**ISMD 2022 Conference** 

## **Beyond NNLO in Global PDF fits MSHTaN<sup>3</sup>LO** Parton Distribution Functions

Jamie McGowan, Thomas Cridge, Lucian Harland-Lang and Robert Thorne **August 2022** 

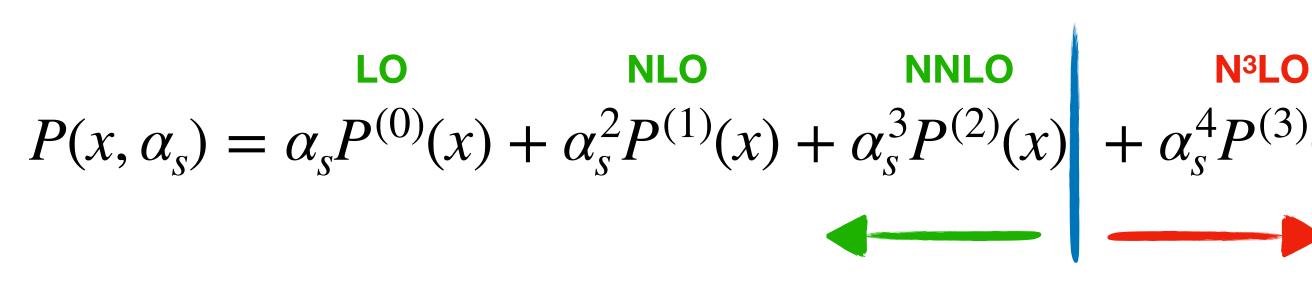


Full information in recent article: JM et. al., <u>2207.04739</u>



### What is a theoretical uncertainty? And also... why do we care?

• Leading source from **Missing High** different areas these occur in  $F_2$ .

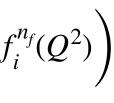


- Current knowledge is up to NNLO, with higher orders unknown.
- Potentially large corrections hiding in higher orders beyond theory truncation.
- Already progress in calculating features at N<sup>3</sup>LO<sup>[1-11]</sup>.

### • Leading source from Missing Higher Orders in perturbation theory - many

$$F_{2}(x,Q^{2}) = \sum_{\alpha \in \{H,q,g\}} \sum_{i \in \{q,g\}} \left( C_{q,\alpha}^{GMVF,n_{f}+1} \otimes A_{\alpha i}(Q^{2}/m_{h}^{2}) \otimes f_{i}^{n_{f}}(Q^{2}) + C_{H,\alpha}^{GMVF,n_{f}+1} \otimes A_{\alpha i}(Q^{2}/m_{h}^{2}) \otimes f_{i}^{A} P^{(3)}(x) + \dots \right)$$

$$\frac{df}{d \ln \mu_{f}^{2}} = P \otimes f$$





# **Theoretical Uncertainties in a Global PDF Fit** $\begin{cases} P(T|D) \propto \exp\left(-\frac{1}{2}M^{-1}(\theta' - \overline{\theta}')^2 - \frac{1}{2}(T' - D)^T H(T' - D)\right) \\ P(\theta') = \frac{1}{\sqrt{2\pi\sigma_{\theta'}}}\exp(-\theta'^2/2\sigma_{\theta'}^2) \end{cases}$

$$P(T|D) \propto \exp\left(-\frac{1}{2}(T-D)^T H_0(T-D)\right)$$

- we have?
- by a prior probability distribution<sup>[12]</sup>.
- close to the behaviour already known).



• Do we need to wait for a full description of the next order to be able to use the knowledge

• Can attempt to parameterise the higher order effects with a nuisance parameter defined

• Allow the fit to move these N<sup>3</sup>LO parameters (with a **penalty attached** to ensure we stay

With these alterations, we follow the **same** practice as set out in the MSHT20 NNLO PDF fit - the exact same global fit is done to approximate  $N^{3}LO$  (a $N^{3}LO$ ).

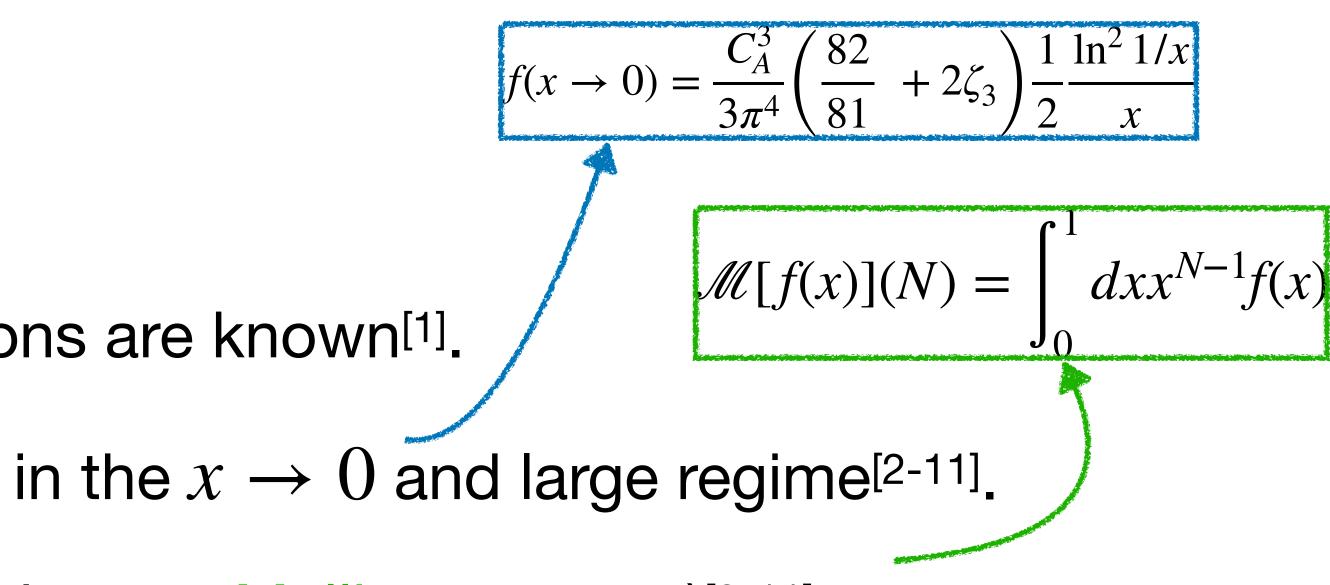




## What do we know?

...and what don't we know?

- Zero-mass N<sup>3</sup>LO coefficient functions are known<sup>[1]</sup>.
- Some knowledge of leading terms in the  $x \rightarrow 0$  and large regime<sup>[2-11]</sup>.
- Some numerical constraints (Low-integer Mellin moments)<sup>[2-11]</sup>.
- **Intuition** from lower orders/expectations from perturbation theory.
- Other parts, we know a very **limited amount** about  $(A_{gg,H}^{(3)})$  and most *K*-factors) [8-10]

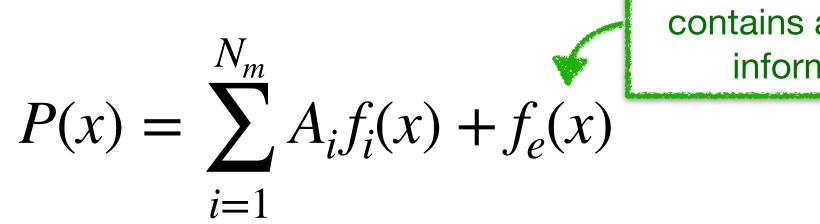






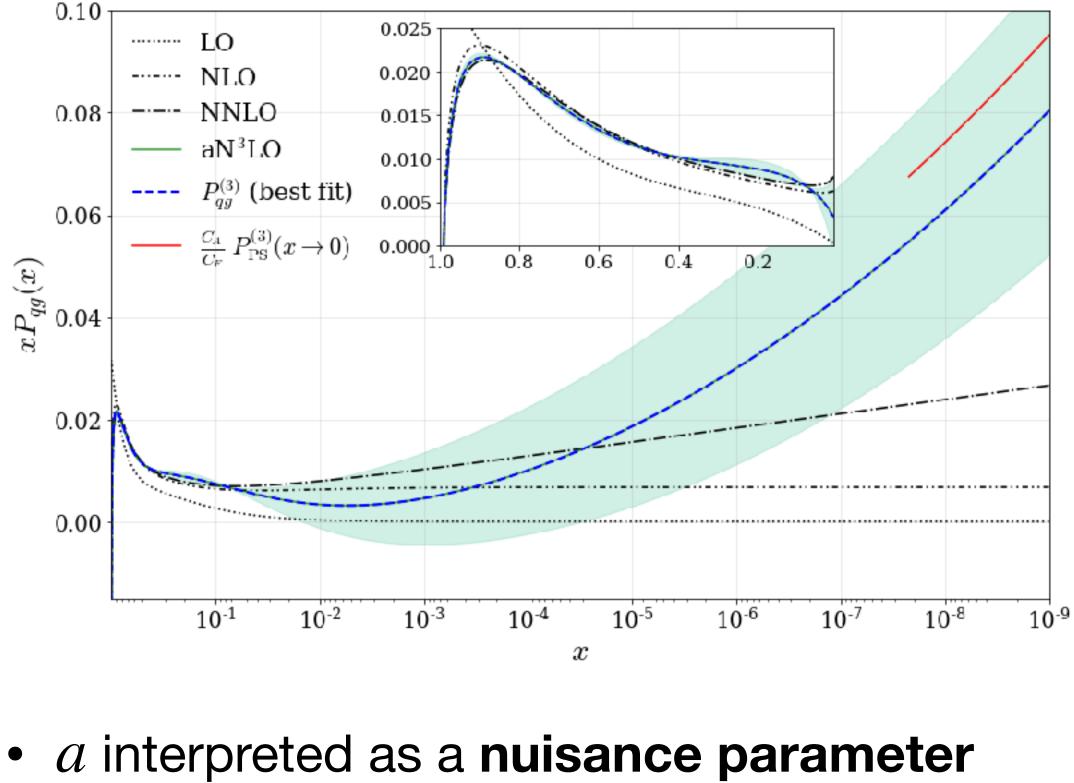
### Splitting Functions up to N<sup>3</sup>LO ...approximately

- Consider we know  $N_m$  Mellin moments<sup>[1-5]</sup>.
- With  $N_m$  constraints, we employ:



- Choose a set of **relevant functions**  $f_i$  and solve for  $A_i$ .
- To allow control of this function, introduce a lacksquaredegree of freedom *a*.  $f_{\rho}(x) \rightarrow f_{\rho}(x, a)$

contains any known information.



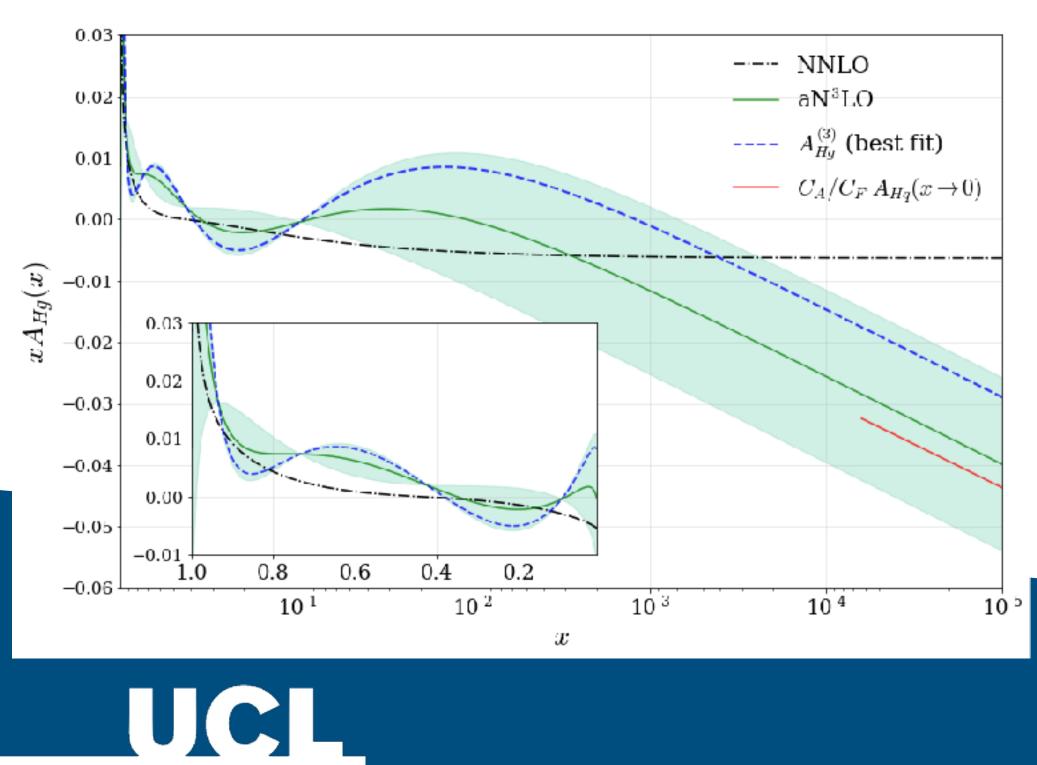
- allowed to vary in a PDF fit.
- In our treatment *a* is the **coefficient of the** most divergent unknown small-x term.

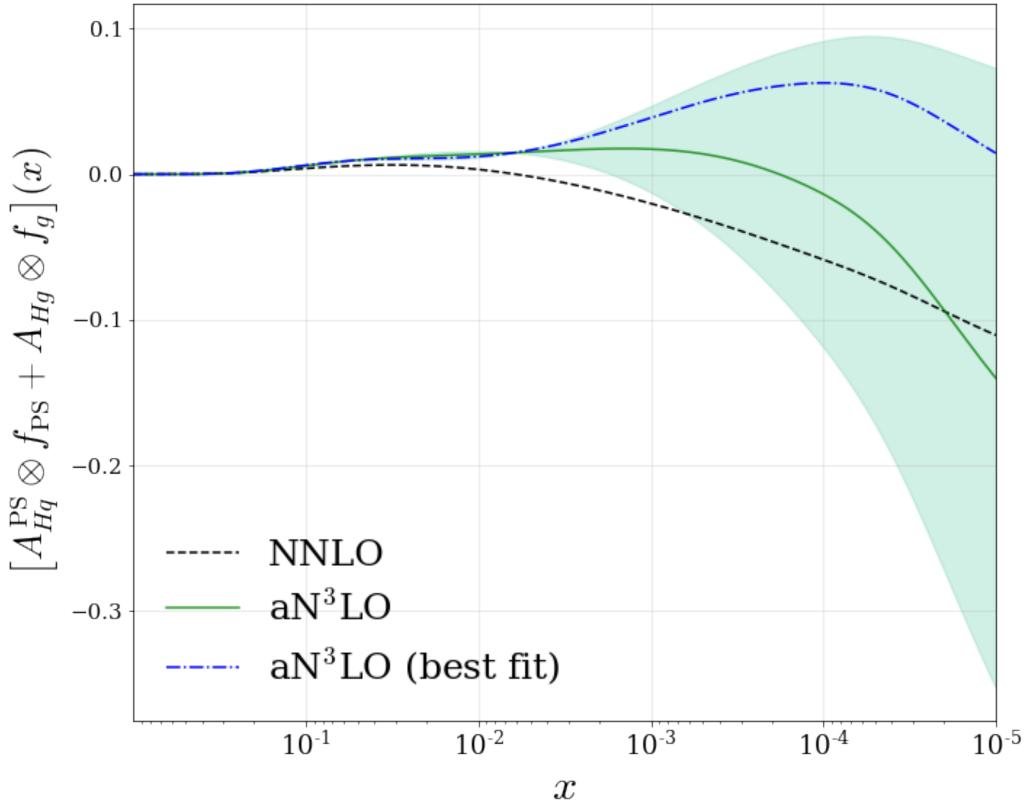




### Transition Matrix Elements up to N<sup>3</sup>LO ...approximately

- Following the same procedure as for the splitting functions.
- $A_{Hg}$  is the **dominant contribution** to the overall form of  $(H + \overline{H})$  shown across.





•  $A_{Hg}$  variation is **comparable** to previous results<sup>[14]</sup>.



## N<sup>3</sup>LO K-factors

- Parameterise the N<sup>3</sup>LO K-factor as a superposition of both NNLO and NLO K-factors.
- Allows the fit to decide on a shape (based on the shapes of preceding orders) and an overall magnitude.
- Center variational parameters  $\hat{a}_1, \hat{a}_2$ about 0, so  $K_{NNLO}$  is the **central** value.

 $K(y) = 1 + \alpha_{\rm s} D(y) + \alpha_{\rm s}^2 E(y) + \alpha_{\rm s}^3 F(y) + \mathcal{O}(\alpha_{\rm s}^4)$ 

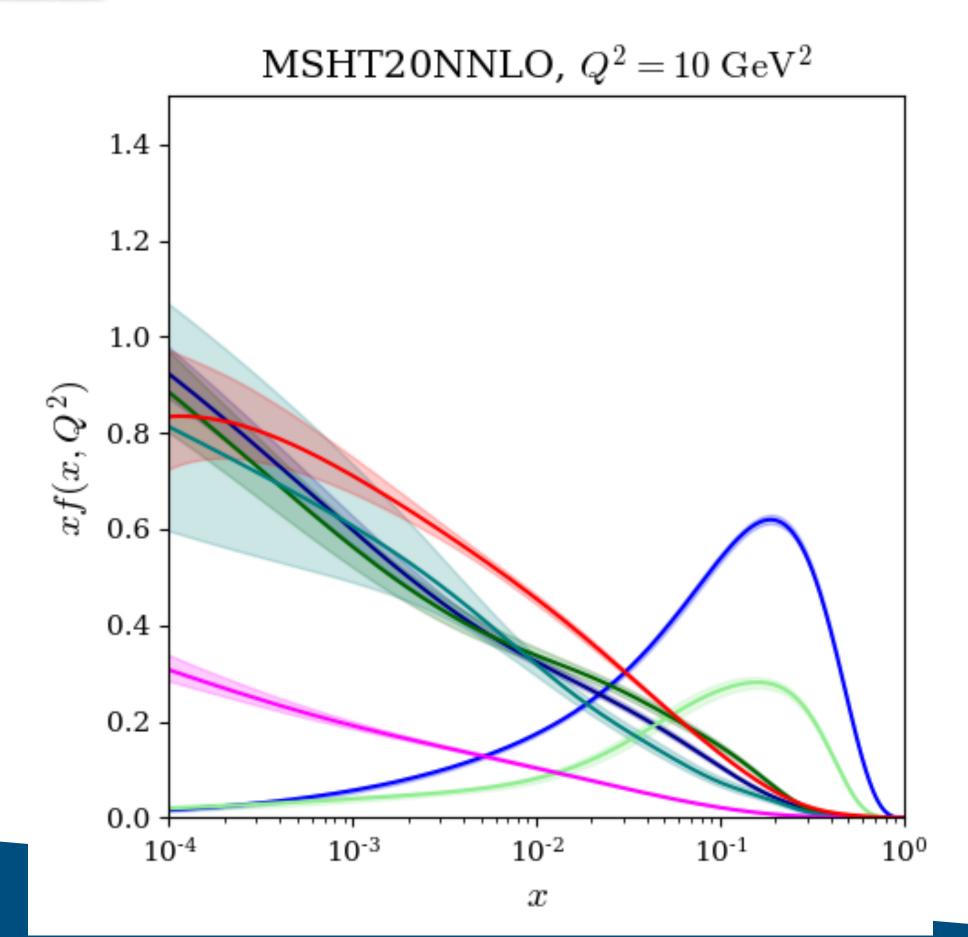
 $K^{\text{N}^{3}\text{LO/LO}} = K^{\text{NNLO/LO}} \left( 1 + \alpha_{s}^{3} \hat{a}_{1} D + \alpha_{s}^{3} \hat{a}_{2} E \right)$ 

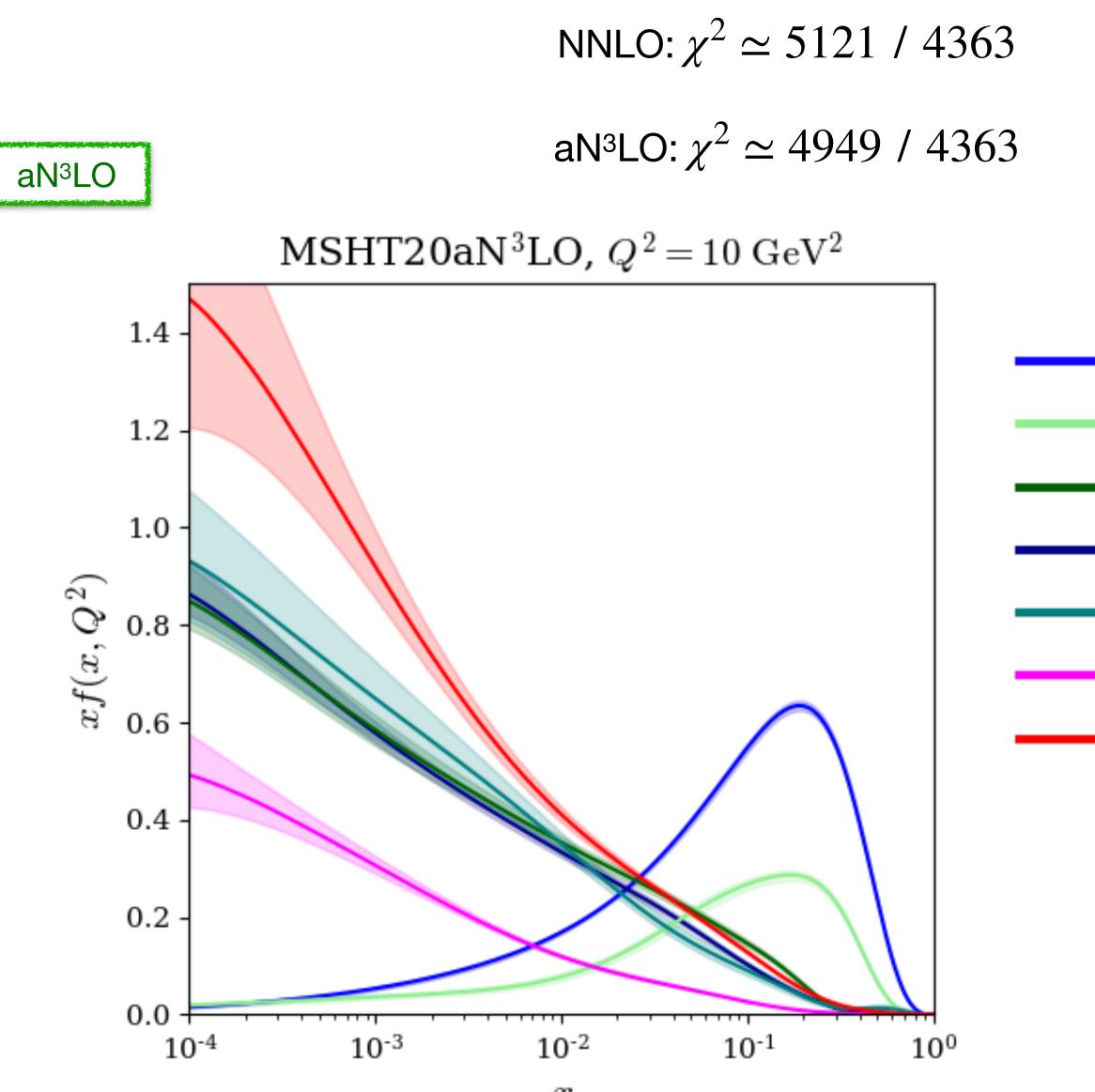
- **Correlated K-factors** for each of the 5 processes: DY, Top, Jets,  $Z p_T \& VB$  Jets and Dimuon (also dijets).
- $\hat{a}_1, \hat{a}_2$  could be included as correlated with **PDF parameters** (incl. other N<sup>3</sup>LO theory parameters) or as **completely decorrelated** from the inclusive DIS process.
  - Ignores some small correlations through DGLAP.



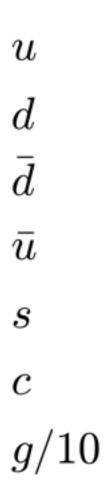


# **NNLO**





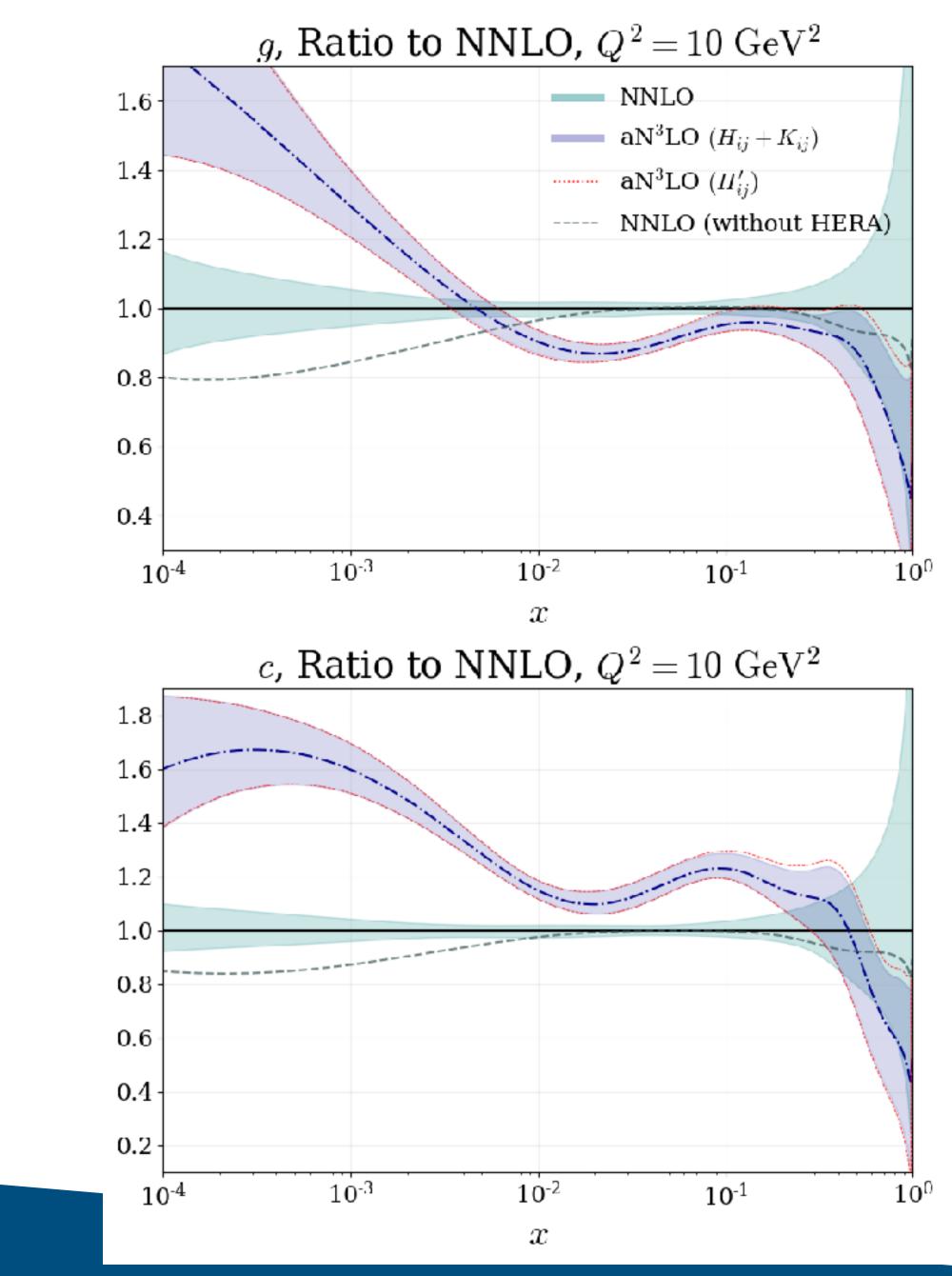
x





## **MSHT aN<sup>3</sup>LO PDFs**

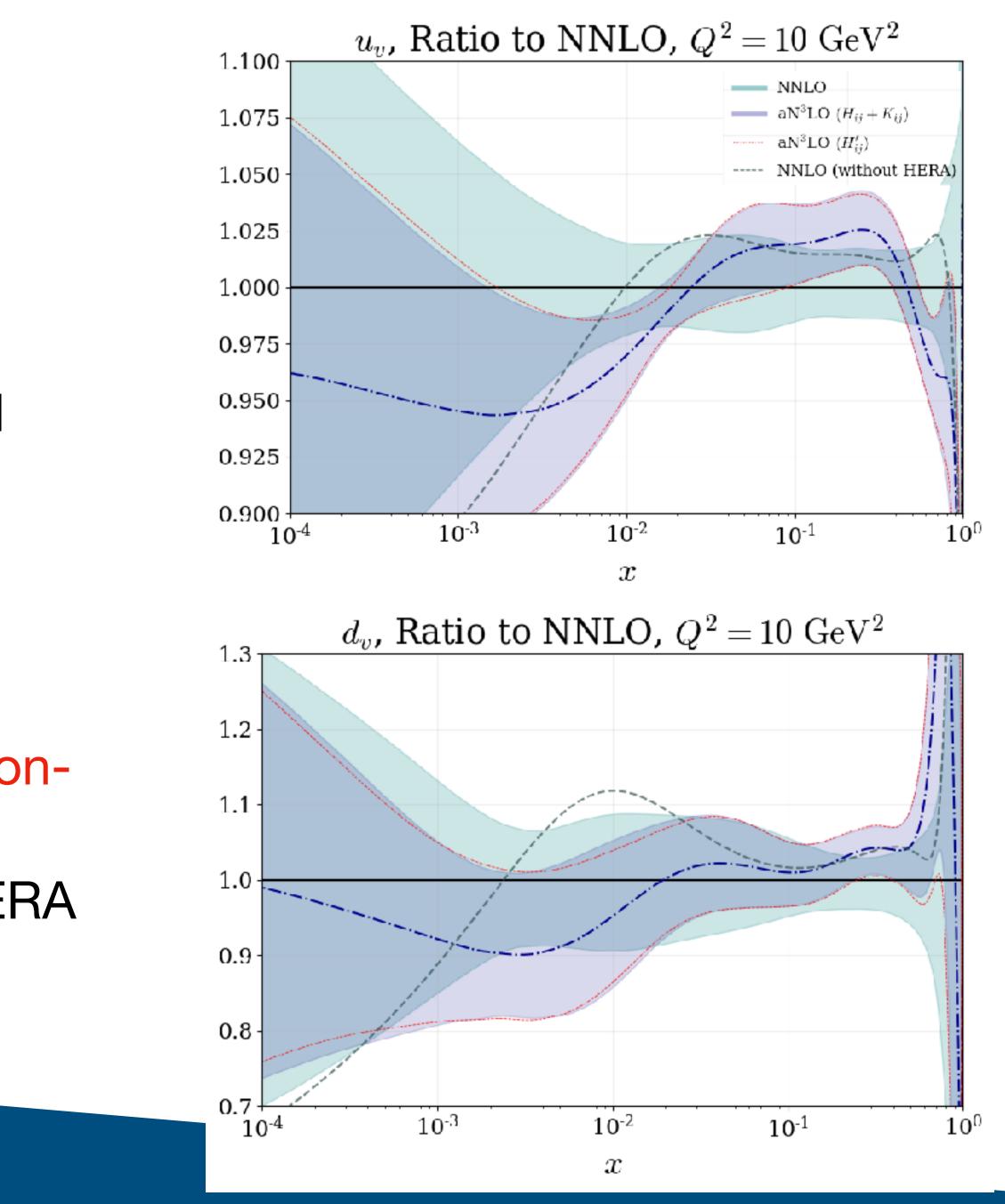
- Gluon is enhanced at small-x due to the large logarithms present at higher orders.
- Charm receives a sizeable contribution from  $A_{H\varrho}^{(3)}$ .
  - $A_{Hg}$  at high-x and the gluon at small-x involved in convolution.
- At high- $Q^2$  there are less drastic effects, however charm remains more similar to CT18 NNLO PDF than MSHT20 NNLO PDF.





## MSHT aN<sup>3</sup>LO PDFs

- Correlated and uncorrelated *K*-factors show consistent uncertainty predictions across all  $(x, Q^2)$ .
- Quarks are reduced at large and small-*x* to accommodate the gluon.
- aN<sup>3</sup>LO follows more closely the NNLO fit to non-HERA datasets at high-*x*, demonstrating a reduction in tension between the small-*x* HERA data and other datasets.

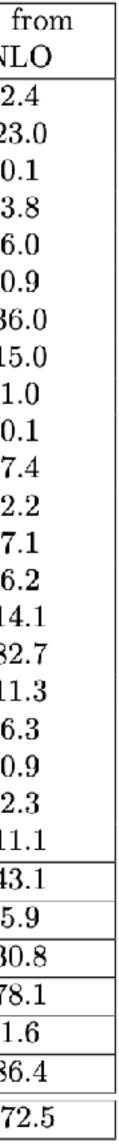






- We see a **reduction** in  $\chi^2$  from NNLO across all datasets ( $\Delta \chi^2 = -172.5$  for 20 extra parameters).
- ATLAS 8 TeV  $Z p_T$ <sup>[15]</sup> sees a huge reduction in  $\chi^2_{NNLO}$ /npts ~ 1.82 to  $\chi^2_{aN^3LO}$ /npts ~ 1.02.
- This is a **similar reduction** found at NNLO when HERA datasets were not included<sup>[17]</sup>.
- In the aN<sup>3</sup>LO fit, we also see a **reduction** in the HERA data  $\chi^2$ .

Dataset	N .	$\sim^2$	$\Delta\chi^2$ f
Dataset	$N_{ m pts}$	$\chi^2$	$\begin{vmatrix} \Delta \chi \\ NNI \end{vmatrix}$
HERA $ep F_2^{charm}$	79	134.7	+2
NMC/BCDMS/SLAC/HERA $F_L$	57	45.5	-23
HERÁ $e^+p$ CC	39	51.8	-0
HERA $e^{-p}$ CC	42	66.3	-3
HERA $e^+p$ NC 820 GeV	75	83.8	-6
HERA $e^-p$ NC 460 GeV	209	247.4	-0
HERA $e^+p$ NC 920 GeV	402	476.7	-36
HERA $e^-p$ NC 575 GeV	259	248.0	-15
HERA $e^-p$ NC 920 GeV	159	243.3	-1
ATLAS $W^+$ , $W^-$ , Z	30	30.0	+0
CMS double diff. Drell-Yan	132	137.1	-7
LHCb 2015 $W, Z$	67	97.2	-2
ATLAS 7 TeV jets	140	214.5	-7
ATLAS 7 TeV high prec. $W, Z$	61	110.5	-6
CMS 7 TeV jets	158	189.8	+14
ATLAS 8 TeV $Z p_T$	104	105.8	-82
m CMS~8~TeV~jets	174	272.6	+11
ATLAS 8 TeV High-mass DY	48	63.4	+6
ATLAS 8 TeV $W + \text{jets}$	30	19.1	+0
ATLAS 8 TeV $W$	22	55.1	-2
CMS 2.76  TeV jet	81	113.9	+11
DY data Total	864	1044.8	-43
Top data Total	71	73.4	-5
Jets data Total	739	972.9	+30
$p_T$ Jets data Total	144	137.1	-78
Dimuon data Total	170	124.6	-1
DIS data Total	2375	2585.2	-86
Total	4363	4948.6	-17







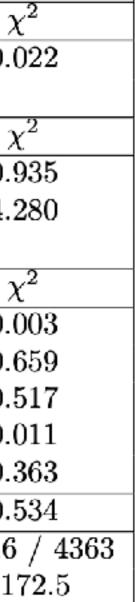
- The overall  $\chi^2$  follows the **general trend** one may expect from perturbation theory.
- Evidence that including aN<sup>3</sup>LO has reduced tensions between small and large-x.
- $\chi^2$  reduction is **mostly due** to new theory, not just from K-factors included in fit.
- Average penalty for included 20 aN<sup>3</sup>LO parameters is ~ 0.53.



	LO	NLO	NNLO	aN <sup>3</sup> LO
$\chi^2/N_{ m pts}$	2.57	1.33	1.17	1.13

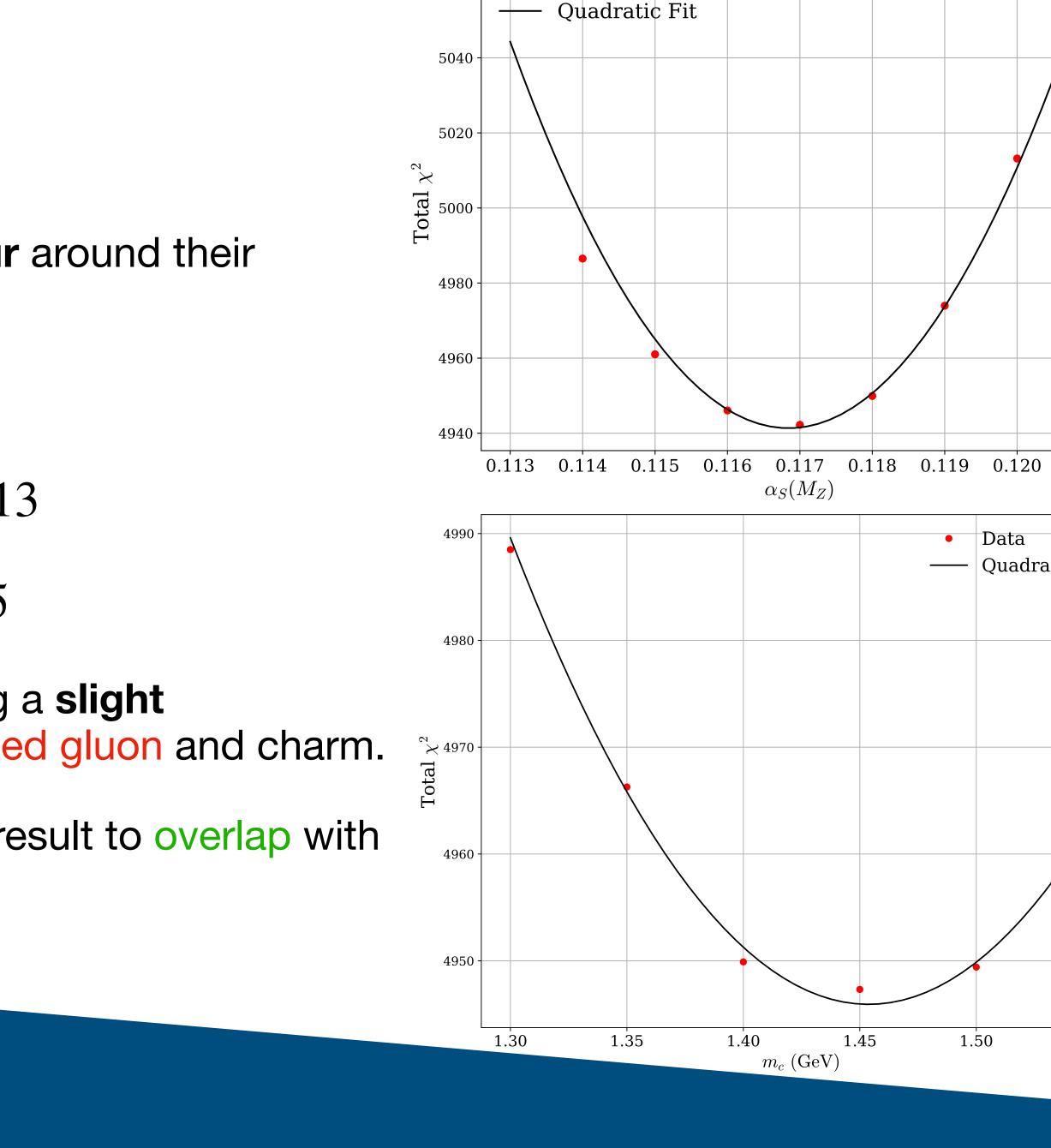
Low- $Q^2$ Coefficient	$\chi^2$	Low- $Q^2$ Coefficient	$\chi^2$
$c_q^{ m NLL}$	0.000	$c_g^{ m NLL}$	0.862
Transition Matrix Elements	$\chi^2$	Transition Matrix Elements	$\chi^2$
$a_{Hg}$	0.526	$a_{qq,H}^{NS}$	0.022
$a_{gg,H}$	1.091		
Splitting Functions	$\chi^2$	Splitting Functions	$\chi^2$
$ ho_{qq}^{NS}$	0.007	$ ho_{gq}$	0.935
$ ho_{qq}^{P_{qq}} ho_{qq}^{PS}$	0.255	$ ho_{gg}$	4.280
$ ho_{qg}$	0.000		
K-factors	$\chi^2$	K-factors	$\chi^2$
DY <sub>NLO</sub>	0.061	DY <sub>NNLO</sub>	0.003
$\mathrm{Top}_{\mathrm{NLO}}$	0.105	Top <sub>NNLO</sub>	0.659
$\rm Jet_{NLO}$	0.063	$\rm Jet_{NNLO}$	0.517
$p_T \mathrm{Jets}_{\mathrm{NLO}}$	0.438	$p_T  m Jets_{NNLO}$	0.011
Dimuon <sub>NLO</sub>	0.481	Dimuon <sub>NNLO</sub>	0.363
N <sup>3</sup> LO Penalty Total	10.7 / 20	Average Penalty	0.534
		Total	4948.6 / 43
		$\Delta \chi^2$ from NNLO	-172.5





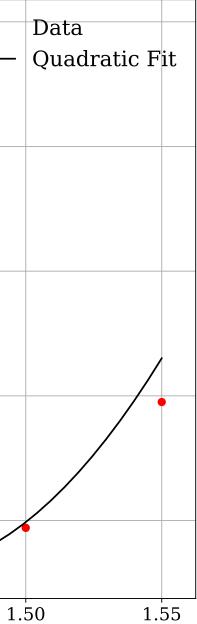


- Both  $\alpha_S(M_Z^2)$  and  $m_c$  show a **quadratic behaviour** around their respective minima.
- Best fit of  $\alpha_{\rm S}(M_{\rm Z}^2)$  is settling  $\simeq 0.1167$ .
  - MSHT20 NNLO:  $\alpha_S(M_Z^2) = 0.1174 \pm 0.0013$
  - MSHT20 NLO:  $\alpha_S(M_Z^2) = 0.1203 \pm 0.0015$
- Both these results suggest that the fit is preferring a **slight** ulletsuppression of the PDFs, particularly the enhanced gluon and charm.
- With a **future full analysis** we expect the aN<sup>3</sup>LO result to overlap with the NNLO world average within uncertainties.



Data







### N<sup>3</sup>LO Drell-Yan Processes

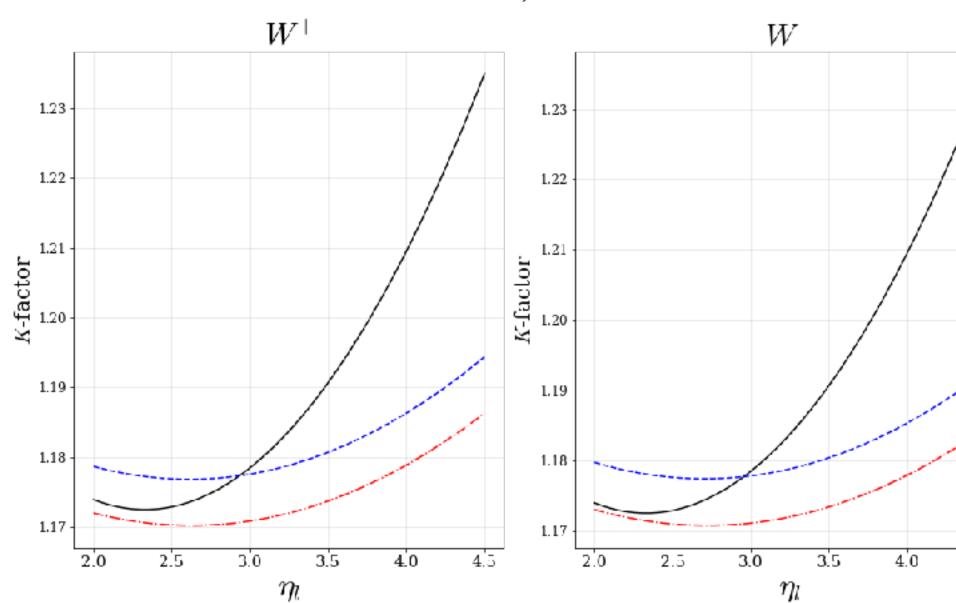
(K-factors up to  $N^{3}LO$ )

• *K*-factors transform the hard cross section between orders.

 $K(y) = 1 + \alpha_s D(y) + \alpha_s^2 E(y) + \alpha_s^3 F(y) + \mathcal{O}(\alpha_s^4)$ 

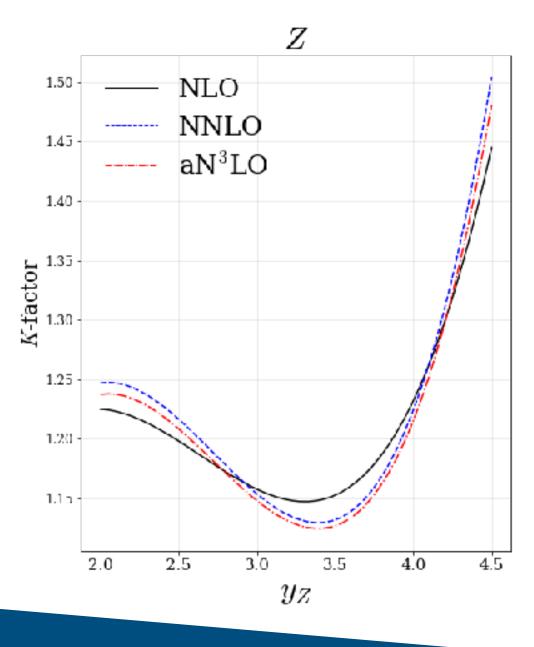
 $K^{\text{aN}^3\text{LO/LO}} = K^{\text{NNLO/LO}} \left(1 + \alpha_s^3 \hat{a}_1 D + \alpha_s^3 \hat{a}_2 E\right)$ 

- Allowed to vary about the NNLO central value.
- Predict a ~1% decrease in the DY K-factors from NNLO.
- In agreement with recent results found using NNLO PDFs with aN<sup>3</sup>LO cross section<sup>[15]</sup>.



LHCb 2015 W, Z dataset results

	NNLO	aN <sup>3</sup> LO
$\chi^2_{ m DY}/N_{ m pts}$	1.26	1.21

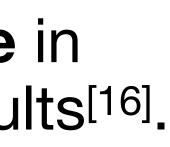


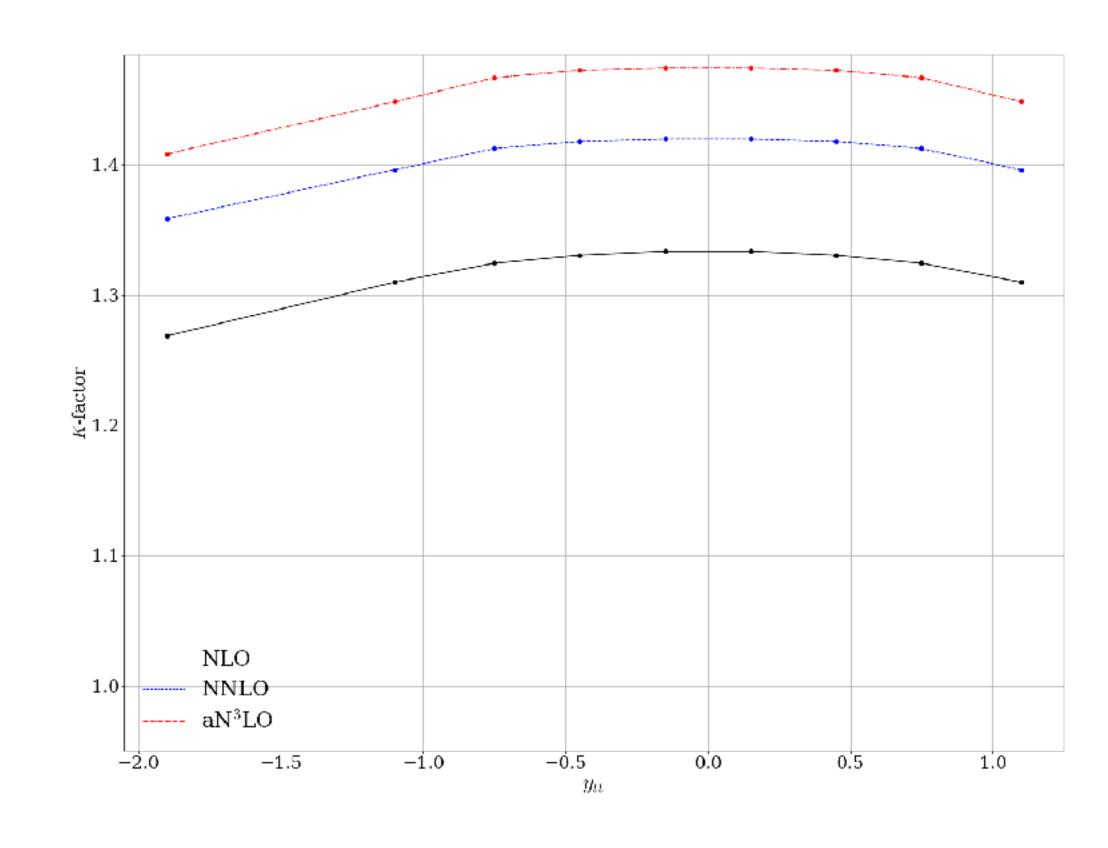




### N<sup>3</sup>LO Top Processes (K-factors up to $N^{3}LO$ )

- Top *K*-factors see an **overall increase** in magnitude, consistent with recent results<sup>[16]</sup>.
- $\chi^2$  results show a **marginally better** fit overall.
- K-factors have successfully accounted for the **theory changes** in the  $F_2$  structure function theory.
- K-factor for CMS 8 TeV single diff.  $t\bar{t}$  shown here.



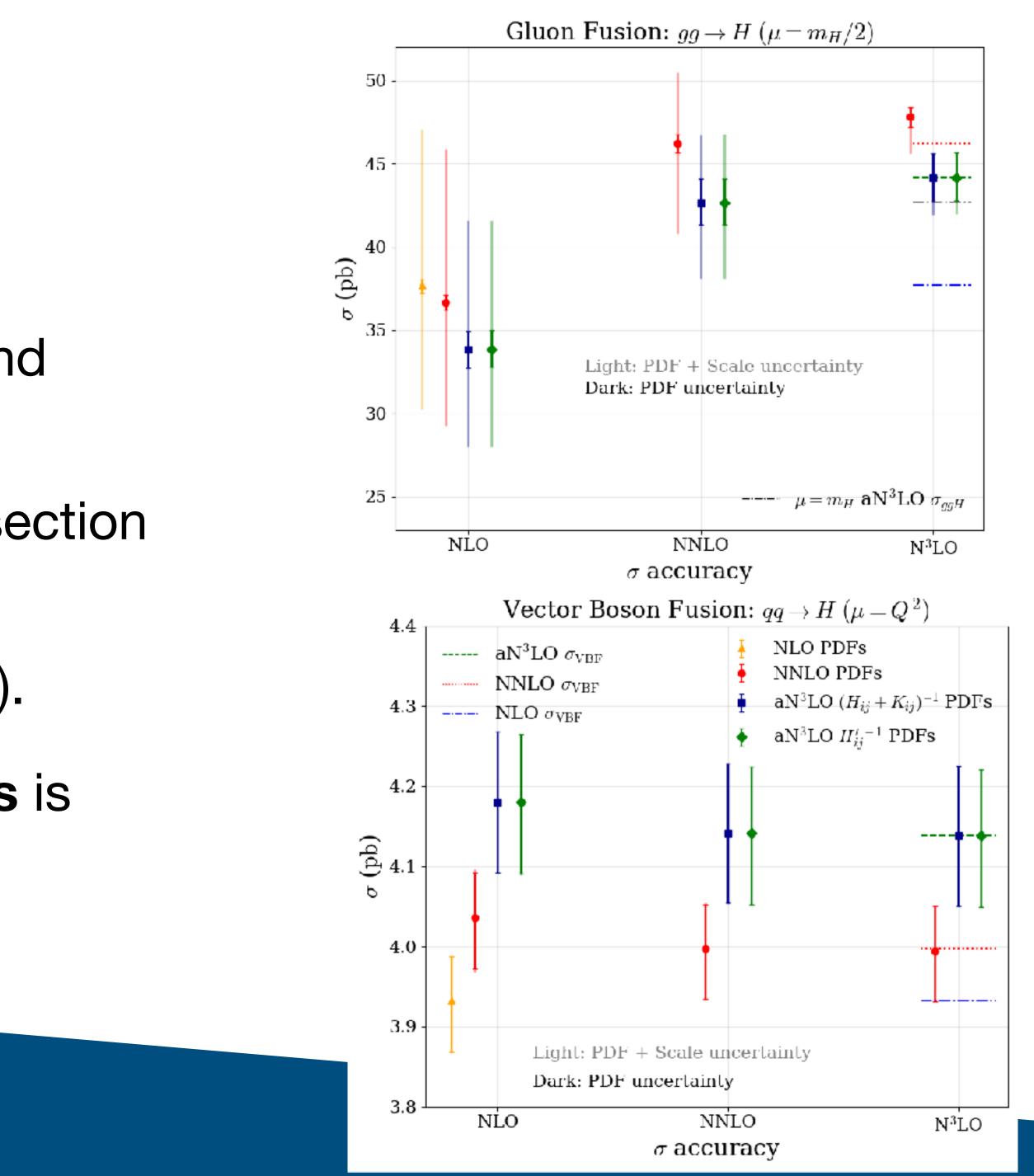


	NNLO	$aN^{3}LO$
$\chi^2_{ m top}/N_{ m pts}$	1.12	1.03



### **Higgs Predictions** For gluon fusion and Vector Boson Fusion (VBF)

- Good agreement between NNLO and aN<sup>3</sup>LO for gluon fusion (top).
- Cancellation between N<sup>3</sup>LO cross section and PDFs not guaranteed.
- Less cancellation for VBF (bottom).
- However variation between orders is smaller for VBF  $\sigma$ .







## **Dijet data at the LHC**

- Fit quality to dijet data at NNLO shows a good **improvement** from jet data.
  - Particularly better fit to  $Z p_T$ .
  - Slightly worse fit to top data.
- Fit quality is also better when fitting to dijet data at aN<sup>3</sup>LO.
  - Fit quality to all **other data** (incl.  $Z p_T$  and top datasets) becomes marginally better  $\Delta \chi^2_{\rm other} \sim -20.$
- Full analysis to follow in the **near future**.



 CMS 8 TeV dijet data is fit with leading colour K-factors - full colour results will follow soon.

### Jets data

	N .	$\chi^2/$	$N_{ m pts}$
	$N_{ m pts}$	NNLO	aN <sup>3</sup> LO
ATLAS 7 TeV jets	140	1.58	1.53
CMS 7 TeV jets	158	1.11	1.20
CMS 8 TeV jets	174	1.50	1.57
Total	472	1.39	1.43

### Dijets data

	N	$\chi^2/$	$N_{ m pts}$		
	$N_{ m pts}$	NNLO	aN <sup>3</sup> LO		
ATLAS 7 TeV dijets	90	1.05	1.14		
CMS 7 TeV dijets	54	1.43	1.40		
CMS 8 TeV dijets	122	1.04	0.84		
Total	266	1.12	1.06		





## Summary

- **Approximate N<sup>3</sup>LO PDFs** are available and we encourage their use.
  - Available as LHAPDF grids at www.hep.ucl.ac.uk/msht/ (see publication for usage instructions).
  - Full information is available in the article JM et. al., <u>2207.04739</u>
- Provide an intuitive and controllable way to include theoretical uncertainties into PDFs.
- Results show good agreement with current N<sup>3</sup>LO predictions.  $\bullet$
- Stay tuned for further developments regarding dijets (and SeaQuest) in an aN<sup>3</sup>LO global fit.



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For an exhaustive list please refer to J. McGowan et. al., (2022) 2207.04739

## Full $\chi^2$ Breakdown

$\begin{array}{c} \text{DOW}  Normal of the state of the $					Dataset	N <sub>pts</sub>	$\chi^2$	$\Delta \chi^2$ from NNLO
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$					$D \emptyset \Pi W \rightarrow \nu e asym.$ [66]	12	29.0	-5.0
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	<b>ceakdown</b>				DØ II $p\bar{p}$ incl. jets [83]	110	113.6	-6.7
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	Ganaovii					30	29.9	-0.0
Lotacet $N_{\rm pb}$ $\lambda$ $N_{\rm N100}$ NUC $N_{\rm N100}$ NUC $24$ $7.5$ $+0.1$ BCDMS $\mu p F_2$ [114]163180.7 $+0.5$ $p_T > 25,30$ GeV [69] $p_T > 25,30$ GeV [60] $p_T$	Detect	21	2	A + 2 (mana		11	7.0	-0.8
BCDMS $\mu p$ F, [114]163180.7+0.5BCDMS $\mu p$ F, [114]151144.0-2.0NMC $\mu p$ F, [113]123119.2-4.9NMC $\mu p$ F, [113]123110.5-6.2SLAC $er p$ F, [116, [117]3732.0-0.0SLAC $er f_{1}$ [116, [117]3821.6-1.4E665 $\mu p$ F, [118]5364.3+4.7E665 $\mu p$ F, [118]5364.3+4.7NuTeV $\nu N$ $r_{1}$ [12]5338.7+0.4NuTeV $\nu N$ $r_{1}$ [13]4234.3+3.6INUC $\mu n / \mu p$ [120]148128.4-2.4UHCb Susce $p D Y$ [60]184208.8-16.2B266 / Nusce $p D Y$ [60]184208.8-16.2CRS $er p$ F, [114]1577-2.6HERA $er p$ Stam<[121]	Dataset	Npts	χ-			24	7.5	+0.1
BCDMS $\mu a$ F, [114]151144.0-2.0LHCb $Z \rightarrow e^+e^-$ 920.6-2.1NMC $\mu p$ F, [113]123119.2-4.9-4.9-4.91012.9+0.4NMC $\mu p$ F, [113]123119.2-4.9-4.9-4.91012.9+0.4SLAC $ep$ F, [116]117]3821.6-1.4-1.4CMS $Z \rightarrow e^+e^-$ [72]3517.3-0.6SLAC $ep$ F, [118]5366.1+1.4-2.4Tevatron, ATLAS, CMS1714.1-0.5E665 $\mu p$ F, [118]5367.1+2.4 $\sigma_t$ [ $\overline{\mathcal{O}}_2$ -[ $\overline{\mathcal{O}}_2$ ]1318.6-0.3NuTeV $vN$ F, [119]4234.3+3.6LHCb 2015 W, Z [ $\overline{\mathcal{O}}_2$ , [ $\overline{\mathcal{S}}_2$ , [ $\overline{\mathcal{S}}_2$ ]6797.1-2.3NMC $\mu n/\mu p$ [120]148128.4-2.4LHCb 2015 W, Z [ $\overline{\mathcal{O}}_2$ , [ $\overline{\mathcal{S}}_2$ , [ $\overline{\mathcal{S}}_2$ ]6797.1-2.3NMC $\mu n/\mu p$ [121]79135.8+3.6LHCb 2015 W, Z [ $\overline{\mathcal{S}}_2$ , [ $\overline{\mathcal{S}}_2$ ]1012.2+3.6NMC $\mu N \rightarrow \mu \mu X$ [113]8669.0+1.3CMS 7 TeV [ets [84]140214.0-7.6CHORUS $vN x_f_3$ [122]2819.5+1.0CMS 7 TeV [ets [81]104106.3-82.2CHORUS $vN x_f_3$ [123]289.5+1.0CMS 7 TeV [ets [82]104106.3-82.2CHORUS $vN x_f_3$ [125]2819.5+1.0CMS 8 TeV sing, diff. ff110.4-6.2HERA $e^+ p$ NC 520 GeV [126]29	BCDMS $uv F_2$ [114]	163	180.7		$p_T > 25,30 \text{ GeV}$ [69]			
$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	,,			1	LHCb $Z \rightarrow e^+e^-$ [70]	9	20.6	-2.1
NMC $\mu d F_{2}$ [113]123106.5-6.2CMS $Z \rightarrow e^{+}e^{-}$ [72]3517.3-0.6SLAC $ed F_{2}$ [116,112]3732.0-0.0<		123		1	LHCb W asym. $p_T > 20$ GeV [71]	10	12.9	+0.4
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	NMC $\mu d F_2$ [115]	123	106.5	-6.2		35	17.3	
$\begin{array}{c c c c c c c c c c c c c c c c c c c $		37	32.0	-0.0				1
$\begin{array}{c c c c c c c c c c c c c c c c c c c $								1
Luco pr $T_2$ [12]S3S0S7.1+1.4NuTeV $vN x_5^3$ [119]4234.3+3.6NMC $\mu n/\mu p$ [120]148128.4-2.4E866 / NuSea $pn$ DY [61]157.7-2.6HERA $ep E^{harm}$ [121]79135.8+3.6NMC/BCDMS/SLAC/HERA5745.5-23.0 $F_L$ [114,113,117,122,12279135.8CCRE $vN \rightarrow \mu\mu X$ [113]8669.0+1.3NuTeV $vN \rightarrow \mu\mu X$ [113]8455.3-3.1CHORUS $vN F_2$ [122]4232.9+2.7CHORUS $vN F_2$ [122]4232.9CHORUS $vN F_2$ [122]4266.3HERA $e^+ p$ CC [126]3951.6HERA $e^- p$ CC [126]4266.3HERA $e^- p$ NC 400 GeV [126]20944.1-5.8HERA $e^- p$ NC 575 GeV [126]259247.9-15.1HERA $e^- p$ NC 902 GeV [126]259247.9-15.1HERA $e^- p$ NC 902 GeV [126]259247.9-15.1CDF II $p\bar{p}$ ind, jets [82]76GB1 II $pri$ ind, jets [82]76GB2 II $pri$ ind, jets [82]76 <td>·</td> <td></td> <td></td> <td>1</td> <td></td> <td></td> <td>11.1</td> <td>0.0</td>	·			1			11.1	0.0
NuTeV $vNxT_5$ [119]4234.3+3.6NMC $un/\mu p$ [120]148128.4-2.4E866 / NuSea $pp$ DY [61]184208.8-16.2E866 / NuSea $pp$ DY [61]157.7-2.6HERA $ep f_{2}^{barm}$ [121]79135.8+3.6NMC/BCDM5/SLAC/HERA5745.5-23.0 $F_L$ [114,115,117,122-124]74.55-23.0CCFR $vN \rightarrow \mu\mu X$ [113]8669.0+1.3NuTeV $vN \rightarrow \mu\mu X$ [113]8455.3-3.1CHORUS $vN xF_B$ [125]4232.9+2.7CHORUS $vN xF_B$ [125]2819.5+1.0HERA $e^+p$ CC [126]4266.3-3.8HERA $e^+p$ NC 200 GeV [126]7584.0-5.8HERA $e^+p$ NC 200 GeV [126]209247.1-1.2HERA $e^+p$ NC 200 GeV [126]159243.4-1.0CDF II $p\bar{p}$ ind, jets [83]7668.7+8.3DØ II Z rap. [63]2839.6+2.5CDF II Z rap. [63]2839.6+2.5DØ II Z rap. [63]2839.6+2.5DØ II Z rap. [63]2839.6+2.5DØ II Z rap. [64]1016.7-0.6CDF II Z rap. [64]1016.7-0.6DØ II Z rap. [64]1016.7-0.6DØ II Z rap. [64]1016.7-0.6DØ II Z rap. [64]1016.7-0.6DØ II Z rap. [64]1016.7-0.6CDF II Z rap. [64]28-0.6						132	136.8	_77
$\begin{array}{c c c c c c c c c c c c c c c c c c c $				1				
E866 / NuSea $pp$ DY [61]184208.8-16.2E866 / NuSea $pd/pp$ DY [61]157.7-2.6HERA $ep p_2^{barm}$ [12]79135.8+3.6NMC/BCDMS/SLAC/HERA5745.5-23.0 $F_1$ [114,113,117,122-124]-0.6CMS 7 TeV $y + c$ [88]10CCFR $vN \rightarrow \mu\muX$ [113]8669.0+1.3NuTeV $vN \rightarrow \mu\muX$ [113]8669.0+1.3CCFR $vN \rightarrow \mu\muX$ [113]8669.0+1.3CHORUS $vN \neq_p$ [125]4232.9+2.7CHORUS $vN x_{F_2}$ [122]2819.5+1.0HERA $e^+p$ CC [126]3951.6-0.4HERA $e^-p$ NC 460 GeV [126]7584.0HERA $e^-p$ NC 20 GeV [126]7584.0HERA $e^-p$ NC 525 GeV [126]209247.1HERA $e^-p$ NC 520 GeV [126]209247.1HERA $e^-p$ NC 200 GeV [126]209247.1HERA $e^-p$ NC 200 GeV [126]209247.1HERA $e^-p$ NC 525 GeV [126]259247.9HERA $e^-p$ NC 520 GeV [126]259247.9HERA $e^-p$ NC 920 GeV [126]159243.4HERA $e^-p$ NC 525 GeV [126]259247.9HERA $e^-p$ NC 526 GeV [126]2816.8HOP II $p\bar{p}$ ind. jets [82]7668.7HERA $e^-p$ NC 920 GeV [126]2816.8HOP II $z$ rap. [62]2816.8HOP II $z$ rap. [62]2816.8HOP II $z$ rap. [62]2816.8HOP II $z$ rap. [62]28	Nullev $vN xF_3$ [119]			1				
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HERA $ep$ $p_{2}^{\text{charm}}$ [12]79135.8+3.6HIM MODELNMC/BCDMS/SLAC/HERA5745.5-23.0CMS 7 TeV $\psi$ + $c$ [88]1011.0+3.6 $F_L$ [114,[115,[117,[122-I22]]8669.0+1.3CMS 7 TeV $\psi$ + $c$ [88]158189.9+14.1NuTeV $vN \rightarrow \mu\mu\chi$ [113]8669.0+1.3DØ W asym. [7]148.8-3.3CHORUS $vN \rightarrow p_{L}\chi$ [125]4232.9+2.7ATLAS 8 TeV Z $p_T$ [87]104106.3-82.2CHORUS $vN x_{T_3}$ [125]2819.5+1.0CMS 8 TeV Z $p_T$ [87]104106.3-82.2CHORUS $vN x_{T_3}$ [125]2819.5+1.0CMS 8 TeV Z $p_T$ [87]104106.3-82.2CHORUS $vN x_{T_3}$ [125]2819.5+1.0CMS 8 TeV Z $p_T$ [87]104106.3-82.2CHORUS $vN x_{T_3}$ [126]29247.1-1.2CMS 8 TeV sing. diff. $t\bar{t}$ [110]2525.0-0.7HERA $e^+p$ NC 820 GeV [126]209247.1-1.2ATLAS 8 TeV sing. diff. $t\bar{t}$ 52.2-1.2HERA $e^-p$ NC 920 GeV [126]159247.9-15.1ATLAS 8 TeV sing. diff. $t\bar{t}$ 11.21523.8+1.3CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3ATLAS 8 TeV double diff. $t\bar{t}$ [112]1523.8+1.3DØ II Z rap. [63]2839.6+2.5CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9DØ II Z rap. [64]2016.7-0.6 <td></td> <td></td> <td></td> <td>1</td> <td></td> <td></td> <td></td> <td></td>				1				
NMC/BCDMS/SLAC/HERA       57       45.5 $-23.0$ ATLAS 7 TeV high prec. W, Z [59]       10 $12.2$ $+3.0$ $F_L$ [114,[115,117,122-124]       CCFR $vN \rightarrow \mu\mu X$ [113]       86       69.0 $+1.3$ ATLAS 7 TeV high prec. W, Z [59]       61 $110.4$ $-6.2$ CCFR $vN \rightarrow \mu\mu X$ [113]       84       55.3 $-3.1$ DØ W asym. [77]       14       8.8 $-3.3$ CHORUS $vN vF_2$ [125]       42 $32.9$ $+2.7$ ATLAS 8 TeV $Z p_T$ [87]       104 $106.3$ $-82.2$ CHORUS $vN vF_2$ [126]       28 $19.5$ $+1.0$ CMS 8 TeV jets [83] $174$ $271.9$ $+10.6$ HERA $e^+p$ NC 820 GeV [126]       42 $66.3$ $-3.8$ ATLAS 8 TeV sing. diff. $t\overline{t}$ [110]       25 $25.0$ $-0.7$ HERA $e^-p$ NC 820 GeV [126]       209 $247.1$ $-1.2$ $dilep.$ [111] $TLAS 8 TeV sing. diff. t\overline{t}$ $5$ $2.2$ $-1.2$ HERA $e^-p$ NC 920 GeV [126]       209 $247.9$ $-15.1$ $TLAS 8 TeV High-mass DY [78]       48 63.8 +6.6         HERA e^-p NC 920 GeV [126]       259       243.4 -1.0 CMS 8 TeV doub$								
$F_L$ [114,113,117,1122-124] CCFR $vN \to \mu\mu X$ [113]8669.0+1.3AILAS / TeV high prec. $W, Z$ [59]61110.4-6.2 $CCFR vN \to \mu\mu X$ [113]8455.3-3.1 $CMS 7 TeV$ jets [81]158189.9+14.1 $NuTeV vN \to \mu\mu X$ [113]8455.3-3.1 $DØ$ W asym. [77]148.8-3.3 $CHORUS vN F_2$ [125]4232.9+2.7 $ATLAS 8 TeV Z p_T$ [87]104106.3-82.2 $CHORUS vN x_5$ [125]2819.5+1.0 $CMS 8 TeV jets$ [83]174271.9+10.6 $HERA e^+ p CC$ [126]3951.6-0.4 $ATLAS 8 TeV sing. diff. tf$ [110]2525.0-0.7 $HERA e^+ p NC$ 820 GeV [126]209247.1-1.2 $ATLAS 8 TeV sing. diff. tf$ 52.2-1.2 $HERA e^- p NC 460 GeV [126]209247.1-1.2ATLAS 8 TeV High-mass DY [78]4863.8+6.6HERA e^- p NC 920 GeV [126]159243.4-1.0CMS 8 TeV double diff. tf112]1523.8+1.3CDF II p\bar{p} incl. jets [82]7668.7+8.3ATLAS 8 TeV W [79]2254.8-2.6DØ II Z rap. [62]2816.8+0.5CMS 8 TeV sing. diff. tf[112]1523.8+1.3CDF II IZ rap. [63]2839.6+2.5CMS 8 TeV sing. diff. tf[12]98.3-4.9DOI II W \to w asym[64]16.7-0.6-0.6-0.6-0.6-0.6-0.6$	• ∠ •							
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NuTeV $vN \rightarrow \mu\mu X$ [113]8455.3-3.1DØ W asym. [77]148.8-3.3CHORUS $vN F_2$ [125]4232.9+2.7ATLAS 8 TeV $Z p_T$ [87]104106.3-82.2CHORUS $vN xF_3$ [125]2819.5+1.0CMS 8 TeV jets [85]174271.9+10.6HERA $e^+p$ CC [126]3951.6-0.4ATLAS 8 TeV sing. diff. $t\bar{t}$ [110]2525.0-0.7HERA $e^-p$ CC [126]4266.3-3.8-3.8ATLAS 8 TeV sing. diff. $t\bar{t}$ 52.2-1.2HERA $e^-p$ NC 460 GeV [126]209247.1-1.2-1.2dilep. [111]4863.8+6.6HERA $e^-p$ NC 920 GeV [126]402476.2-36.5ATLAS 8 TeV High-mass DY [78]4863.8+6.6HERA $e^-p$ NC 920 GeV [126]159243.4-1.0CMS 8 TeV double diff. $t\bar{t}$ [112]1523.8+1.3CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3ATLAS 8 TeV W [79]2254.8-2.6DØ II Z rap. [63]2839.6+2.5CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9DO II W $w \rightarrow ww$ seem [64]1016.7-0.6CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9		86	69.0	+1.3	CMS 7 TeV jets [81]	158	189.9	+14.1
CHORUS $vN F_2$ [125]4232.9+2.7ATLAS 8 TeV $Z p_T$ [87]104106.3-82.2CHORUS $vN xF_3$ [125]2819.5+1.0CMS 8 TeV jets [85]174271.9+10.6HERA $e^+ p$ CC [126]3951.6-0.4ATLAS 8 TeV sing. diff. $t\bar{t}$ [110]2525.0-0.7HERA $e^- p$ CC [126]4266.3-3.8ATLAS 8 TeV sing. diff. $t\bar{t}$ 52.2-1.2HERA $e^- p$ NC 820 GeV [126]209247.1-1.2ATLAS 8 TeV sing. diff. $t\bar{t}$ 52.2-1.2HERA $e^- p$ NC 920 GeV [126]209247.1-1.2ATLAS 8 TeV ligh-mass DY [78]4863.8+6.6HERA $e^- p$ NC 575 GeV [126]259247.9-15.1ATLAS 8 TeV W + jets [89]3019.2+1.1HERA $e^- p$ NC 920 GeV [126]159243.4-1.0CMS 8 TeV double diff. $t\bar{t}$ [112]1523.8+1.3CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3ATLAS 8 TeV W [79]2254.8-2.6DØ II Z rap. [62]2830.6+2.5CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9DØ II W $w$ W W [64]1016.7-0.6CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9		84			DØ W asym. [77]	14	8.8	-3.3
HERA $e^+ p$ CC [126]3951.6 $-0.4$ ATLAS 8 TeV sing. diff. $t\bar{t}$ [110]2525.0 $-0.7$ HERA $e^- p$ NC 820 GeV [126]4266.3 $-3.8$ ATLAS 8 TeV sing. diff. $t\bar{t}$ 52.2 $-1.2$ HERA $e^- p$ NC 820 GeV [126]209247.1 $-1.2$ ATLAS 8 TeV sing. diff. $t\bar{t}$ 52.2 $-1.2$ HERA $e^- p$ NC 920 GeV [126]209247.1 $-1.2$ ATLAS 8 TeV High-mass DY [78]4863.8 $+6.6$ HERA $e^- p$ NC 920 GeV [126]259247.9 $-15.1$ ATLAS 8 TeV High-mass DY [78]4863.8 $+1.1$ HERA $e^- p$ NC 920 GeV [126]159243.4 $-1.0$ CMS 8 TeV double diff. $t\bar{t}$ [112]1523.8 $+1.3$ CDF II $p\bar{p}$ incl. jets [82]7668.7 $+8.3$ ATLAS 8 TeV W [79]2254.8 $-2.6$ DØ II Z rap. [62]2816.8 $+0.5$ CMS 2.76 TeV jet [86]81113.7 $+10.8$ CDF II Z rap. [63]2839.6 $+2.5$ CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3 $-4.9$	,,	42	32.9	+2.7	ATLAS 8 TeV Z $p_T$ [87]	104	106.3	-82.2
HERA $e^{-p}$ CC [126]4266.3 $-3.8$ ATLAS & TeV sing. diff. $t\bar{t}$ 125126136HERA $e^+p$ NC 820 GeV [126]7584.0 $-5.8$ ATLAS & TeV sing. diff. $t\bar{t}$ 52.2 $-1.2$ HERA $e^-p$ NC 460 GeV [126]209247.1 $-1.2$ $dilep. [11]$ ATLAS & TeV sing. diff. $t\bar{t}$ 52.2 $-1.2$ HERA $e^-p$ NC 920 GeV [126]402476.2 $-36.5$ $ATLAS & TeV High-mass DY [78]4863.8+6.6HERA e^-p NC 575 GeV [126]259247.9-15.1ATLAS & TeV W + jets [89]3019.2+1.1HERA e^-p NC 920 GeV [126]159243.4-1.0CMS & TeV double diff. t\bar{t}112]1523.8+1.3CDF II p\bar{p} incl. jets [82]7668.7+8.3ATLAS & TeV W [79]2254.8-2.6DØ II Z rap. [62]2816.8+0.5CMS 2.76 TeV jet [86]81113.7+10.8CDF II Z rap. [63]2839.6+2.5CMS 8 TeV sing. diff. t\bar{t}91]98.3-4.9$	CHORUS $\nu N x F_3$ [125]	28	19.5	+1.0	CMS 8 TeV jets [85]	174	271.9	+10.6
HERA $e^-p$ CC [126]4266.3 $-3.8$ HERA $e^+p$ NC 820 GeV [126]7584.0 $-5.8$ HERA $e^-p$ NC 460 GeV [126]209247.1 $-1.2$ HERA $e^-p$ NC 920 GeV [126]402476.2 $-36.5$ HERA $e^-p$ NC 575 GeV [126]259247.9 $-15.1$ HERA $e^-p$ NC 920 GeV [126]159243.4 $-1.0$ CDF II $p\bar{p}$ incl. jets [82]7668.7 $+8.3$ CDF II $p\bar{p}$ incl. jets [82]7668.7 $+8.3$ CDF II Z rap. [62]2816.8 $+0.5$ CDF II Z rap. [63]2839.6 $+2.5$ CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3PO(11 W) $\rightarrow$ tw asym1016.7PO(11 W) $\rightarrow$ tw asym1016.7		39	51.6	-0.4	ATLAS 8 TeV sing. diff. tt [110]	25	25.0	-0.7
HERA $e^+p$ NC 820 GeV [126]7584.0-5.8dilep. [111]HERA $e^-p$ NC 460 GeV [126]209247.1-1.2ATLAS 8 TeV High-mass DY [78]4863.8+6.6HERA $e^+p$ NC 920 GeV [126]402476.2-36.5ATLAS 8 TeV High-mass DY [78]4863.8+6.6HERA $e^-p$ NC 575 GeV [126]259247.9-15.1ATLAS 8 TeV W + jets [89]3019.2+1.1HERA $e^-p$ NC 920 GeV [126]159243.4-1.0CMS 8 TeV double diff. $t\bar{t}$ [112]1523.8+1.3CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3ATLAS 8 TeV W [79]2254.8-2.6DØ II Z rap. [62]2816.8+0.5CMS 2.76 TeV jet [86]81113.7+10.8CDF II Z rap. [63]2839.6+2.5CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9DØ II W $\rightarrow$ W asym1016.7-0.6CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9						5	2.2	-1.2
HERA $e^{-p}$ NC 480 GeV [126]209247.1-1.2HERA $e^{+p}$ NC 920 GeV [126]402476.2-36.5HERA $e^{-p}$ NC 575 GeV [126]259247.9-15.1HERA $e^{-p}$ NC 920 GeV [126]159243.4-1.0HERA $e^{-p}$ NC 920 GeV [126]159243.4-1.0CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3CDF II $z$ rap. [62]2816.8+0.5CDF II Z rap. [63]2839.6+2.5DØ II Z rap. [63]2839.6+2.5DØ II W $\rightarrow$ 1W asyme [64]1016.7DØ II W $\rightarrow$ 1W asyme [64]1016.7DØ II W $\rightarrow$ 1W asyme [64]10CMS 8 TeV sing. diff. $t\bar{t}$ [91]98 TeV sing. diff. $t\bar{t}$ [91]98 TeV sing. diff. $t\bar{t}$ [91]								
HERA $e^- p$ NC 920 GeV [126]259247.9-15.1ATLAS 8 TeV W + jets [89]3019.2+1.1HERA $e^- p$ NC 920 GeV [126]159243.4-1.0CMS 8 TeV double diff. $t\bar{t}$ [112]1523.8+1.3CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3ATLAS 8 TeV W [79]2254.8-2.6DØ II Z rap. [62]2816.8+0.5CMS 2.76 TeV jet [86]81113.7+10.8CDF II Z rap. [63]2839.6+2.5CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9					I	48	63.8	+6.6
HERA $e^-p$ NC 920 GeV [126]159243.4-1.0CMS 8 TeV double diff. $t\bar{t}$ [112]1523.8+1.3CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3ATLAS 8 TeV W [79]2254.8-2.6DØ II Z rap. [62]2816.8+0.5CMS 2.76 TeV jet [86]81113.7+10.8CDF II Z rap. [63]2839.6+2.5CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9DØ II W $\rightarrow tW$ asym1016.7-0.6CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9								1
CDF II $p\bar{p}$ incl. jets [82]7668.7+8.3ATLAS 8 TeV W [79]2254.8-2.6DØ II Z rap. [62]2816.8+0.5CMS 2.76 TeV jet [86]81113.7+10.8CDF II Z rap. [63]2839.6+2.5CMS 8 TeV sing. diff. $t\bar{t}$ [91]98.3-4.9DØ II W $\Rightarrow tw asym [64]1016.7-0.6CMS 8 TeV sing. diff. t\bar{t} [91]98.3-4.9$	· · · · · · · · · · · · · · · · · · ·							
DØ II Z rap. [62]       28       16.8 $+0.5$ CMS 2.76 TeV jet [86]       81       113.7 $+10.8$ CDF II Z rap. [63]       28       39.6 $+2.5$ CMS 8 TeV sing. diff. $t\bar{t}$ [91]       9       8.3 $-4.9$ DØ II W $\Rightarrow$ ww asym       16.7 $-0.6$ CMS 8 TeV sing. diff. $t\bar{t}$ [91]       9       8.3 $-4.9$								-
CDF II Z rap. [63]       28       39.6       +2.5       CMS 8 TeV sing. diff. $t\bar{t}$ [91]       9       8.3       -4.9         DØ II W $\rightarrow t''$ asym       10       16.7       -0.6       -0.	· · · · ·							
$DO[W] \rightarrow ww [64]$ 10 167 -06 -06 -0.5 -0.5 -0.5 -0.5 -0.5 -0.5 -0.5 -0.5				1				1
	· · · · · · · · · · · · · · · · · · ·					-		
CDF II W asym. [65] 13 20.1 +1.1 ATLAS 8 lev double diff. Z [80] 59 81.5 -4.1					ATLAS 8 TeV double diff. Z [80]	59	81.5	-4.1





### **Comparison with/without** *K***-factors**

Dimuon Dataset	$\chi^2$	$\Delta \chi^2$	$\Delta \chi^2$ from NNLO
		from NNLO	(NNLO K-factors)
CCFR $\nu N \rightarrow \mu \mu X$ [113]	69.0 / 86	+1.3	+2.6
NuTeV $\nu N \rightarrow \mu \mu X$ [113]	55.3 / 84	-3.1	-3.1
Total	124.3 / 170	-1.8	-0.5

DY Dataset	$\chi^2$	$\Delta \chi^2$	$\Delta \chi^2$ from NNLO
		from NNLO	(NNLO K-factors)
E866 / NuSea pp DY [60]	208.8 / 184	-16.2	-11.6
E866 / NuSea pd/pp DY [61]	7.7 / 15	-2.6	-2.9
DØ II Z rap. [62]	16.8 / 28	+0.5	+0.3
CDF II Z rap. [63]	39.6 / 28	+2.5	+1.3
DØ II $W \rightarrow \nu \mu$ asym. [64]	16.7 / 10	-0.6	-0.5
CDF II W asym. [65]	20.1 / 13	+1.1	+0.8
DØ II $W \rightarrow ve$ asym. [66]	29.0 / 12	-5.0	-5.3
ATLAS $W^+$ , $W^-$ , Z [67]	29.9 / 30	-0.0	+0.3
CMS W asym. $p_T > 35 \text{ GeV}$ [68]	7.0 / 11	-0.8	-0.6
CMS W asym. $p_T > 25,30$ GeV [69]	7.5 / 24	+0.1	-0.1
LHCb $Z \rightarrow e^+e^-$ [70]	20.6 / 9	-2.1	-1.6
LHCb W asym. $p_T > 20$ GeV [71]	12.9 / 10	+0.4	+1.0
$CMS Z \rightarrow e^+e^-$ [72]	17.3 / 35	-0.6	-0.6
ATLAS High-mass Drell-Yan [73]	18.6 / 13	-0.3	-1.1
CMS double diff. Drell-Yan [74]	136.8 / 132	-7.7	+11.9
LHCb 2015 W, Z [57, 58]	97.1 / 67	-2.3	-2.8
LHCb 8TeV $Z \rightarrow ee$ [75]	26.9 / 17	+0.7	-0.2
CMS 8 TeV W [76]	12.1 / 22	-0.6	+0.2
ATLAS 7 TeV high prec. $W_{J}Z$ [59]	110.4 / 61	-6.2	-18.7
DØ W asym. [77]	8.8 / 14	-3.3	-1.8
ATLAS 8 TeV High-mass DY [78]	63.8 / 48	+6.6	+2.8
ATLAS 8 TeV W [79]	54.8 / 22	-2.6	-1.1
ATLAS 8 TeV double diff. Z [80]	81.5 / 59	-4.1	-1.9
Total	1044.6 / 864	-43.2	-32.1

Jets Dataset	$\chi^2$	$\Delta \chi^2$	$\Delta \chi^2$ from NNLO
		from NNLO	(NNLO K-factors)
CDF II $p\bar{p}$ incl. jets [82]	68.7 / 76	+8.3	+0.6
DØ II $p\bar{p}$ incl. jets [83]	113.6 / 110	-6.7	-3.5
ATLAS 7 TeV jets [84]	214.0 / 140	-7.6	+2.4
CMS 7 TeV jets [81]	189.9 / 158	+14.1	+14.5
CMS 8 TeV jets [85]	271.9 / 174	+10.6	+22.9
CMS 2.76 TeV jet [86]	1 <b>13.7 / 8</b> 1	+10.8	+13.5
Total	971.7 / 739	+29.6	+50.3



DIS Dataset	$\chi^2$	$\Delta \chi^2$	$\Delta \chi^2$ from NNLO
		from NNLO	(NNLO K-factors)
BCDMS $\mu p F_2$ [114]	180.7 / 163	+0.5	+0.1
BCDMS $\mu d F_2$ [114]	144.0 / 151	-2.0	-1.1
NMC $\mu p F_2$ [115]	119.2 / 123	-4.9	-7.0
NMC µd F <sub>2</sub> [115]	106.5 / 123	-6.2	-10.2
SLAC <i>ep F</i> <sub>2</sub> [116, 117]	32.0 / 37	-0.0	+0.5
SLAC ed F <sub>2</sub> [116, 117]	21.6 / 38	-1.4	-1.4
E665 µp F <sub>2</sub> [118]	64.3 / 53	+4.7	+5.7
E665 $\mu d F_2$ [118]	67.1 / 53	+2.4	+2.8
NuTeV $\nu N F_2$ [119]	38.7 / 53	+0.4	+1.7
NuTeV $\nu N x F_3$ [119]	34.3 / 42	+3.6	+1.9
NMC μn/μp [120]	128.4 / 148	-2.4	-2.6
HERA $ep F_2^{charm}$ [121]	135.8 / 79	+3.6	+9.1
NMC/BCDMS/SLAC/HERA	45.5 / 57	-23.0	-23.3
$F_L$ [114, 115, 117, 122–124]			
CHORUS $\nu N F_2$ [125]	32.9 / 42	+2.7	+3.0
CHORUS $\nu N x F_3$ [125]	19.5 / 28	+1.0	+1.1
HERA <i>e</i> <sup>+</sup> <i>p</i> CC [126]	51.6 / 39	-0.4	+0.3
HERA $e^-p$ CC [126]	66.3 / 42	-3.8	-3.0
HERA e <sup>+</sup> p NC 820 GeV [126]	84.0 / 75	-5.8	-5.5
HERA <i>e<sup>-</sup>p</i> NC 460 GeV [126]	247.1 / 209	-1.2	-0.4
HERA e <sup>+</sup> p NC 920 GeV [126]	476.2 / 402	-36.5	-33.3
HERA <i>e<sup>-</sup>p</i> NC 575 GeV [126]	247.9 / 259	-15.1	-14.4
HERA <i>e<sup>-</sup>p</i> NC 920 GeV [126]	243.4 / 159	-1.0	-1.0
Total	2587.0 / 2375	-84.7	-76.8

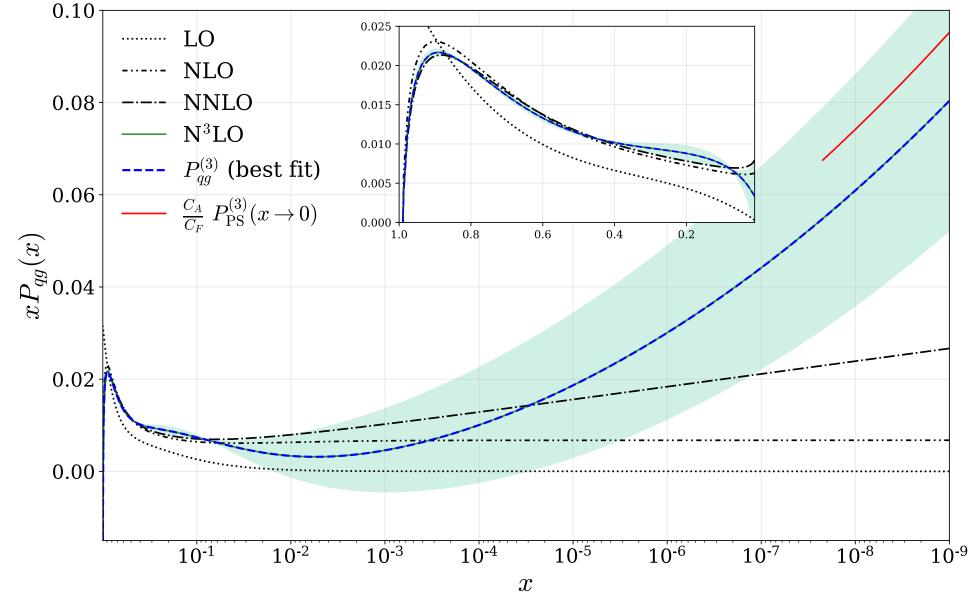
		Top Dataset	$\chi^2$	$\Delta \chi^2$	$\Delta \chi^2$ from NNL(
		-		from NNLO	(NNLO K-factor
	Tevatro	Tevatron, ATLAS, CMS $\sigma_{t\bar{t}}$ [97–109]		-0.5	-0.7
	ATLA	S 8 TeV single diff. <i>t</i> [110]	25.0 / 25	-0.7	-0.0
	ATLAS 8	TeV single diff. <i>tt</i> dilep. [1]	11] 2.2 / 5	-1.2	-0.7
	CMS	8 TeV double diff. $t\bar{t}$ [112]	23.8 / 15	+1.3	+4.9
	CM5	58 TeV single diff. <i>t</i> $\overline{t}$ [91]	8.3 / 9	-4.9	-5.4
		Total	73.3 / 71	-6.0	-2.0
ata Datasat	2	A <sup>2</sup> A <sup>2</sup> C			
ets Dataset	x-	$\Delta \chi^2 = \Delta \chi^2 \text{ fr}$	om NNLO		

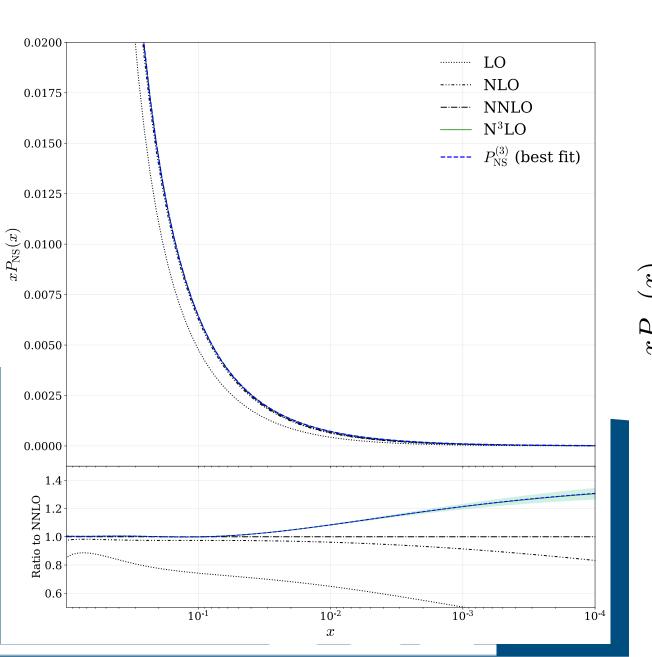
$p_T$ Jets Dataset	$\chi^2$	$\Delta \chi^2$	$\Delta \chi^2$ from NNLO	
		from NNLO	(NNLO K-factors)	
IS 7 TeV W + c [88]	12.2 / 10	+3.6	+1.3	
.AS 8 TeV Z p <sub>T</sub> [87]	106.3 / 104	-82.2	-52.5	
S 8 TeV W + jets [89]	19.2 / 30	+1.1	+0.4	
Total	137.7 / 144	-77.5	-50.9	

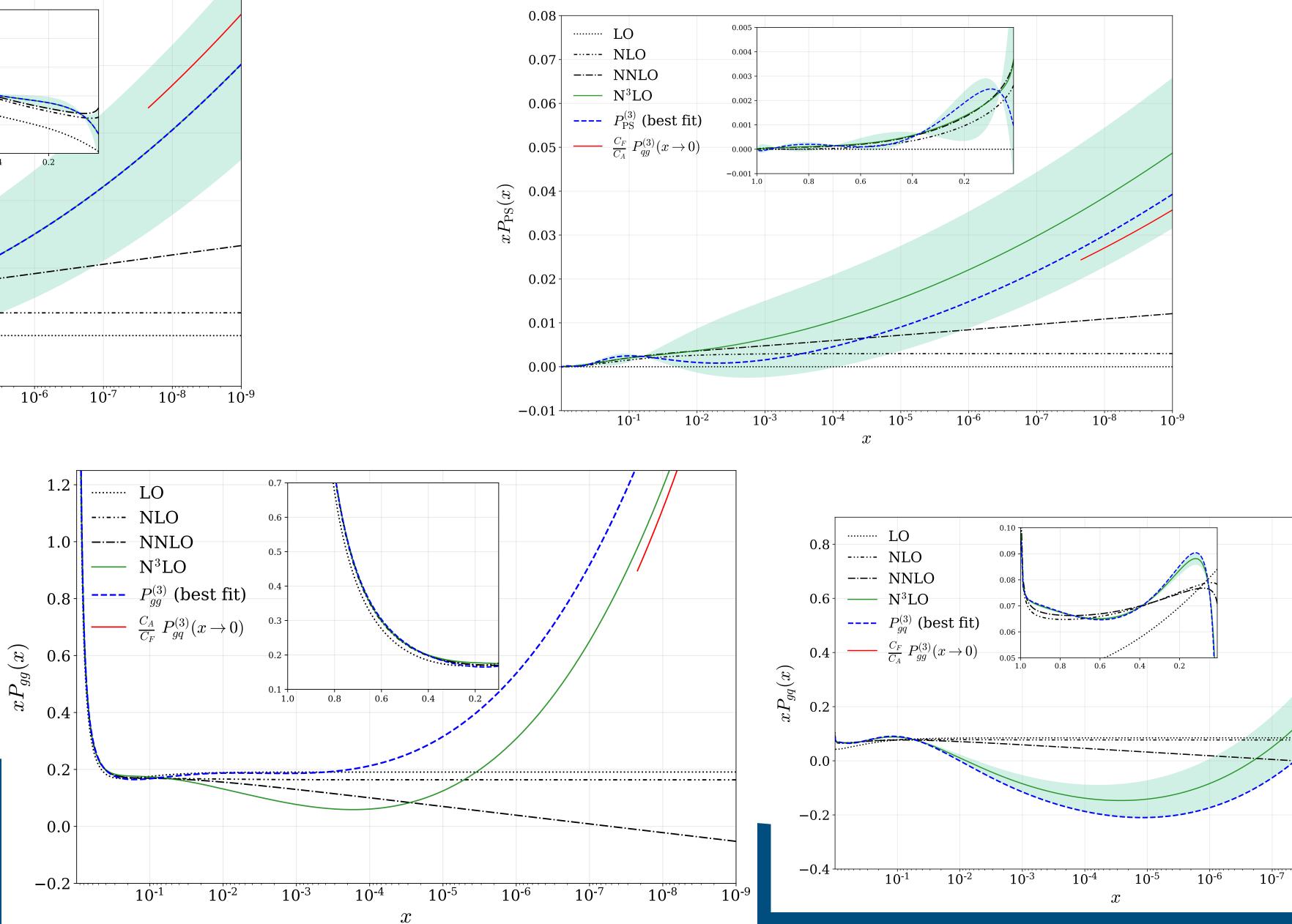


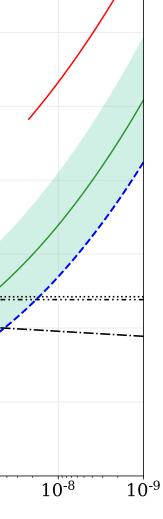


## **Approximate N<sup>3</sup>LO Splitting Functions**

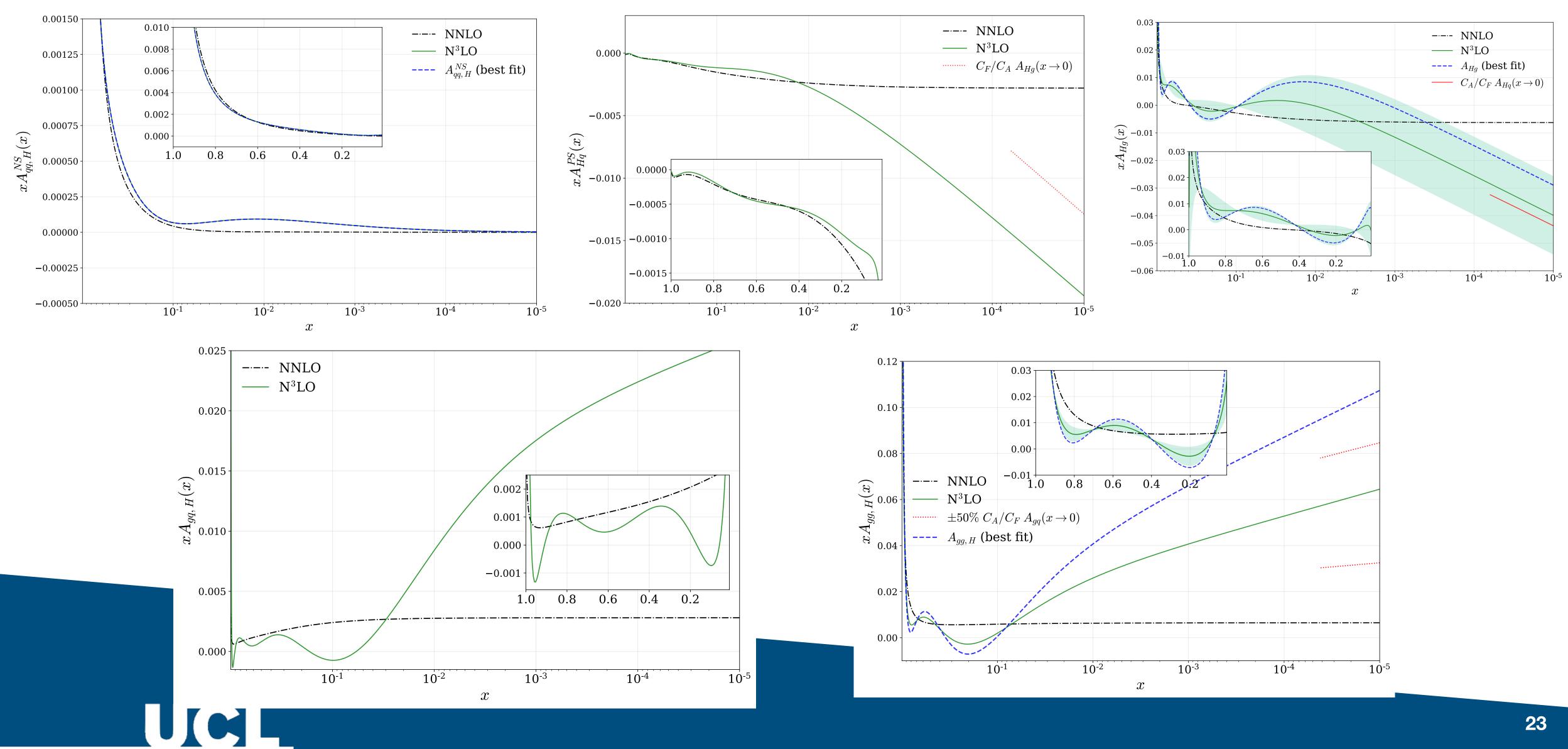




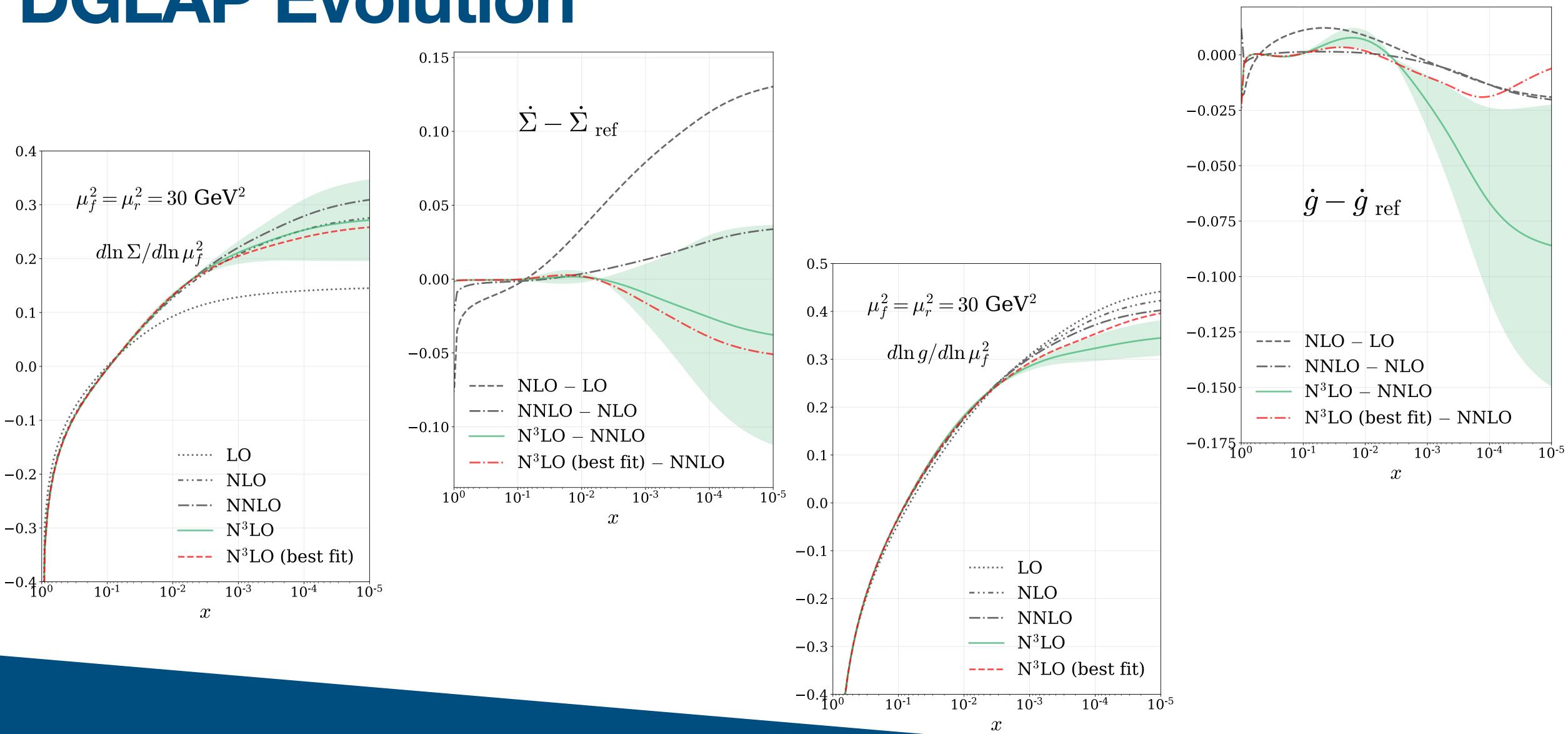




### **Approximate N<sup>3</sup>LO Transition Matrix Elements**



## **DGLAP** Evolution





## Usage of aN<sup>3</sup>LO PDFs

- For DIS processes, using the standard PDF set is advised.
- For any of the other 5 processes included in the fit (which we fit K-factors for), we provide the full details of these fitted aN<sup>3</sup>LO K-factors.
- For processes not included in the fit, this will be a little more involved.
  - Full details and instructions are provided in J. McGowan et. al., (2022) 2207.04739
    - Feel free to contact us with questions about usage.





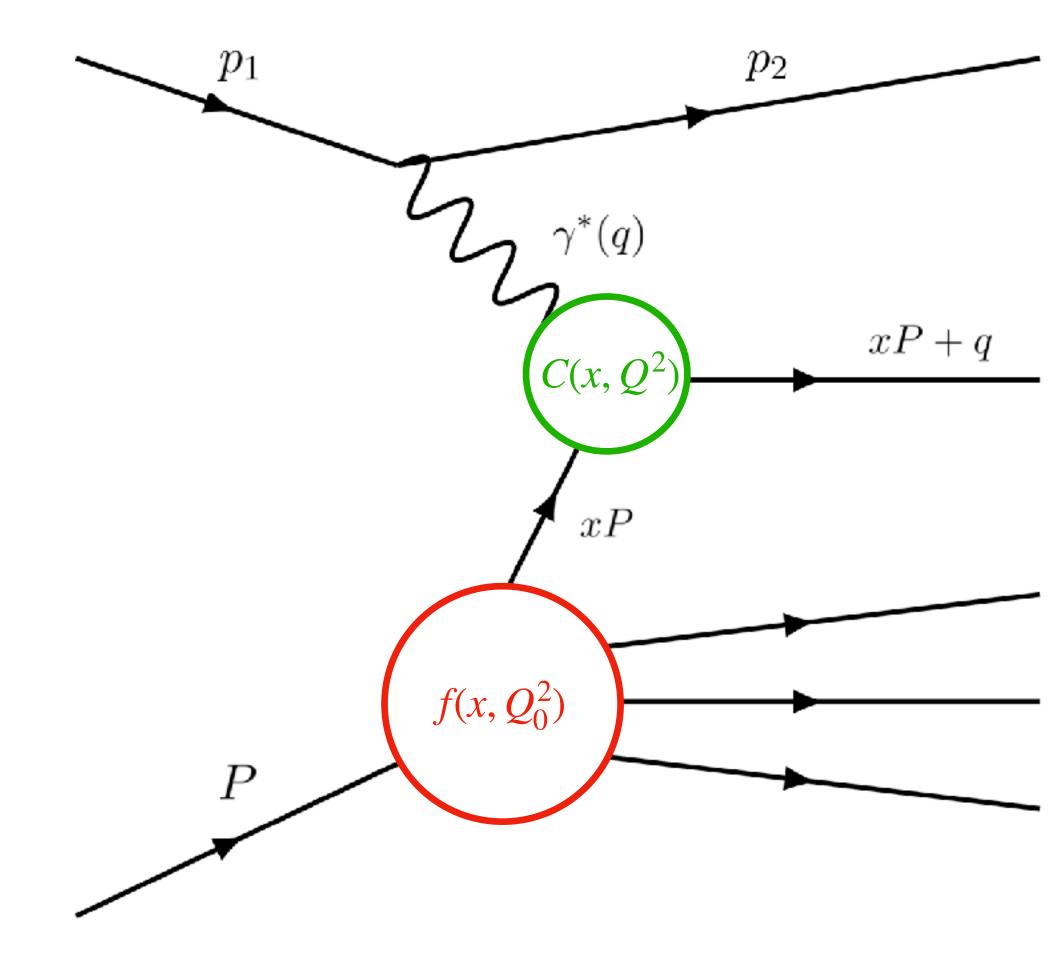
### **Deep Inelastic Scattering**

- PDFs probability of a parton fluctuating out of proton.
- Coefficient function perturbatively calculated.

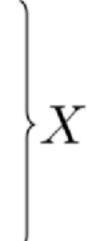
 $C(x,Q^2) = C^{(0)}(x,Q^2) + \alpha_s C^{(1)}(x,Q^2) + \alpha_s^2 C^{(2)}(x,Q^2) + \alpha_s^3 C^{(3)}(x,Q^2) + \dots$ 

 PDFs are determined from experiment using complex parameterisations.





• 'Global' fit using many different data sets and processes.







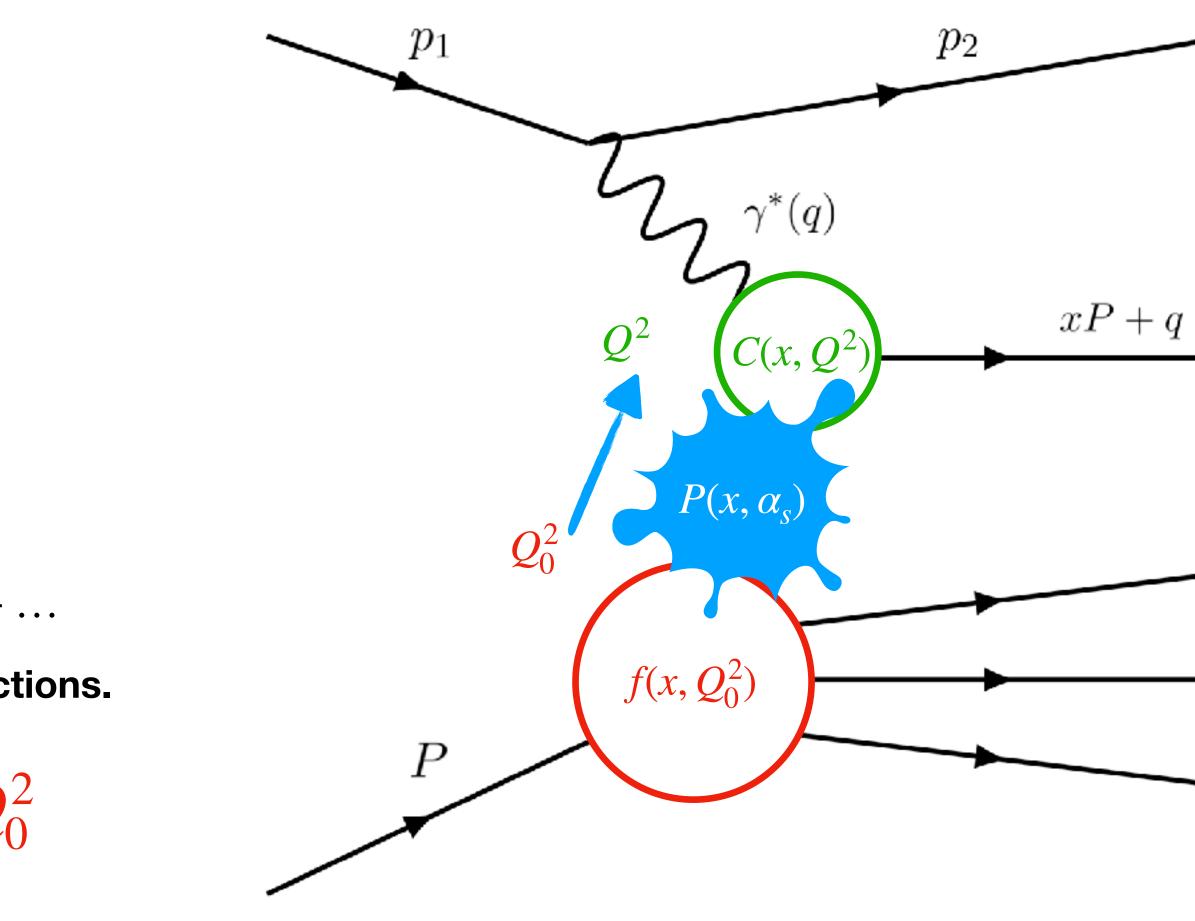
### **Deep Inelastic Scattering**

• Scale dependence of PDFs is also calculable in QCD perturbation theory!

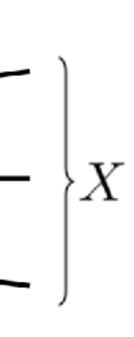
$$\mu^2 \frac{d}{d\mu^2} f(x,\mu^2) = P(x,\alpha_s(\mu^2)) \otimes f(x,\mu^2)$$

$$P(x,\alpha_s) = \alpha_s P^{(0)}(x) + \alpha_s^2 P^{(1)}(x) + \alpha_s^3 P^{(2)}(x) + \alpha_s^4 P^{(3)}(x) +$$
where  $P(x,\alpha_s)$  are the splitting functions

• PDFs parameterised at a starting scale  $Q_0^2$ and **evolved** to a desired scale  $Q^2$ .



**Takeaway:** Perturbatively calculable quantities are essential ingredients for PDF determination (and making predictions using PDFs).

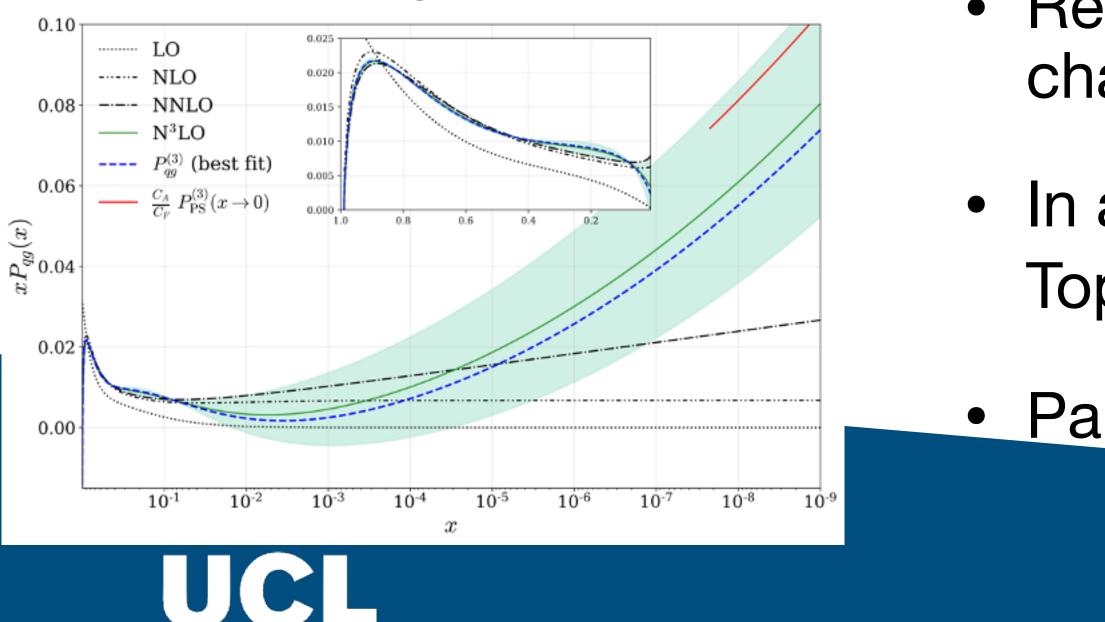


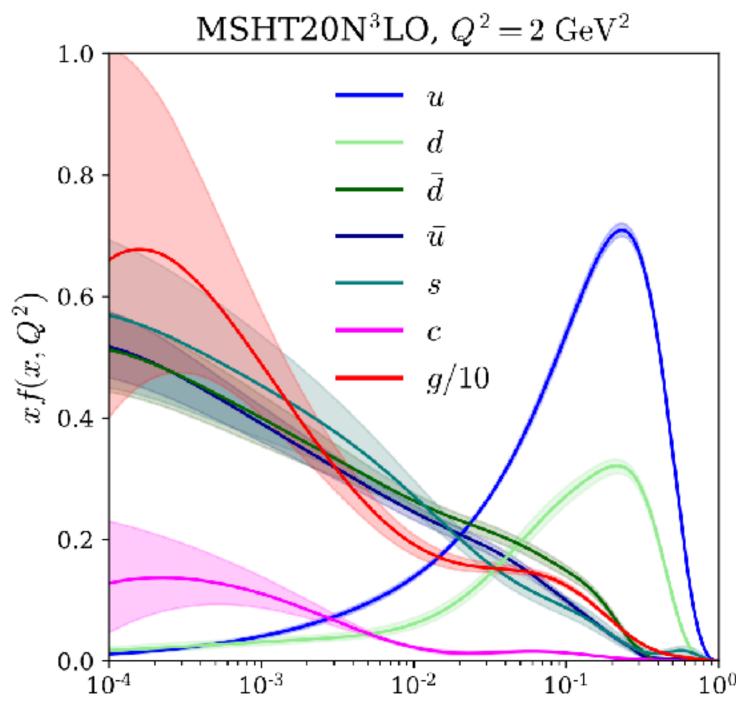




## **MSHT Approximate N<sup>3</sup>LO PDFs**

- MHOUs are leading source of theoretical uncertainty.
- Parameterisation of N<sup>3</sup>LO  $F_2$  structure function (incl. N<sup>3</sup>LO splitting functions) and N<sup>3</sup>LO K-factors for a consistent aN<sup>3</sup>LO fit.
- Overall better fit to data reduced tensions between small and large-*x*.





- Results show a harder gluon  $\rightarrow$  enhanced charm.
- In agreement with recent N<sup>3</sup>LO results DY and Top process *K*-factors. arXiv: 2107.09085, 2203.03698
- Paper and PDF sets available (very) soon.



