Novel ways to Probe Dark Sector at Colliders

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4th World Summit on Exploring The Dark Side of the Universe

7-11 novembre, 2022 La Réunion, France





Prologue

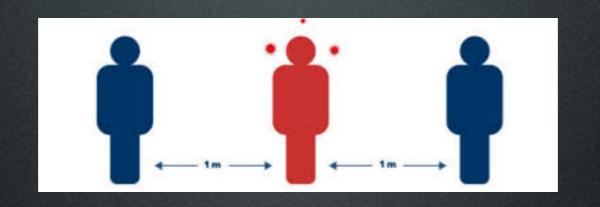
Novel?

Unexplored signatures ignored so far ...

All we see is this

.... whereas this may be hiding!

Principle of (social) distancing in object reconstruction!



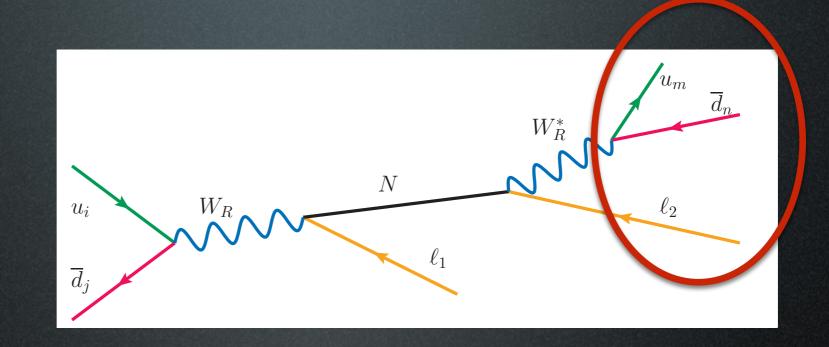
we followed it before it was cool ...

Principle of (social) distancing in object reconstruction!

- Object reconstruction algorithms run independent of one another
- Same detector signature can result in multiple objects being reconstructed, results in fakes!
- Electrons as jets, and vice versa (jets contain neutral pions!)
- Overlap removal to address the double counting

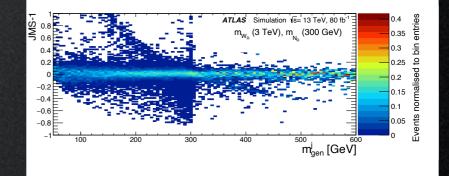
But who ordered that?

Covered in the talk By Antonia Struebig on Monday



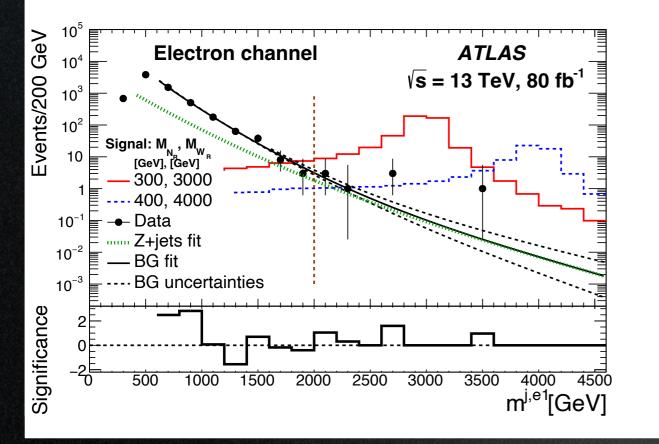
Boosted heavy neutrino search: electron in a large-radius jet

In ATLAS electron reconstruction assumed no nearby real jet, and applies implicit isolation requirement. That reduces signal efficiency, and the presence of such a jet affects the electron performance numbers 7

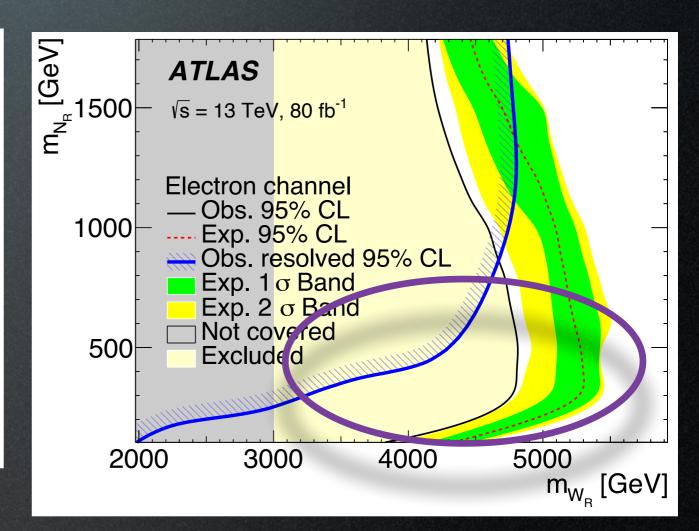


JMS well modelled

Boosted Heavy Neutrino Search



Very small background due to the extreme topology



Complementary strength from resolved analysis

Why stop at electrons?



Here's a llama There's a llama And another little llama Fuzzy Llama Funny Llama Llama Llama duck

Half a llama Twice a llama Not a llama

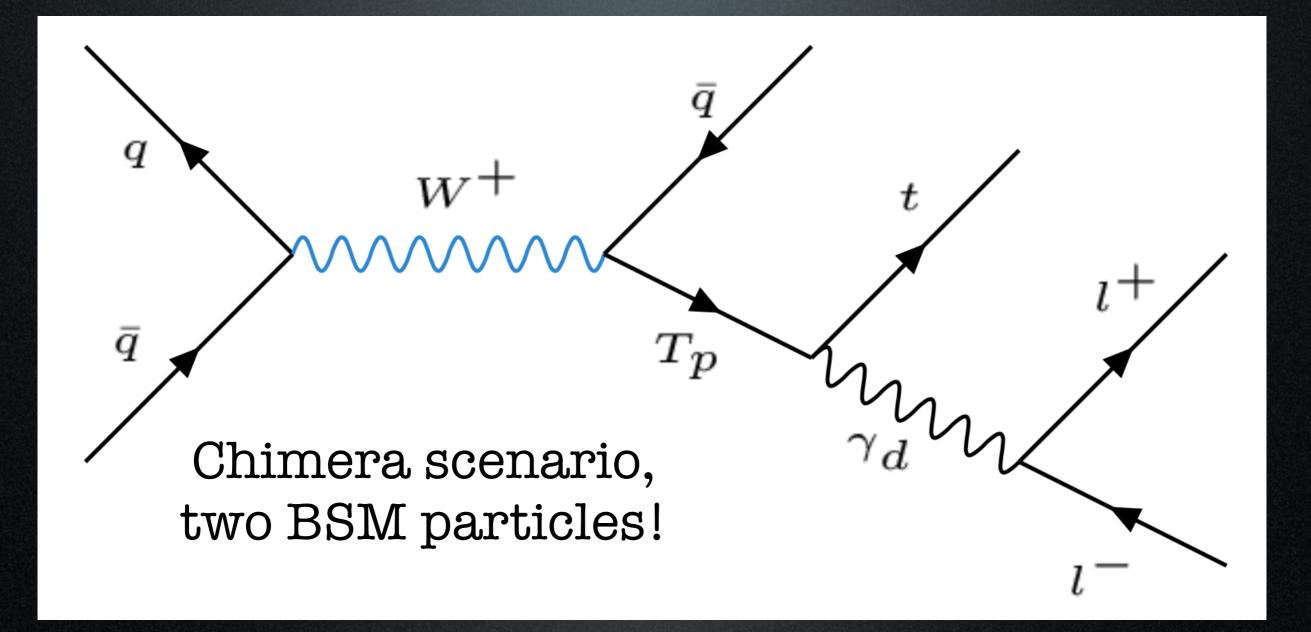


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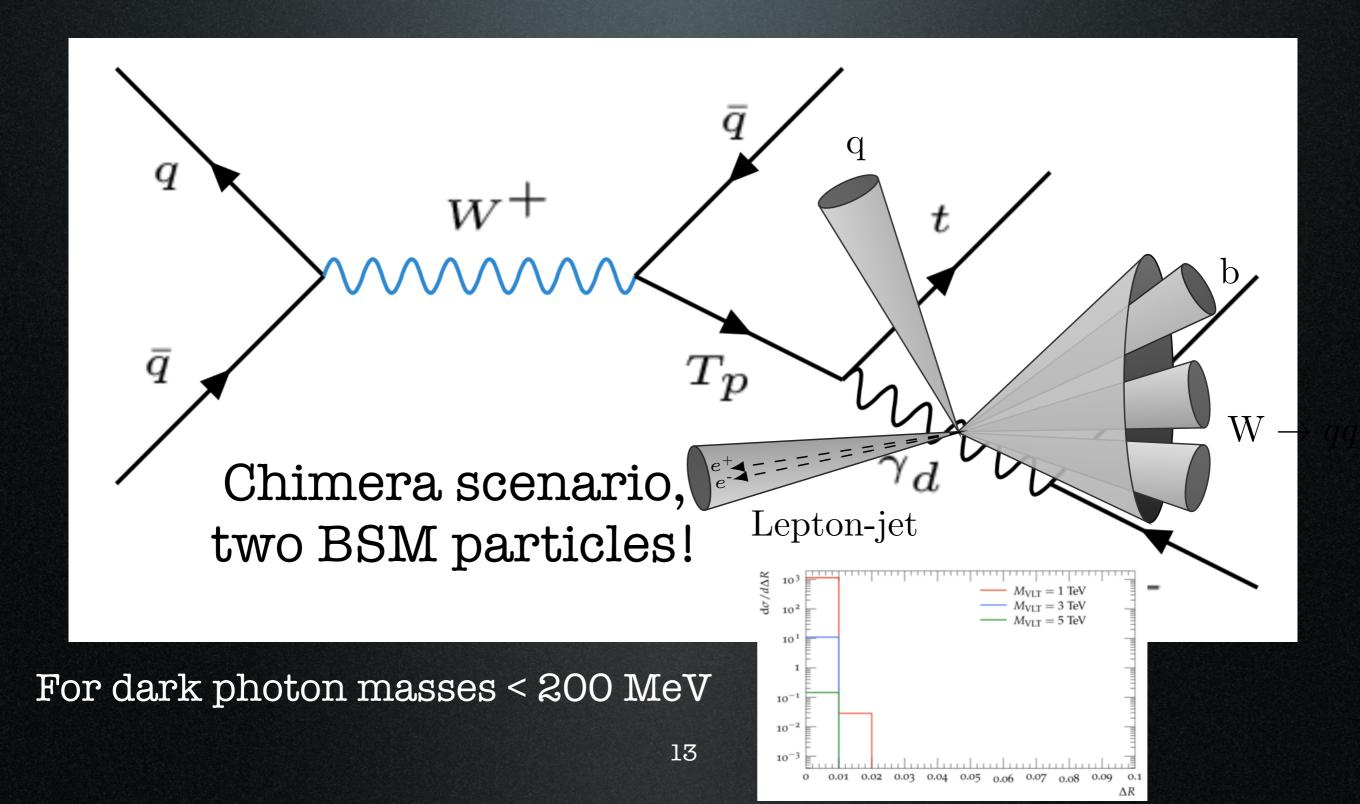
Lepton jet from dark photon

... with Karien du Plessis, M. Flores, S. Sinha, and H. van der Schyf, SciPost Phys. 13, 018 (2022)



SciPost Phys. 13, 018 (2022)

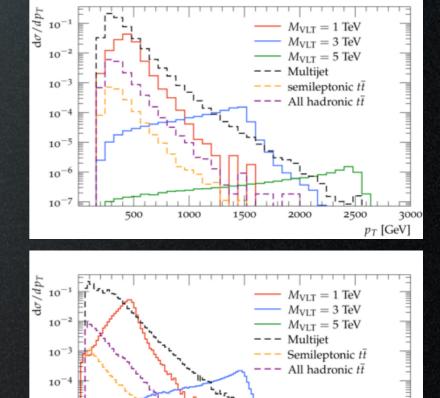
Lepton jet from dark photon



ciPost Phys. 13, 018 (2022)

Search Strategy (hadronic channel)

- Light lepton jet and heavy top jet, both boosted!
- Electron multiplicity -> misleading
- Cannot reconstruct the lepton jet mass
- Largest background: multijet



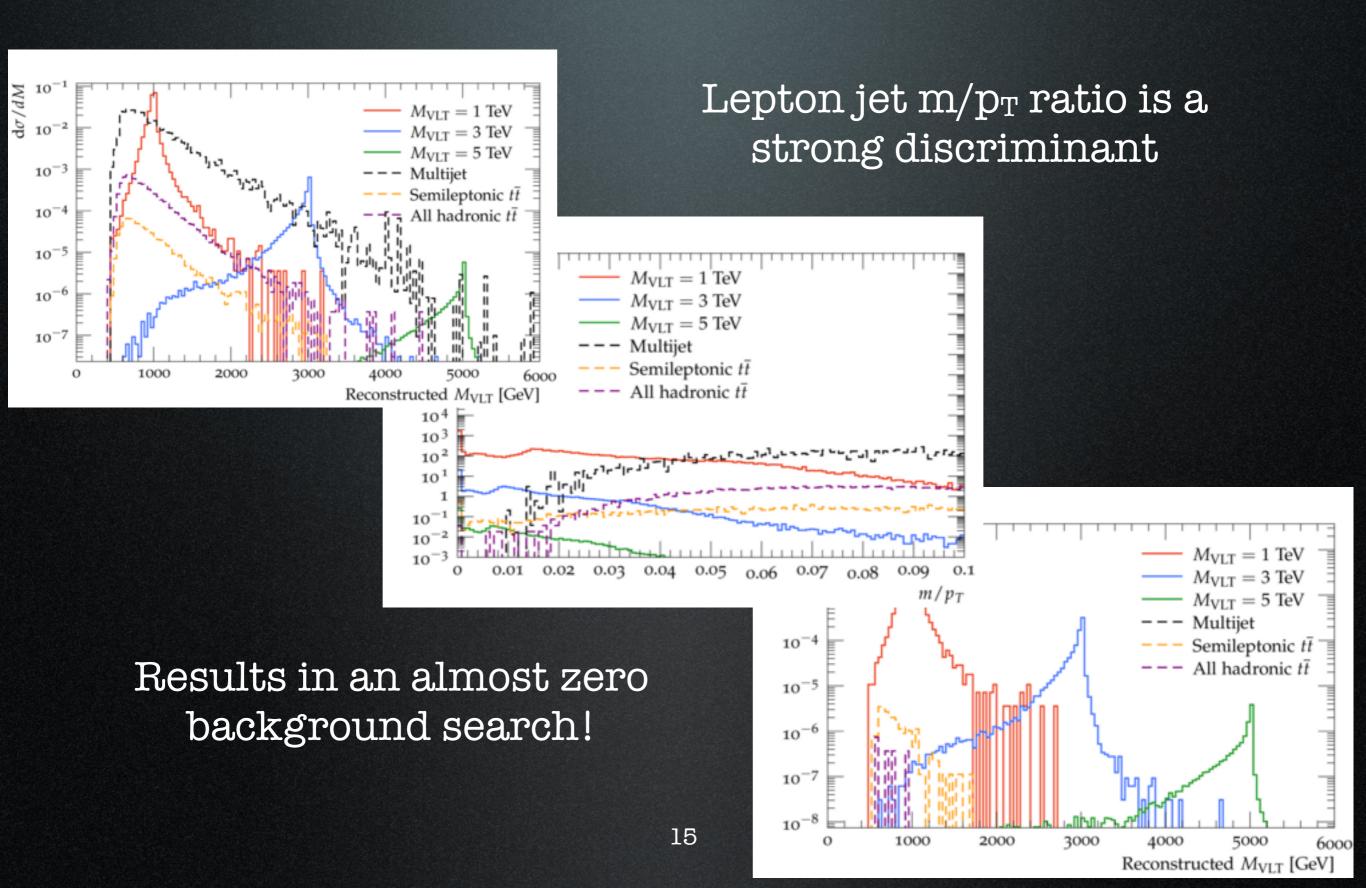
0 3000 p_T [GeV]

 10^{-5}

Not covering the lepton channel result, but mostly same considerations apply

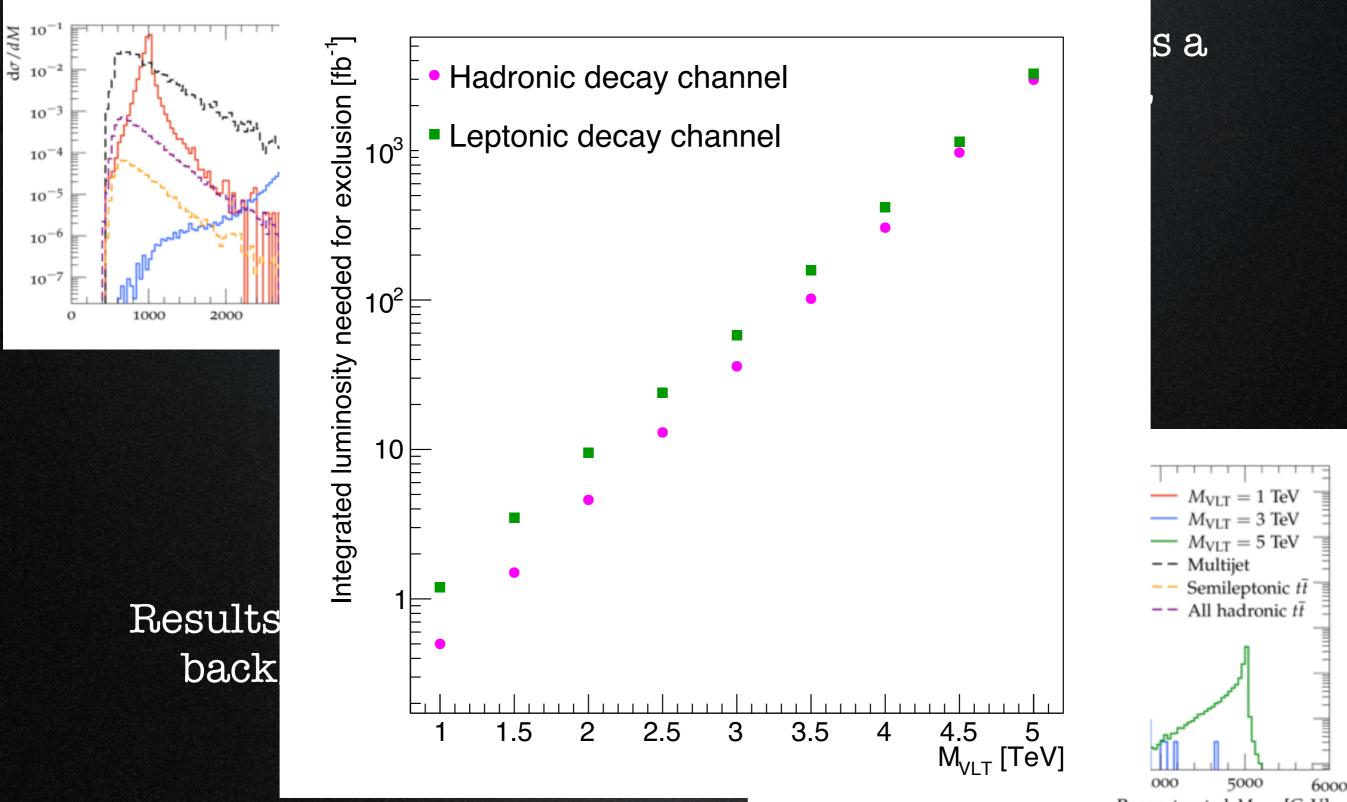
SciPost Phys. 13, 018 (2022)

Then what?



SciPost Phys. 13, 018 (2022)

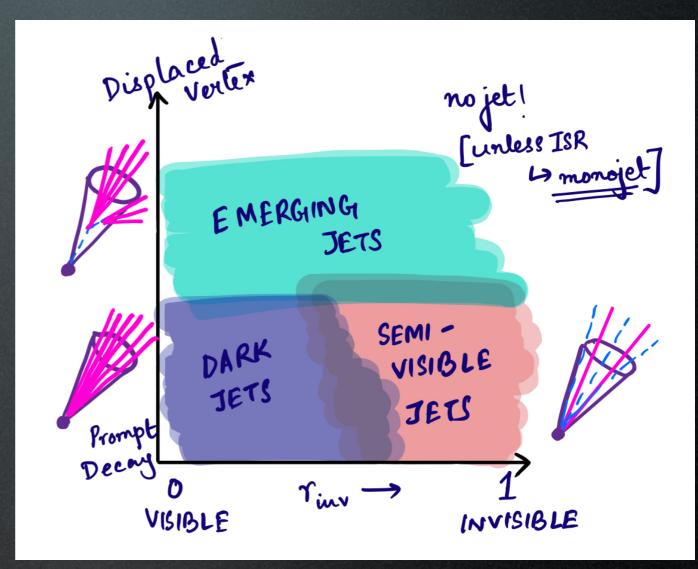
Then what?



Reconstructed M_{VLT} [GeV]

Dark and semi-visible jets

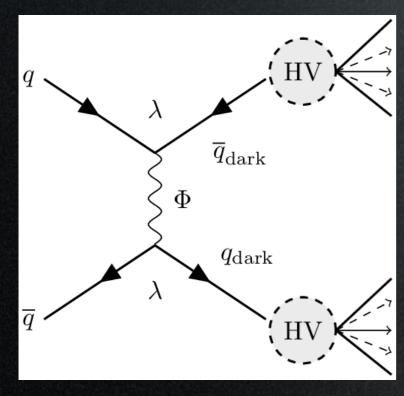




ATLAS result covered in the talk by Sukanya Sinha on Wednesday: ATLAS-CONF-2022-038 Dark hadrons decaying in a QCD-like fashion, fully (dark jets) or partially back to visible sector (semi-visible jets, based on Cohen et al)

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A quick detour: the topology and the challenges for SVJ



ATLAS-CONF-2022-038

Same fraction of dark hadrons In each jet

Why any MET?

A quick detour: the topology and the challenges



A real event will look like this!

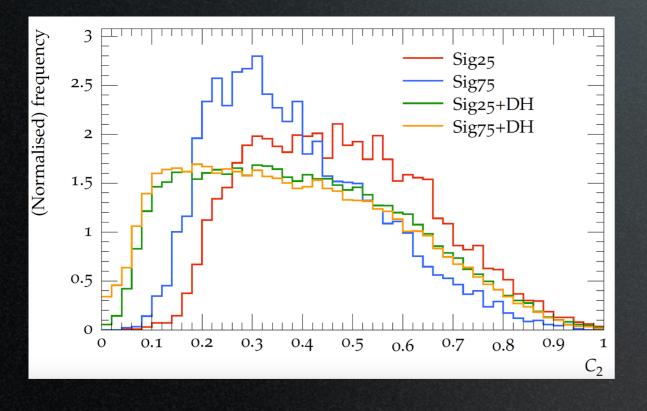
Quantum fluctuations, and boost by extra jets

Therefore **MET**

Not detector noise? Not an easy answer, but check mismodelling in different multijet VRs of lower H_T or MET thresholds

Exploratory study using jet substructure observables

.. with S. Sinha, SciPost Phys. 10, 084 (2021)



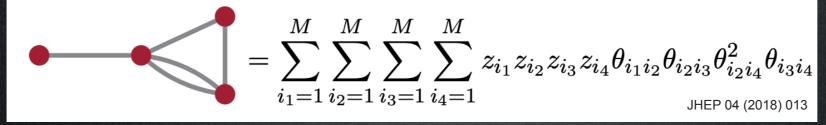
Substructure becomes less twopronged with visible and dark hadrons in them, and the absence of the dark hadrons create the two-pronged structure → The substructure is created by the interspersing of visible hadrons with dark hadrons.

Can use ML algorithms, jet images ...

Use of Energy Flow Polynomials

... with A. Buckley and S. Sinha, arXiv:2209.14964

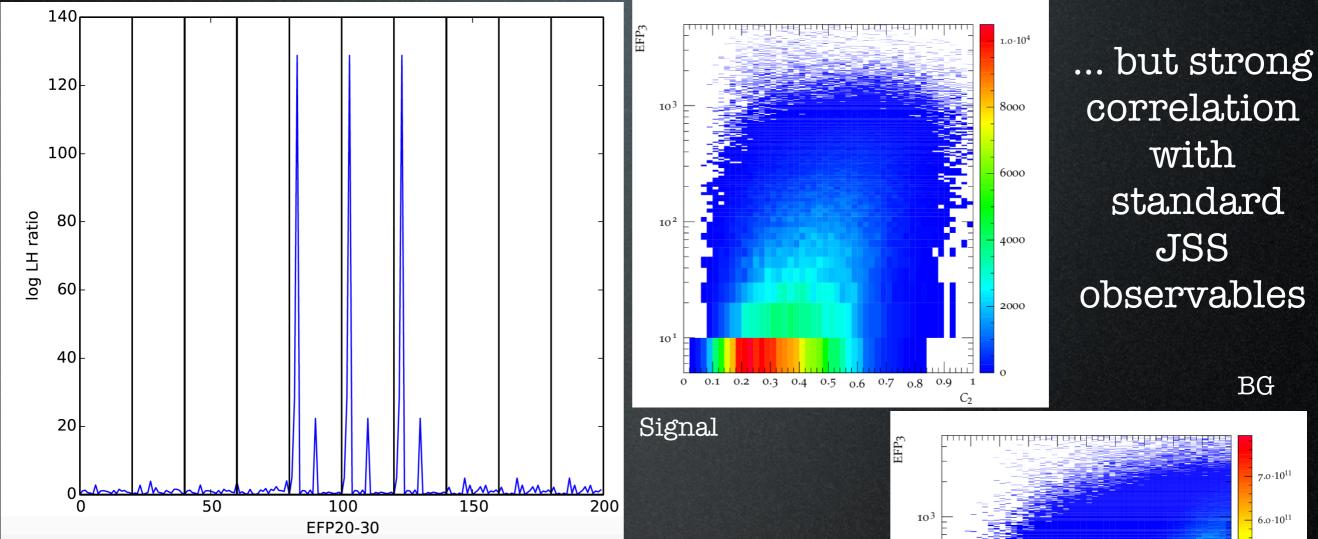
• A complete IRC-safe linear basis:



- Use some sensible truncation scheme
- Check if they offer any discriminating power (and compare to usual JSS observables)
- ... and possibly construct new (JSS-like) observables!

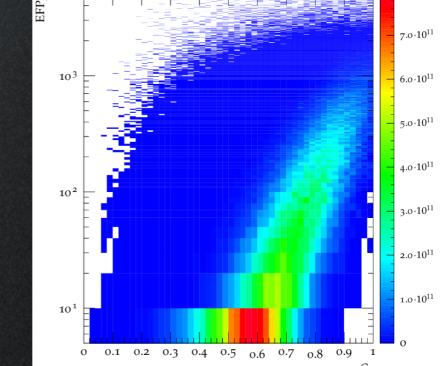
arXiv:2209.14964

Use of Energy Flow Polynomials



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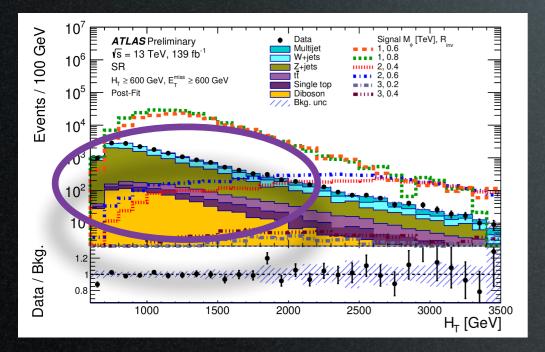
Certain EFP diagrams seem to have some bins that multijet background does not populate at all, in which the SVJ signal dominates...



SVJ with Heavy Flavour

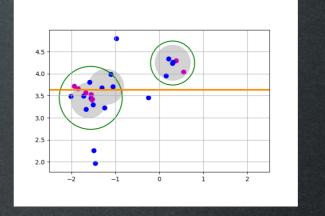
with S. Sinha, arXiv:2207.01885

ATLAS-CONF-2022-038

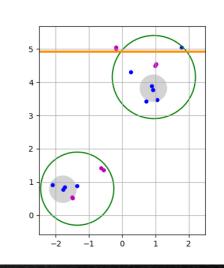


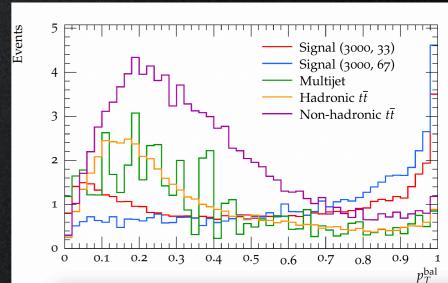
Can we reduce non dominant backgrounds?

What if SVJ is produced exclusively with b-jets? Turns out it is a theoretically well motivated scenario! Better handle on identifying/ reconstructing the SVJ!



-> Better reconstruction with Variable Radius jets (JHEP 0906:059,2009)





Summary

We have this beautiful LHC dataset, with potential doubling of integrated luminosity in a few years.

Let us explore the unexplored, the dark stuff hidden in plain day light!

Supporting Material



Over all constituents (beta: angular exponent):

$$\begin{aligned} \operatorname{ECF}(1,\beta) &= \sum_{i} p_{Ti} \\ \operatorname{ECF}(2,\beta) &= \sum_{i < j} p_{Ti} \, p_{Tj} \, (R_{ij})^{\beta} \, \leftarrow \, \begin{bmatrix} \operatorname{see Banfi, Salam, Zanderighi;} \\ \operatorname{Jankowiak, Larkoski} \end{bmatrix} \\ \operatorname{ECF}(3,\beta) &= \sum_{i < j < k} p_{Ti} \, p_{Tj} \, p_{Tk} \, (R_{ij} R_{jk} R_{ki})^{\beta} \\ \operatorname{ECF}(N,\beta) &= \sum_{\operatorname{sets of } N} (N \, \operatorname{energies}) \times \left(\begin{pmatrix} N \\ 2 \end{pmatrix} \operatorname{angles} \right)^{\beta} \end{aligned}$$

ECF(N+1) << ECF(N) for N subjets

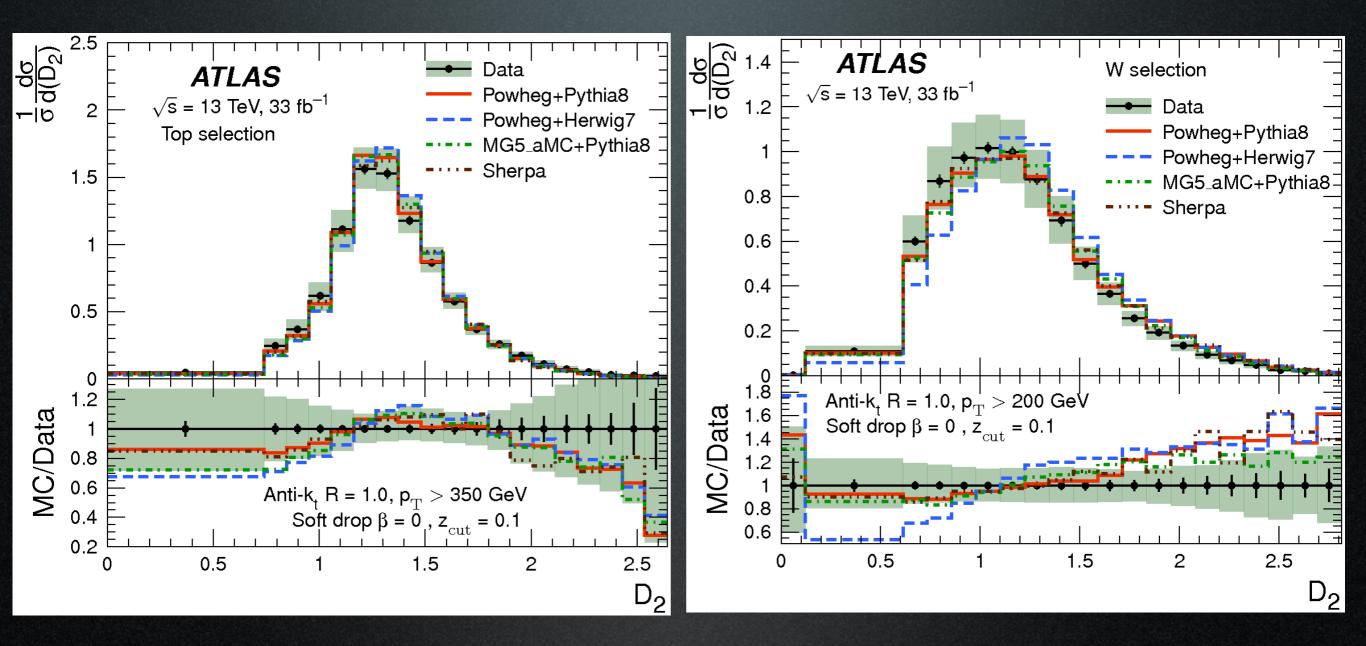
Define (double) ratio = [ECF(N+1)/ECF(N)]/[ECF(N)/ECF(N-1)]

$$C_N^{(\beta)} = \frac{\mathrm{ECF}(N+1,\beta) \mathrm{ECF}(N-1,\beta)}{\mathrm{ECF}(N,\beta)^2}$$

Analogous to Nsubjettiness ratio

Large C_N: more than N subjets, extra radiation is not correlated with leading order N subjets. For small C_N: the additional radiation is soft/collinear





Again shifted peak in W, models overestimating gluon radiation

How to compute EFPs?

$$ullet_{j} \iff \sum_{i_{j}=1}^{M} z_{i_{j}} \quad k$$
 ———— $\ell \iff heta_{i_{k}i_{\ell}}$

Each edge (k,l) in a multigraph is in one-to-one correspondence with a term θ in an angular monomial

Each vertex j in the multigraph corresponds to a factor of z and summation over i, in the EFP

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$$\bullet = \sum_{i_1=1}^{M} \sum_{i_2=1}^{M} \sum_{i_3=1}^{M} \sum_{i_4=1}^{M} \sum_{i_4=1}^{M} z_{i_1} z_{i_2} z_{i_3} z_{i_4} \theta_{i_1 i_2} \theta_{i_2 i_3} \theta_{i_2 i_4}^2 \theta_{i_3 i_4}$$

4 particles/constituents in a jet ---> 4 energy fractions, 5 angularity values ----> degree 5 polynomial

Because the EFP basis is infinite, a suitable organization and truncation scheme is necessary to use the basis in practice.

	d	1	2	3	4	5	6	7	8	9	10
	2	1	1	1	1	1	1	1	1	1	1
	3		1	2	3	4	6	7	9	11	13
	4			2	5	11	22	37	61	95	141
	5				3	11	34	85	193	396	771
$ _N$	6					6	29	110	348	969	2445
	7						11	70	339	1318	4457
	8							23	185	1067	4940
	9								47	479	3294
	10									106	1279
	11										235

How to compute EFPs?

Several combinations for diagrams possible

How to compute EFPs?

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Several combinations for diagrams possible

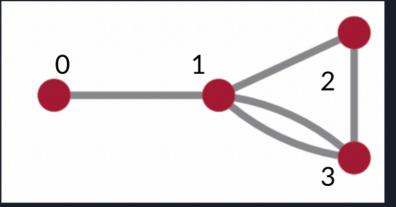
Restricting ourselves to this regime...

The linked paper discusses in great detail how different EFP combinations lead to well-known jss observables

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What we ideally want to achieve with EFPs?

- Implement EFPs in Rivet and see if any particular combination of EFPs helps to distinguish between standard q/g jets and more unconventional jets
 - Might lead to a new jet-substructure observable for dark shower discrimination
- Setup working (computing EFP multigraphs till N = 7, d = N 1, N, N + 1) → python code taking into account the different possible orientations of the input "particles" and designing an array of possible EFP diagrams as a grid



EFP diagram to "particle" pair translation

 $\{\{0, 1\}, \{1, 2\}, \{2, 3\}, \{1, 3\}, \{1, 3\}\}$

N=4,d = N - 2	0	1	2	3
0	-	-	-	-
1	1	-	1	2
2	-	-	-	-
3	-	-	1	-