

Evolution of Ly α halos between low and high redshift

Possible sign of changes in the CGM

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Lya is a common tool
understanding it is critical - but hard

Lya is among the most common tools for finding high redshift galaxies

- Large FOV Integral field spectrographs (MUSE)

It is also a potential diagnostic of many things:

- Star formation
- Neutral gas around galaxies
- Reionization

We need low-z observations
to get at the physics

What galaxies are emitting the Ly α ?

- how does it relate to the properties of galaxies

Optical emission lines are unobservable at high-z
but obligatory JWST reference

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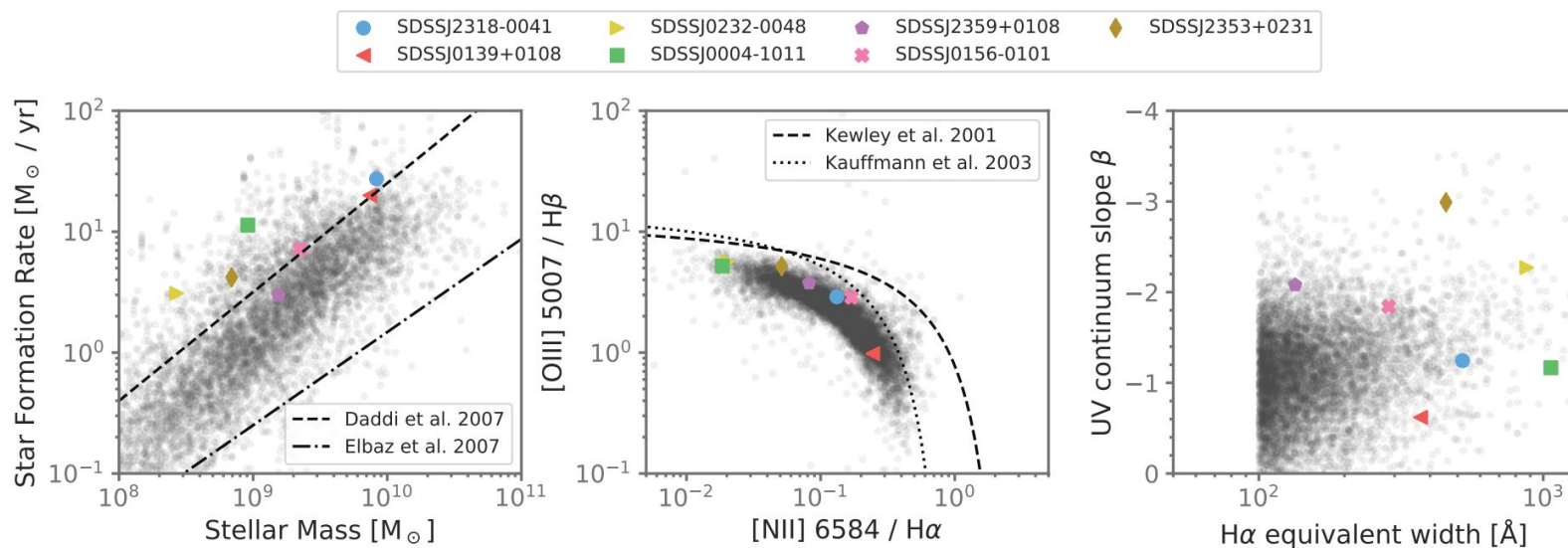
But low-z brings some unique challenges:

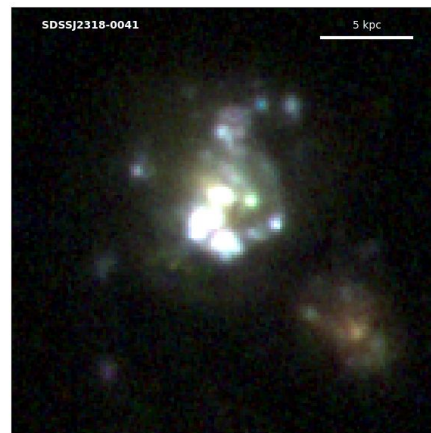
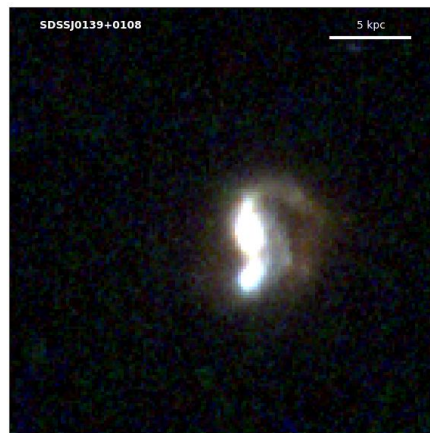
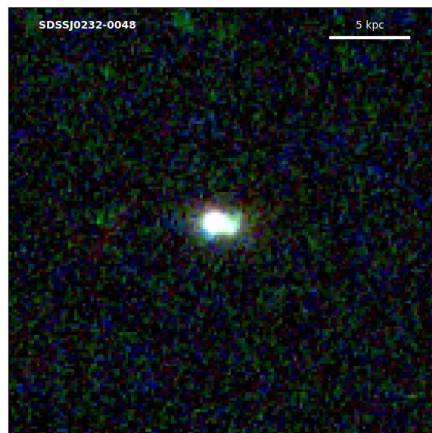
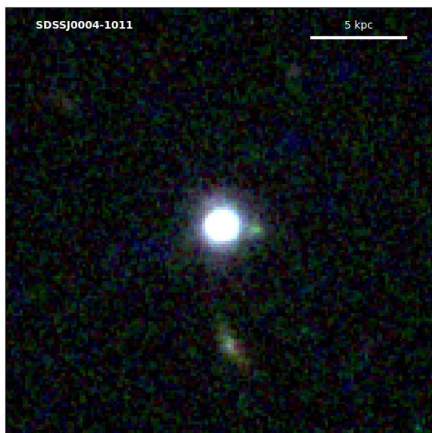
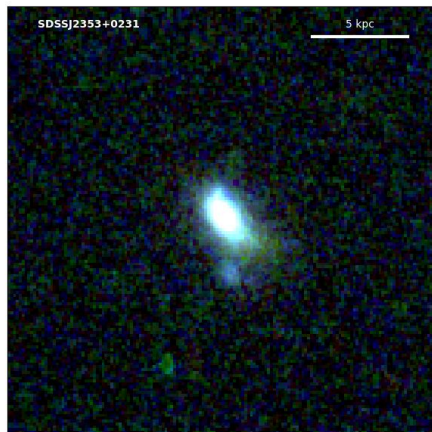
- $\lambda = 1216$ -> need UV observations -> HST
- Backgrounds: Geocoronal emission (Ly α or [OI])
- Galaxy sizes
- Dark currents

We have a new systematics controlled sample

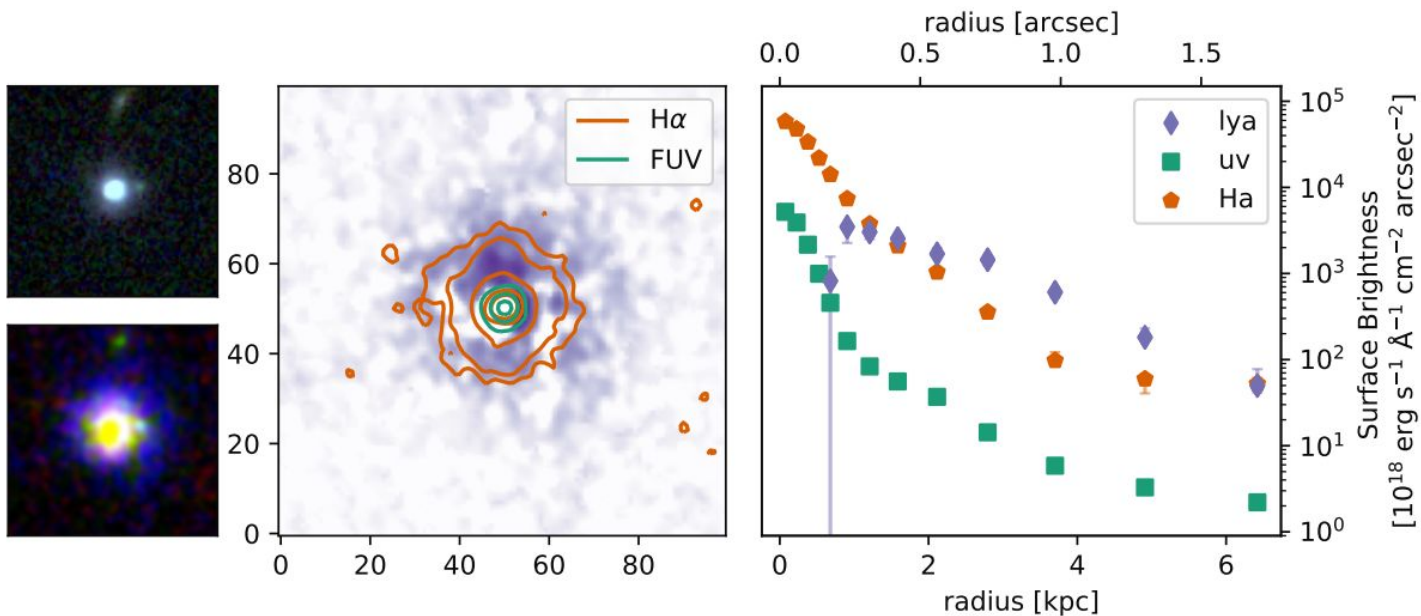
We can sacrifice some spatial resolution to control for systematics and backgrounds

We also match the galaxies in properties to high-z samples

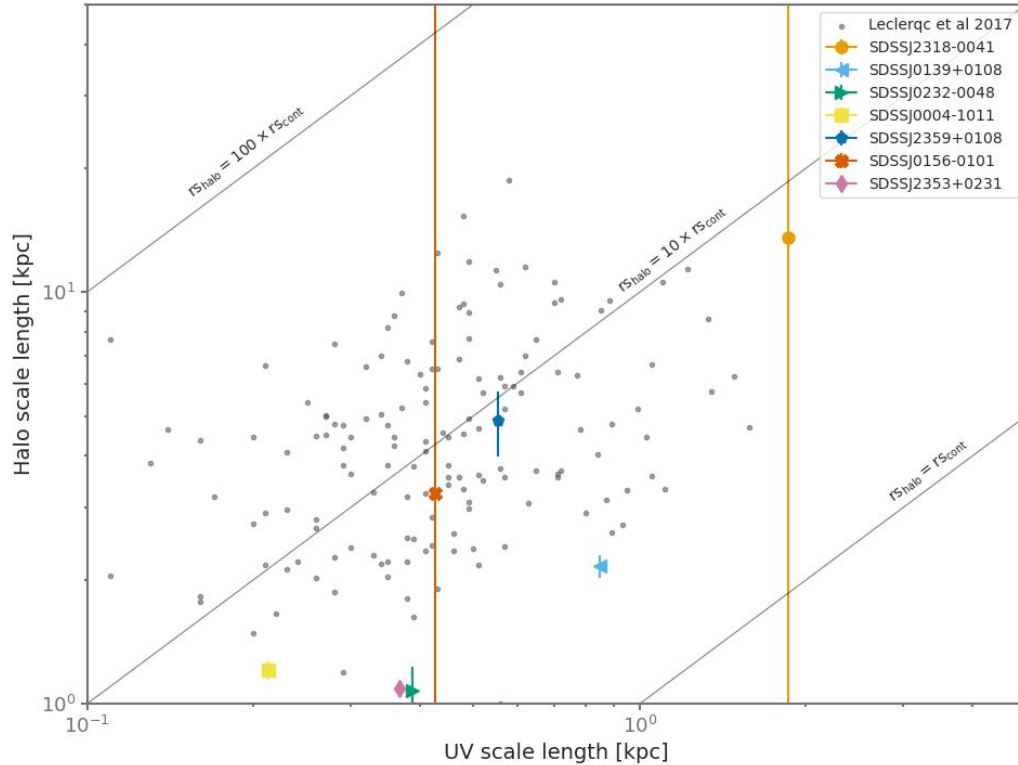




We want to characterize the extended Ly α emission

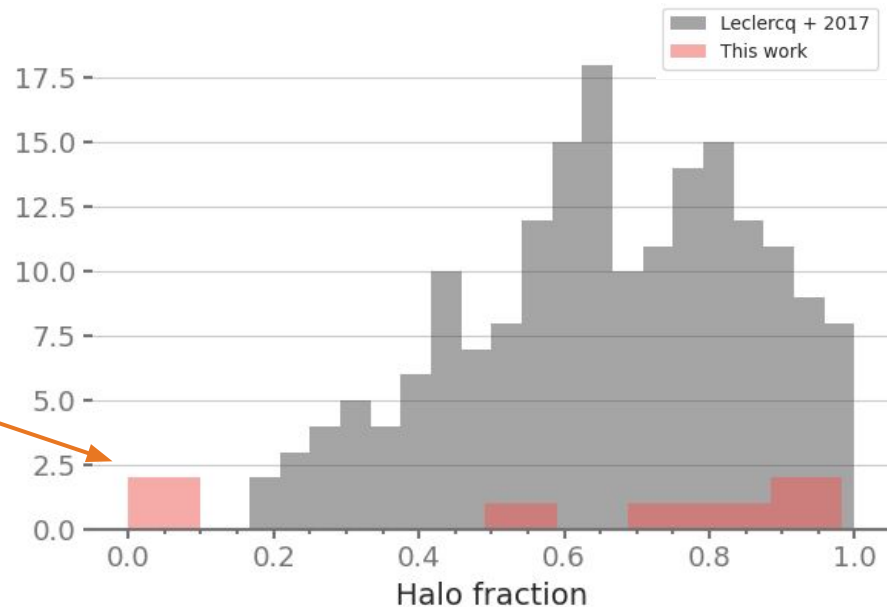


Despite being selected to be similar
populates only the smaller halo space



And some galaxies see smaller fraction of halo emission than any MUSE galaxy

2 out of 7 galaxies
consistent with only central
Iya emission



Difficult to say why this should be

Bias in emission line selection at high redshifts?

- **Not likely:** The most centrally concentrated emitters are also the brightest

No scattering gas?

- Surveys of neutral gas around galaxies show that there is plenty of that

We need to examine the correlations with other galaxy properties more in detail

In conclusion: Tentative hints that low-z Iya emission differs from high-z

Trying to determine whether the effect is real or not is crucial.

Then we can see if there are actually some properties of the galaxies themselves that change

We can use JWST for that!