## **MDI and Detector**

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## **Interaction Region and MDI Design**

The high luminosity requires:

- Low beta-function at the IP (few cm)
- **•** High number of muons per bunch  $(N_{\mu} \sim 2 \cdot 10^{12})$

Muons decay particles…back of the envelope evaluation:

beam 1.5 TeV  $\lambda = 9.3 \times 10^6$ m, with  $2 \times 10^{12}$   $\mu$ /bunch  $\Rightarrow 2 \times 10^5$  decay per meter of lattice.

Beam induced background, if not properly treated, could be critical for:

- Magnets, they need to be protected.
- People, due to neutrino induced radiation.
- **Detector, the performance depends on the rate of background particles arriving to each subdetector.**

A holistic approach is needed, tight together the development of the IR optics, the magnets and the shielding strategies (magnets and detector).

### **Optimization of Interaction Region at**  $\sqrt{s} = 1.5 \text{ TeV}$ These issues as well as well as well as well as well as well as  $\omega$  as  $\omega$  as  $\omega$

Fermilab-Conf-11-094-APC-TD Y.I. Alexahin et al. Muon Collider Interaction Region Design FERMILAB-11-370-APC N.V. Mokhov et al. *Muon collider interaction region and machine-detector interface design* 





in the coils and in the inner part of the detector.  $\mathbf{p}$ Quadrupoles in  $Nb<sub>3</sub>Sn$  characteristics in the papers. Dedicated dipoles to minimize the number of decay electrons

 $p)$  and

 $s(m)$ 



## **Optimization of Interaction Region at**  $\sqrt{s} = 1.5$  **TeV with absorbers**

Important role is played by the absorber materials



Deposited power density in Q1 (mW/g)

nozzle inside quadrupoles oles of 0.1 of their aperture FIG. 6 (color). Deposited power density in Q1 (mW/g) for three cases: "standard" (left), with absorbers inside (center) FIG GGGdgr). Deposited power density in UmpWe for three cases: "standard" (left z with absorbers inside (center nozzle

 $\mathcal{A}$  it follows from Table 4, the traditional cos  $\mathcal{A}$  $\mu$ ponatena Lucchesi

## **IR and Detector background**



A: no masks between magnets, 6° cone with a 5σ radius liner up to 2 m from IP; **B:** 5σ masks inserted between FF quads, cone angle 10°, 5σ liner up to 1 m from IP; **C:** same as above plus FF quad displacement.

### Results:

- Masks and increased cone angle reduce  $e^{\pm}$  and  $\gamma$  fluxes by factors 300 and 20.
- Displacing FF quads increases  $e^{\pm}$  flux by up to 50% decreases  $\gamma$  flux by factor 15

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## **Further Detector Nozzle Optimization**

 $\frac{1}{2}$  and  $\frac{1}{2}$  a For example, gamma flux as a function of the angle of inner cone opening towards IP at the outer cone angle of 10°



FIG. 9 (color). Gamma flux vs. inner cone angle at



10<sup>0</sup> nozzle geometry **General view**  $(600, 60)$  $R, cm<sub>1</sub>$ Zoom in beam pipe  $(Z,R)$ 600.50  $Z, cm$  $(6.25, 2.2875)$ <br> $(6.25, 2.24)$ **BCH2** 13.762. 2.37322  $(0, 2.24)$ 3.2334, 2.27969)  $(6.25, 2.2)$  $(0, 2.2)$  $(100, 17.6)$  $(200, 17)$  $0.151$ W W  $(600, 1.78)$ Be  $(100, 0.3)$  $W$  – tungsten R.cm. **Be-beryllium**  $(Z,R)$ **BCH2** – borated polyethylene  $(15, 0.6)$ Z.cn

These studies have brought to the final nozzle configuration

#### **Beam-Induced Background study Deam-Induced Dackground study**







Produced with MARS15: particles arriving to the detector.

## **Further Detector Nozzle Optimization**



### IR elements and geometry produced by LineBuilder visualized by FLAIR

Beam Pipe aperture, coil transverse dimensions, materials...



- MAP optic files.
- Details on magnets material and passive elements are also needed

## **Visualization of all the elements of the IR**



**NozzleBR** 

## **Complete IR Design**





#### **Comparison MARS15 - FLUKA** computed the iterative task of MDI optimization will rely on a problem interpretation of the Monte State of



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## **Comparison MARS15 - FLUKA**



One beam,  $\mu^-$  of 750 GeV with  $2 \cdot 10^{12}$  particles/bunch

## **Results in details**

One beam,  $\mu^-$  of 750 GeV with  $2 \cdot 10^{12}$  particles/bunch



## **Beam-Induced Background Origin**

One beam,  $\mu^-$  of 750 GeV with  $2 \cdot 10^{12}$  particles/bunch





# **Given the BIB, how do we design the detector?**

## **Tracker at**  $\sqrt{s}$  = 1.5 TeV



Tracking performance have been studied applying timing and energy cuts on clusters reconstruction compatible with IP time spread.

Vertex detector properly designed to not  $\frac{1}{\sqrt{1-\frac{1$ overlap with the BIB hottest spots around the interaction region.



## **Calorimeter at**  $\sqrt{s}$  =1.5 TeV



BIB deposits large amount of energy in both ECAL and HCAL



# **Is this the end of the story?**



## **Table 2 1012. Number of BIB particles from MARS15 file "mumi-1e3x500-26"**



Previous St Detector Bac Collider: MA factor  $\sim$ 2 He

Sources of differences:

- Different materials between MARS and FLUKA?
- Passive materials, do we have all the absorbers?
- Intrinsic differences MARS15 vs FLUKA? Generate BIB

#### Lattice improvements

- Add collimators to remove secondary muons coming from very fa
- Further optimization of absorbers and nozzle?
- **•** Further optimization of magnets aperture and liner?



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## **To conclude**

- $\triangleright$  Simulation and analysis tools to optimize MDI are ready and well tested.
- Ø Benchmark Monte Carlo? Test with data?
- $\triangleright$  Nest step is nozzle optimization at  $\sqrt{s} = 3$  TeV where the IR lattice is well tuned. Work is already in progress.

Strong collaboration between accelerator and detector physicists is mandatory for the proper MDI design.

