Dark Matter

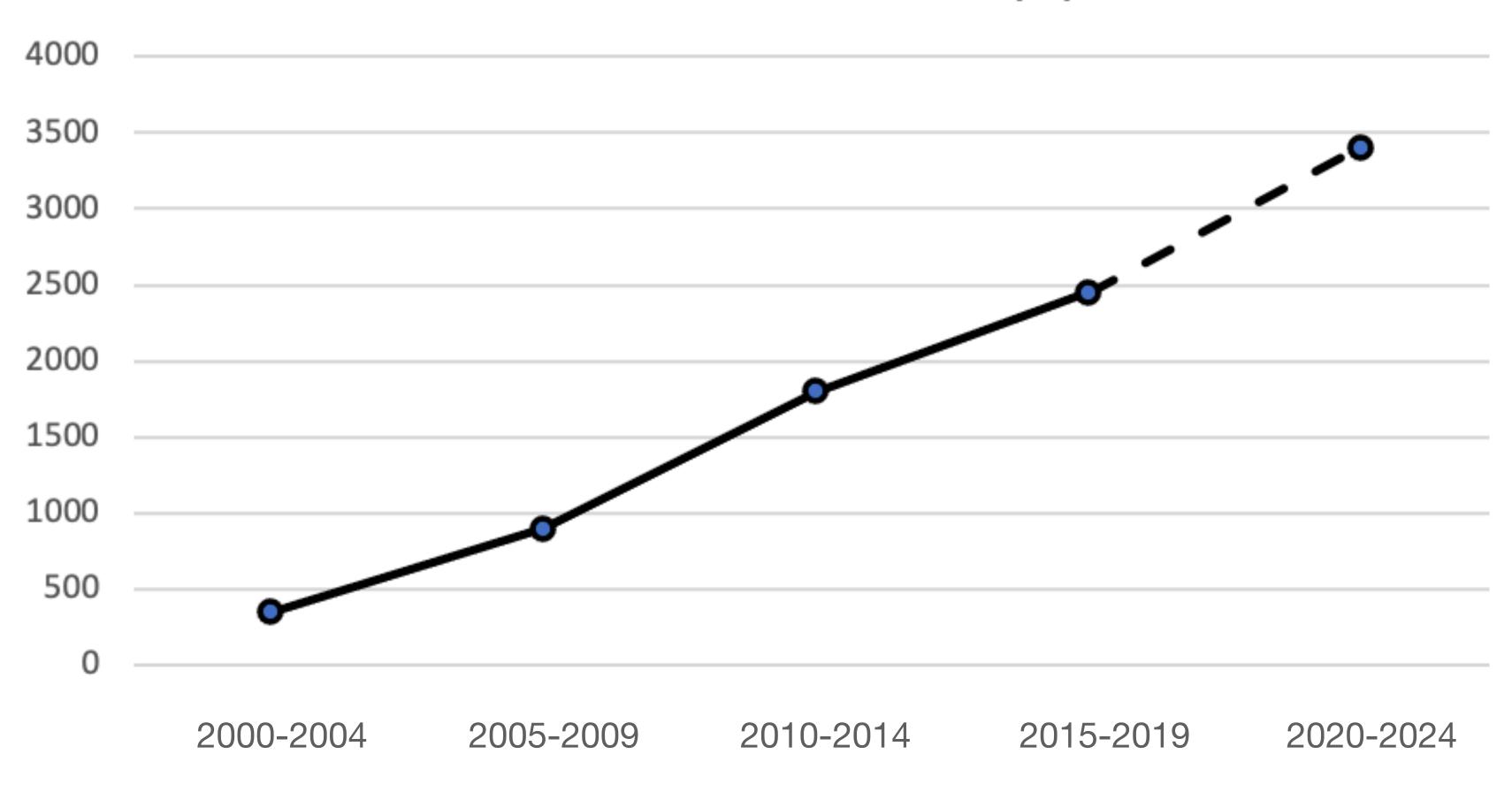
# Theory and Paradigmatic Models

... and EuCAPT's activities

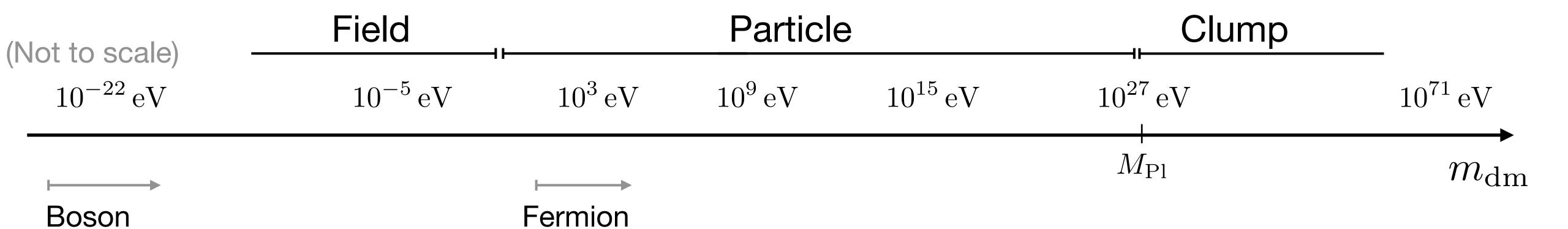


### Theory papers on dark matter

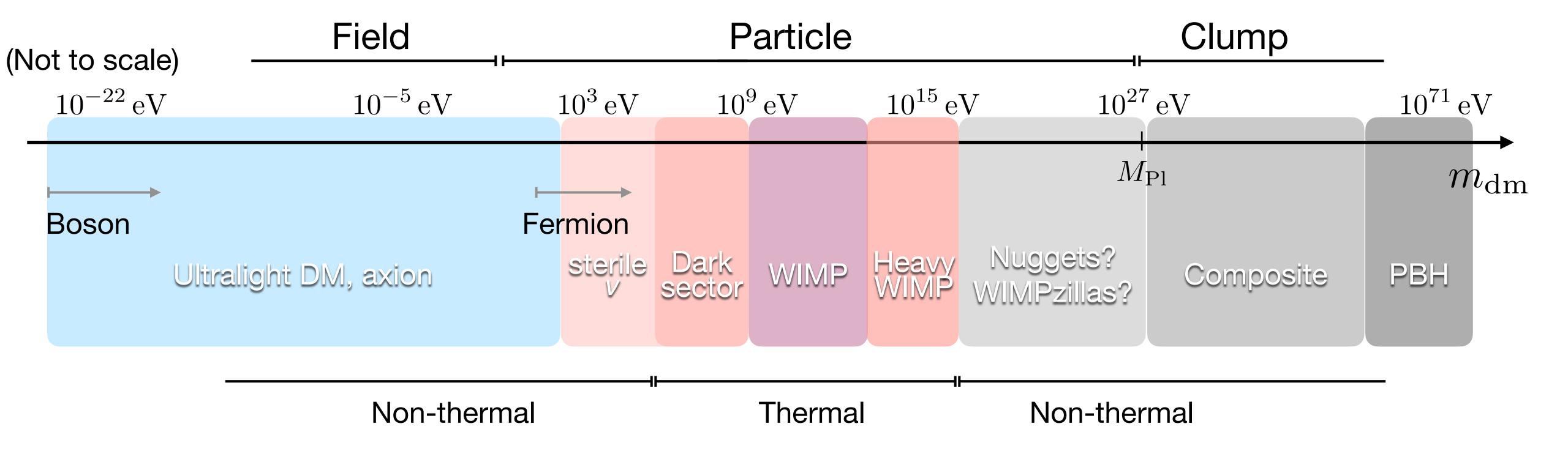
"dark matter" titles on hep-ph



## Model independent results

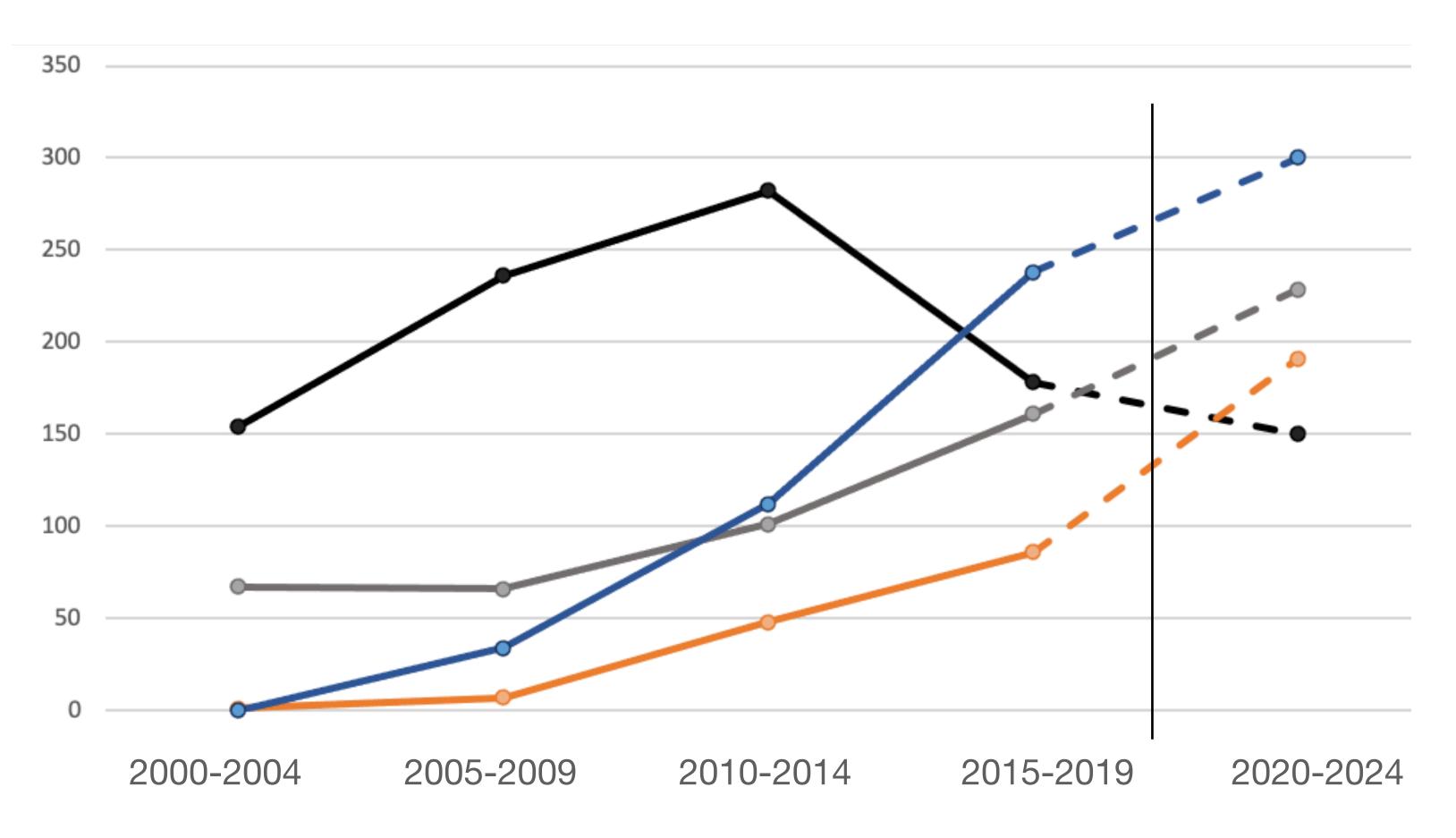


## Paradigmatic models



## What models do people work on?

(caveats apply)



WIMPs, supersymmetric dm, neutralino dm

Dark sector, portal dm

Sterile neutrino

Axion dm

#### Point 1:

Thermal models are still going strong

#### Point 2:

Non-thermal models are increasingly tested

#### Point 3:

The theory community strives to be open-minded and inter-connected

#### Point 1:

Thermal models are still going strong

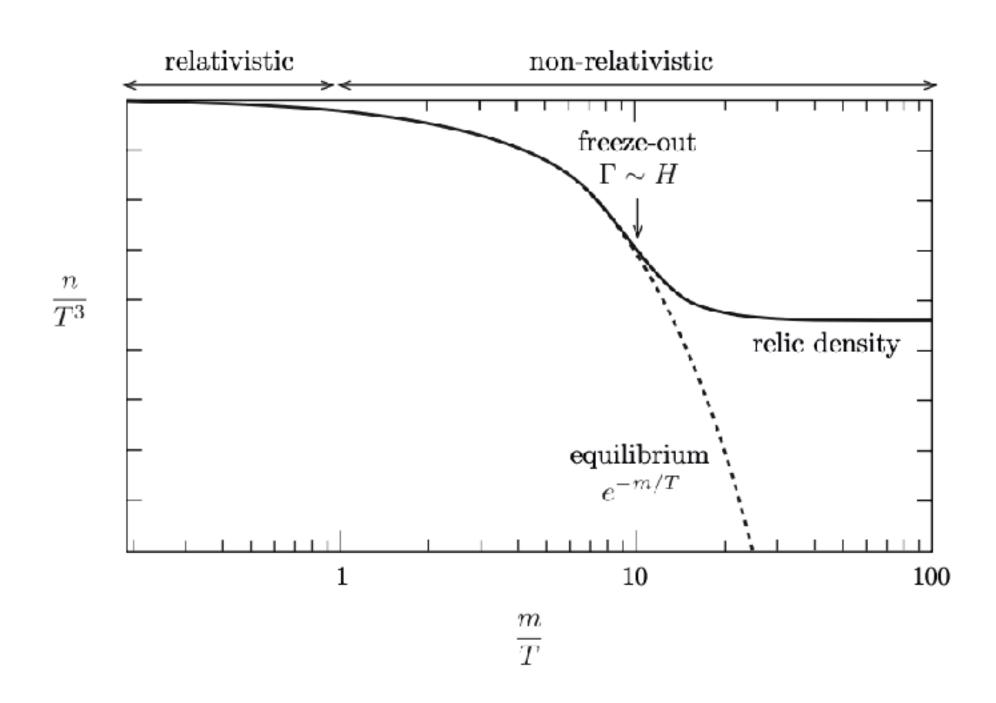
## WIMPs

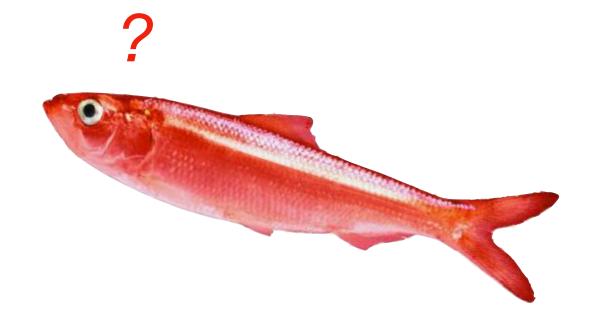
Thermal freeze-out:  $\Omega_{dm} \approx 0.2$  if

$$\langle \sigma_{\rm ann} v \rangle \sim \frac{\alpha_w^2}{m^2} \sim \frac{1}{({\rm TeV})^2}$$

Traditionally  ${
m GeV} \lesssim m \lesssim {
m TeV}$ 

Works well with low-energy supersymmetry



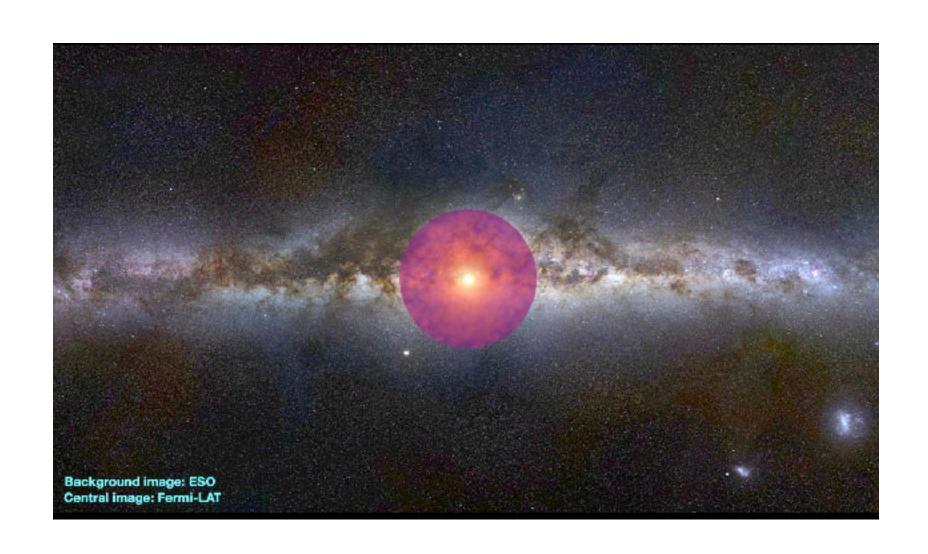


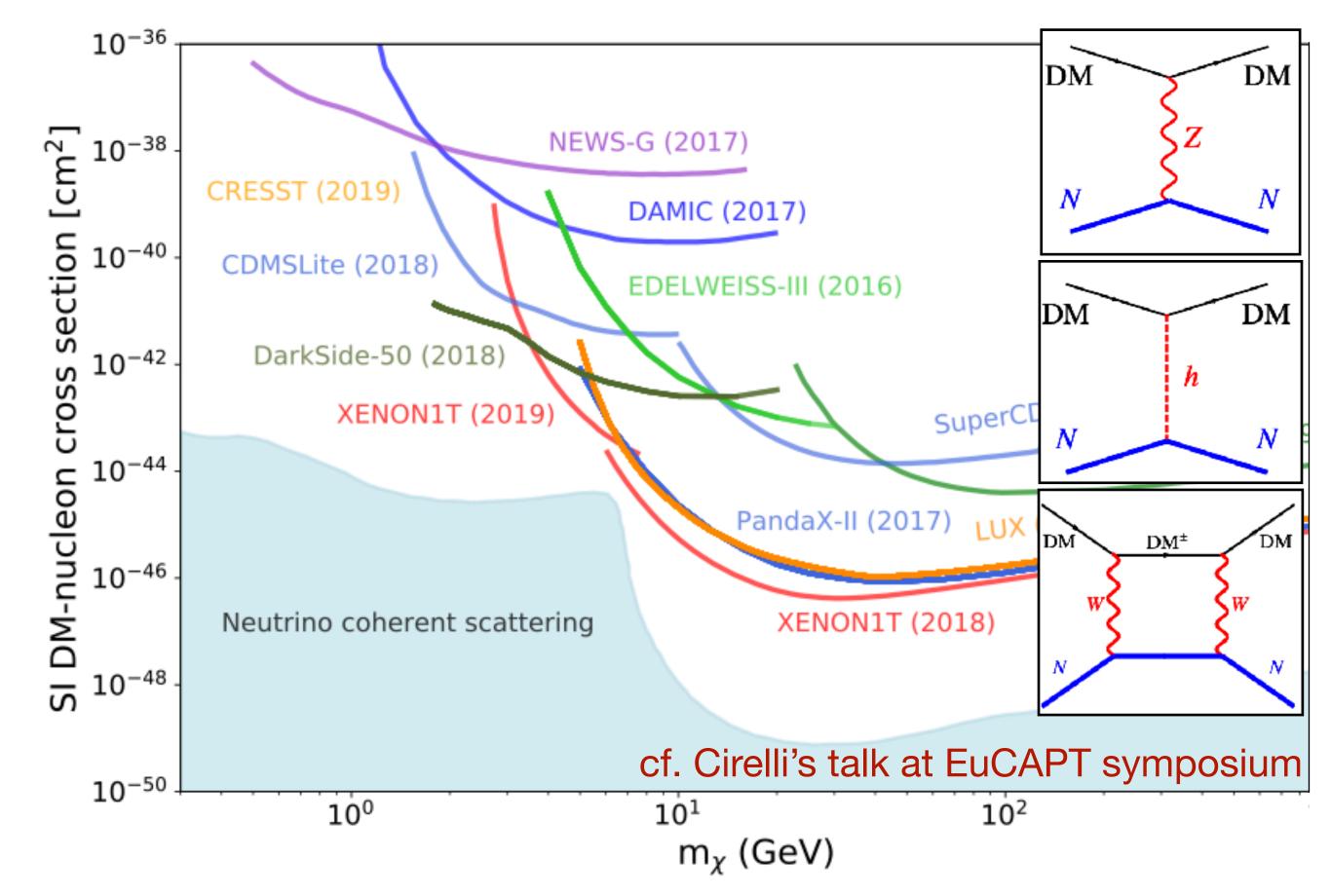
## WIMPs

Viable models down to & within the neutrino "fog"

cf. Baudis talk

## Hints from indirect detection (subject to astrophysical uncertainties) inspire sharpened model building





#### Challenges & developments:

WIMPs that do more

Highlight underlying assumptions

## Dark sectors

Thermal dark matter beyond WIMPs

$$\langle \sigma_{
m ann} v \rangle \sim rac{lpha_x^2}{m^2}$$
 New mediator e.g. vector that kinematically mixes with Z,  $\gamma$ 

Feng, Kumar 0803.4196

Broad viable mass range:

$${\rm keV} \lesssim m \lesssim 100 \,{\rm TeV}$$

Simplified models, effective field theory parametrisation

Review: Lin 1904.07915

Many realisations, e.g. SIMP dark matter:

$$m \sim 100 \, \mathrm{MeV}$$
 $\alpha_x = 1$ 

## Dark sectors

Both the dark matter and the mediator can be searched for

Synergetic with many new proposals for direct detection of light dark matter

cf. Privitera's talk

Many models have additional cosmological constraints, e.g. from N<sub>eff</sub>

Astrophysical constraints among the most stringent

#### Challenges & developments:

Model building and phenomenology alongside experimental and observational efforts

#### Point 2:

Non-thermal models are increasingly tested

## Sterile neutrinos

Singlet fermions that mix with neutrinos

$$\frac{1}{2} \left( \bar{\nu}_L, \bar{\nu}_R^c \right) \begin{pmatrix} 0 & m_D \\ m_D & M \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix} + \text{h.c.}$$

See-saw mass, and small mixing  $\theta = \frac{m_D}{M}$ 

Narrow mass range:  $\text{keV} \lesssim M \lesssim \text{few} \times 10 \, \text{keV}$ 

Review: Boyarsky et al. 1807.07938

Production: thermal non-thermal

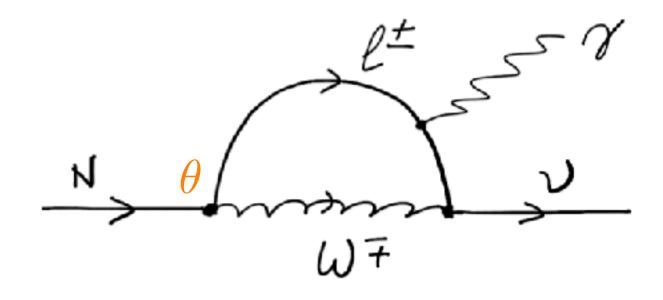
Always there

MSW resonance

cannot produce all dm

in-medium mass degeneracy, if L is sufficiently large

## Sterile neutrinos



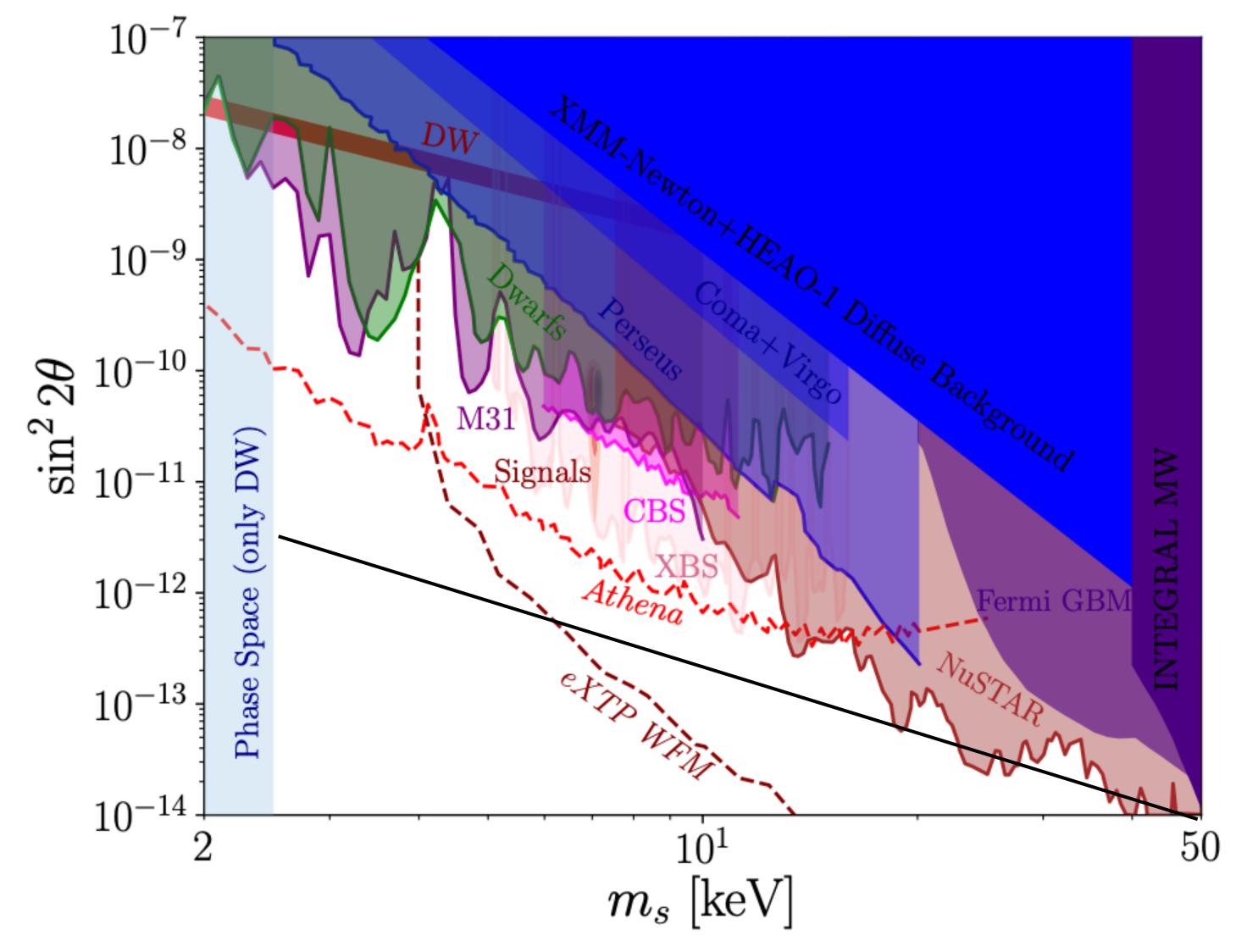
Test: X-ray line signal

#### Challenges & developments:

Robust calculations of spectrum

How to generate large enough *L*?

Non-minimal models



## Axions

A class of naturally light (pseudo-) scalar dark matter candidates

cf. Majorovits talk

#### The QCD axion:

- Naturally light
- Couples to photons
- Couples to gluons
- Nulls neutron EDM
- One parameter:  $m_a$

#### **Axion-like particles (ALPs):**

- Naturally light
- Does not couple to gluons
- All Wilson coefficients are parameters:  $m_a$  ,  $g_{a\gamma}$
- Ubiquitous in string theory

Excellent cold dark matter candidates
Behaves like classical field

## Axions

Viable mass range for QCD axion dark matter depends on cosmology

#### I. Homogeneous across the universe

"Pre-inflationary scenario" axion misalignement

#### II. Random initial conditions when formed

"Post-inflationary scenario" axion misalignement + topological defects

ALPs can comprise ultralight dark matter:

$$m_a \gtrsim 10^{-22} \,\mathrm{eV}$$

$$m_a \le 0.15 \,\mathrm{meV}$$
  $m_a \gtrsim 0.1 \,\mu\mathrm{eV} \quad (\mathrm{or} \ 10^{-12} \,\mathrm{eV})$ 

$$m_a = (25.2 \pm 11.0) \,\mu\text{eV}$$

A single viable value, but hard to compute

Buschman et al. [1906.00967] Review: Di Luzio et al. [2003.01100]

## Axions

Most interesting parameter space is now probed experimentally

Strong synergy with astroparticle physics

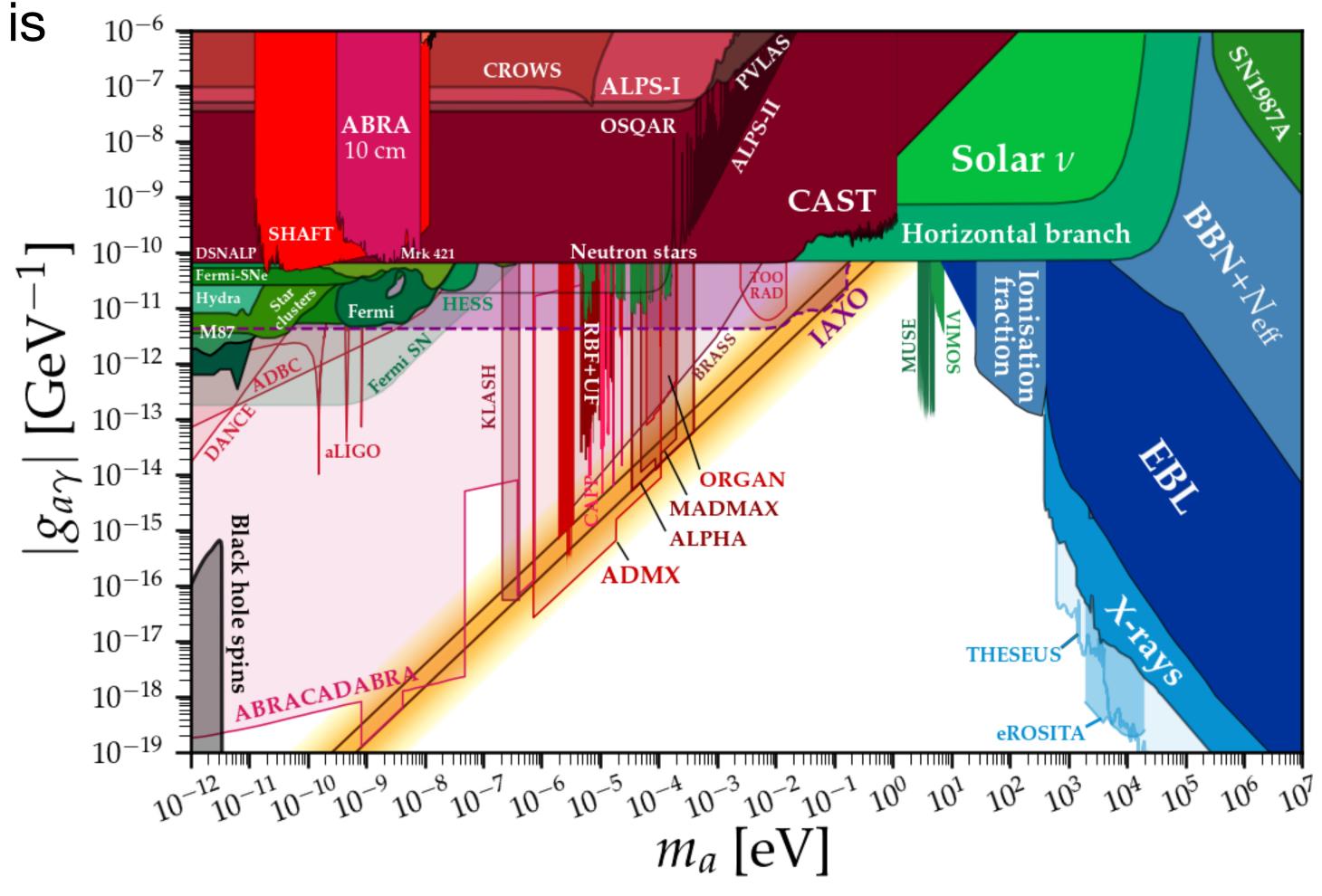
#### Challenges & developments:

Get predictions right

New ways to probe

Populating the parameter space

Theoretical expectations for ALP models

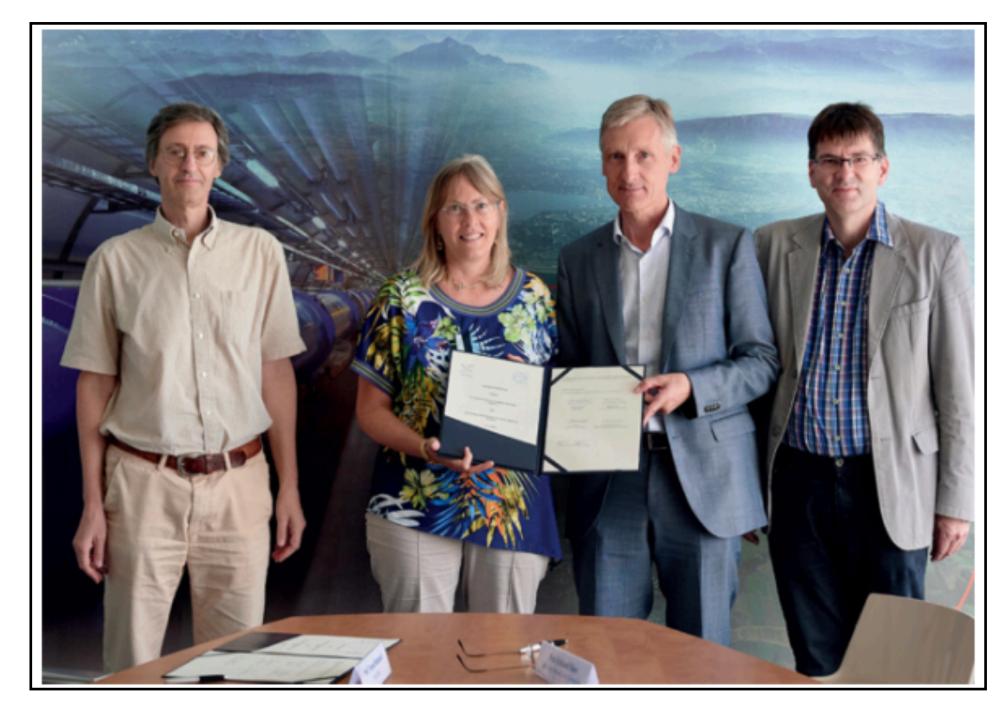


#### Point 3:

The theory community strives to be open-minded and inter-connected

## EuCAPT aims to:

- bring together the European community of theoretical astroparticle physicists and cosmologists
- increase the exchange of ideas and knowledge
- coordinate scientific and training activities
- help scientists attract adequate resources for their projects
- promote a stimulating, fair and open environment in which young scientists can thrive



Fonded in 2019 by APPEC and CERN.

## Team

#### Steering committee:



Gianfranco Bertone (director)



Piero Ullio



Philippe Brax



Licia Verde



Vitor Cardoso



Gian Giudice



**David Marsh** 



**Andrew Taylor** 



Toni Riotto



**David Langlois** 



Silvia Pascoli



Nikolina Sarcevic





## Activities

Meetings & talks:

Community building:

Communication:

Newsletter

Annual symposium

White paper

Thematic workshops

Training

Webpage

List of virtual seminars

Community profiles

Research highlights

Code of conduct

Census

Virtual colloquia



## Summary

Dark matter theory isn't a topic, but a wide field

Models offer a way to organise huge span of possibilities; suggest paths for detection; highlight tacit assumptions

Each sub-field has its own challenges; developments on many fronts

Point 1:

Thermal models are still going strong

Point 2:

Non-thermal models are increasingly tested

Point 3:

The theory community strives to be open-minded and inter-connected

Several sub-fields develop closely with experiments and observations

Several long-standing paradigms are facing critical tests