

Dark sector searches with NA64 experiment @CERN

Laura Molina Bueno
laura.molina.bueno@cern.ch

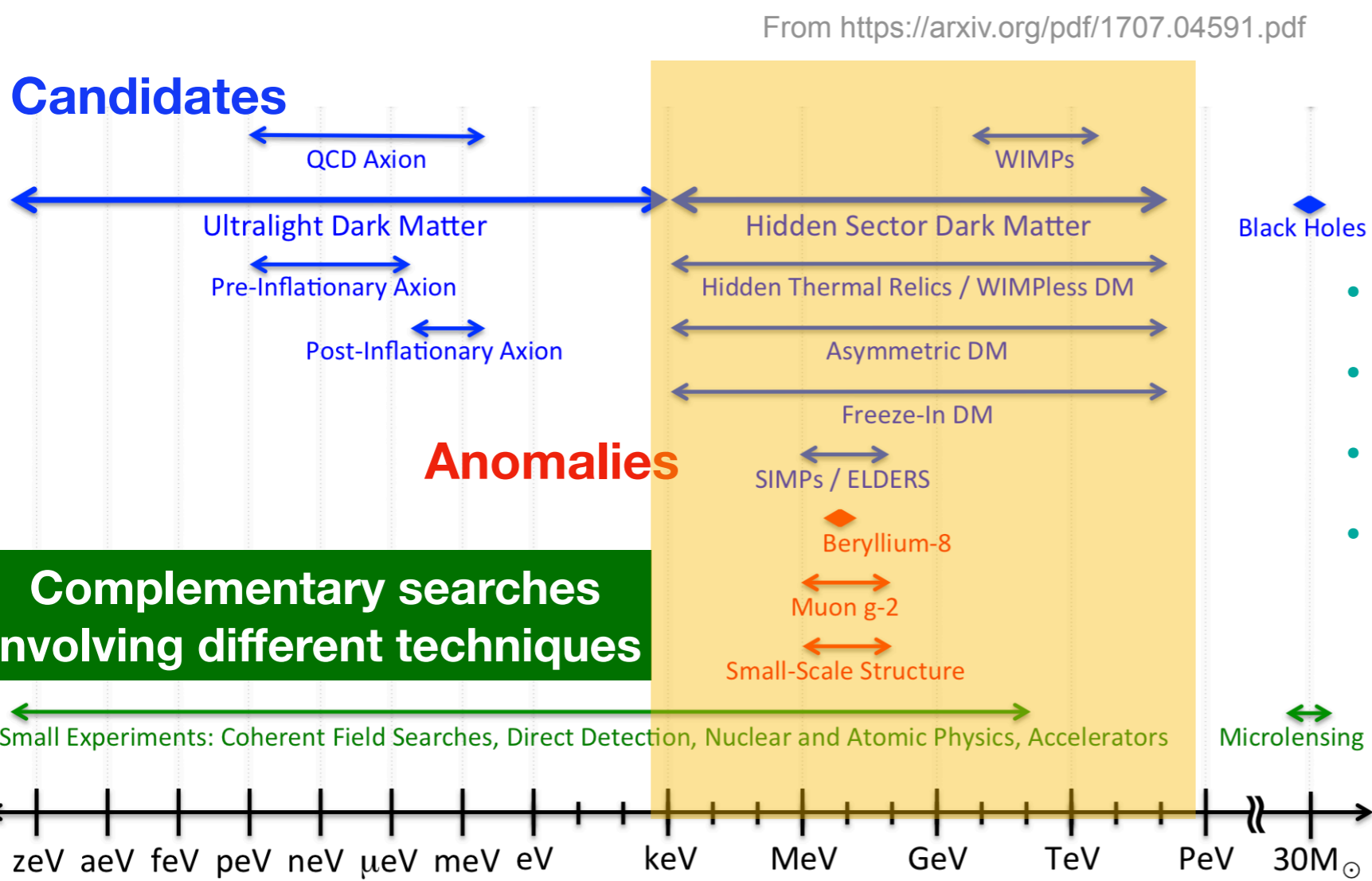
“A rainbow of Dark sectors”
Aspen online winter conference
21st March-1st April 2021

Dark sectors



An interesting framework to explain the origin of dark matter

Existence of *dark sectors* which couple weakly with standard model particles and can decay into dark matter candidates.



$$L_{Total} = L_{SM} + L_{DS} + L_{Portal}$$

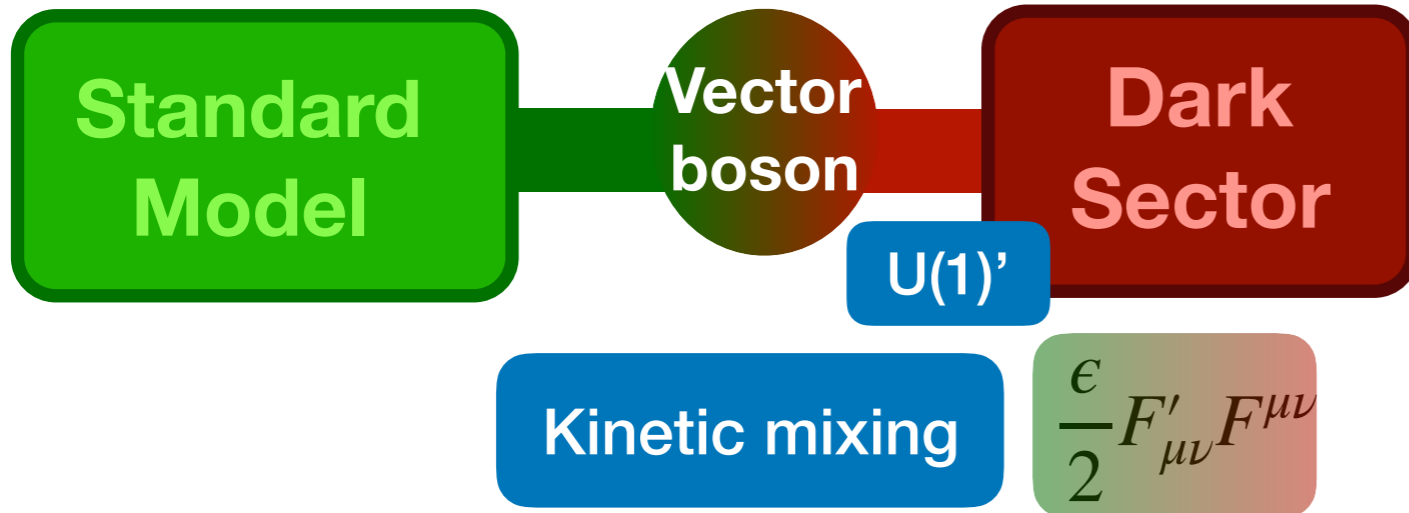
- Vector: **Dark Photon**
- Scalar: **Dark Higgs**
- Fermion: **Heavy neutral lepton**
- Pseudo-scalar: **Axion**

Recent review:
G.Lanfranchi, M.Pospelov and P.Schuster
arxiv:2011.02157

See N.Toro's talk!



NA64 target: *the vector portal*

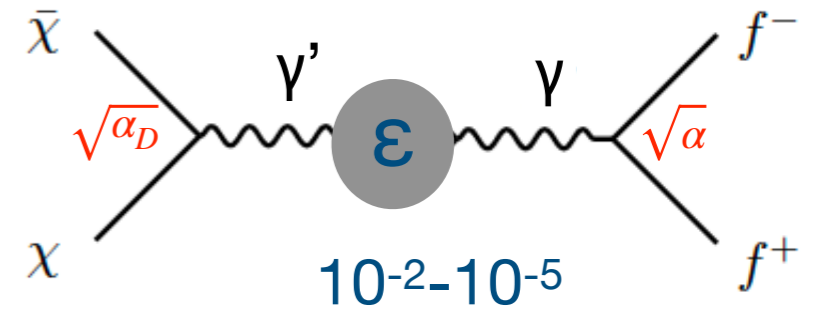
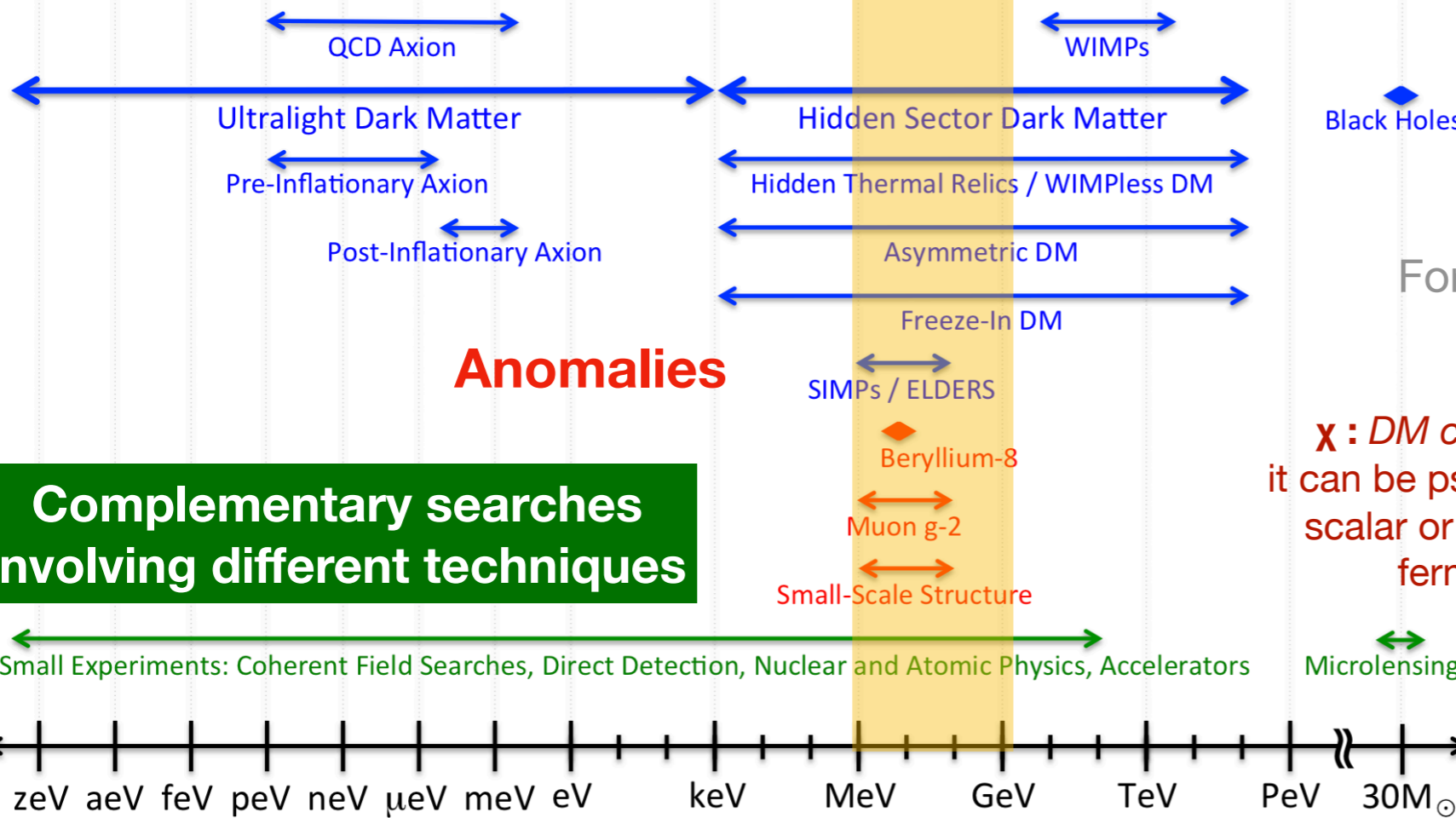


An interesting framework to explain the origin of dark matter

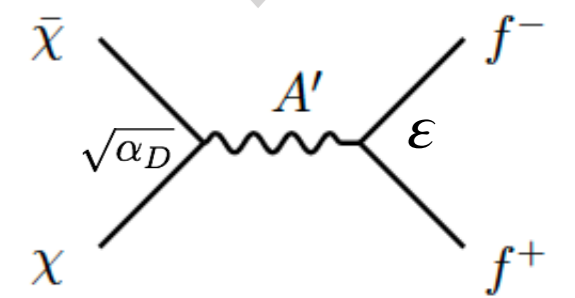
$$L_{Total} = L_{SM} + L_{DS} + L_{Portal}$$

• Vector: *Dark Photon*

Candidates



For simplicity during the talk



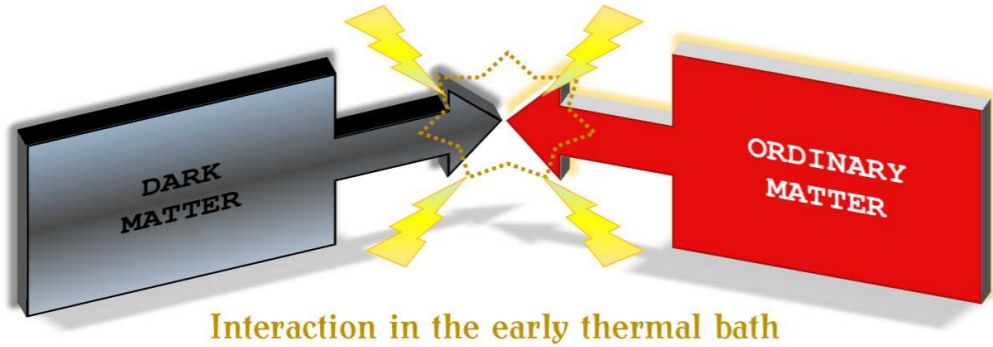
Parameter space defined by $(mA', m\chi, \epsilon, \alpha_D)$

See N.Toro's talk!

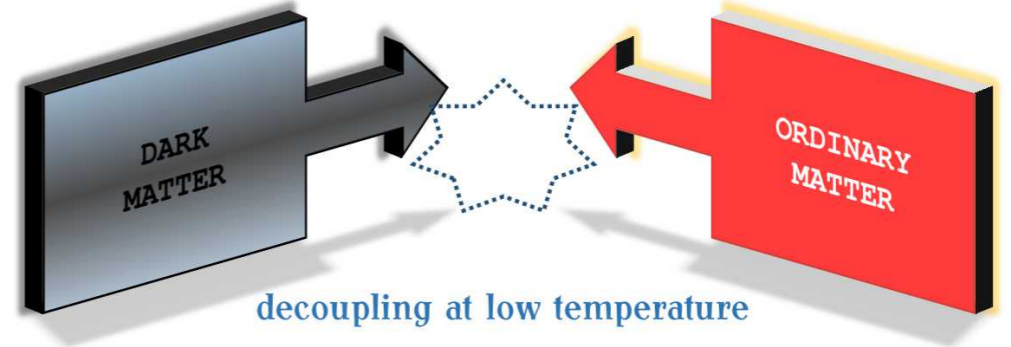


Light thermal dark matter

THE EARLY UNIVERSE

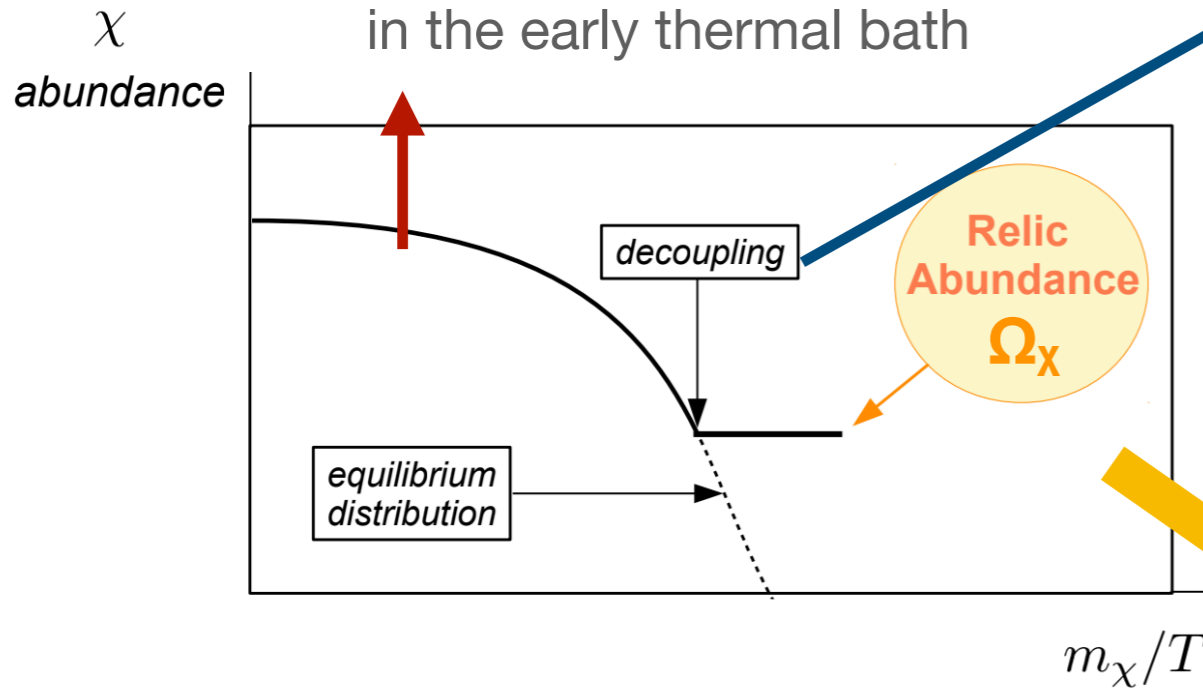


THE LATE UNIVERSE



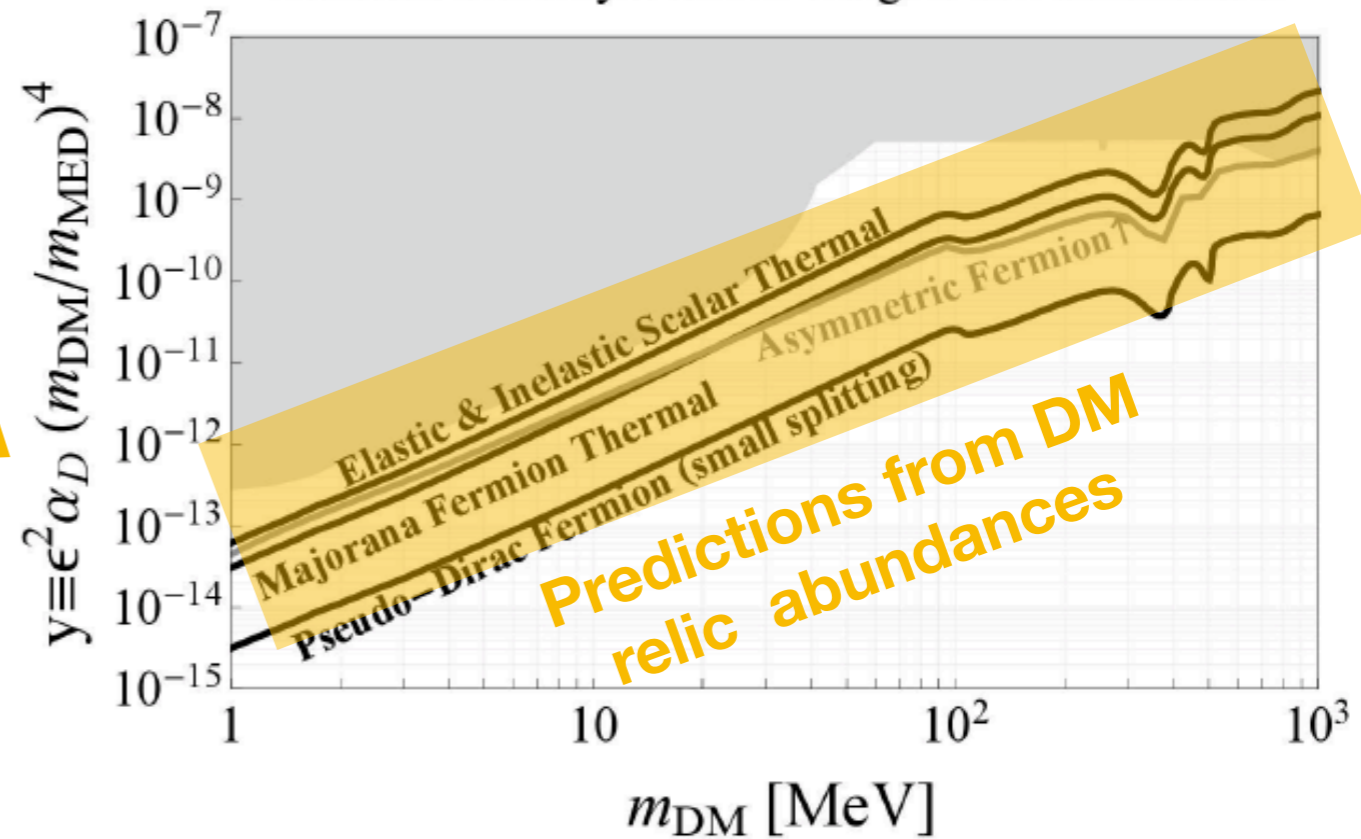
Freeze out

Non-gravitational interactions in the early thermal bath

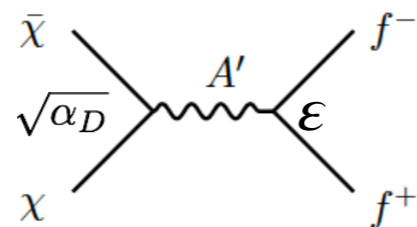


This implies a given annihilation rate to predict

Thermal and Asymmetric Targets at Accelerators



$$\Omega_\chi \propto \frac{1}{\langle \sigma v \rangle} \sim \frac{m_\chi^2}{g_\chi^4}$$



$$\sigma v(\chi\chi \rightarrow A'^* \rightarrow ff) \propto \epsilon^2 \alpha_D \frac{m_\chi^2}{m_{A'}^4} = \frac{y}{m_\chi^2}$$

J.Feng, J. Kumar

Phys. Rev. Lett.101231301

Light Dark Matter experiment arXiv:1808.05219v1

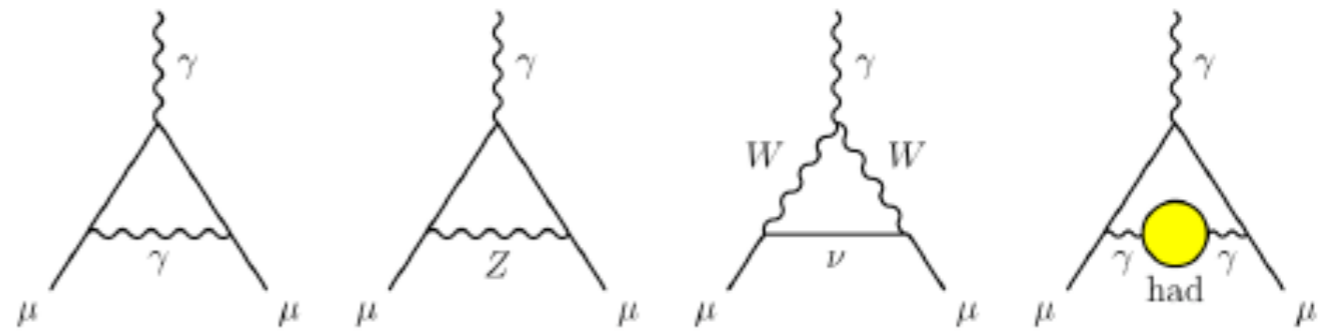
See N. Toro's and L. Tompkins's talks!



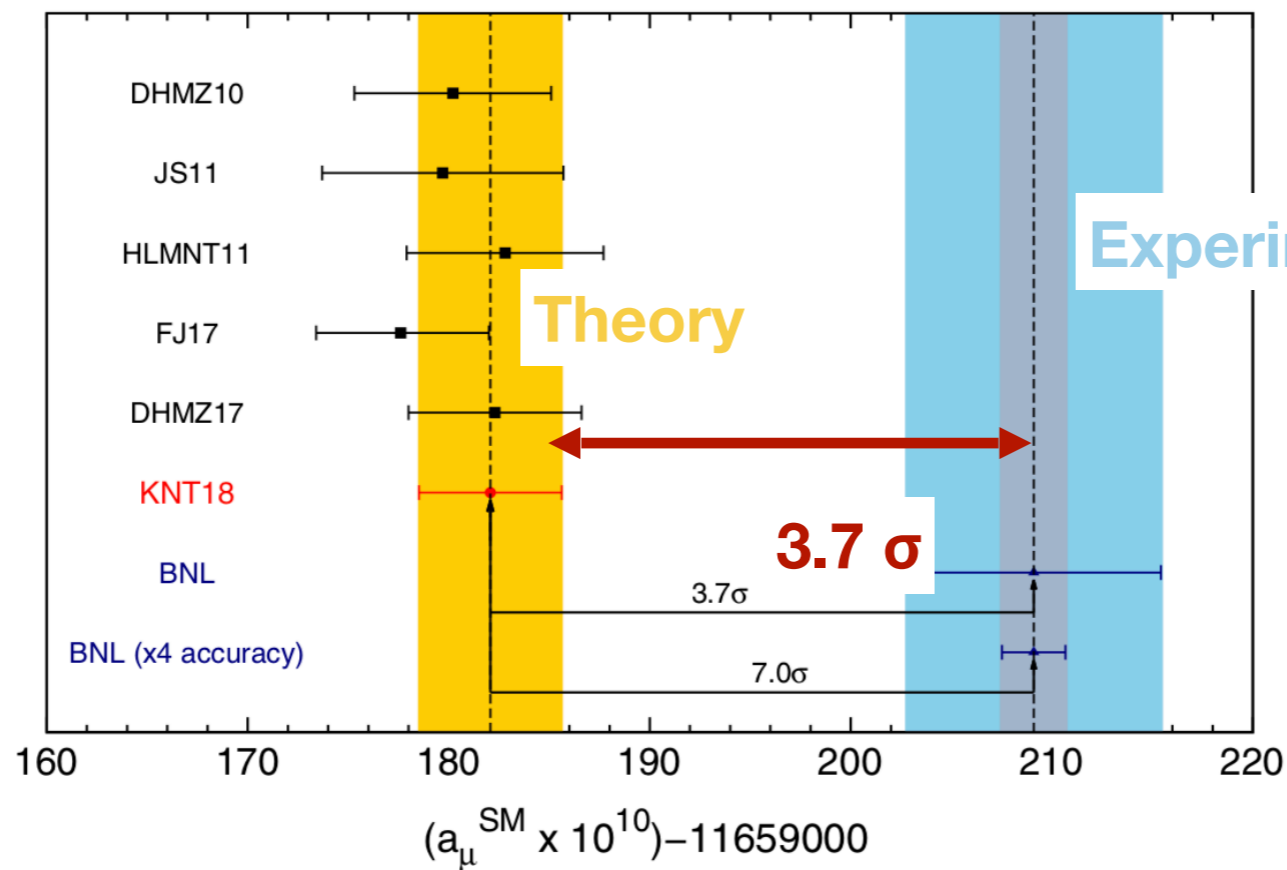
$(g-2)_\mu$ anomaly: an additional motivation

$$a_\mu = \frac{g_\mu - 2}{2}$$

Anomalous muon magnetic moment



$$a_\mu^{TH} = a_\mu^{QED} + a_\mu^{EW} + a_\mu^{HAD}$$

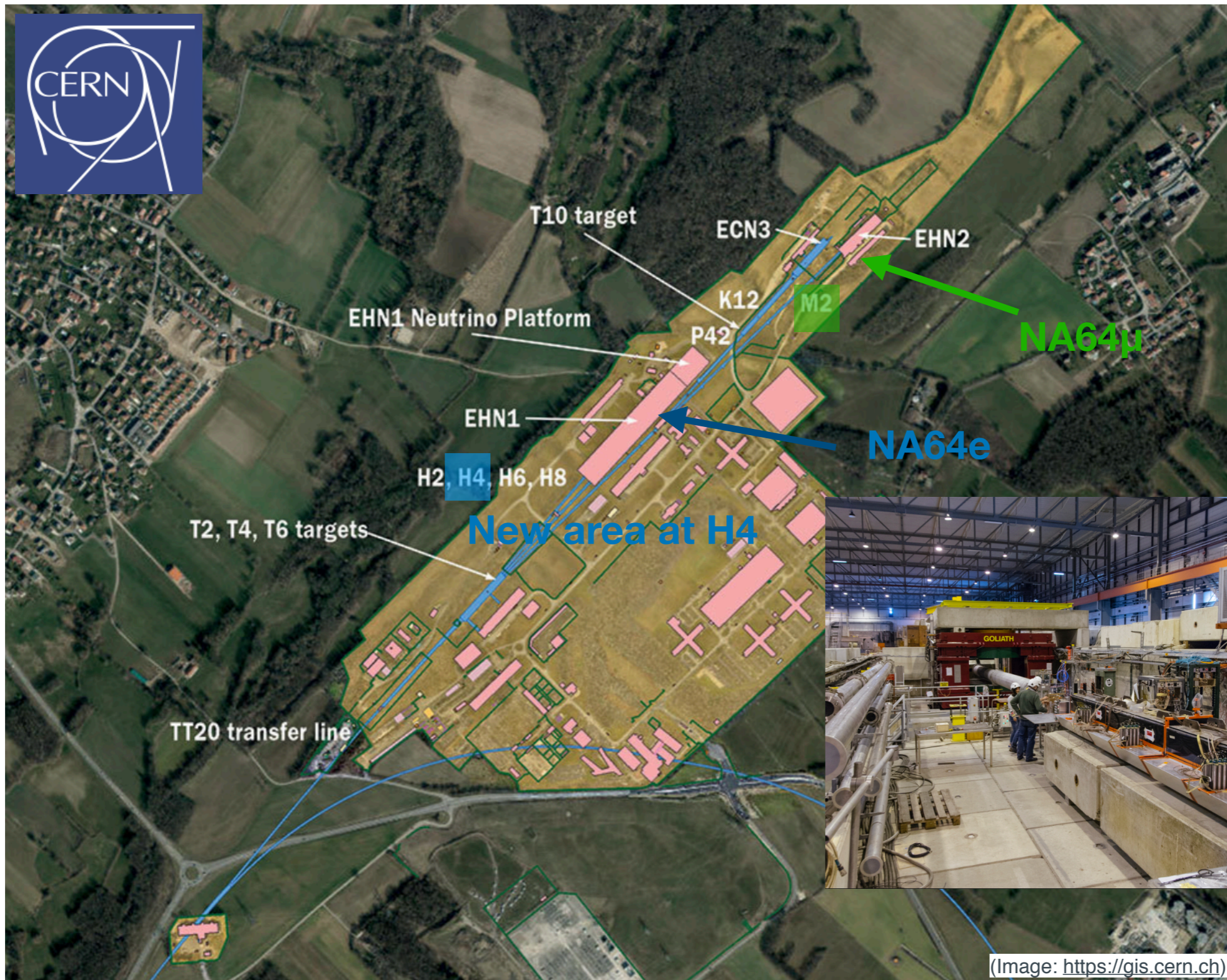


New experiment at Fermilab
<http://muon-g-2.fnal.gov/2-the-physics-of-g-2.html>

Keshavarzi, A., Nomura, D. & Teubner, T. *Phys. Rev. D* **97**, 114025 (2018).



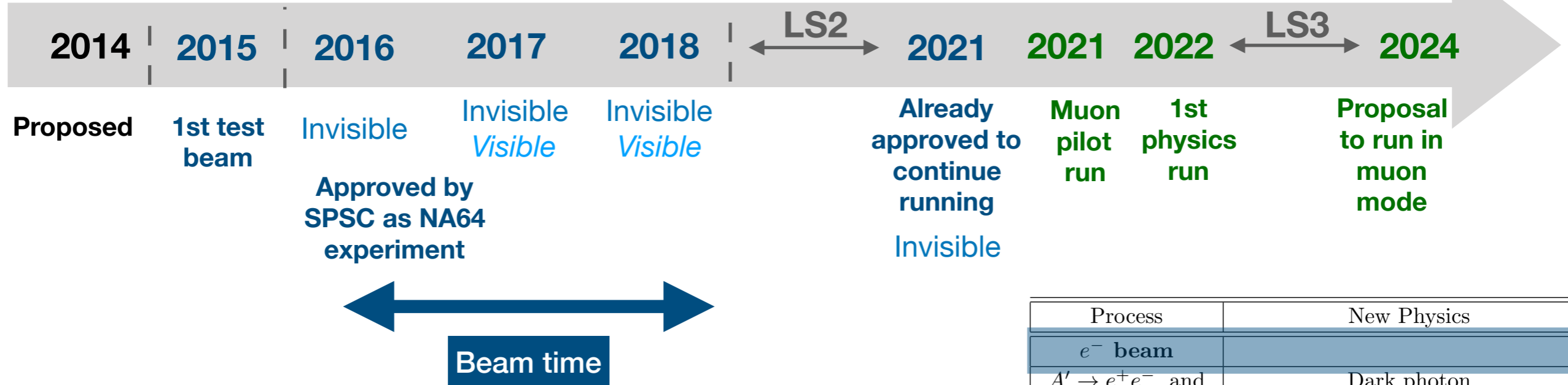
The NA64 experiment and its physics program





The NA64 experiment and its physics program

Fixed target experiment at the CERN SPS designed to probe Dark sector physics



Broad physics program →

International collaboration: **46 researchers** from **14 institutions**



CERN-PBC-REPORT-2018-007



Process	New Physics
e^- beam	
$A' \rightarrow e^+e^-$, and $A' \rightarrow invisible$ $A' \rightarrow \chi\bar{\chi}$	Dark photon sub-GeV Dark Matter (χ)
$X \rightarrow e^+e^-$ milliQ particles $a \rightarrow \gamma\gamma, invisible$	new gauge X - boson Dark Sector, charge quantisation Axion-like particles
μ^- beam	
$Z_\mu \rightarrow \nu\nu$ $Z_\mu \rightarrow \chi\bar{\chi}$ milliQ $a_\mu \rightarrow invisible$ $\mu - \tau$ conversion	gauge Z_μ -boson of $L_\mu - L_\tau$, $< 2m_\mu$ $L_\mu - L_\tau$ charged Dark Matter (χ) Dark Sector, charge quantisation non-universal ALP coupling Lepton Flavour Violation
π^-, K^- beams	
$\pi^0 \rightarrow invisible$ $\eta \rightarrow invisible$ $\eta' \rightarrow invisible$ $K_S^0 \rightarrow invisible$ $K_L^0 \rightarrow invisible$	Current limits, PDG'2018 $Br(\pi^0 \rightarrow invisible) < 2.7 \times 10^{-7}$ $Br(\eta \rightarrow invisible) < 1.0 \times 10^{-4}$ $Br(\eta' \rightarrow invisible) < 5 \times 10^{-4}$ no limits no limits
e^+ beam	

Resonant A' production



NA64 technique for A' decays and its signatures

Initial e^- beam

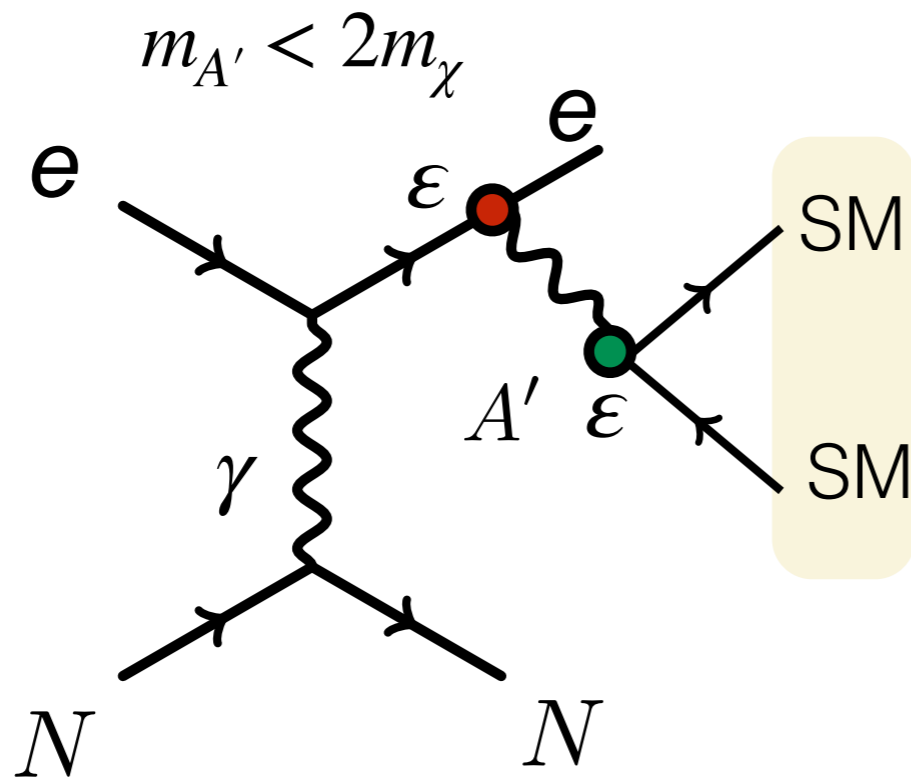


Active Dump
 A'

Focus of my talk

Setup:

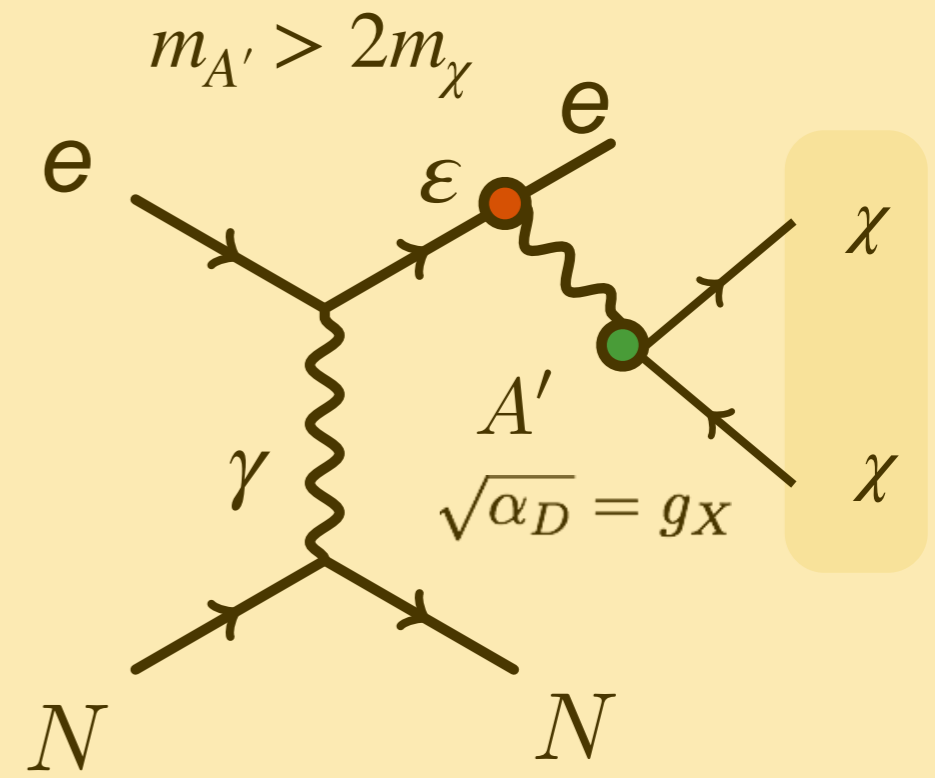
Visible mode



Signature:

SM particles
pair production

Invisible mode

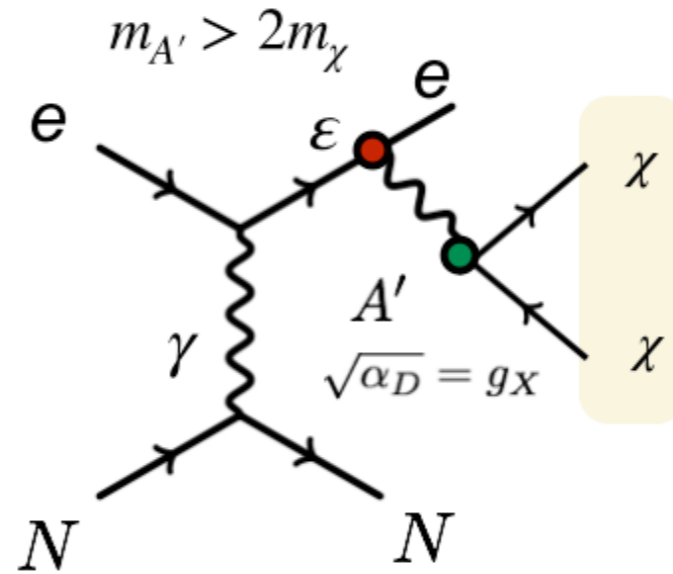


Missing energy

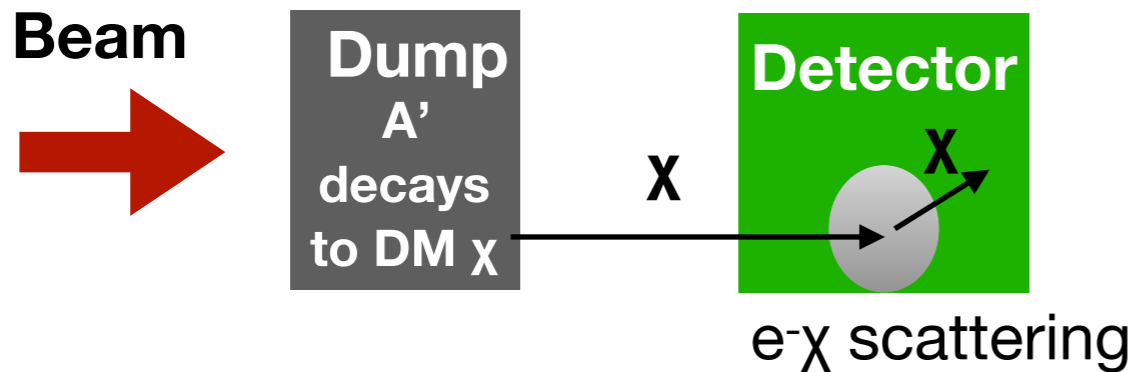


NA64 invisible searches

Invisible mode



1) BEAM DUMP APPROACH
(MiniBooNE, LSND, NA62,
SHIP, T2K, DUNE...)



Flux of X generated by decays of A's produced in the dump.

Signal: X scattering in far detector

$$\sigma \propto \epsilon^4 \alpha_D$$

2) NA64/LDMX APPROACH



Produced A's carry away energy from the active dump.

Signal: Missing energy/momentum

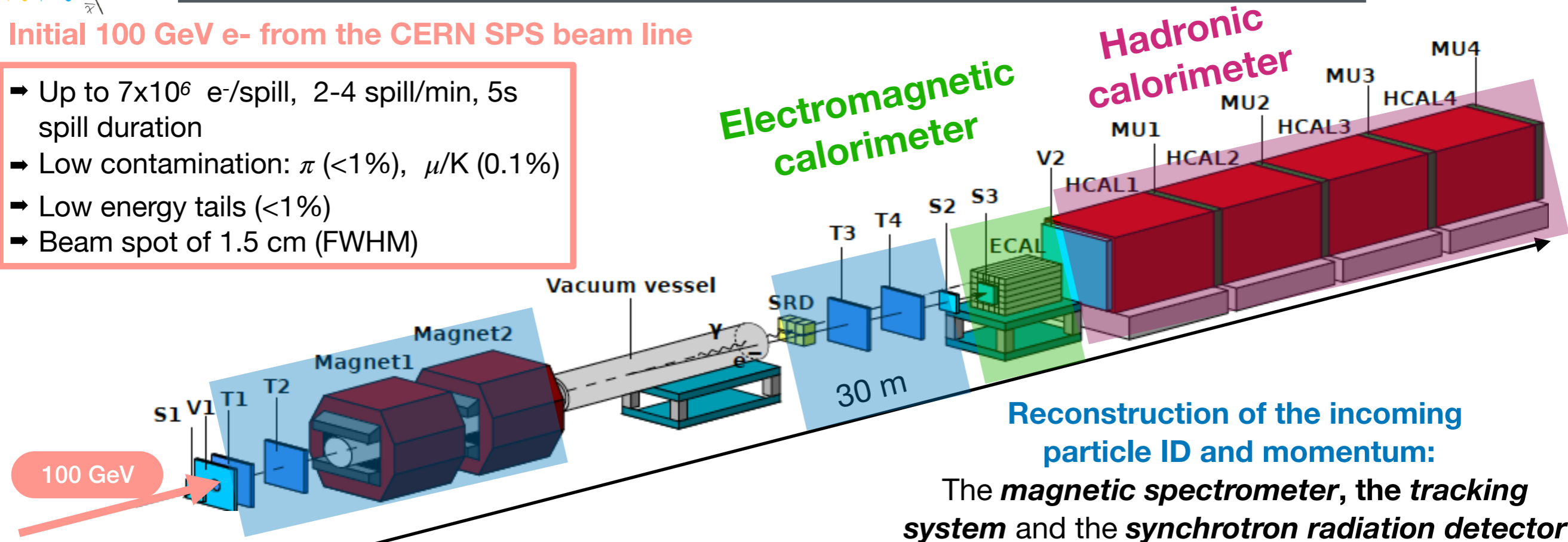
$$\sigma \propto \epsilon^2$$



NA64 invisible searches: the setup

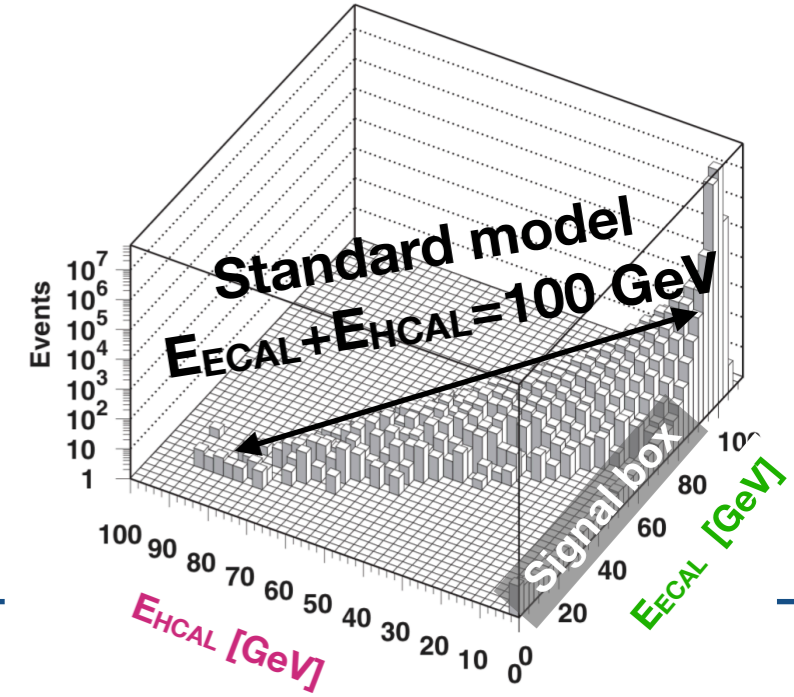
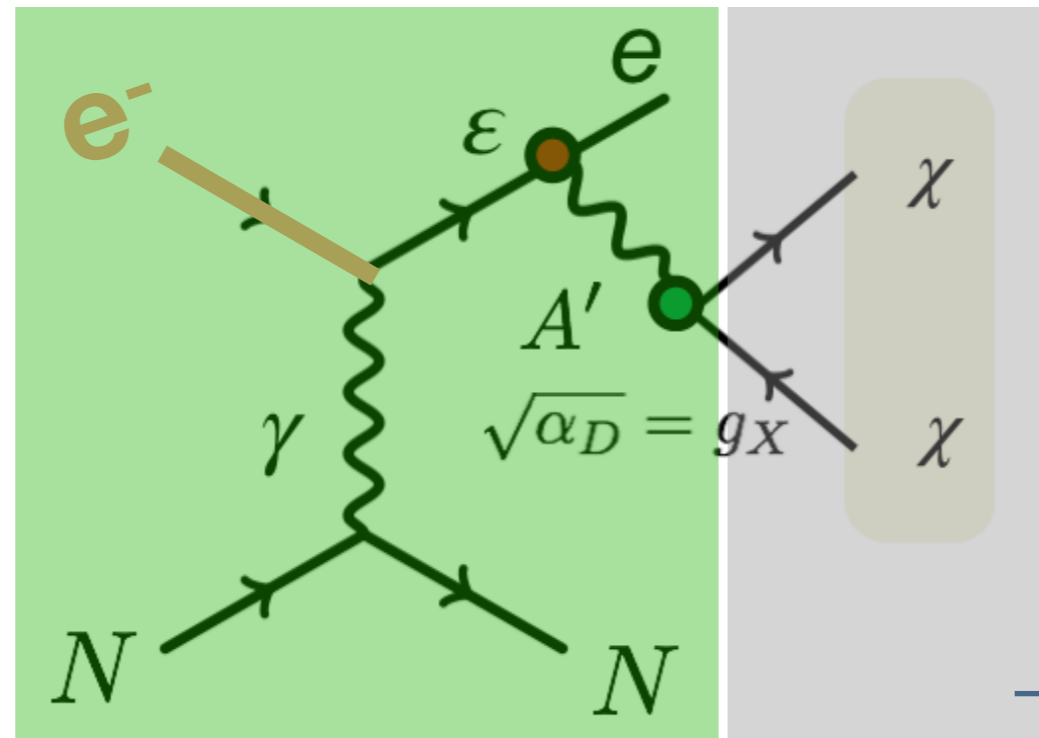
Initial 100 GeV e⁻ from the CERN SPS beam line

- Up to 7×10^6 e⁻/spill, 2-4 spill/min, 5s spill duration
- Low contamination: π (<1%), μ/K (0.1%)
- Low energy tails (<1%)
- Beam spot of 1.5 cm (FWHM)



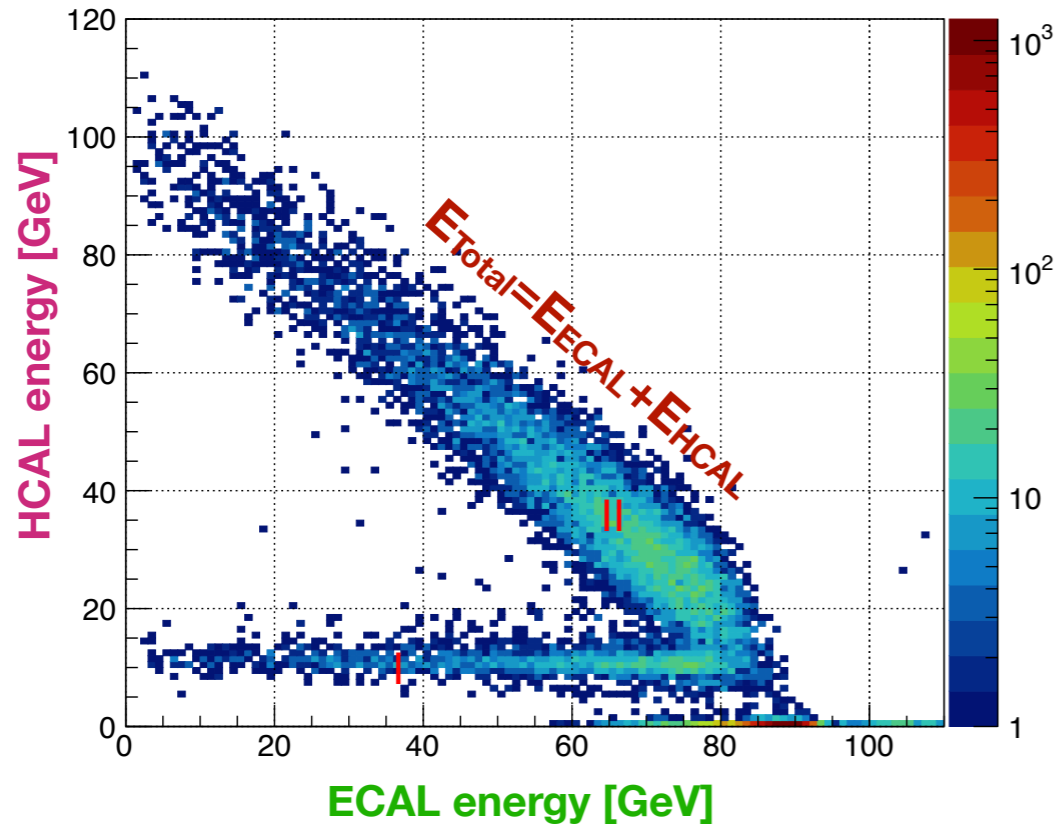
Missing energy: A' signal
 $E_{ECAL} < 50$ GeV
 $E_{HCAL} \text{ energy} < 2$ GeV

Active target



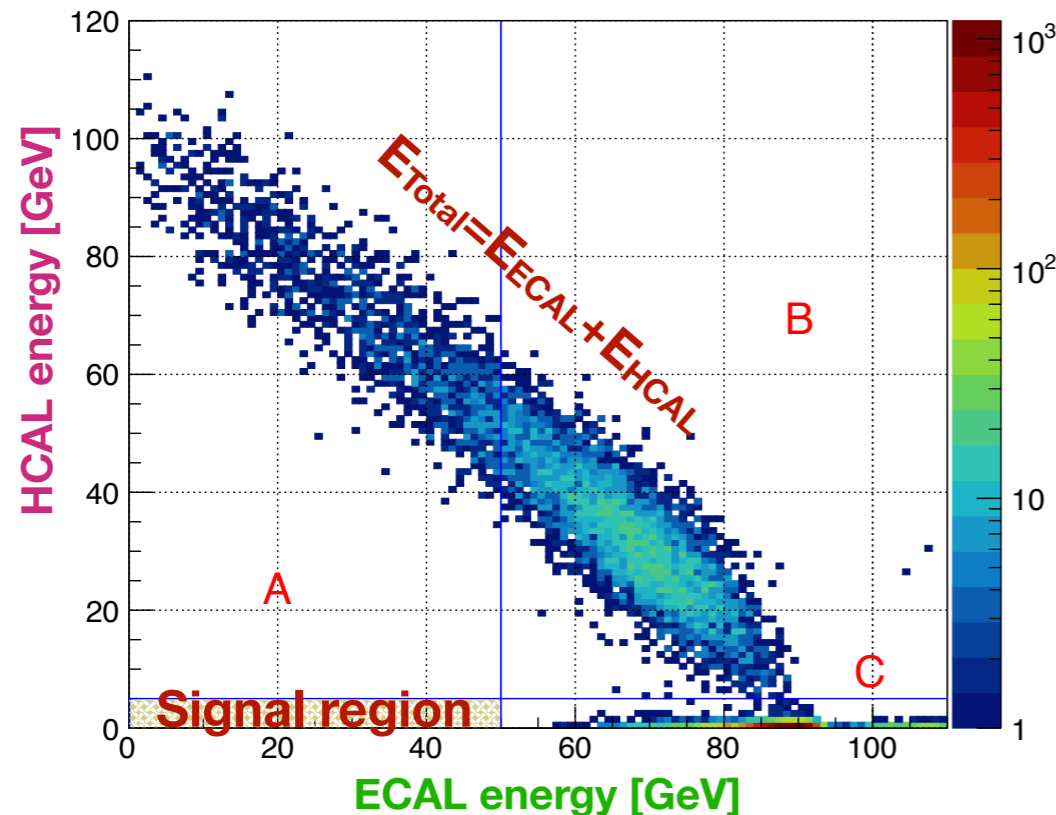


NA64 invisible searches: results



Full 2016-2018 data:
 2.84×10^{11} EOT

- ➔ **Region I:** $e^- Z \rightarrow e^- Z \gamma; \gamma \rightarrow \mu^+ \mu^-$
→ benchmark for MC
- ➔ **Region II:** SM events
 $E_{\text{ECAL}} + E_{\text{HCAL}} \approx 100 \text{ GeV}$



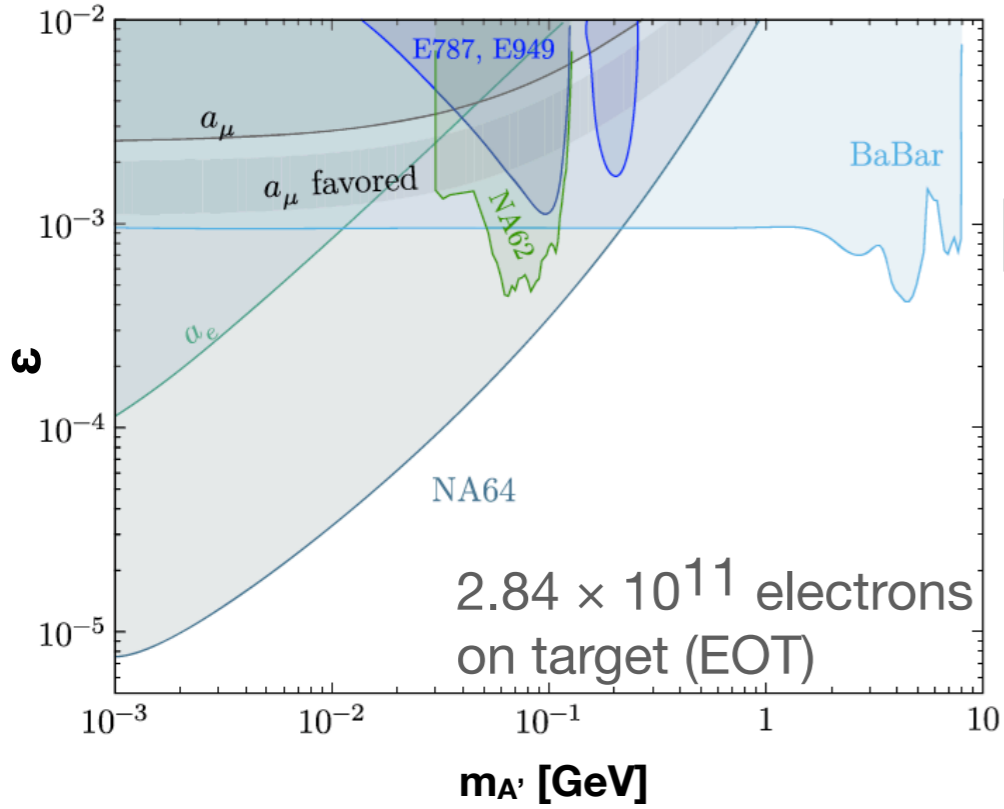
Event Selection Criteria:

- ◆ *Timing information* → Pile up and noise suppression.
- ◆ *Clean incoming track:* angle + single hit in all trackers, momentum $\sim 100 \text{ GeV}$
- ◆ *Electron identification:*
 - Synchrotron radiation
 - Shower profile compatible with e^- in ECAL → Hadron suppression
- ◆ *No punchthrough:* No activity in Veto and in HCAL



Results

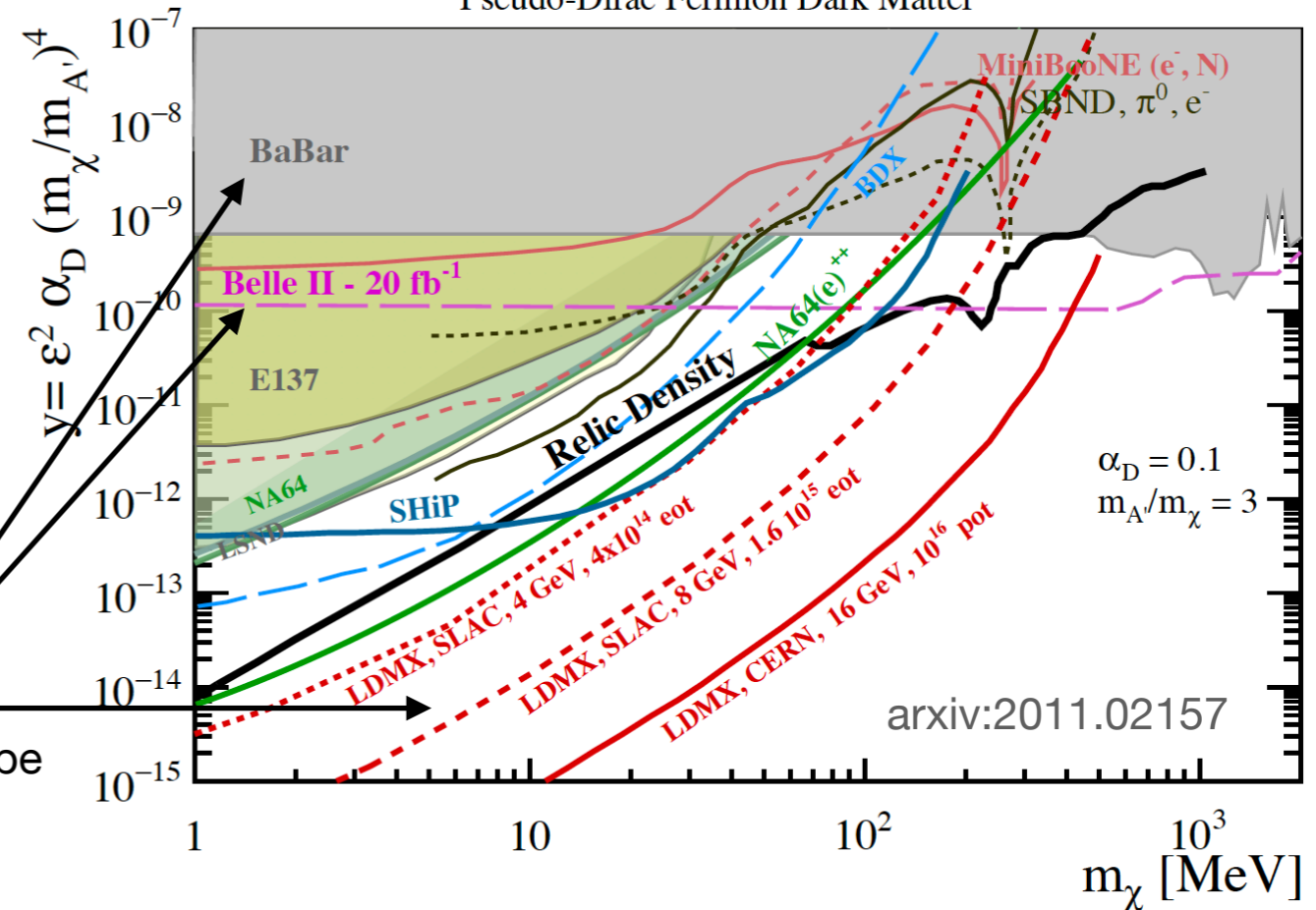
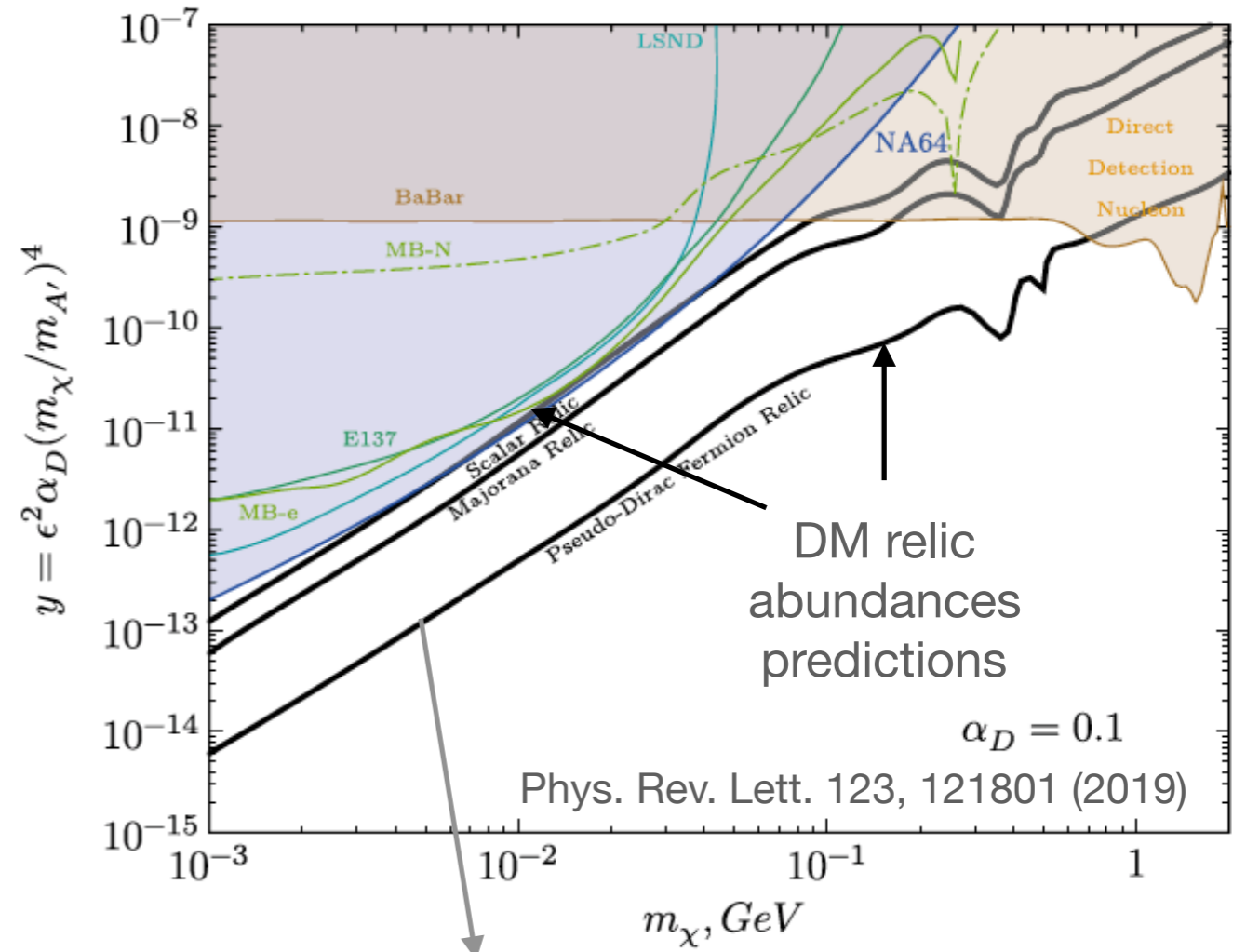
Constraints on LTDM models



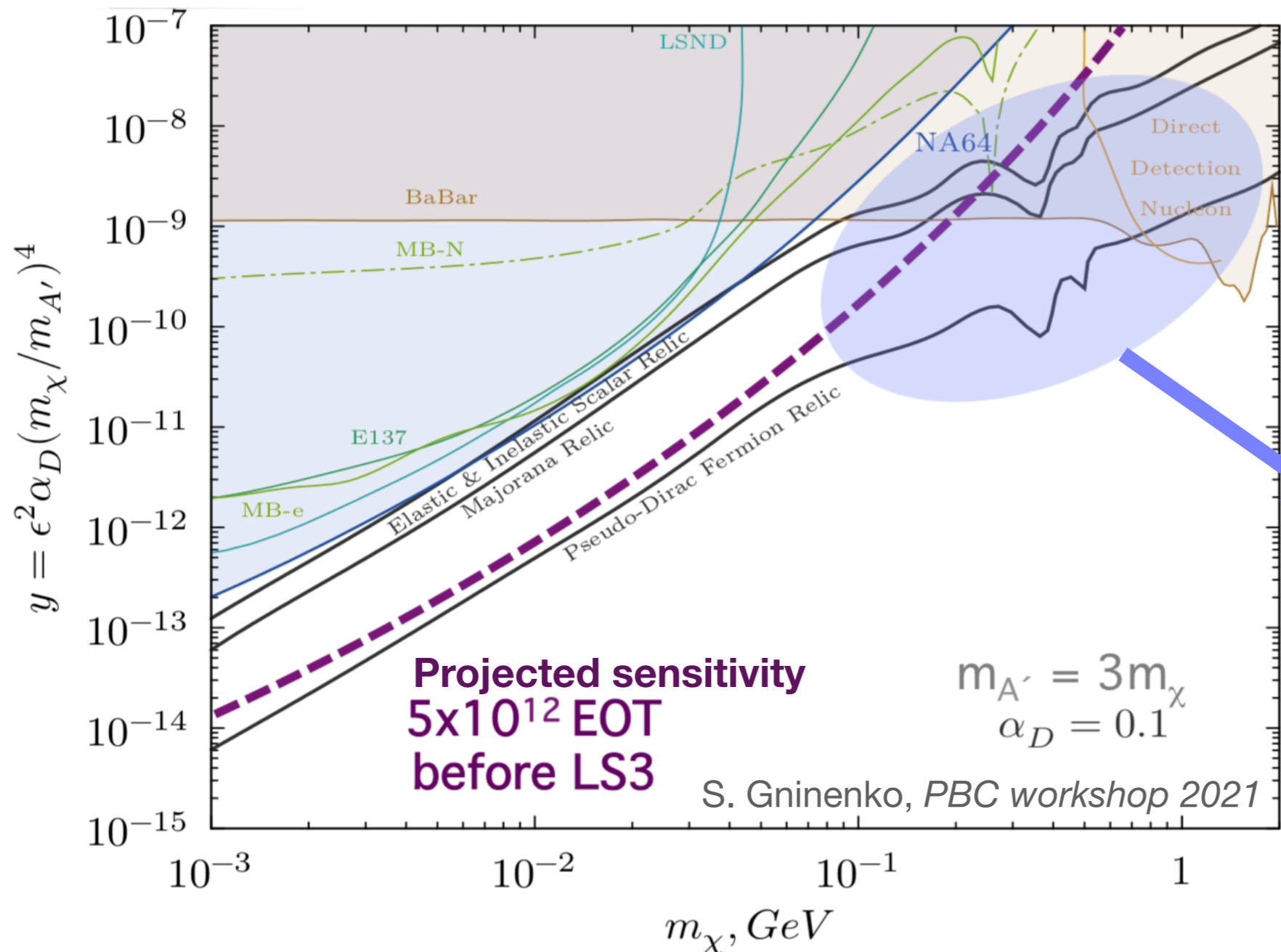
Proportional to **DM \leftrightarrow SM** annihilation cross-section

For the first time NA64 constraints better than in previous experiments and very close to LTDM benchmark models

See talks by B.Shuve, L. Tompkins and J.Strube



Future prospects in 2021



Resume data taking this year

- New fixed location at H4 beam line
- Beam and setup upgrade

How can we enlarge the sensitivity at higher masses?

New ideas:

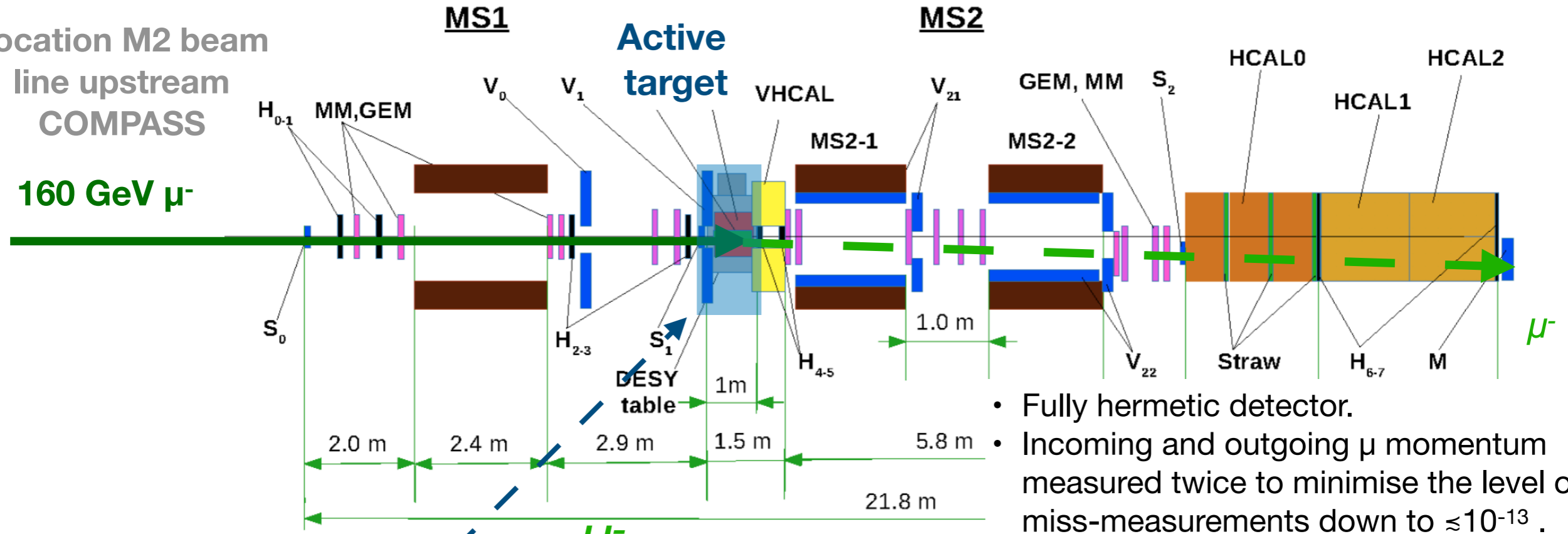
- **A' resonant production**
L. Marsicano et al. Phys. Rev. Lett. 121, 041802
- **Use a muon beam**
S.Gninenko et al. PLB796, 117 (2019)

Future prospects: NA64e + NA64 μ

Exploring dark sector physics weakly coupled to muons

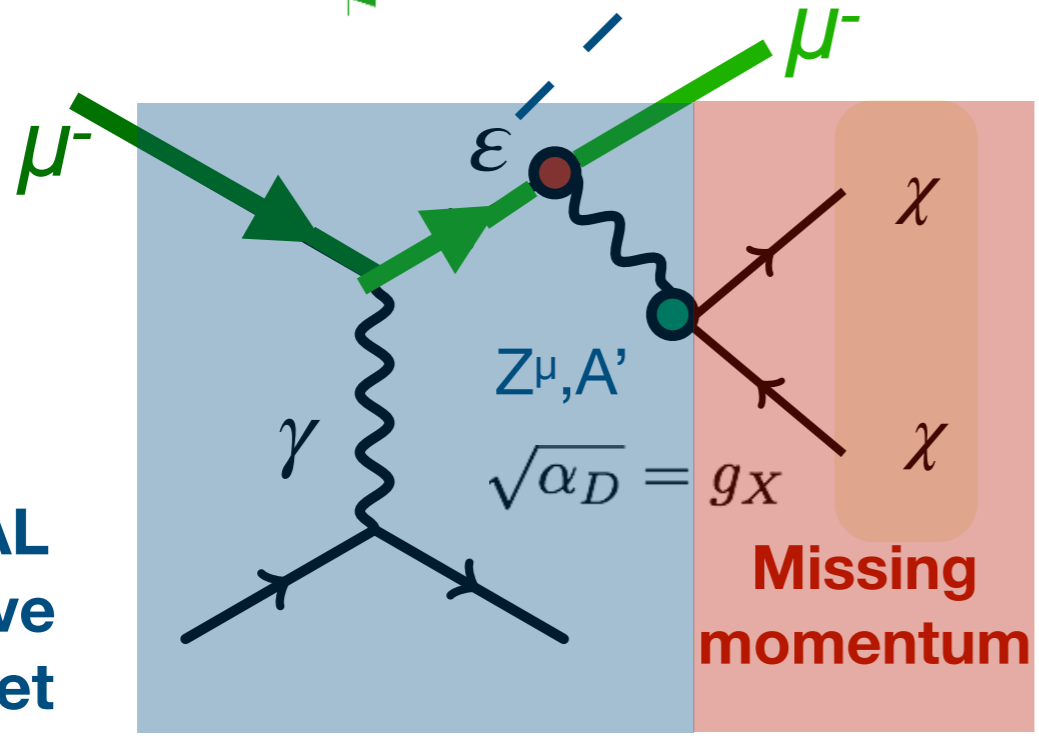
Location M2 beam line upstream COMPASS

160 GeV μ^-



- Fully hermetic detector.
- Incoming and outgoing μ momentum measured twice to minimise the level of its miss-measurements down to $\lesssim 10^{-13}$.

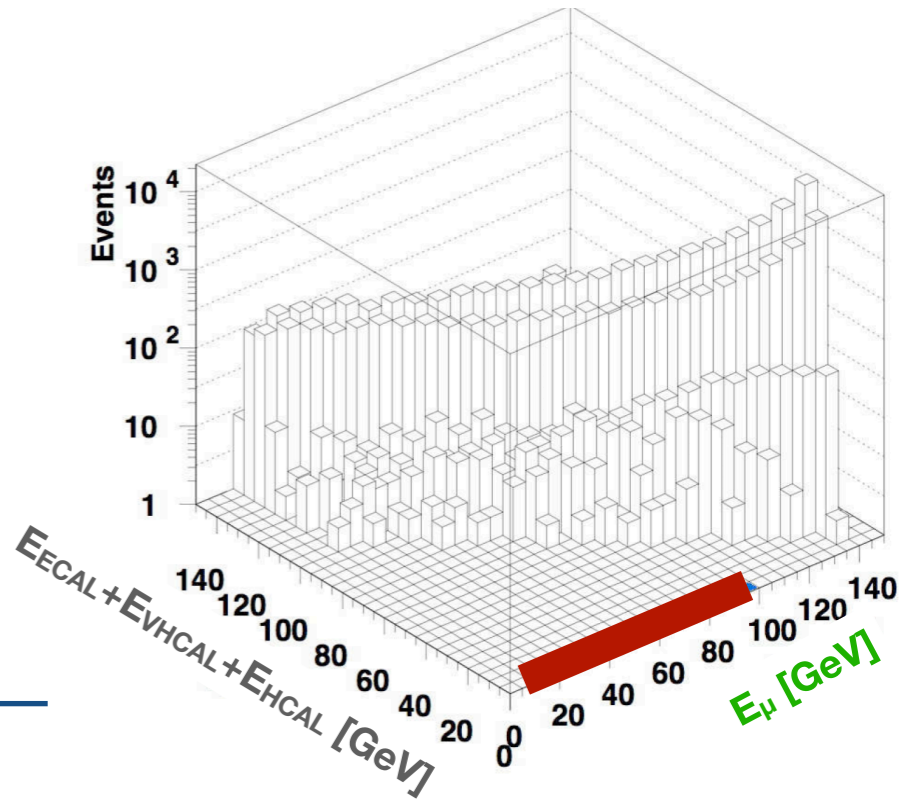
ECAL active target



Z^{μ}, A' decaying to DM particles

Signature

- Missing momentum (Deflected μ^- energy < 80 GeV).
- No energy on ECAL, VHCAL and HCAL (meaning compatible with a MIP energy).





Future prospects: NA64_μ physics goals

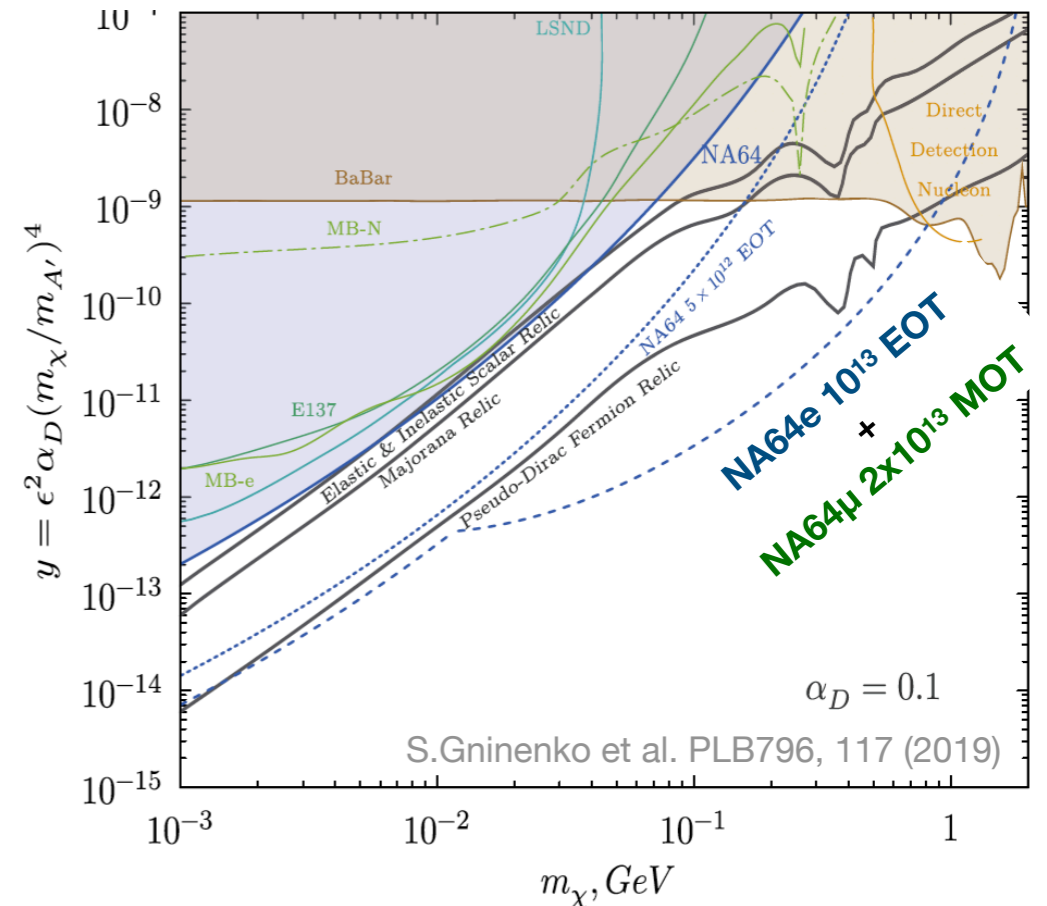
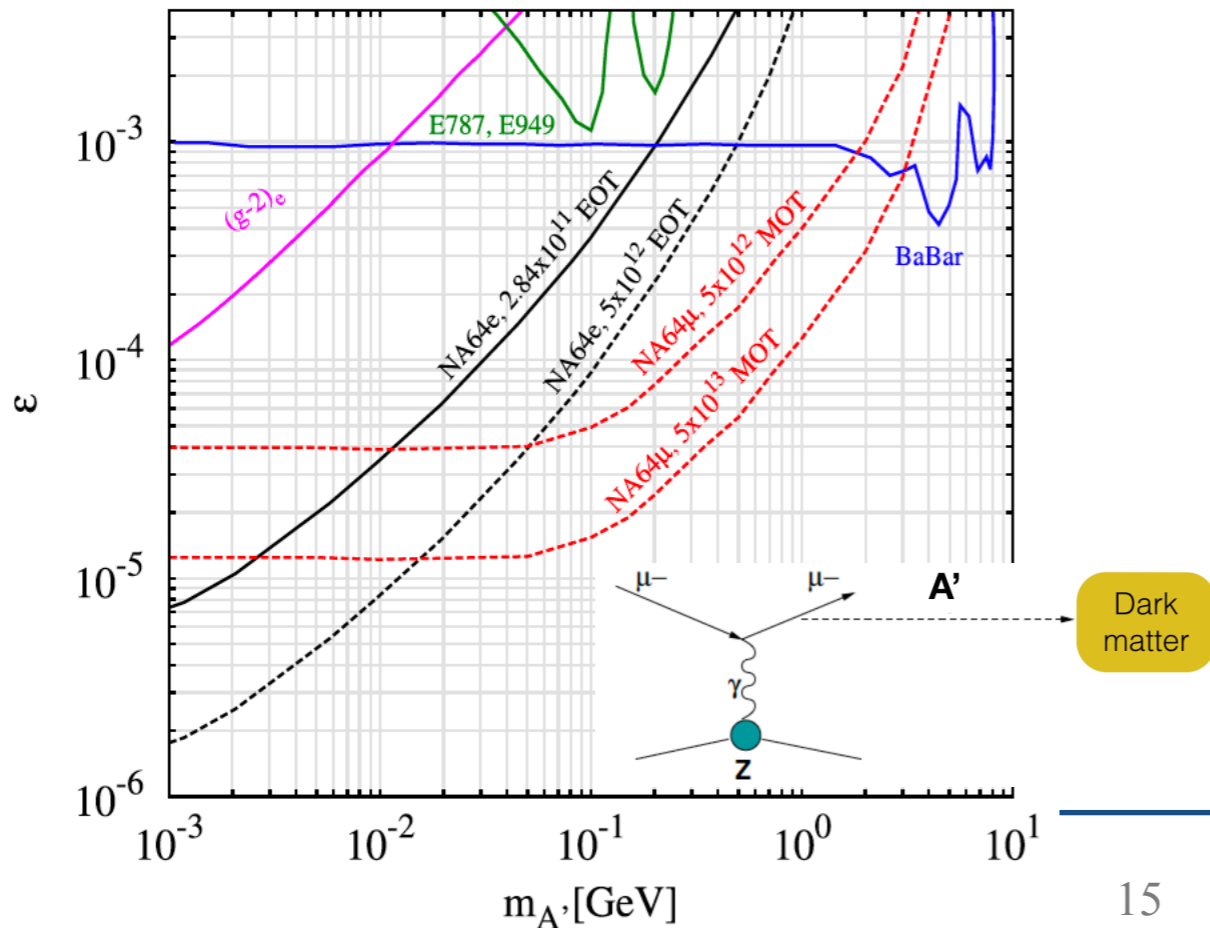
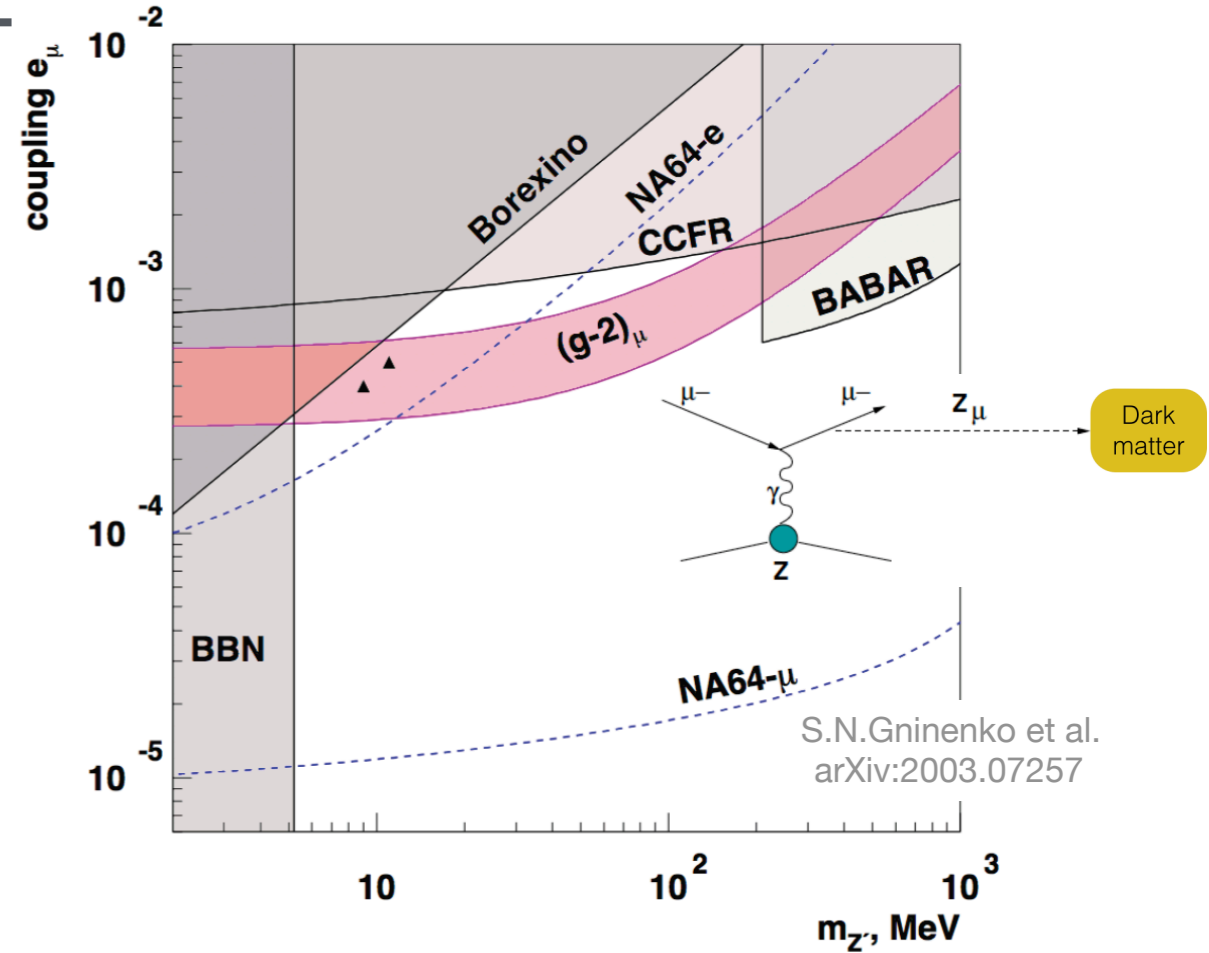
1. **Light Z' coupled to the muon**, as a remaining low mass explanation of the $(g-2)_\mu$ (the muon anomaly).
2. **Light Dark Matter** interacting with the Standard Matter via dark photon A' in the A' mass region ≥ 0.1 GeV (complementary search to NA64e).
3. **Scalar, ALPs** coupled to the muon, **millicharged** particles,
4. **Lepton Flavour Violation** in $\mu Z \rightarrow \tau Z$ conversion in flight.



Already approved Feasibility studies

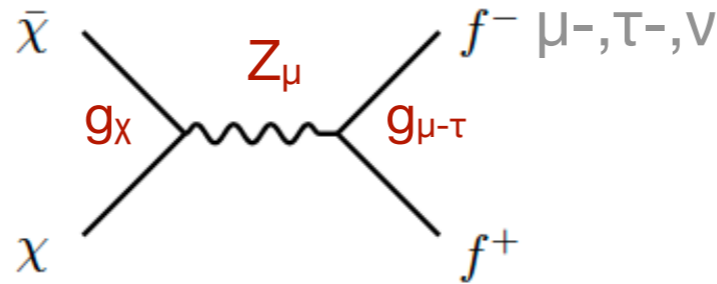
Expected first physics runs to probe Z' as $(g-2)_\mu$ explanation

cover high A', Z' masses



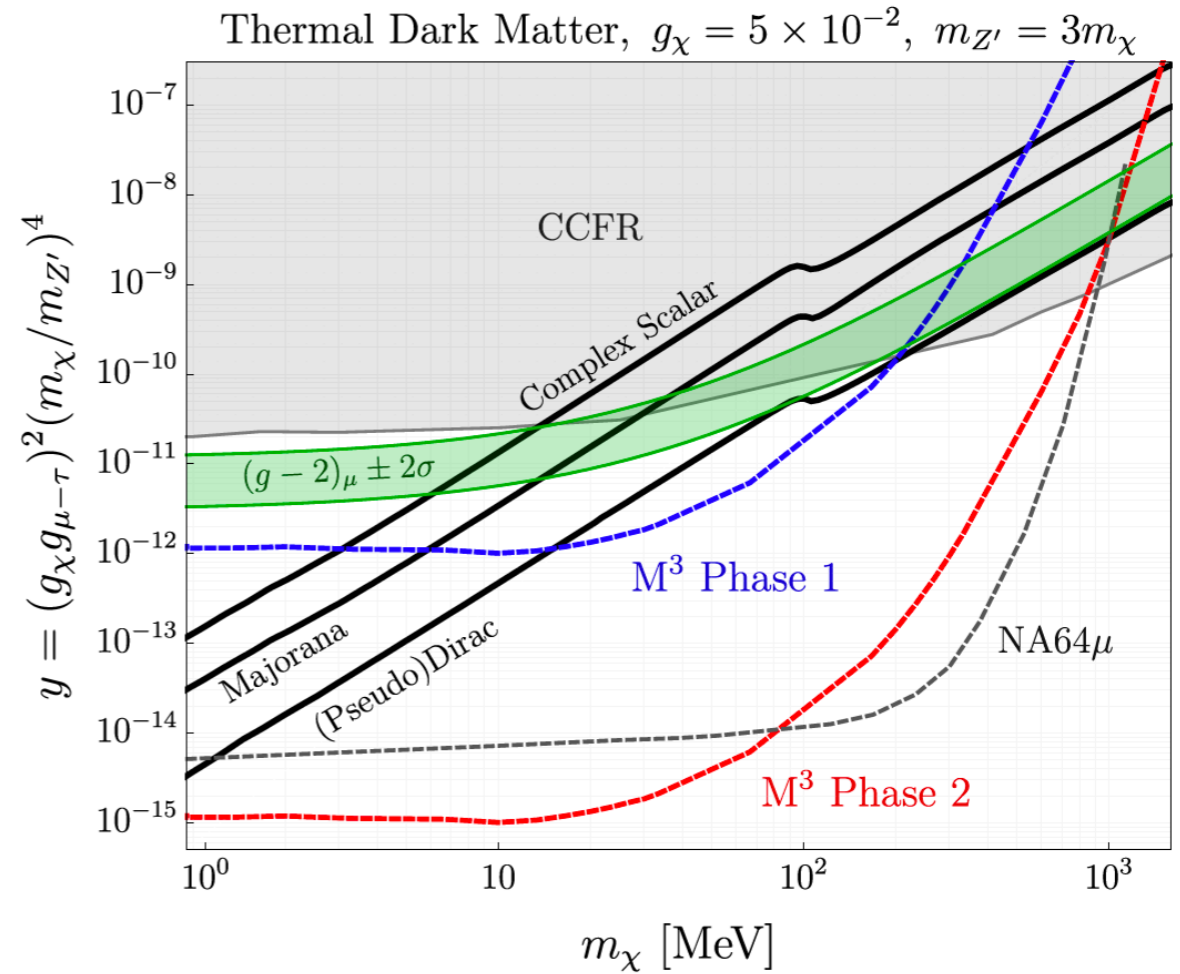
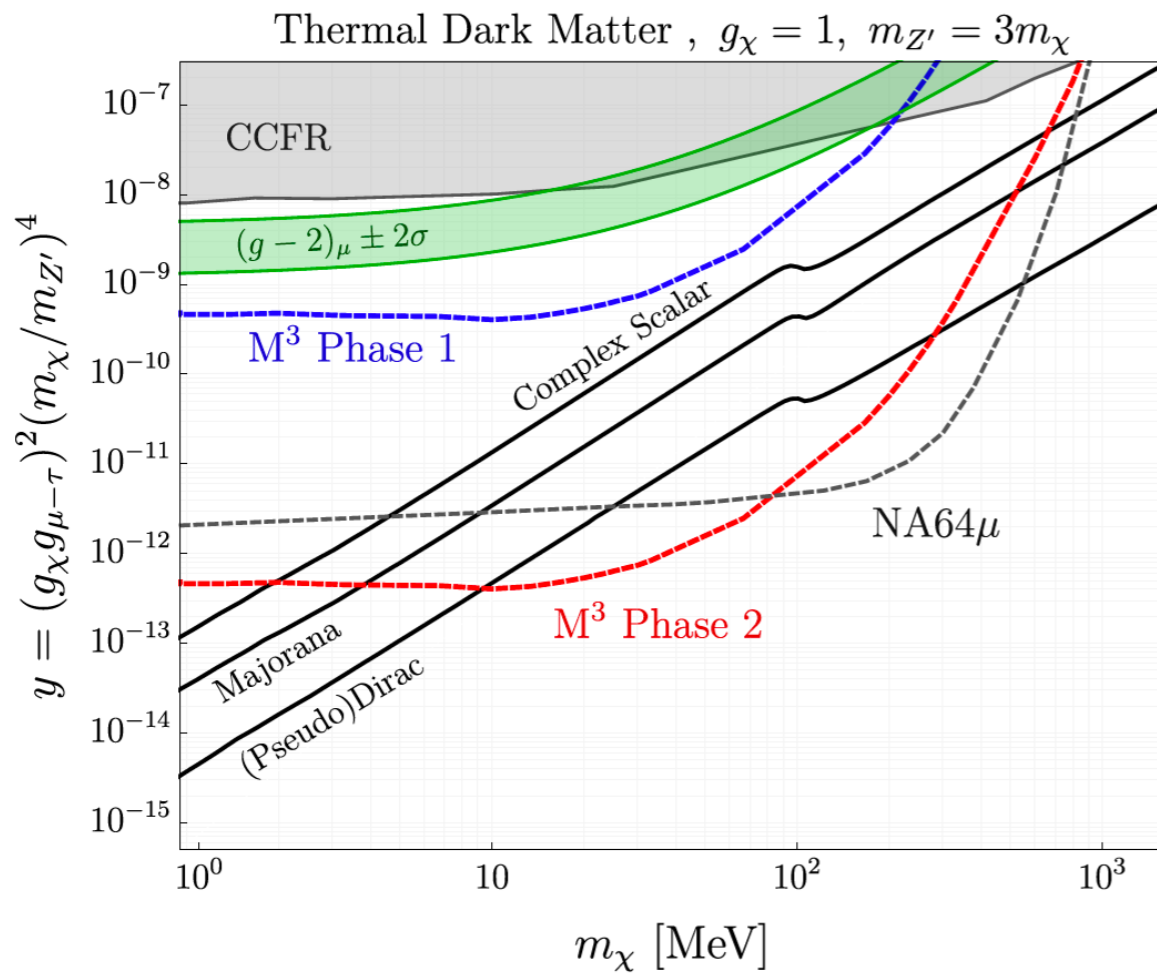
Future prospects: NA64_μ physics goals

Dark Matter generated via a new massive Z_μ boson from broken $U(1)'_{L\mu-L\tau}$



$$\langle \sigma v \rangle \propto g_\chi^2 g_{\mu-\tau}^2 \frac{m_\chi^2}{m_{Z'}^4} = \frac{y}{m_\chi^2}$$

$$m_{Z_\mu} > 2m_\chi: Z_\mu \rightarrow \chi\chi$$



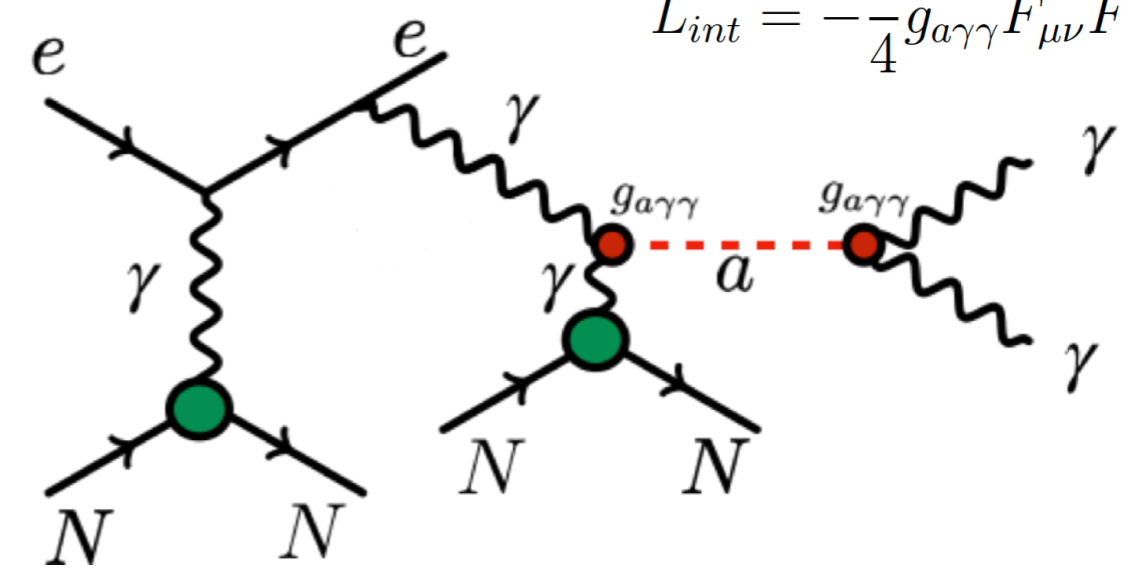
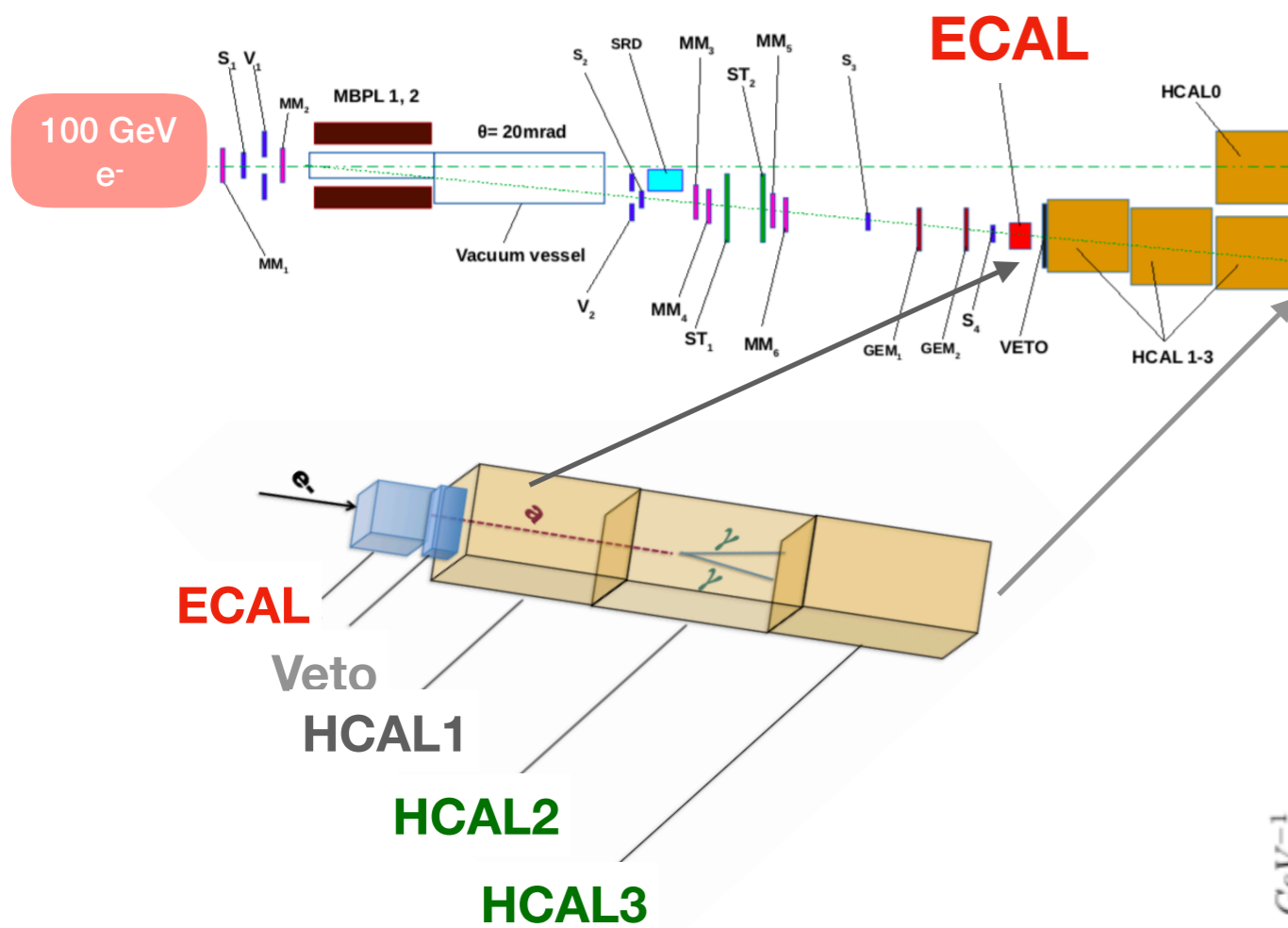
"M³: A New Muon Missing Momentum Experiment to Probe $(g-2)_\mu$ and Dark Matter at Fermilab", arxiv:1804.03144

ALPs searches @NA64

ALPs predominantly coupled to photons produced via Primakoff effect

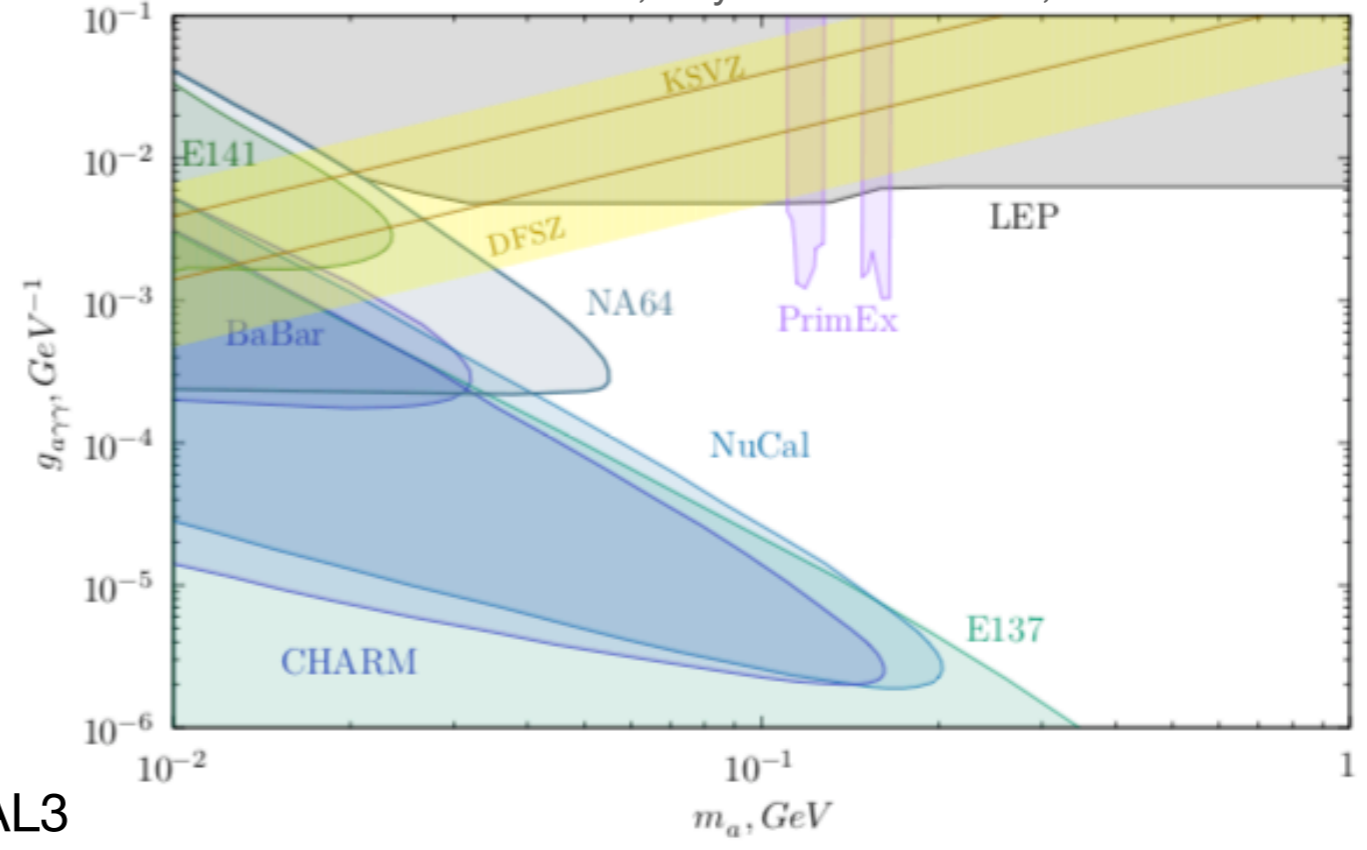
$$L_{int} = -\frac{1}{4}g_{a\gamma\gamma}F_{\mu\nu}\tilde{F}^{\mu\nu}a$$

NA64 invisible setup



Main goal: to probe the gap in the parameter space between the beam-dump and LEP searches

NA64 collaboration, Phys. Rev. Lett. 125, 081801



Signature:

No signal on veto and HCAL1

+

- A. **Visible Decay into $\gamma\gamma$** on HCAL2 || HCAL3
- B. **Decays after HCAL3:** no activity on HCAL2 & HCAL3

Constraints on new physics in $(g-2)_e$

$$\Delta a_e = a_e^{exp} - a_e^{LKB} = (4.8 \pm 3.0) \times 10^{-13} \quad (1.6\sigma)$$

$$\Delta a_e = a_e^{exp} - a_e^B = (-8.8 \pm 3.6) \times 10^{-13} \quad (-2.4\sigma)$$

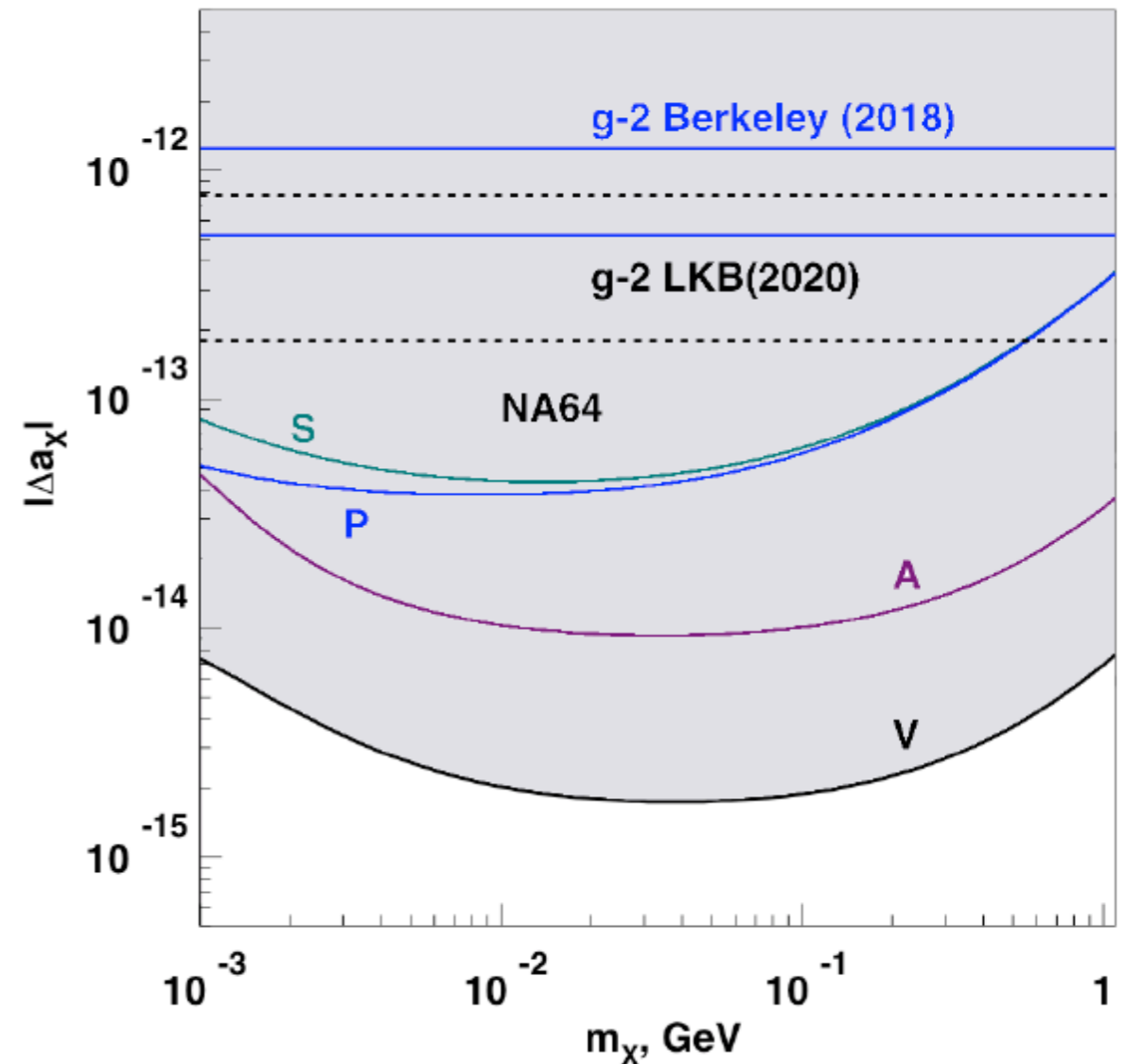
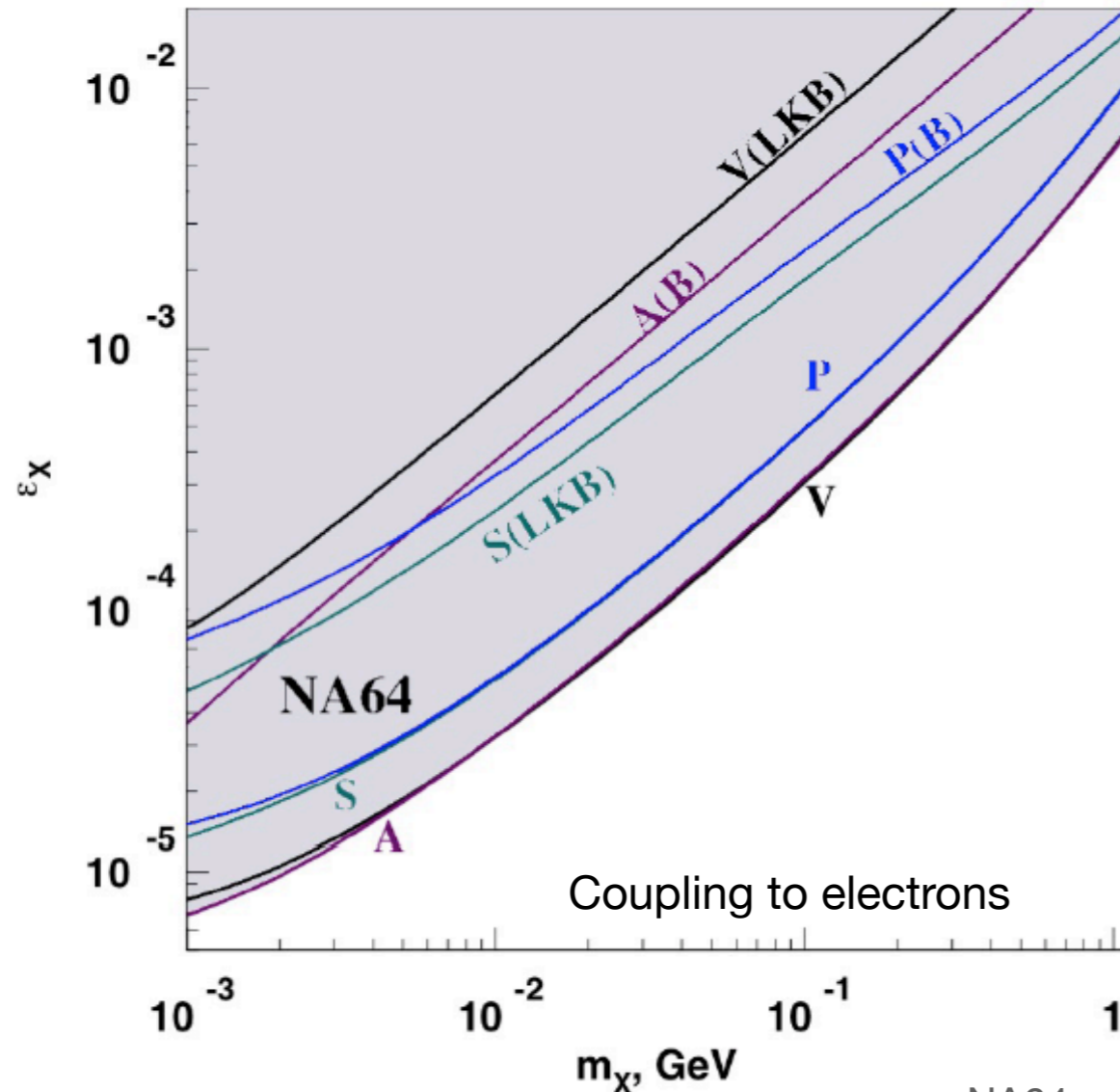
Results from high precision measurements of α at LKB and Berkley

Differences due to unknown experimental error at present or might be an indication of new physics?

Could a new generic boson contribute to $(g-2)_e$?

$e^-Z \rightarrow e^-ZX; X \rightarrow \text{invisible}$ X could be S,P, V or A

1-loop contributions of S,P,V or A to Δa_e



NA64 collaboration, arXiv:2102.01885



NA64 visible searches

Initial e⁻ beam

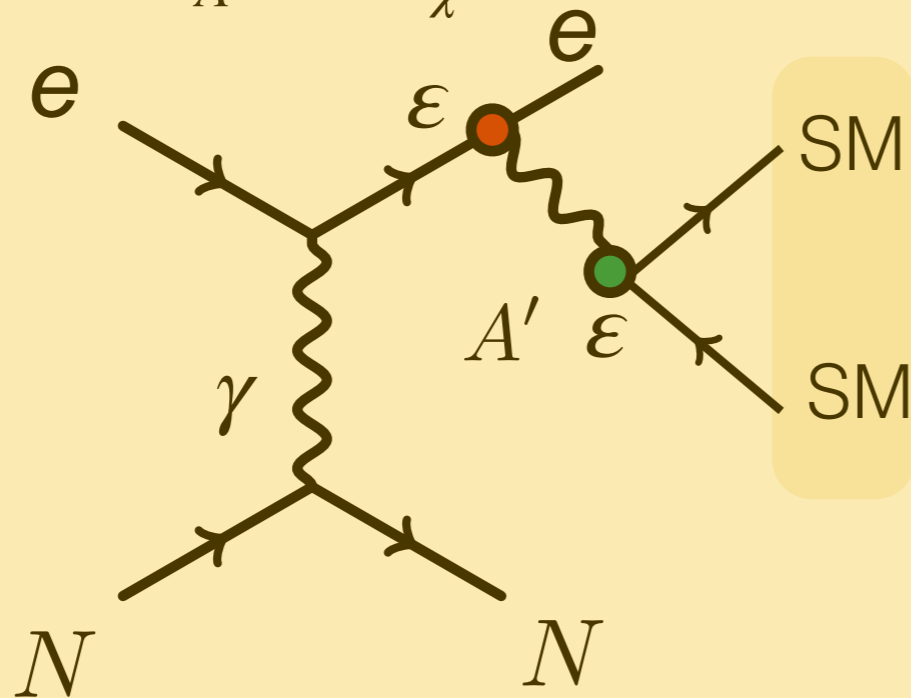


Active Dump
A'

Setup:

Visible mode

$$m_{A'} < 2m_\chi$$

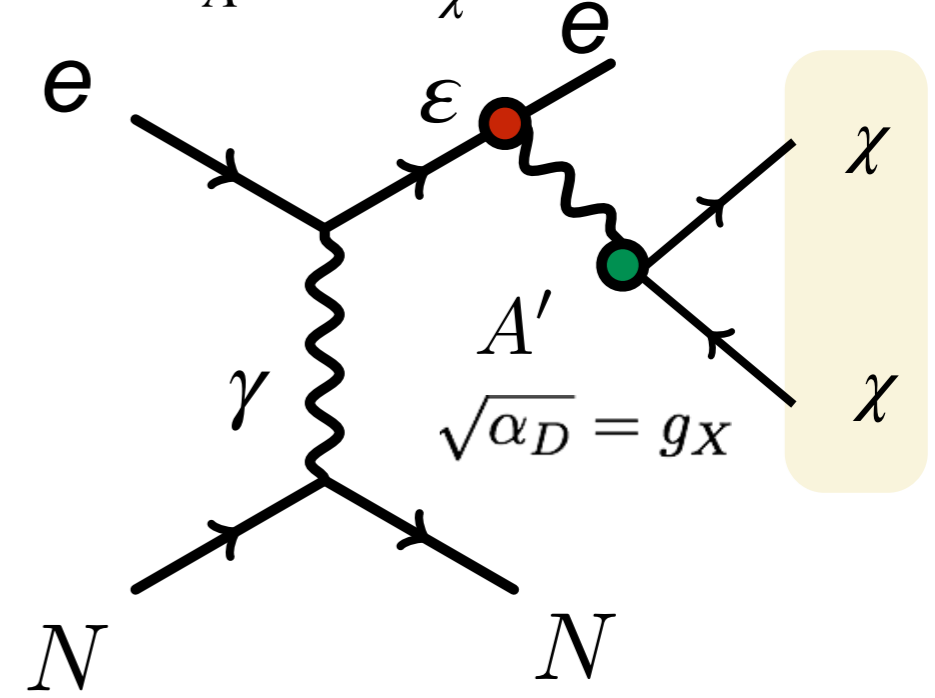


Signature:

SM particles
pair production

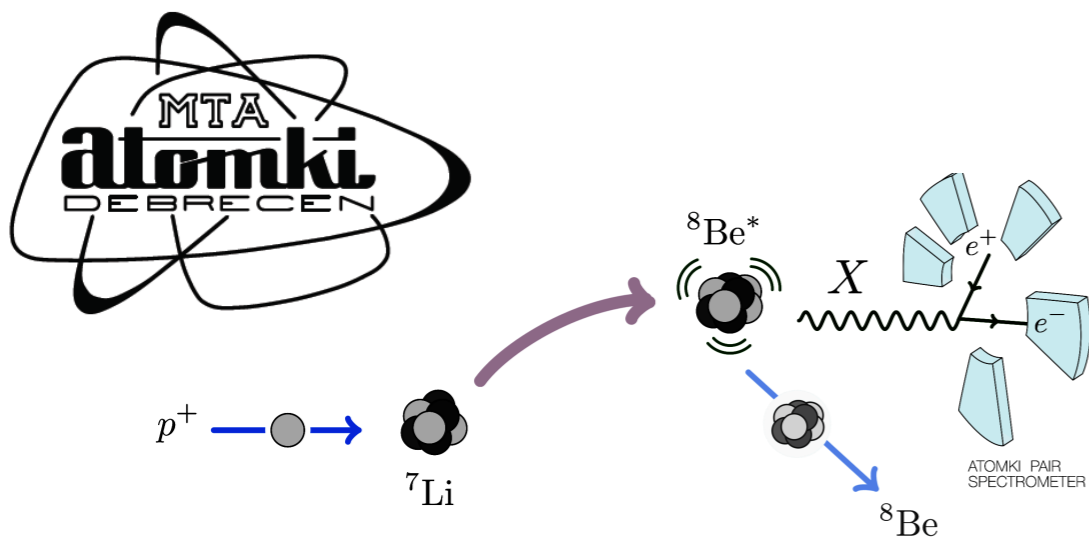
Invisible mode

$$m_{A'} > 2m_\chi$$

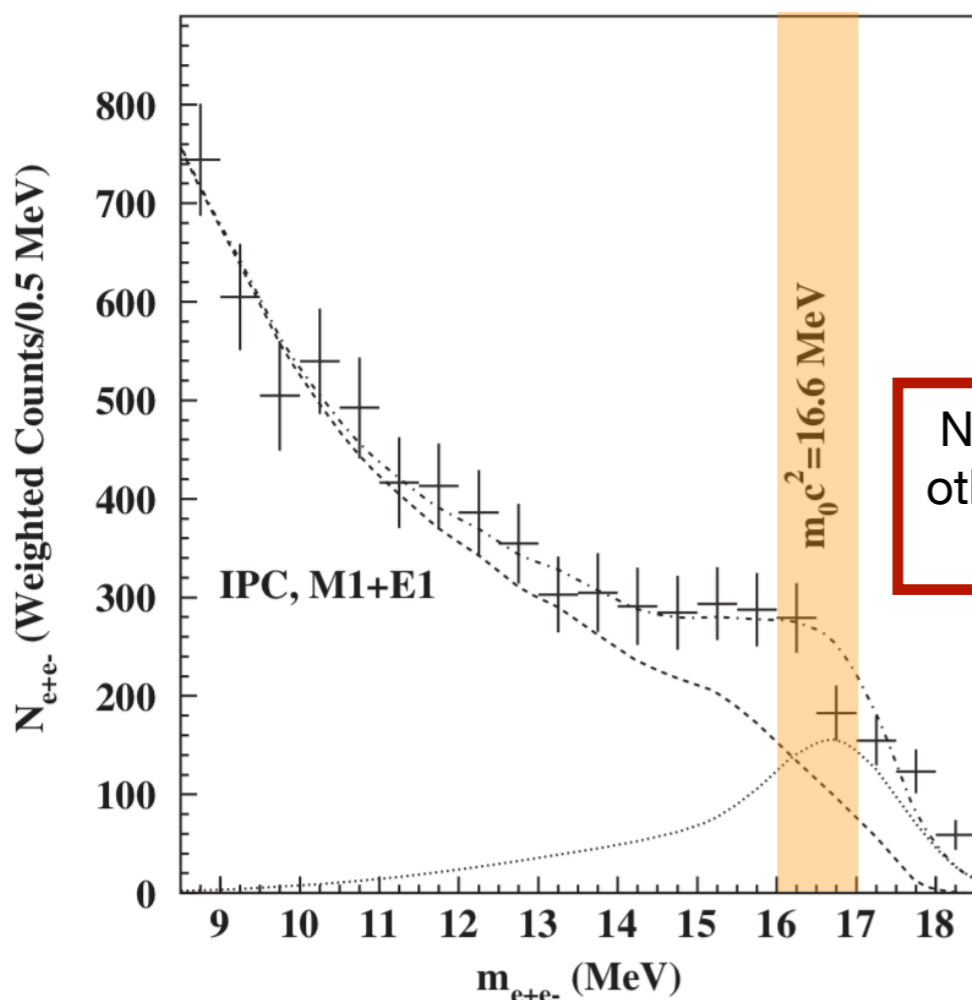


Missing energy

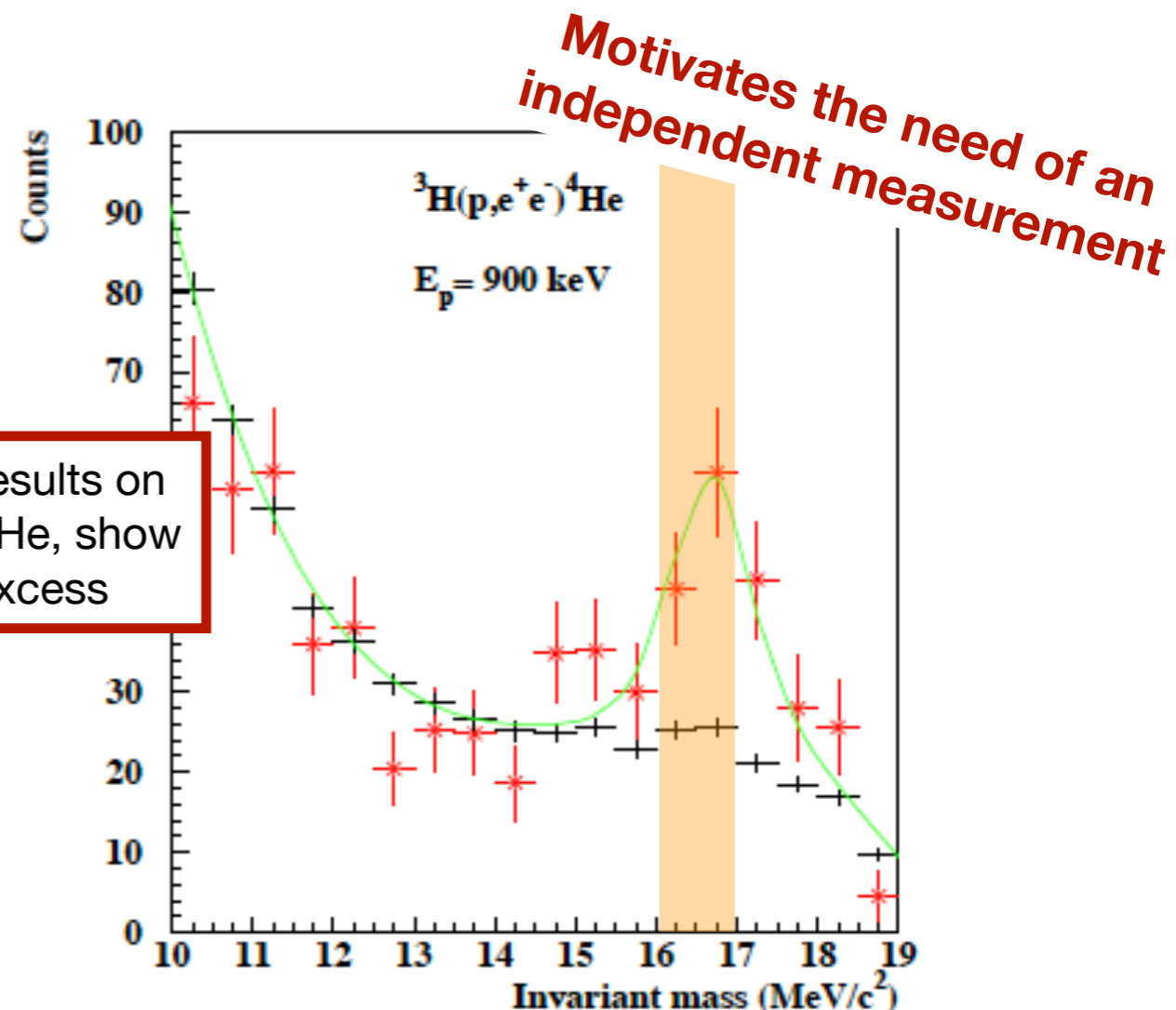
Additional motivation: ^8Be anomaly



- Scalar, pseudo-scalar, vector, axial-vector models could explain the anomaly (large literature)
- NA64 addresses the search for X17 in a model independent way, just assuming its non-zero coupling with electrons.
- Vector model used as benchmark.



A.J. Krasznahorkay et al. Phys. Rev. Lett.116, 042501 (2015)



A. J. Krasznahorkay et. Al Arxiv:1910.10459 (2019)

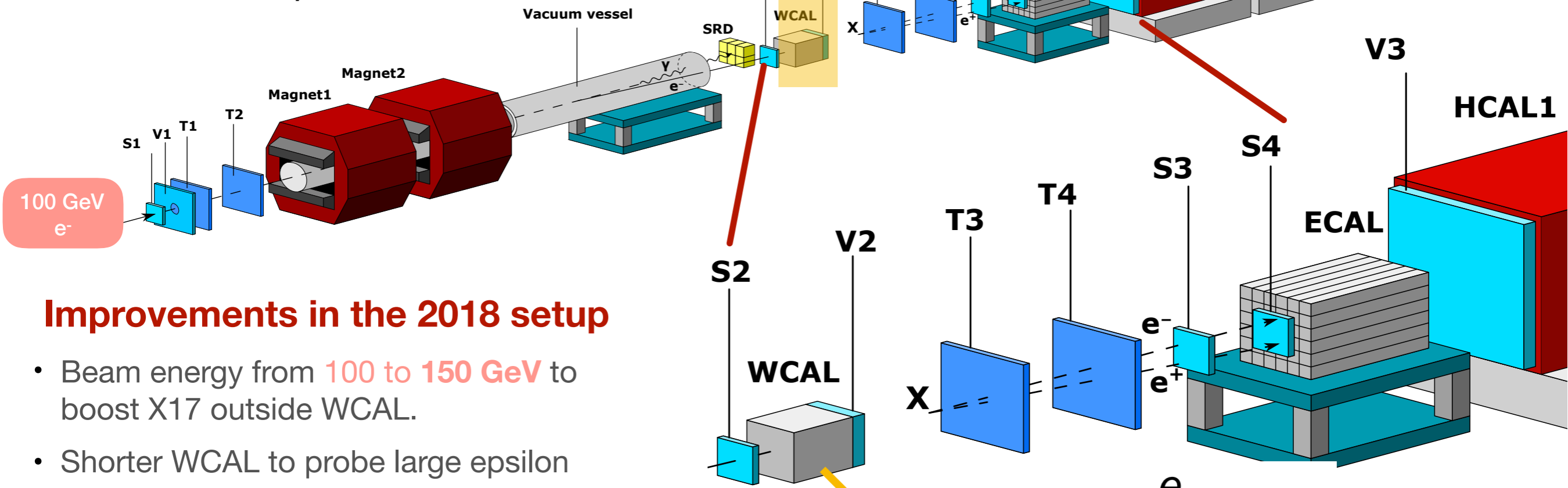
New recent results on other nuclei, ^4He , show a similar excess



First run in 2017

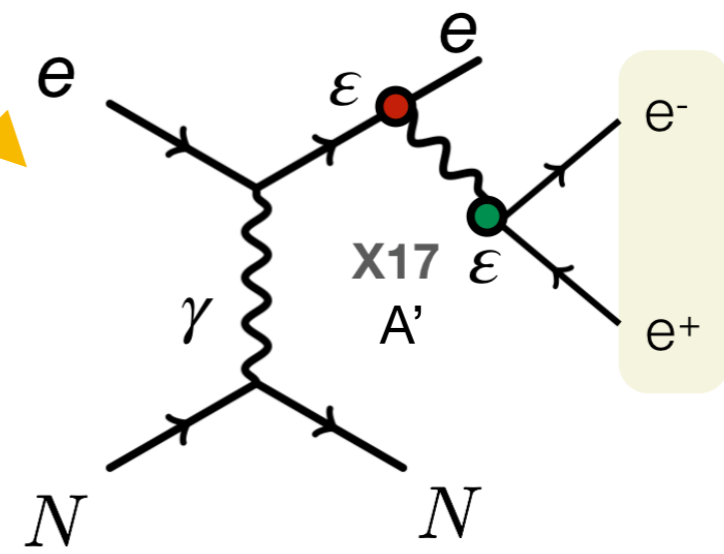
NA64 invisible setup reused as much as possible

Addition of compact Tungsten calorimeter



Improvements in the 2018 setup

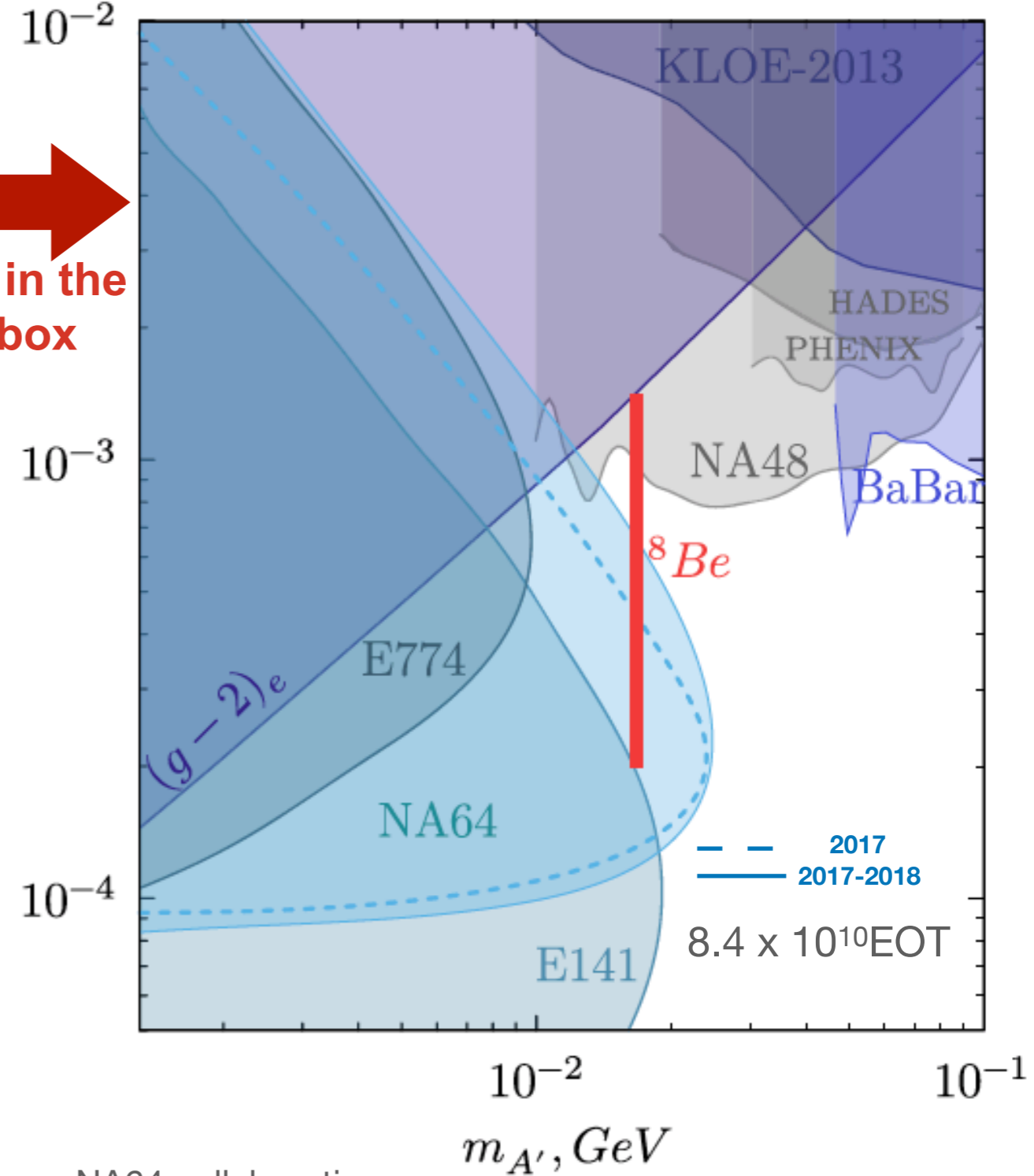
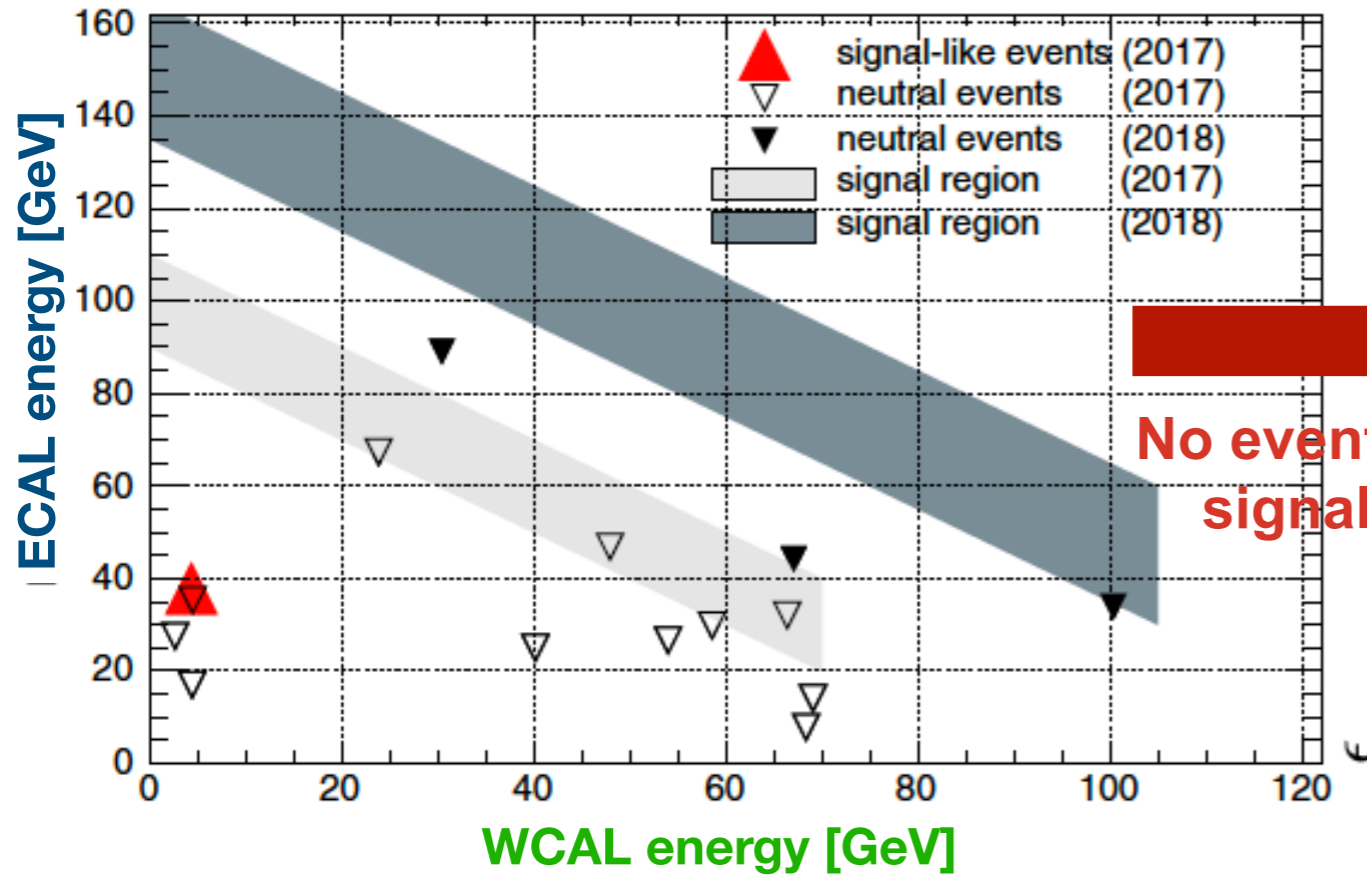
- Beam energy from 100 to 150 GeV to boost X17 outside WCAL.
- Shorter WCAL to probe large epsilon
- Thinner veto (W2) after WCAL to minimise the probability that X17 decays in it.
- For Background suppression: vacuum pipe installed+increased WCAL-ECAL distance
- Additional trackers





Combined results for 2017 and 2018

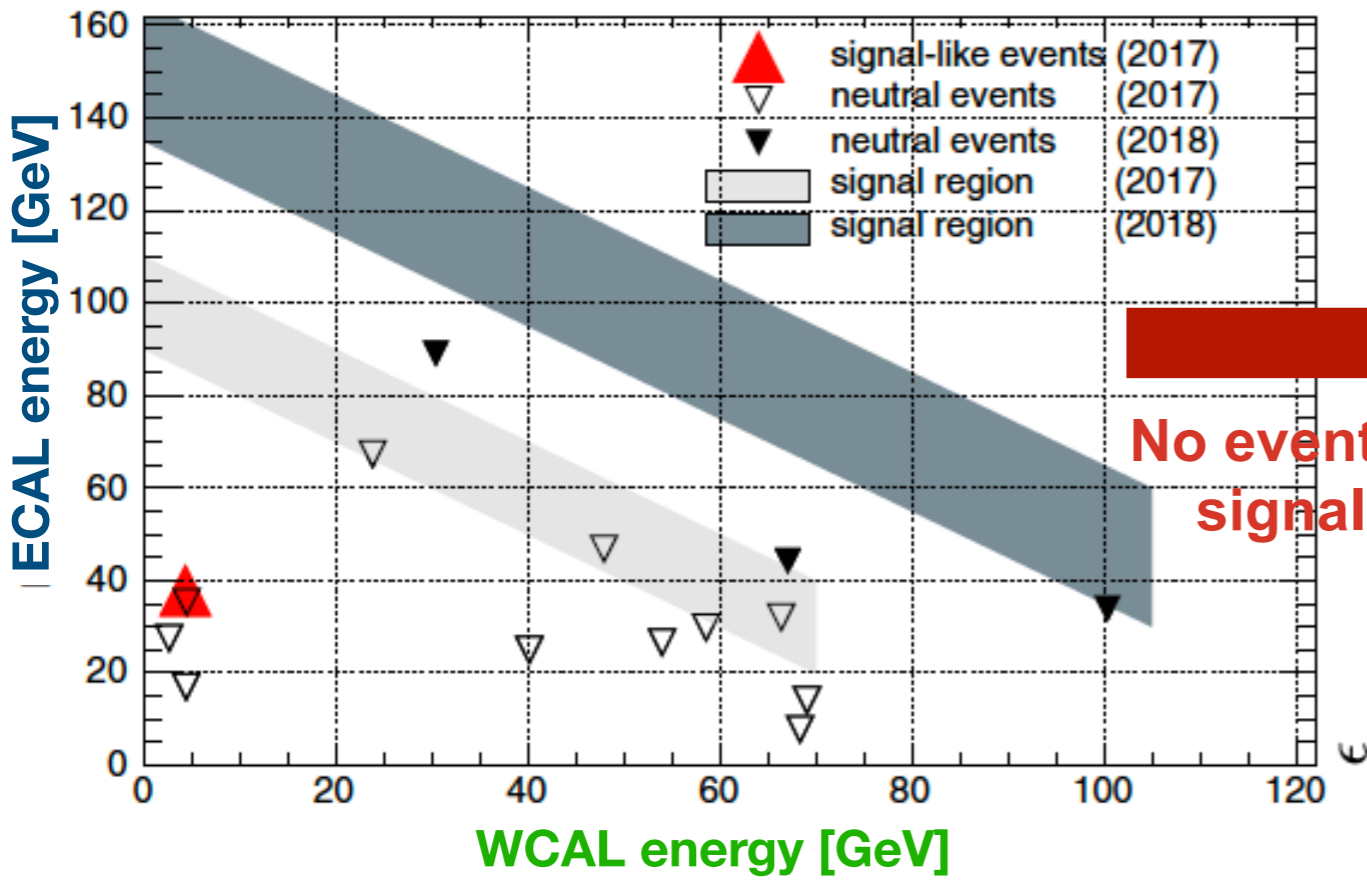
Without the 2018 setup optimisation ϵ would have only increased logarithmically with the number of EOTs.



NA64 collaboration,
 PRL 120, 231802 (2018), PRD 107, 071101 (R) 2020

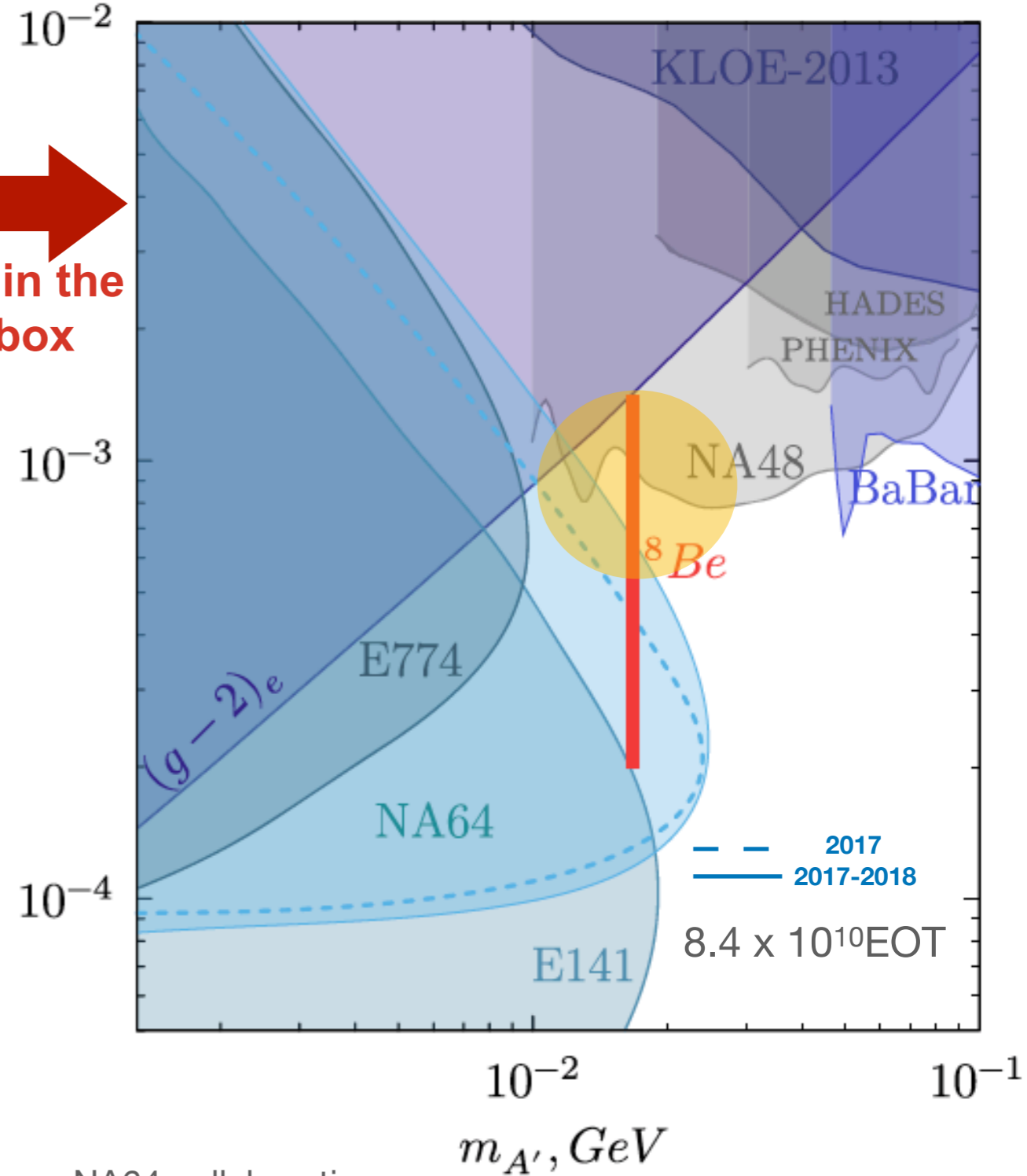


Future prospects for 2022

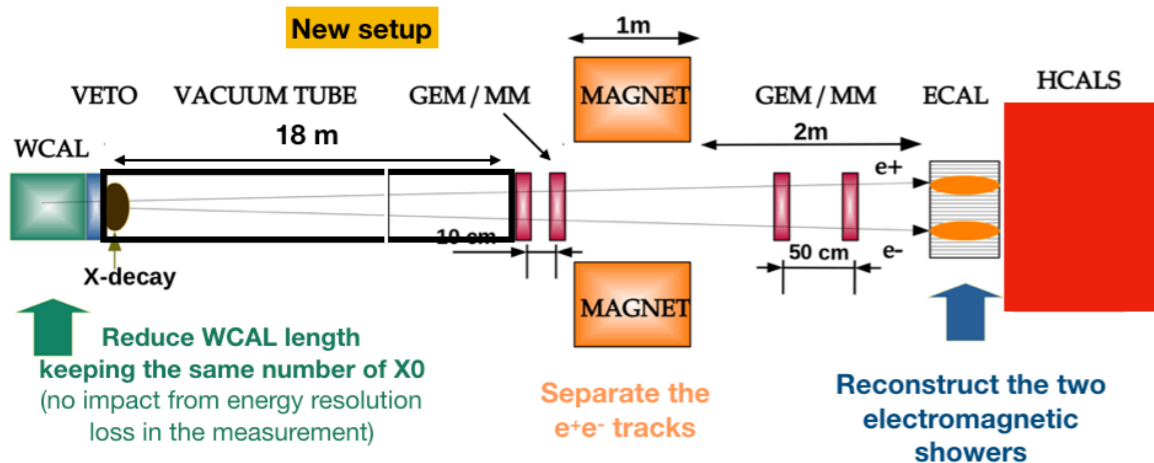


No event in the signal box

Without the 2018 setup optimisation ϵ would have only increased logarithmically with the number of EOTs.

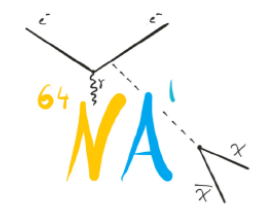


Full parameter space
Invariant mass reconstruction



NA64 collaboration, EPJ C 80 (2020) 12, 1159

NA64 collaboration,
PRL 120, 231802 (2018), PRD 107, 071101 (R) 2020



Summary

Dark sector physics interesting framework to explain dark matter

The European Strategy Group recommended to support searches for dark sector candidates.

NA64 is an ideal experiment for testing benchmark scalar, Majorana and pseudo-Dirac thermal sub-GeV dark matter models

Combined *invisible analysis* data 2016-2018
with 2.84×10^{11} EOT

$A' \rightarrow \chi\bar{\chi}$: Results exceeded sensitivity of previous experiments to thermal sub-GeV dark matter.

Combined *visible* data 2017-2018 with 8.4×10^{10} EOT

$X17 \rightarrow e^+e^-$: Vector coupling with electrons at $\epsilon < 6.8 \times 10^{-4}$ and a mass of 16.7 excluded

Future prospects after LHC long shutdown 2 for 2021-2024

- NA64 setup upgrade to run at high intensity
- Main goal to explore the parameter space of benchmark light dark matter with **> goal 5×10^{12} EOT**
- Substantially increase sensitivity to $A' \rightarrow e^+e^-$ decays and explore remaining parameter space for $X17 \rightarrow e^+e^-$
 - *In case of X17 signal-like events reconstruct the invariant mass with precision at few percent level.*
- **Start searches of dark sectors weakly coupled to muons with NA64 μ :**
 - $(g-2)_\mu$ and $L_\mu-L_\tau Z'$: pilot run in 2021 at M2 beamline at CERN SPS (160 GeV/c muon)
 - Probing light dark matter parameter space for $m_{A'} > 100$ MeV
- Improve sensitivity for ALPs (a) and scalars (s) to probe the challenging region of mass $m_{a,s} \sim m_{\pi^0}$

THANKS!

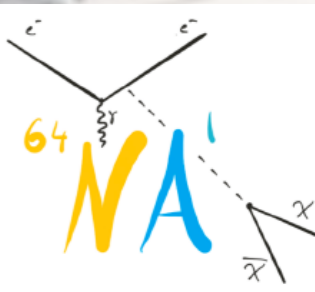
Acknowledgements

NA64 collaboration in particular *P.Crivelli* and *S.Gninenko*

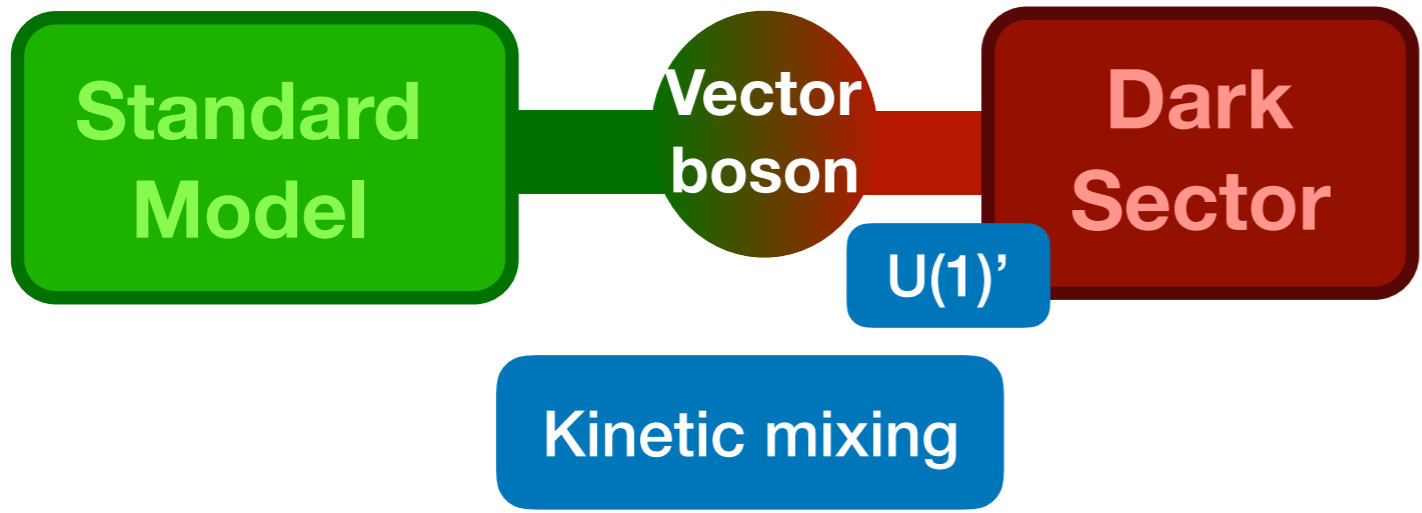
ETH Zürich group in particular *P. Crivelli, B.Bento, E. Depero, H.Sieber*



SNSF Ambizione grant: PZ00P2_186158



NA64 target: *the vector portal*



An interesting framework to explain the origin of dark matter

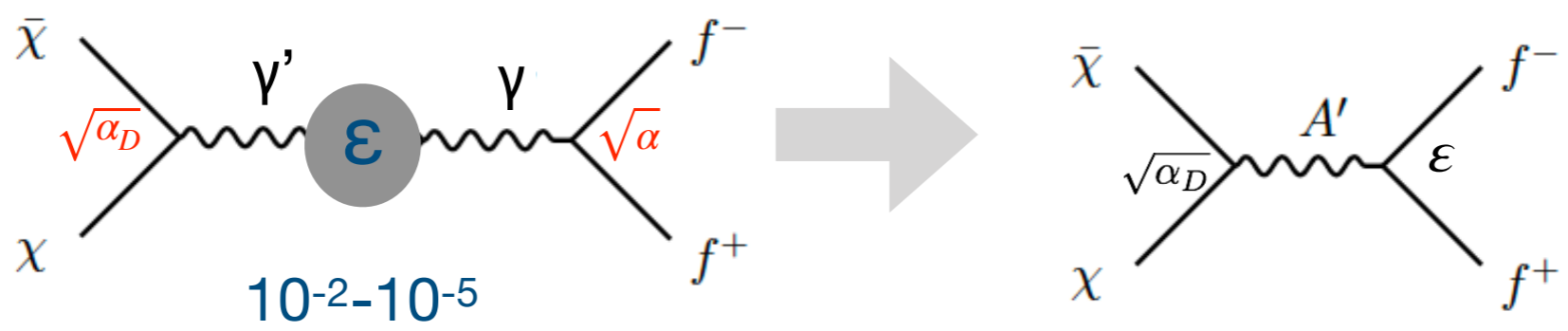
$$L = L_{SM} - \frac{1}{4} F'_{\mu\nu} F'^{\mu\nu} + \frac{m_{A'}^2}{2} A'_\mu A'^\mu + \frac{\epsilon}{2} F'_{\mu\nu} F^{\mu\nu} + i\bar{\chi}\gamma^\mu \partial_\mu \chi - m_\chi \bar{\chi}\chi - e_D \bar{\chi}\gamma^\mu A'_\mu \chi$$

Standard Model Lagrangian

A' massive vector field associated with the $U'(1)$ broken symmetry

Kinetic mixing term with the *standard photon* γ

χ : DM candidate, it can be pseudo-Dirac, scalar or Majorana fermion

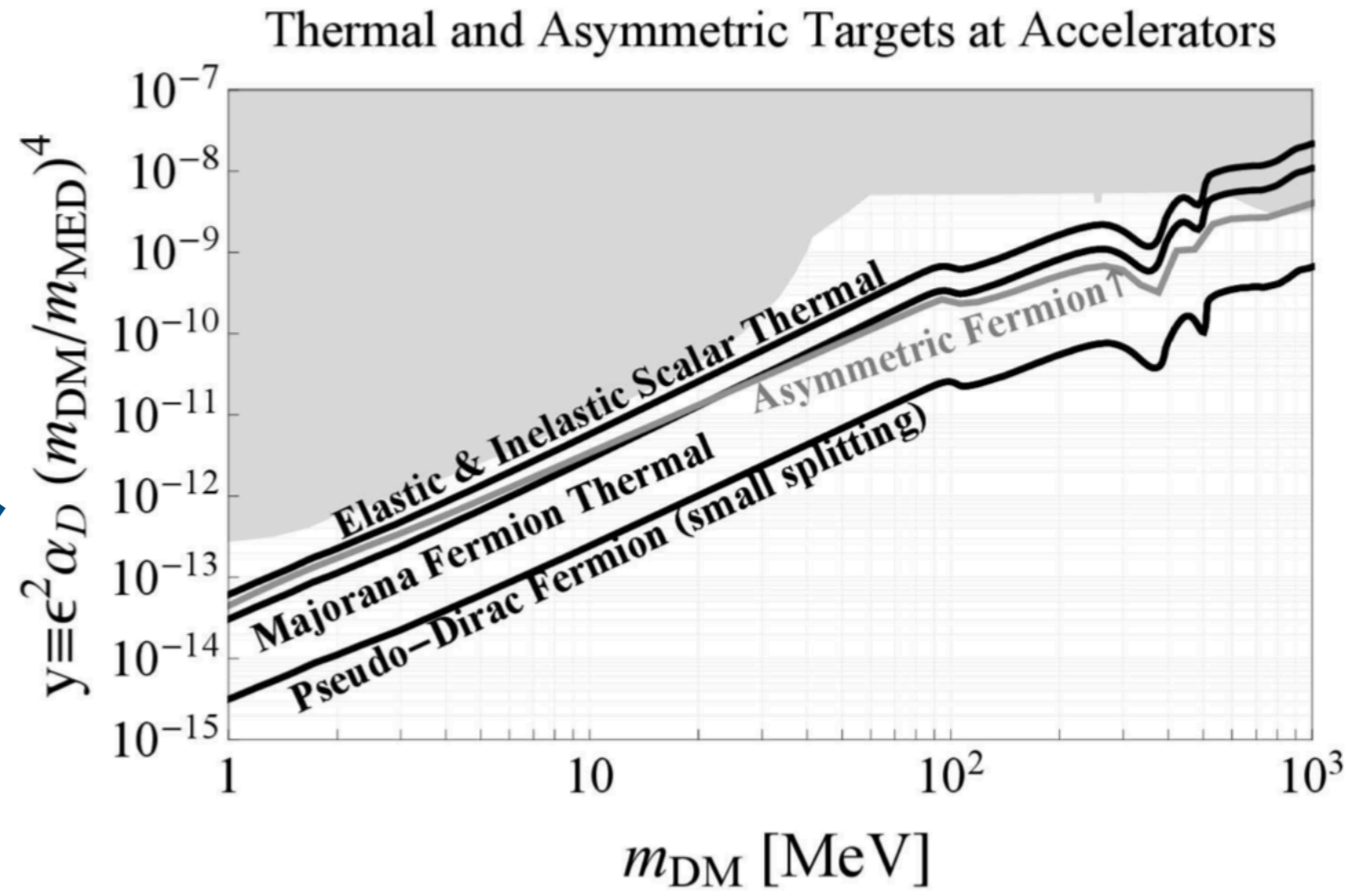
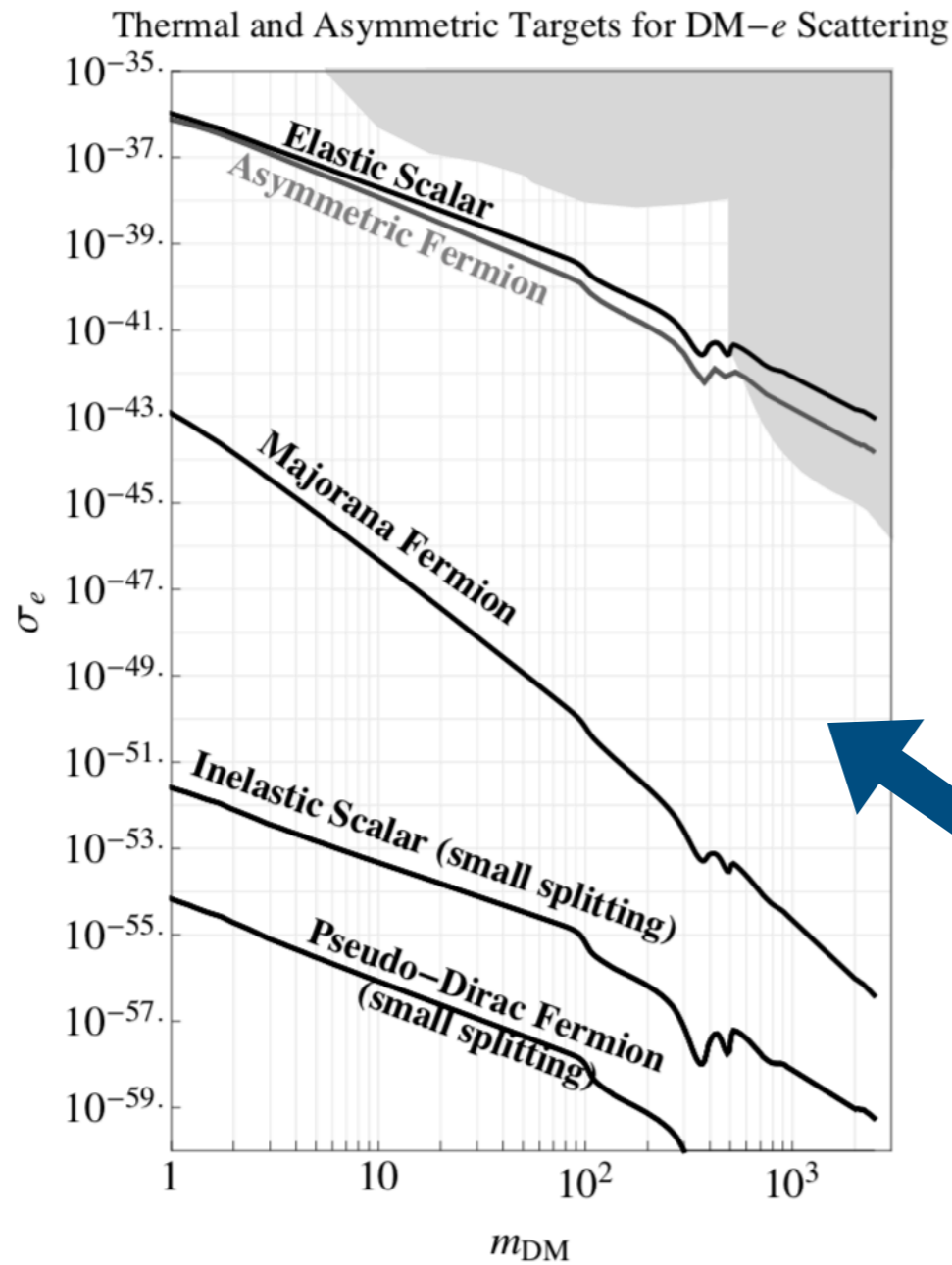


For simplicity during the talk

Parameter space defined by $(m_{A'}, m_\chi, \epsilon, \alpha_D)$

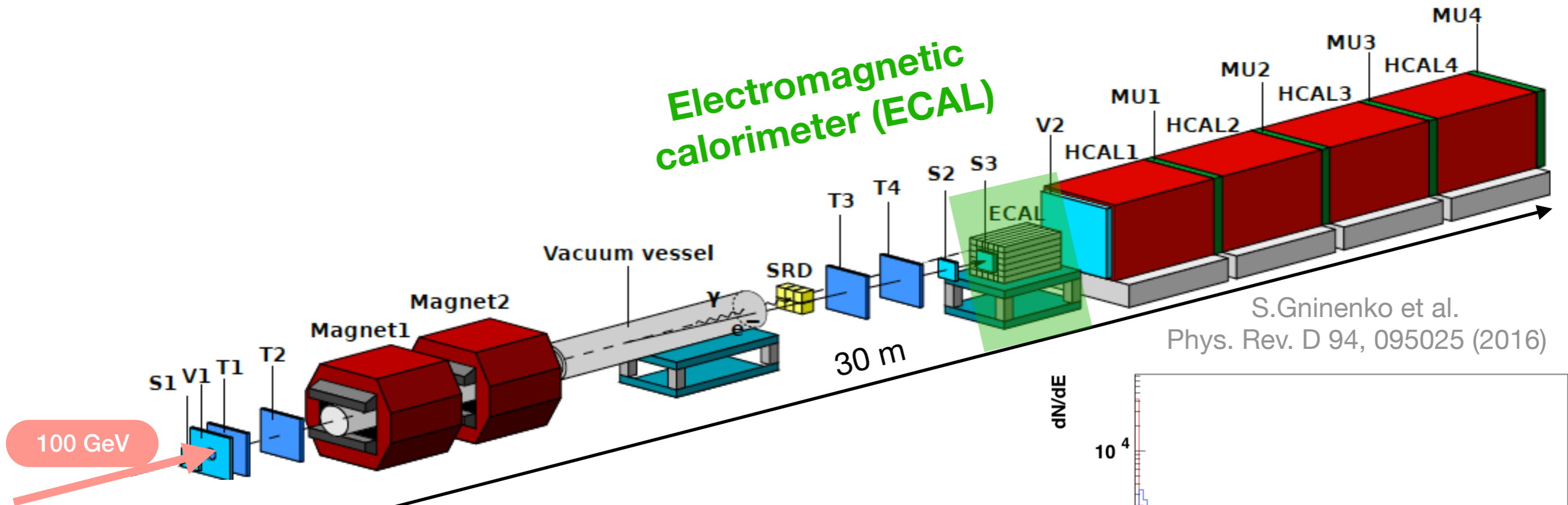
Light thermal dark matter

Equivalent parameter space for e-DM scattering probed by direct-detection experiments



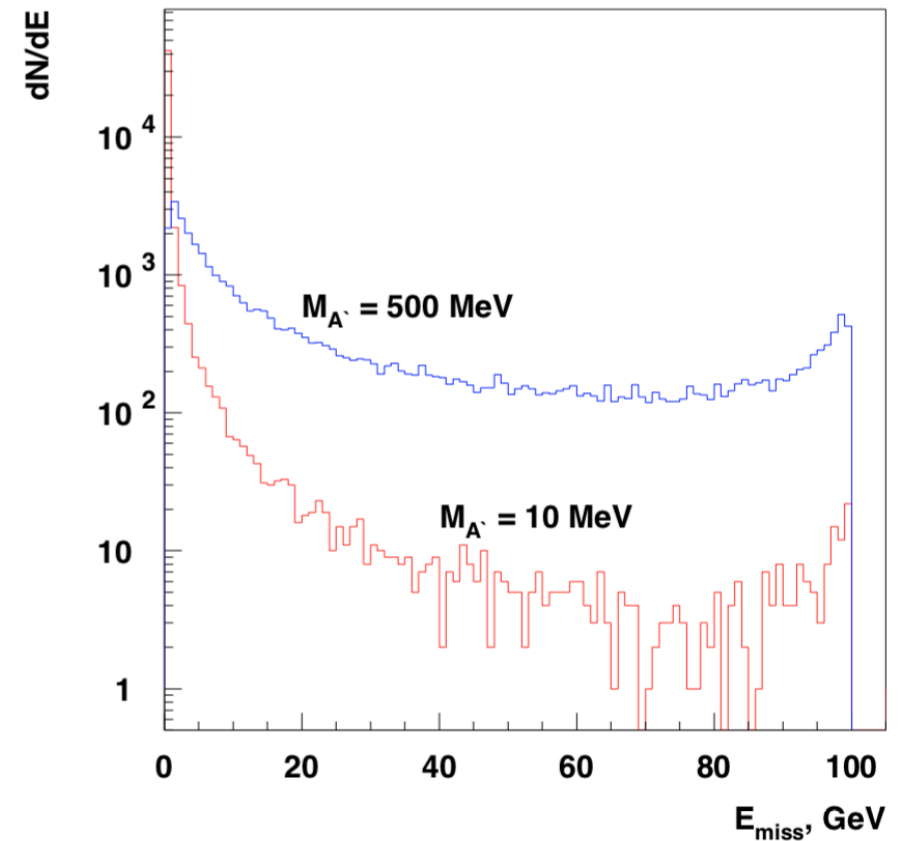
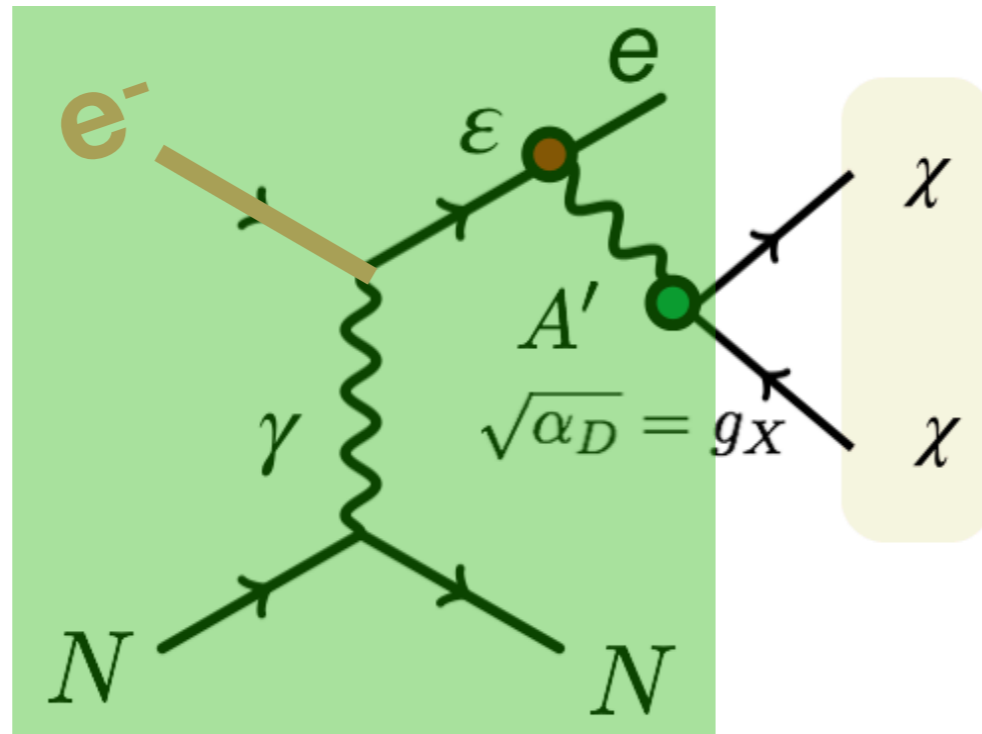


NA64 invisible searches: the setup



S.Gninenko et al.
Phys. Rev. D 94, 095025 (2016)

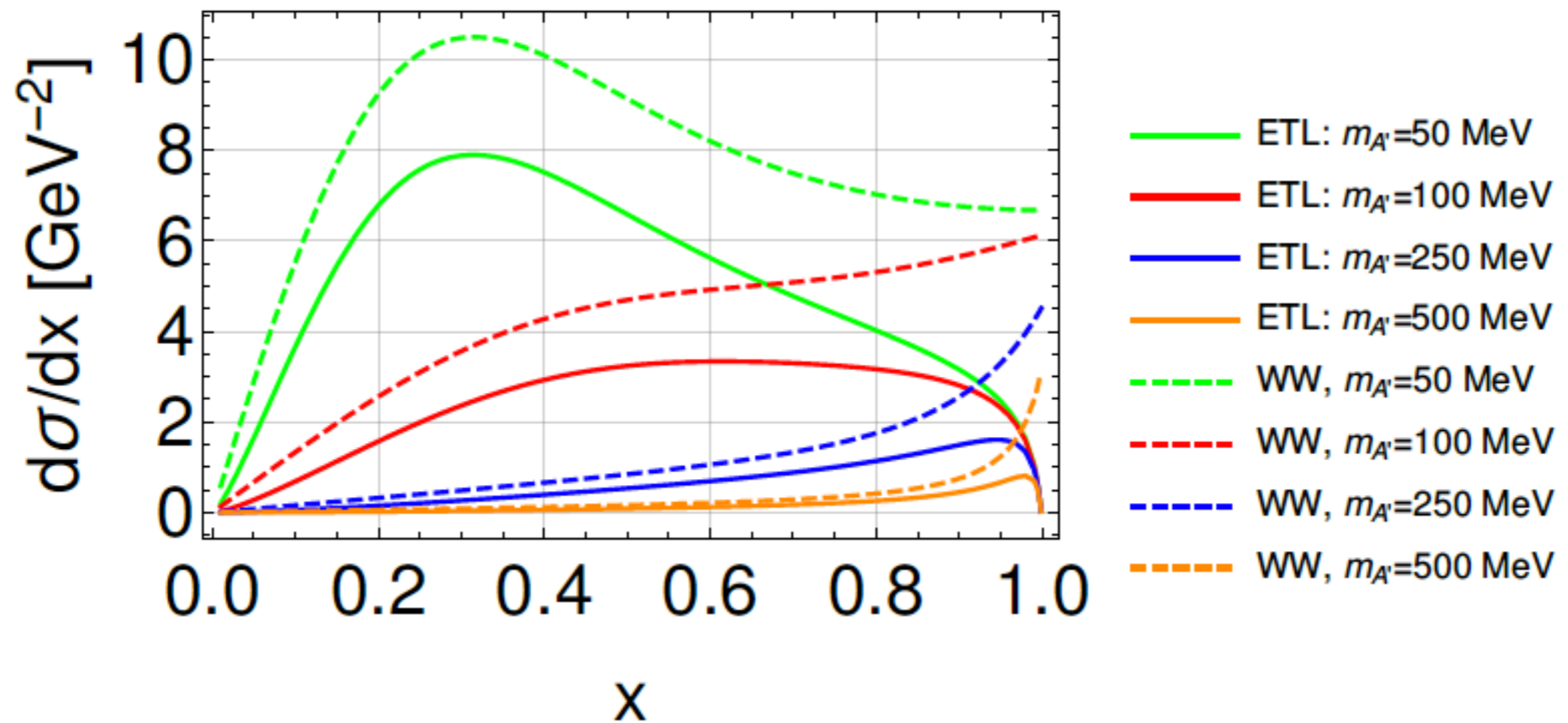
Active target



Bremsstrahlung of A'



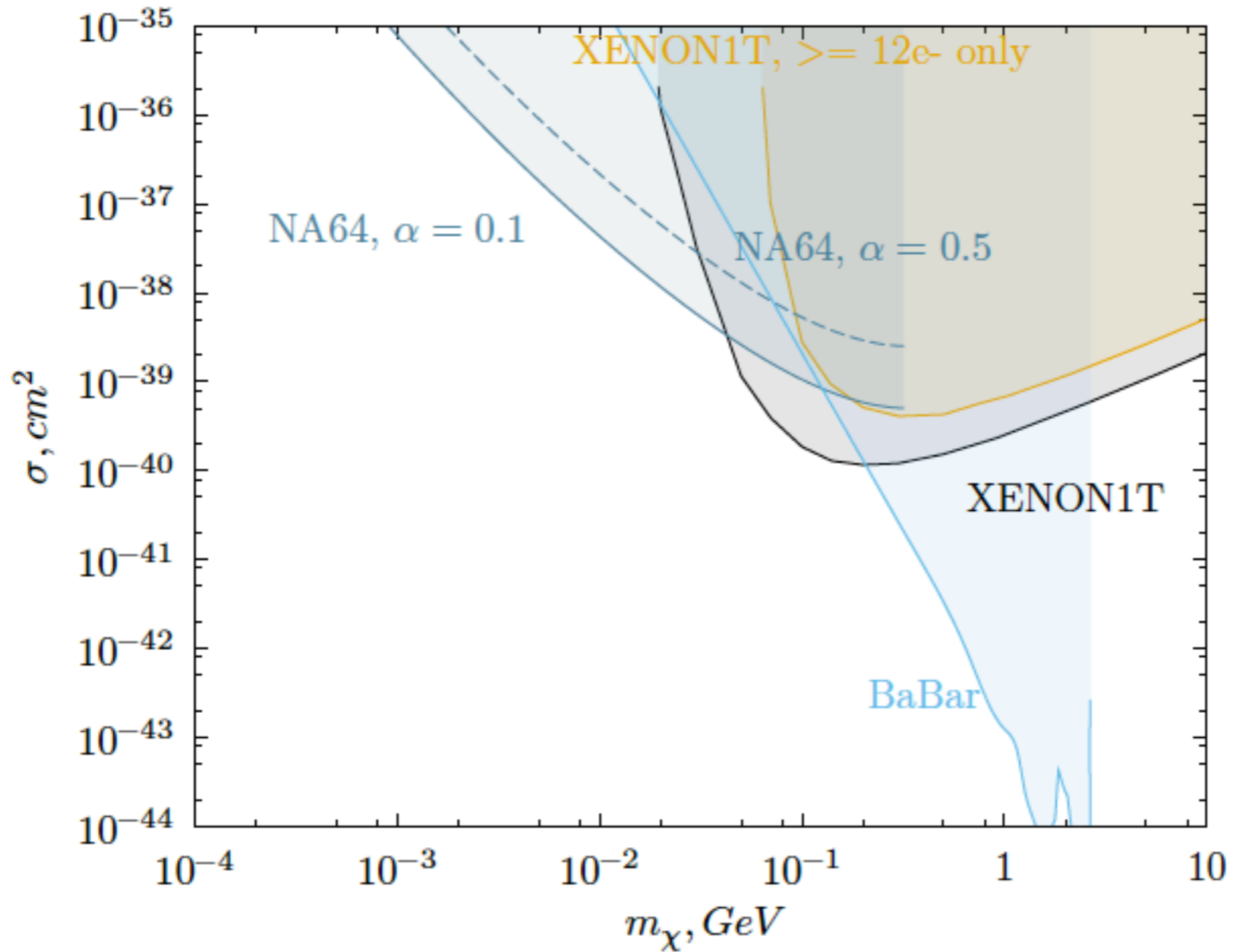
Exact tree level cross-section calculation





NA64 invisible searches: **results**

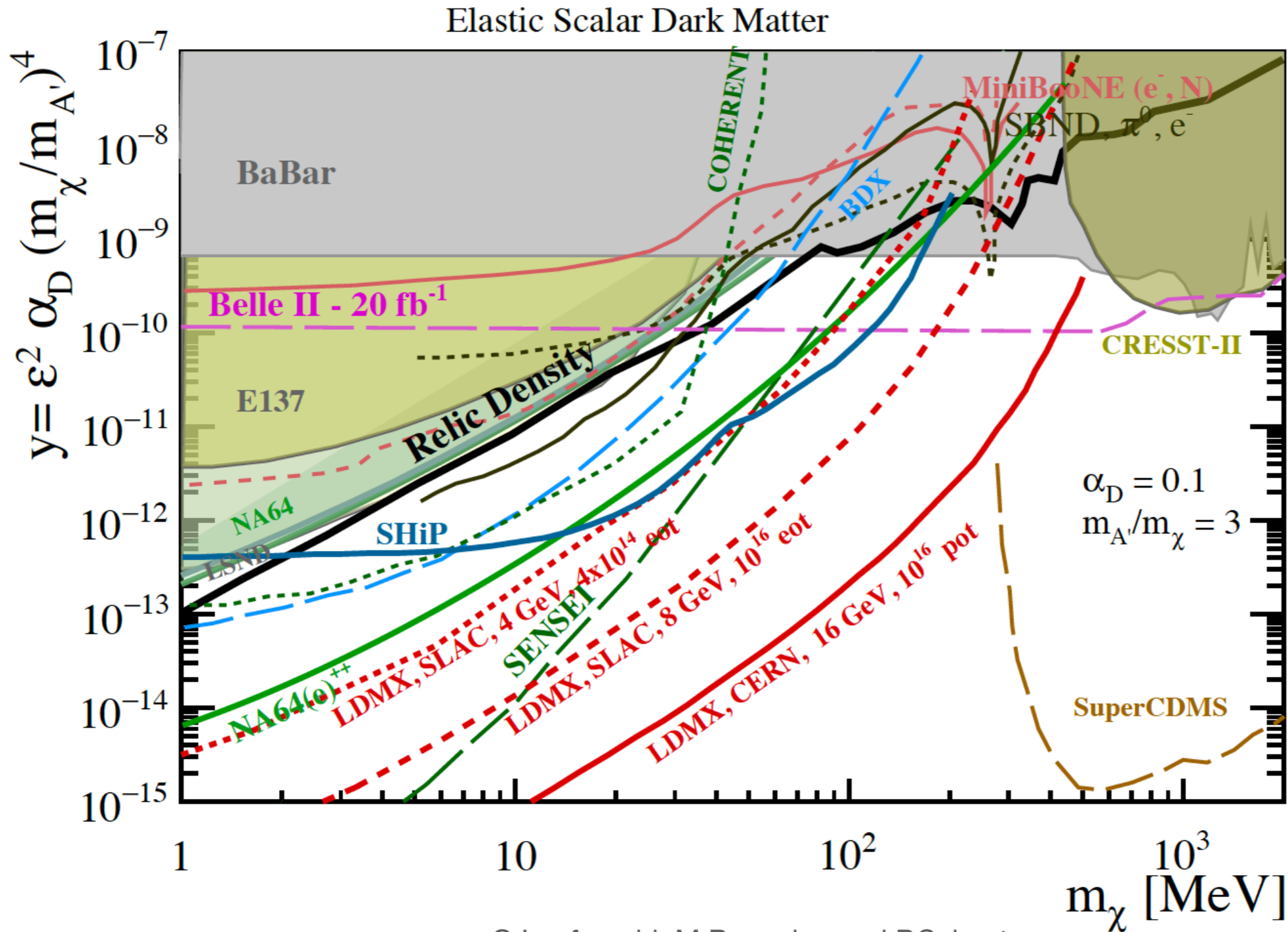
Complementarity with direct DM searches



XENON Collaboration. arXiv:1907.11485
S.N.Gninenko, N.V. Krasnikov and V.A.Matveev, arXiv:2003.07257



Invisible searches



G.Lanfranchi, M.Pospelov and P.Schuster
arxiv:2011.02157



Invisible searches

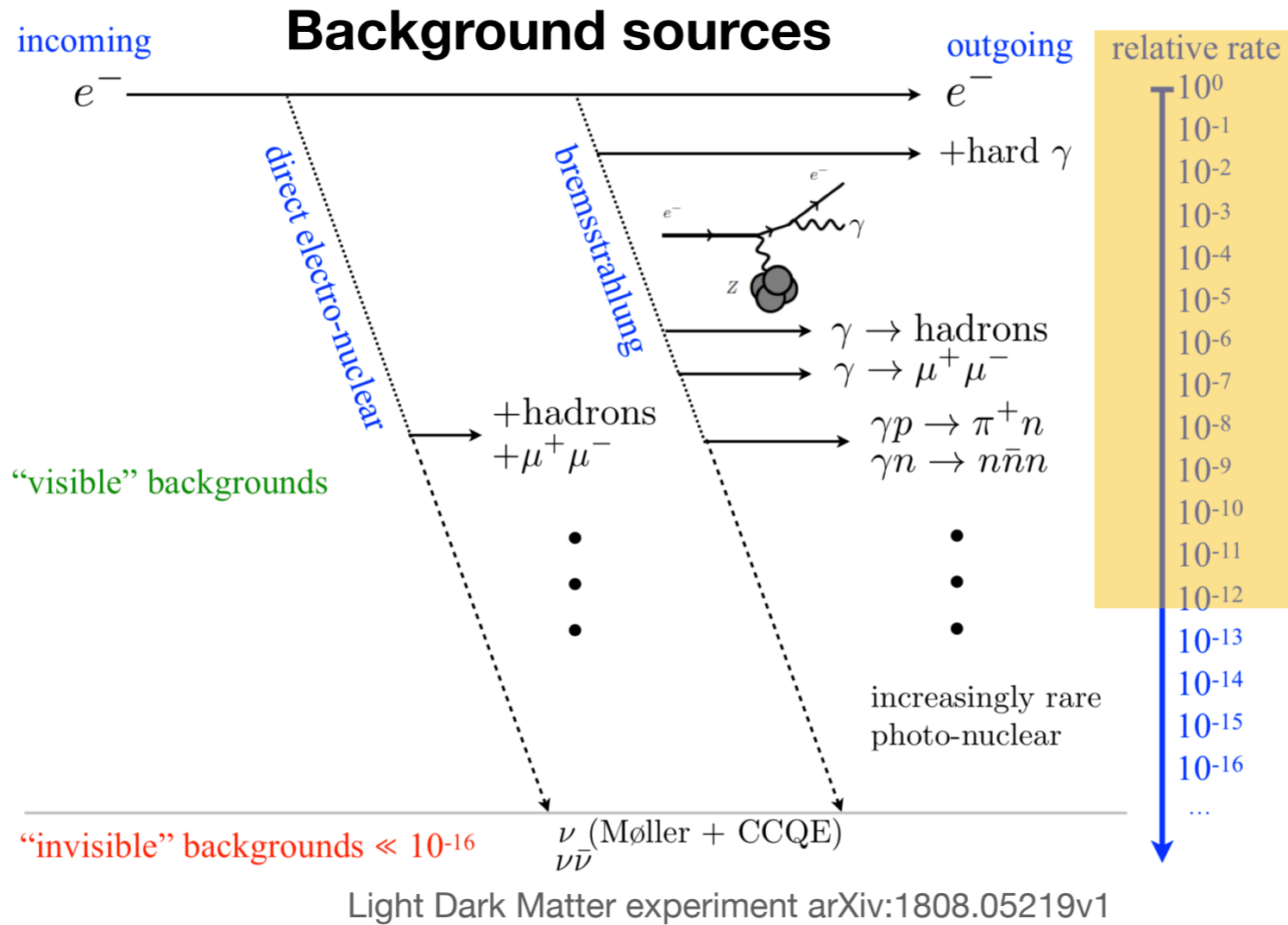
The observed abundance of DM relic density, with the requirement of the thermal freeze-out of DM annihilation into visible matter through $\gamma - A'$ mixing allows one to derive a relation between the parameters:

$$\alpha_D \simeq 0.02 f \left(\frac{10^{-3}}{\epsilon} \right)^2 \left(\frac{m_{A'}}{100 \text{ MeV}} \right)^4 \left(\frac{10 \text{ MeV}}{m_\chi} \right)^2$$

$m_{A'}/m_\chi = 3$ for scalar and $f \leq 10$ for fermion



NA64 invisible searches: results



Main source: electro-nuclear interactions along the beam line

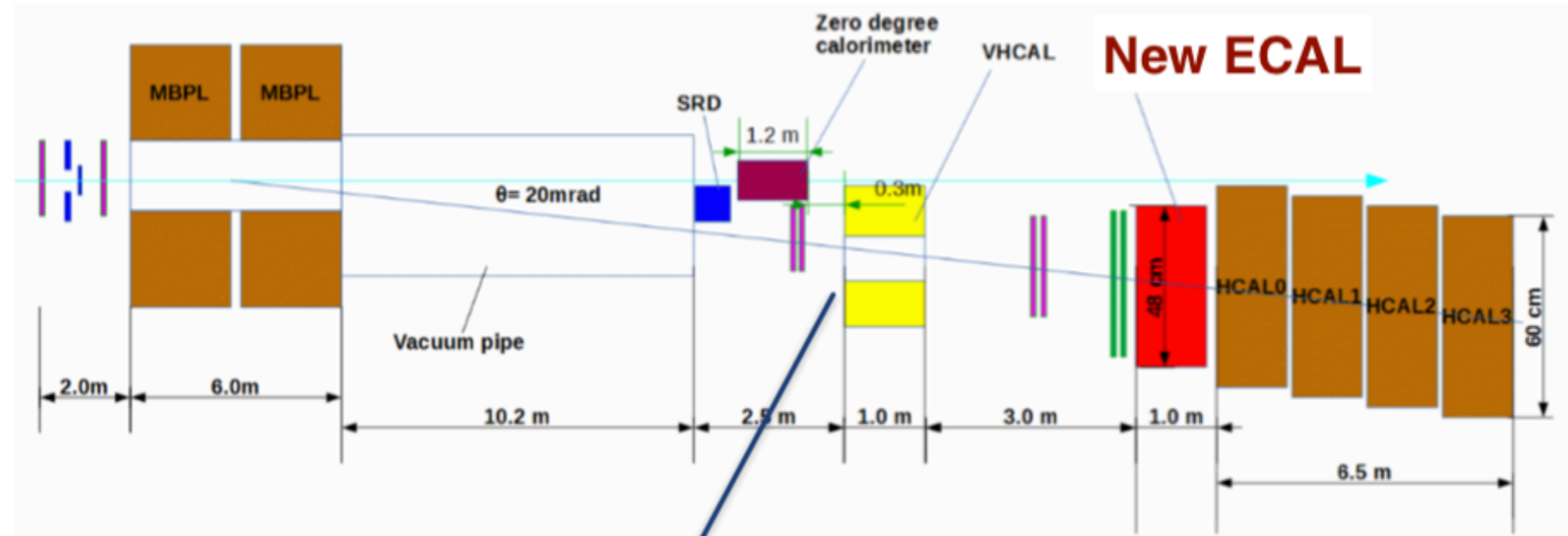
2.84×10^{11} electrons on target (EOT)

Background source	Background number, n_b
punchthrough γ 's, cracks, holes	< 0.01
loss of dimuons	0.024 ± 0.007
$\mu \rightarrow e\nu\nu$, π , $K \rightarrow e\nu$, K_{e3} decays	0.02 ± 0.01
e^- interactions in the beam line	0.43 ± 0.16
μ, π, K interactions in the target	0.044 ± 0.014
accidental SR tag and μ, π, K decays	< 0.01
Total n_b	0.53 ± 0.17

Extrapolated from data

Setup upgrade

GOAL: increase the overall performance and improve background rejection



New VHCAL: to improve detector hermiticity and reject high- p_t hadronic secondaries from beam interactions stream the ECAL dump. Search expected to be BKG free up to $\sim 10^{13}$ EOT

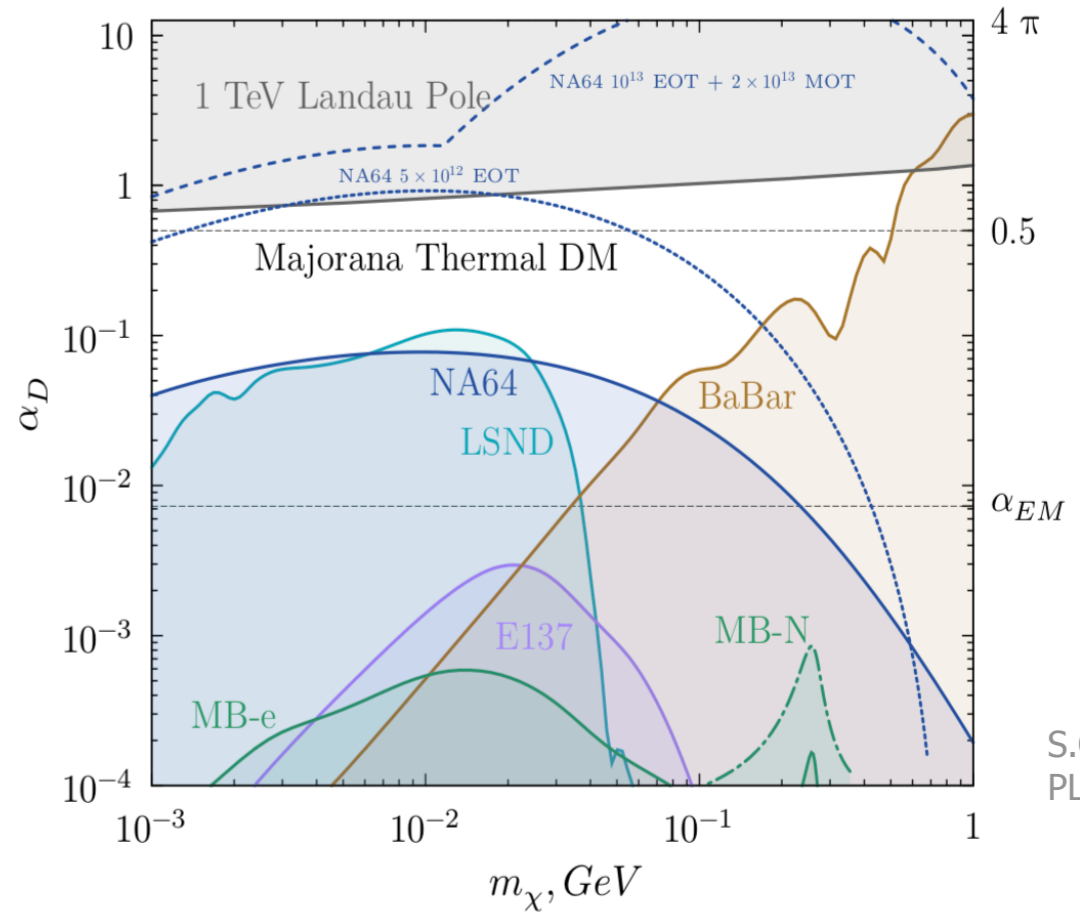
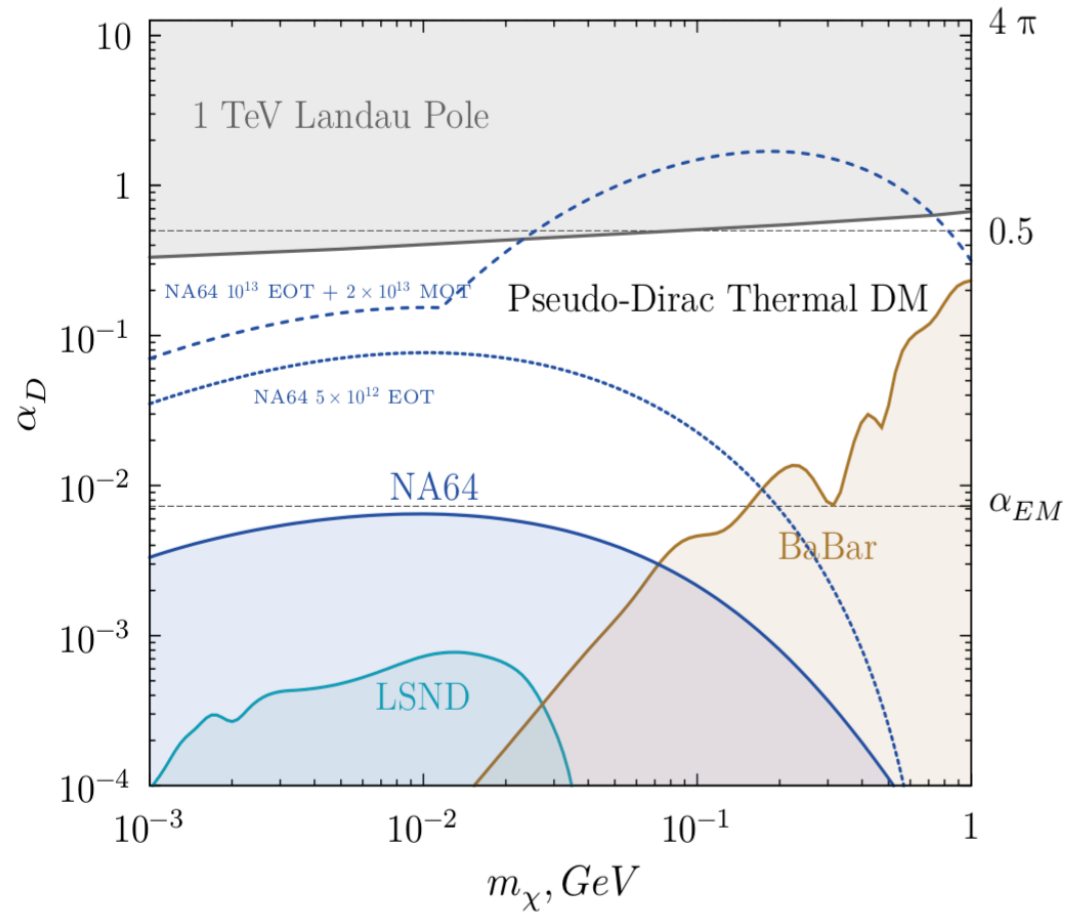
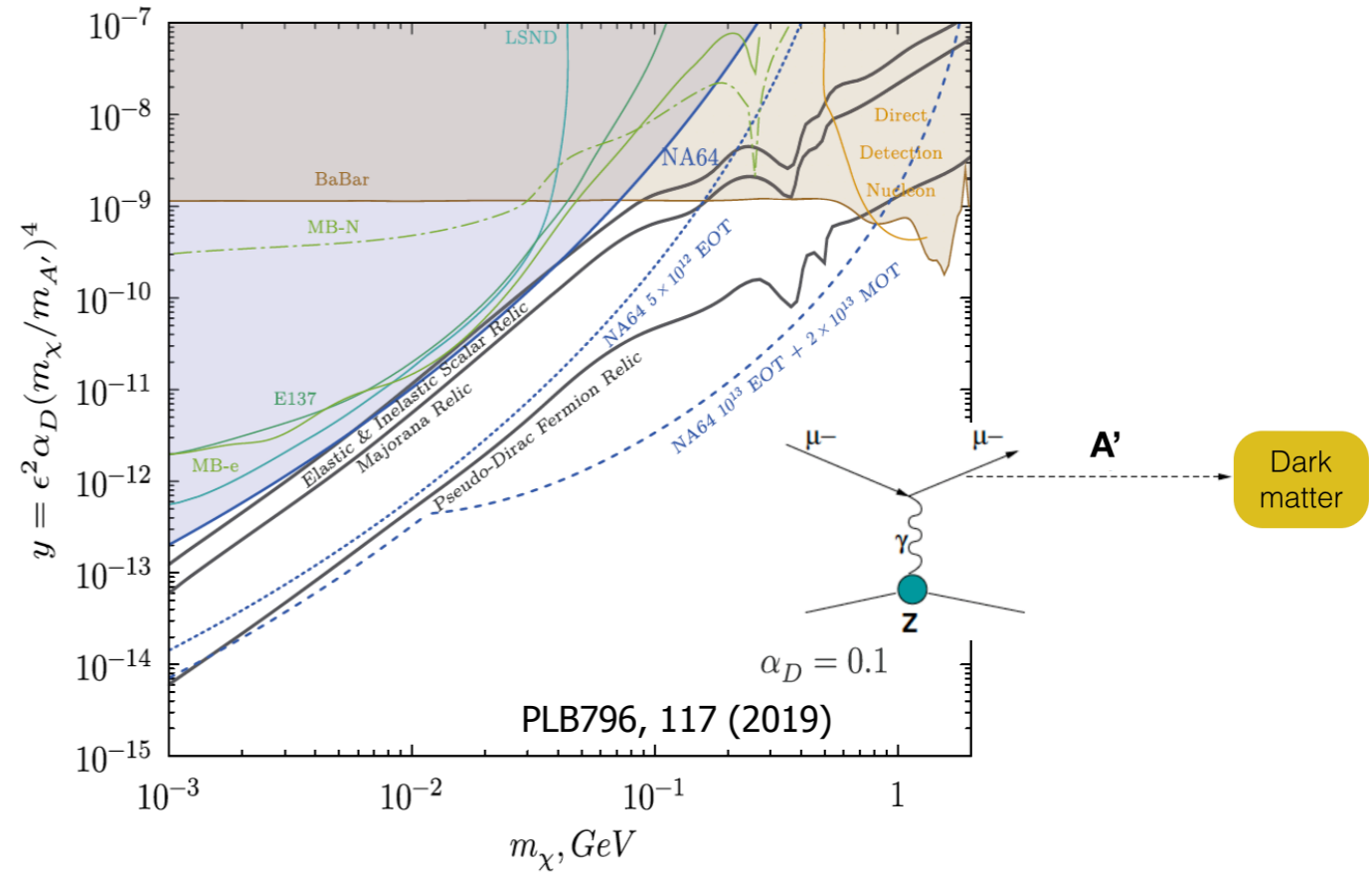
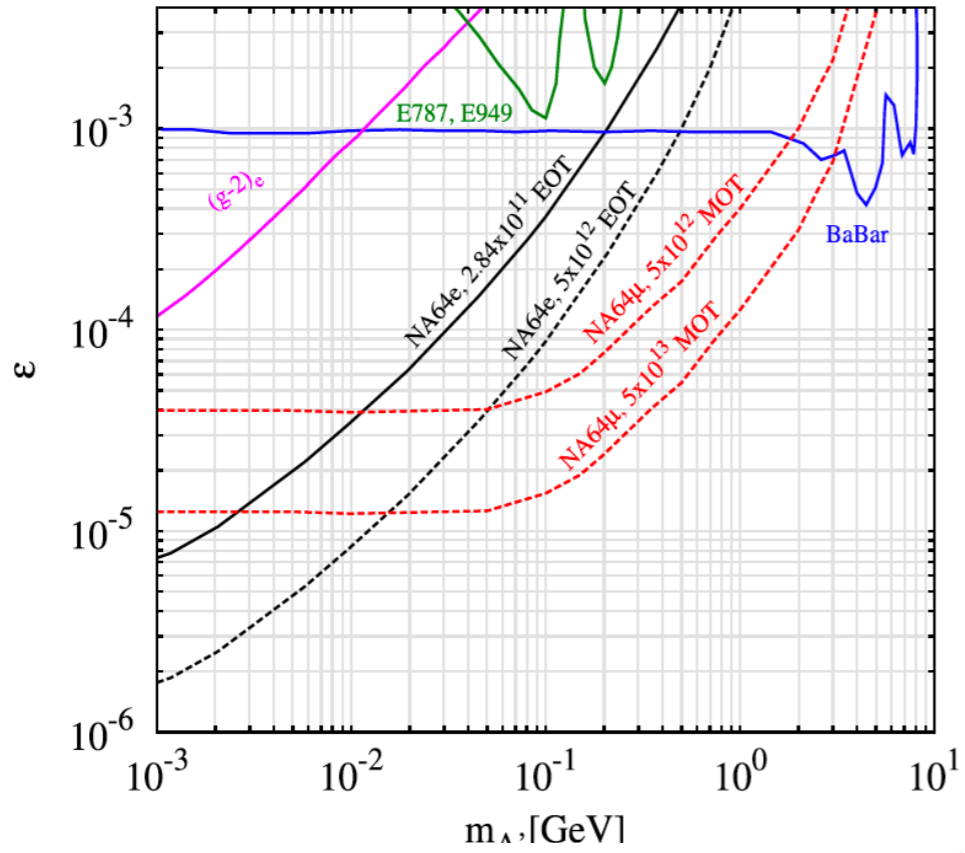


- Dimensions $\sim 50 \times 50 \text{ cm}^2$, 16 cells, matrix 4x4 cells
- Central hole size $12 \times 6 \text{ cm}^2$
- Cell size $12 \times 12 \text{ cm}^2$
- Length $\sim 100 \text{ cm}$, 5λ
- 30 layers, 25 mm copper + 2 mm scintillator
- Read out WLS fiber, 12 fibers per scintillator
- Light yield ~ 15 photoelectrons per MIP

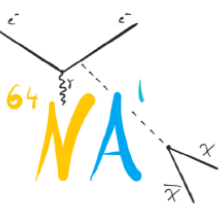
From P.Crivelli, 138th SPSC meeting June 2020



NA64e + NA64μ



S.Gninenko et al.
PLB796, 117 (2019)

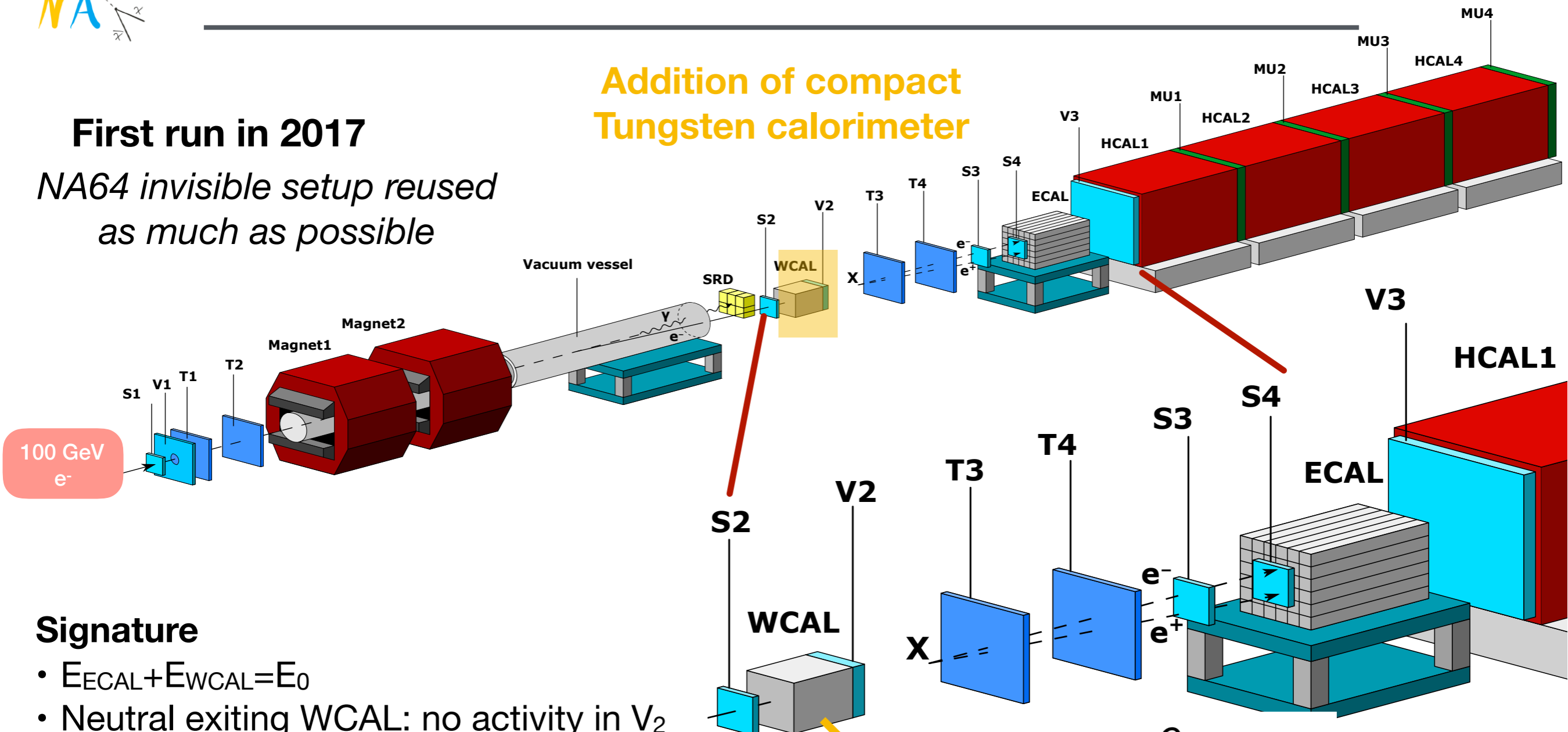


NA64 *visible* searches

Addition of compact Tungsten calorimeter

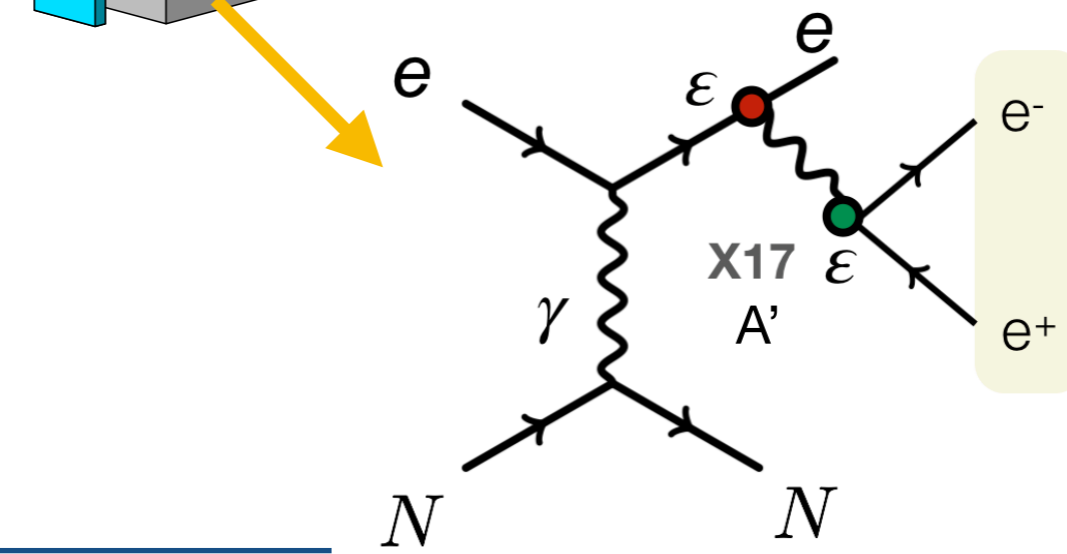
First run in 2017

NA64 invisible setup reused as much as possible



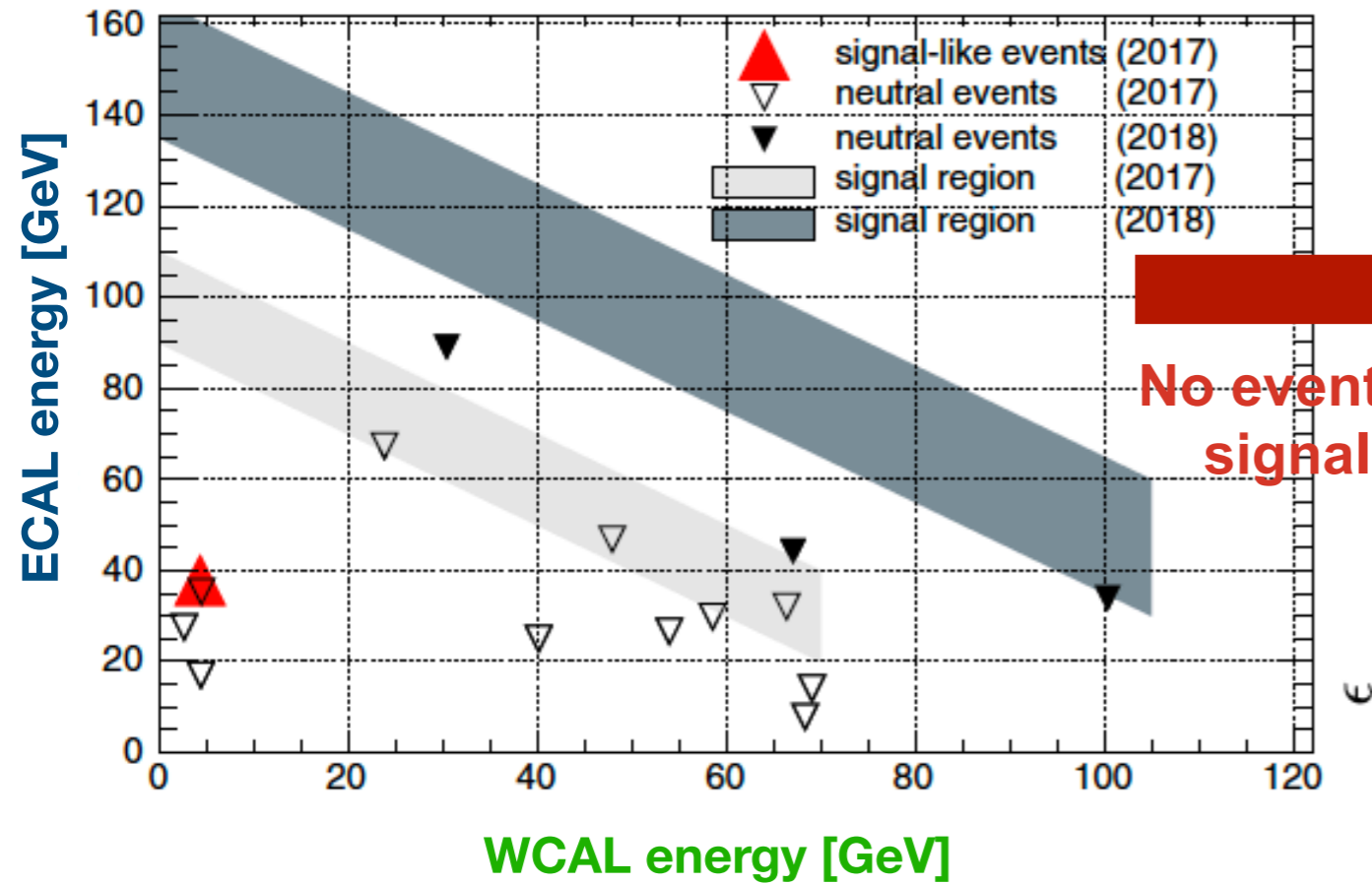
Signature

- $E_{\text{ECAL}} + E_{\text{WCAL}} = E_0$
- Neutral exiting WCAL: no activity in V₂
- Leaking in decay volume: single electromagnetic shower
- Charged particle in decay volume: signal on S_{3,4}
- No hadron large scattering: no activity in HCAL/V₃

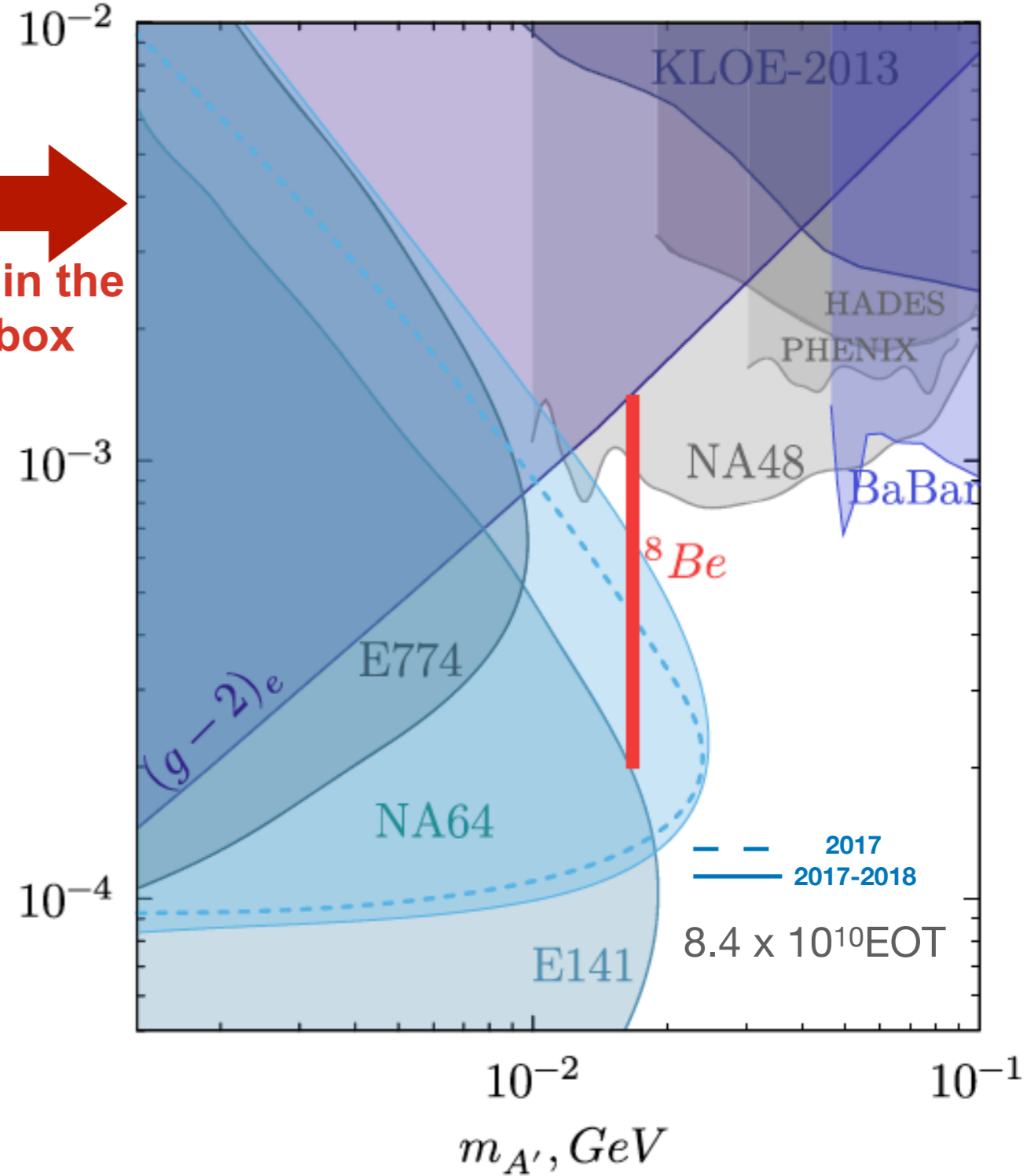




Combined results for 2017 and 2018



Without the 2018 setup optimisation ϵ would have only increased logarithmically with the number of EOTs.



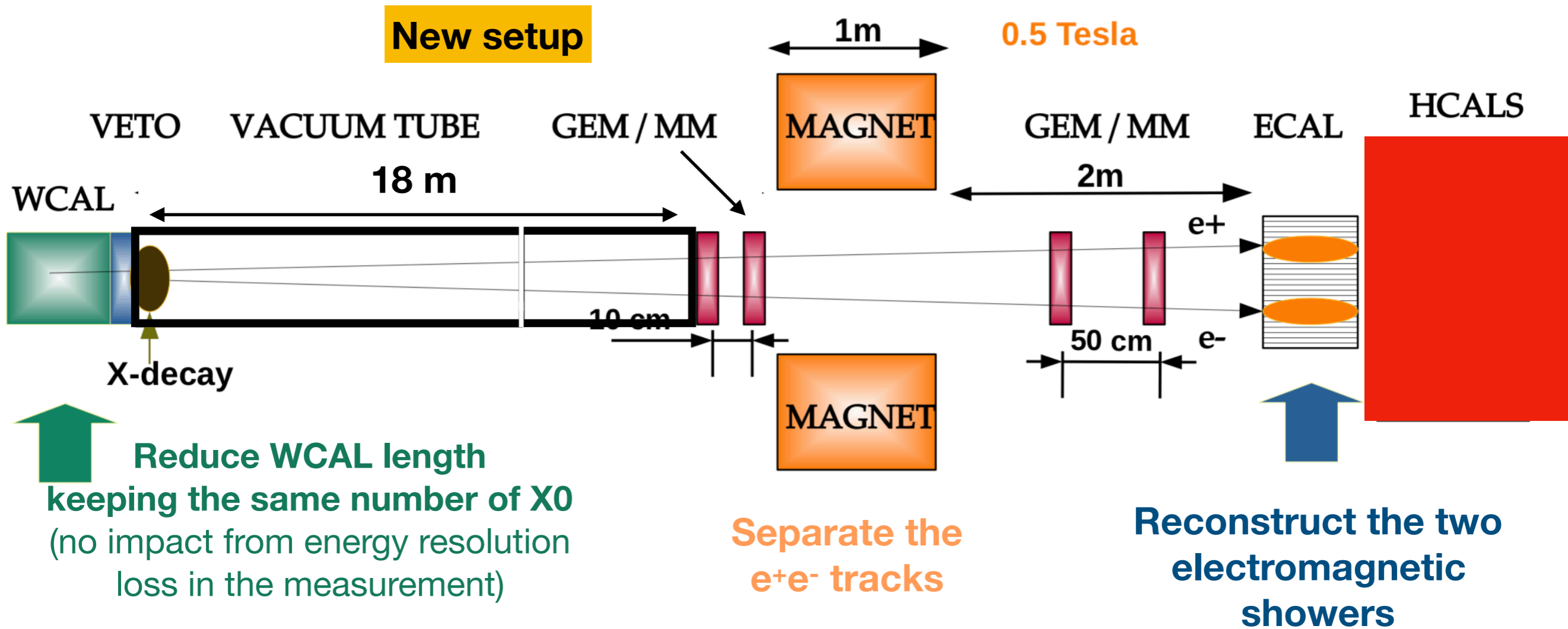
Source of background	2017 data	2018 data
$K_S^0 \rightarrow 2\pi^0$	0.06 ± 0.034	0.005 ± 0.003
$\pi N \rightarrow (\geq 1)\pi^0 + n + \dots$	0.01 ± 0.004	0.001 ± 0.0004
Punchthrough π^-	0.0015 ± 0.0008	0.0007 ± 0.0004
Punchthrough γ	<0.001	<0.0005
$\pi, K \rightarrow e\nu, K_{e4}$ decays	<0.001	<0.001
$eZ \rightarrow eZ\mu^+\mu^-; \mu^\pm \rightarrow e^\pm\nu\bar{\nu}$	<0.001	<0.001
Total	0.07 ± 0.035	0.006 ± 0.003

NA64 collaboration, PRL 120, 231802 (2018), PRD 107, 071101 (R) 2020



Future prospects for 2022 run

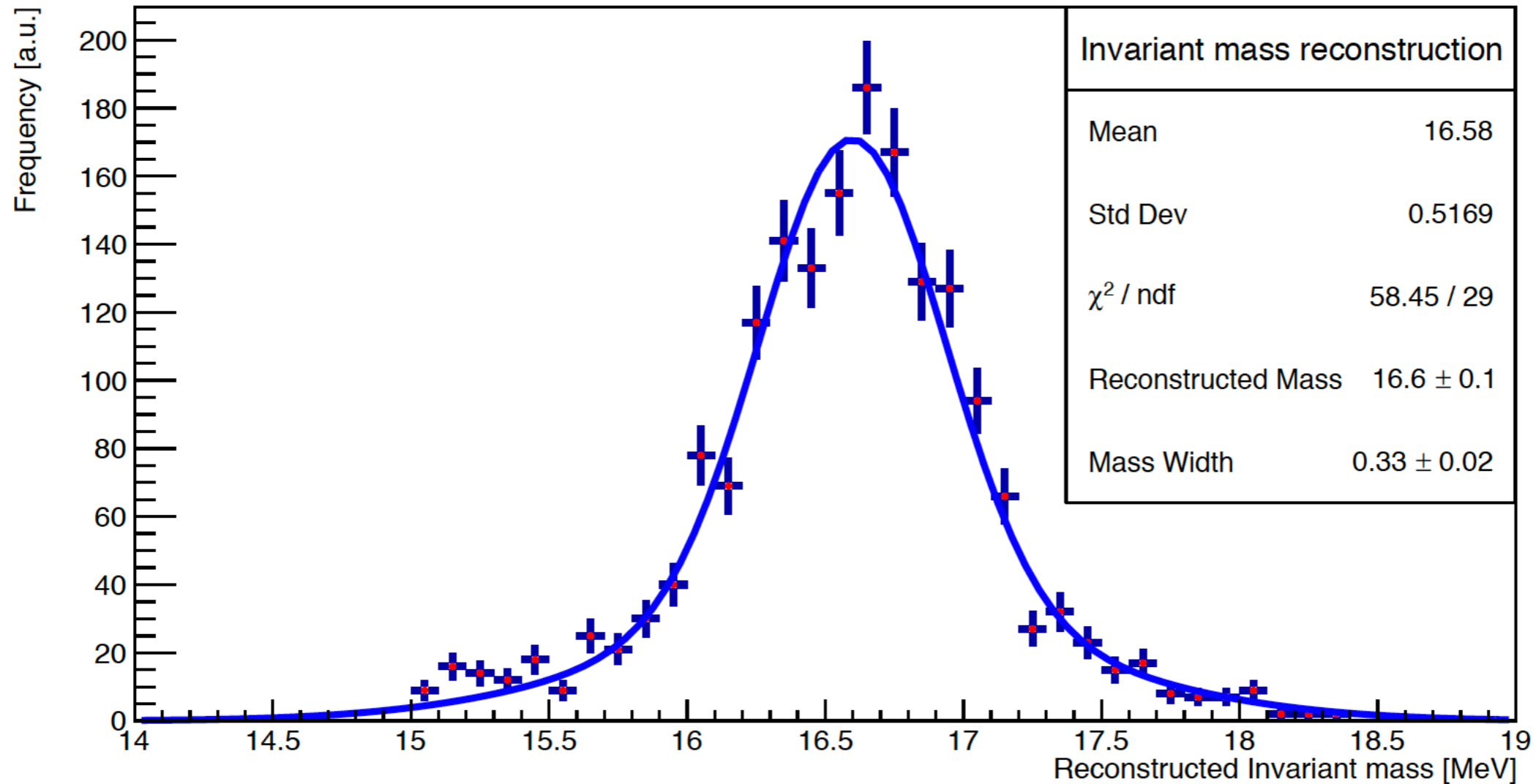
Explore the **full X17 allowed parameter space** to explain the ^8Be anomaly
Invariant mass reconstruction in case signal-like events are founded





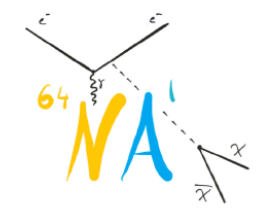
Future prospects for 2022 run

Explore the **full X17 allowed parameter space** to explain the ^8Be anomaly
Invariant mass reconstruction in case signal-like events are founded

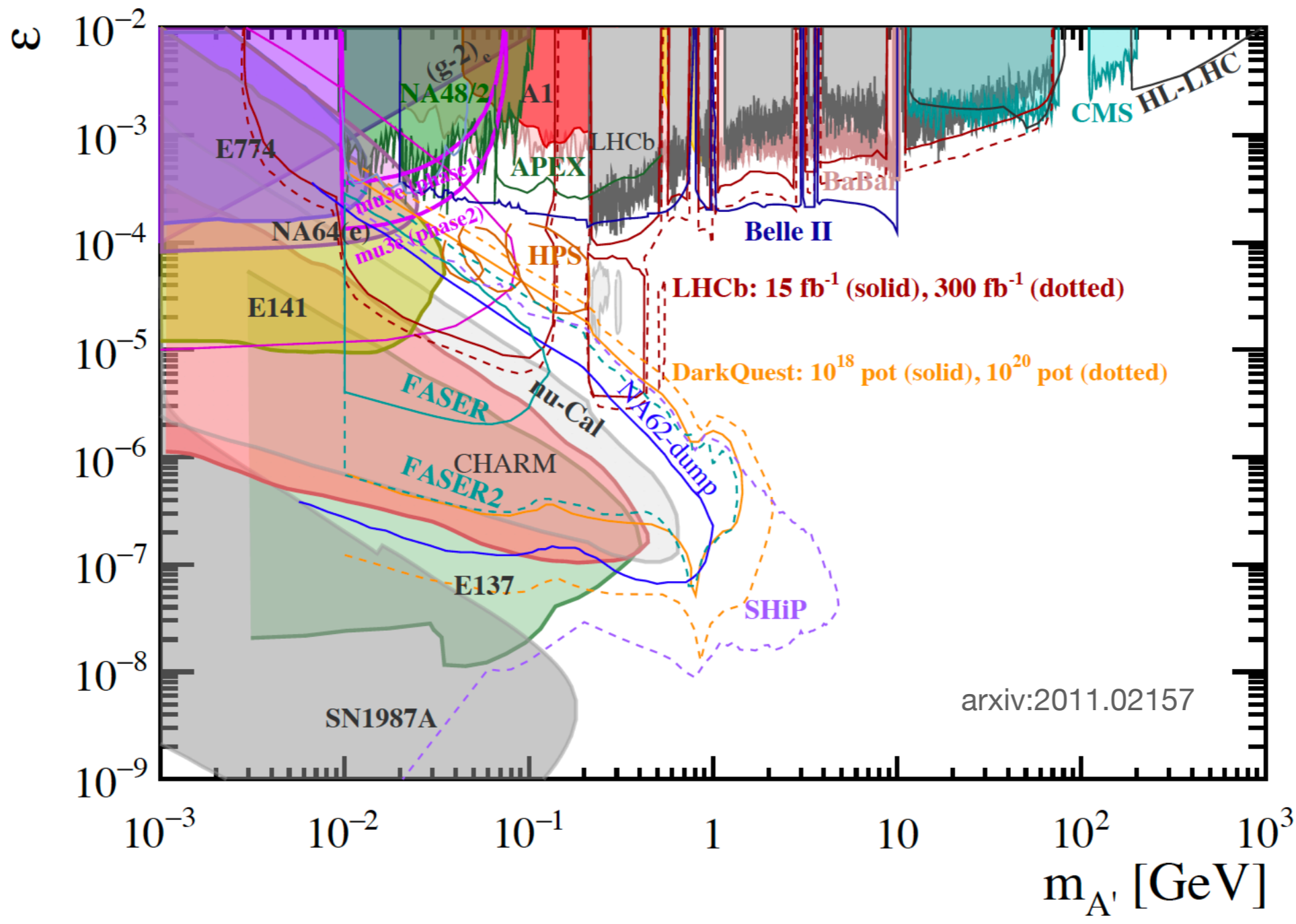


Invariant mass reconstruction precision at the level of 2%

NA64 collaboration, EPJ C 80 (2020) 12, 1159



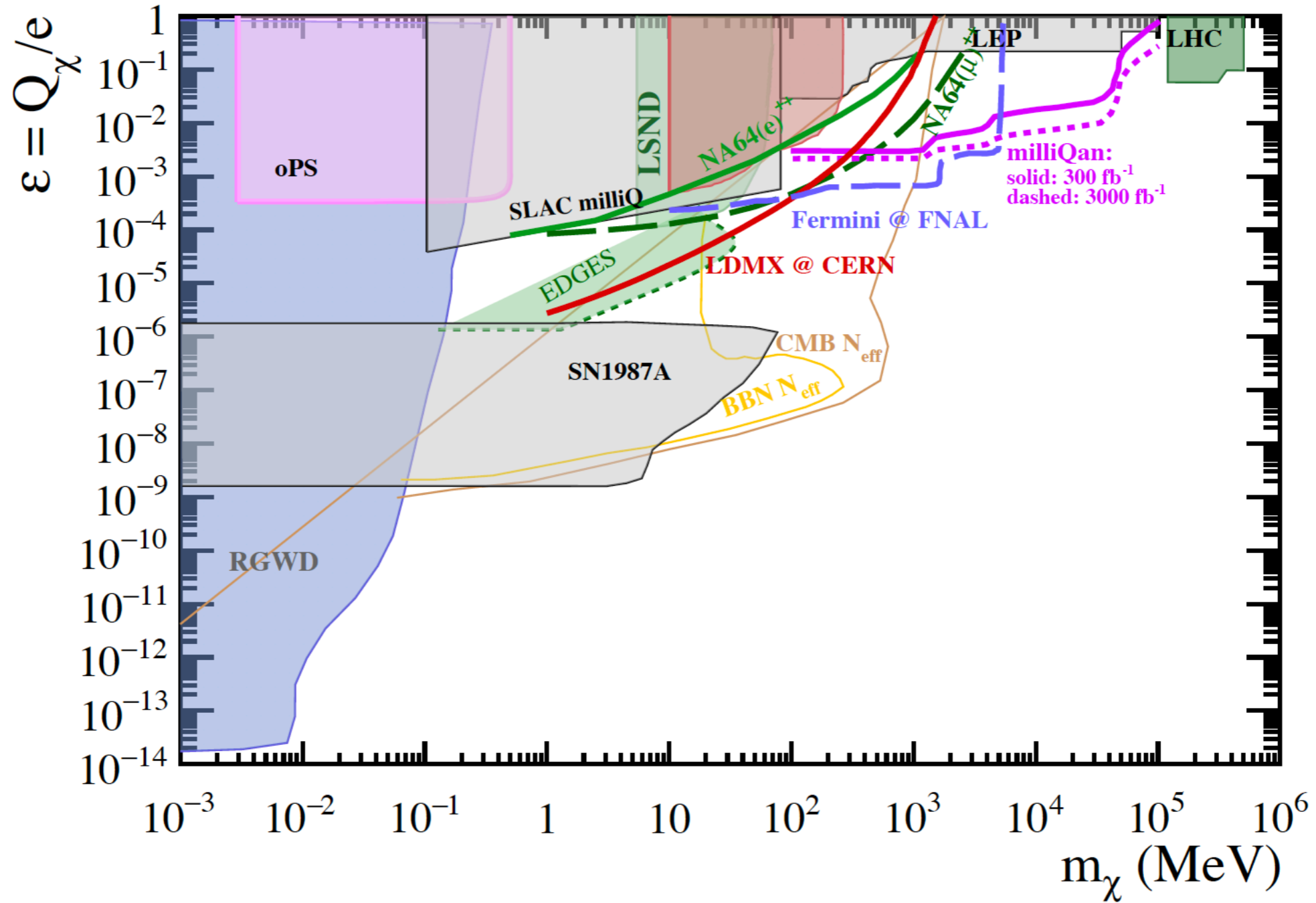
Dark photon visible searches



G.Lanfranchi, M.Pospelov and P.Schuster
arxiv:2011.02157



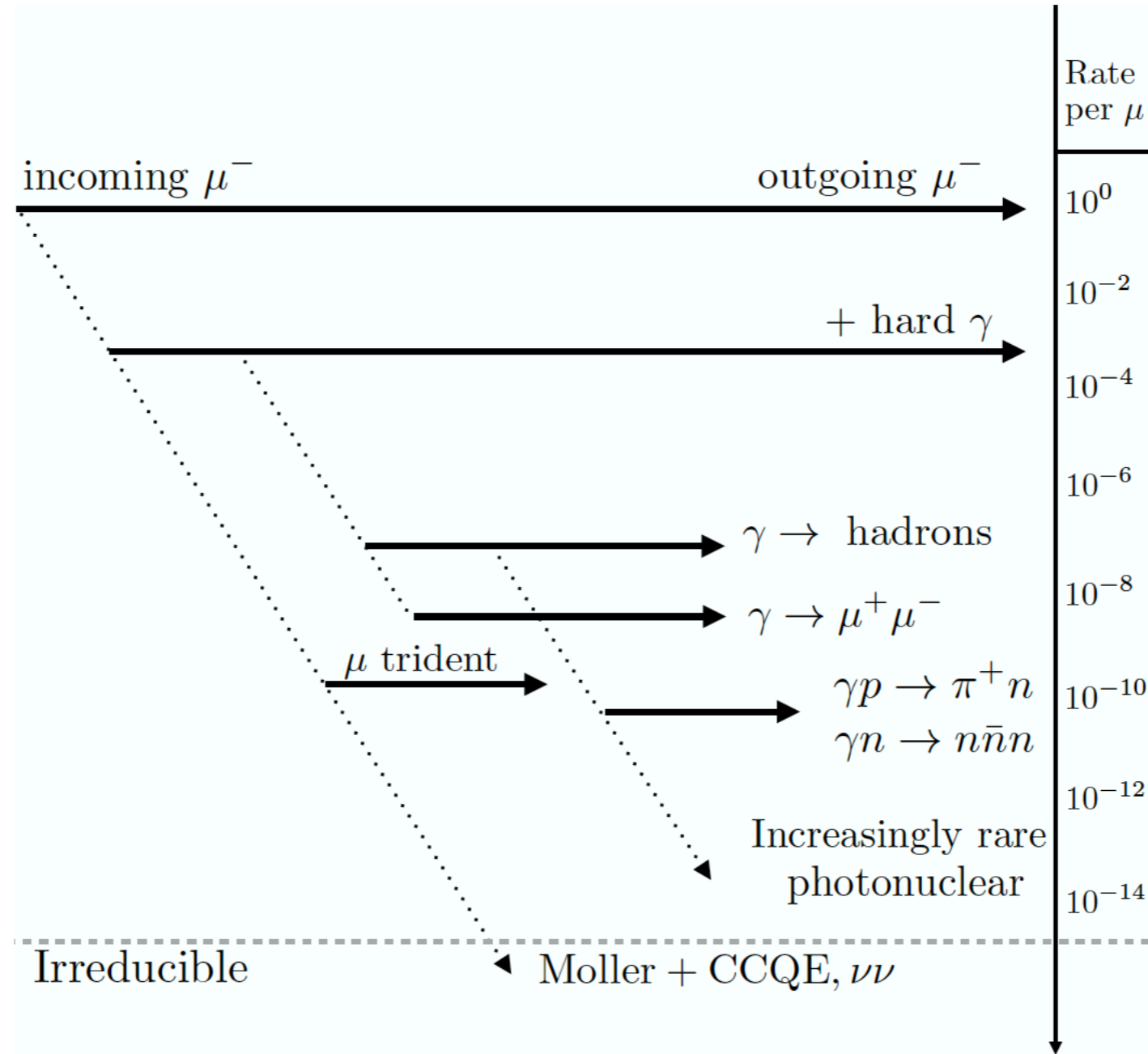
Millicharged particles



G.Lanfranchi, M.Pospelov and P.Schuster
arxiv:2011.02157



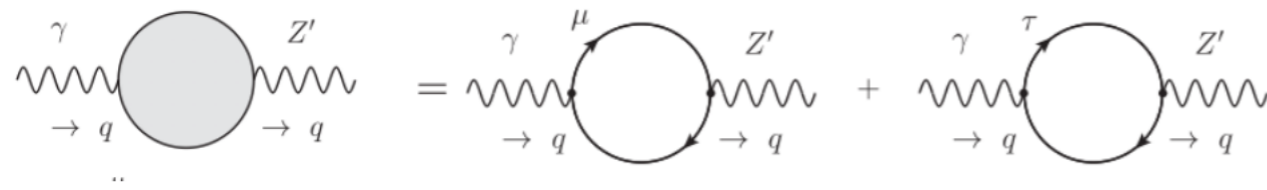
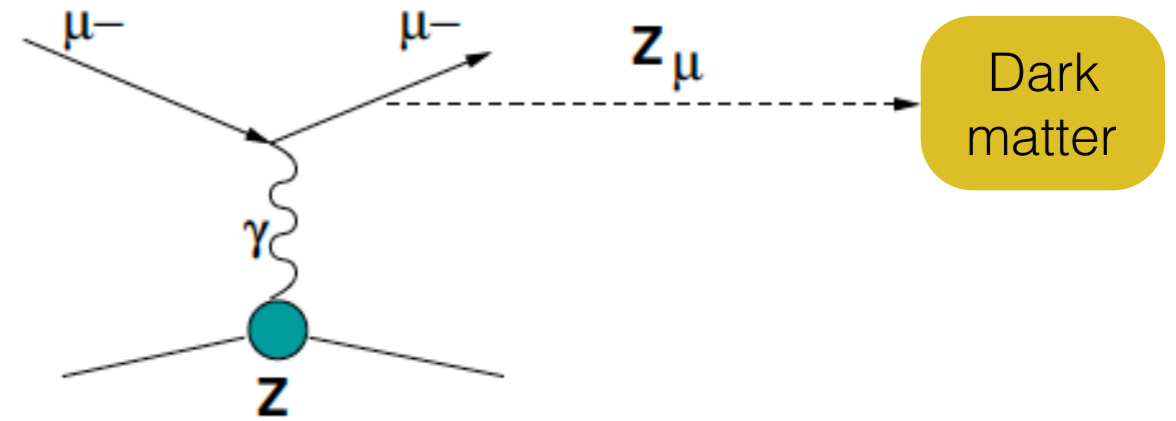
Background sources



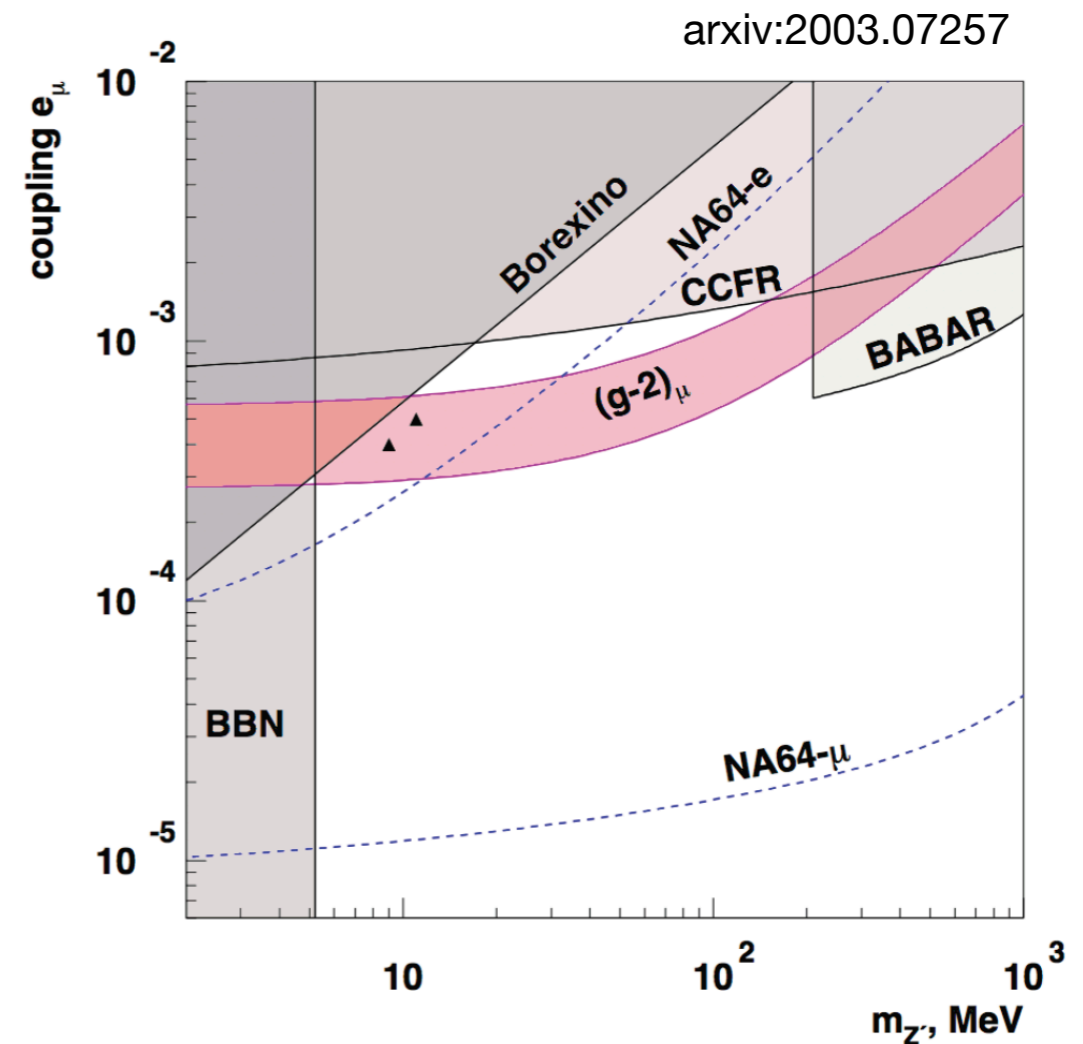
"M³: A New Muon Missing Momentum Experiment to Probe $(g-2)_\mu$ and Dark Matter at Fermilab", arxiv:1804.03144



Future prospects: NA64_μ physics goals

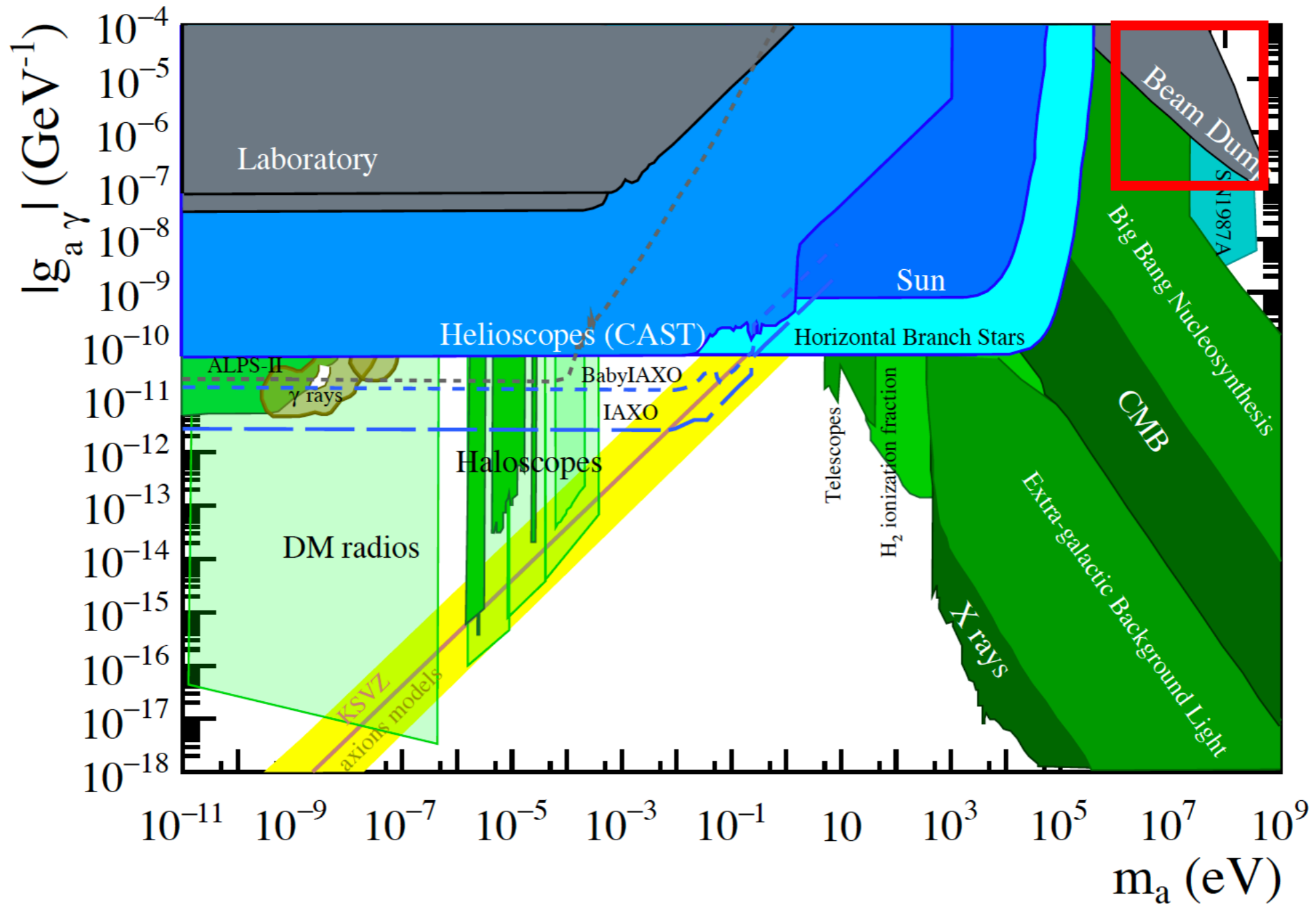


- γ - Z' mixing via loop
- **Complementarity between muon and electron mode:** If Z' is observed in muon mode should also be observed in e^- mode



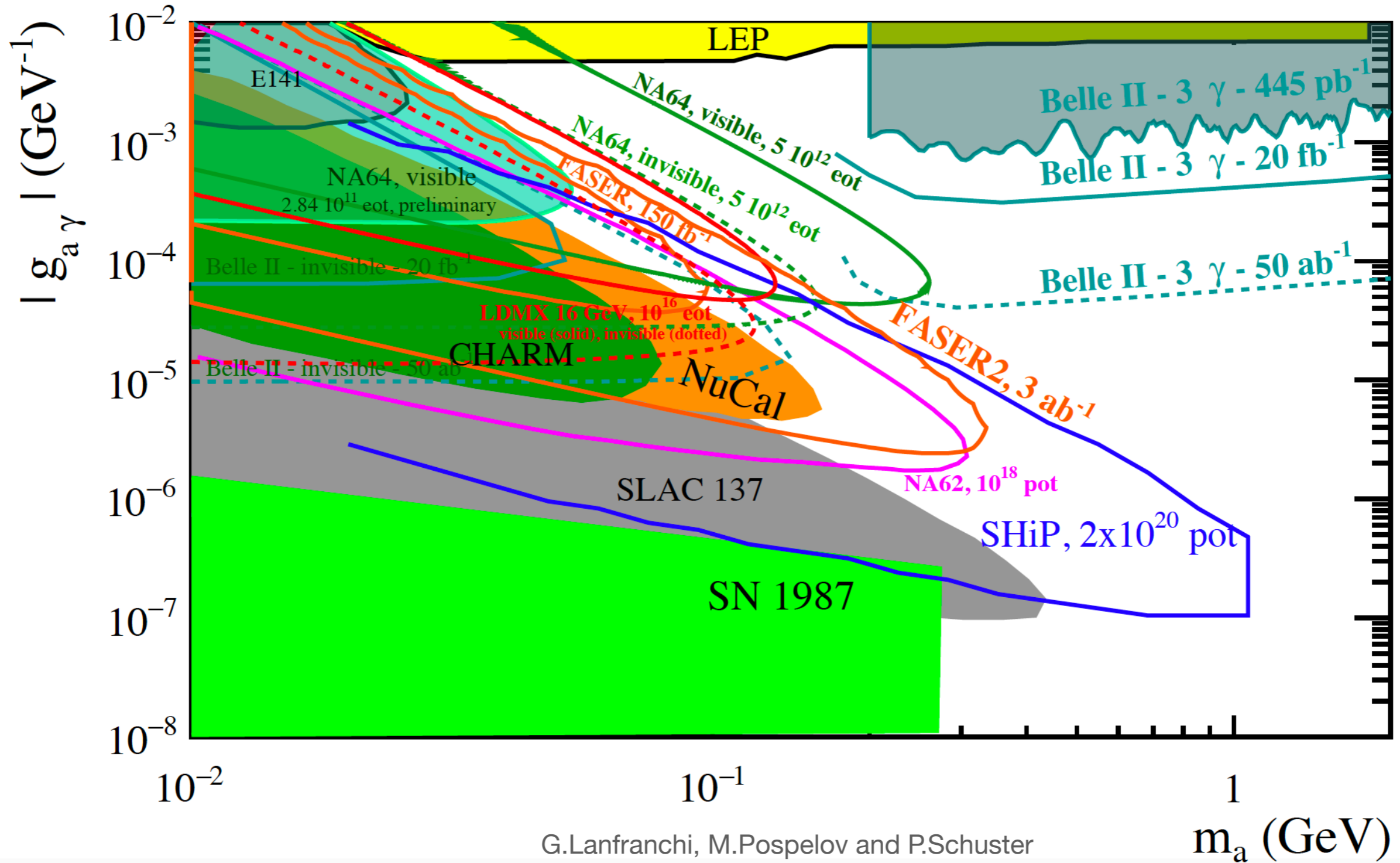


ALPs searches



G.Lanfranchi, M.Pospelov and P.Schuster
arxiv:2011.02157

Projections for ALPs



G.Lanfranchi, M.Pospelov and P.Schuster
arxiv:2011.02157