

Laser Interferometer Space Antenna

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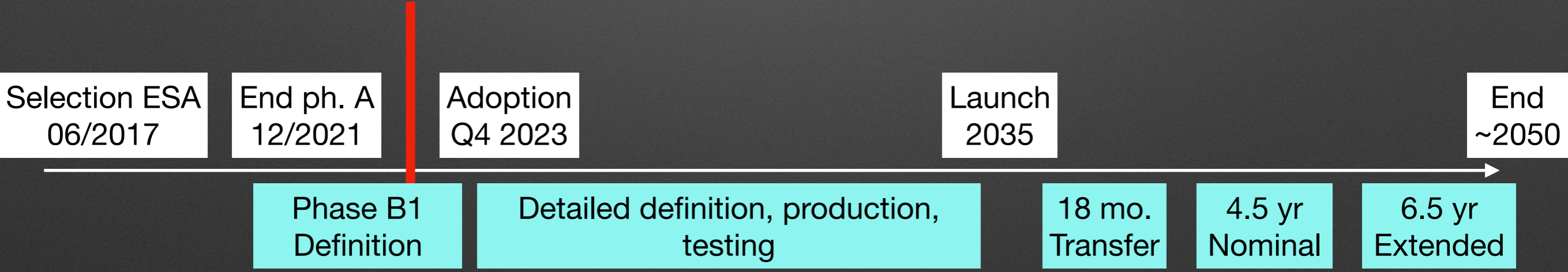
EUROPEAN SPACE AGENCY L3 MISSION

First ESA/NASA proposal in the 90's

Selected as flagship mission for the « Gravitational Universe » in 2017 (Cosmic Vision program), cost cap 1.3 Billion Euros

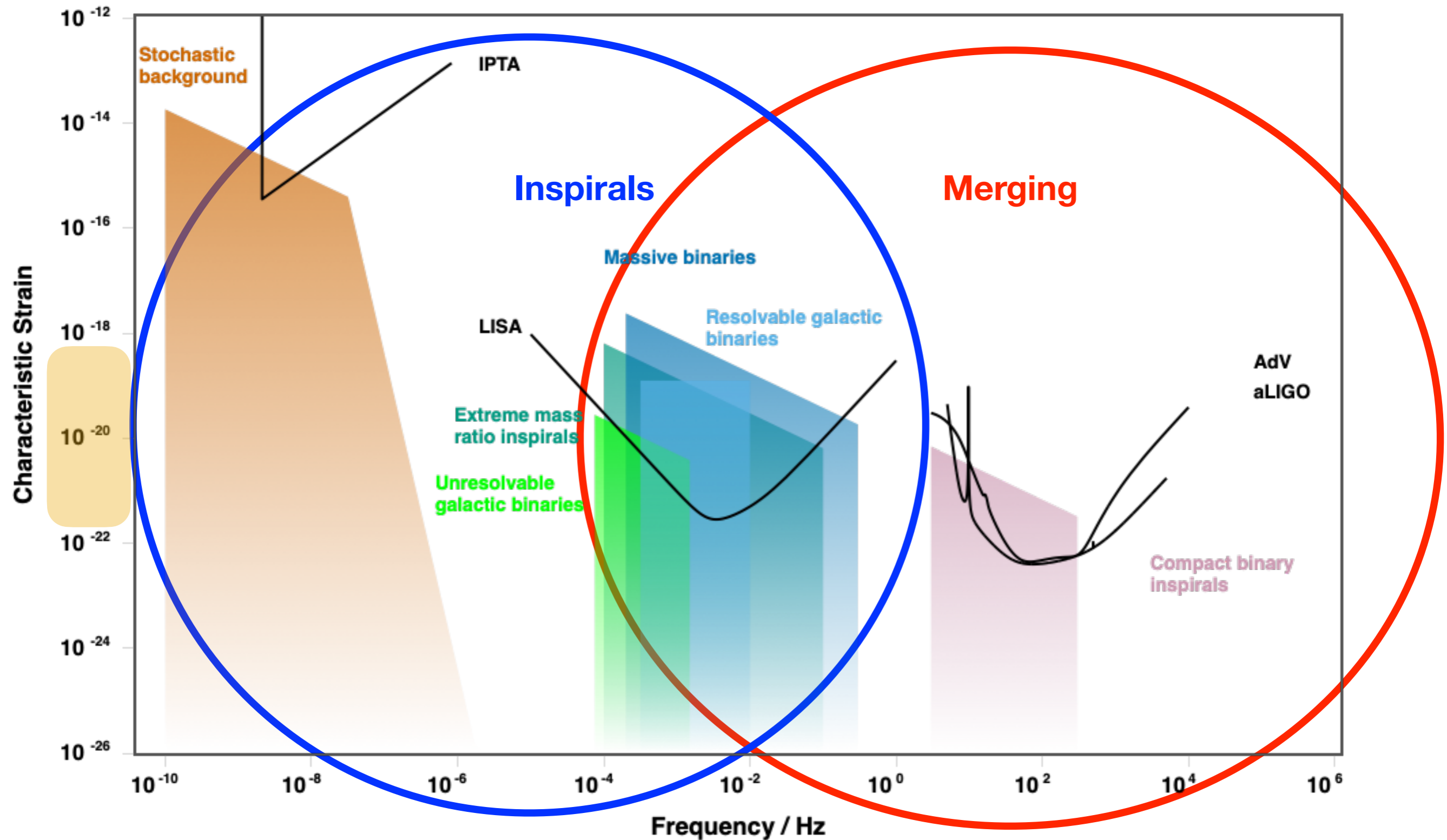
Expected launch mid-2030's

Adoption (initially) planned 2024 -> now 2023



Germany, France, Switzerland, Italy, UK, US, Belgium, Denmark, Spain, Finland, Hungary, Netherlands, Portugal, Romania, Sweden

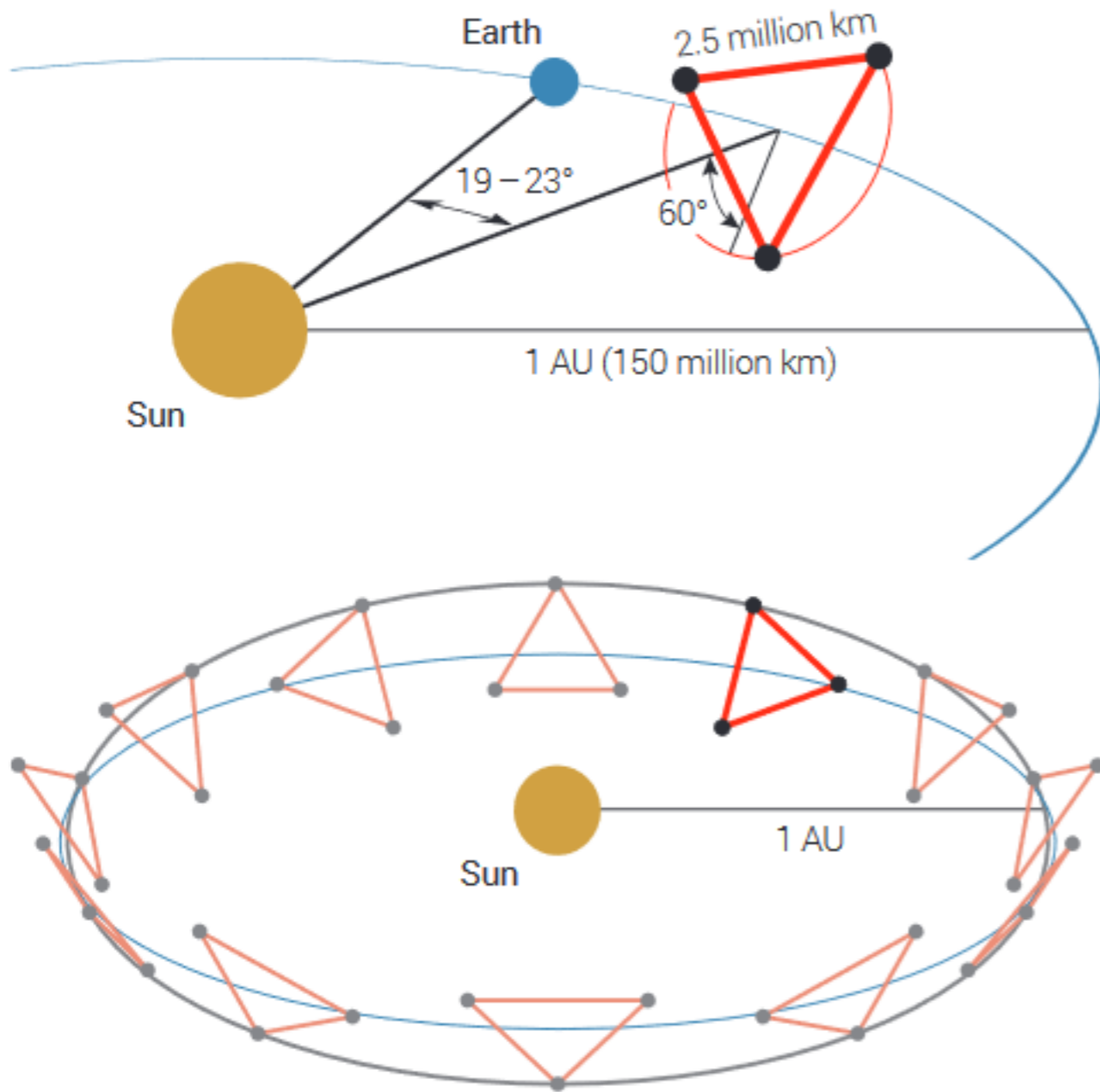
FREQUENCIES OF GRAVITATIONAL WAVES



$$f_{GW} = 2f_{orb}$$

$$A_{GW} \propto \frac{M_c^{5/3} f^{2/3}}{D}$$

THE LISA CONSTELLATION



3 spacecraft 20 days behind the Earth

2.5 million km separation

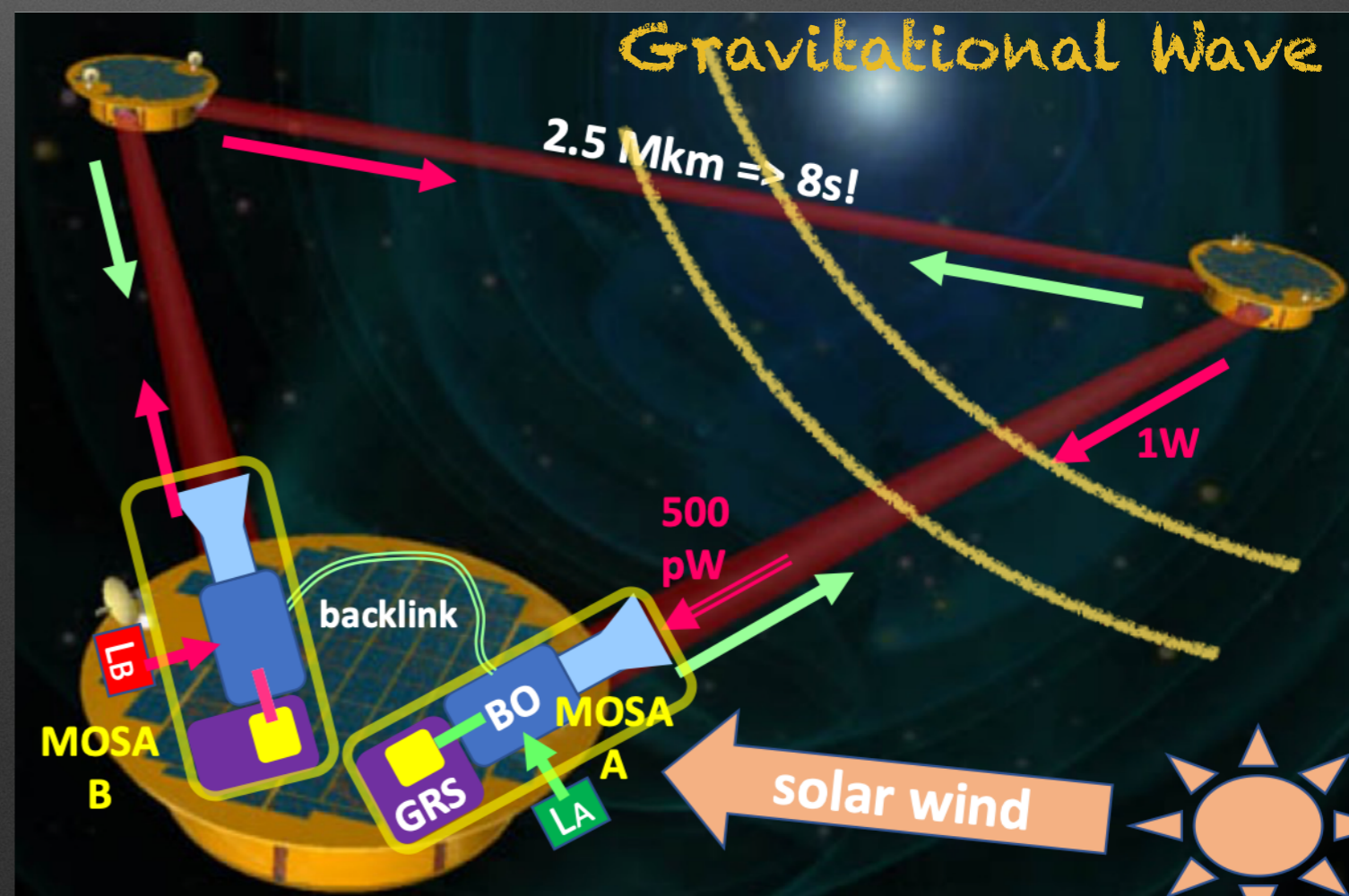
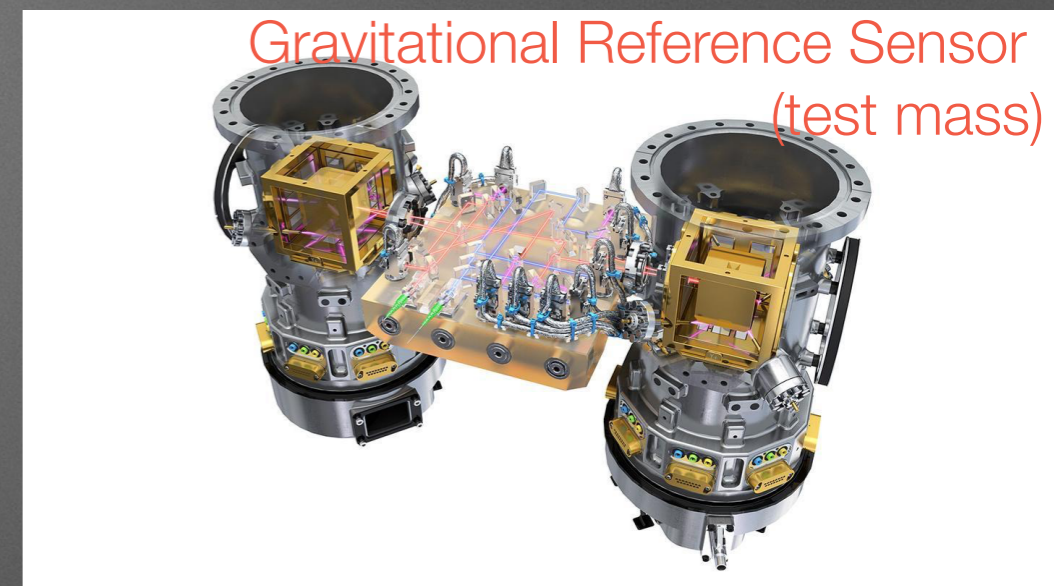
Distance measurements precision: 10^{-12} m

MEASUREMENT CONCEPT

Principle: measure distance variation between test masses

-> with laser interferometer

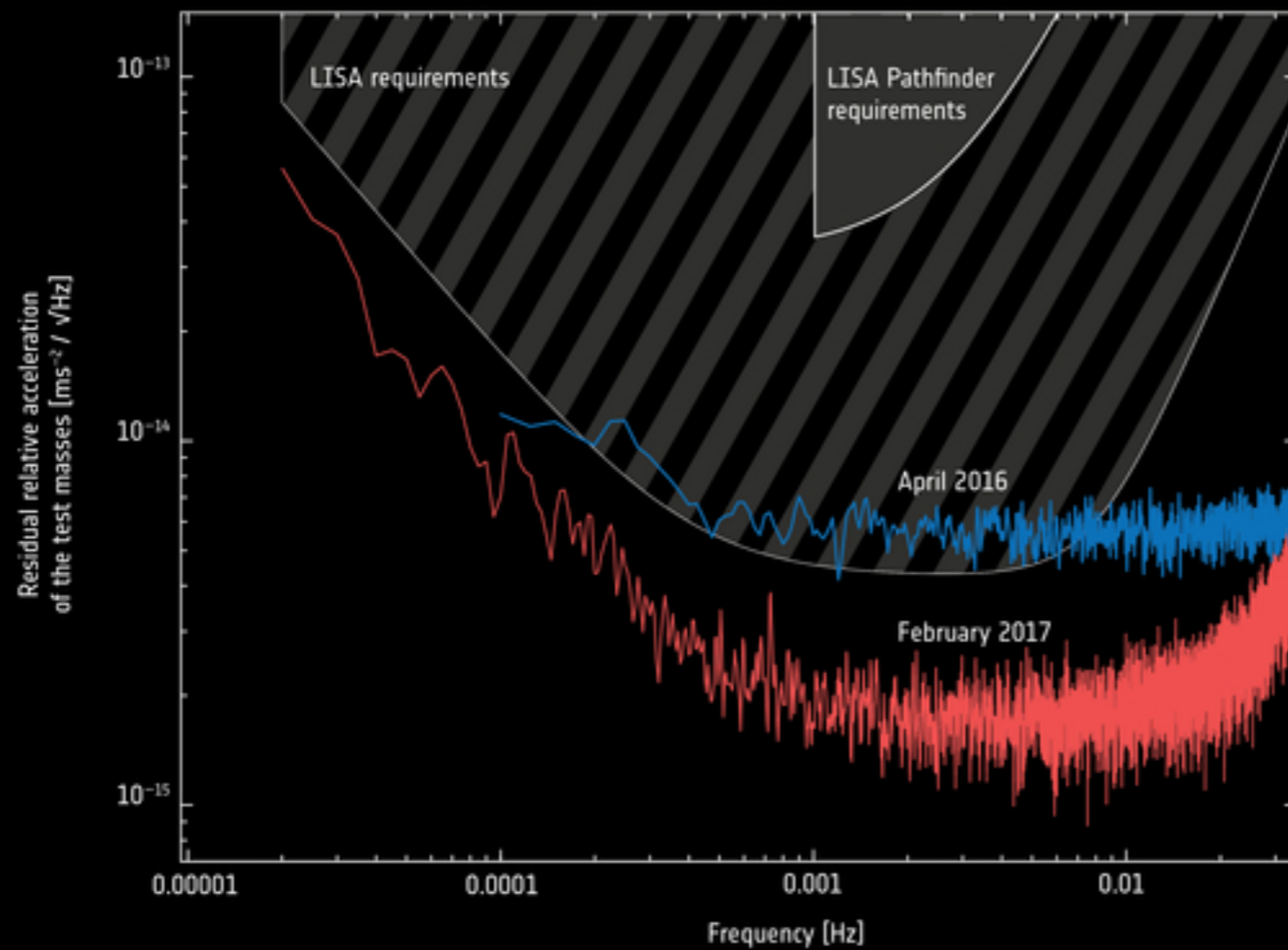
6 laser links: smart combination to extract GW signal (Time Delay Interferometry)



MOSA: Moving optical sub-assembly

WE KNOW IT WORKS !

LISA Pathfinder mission (2015) : following the path of the test masses works



LISA Pathfinder collaboration

LISA DATA ANALYSIS

Millions of superposed sources, very different from (current) ground-based detectors

Dream of a beautiful symphony



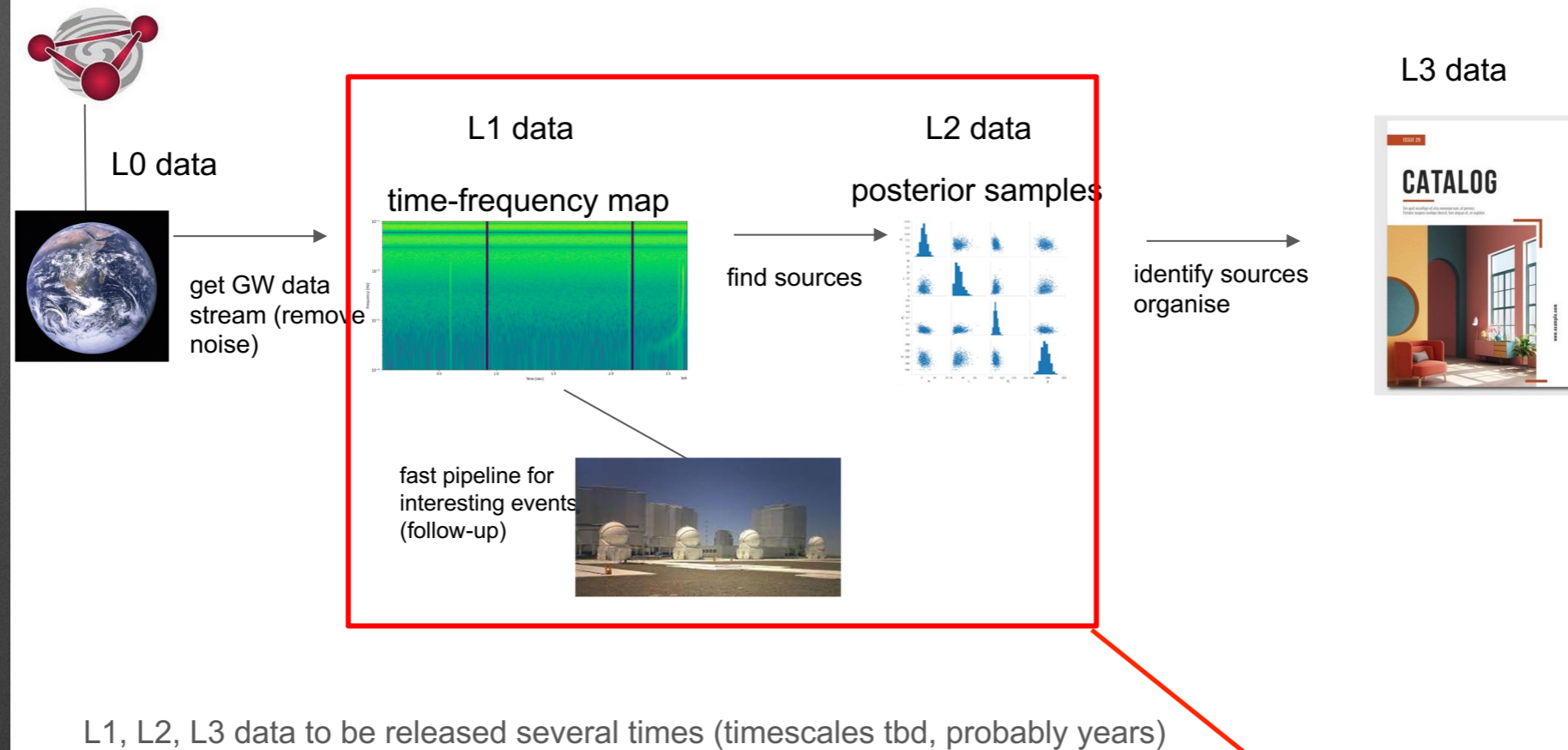
LISA DATA ANALYSIS

Millions of superposed sources, very different from (current) ground-based detectors



LISA DATA ANALYSIS

Global pipeline structure (L0-L3)

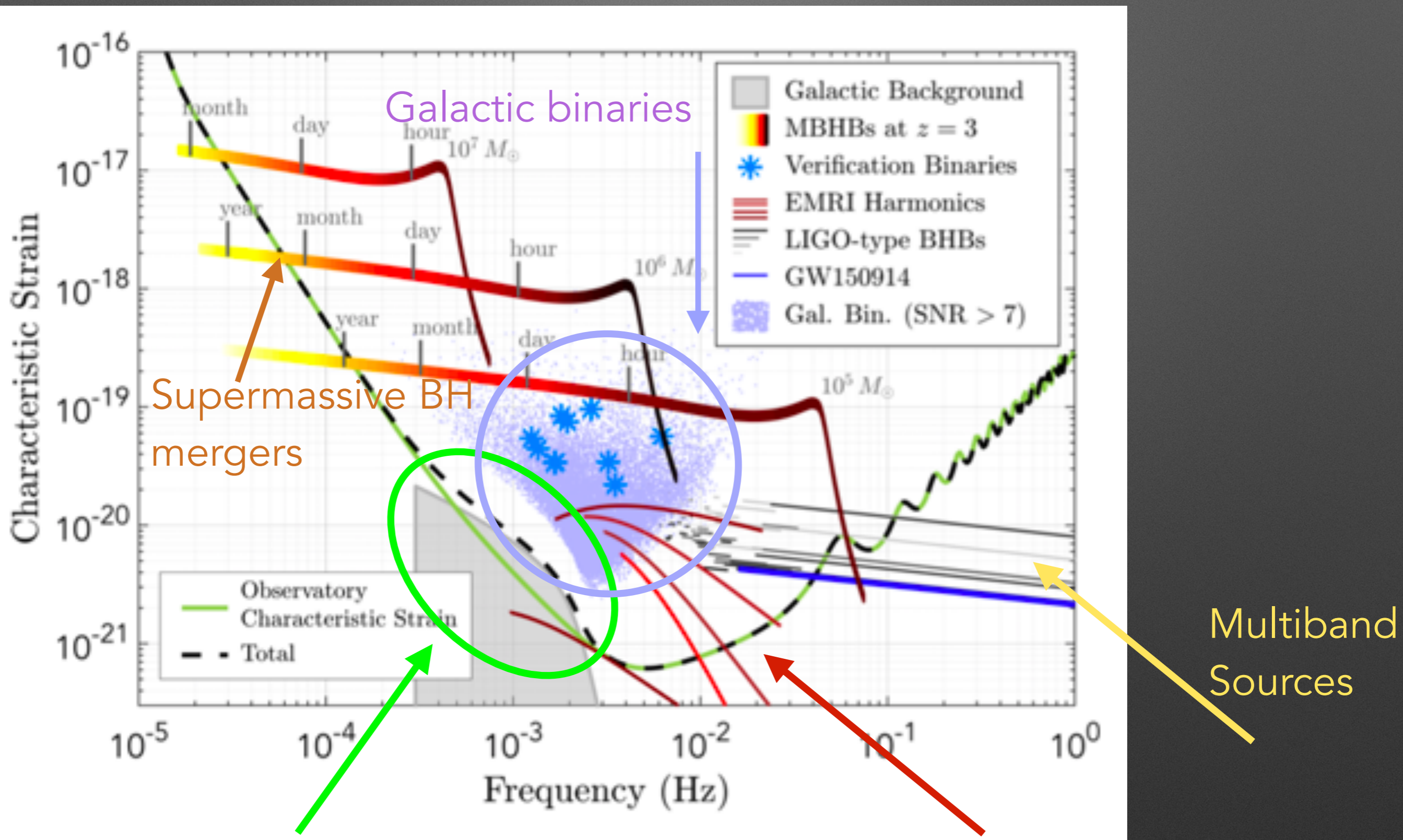


Global Bayesian Fit

Ongoing LISA Data Challenges

Coordinated by France

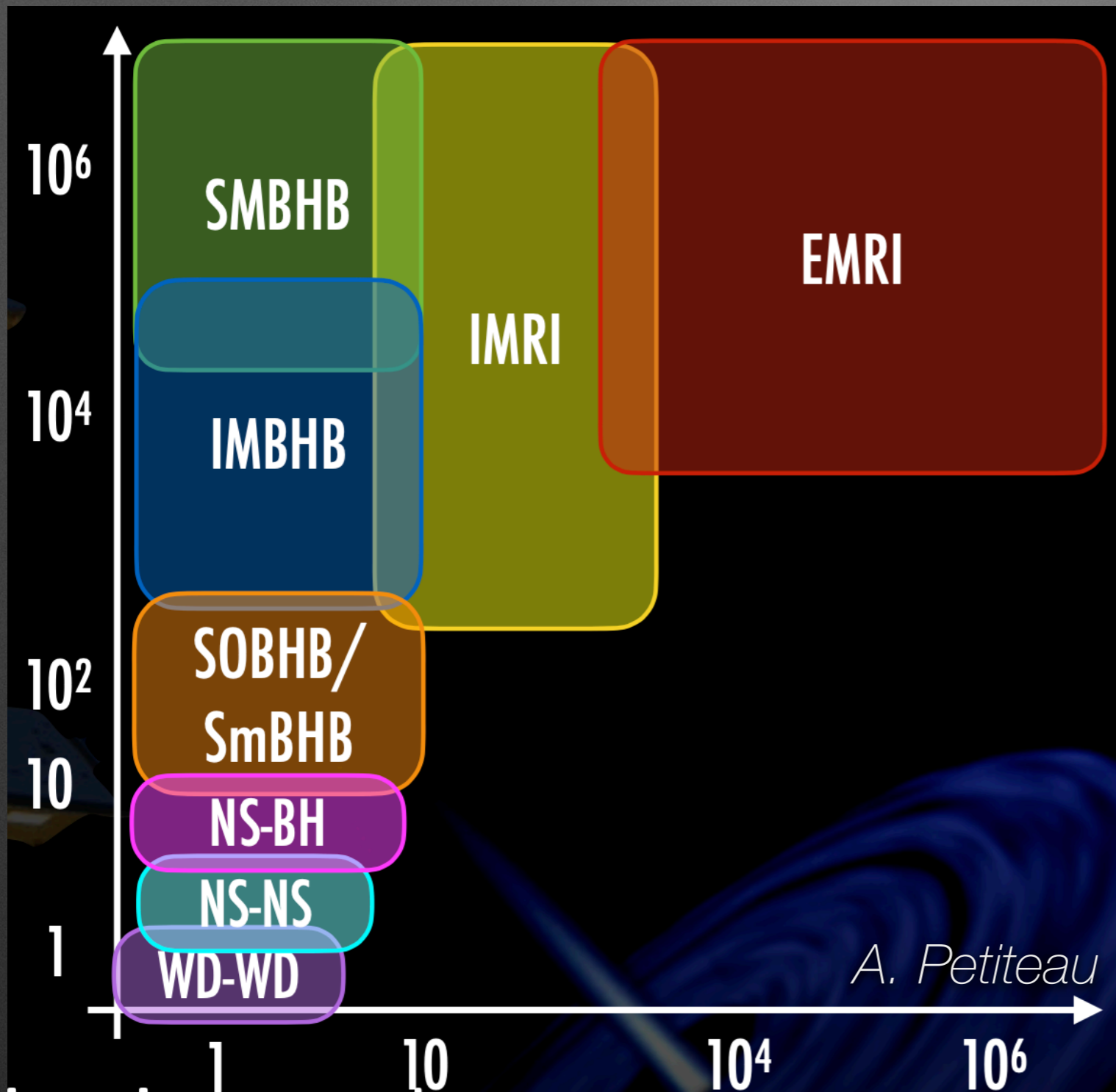
A WIDE VARIETY OF SOURCES



Unresolved sources:
Confusion noise

Extreme Mass Ratio Inspirals

A VERY WIDE RANGE OF COMPACT BINARIES



LISA SCIENCE OBJECTIVES

Astrophysics

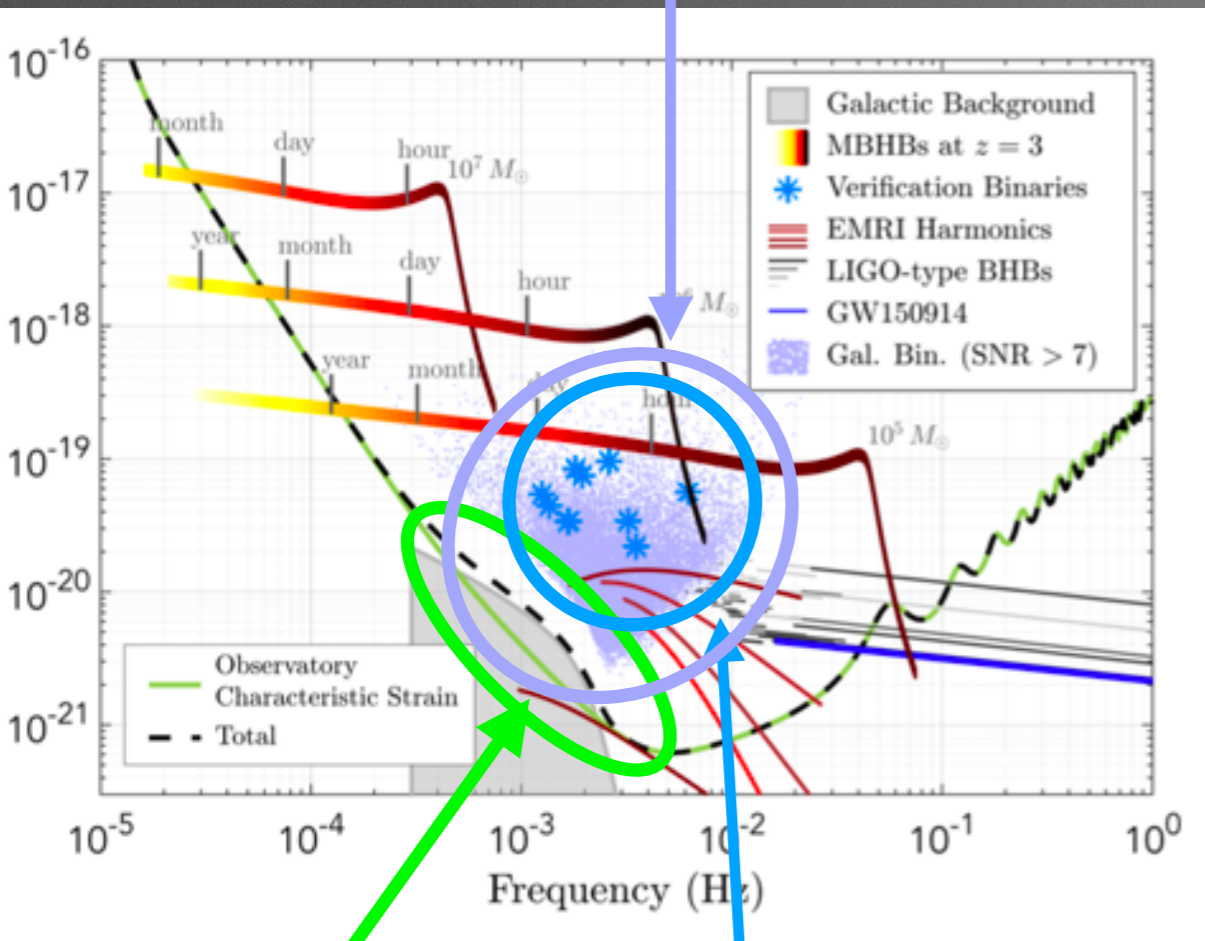
Fundamental Physics

Cosmology

- SO1: Study the formation and evolution of compact binaries in the Milky Way
- SO2: Trace the origin, growth and merger history of massive black holes across cosmic ages
- SO3: Probe the dynamics of dense nuclear clusters using EMRIs
- SO4: Understand the astrophysics of stellar origin black holes
- SO5: Explore the fundamental nature of gravity and black holes
- SO6: Probe the rate of expansion of the Universe
- SO7: Understand stochastic GW backgrounds and their implications for the early Universe and TeV-scale particle physics
- SO8: Search for GW bursts and unforeseen sources

GALACTIC SOURCES

White dwarf binaries



Unresolved sources:
Confusion noise

Verification
binaries

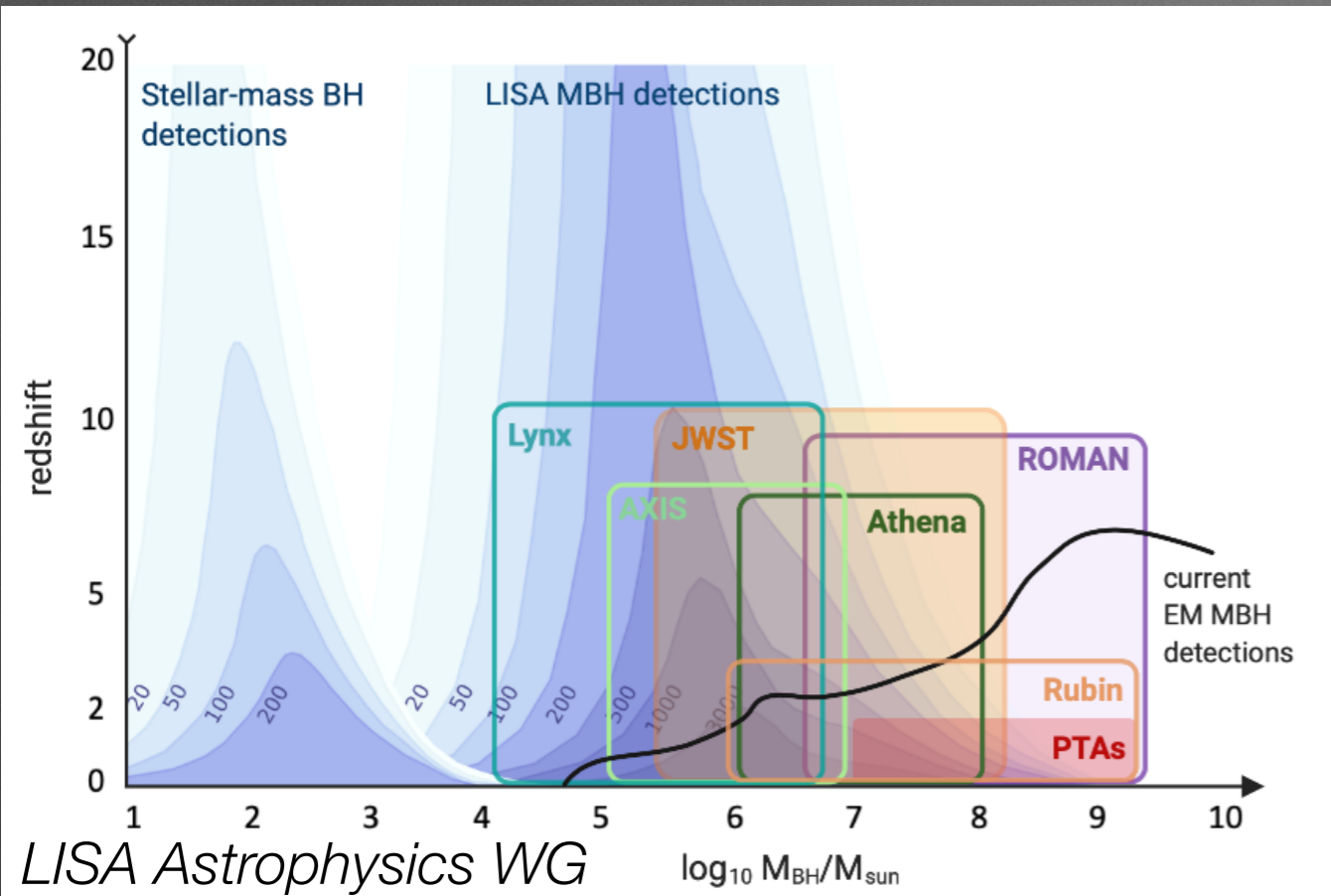
Tracing stellar remnants in the Milky Way

$\sim 0 \cdot 10^4$ resolved
white dwarf binaries -> guaranteed
sources

Fraction with mass measurements
And/or « good » sky localisation ->
multi-messenger sources

Some binaries with neutron stars, black
holes, and/or « living stars », but will be
hard to identify

SUPERMASSIVE BLACK HOLES

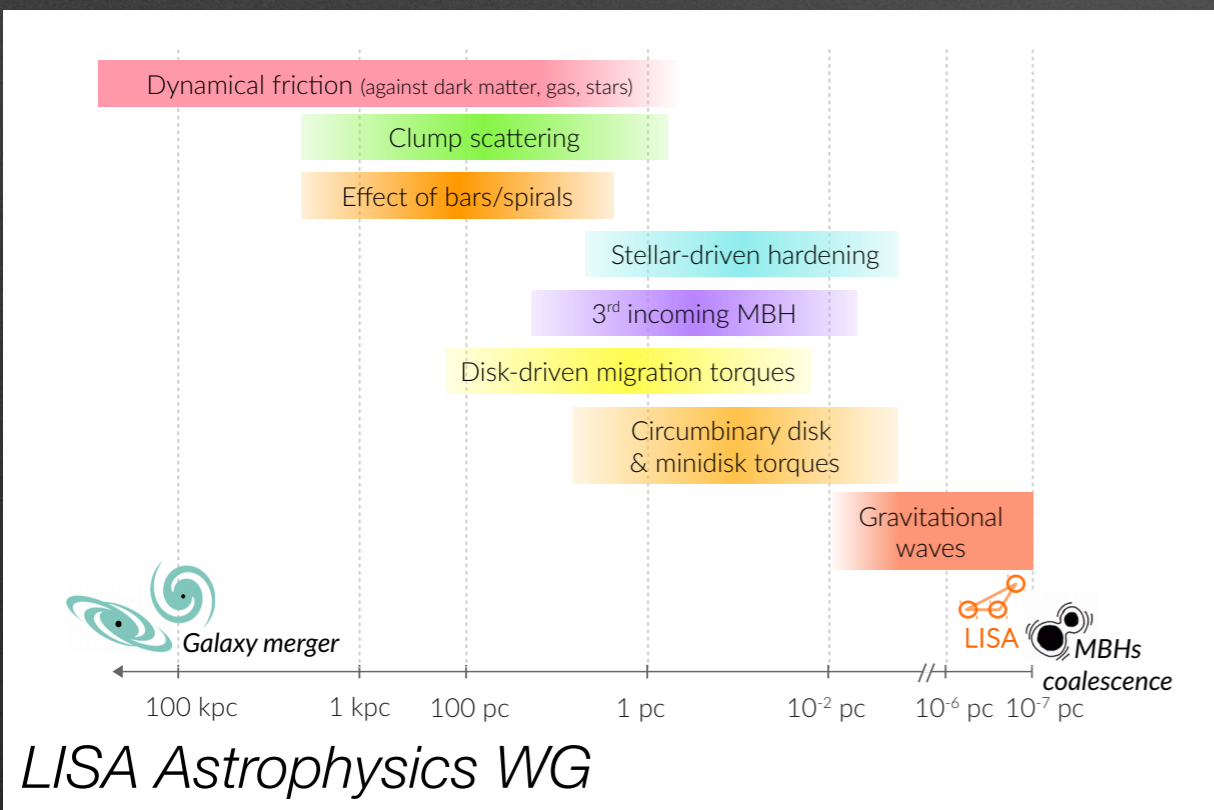


Understand formation and growth of supermassive black holes over cosmic history

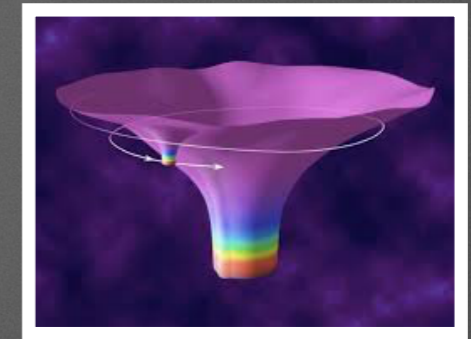
Inspiral (up to months before), merger and bringdown

Likely multi messenger sources - > low-latency EM follow-up

Major uncertainties: Event Rate hard to predict: 10-100 year



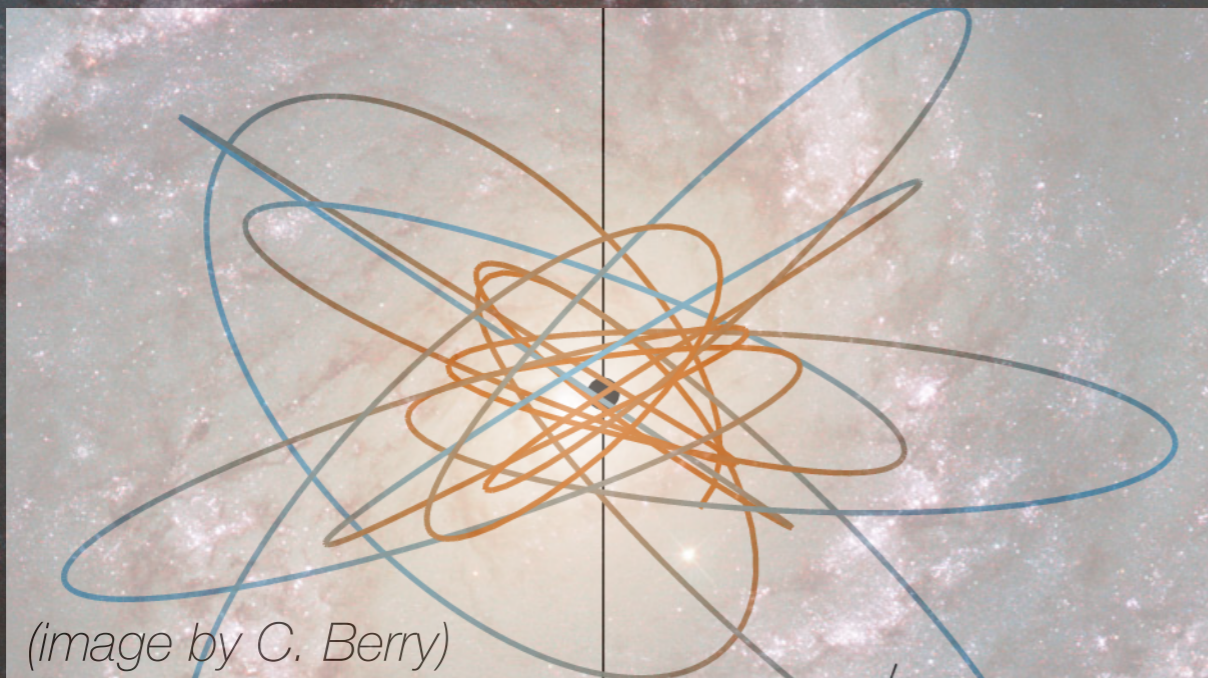
EXTREME MASS RATIO INSPIRALS



Stellar black hole
around supermassive
black hole
-> probe gravity under
extreme conditions

Very complex sources
in terms of waveforms

Very uncertain rates
(few/yr -> 100s /yr)



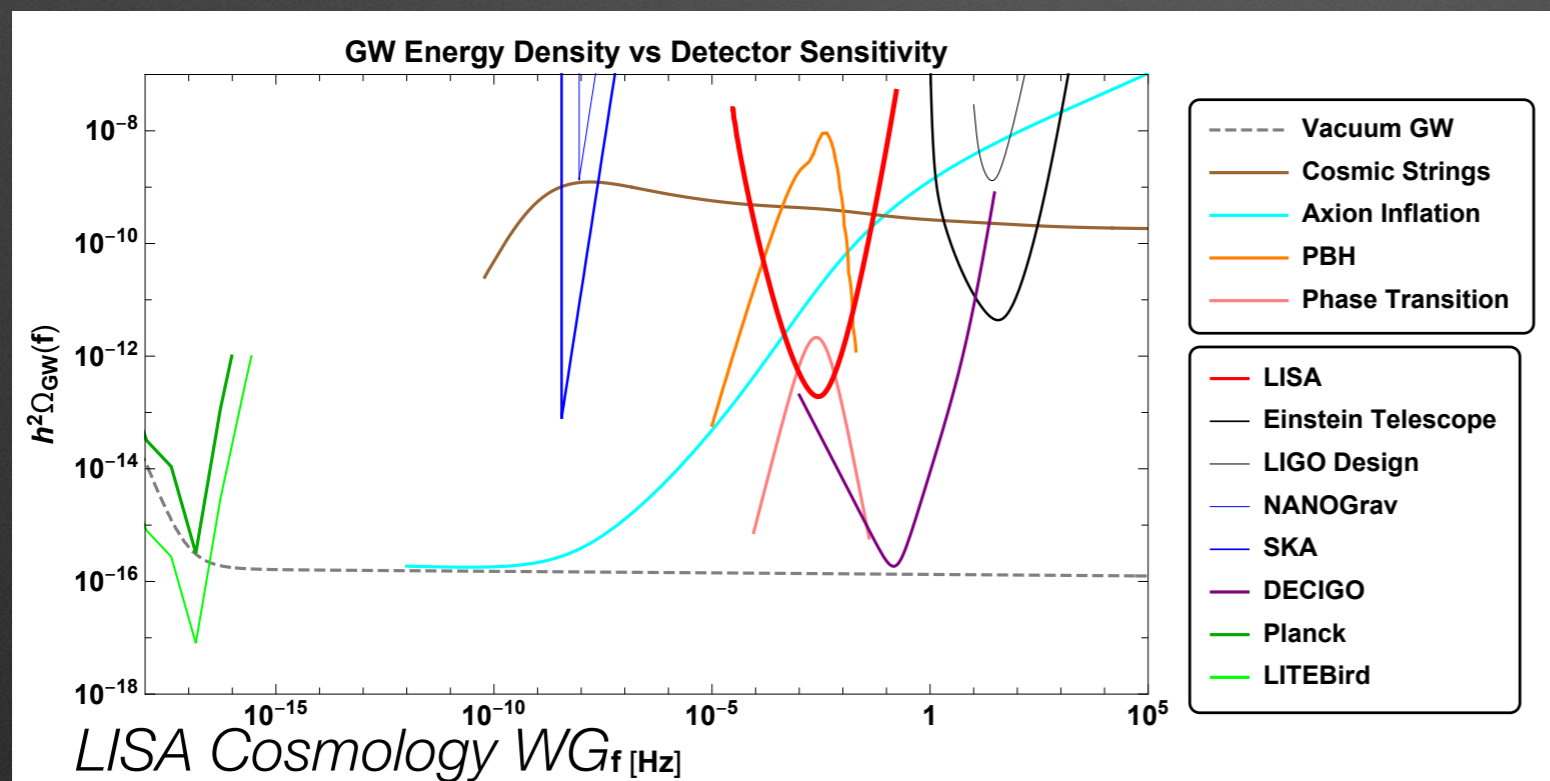
(image by C. Berry)

COSMOLOGICAL BACKGROUND(S)

Probe regions beyond EM and neutrinos

Connect to constraints from PTA and kHz detectors

Need to disentangle from astro background/foreground



Recent White Papers & extra information

- Living Reviews in Relativity
- Astrophysics with the Laser Interferometer Space Antenna (Amaro-Seoane+22)
- Cosmology with the ... (Auclair+22)
- New Horizons for fundamental Physics with LISA (Arun+22)
- Waveforms (in prep.)
- Data Challenge (in prep.)
- <https://lisa-ldc.lal.in2p3.fr/challenge2a>
- elisascience.org
- Consortium with Full Members and Associate Members

GRAVITATIONAL WAVE OBSERVATIONS IN SPACE

LISA: ESA L3 mission

Expected to launch in 2035, expected adoption next year

Opening of a new frequency window of GWs

Major impact in astrophysics, cosmology and fundamental physics

Main sources:

Supermassive black hole mergers at cosmological distances

White dwarf binary inspirals in the Milky Way

Extreme mass ratio inspirals

Foregrounds and background