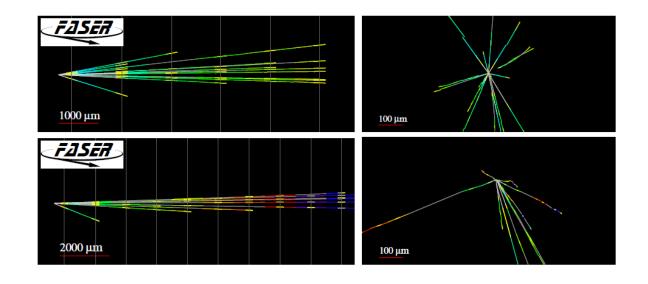


FASER v update and first neutrino interaction candidates

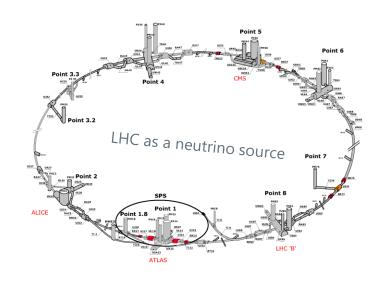
Tomoko Ariga (Kyushu University)
on behalf of the FASER Collaboration



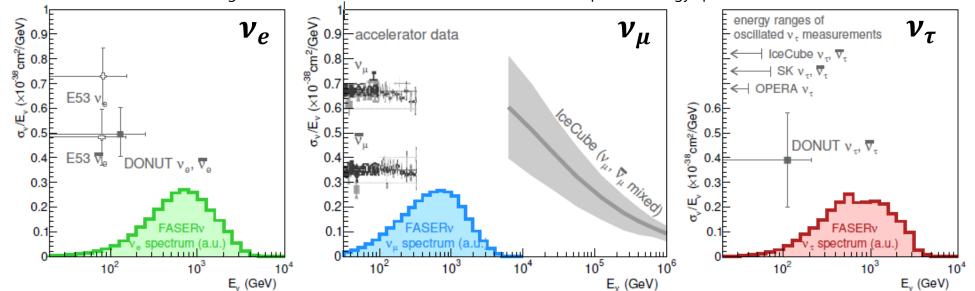
First neutrino interaction candidates at the LHC, <u>arXiv:2105.06197</u>

Physics motivations

- **Studying neutrinos in unexplored energy regime (TeV energies)**
 - Neutrinos from the LHC
 - First detection of collider neutrinos
 - High energy frontier of man-made neutrinos
 - Cross section measurements of different flavors at high energy
 - Probing neutrino-related models of new physics
 - From the other perspective, measurements of forward particle production

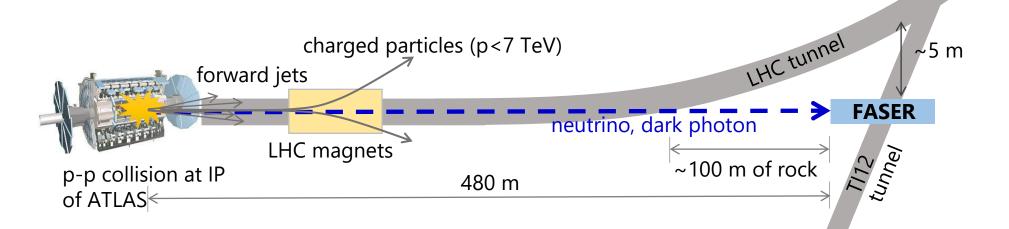






The FASER experiment

- FASER is a small and fast experiment at the LHC.
 - Will take data in LHC Run-3 (2022-2024).
- FASER (new particle searches) approved by CERN in Mar. 2019.
 - Targeting light, weakly-coupled new particles at low p_T .
 - Funded by the Heising-Simons and Simons Foundations with support from CERN.
- **FASER***v* (neutrino measurements) approved by CERN in Dec. 2019.
 - First measurements of neutrinos from a collider and in unexplored energy regime.
 - Funded by the Heising-Simons Foundation, ERC, JSPS and the Mitsubishi Foundation.



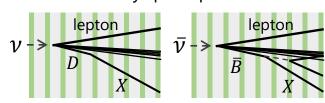


Physics potential:

high-energy neutrino interactions

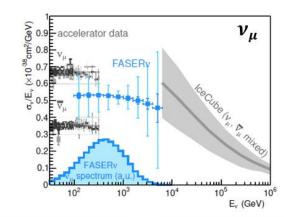
- Primary goal: cross section measurements of different flavors at TeV energies
 - where no such measurements currently exist.
- NC measurements
 - Could constrain neutrino non-standard interactions (NSI).
- Neutrino CC interaction with charm production $(vs \rightarrow lc)$
 - Study the strange quark content.
 - Probe inconsistency between the predictions and the LHC data [Eur. Phys. J. C77 (2017) 367].
- Neutrino CC interaction with beauty production
 - Has never been detected.

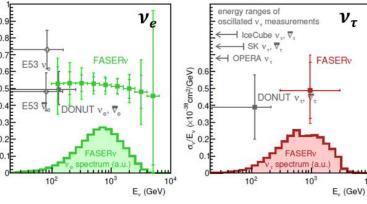
CC heavy quark production

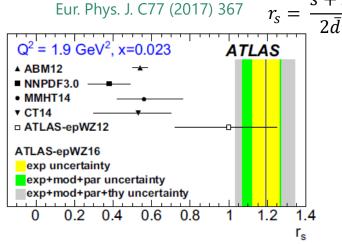


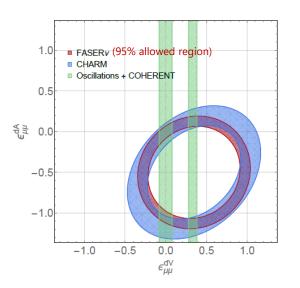
FASER Collaboration, <u>Eur. Phys. J. C 80 (2020) 61</u>, arXiv:1908.02310

A. Ismail, R.M. Abraham, F. Kling, <u>Phys. Rev. D 103, 056014 (2021)</u>, arXiv:2012.10500



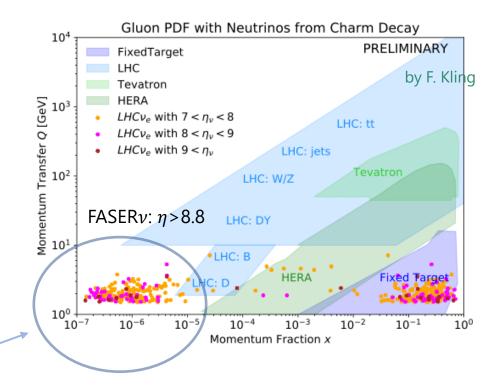




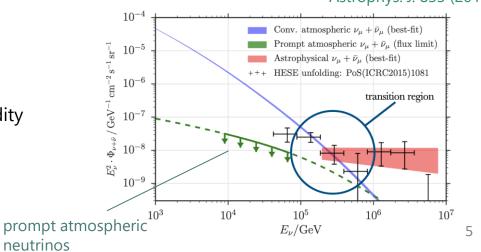


Physics potential: forward particle production

- Neutrinos produced in the forward direction at the LHC originate from the decay of hadrons, mainly pions, kaons, and charm particles.
- Forward particle production is poorly constrained by other LHC experiments.
- FASER ν 's measurements provide novel input to validate/improve generators.
 - First data on forward kaon, hyperon, charm
- Neutrinos from charm decay could allow to
 - test transition to small-x factorization, see effects of gluon saturation, constrain low-x gluon PDF, probe intrinsic charm.
- Relevant for neutrino telescopes (such as IceCube).
 - In order for IceCube to make precise measurements of the cosmic neutrino flux, accelerator measurements of high energy and large rapidity charm production are needed.
 - As 7+7 TeV p-p collision corresponds to 100 PeV proton interaction in fixed target mode, a direct measurement of the prompt neutrino production would provide important basic data for current and future high-energy neutrino telescopes.



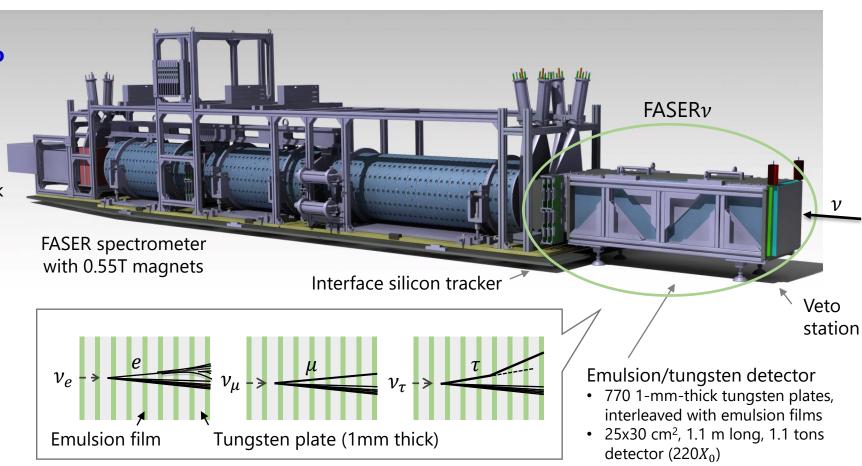
IceCube Collaboration. Astrophys. J. 833 (2016)



neutrinos

The FASER v detector for LHC Run-3

- Emulsion/tungsten detector, interface silicon tracker, and veto station will be placed in front of the FASER main detector.
- Allow to distinguish all flavor of neutrino interactions.
 - **Muon identification** by their track length in the detector $(8\lambda_{int})$
 - **Muon charge identification** with hybrid configuration \rightarrow distinguishing ν_{μ} and $\bar{\nu}_{\mu}$
 - Neutrino energy measurement with ANN by combining topological and kinematical variables



Expected neutrino event rate in LHC Run-3

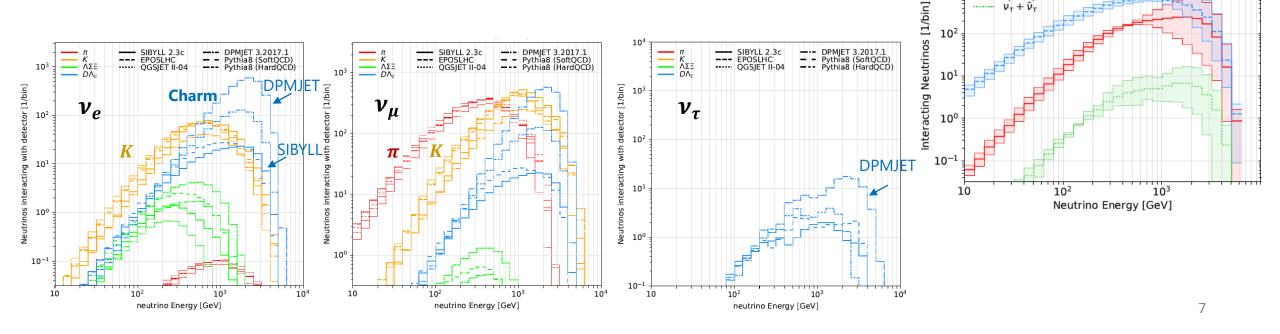
- A high-intensity beam of neutrinos will be produced in the far-forward direction.
- FASERν will be centered on the LOS (in the FASER trench) to maximizes fluxes of all neutrino flavors.

Expected number of CC interactions in FASER ν during LHC Run-3 (150 fb⁻¹)

		1		
Generators		$\mathrm{FASER} u$		
light hadrons	heavy hadrons	$ u_e + \bar{\nu}_e $	$ u_{\mu} + \bar{ u}_{\mu} $	$ u_{\tau} + \bar{\nu}_{\tau} $
SIBYLL	SIBYLL	1343	6072	21.2
DPMJET	DPMJET	4614	9198	131
EPOSLHC	Pythia8 (Hard)	2109	7763	48.9
QGSJET	Pythia8 (Soft)	1437	7162	24.5
Combin	ation (all)	2376^{+2238}_{-1032}	7549^{+1649}_{-1476}	$56.4_{-35.1}^{+74.5}$
Combination (w/o DPMJET)		1630^{+479}_{-286}	7000^{+763}_{-926}	$31.5^{+17.3}_{-10.3}$

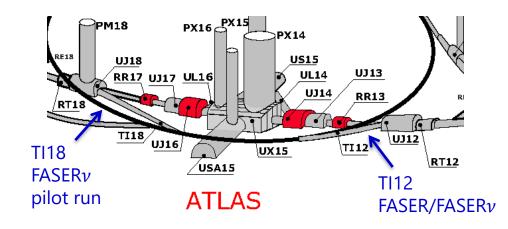
Differences between the generators checked with the same propagation model (RIVET-module)

FASERV

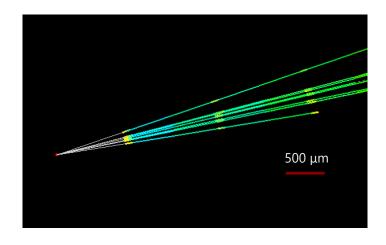


Pilot run in 2018 (LHC Run-2)

Aiming to demonstrate neutrino detection at the LHC for the first time

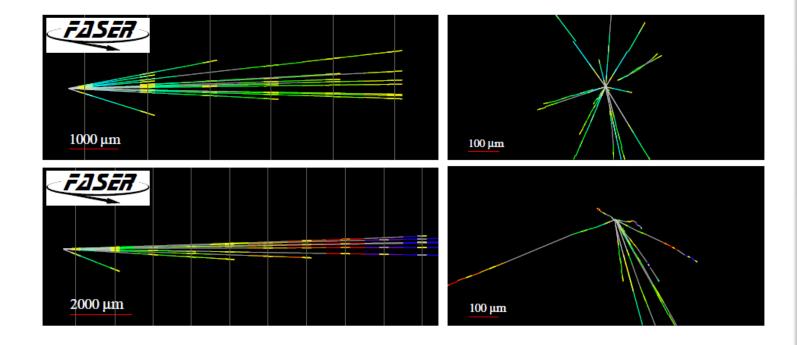






- Aims: charged particle flux measurement and neutrino detection
- We performed measurements in the tunnels TI18 and TI12, 480 m from the ATLAS IP.
- For neutrino detection, a 30 kg emulsion detector was installed in TI18 and 12.2 fb⁻¹ data was collected.

Neutrino interaction candidates



First neutrino interaction candidates at the LHC, arXiv:2105.06197

UCI-TR-2021-04, KYUSHU-RCAPP-2020-04, CERN-EP-2021-087

First neutrino interaction candidates at the LHC

Henso Abreu, ¹ Yoav Afik, ¹ Claire Antel, ² Akitaka Ariga, ^{3,4} Tomoko Ariga, ^{5,*} Florian Bernlochner, ⁶ Tobias Boeckh, 6 Jamie Boyd, 7 Lydia Brenner, 7 Franck Cadoux, 2 David W. Casper, 8 Charlotte Cavanagh, 9 Francesco Cerutti, 7 Xin Chen, 10 Andrea Coccaro, 11 Monica D'Onofrio, 9 Candan Dozen, 10 Yannick Favre, 2 Deion Fellers, 12 Jonathan L. Feng,⁸ Didier Ferrere,² Stephen Gibson,¹³ Sergio Gonzalez-Sevilla,² Carl Gwilliam,⁹ Shih-Chieh Hsu,¹⁴ Zhen Hu, 10 Giuseppe Iacobucci, 2 Tomohiro Inada, 10 Sune Jakobsen, 7 Enrique Kajomovitz, 1 Felix Kling, 15 Umut Kose, 7 Susanne Kuehn, 7 Helena Lefebvre, 13 Lorne Levinson, 16 Ke Li, 14 Jinfeng Liu, 10 Chiara Magliocca, 2 Josh McFayden, 17 Sam Meehan, 7 Dimitar Mladenov, 7 Mitsuhiro Nakamura, 18 Toshiyuki Nakano, 18 Marzio Nessi, 7 Friedemann Neuhaus, 19 Laurie Nevay, 12 Hidetoshi Otono, 5 Carlo Pandini, 2 Hao Pang, 10 Lorenzo Paolozzi, 2 Brian Petersen, Francesco Pietropaolo, Markus Prim, Michaela Queitsch-Maitland, Filippo Resnati, Hiroki Rokujo, 18 Marta Sabaté-Gilarte, Jakob Salfeld-Nebgen, Osamu Sato, Paola Scampoli, Kristof Schmieden, Matthias Schott, 19 Anna Sfyrla, 2 Savannah Shively, 8 John Spencer, 14 Yosuke Takubo, 21 Ondrej Theiner, 2 Eric Torrence, 12 Sebastian Trojanowski, 22 Serhan Tufanli, 7 Benedikt Vormwald, 7 Di Wang, 10 and Gang Zhang 10

(FASER Collaboration) Department of Physics and Astronomy, Technion—Israel Institute of Technology, Haifa 32000, Israel Département de Physique Nucléaire et Corpusculaire, University of Geneva, CH-1211 Geneva 4, Switzerland

³Albert Einstein Center for Fundamental Physics, Laboratory for High Energy Physics,

University of Bern, Sidlerstrasse 5, CH-3012 Bern, Sustzerland

⁴Department of Physics, Chiba University, 1-33 Yayoi-cho Inage-ku, Chiba, 263-8522, Japan Kyushu University, Nishi-ku, 819-0395 Fukuoka, Japan
 Universität Bonn, Regina-Pacis-Weg 3, D-53113 Bonn, Germany

**Department of Physics, Tsinghua University, Beijing, China

**Department of Physics and Astronomy, University of California, Irsine, CA 92697-4575, USA

**University of Liverpool, Liverpool L69 3BX, United Kingdom

**Department of Physics, Tsinghua University, Beijing, China ¹¹INFN Sezione di Genova, Via Dodecaneso, 33-16146, Genova, Italy ¹²University of Oregon, Eugene, OR 97403, USA

Tuniversity of Oregon, Eugene, OR 97403, USA
 Royal Holtoway, University of Dandon, Esham, TW20 0EX, UK
 Department of Physics, University of Washington, PO Box 33:560, Scattle, WA 93:95-1560, USA
 SLAC National Accelerator Laboratory, 2575 Sand Hill Road, Menlo Park, CA 94025, USA
 Department of Particle Physics and Astrophysics,

Weizmann Institute of Science, Rehovot 76100, Israel

17 Department of Physics & Astronomy, University of Sussex. Sussex House, Falmer, Brighton, BN1 9RH, United Kingdom

¹⁸ Nagoya University, Furo-cho, Chikusa-ku, Nagoya 464-8602, Japan ¹⁹Institut für Physik, Universität Mainz, Mainz, Germany

²⁰ Dipartimento di Fisica "Ettore Pancini", Università di Napoli Federico II, Complesso Universitario di Monte S. Angelo, 1-80126 Napolo, Italy ²¹Institute of Particle and Nuclear Study, KEK, Oho 1-1, Tsukuba, Ibaruki 305-0801, Japan ²² Astrocent, Nicolaus Copernicus Astronomical Center Polish Academy of Sciences, ul. Bartycka 18, 00-716 Warsaw, Poland (Dated: May 14, 2021)

 $FASER\nu$ at the CERN Large Hadron Collider (LHC) is designed to directly detect collider neutrinos for the first time and study their cross sections at TeV energies, where no such measurements currently exist. In 2018, a pilot detector employing emulsion films was installed in the far-forward region of ATLAS, 480 m from the interaction point, and collected 12.2 fb⁻¹ of proton-proton collision data at a center-of-mass energy of 13 TeV. We describe the analysis of this pilot run data and the observation of the first neutrino interaction candidates at the LHC. This milestone paves the way for high-energy neutrino measurements at current and future colliders.

I. INTRODUCTION

There has been a longstanding interest in detecting neutrinos produced at colliders [1-6], but to date no col-

. Corresponding author: tomoko.ariga@cern.ch

07

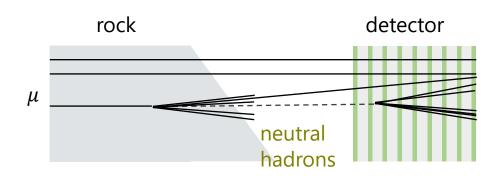
lider neutrino has ever been directly detected. Protonproton (pp) collisions at a center-of-mass energy of 14 TeV during LHC Run-3, with an expected integrated luminosity of 150 fb⁻¹, will produce a high-intensity beam of $\mathcal{O}(10^{12})$ neutrinos in the far-forward direction with mean interaction energy of about 1 TeV. FASER [7] is designed to detect these neutrinos and study their prop

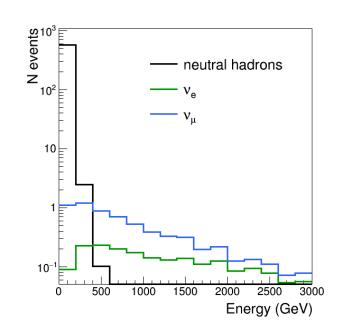
Background estimation

- The pilot detector lacked the ability to identify muons given its depth of only $0.6\lambda_{int}$, much shorter than the $8\lambda_{int}$ of the full FASER ν detector.
- → Separation from neutral hadron BG (produced by muons) is much harder than the physics run.
- Muons rarely produce neutral hadrons in upstream rock, which can mimic neutrino interaction vertices.
- The produced neutral hadrons are low energy → discriminate by vertex topology.

The production rates of neutral hadrons per incident muon

	Negative Muons	Positive Muons
K_L	3.3×10^{-5}	9.4×10^{-6}
K_S	8.0×10^{-6}	2.3×10^{-6}
n	2.6×10^{-5}	7.7×10^{-6}
$ar{n}$	1.1×10^{-5}	3.2×10^{-6}
Λ	3.5×10^{-6}	1.8×10^{-6}
$ar{\Lambda}$	2.8×10^{-6}	8.7×10^{-7}





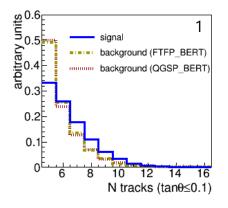
Variables for the BDT analysis

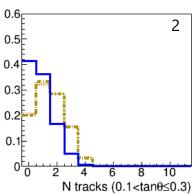
5 variables used in the analysis

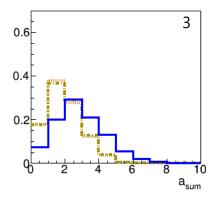
- 1. the number of tracks with $\tan \theta < =0.1$ with respect to the beam direction
- 2. the number of tracks with $0.1 < \tan \theta < = 0.3$ with respect to the beam direction
- 3. the absolute value of vector sum of transverse angles calculated considering all the tracks as unit vectors in the plane transverse to the beam direction (a_{sum})
- 4. for each track in the event, calculate the mean value of opening angles between the track and the others in the plane transverse to the beam direction, and then take the maximum value in the event (ϕ_{mean})
- 5. for each track in the event, calculate the ratio of the number of tracks with opening angle \leq 90 degrees and \geq 90 degrees in the plane transverse to the beam direction, and then take the maximum value in the event (r).

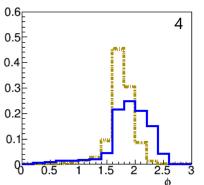
Concepts

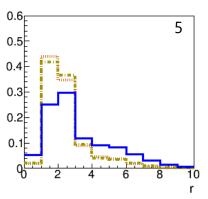
- The neutrino energy is higher than the neutral hadron energy. Higher energy, more particles are produced in forward direction, i.e. tan(theta)<0.1. → variable 1, 2
- Momentum in the transverse plane is more balanced in hadron interactions than neutrino CC and NC interactions. Outgoing leptons in neutrino interactions take a major energy, which distorts this variable.
 → variable 3
- For CC interactions, we expect the outgoing lepton and hadron system are back to back in the transverse plane. → variable 4, 5





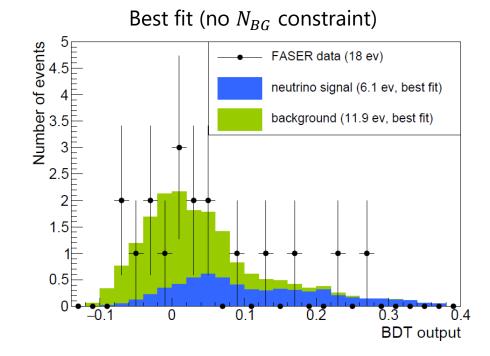






Results

- Analyzed target mass 11 kg
- 18 neutral vertices were selected
 - by applying # of charged particle \geq 5, etc.
 - Expected signal $3.3^{+1.7}_{-0.9}$ events, BG 11.0 events
- In the BDT analysis, an excess of neutrino signal is observed. Statistical significance 2.7σ from null hypothesis
- This result demonstrates detection of neutrinos at the LHC.



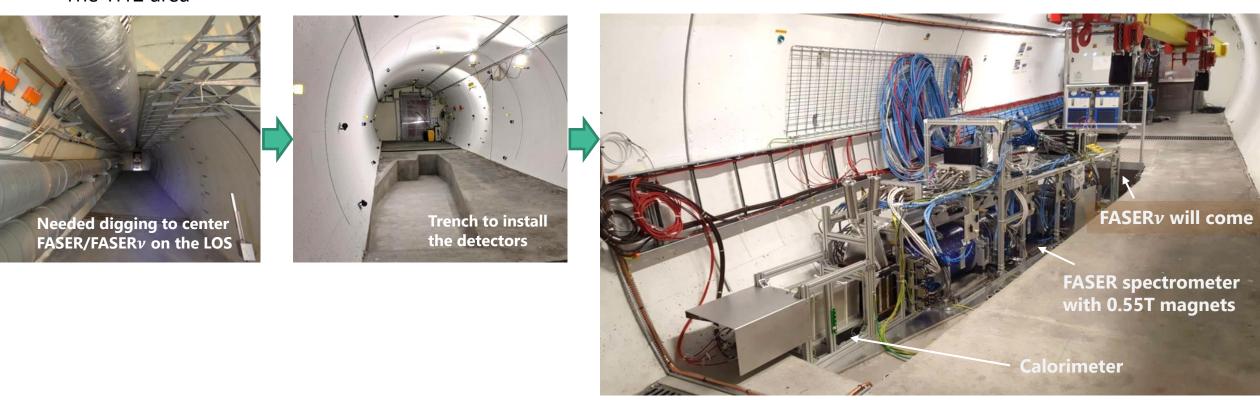
We are currently preparing for data taking in LHC Run-3. With a deeper detector and lepton identification capability, FASER ν will perform better than this pilot detector.

Schedule



Preparation towards LHC Run-3

The TI12 area



The FASER main detector was successfully installed into the TI12 tunnel in March 2021. Acknowledge great support from many CERN teams involved in the work

FASER_v installation test





Emulsion detector preparation

- Emulsion gel and film production facilities in Nagoya have been set up in 2020.
- We are testing mass production.

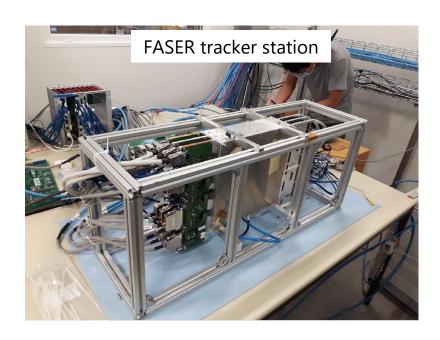


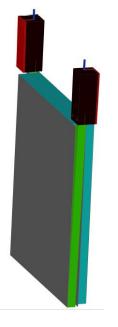




Interface tracker (IFT) and veto system

- **IFT** will use the same design as the tracker station in the FASER spectrometer.
 - Silicon strip detector with ATLAS SCT barrel modules.
 - Machining of the tracker planes was completed.
 - Mounting SCT modules to the planes is in progress.
 - Commissioning of the planes, assembly to the station, commissioning of the station will be performed by early-July.
- Veto station consists of two 2-cm scintillators and WLS (Wave Length Shifting) bars with two PMTs.
 - Scintillators, WLS bars and PMTs were ordered, now all at CERN, and PMTs were tested.
 - Assembly of the scintillator units are ongoing.
 - Commissioning with cosmic rays is planned during the summer.





Veto dimension 30×35 cm²



Summary and prospects

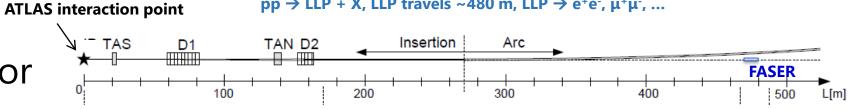
- FASERν at the CERN LHC is designed to directly detect collider neutrinos for the first time and study their properties at TeV energies.
- We have detected first neutrino interaction candidates at the LHC in the 2018 pilot run data.
 - arXiv:2105.06197
- We expect to collect ~10000 CC interactions (distinguishing the flavors) in LHC-Run3 (2022-2024). Preparation for the data taking is in progress.
- Also planning neutrino measurements in the HL-LHC era.
 - A large detector for precision v_{τ} physics with 10-30 tons of target



Backup

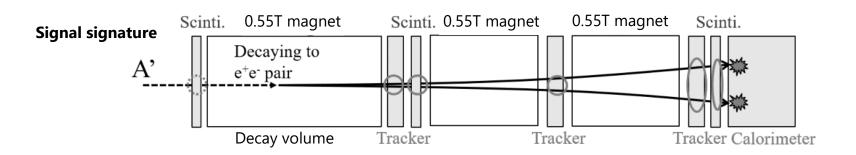
The FASER main detector

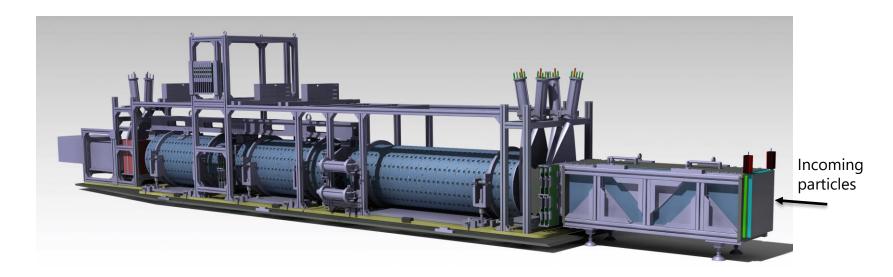




Technical proposal: FASER, CERN-LHCC-2018-036; LHCC-P-013

FASER's physics reach for long-lived particles, Phys. Rev. D 99, 095011





The detector consists of:

- Scintillator veto
- 1.5 m long decay volume
- 2 m long spectrometer
- EM calorimeter

FASER ν : $\nu_e + \bar{\nu}_e$ 10^{11} Neutrinos [1/bin] 10¹⁰ FASER ν : $\nu_{\mu} + \bar{\nu}_{\mu}$ Shower 10¹² Neutrinos [1/bin] 10₁₀ 10^{9} FASER ν : $\nu_{\tau} + \bar{\nu}_{\tau}$ 10¹⁰ Neutrinos [1/bin] 10^{7} 10² 10^{3} 10 Neutrino Energy [GeV]

Energy distributions of neutrinos passing though FASER ν

F. Kling, Forward Neutrino Fluxes at the LHC, arXiv:2105.08270

Radial spectrum

