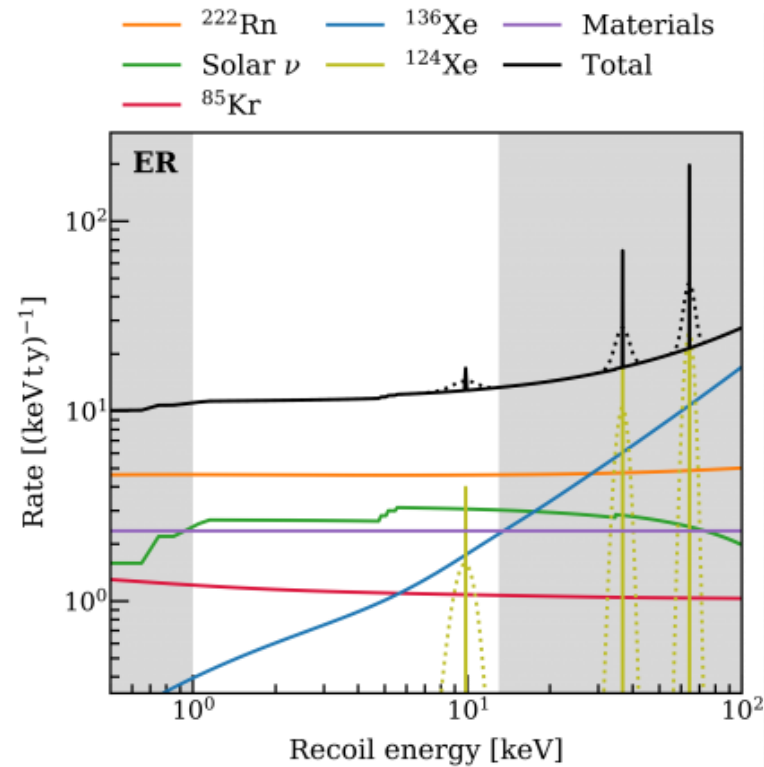
Eur. Phys. J. C (2020) 80:785





# ER and NR background

- MC projections for XENONnT
- JCAP11(2020)031

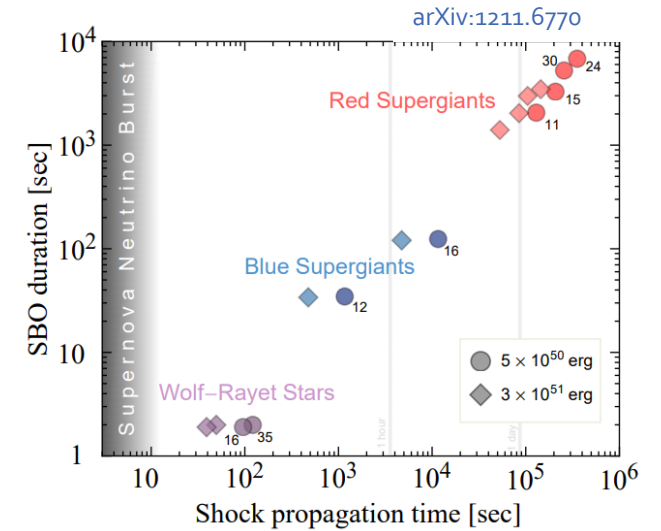
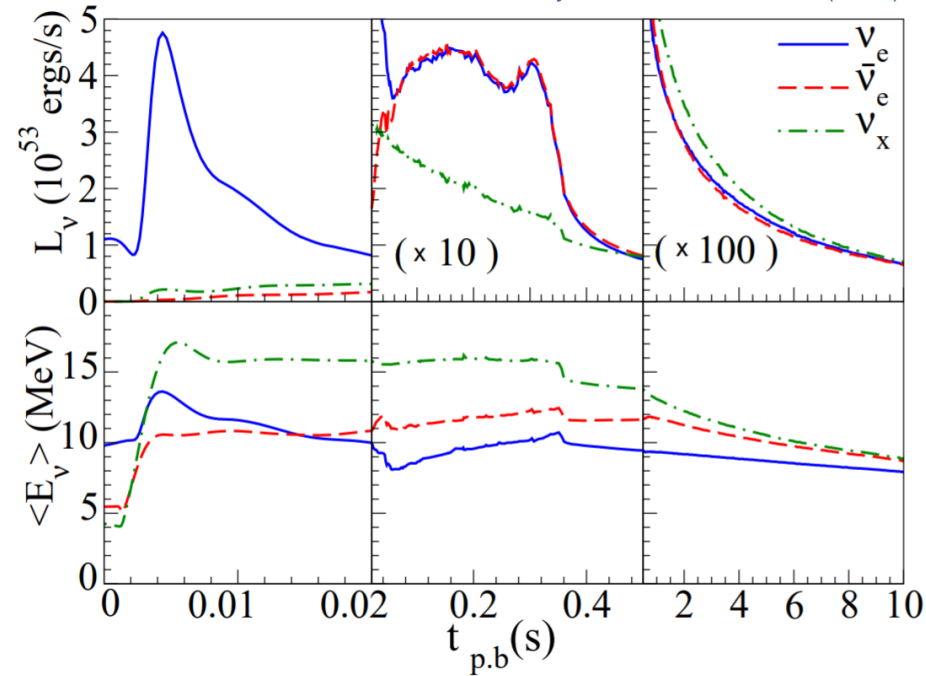




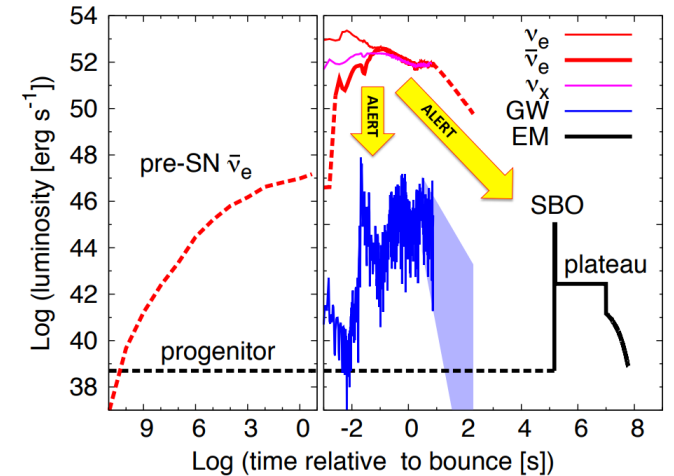
## XENONnT as supernovae observatory

- Core-collapse supernovae produce an enormous amount of  $O(10)$  MeV neutrinos ( $2.2 \times 10^{53}$  erg in SN1987a)
- SN neutrinos precede the EM radiation from minutes up to days
- Neutrino signal can be used as an early warning
- Signals from obscured SNe, in common with GW

Phys. Rev. D 94, 103009 (2016)



arXiv:2011.00035



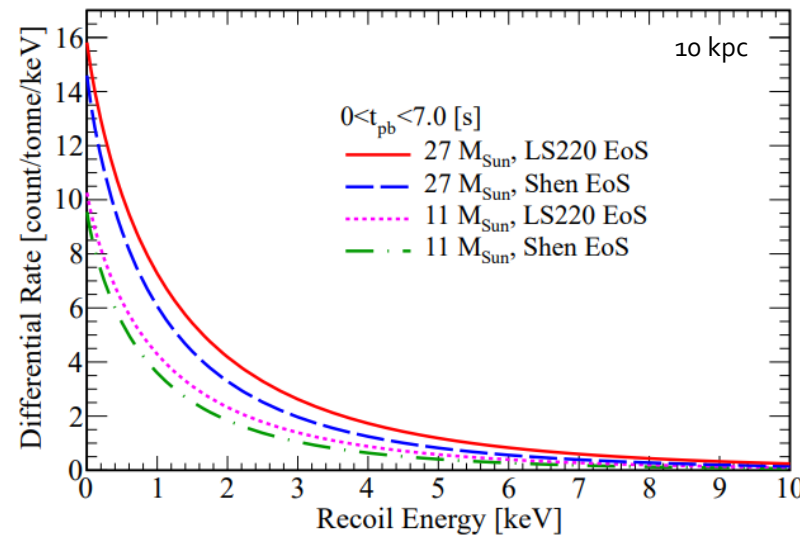


## XENONnT as a SN observatory

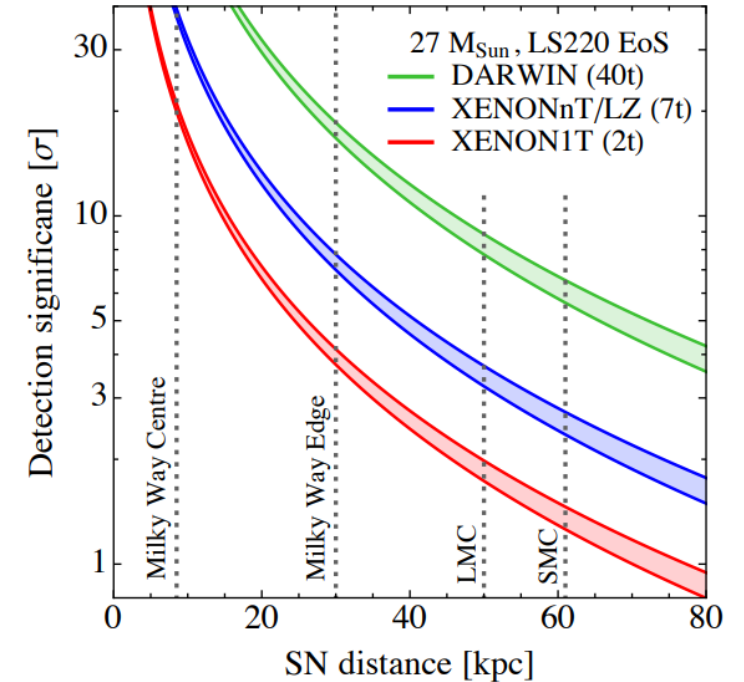
- In XENONnT, SN neutrinos mostly interact through CEvNS, a flavour independent channel
- Low-NR, O(1 keV), enhanced by ionization-only channel
- Constrained by mass, energy threshold and single-electron background
- Expecting ~80 events in the TPC for a 27Mo@10kpc

$$\frac{d\sigma}{dE_r} = \frac{G_F^2}{4\pi} Q_w^2 M \left( 1 - \frac{ME_r}{2E_\nu^2} \right) F(E_r)^2$$

$$Q_w^2 = \left[ \left( \frac{1}{2} - 2\sin^2(\theta_w) \right) Z - \frac{1}{2} N \right]^2$$



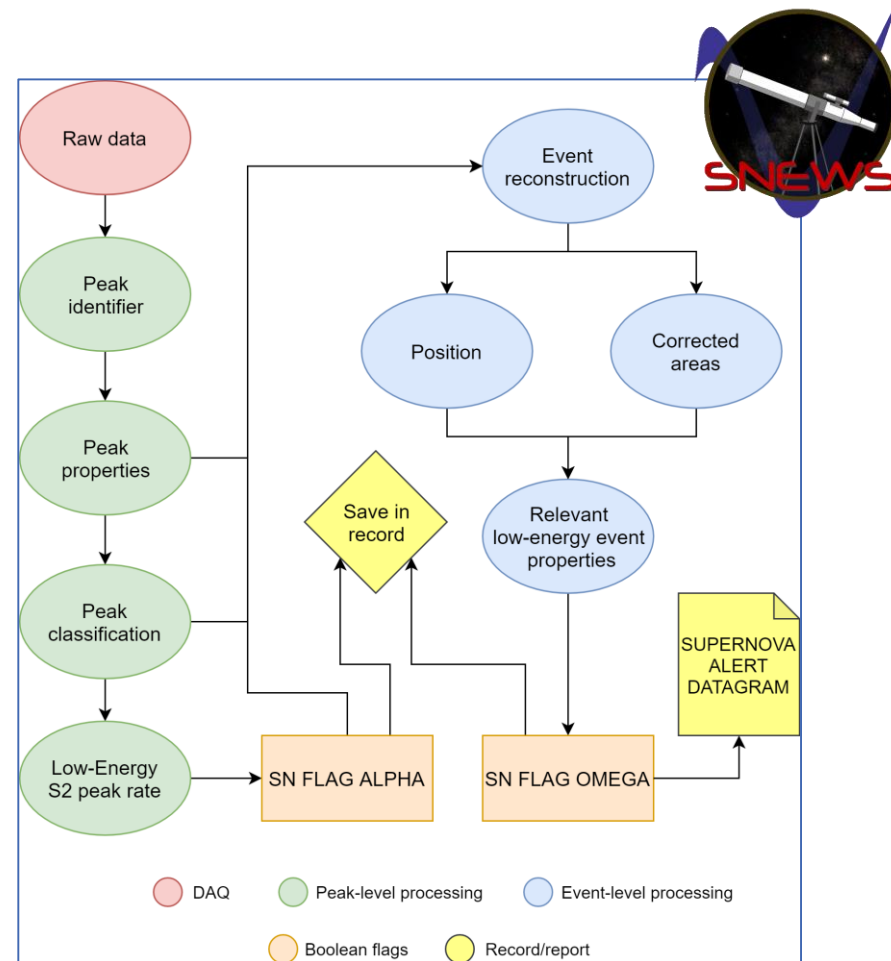
Phys. Rev. D 94, 103009 (2016)





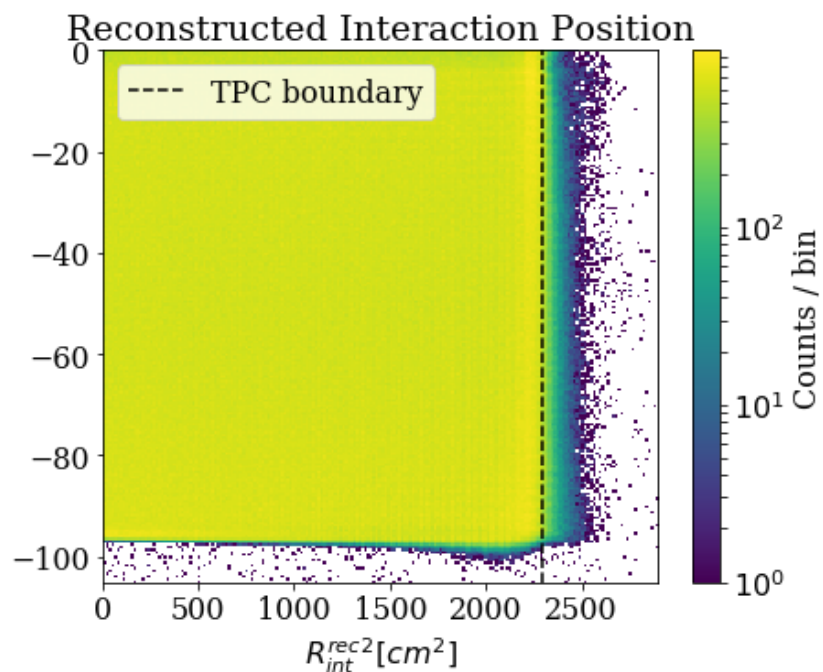
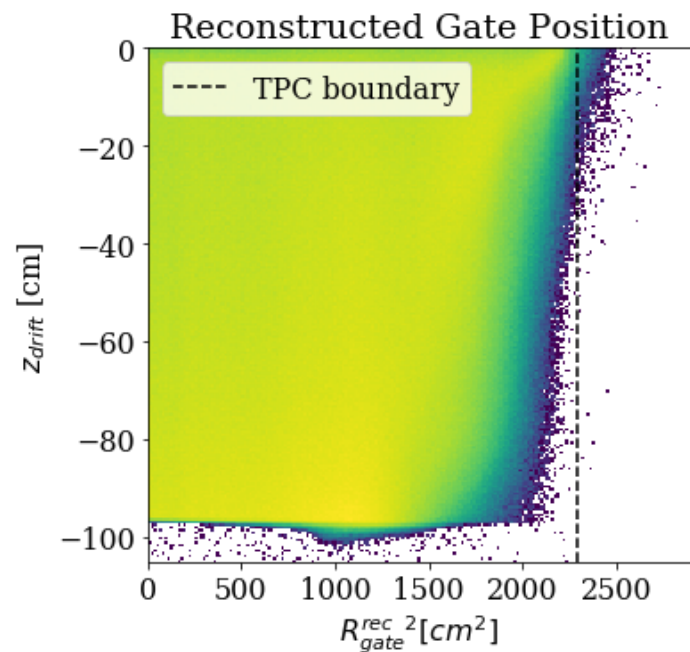
## A collaborative effort

- XENONnT will receive alerts from the SNEWS network and act on its data accordingly
- No pointing available, only timing
- Prompt response to SNe signals under study
- Actively contributing to SNEWS 2.0 (arXiv:2011.00035) under consideration





# Field Corrections





# Light and Charge yields

